THE THEORY OF RELATIVITY

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## THE THEORY OF RELATIVITY

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## PREFACE

This introduction to the Theory of Relativity is based in part upon a course of lectures delivered in University College, London, 1912-13. The treatment, however, has been made much more systematical, and the subject matter has been extended very considerably; but, throughout, the attempt has been made to confine the reader's attention to matters of prime importance. With this aim in view, many particular problems even of great interest have not been touched upon. On the other hand, it seemed advantageous to trace the connexion of the modern theory with the theories and ideas that preceded it. And the first three chapters, therefore, are devoted to the fundamental ideas of space and time underlying classical physics, and to the electromagnetic theories of Maxwell, Hertz-Heaviside and Lorentz, from the last of which Einstein's theory of relativity was directly derived. In the exposition of the theory itself free use has been made not only of the matrix method of representation employed by Minkowski, but even more of the language of quaternions. Very little indeed of these mathematical methods is required to follow the exposition, and this little is given in Chapter V., in a form which will be at once accessible to those acquainted with the elements of the ordinary vector algebra.

It is hoped that the book will give the reader a good insight into the spirit of the theory and will enable him easily to follow the more subtle and extended developments to be found in a large number of special papers by various investigators.

I gladly take the opportunity of expressing my thanks to Mr . Willian Francis and Dr. T. Percy Nunn for their kindness in reading a large portion of the MS., to Prof. Alfred W. Porter, F.R.S., for reading all the proofs and for many valuable suggestions, and to the Publishers and the Printers for the care they have bestowed on my work.
L. S.

Lonion, Atril, 1914.

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## CHAPTER I.

## OLASSICAL RELATIVITY.

Before entering upon the subject proper of this volume, namely, the modern doctrine of Relativity and the history of its origin and developsment, it seems desirable to dwell a little on the more familiar ground of what might be called the classical relativity, and to eonsider two particular points which are of fundamental importance, not only for the appreciation of the whole subject to follow, but also for an adequate understanding of almost all physico-mathematical considerations. What I am alluding to are the following questions: $x^{\prime \prime}$ the choice of a framework of axes or, more generally, of a system of reference in space, and $2^{4}$ the definition of physical time, or the selection of a clock or time-keeper, to be empleyed for the quantitative determination of a succession of physical events.

Both of these questions existed and were solved, at least implicitly, a long time before the invention of the noodern Principle of Redativity, in fact centurics ago, in their essence as carly as Copernicus founded his system.*

The question of a spacc-framework is obvious enough and widely known ; it will require therefore only a few simple remarks.

The most superficial observation of everyday life would suffiee to show that the form and the degree of simplicity of the statement of the laws of physical phenomena, more cspecially of the laws of motion of what are called material bedics, depend essentially on our selece. tion of a system of reference in space. Certain frameworks of reference are peculiarly fitted for the representation of a particular

[^0]instance of motion of a particular body or also of almost any observable motion of bodies in general, leading to a high degree of completeness, exactness and simplicity, while other frameworks (moving in an arbitrary manner relatively to those) give of the same phenomena a most complicated, intricate and confused picture.*
Suppose that somebody, ignorant of the work of Copernicus, Galileo and Newton, but otherwise gifted with the highest experimental abilities and mathematical skill (a quite imaginary supposition, being hardly consistent with the first one), chooses the interior of an old-fashioned coach, driven along a fairly rough road, as his laboratory and tries to investigate the laws of motion of bodies enclosed together with him in the coach-say, of a pendulum or of a spinning top-and selects that vehicle as his system of reference. Then his tangible bodies and his conceptual 'material points,' starting from rest or any given velocity, would describe the most wonderful paths, in incessant shocks and jerky motions; the axis of his' 'free gyroscope' would oscillate in a most complicated way, -never disclosing to him the constancy of the vector known to us as the 'angular momentum,' i.e. the rotatory. analogue of Newton's first law of motion. Nor would the uniform translational motion have for him any peculiarly simple or generally noteworthy properties at all. His mechanical experience being, in a word, full of surprises, he would soon give up his task of stating any laws of motion whatever with reference to the coach. Getting out of it on to firm ground, he will readily find out that the earth is a much better system of reference. With this framework, smoothness and simplicity will take the place of hopeless irregularity. Undoubtedly, this property must have been remarked in a very early stage of man's history, and the above example will appear to the least trained student of mechanics of our present times trivial and simply ridiculous. 'Of course,' he would say, 'the motions of material bodies relatively to that coach are so very complicated, for that vehicle is itself moving in a highly complicated way.' He would hardly consider it worth while to add 'relatively to the earth.' The coach being such a small, insignificant thing in comparison with the terrestrial globe, it would seem extravagant to our interlocutor, if somebody insisted rather on saying that it is the earth which moves in such a complicated way relatively

* And as to 'absolute motion,' regardiess of any system of reference, it is needless to mention that it is devoid of meaning in exactly the same way as 'absolute position.'
very feeble argument (as we shall see presently, from another example).

At any rate the earth, the 'firm ground,' allowance being made for occasional large shocks and for very small but incessant oscillations of every part of its surface,* has proved to be an excellent system of reference for almost all motions, especially those on a small scale with regard to space and time, and practically without any reservation for all pieces of machinery and technical contrivance. In fact, the earth as a system of reference offered at once the advantage of a high degree of simplicity of description of states of equilibrium and motion, opening a wide field for the application of Newton's mechanics, at least as regards purely terrestrial observations and experiments. $\dagger$ The earth is then a reference-system which is constantly used also by the most advanced modern student of mechanics.

But things become altogether different when we look up to the sky and desire to bring into our mechanical scheme also the motions of those luminous points, the celestial bodies, including, of course, our satellite, the moon, and our sun. Then the earth loses its privilege as a framework of reference. If it were only for the serealled 'fixed stars,' which form the enormous majority of those lutuinots points (and the moon too), we could still satisfy our vanity and continue to consider our globe as an miversal mechanical system of reference, the system of reference, as it were. On our plane drawings, or in our three-dimensional models, we could then represent the earth by a fixed disc, or sphere, respectively, with a smaller sphere moving round it in a circular orbit, to imitate our moon, the whole surrounded by a large spherical shell of glass sown with millions of tiny stars, spinning gently and uniformly round the carth's axis,--very

[^1]much like, in fact, some primitive mental pictures of the universe.* But the case becomes entirely different when we come to consider the far less numerous class of luminous points or little dises, the planets, and the comets, moving visilly among the 'lixed' shining points in a complicated way. Then, even before touching any dynamical part of the celestial problem, we are compelled to give up our earth as a system of reference and replace it by that of the 'fixed stars,' originally so inconspicuous, or-what turns out to be equally goodby a framework of axes pointing from an initial Ioint fixed in the sun towards any given triad of fixed stars. It is needless to tell here again the long story of that admirable and ingenious system which was founded by Peolemy (born about 40 B.e.:), which hold the field during fourteen centuries, to be replaced fually and definitely by the system of Copernicus ( $1473+1543$ ), which transferred to the sun the previous dignity of the etrth. $\dagger$ The Copernican system of reference had the enormous advantage of simplicity, quite independently of any mechanical, i.ce. (to put it more strictly) dynamical considerations. Its supuriority to the geocentric system manifested itself already in the simplicity it gave to the paths of the solar family of bodies, the wonderfully simple shapese of the orbits of the planets. In the geocentric srheme we had the complicated system of 'excentrics and epicycles' of I'tolemy, whereas taking, in our drawing or model, the sun as fixect, the orbits of the planets became simple circles, which in the next step of approximation turned out to be slightly elliptic. Thus the Coperniern system of reference had its enormous advantages before any properly mechanical point of the subject was entered upon. Histerically, in fact, the mechanics of Gaileo and Newton came a long time after Copernicus, so that the

[^2]privilege of reference-system was taken away from our earth and transferred to the sun on the ground of purely kinematical considerations of simplicity, a few centuries before Newton. But afterwards the Copernican or the 'fixed-stars' system of reference appeared to be wonderfully appropriate to Newtonian mechanics, both in its original shape and in its later (chiefly formal) development by Laplace for celestial and by Lagrange for terrestrial and general problems. It soon became the final reference-system of mechanics. It is relatively to this 'fixed-stars' system of reference that the law of inertia has proved to be valid. We will call it, therefore, following the modern habit, the inertial system, or somctimes, also, the Nezetonian system of reference.* It is relatively to this system that spinning bodies behave in the characteristically simple manner which has led many authors to speak of their property of 'absolute orientation.' Or, to put it in less obscure words, it is relatively to the inertial system that the vector called angular momentum is preserved, both in sizc and in direction,-this property being a consequence of the fundamental laws of Newton's mechanics, and, at the same time, a perfeet and most instructive analogue to Newton's First Law of motion. $\dagger$ The most immediate and tangible manifestation of this property is that the axis of a free gyroscope (practically coinciding in direction with its angular momentum) points always towards the stume fixed star ; thus having the simplest relation to the inertial system, since it is invariably orientated in this system of reference. Notice that it would, therefore, be more extravagant to say that the axis of such at gyroscope moves relatively to the carth than vice versct,--whough apparently, bodily, the gyroscope of human make is such an inconspicuous tiny thing in comparison with our planct. The conservation of the angular momentum, or moment of momentum, $\mathbf{\Sigma m V r v} \ddagger$ of the whole solar system, which is best known in connexion with Laplace's 'invariable plane,' is but the same thing on a larger seale than that exhibited by our spinning tons. But this only by the

[^3]way. What mainly concerns us here is that the 'fixed-stars' system -or, more rigorously, any one out of the $\infty^{8}$ multitude of equivalent inertial systems-has gradually turned out to be peculiarly fitted as a system of reference for the representation of the motion of material bodies.

But also with this system of reference the laws of motion have their simple, Newtonian form only for a $t$ measured in a certain way, i.e. for a certain clock or time-keeper, e.g. approximately the earth in its diurnal rotation, or, more exactly (in connexion with what is known as the frictional retarding effect of the tides), a time-keeper slightly different from the rotating earth. This is equivalent to defining as equal intervals of time those in which a body not acted on by 'external forces,' i.e. very distant from other bodies or otherwise suspected sources of clisturbance, describes equal paths.* In maintaining the motion of such and such a body in such and such circumstances to be uniform, we do not make a statement, but rather are defining what we strictly mean by equal intervals of time. Selecting quite at random a different time-keeper, we could not, of course, expect the same simple laws to hold, with respect to the inertial system of reference. But with another space-framework of reference another time-keeper might do as well.
Thus we see that, to a certain extent, the choice of a system of reference in space has to be made in conjunction with the selection of a time-keeper. Our $x, y, z, t$, the whole tetrad, the space and time framework must be selected as one whole. That kind of 'union' emphasized by the late Hermann Minkowski, the joint selection of $x, y, z, t$, manifesting itself in the modern relativistic theory by the consideration of a four-dimensional 'world' (instead of time and space, separately), is not altogether such an entirely new and revolutionary idea as is generally believed; for to a certain extent, and in a somewhat different sense, it is as well a requirement of Newtonian mechanics, and, more generally, of the classical kind of Plysics, as of modern Relativity. What difference there really is between the two we shall see in the following chapters.

[^4]Meanwhile we have touched, in passing, the fourth variable $t$, and this brings us to our second point, namely, the definition of physical time, the selection of 'the independent variable $t$ ' of our physico-mathematical equations, but viewed more generally, and more carefully, than above, where we have touched it only incidentally.

To explain this question, of capital importance for almost every quantitative physical research, I must ask you to direct your attention to the following considerations.

Suppose we do not limit ourselves to the investigation of motion only, but are concerned with every possible kind of physical phenomena, such as conduction of heat or electricity, diffusion of liquids or gases, melting of ice, evaporation of a liquid, etc., etc., and that we propose to describe the progress of these phenomena in time, to trace their history, past and future. How are we, then, to select our timequantity $t$ ?

First of all, we cannot, of course, take it to be Newton's 'absolute time,' which is defined, according to a quotation from Maxwell,* as follows :
'Absolute, true, and mathematical Time is conceived by Newton as flowing at a constant rate, unaffected by the speed or slowness of the motions of material things. It is also called Duration.'

For, supposing there is such a thing, $\dagger$ we do not know how to find or to construct a clock which measures this 'absolute time,' even approximately; that is to say, we have no criterion to distinguish such a clock from a 'wrong' one. And thus, certainly, we cannot use this kind of definition for physical purposes. How are we then to measure our $t$ ? Granting that the selection of a chronometer indicating our $t$ is (at least within certain wide limits) arbitrary or free, what is the requirement on which we have to base our choice?

Now, it seems to me that the first and most general requirement, which may also be seen to be tacitly assumed in all the investigations of both the more recent and classical natural philosophers, especially physicists and astronomers, is
that our differential equations, representing the laws of physical (and other) phenomena, should not contain the time, the variable $t$, explicitly,

* Matter and Motion, page 19.
+ But, as a matter of fact, the phrase 'flowing at a constant rate' is simply meaningless.

i.e. . that for "any sufficiently comprehensive physical system, of which the instantaneous state is defined, say, by $p_{1}, p_{2}, \ldots p_{n}$, the differential equations shoridl be of the form

$$
\left.\begin{array}{c}
\frac{d p_{i}}{d t}=f_{i}\left(p_{1}, p_{2}, \ldots p_{n}\right),  \tag{A}\\
i=1,2, \ldots n .
\end{array}\right\}
$$

This requirement is also intimately connected with a certain form of what Maxwell* calls 'the General Maxim of Physical Science' and what is commonly called the Principle of Causality.

To make my above statement more intelligible to a wider circle of (non-mathematical) readers, let us consider some very simple examples which will enable us also to see the exact meaning of instantaneous 'state' of a system and to learn to distinguish between two very important and large classes of systems: 1) complete or 'undisturbed,' and 2) incomplete or 'disturbed' systems.

Suppose we have a small metallic sphere, $\dagger$ suspended somewhere in a large dark cellar kept at constant temperature $a$, receiving no heat, radiant or other, from without. Suppose we heated the sphere to $100^{\circ} \mathrm{C}$., which is to be $>a$ (say, $a=0^{\circ} \mathrm{C}$.), and from that instant left it to its own fate. We return to it after an hour, as measured, say, on one of our common clocks (i.e. rotating earth as time-keeper), and we find it has cooled down, say, to $90^{\circ}$. Thus :

$$
\left.\begin{array}{lc}
t \quad \theta \\
t_{0} & \mathrm{r} 00^{\circ}, \\
t_{0}+\mathrm{rh} . & 90^{\circ} .
\end{array}\right\} \quad \begin{aligned}
& \therefore \Delta \theta=-10^{\circ} \\
& \quad \text { for } \Delta t
\end{aligned}=\mathrm{xh}
$$

Now, if we repeated the whole experiment to-morrow or next week, we should find that during one hour the fall of temperature of our suspended sphere would again be from $100^{\circ}$ to $90^{\circ}$, i.e. $\Delta \theta=-10^{\circ}$ for $\Delta t=\mathrm{rh}$. We could make similar observations for any other stage of the cooling process of our little sphere (say down from $50^{\circ}$ instead of $100^{\circ}$ ) and for other time-intervals (say $\frac{1}{2} \mathrm{~h}$. instead of I .), arbitrarily small, $\ddagger$ and, repeating our observations, we should find again and again the same permanency of results,-only with different values of $\Delta \theta$ for different intervals $\Delta t$ and for different starting temperatures.

[^5]
## COMPLETE AND INCOMPLETESSYSTEMS

Thus, the temperature $\theta$ of our sphere, placed in in specified conditions of its environment, varies withotime (ordinary do (oxime) in a certain determinate way, namely, so adstarthy nom jj en
 alivays one and the same, that is to say, no matter when this battens, independently of $t_{1}, t_{2}$, but depending only on

$$
t_{2}-t_{1}=\Delta t
$$



Now, such a system, ie. the sphere in its above environment, $I_{3}$ propose to call an undisturbed or, what for the beginning is more cautious, complete system. And, in this case $\theta$ being the only quantity on whose instantaneous value the whole (thermal) future history of our sphere depends, we shall say, in accordance with general use, that the instantaneous value of the temperature $\theta$ defines the instantaneous state of our system ( $\alpha$ being supposed given once and for ever). In the case before us we have a one-dimensional system, which may be called also a system of one degree of freedom.*

Take the limit of the mean rate of change $\Delta \theta / \Delta t$ for $\Delta t \rightarrow 0$; then the differential equation of our simple system will be of the form

$$
\begin{equation*}
\frac{d \theta}{d t}=f(\theta), \tag{I}
\end{equation*}
$$

which may be read: the instantaneous time-rate of change of the temperature is a function of its instantaneous value only. $\dagger$ We know in this case that $f(\theta)=-h(\theta-\alpha)$ approximately, when $\theta-\alpha$ is small, where $h$ is a positive constant; but the particular form of the function $f$ is for our present purposes a matter of indifference.

Let us, on the other hand, consider a similar sphere suspended, say, in a window, exposed south, in a land in which the sun is wont to shine often. Then, for the same starting value $\theta$ and same $\Delta t$, the change $\Delta \theta$ will be different at different times of the day, ecg. larger from 7 till 8 a.m. than from 2 till 3 p.m., larger in winter than in summer, and so on. Now, a system such as this sphere we will call a disturbed' system or a system ' exposed to external agents,' or better an incomplete system, for this concept does not presuppose the knowledge of what is meant by 'action' of one system upon another.

[^6]In the present case the differential equation of our system will be of the form

$$
\begin{equation*}
\frac{d \theta}{d t}=g(\theta, t), \tag{2}
\end{equation*}
$$

$t$ being again measured with the ordinary (earth-)clock, and $g$ being some function involving $t$ in a very complicated manner.

Now, according to the above general requirement, our $t$-clock would be the right one, the peculiarly fitted one, for our first physical system, (I), but not for the second, (2).

By selecting a different time-keeper we might possibly convert some (not all) 'disturbed' into 'undisturbed' or complete systems; but then we should spoil the completeness of ( r ). Let us see, first of all, what other clocks we can take instead of our original one without spoiling the simple property of ( I ). Instead of $t$, take

$$
T=\phi(t) ;
$$

then ( I ) will be transformed into

$$
\frac{d \theta}{d T}=f(\theta) / \dot{\phi}(t), \text { say }=\psi(T) \cdot f(\theta) .
$$

Thus, if the property of completeness is to be preserved, $\dot{\phi}(t)$ must be a constant, and consequently $T$ a linear function of $t$, say

$$
T=t_{0}+a t,
$$

amounting only to a different initial point of time-reckoning and to the choice of a different time unit.

Now (2), the equation of our second sphere, is not of the form $d \theta \mid d t=\psi(t) \cdot f(\theta)$, but rather of the form

$$
\frac{d \theta}{d t}=f[\theta-a(t)]+G(t) ;
$$

consequently, if we wished also to sacrifice the completeness of ( I ), we certainly cannot transform (2) into an undisturbed or complete system, by any $T=\phi(t)$. Hence the moral: certain incomplete systems cannot be made complete by merely selecting a new clock instead of the old one, and such systems I propose to call essentially incomplete systems.

But suppose we had a system obeying a law of the form

$$
\begin{equation*}
\frac{d \theta}{d t}=-h(t) \cdot(\theta-a), \tag{3}
\end{equation*}
$$

i.e. a sphere as in ( x ), but having a coefficient $h$ (coefficient of what Fourier called external conduction, divided by specific thermal capacity), which due to some visible changes of the sphere's surface, such as oxidation, is variable, instead of being constant. Then we could represent it as a complete system by taking instead of the $t$-clock another clock indicating the time

$$
T=\int_{0}^{t} h(t) d t, \text { say }=F(t) \text {; }
$$

but, $F(t)$ not being a lincar function of the old time, this innovation would at once spoil the completeness of (r).

At this stage we would find ourselves in face of an alternative : which of the two systems, ( $\mathbf{r}$ ) or (3), is to be saved, which is to be sacrificed? And, correspondingly: which of the two clocks, the $t$-clock or the $T$-clock is to be sclected as time-keeper? If we could detect no differences between the spheres ( x ), (3)-besides that of their respective thermal histories-the choice would be difficult, or rather arbitrary, quite a matter of taste or caprice. But, say, the latter sphere, (3), gets oxidized, shrinks or expands, and what not, and the former, ( r ), remains sensibly unaffected by the process of repeated cooling and heating. Therefore, following the maxim or principle of causality, we would conserve our $t$-clock, best fitted for ( x ), and would try to convert (3) into a complete system in a different way, namely, by taking account explicitly of the oxidation of the sphere's surface, of its dilatation, and so on, i.e. by introducting besides $\theta$ other quantities, say, the amount $m$ of free oxygen present in the enclosure and the radius $r$ of the sphere, and by defining the state of the system by the instantaneous values of $\theta, m r$.

In this way, retaining our old clock, we should have converted the originally disturbed system of one degree of freedom into a complete system of three or more degrees of freedom. As a rule, we do not reject our traditional time-keeper at once. Encountering an incomplete or disturbed system, every physicist will, first of all, try to throw the 'disturbances' on some 'external agent' rather than on his clock. He will look round him for external agents, almost instinctively following the voice of the maxim of causality, whispering to him, as Maxwell puts it (Matter and Motion, p. 21): 'The difference between one event and another does not depend on the mere difference of the times.' And finding nothing particularly suspect in the nearest
neighbourhood, he will look farther round, or deeper into, the system in cjuestion.

Similarly, if we amplified the system of our second example (the sphere cooling before an opern window), taking in the sun varying in position, the atmosphere, and possibly a host of other things, we would obtain a larger, more comprehensive system which, though more complieated than the original one, would satisfy us as being undisturbed, with our old $t$-clock.

So it is in many other eases. Thus, we can say:
Adding to a given fragment of nature (system), which in terms of a certain $t$-clock behaves like a disturbed or incomplete system ( $p_{1}, p_{y}, \ldots p_{n}$ ), i.e. obeys the equations

$$
\begin{gather*}
d p_{1} f\left(p_{1}, p_{2}, \ldots p_{n}, t\right),  \tag{4}\\
d t=1,2, \ldots n,
\end{gather*}
$$

fresh frayments of nature (with the corresponding parameters $p_{n+1}, \ldots p_{n+n}$ ), we often obtain a now, larger,* system which, still with the same $t$, is undisturbed or complete:

$$
\begin{gather*}
d p_{i} \quad F_{i}\left(p_{1}, p_{1}, \ldots p_{n}, p_{n+1}, \ldots p_{n+m}\right),  \tag{5}\\
i=1,2, \ldots n+m .
\end{gather*}
$$

In short, we complete the system $S_{n}$ to $S_{n+m}$. The $t$, implied here, is practically the time indieated by that clock which proved peculiarly fitted for the deseription of our previous stock of experience. 'Thus, for example, Fourier's theory of conduction of heat was preceded by the triumphs of classical mechanics; and if asked what the 6 in his fundamental ecpuation

$$
\begin{aligned}
& i \theta=a^{2} \Gamma^{2} \theta \\
& i t=0
\end{aligned}
$$

meant, he would, doubtless, answer that it is to be measured by the rotating earth as time-keeper, though he hardly ever stopped in his researches to consider this matter explicitly.

Thus, generally, we do not reform our traditional clocks, but we make our systens complete as in (5), by amplifying them. But

[^7]sometimes, when we think that we have made our system $S_{n+m}$ sufficiently comprehensive, that we have exhausted all reasonably suspected material as possible 'external agents,' and when $S_{n+m}$ nevertheless continucs to behave as an incomplete system, i.e. when still
\[

$$
\begin{equation*}
\frac{d p_{i}}{d t}=f_{i}\left(p_{1}, \ldots p_{n+m}, t\right) \tag{6}
\end{equation*}
$$

\]

then, to make it finally complete, we decicle ourselves to change our $t$, our traditional clock,-especially if the change required is a slight one. 'This procedure, of course, is possible only when the $F_{i}$ 's in (6) are all of the form

$$
\begin{equation*}
F_{i}=\phi(t), M M_{i}\left(p_{1}, \ldots p_{n+m}\right) . \tag{7}
\end{equation*}
$$

Otherwise, we feel obliged to help the matter by introducing yet fresh parameters $p_{n+m+1}, p_{n+m+4}$, cts:, and not finding real (perceivable) supplementary material round us, we introduce fictitious supplements, which sometimes turn out to be real afterwards, thus leading to new discoveries.

From this it is also manifust that the Principle of Causality has the true character of a maxim; though of inestimable value both in science and in everyday life, it is not a law of nature, but rather a maxim of the naturalist.
We have classical examples of hoth the procedures sketched above, viz. of reforming our clocks and of supplementing or amplifying a system with the view of securing its completeness. In the first place, to get rid of one of the inequalities in the motion of the moon round the earth, astronomers have had recourse to the supposition that there is a gradual slackening in the speed of the earth's rotation. Of course, they did it in connexion with the tides and with immediate regard to the fundamental principles of mechanics, implying also the law of gravitation. But at any rate, in doing so, and in declaring that the earth as a clock is losing at the rate of 8.3 , or (according to another estimate) of 22 seconds fer century, they gave up the earth as their time-keeper and substituted for the sidercal time $t$ a certain function $T \equiv \phi(t)$, slightly differing from $t$, as their new 'kinetic time,' as Prof. Love calls it.* Secondly, as is widely known, the perturbations of the planet Uranus have led Adams and Le Verrier
*A. E. II. Love, Theoretical Mechanics, second edition, Cambridge, 1906, page 358. In connexion with our sulject, the whole of Chapter XI. of Prof. Love's book may be warmly recommended to the reader.
(working independently) to complete the system by a celestial body, at first fictitious, but then, thanks to admirable calculations based on the $\frac{\mathrm{T}}{\gamma^{2}}$-law, actually discovered and called Neptune. Notice that both kinds of procedure have essentially the character of successive approximations.
Any future researches of mechanical, thermal, electromagnetic and other phenomena, either new or old ones but treated with increasing accuracy, if leading to 'disturbed' systems, obstinately withstanding the supplementing procedure (i.e. that consisting in the introduction of fresh parameters $p_{n+1}$, etc.), may oblige us to reform also the newer, slightly corrected earth-clock, to give up the 'kinetic time' of modern astronomy for a better one, more exactly fitted for the representation of a larger field of phenomena, and so on by successive approximation. It may well happen that we shall have to give up the kinetic time for the sake of the 'electromagnetic time,' if I may so call the variable $t$ entering in Maxwell's differential equations of the electromagnetic field.* For suppose for a moment that some future experimental investigations of high precision were to prove that the variable $t$ in

$$
\frac{\partial \mathbf{E}}{\partial t}=c \cdot \operatorname{curl} \mathbf{M}, \quad \frac{\partial \mathbf{M}}{\partial t}=-c \cdot \operatorname{curl} \mathbf{E}
$$

is not proportional to the kinetic time ; then the electricians would hardly give up these admirably simple and comprehensive equations; they would rather sacrifice the kinetic time. Thus, in the struggle for completeness of our physical universe, we shall have always to balance the mathematical theory of one of its fragments, or sides, against that of another. A great help in this struggle is to us the circumstance that, though, rigorously, all parts of what is called the universe interact with one another, yet we are not obliged to treat at once the whole universe, but can isolate from it relatively simple

[^8]parts or fragments, which behave sensibly as complete systems, or are easily converted into such.

Herewith I hope to have explained to you, at least in its fundamental points, the question of selection of a time-keeper.

Thus, we know, essentially, how to measure our $t$, at least in or round a given place (taken relatively to a certain space-framework). We do not yet know what is the precise meaning of simultaneous events occurring in places distant from one another. But the notion of simultaneity, especially for systems moving relatively to one another, belongs to the modern Theory of Relativity, and is, in fact, a characteristic point in Einstein's rcasoning. Therefore it will best be postponed until we come to treat of the principal subject of this volume.

We could now pass immediately to the history of the electromagnetic origin of the modern principle of relativity, extending from Maxwell to Lorentz. But since we alrcady have come to touch, more than once, Newtonian or classical mechanics, let us dwell here another moment upon this subject.

Let us call $\Sigma$ one of the 'incrtial' systems of reference, say the system of 'fixed' stars, and let $x_{l}, y_{l}, z_{l}$ be the rectangular coordinates of the $i$-th particle* of a material system, relatively to $\Sigma$, at the instant $t$. Then the Newtonian equations of motion are

$$
\begin{equation*}
m_{i} \frac{d^{2} x_{i}}{d i^{2}}=X_{l}, \text { ctc. } \tag{8}
\end{equation*}
$$

or

$$
\begin{gathered}
\frac{d x_{i}}{d t}=u_{t}, \quad d y_{i}=z_{i}, \quad d z_{i}=z v_{l}, \\
m_{l} \frac{d t}{d t}=X_{i}, \quad m_{i} d v_{i}=Y_{i}, \quad m_{l} \quad d z v_{i}=Z_{i},
\end{gathered}
$$

where $m_{l}$, the masses, are constant scalars belonging to the individual particles, $t$ is the 'kinetic time' and $X_{l}$, etc., are functions of the instantaneous state of the material system, i.e. of the instantaneous configuration and (in the most general case) of the instantaneous velocities of the particles relatively to one another, which for certain systems may, but for a sufficiently comprehensive system do not, contain explicitly the time $t$. If the material system is subject to constraints, say

$$
\phi=0, \quad \psi=0, \quad \text { ctc. }
$$

* The material 'particle' may also play the part of a planet or of the sun, as in celestial mechanics.
then $X_{i}$, etc., contain, besides the compronents of what are called the impressed forces, also terms like

$$
\lambda_{i \psi}^{i}+x_{i}^{2} \psi x_{i}^{i}+\ldots
$$

which depend only upon the relative positions and relative velocities of the parts of the system (i.c. of the mass-particles) to one another or to the surfaces or lines on which they are constrained to remain, or to the points of suppert or suspension entering in such constraints. Thus the bob of a penclulum is constrained to remain at a constant distance relatively to the point of suspension, the friction of a body moving on a rough surface depends on its velocity relative to that surface, and so on. Consequently, if instead of $\searrow$ any other system of reference $\mathcal{Z}^{\prime \prime}\left(x^{\prime}, y^{\prime}, z^{\prime}\right)$ is taken, having relatively to $\Sigma^{\prime}$ a purely translational, uniform, rectilinear motion, $X_{i}, Y_{i}, Z_{l}$ are not changed. And the same thing is true of the left-hand sides of the equations of motion. For, if $x_{i}^{\prime}$, ctce, be the coordinates of the $i$-th particle relatively to $\Sigma^{\prime \prime}$ at the instant 1 , and if we take, for simplicity, the axes of $x^{\prime}, y^{\prime}, z^{\prime}$ parallel to and concurrent with those of $x, y^{\prime}, z$ respectively, then

$$
\left.\begin{array}{c}
x_{i}^{\prime}=x_{l}-u t, \quad y_{i}^{\prime}-s_{t}-v t, \quad z_{i}^{\prime}=z_{l}-2 v t,  \tag{9}\\
t_{t}^{\prime}=t,
\end{array}\right\}
$$

where $(4, z, z)$ is the constant velocity of $\Sigma$ relatively to 2 , and where the fourth equation is added to emphasize that the old time $t$ is retained in the transformation. Consequently,

$$
u_{i}^{\prime}=\frac{d x_{i}^{\prime}}{d t^{\prime}} \frac{d x_{i}}{d t}-u=u_{i}-u, \quad \text { ctc. }
$$

(and for any pair of particless $u_{i}^{\prime}-u_{j}^{\prime}=u_{i}-u_{j}$, cte.), and
which proves the statement.
Thus, the equations of motion (8), or, in vector form,

$$
\begin{equation*}
m_{l}^{d^{4} \mathbf{x}_{1}} d t^{2=1} \mathbf{F}_{l} \tag{8a}
\end{equation*}
$$

remain unchanged by the transformation (9), or, written vectorially, by the transformation

$$
\left.\begin{array}{c}
r_{i}^{\prime}=r_{i}-\nabla t_{j}  \tag{9a}\\
t^{\prime}=\left\{t^{\prime}\right.
\end{array}\right\}
$$

where $\boldsymbol{\nabla}$, the resultant of the above $u, v, z$, is the vector-velocity of $\Sigma^{\prime}$ relatively to $\Sigma$. As regards the time, we could write also $t^{\prime}=a t+b$ ( $a, b$ being constants), but this would amount only to a change of units and shifting of the beginning of time-reckoning.
In view of the above property, the linear transformation (y) or ( $9(a)$, $\nabla$ being any constant vector, is called the Newtonian (and by some: authors the Galileian) transformation. Thus we can say, shortly:

The equations of classical mechanics are invariant reith restect to the Newtonian transformation.
Notice that $\nabla$ being quite arbitrary, both as regards its size (or tensor) and direction, we have in ( $9 a$ ) a manifold of $\infty^{3}$ transforma tions, and all of these form a group of transformations. For, if
and

$$
\begin{align*}
& \mathrm{r}_{i}^{\prime}=\mathrm{r}_{l}-\nabla_{1} t ; \quad t^{\prime}=t, \\
& x_{l}{ }^{\prime \prime}=r_{i}{ }^{\prime}-\nabla_{2} t^{\prime} ; t^{\prime \prime}=t^{\prime}, \\
& \mathbf{r}_{i}^{\prime \prime}=\mathbf{r}_{l}-\boldsymbol{\nabla} t ; t^{\prime \prime}=t^{\prime}, \\
& \boldsymbol{v}=\mathrm{v}_{1}+\mathrm{v}_{2} .
\end{align*}
$$

We shall refer sometimes to ( 9 ) or ( $g e r$ ) as the Nowtonian group.
Notice the simple additive property ( ro ), to be compared later on with a less simple property of the corresponding group, in modern Relativity.

Thus, there is no unique frame of reference for classical mexhanies: if the Newtonian equations of motion are strictly valid relatively to the framework $\grave{\prime}$ of the 'fixed' stars, they are equally valid relatively to any other out of the $\infty^{8}$ frameworks $\Sigma^{\prime \prime}$, comected with $\Sigma$ by (9), say relatively to the solar-system frame, which has relatively to $\Sigma$ a uniform velocity of something like 25 kilometres per second, towards the constellation of Hercules.* Therefore, hy purely inm ternal mechanical experiment and obscrvation, i.e. not looking outside to external systems, we could never distinguish the solar frame $\Sigma^{\prime}$ from $\Sigma$, that is to say, $\Sigma^{\prime}$, like $\Sigma$, does not show any anisotropy with regard to mechanical phenomena. The same remark applies, with sufficient approximation, to the earth's amual motion: it is not ascertainable by purely terrestrial mechanical experiments.
Physicists hoped to detect this motion which they called also 'the motion relative to the aether,' by the means of purely terrestrial

[^9]s.r.
optical or electromagnetic experiments,-we shall see later how unsuccessfully.

In other words, seeing that there is no unique 'kinetic' spaceframework, they tried to find a unique 'optical' or 'electromagnetic' reference-system, the 'aether,' or rather to show that this wonderful medium, already invented for other purposes, was such a unique frame of reference. But the results of all experiments of this kind have been obstinately negative.

It is chiefly this which has led to the construction of the new theory of relativity.

## NOTES TO CHAPTER $I$.

Note 1 (to page 9). To show, generally, the connexion between the integral form of the properties of a complete system, as stated in the above illustrations, and its differential form, of which $\mathrm{eq} .(\mathrm{I})$ is an example, let us consider such a system of $n$ degrecs of frecdom. Let its state at any instant $t$ be determined by

$$
p_{1}(t), p_{2}(t), \ldots p_{n}(t)
$$

Then, $t_{0}=0$ being any other, say, past instant,

$$
p_{i}(t)=P_{i}\left[p_{1}(0), \ldots p_{n}(0) ; t\right], \quad i=\mathrm{r}, 2, \ldots n,
$$

where $P_{i}$ is a symbol of an operation or a function, implying besides the 'initial' state $p(0)$ the time-interval $t=t-t_{0}$ clapsed, but independent of the choice of the initial instant. This is the finite or integral way of expressing that the system is complete. Now let $t=\boldsymbol{c}$ be any particular instant and $t=c$ another instant of time, such that

Then

$$
c=a+b .
$$

$$
p_{i}(c)=P_{i}\left[p_{1}(v), \ldots p_{n}(t) ; b\right]=P_{i}\left[p_{1}(0), \ldots p_{n}(0) ; c\right],
$$

so that the transformations corresponding to the passage of the system from any of its states to its successive states form a group of transformations, $t$ being the (only) 'parameter' of the group. Thus we can imitate Lie's general proof of his Theorem 3 (Sophus Lic, Theorie der Transformationsgruphen, Leipzig, 1888 ; Vol. I.) for this simplest case of one
parameter. Considering $p_{1}(0), \ldots p_{n}(\circ), \alpha, c$ as independent variables, differentiate $p_{i}(c)$ with respect to $\alpha$; then

$$
\begin{aligned}
& =\frac{\partial p_{l}(c)}{\partial a}=0 ;
\end{aligned}
$$

but $\partial b / \partial c=-\mathrm{r}$; thercfore

Now $\not_{1}(c), \ldots p_{n}(c)$ are mutually independent; otherwise less than $n$ quantities $p$ would suffice for the determination of the state of the system, contrary to the supposition. Therefore the functional determinant

$$
\left\lvert\, \begin{array}{lll}
\partial p_{1}(c) & & \partial p_{n}(c) \\
\partial p_{1}(a)
\end{array}\right., \cdots, \frac{\partial p_{n}(r)}{}(t)
$$

cloes not vanish identically, and the above system of $n$ equations catn be solved with respect to $d p_{1}(a) / d u$, ete., leading to

$$
\frac{d p_{i}(r)}{d r}=F_{i}\left[p_{1}(r), \ldots p_{n}(r) ; i n\right], \quad i=1,2, \ldots n
$$

But these equations must be valid for all values of the momatly in dependent magnitudes $b$ and $a$. Giving therefore to bany constant valuc, and writing $t$ instead of $a$, we oftain for any $/$,

$$
\frac{\left\langle p_{1}(t)\right.}{d t}=f_{i}\left[\not p(t), \ldots p_{2}(t)\right]_{1} \quad i \quad i, a_{1}, \ldots \mu_{1}
$$

and this is the differential form alluded to, $f_{n}, f_{y}, \ldots f_{1}$ being funations of the instantaneous state only.
It is instructive to consider the instantaneous state of a system as a point in the $n$-dimensional space, or domutn if states $S_{n}\left(\rho_{1}, h_{1} \ldots A_{n}\right)$ and to trace in this 'space' the dines of states, i.e. the linear cominua of states assumed successively by different copies (exemplats) of the system, starting from given initial states. The differential rquations of these lines of states, or, as Lie calls them, the 'paths (Buthowrath) of the corresponding infinitesimal transformation,' are

$$
\frac{d p_{1}}{f_{1}}=\frac{d p_{3}}{f_{2}} \ldots=\frac{d p_{n}}{f_{n}} .
$$

A complete system may then be characterized by saying that the lines of states are fixed in the corresponding space $S_{n}$, like the line of flow of an incompressible fluid in steady motion. A cony of the system, or rather its representative point, placed on one of these lines remains on
it, moving along it in a determined sense. (For particulars of physical application of these concepts, see the author's paper in Ostwald's Annalen d. Naturphilosophie, Vol. II. pp. 201-254.)

Note 2 (to page 12). Systems obeying partial differential equations, as for instance that of Fourier,

$$
\frac{\partial \theta}{\partial t}=a^{2} \nabla^{2} \theta=c^{2}\left(\frac{\partial^{2} \theta}{\partial x^{2}}+\frac{\partial^{2} \theta}{\partial y^{2}}+\frac{\partial^{2} \theta}{\partial z^{2}}\right)
$$

adduced in the text, may be considered as systems of infinite degrees of freedom. The instantaneous state of such a system implies an infinite number of data $p_{i}$, or say $p=p(x, y, z)$, given as a function of $x, y, z$ for every point of a portion of space coextensive with the system, as for example the instantaneous temperature for every point of a cooling body of finite dimensions, in which case the system will have $\infty^{3}$ degrees of freedom. Instead of one we may have also two or more functions of $x, y, z$, defining the instantaneous state, as for example two vectors, amounting to six scalars, for an electromagnetic system (field), the differential equations being in this case those of Maxivell,

$$
\frac{\partial \mathbf{E}}{\partial t}=c \cdot \operatorname{curl} \mathbf{M}, \quad \frac{\partial \mathbf{M}}{\partial t}=-c \cdot \operatorname{curl} \mathbf{E} .
$$

Here, as in the above example, the right-hand sides do not contain the time explicitly, but depend only on the space-distribution of magnitudes referring to the instantaneous state. If such be the clifferential equations and if also the limit or surface-conditions do not contain the variable $t$ explicitly, the system of infinite degrees of freedom will be a complete or undisturbed one, in the sense of the word adopted throughout the chapter. Thus a heat-conducting sphere, of finite radius $R$, obeying in its interior Fourier's equation and whose surface is thermally isolated or radiates heat into free space, will be a complete system; for its boundary conditions, viz.

$$
\frac{\partial \theta}{\partial r}=0
$$

or

$$
\frac{\partial \theta}{\partial t}=\text { const. } \times(\theta-\text { const. })
$$

respectively, do not contain the time explicitly. But a sphere (like the earth), whose surface is kept at a generally variable temperature by means of external sources (like the sun), will be an incomplete system, unless we amplify it by taking in the 'sources' themselves.

## CHAPTER II.

## MAXWELLIAN EQUATIONS FOR MOVING MEDIA AND FRESNEL'S DRAGGING COEFFICIENT. LORENTZ'S EQUATIONS.

THE modern principle of relativity arose on the ground of Lerenta's electrodynamics and optics of moving bodics. Einstein's work, in fact, consisted mainly in deducing logically, on the basis of plausible and sufficiently general considerations, certain formulae of space and time transformation, which in Lorentz's theory had partly a purels; mathematical meaning and partly the character of an hypothesis invented ad hoc ('local time' and the contraction hypothesis, respectively). In a word, Linstein has given a plausible support to, and a different interpretation of, what appeared already in the theory of the great Dutch physicist. In its turn, the theory of loorent, based on the macroscopic treatment of a crowd of electrons (though later supported and made vital by physical evidence of an entirely different kind), was constructed by its author chiefly with the purposeof accounting for optical phenomena in moving bodies, which may be best grouped summarily under the head of Fresnel's 'drugging eocefficient' and with which the equations of Maxwell and of HertzHeaviside have proved to be in complete disagreement.
Now, it seems to me that the best, most natural and most efficient way of propagating new ideas (if incleed there is such a thing arising in the collective mind of humanity) is to show their intimate connexion with older ones, and the more so when the now ideas have the reputation, widespread but partly unjustified in our catse, of being of a very revolutionary character. It will be advisable, therefore, before entering upon our proper subject, to turn back to Lorent\% and Maxwell. In doing so, I must warn the reader at the outset that the new Relativity, though grown on electromagnetic soil, does not- in spite of a current opinion-require us at all to adopt an electros

magnetic view of all natural phenomena. Nor does it force upon us a purely mechanistic view, which till recently held the field, before the pan-electric tendencies arose. Modern Relativity is broader than this: it subordinates mechanical, electromagnetic and other images to a much wider Principle which is colourless, as it were.

Thus, the reason of returning here to Maxwell is, in the first place, of an historical (and partly didactic) character. But we have yet another reason for dwelling in the present chapter upon the great inheritance left to Science by Clerk Maxwell. It is widely known that but a few things of the old system of physics have remained untouched by the modern principle of relativity, though the changes required are generally but very slight. In fact, almost nothing of the old structure has been spared by the new theory of relativity; but Maxwell's fundamental equations, namely those known as his equations for 'stationary' media, have been spared. More than this: not only have they been preserved entirely in their original form, without the slightest modification of any order of magnitude whatever, but they form one and the best secured of the actual possessions of the new theory, the largest and brightest patch of colour, as it were, on the vast and as yet mostly colourless canvas contained within the frame of the new Principle. Moreover, a peculiar union or combination of the electric and magnetic vectors which appear in Maxwell's equations of the electromagnetic field became the standard and prototype (not as regards physical meaning, but mathematical transformational properties) of a very important class of entities admitted by the new theory (the so-called 'world-six-vectors' or 'physical bivectors').

So much to justify the insertion of the following topics of the present chapter.

Maxwell's fundamental laws of the electromagnetic field in a 'fixed' or 'stationary' non-conducting dielectric medium* may be expressed in integral form as follows:
I. Electric displacement-current through any surface $\sigma$ bounded by the circuit $s=c \times$ line integral of magnetic force $\mathbf{M}$ round $s$.
II. Magnetic current through $\sigma=-c \times$ line integral of electric force $\mathbf{E}$ round $s$,

* Practically, fixed with respect to the earth, or, if not, then with respect to a definite system of reference $S$, to be ascertained on further examination.
i.e. in mathematical symbols:

$$
\begin{align*}
& \frac{d}{d t} \int(\mathscr{F} \mathrm{n}) d v=c \int_{(x)}(\mathbf{M} d \mathrm{~s})  \tag{I.}\\
& \frac{d}{d t} \int(\mathfrak{f l l n}) d \sigma=-c \int_{(*)}(\mathbf{E} d \mathbf{s}),
\end{align*}
$$

II.
here $\mathfrak{F}, \mathfrak{A l}$ denote the diclectric displacement or polarization and ie magnetic induction respectively, $c$ a scalar constant, the velocity fight in vacuum, $n$ a unit vector normal to $r$, the sense of the ategration round $s$, of which $d s$ is a vectorial clentent, being clockise for a spectator looking along $n$ (see Fig. s). Here, as throughut


Fla. r
evolume, ( $(\mathfrak{H n})$, ctc., generally ( AB ), in round parentheses, denote ie scalar product of a pair of vectors:

$$
(\mathbf{A B})=A B \cos (\mathbf{A}, \mathbf{B})
$$

, $B$ being the sizes or absolute values of the vectors $\mathbf{A}, \mathbf{B}$.* Thus, e surface element $d \sigma$ being considered as an ordinary scalar, the arface integral $\int(\mathbb{E} n) \mathrm{d} \sigma$ stands for the total number of Faraday thies nit tubes) crossing $\sigma$, and the surface integral in in. has a similar eaning with respect to the tubes of magnetic induction.

* If it were only for purely vectorial algehra and analyais, we could write, atter eaviside, for the scalar produet simply AB. I3at since we shall have to reent in the quel to Hamilton's quaternionic calculus, we reserve AB for the foll quaterniomir oduct, and write therefore (AB) for the senlar product, i.c. for the notative alar part of the Hamiltonian product, and VAB for the vector produet, thus

$$
\begin{aligned}
A B & =S \cdot A B+V \cdot A B \\
& =-(A B)+V A B .
\end{aligned}
$$



Remembering the definition of 'curl' by means of the line integral, we may write I . and 1. at once in differential form,

$$
\left.\begin{array}{c}
\partial 0 \mathrm{Et}  \tag{1}\\
\partial t=c \cdot \operatorname{curl} \mathbf{M}, \\
\frac{\partial \mathrm{gl}}{\partial \mathrm{t}}=-\operatorname{coc} \operatorname{curl} \mathrm{E},
\end{array}\right\}
$$

or, in Cartesian expansion,

$$
\begin{gather*}
\frac{1}{c} \frac{\partial \mathbb{E}_{1}}{\partial t}=\frac{\partial M_{3}-\partial M_{g}}{\partial t}-\quad \text { ctc., } \\
\frac{\partial z}{c} \frac{\partial A l_{1}}{\partial t}=\frac{\partial E_{4}}{\partial z}-\frac{i E_{3}}{\partial y_{3}},
\end{gather*}
$$

Every point or surface element of or being fived, relatively to the system of coordinates $x, i, z$, round its have been written on the left hand to express partial differentiations with respece to $t$, i.e. local time-rates of change of the corresponding vectors.
( $\mathrm{a} a$ ) or ( x ) is the Herts-Heaviside form of Maxwell's differential equations, although, if 1 an not nistaken, Maxwell himself on one occasion employed this form. At any rate, the Mertz-I Ienviside equations for a stationary medium differ only formally from the equations of Maxwell as given in his monumental 'Treatise' and in several papers; the auxiliary potentials being easily climinated.

As regards the relations obtaining between $\mathbb{E}$, fit and E, M respectively, it will be enough to remember here that the first pair of vectors are linear functions of the second, say,

$$
\begin{equation*}
\mathbb{C E}=\mathcal{K}^{\prime} \mathrm{E} \text { and }, \mathfrak{A l}=\mu \mathrm{M} \text {, } \tag{2}
\end{equation*}
$$

where $K^{*}, \mu$ are in the general case, of crystalline bodies, symmetrical or self-conjugate linear vector operators, which in the simplest case of an isotropic medium degenerate into cordinary scalar coefficients, the diclectric 'constant' or the permiltivith, and the magnetic permeability or the inductivity,--to adopt Meaviside's nomenclature.*

Notice that, using the relations (2), $K^{\circ}$ and $\mu$ being supposed given, we have in ( $x$ ) two vectorial equations of the first order for two vectors, so that if the 'initial' state, saly $\mathbf{E}_{\mathbf{a}}, \mathbf{M}_{0}$, and cventually the limit-conditions, be given, the whole history of the fied, past

[^10]and future, is uniquely determined,-though in most cases the mathematician may have the greatest difficulties in finding it out. The electromagnetic field, as far as it obeys these equations, is at any rate a complete system in the sense of the word previously explained. It will be noticed later that the fundamental equations of the electron theory do not possess this simple property.

From 1., II. we see immediately that the total current, electric or magnetic, through all possible surfaces $\sigma$ bounded by one and the same circuit ( $s$ ), has the same value. Taking therefore a pair of such surfaces $\sigma_{1}, \sigma_{2}$, which together form a surface ( $\sigma$ ), enclosing completely a certain portion $\tau$ of the medium, and inverting one of the normals of the component surfaces (Fig. 2), so that the

normal n is directed everywhere outwards (or everywhere inwards) with respect to the enclosed space, we see that, for any closed surface ( $\sigma$ ),

$$
\int_{(\sigma)}(\mathbb{E n}) d \sigma, \quad \int_{(\sigma)}(f t \mathrm{n}) d \sigma=\text { const. in time },
$$

the second constant being everywhere equal to zero, by experience. In other words, the total electric charge enclosed by $(\sigma)$ does not vary in time, its magnetic analogue being constantly non-existent. The same property being valid for any volume $\tau$, and remembering that 'div' or divergence is defined as the surface integral of a vector per unit of enclosed volume, we may write also, in differential form,

$$
\begin{aligned}
\operatorname{div} \mathbb{E} & =\rho=\text { const. } \\
\operatorname{div} \mathfrak{A} & =\text { const. }=0 ;
\end{aligned}
$$

$\rho$ is the volume density of (true) electricity. The second property is commonly expressed hy saying that the tubes of magnetic induction are always closed, or that $f_{t l}$ has a purely solemoided distribution. The invariability of both divergencess may be seen with ecpual case from ( r ), remembering that the operations div and a/at are commutative, while diveurl $=0$, identically:
Thus, the full system of Maxwell's exfuations for a stationary dielectric, which we will put here together for future reference, is
the equation

$$
\rho \cdot \operatorname{dive}
$$

being here considered as the defintion of the alensity $p$ of wectric charge. Notice, in passing, that the 'clectric charges' have been driven to the background by the Maxwellian thenry (especeially as propagated by Hert\%, Itaviside and Emil (ohn), as rather secondary derivate entities, but to return later with incrased vigour and to reacquire their dominant position, viz. as fundamental elements of the electron theory.

We shall not stop here to consider the general Maxwellian expressions of energy, ponderomotive force and of the corresponding stress.

In vacuo, and practically also in air under ordinary conditions,

$$
\kappa_{\approx} \approx \mu
$$

so that Maxwell's equations (3) become
(4)
to which in the present case may be added also

$$
\begin{equation*}
\operatorname{div} \mathbf{E}=0 \tag{1}
\end{equation*}
$$

expressing the absence of electrue charge. Notice in passing that these equations are not invariant with respect to the Newtonian
unsformation. The transformation which dees preserve their form of a different kind, as will be seen later.
The independent variable $t$ appearing in Maxwll's equations ( $t$ ) e empty space may be taken, provisionally at least, as far as perience goes, to be the ordinary or the kinctie time. Aud ats regards e (or a) space-framework, with respect to which they are intended be rigorously valid, let us call it once and for ever the system. S; natever it may be. If the reader wants to fix his ideas he may think $S$ as the 'fixed-stars' system; but as yet we crunot and need not scuss this point thoroughly, being forced by the very nature of the estion to postpone it to a later chapter. At first sight it might em that (4) are wholly independent of a space-frame of reforence; r the curls and div's can be, and primarily are, clefined in terms line integrals and surface integrals respectively, and thas depemd aly upon the distributional peculiarities of the respestive vector alds. But this means only that the equations in question are dependent of the choice of axes ( $x, y, s$ ) wilhin $S$, the omly erombi on being that they must be immovable relatively to $S$; in other ords, curl E, curl M are vectors as goocl as E, M themselves *and v $\mathbf{E}$, div $\mathbf{M}$ are true scalars like a volume, for instanes. Noties, wever, that, on the left hand of the equations, iy, is to be the fal time rate of change of $\mathbf{E}$ or $\mathbf{M}$, i.e. the variation in a point $/$ ept fixed. Now, this would be altogether meaningless if it not explained with respecet to what frame the pesint $/ P$ is to fixed. It would not belp us very much if somobordy told that $P$ is to be a fixed point of the field or of a Faraday be; for we have no means of identifying such a pwint. The: ath of what has just been said may be seen even more im. ediately from the integral form of Maxwell's equations, i. and it, here for the present case $\mathbb{E}$, git are to be identifed with $\mathbf{E}, \mathbf{M}$; r the circuit ( $s$ ) is to be kept 'fixed,' ic. fixed with respeet to mething. $\dagger$ 'Therefore we necessarily want a frame of referewe call it $S$.

* The distinction of what are called axia/ and polar vectors deves wht concron two re.
+ In the more general case of a ponderable medium, say in a piere of glase, the cenit (s) is, of course, to be fixed in the glass ; but this would not le enuugh : the nole piece of glass, as will be explained presently, must not moves in an arhitaty anner relatively to some external frame or other, if the laws 1., at ate to he: lid, whether the observer does or thees not share its mutium.


To see the property of the scalar constant $c$, eliminate, in the usual way, $\mathbf{E}$ or $\mathbf{M}$, employing their solenoidal properties; then

$$
\begin{equation*}
\because \quad \because^{i} \quad \frac{I}{c^{2}} \partial t^{2} \phi=\nabla^{2} \phi \tag{5}
\end{equation*}
$$

where $\phi$ mieans $\mathbf{E}$ or $\mathbf{M}$, or any one of their Cartesian components $E_{1}, \ldots M M_{3}$; hence, in the case of plane waves, for example,

$$
\frac{\mathrm{r}}{c^{2}} \partial^{2} \phi t^{2}=\frac{j^{2} \phi}{\partial x^{2}}
$$

tr
$\phi=f(x \pm c t)$, ..... T

$$
\phi=f(x \pm c t)
$$sl

$f$ being an arbitrary function of the linear argument. Thus $c$, ..... gi
in round figures $3.10^{10} \mathrm{~cm}$. sec. ${ }^{-1}$, is the velecity of propagation ..... a
in empty space, relatively to $S$, of transversal electromagnetic waves ..... w
or disturbances, their transversality being an immediate consequenceof the solenoidal conditions, which, in the present case, reduce to$\partial E_{1} / \partial x=0, \partial M_{1} / \partial x=0$. Henceforth $c$ will be referred to shortlyas 'the velocity of light,' and sometimes as the 'erilical' velocity.

What is properly called a wave is a nom-stationary surface of discon-
and
Notice that this property is quite independent of the direction of ..... f
the wave normal, i.e. of its orientation with respecet to any axes drawn ..... i:
in $S$. In other words: ..... (

Maxwell's equations imply isotropic as well as uniform * propaga-
tion in empty space relatively to S, i.e. to that system in which ..... t
they are valid. There are no privileged places or directions forthe electromagnetic disturbances.

Thus a continuous train of spherical waves, with centre $O$, will remain spherical for ever, which may be seen also from (5). For a

[^11]urticular integral of that cquation, adaptabforany initial state ,$=\frac{1}{r} f(r)$, is $\phi=\frac{1}{r} f(r \pm c t)$, $r$ thay tha scalar efoty measured om $O$. Again-which is more sall futpory-tir be ar anginstant spherical surface of transversal discoftifitity or propeteledroagnetic wave, then, expanding (or shrinking) with time, i main spherical for ever, with centre $O$ coinciding alpays at of the original $\sigma$, fixed once and for ever zeith respect to the " ame $S$,-quite independently of whether and how the material zurce was moving at the instant when it originated that yave. 'hus a 'point-source' (and notice that a physical source offany lape or finite dimensions may be regarded as such, provided we ว away from it far enough) producing a solitary disturbance, say flash of light, at the instant $t_{0}$, will originate a wave which always ill be spherical of radius
$$
R=c\left(t-t_{0}\right),
$$
aving its centre where the source was at the instant $t_{0}$, no matter ,hither it went afterwards or whence it came, or how swiftly it flashed mough that place.
We shall have to return to this argument, of capital importance, zore than once; but meanwhile we must leave it.
As has been already remarked, Maxwell's expuations for 'stationary' lielectrics, i.e. I. and It, with their supplements as given together rith their differential form under (3), have not only survived the eneral massacre, but have very substantially enriched the new theory. n fact, both the most particular and simple equations (4) for the acuum and the more general ones, (3), for ponderable media have reen incorporated into the possessions of modern Relativity, the ormer in a quite casy way by Einstein (1005), and the latter n a less easy and very ingenious way by Minkowski ( x 907 ). On the other hand, it is needless to tell here again about the side field of experience covered by these equations and about heir numerous and successful applications in proper Electronagnetism, to say nothing about the electromagnetic theory of light which soon after its creation proved to be much superior to the slastic theory.
Serious difficulties arose only in connexion with the electrolynamics, and more especially with the optics of moving media, a ong time before the dates just quoted.

There are two different sets of what are commonly called Maxwellian equations for moving media: $\mathrm{I}^{\circ}$ a system of equations which may be gathered together from different chapters of Maxwell's 'Treatise,' and which we shall call shortly the equations of Maxzeell, though it can be reasonably doubted whether Maxwell himself would consent to attribute to them general validity, especially with the inclusion of optics; and $2^{\circ}$ a system of equations which Hertz obtained by a certain, apparently the most obvious, extension of the meaning of the form i., II., and which Heaviside, independently, constructed by introducing into Maxwell's equations a supplementary term dictated by reasons of electro-magnetic symmetry; these are widely known as the Hertz-Heaviside equations for moving bodies.

We shall use for $\mathrm{r}^{\circ}$ and $2^{\circ}$ the abbreviations ( Mx ), ( HH ). Neither has been able to stand the test of experience. Though contrary to the historical order, it will be more instructive to consider first the latter and then the former system of equations.

Let us return to the semi-integral form of electromagnetic laws I . and 11., given, in words and symbols, on pp. 22-23. These are valid for a ponderable dielectric medium or body, stationary with respect to our frame $S$, and for any surface $\sigma$ which, together with its bounding circuit $s$, is fixed in the body. Thus the surface $\sigma$, through which the 'current' is to be taken, is itself fixed in $S$. Now, what Hertz did in order to obtain the required extension, was simply to suppose that I. and II. are still valid for a body, rigid or deformable, moving with respect to $S$ in any arbitrary manner, provided that the currents on the left-hand side of these equations are taken through a surface composed always of the same particles of the body, or-to put it shortly-through an individual $\sigma$, together with its $s$. This gives for the current per unit area of $\sigma$, instead of the local time-rate of change $\partial \mathbb{E} / \partial t$, if $\mathbf{v}$ be the velocity of a particle relatively to $S$,

$$
\begin{equation*}
\frac{\partial \mathbb{F}}{\partial t}+\mathbf{v} \operatorname{div} \mathbb{E}+\operatorname{curl} \operatorname{VC} \mathbb{E} \mathbf{v} \tag{6}
\end{equation*}
$$

and a similar expression for the magnetic current,* while the righthand sides of r., ir., containing only the instantaneous values of line integrals, remain obviously unaffected by the Hertzian requirement. The distribution of $\mathfrak{f t t}$ being supposed solenoidal, as
previously, the second term in the above expression is absent in the magnetic current. Thus, transferring the curl-terms of the currents to the right-hand sides, we obtain the required equations

$$
\left.\begin{array}{l}
\frac{\partial \mathbb{E}}{\partial t}+\rho \mathbf{v}=c \cdot \operatorname{curl}\left(\mathbf{M}-\frac{\mathrm{I}}{c} \mathrm{~V} \mathbb{C} \mathbf{v}\right)  \tag{HH}\\
\frac{\partial \mathfrak{A t}}{\partial t}=-c \cdot \operatorname{curl}\left(\mathbf{E}-\frac{\mathrm{I}}{c} \mathrm{Vv} \mathbb{V} t\right) .
\end{array}\right\}
$$

Heaviside calls VEEV/c the 'motional magnetic force' and Vvefl $/ c$ the 'motional electric force,' considering them as a kind of impressed forces.

In what we have called Maxwell's equations, the former of these 'motional forces' and the convection current $\rho \mathbf{v}$ are absent ; otherwise they are as $(\mathrm{HH})$; thus

$$
\left.\begin{array}{l}
\frac{\partial \mathbb{E}}{\partial t}=c \cdot \operatorname{curl} \mathbf{M}  \tag{Mx}\\
\frac{\partial \mathfrak{A t}}{\partial t}=-c \cdot \operatorname{curl}\left(\mathbf{E}-\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{v} \mathbb{C} t\right) \cdot
\end{array}\right\}
$$

The connexions between $\mathfrak{E}, \mathfrak{d}$ and $\mathbf{E}, \mathbf{M}$ are as in (3), except that $K, \mu$ may undergo continuous variations due to the strain of the material medium. Also, div $\mathfrak{f t t}=0$, as in (3). Notice, in passing, that the first of ( HH ) gives

$$
\frac{\partial \rho}{\partial t}+\operatorname{div}(\rho \mathbf{v})=0
$$

or

$$
\frac{d \rho}{d t}+\rho \operatorname{div} \mathbf{v}=0,
$$

where $d_{\rho} / d t=\partial_{\rho} / \partial t+(\nabla \nabla) \rho$ is the variation at an individual point of the body. Now, divv being the cubic dilatation, per unit time and per unit volume, the last equation may at once be written

$$
\frac{d}{d t}(\rho d \tau)=0,
$$

where $d \tau$ is an individual volume-element of the material medium, i.e. an element composed always of the same particles. Thus the charge $\rho d \tau$ of any such element remains invariable, being attached to it once and for ever. The charge, being preserved in quantity, moves
with the body. In this respect it behaves like the mass, according to classical mechanies. As regards the equations (Mx), they must be considered as referring to the particular case of an uncharged body; Maxwell happened not to consider explicitly charges in motion; otherwise he would doubtless have brought in the term $\rho \mathbf{v}$.

Now, both of these systems of equations, (Mx) as well as (HHP), are in full disagreement with experience, especcially with optical experience, terrestrial and astronomical, ici. with experiments on the propagation of electromagnetic waves (light) in bodies moving relatively to the observer, and also in bodics moving with the

body move as a whole with uniform translational velocity $\nabla$ with respect to $S$, and let plane waves of light be propagated in it along the positive direction of $\nabla$ (Fig. 3). If the unit-vector $i$ be the wave normal, concurrent with the propagation, then $\nabla=$ vi. Let $\mathfrak{v}$ be the scalar velocity of propagation of the waves, when the material medium

[^12]ationary in $S$, and $\mathfrak{b}$ their velocity of propargation, as judged the $S$-standpoint, when the medium is moving with its actual city. What is the relation between $b$ and $b^{\prime}, v$ ? If we were erned with waves of sound, instead of light waves, then $\mathfrak{b}$ would imply the sum of $\mathfrak{b}^{\prime}$ and of the whole $v$; the waves would be ely dragged by the medium, say air or water, with its full city. But the case before us is different. Write, generally,
\[

$$
\begin{aligned}
& \mathfrak{v}=\mathfrak{b}^{\prime}+\kappa v \\
& \kappa=\frac{\mathfrak{b}-\mathfrak{v}^{\prime}}{v} ;
\end{aligned}
$$
\]

$\kappa$, whatever its value, will be what is called the dragging coemolont, cating the fraction (if it happens not to be the whole) of the ium's velocity conferred upon the waves. What is, then, the ging coefficient in the case of clectromagnetic, and enpecially of inous waves?
ccording to (HH) it is, obviously, equal to unity. 'To sere his ave no need to integrate these differential equations,* hut simply emember Hertz's interpretation of the laws 1., ti., which furnished with these equations (p. 30). lior according to that interpretaand extension, of 1. , 1r., the electromagnetic disturbanees behave ively to the material medium (gencrally in carlh of its demems, in the present case, of rigid translation, throughout the whole ium) just as if it were stationary. Hence, on the: ground of sical kinematics of course, the velocity of the medium is simply ed to that of the waves, preecisely as in the case of soumed. Thus, , according to ( HH ).
et us now see what is the value of the drageing corfficiont rding to (Mx). Take the simplest case of an isotrupic mediom ;

$$
\mathfrak{v}^{\prime}=\stackrel{C}{K}^{\prime}
$$

$e$, by the way, $\mu=x$ for light waves. Measuring $x$ atong $i$ in the
 tion of the argument $x-\mathfrak{b}$, so that $\mathfrak{b}$ will be the velocity of
hough the reader, to satisfy himself, may do so. I'rocectling similaty as in ase of (Mx), worked out in Note 3 at the end of this chapter, lie will ackn that $\mathfrak{b}=\mathfrak{b}^{\prime}+v$.

or

$$
\begin{equation*}
\mathfrak{b}=\mathfrak{v}^{\prime}\left(\mathrm{x}+\frac{1}{4} n^{2} \beta^{2}\right)^{\frac{1}{2}}+\frac{1}{2} v \tag{7}
\end{equation*}
$$

where $\beta=v / c$ and where $n=c / b^{\prime}$ is the index of refraction of the medium. Now, in all actual experiments, by means of which the dragging of light can be determined, $\beta$ is a small fraction, viz. $10^{-4}$ in the case of Airy's astronomical, and much smaller in that of Fizeau's terrestrial experiment, both to be considered later. Therefore terms of the order of $\beta^{4}$ can certainly be rejected, so that

$$
\mathfrak{v}=\mathfrak{b}^{\prime}+\frac{1}{2} v+\frac{1}{\delta} \leadsto c \beta^{2}
$$

and

$$
\begin{equation*}
\kappa=\frac{1}{2}\left(\mathrm{I}+\frac{n}{8} \beta\right) \text {; } \tag{7a}
\end{equation*}
$$

but here even the $\beta$-term may be safely omitted, so that finally

$$
\kappa \doteqdot \frac{1}{5} . *
$$

Thus, we have for the dragging coefficient according to (HH) and ( Mx ), respectively,

$$
\begin{align*}
& \kappa=1,  \tag{HH}\\
& \kappa \doteqdot \frac{1}{2} . \tag{Mx}
\end{align*}
$$

Now, both of these are radically wrong, the true one, i.e. that showing excellent agreement with experiment, being Fresnel's widely known dragging coefficient (coéficient d'entrainement)

$$
\begin{equation*}
\kappa=\mathrm{I}-\frac{\mathrm{I}}{n^{2}}, \tag{Frsnl}
\end{equation*}
$$

where $n$ is the index of refraction. It is, for more than one reason, worth our while to dwell here upon the interesting history of Fresnel's coefficient.

The phenomenon of stellar aberration, discovered by Bradley in r728, found its immediate explanation when the assumption was made that the light-waves do not share in the earth's orbital motion,

[^13]consequently, in the motion of the tube of the telescopse: led with air or empty). In fact, making this assumption, the tional formula
\[

$$
\begin{equation*}
\frac{v}{c}=\frac{\sin \phi}{\sin \theta} \quad \text { (sce Fig. 4) } \tag{s}
\end{equation*}
$$

\]

for $\theta=\pi / 2$,

$$
\begin{equation*}
\frac{v}{c}=\sin \phi \doteqdot \tan \phi, \tag{8ta}
\end{equation*}
$$

sily obtained by using the widely known analogy of a ship, in n pierced by a shot fired from a gun on the shore.
rmula (8) gave, from Bradley's observations ( $\phi=20^{\prime \prime} \cdot 44$ ) and the known velocity $v$ of the earth's motion ( 30 kilom. per second), ae for $c$, the velocity of propagation of light, which agreed very


Fitc. 4.
ly with that obtained by Romer in 1676 from ohservations of clipse of Jupiter's satellites. Thus (8) was verified. 'T'u stateare facts, it would have been enough to say simply that the tulne e telescope, or the air contained in it, does not carry with it the coming from the star, whatever it may consist in (corpustles or 5). But to make the statement more tangible, it has been suid the 'corpuscles' or the 'acther,' respectively, do not share in elescope's motion. Whereas aberration was explained by its verer in terms of the corpuscular theory (each corpuscle of light ponding then most immediately to the shot in the above gy), it was Young who first showed ( 180.4 ) how it may be ined on the wave-theory of light and on the hypothesis that the $r$ 'pervades the substance of all material boolies with litue or m , ance, as freely perhaps as the wind passes through a grove of

trees.' 'This pinturemper analugy liturd altugether the case of air, which behaves wery narly like a vacum, hat not glass or water, for which the 'growe of trew' had to he rplated by a rather dense thicket. Hut any rate the alone words of loung hit very near the truth.

To pat it shorty, in the tave of air the drageng is mil, or nearly ser, $\kappa$; o.

Hut the case is different for untically demser media, having, for light of a given fretpeney, an index of refraction $n$, sensibly different from tuity. For if o were nil alse for such media, we should have to replate $c$ in ( 8 ) by the sualler velomity of propagation $c / a$, so that the angle of aberration weotd he differen for optically different media, whereas it has beon proved evperimentally to be just the same as in the case of air. More wernealy, Arago comeluded from his experimonts on the light of stars that the carth's motion has no sensible influene on the refraction (and refletion) of the rays emitted by these light shures, fro that the ray, toming from a star behave, say, in the case of a prisu or a dah of ghas, prectisely as they would if the star were situated at the prime in which it apmars to us in consequence of ordinary Dradleym (air tolewrox aberration, and the earth were at rest relatively to the sar. Araze himself tried to explain this result of his experiments on the conpustular theory, and on the supplementary hopuhesis that the sourre of light inpress upon the roppuscles an imhing of difieront velowition, and that out of these none but these codowed with a certain velocity $( \pm .01 \%)$ have the power of exciting our organ of sight. Hut this strange hypothesis
solu solic Her
 from other differutios, deses not sem to have satisfied its author. At any rate, Arago propered to frenel to investigate whether the above result of his ohservations cothd not be more casily reconciled with the wave theory of lisht.

It was in answer to this invitation that Fresen wrote in 1818 his celebrated letter to Arago 'on the influence of the earth's motion upon certain optical phenomem,' 1 in which he gives a beautiful






on of the problem, and which has since become one of the most supports of modern inquiry into the optics of moving media. appears for the first time his 'coéfficient d'entrainement,' ly mentioned above. Fresnel based the theory of aberration, ssociated matters, on the following hypothesis, which turned out a very happy guess indeed:
Fresnel supposed that the excess, and only the excess, of the her contained in any ponderable body over that in an equal ume of free space is carried along zuith the full velocith, 7 , of the $y$; while the rest of the aether within the space occupied by the dy, like the whole of the free aether outside, is stationary;-with pect to the fixed stars, of course.
s amounts * to supposing that the velocity of propagation of the waves is augmented only by the velocity of the 'centre of 'y' (centre of mass) of the whole mass of the aether contained e body. This velocity will, generally, be but a fraction of $\%$ $\mathrm{t} \kappa v$; then $\kappa$ will be what has above been called the dragging cient. Let $\rho_{0}$ be the density of the acther outside the and $\rho$ its density within the body; then, by fremell's hesis,

$$
\begin{gathered}
\left(\rho-\rho_{0}\right) v=\rho . \kappa v \\
\kappa=r-\rho_{0} / \rho .
\end{gathered}
$$

$e$ being the coefficient of elasticity of the auther within the and $e_{0}$ that of the free aether, the body's refractive index $n$ is by

$$
n^{2}=\left.\frac{e_{0}}{\rho_{0}}\right|_{\stackrel{e}{\rho}} ^{\rho}
$$

Fresnel's aether has throughout the same elasticity; within erable bodies and interplanetary space, so that ern and $/ \rho_{0}$.
us we obtain Fresncl's celebrated formula for the: dragking cient:

$$
\begin{equation*}
\kappa=\mathrm{r}-\frac{\mathrm{I}}{n^{2}} . \tag{Final}
\end{equation*}
$$

ice that considering the excess of the acthor, i.e. $f \cdots p_{0}$ prr rolume, as a permanent part of material boclies, it can lo. sain that the aether proper is not moved at all, that it is cutirely the letter in question, p. 63 r of reprint in V(o). II. of trimers cometrer.

uninfluenced by the moving bodies. Fresnel's theory is therefore
hough Fresnel's own reasoning, reprinted at the end of the ent chapter (Note 4), exhausts the subject entirely, let us yet il upon it a moment.
the aether behaved in optically denser bodices as in air, i.e. if e were no dragging at all, we shoukd have, by the ship and shot ogy, instead of (8),

$$
\frac{v}{c / n}=\frac{\sin \phi}{\sin \theta}
$$

reing the velocity of propagation of light in water, or in any other lium filling the tube of the telescope. 'Then Airy's experiment ld have given a positive result. But he obtained precisely the e $\phi$ as for air. This negative result suggested to him (at least as s usually represented in text-books) the supposition that the ter carries with it the acther' with only a certain part of its city, namely such that, in the above formula, we have to write: istead of $v$, where

$$
v=v / n
$$

hat

$$
\begin{aligned}
& \sin \phi \\
& \sin \theta
\end{aligned}=\frac{v}{c / n}=\vec{c}
$$

or air. In reality the process of compensation is not so simple as ; but in Airy's experiment the compensation-sensibly completroduced in a slightly different way. Comsidering a shab of water ing perpendicularly to its axis, and neglecting second order terms $\left.v^{2} / c^{2}=10^{-x}\right)$, you will casily obtain *

$$
\begin{equation*}
\frac{\sin \phi}{\sin \theta}=(v-z) c^{2} / h^{2}=(1-\kappa){ }_{i}^{v}, \tag{y}
\end{equation*}
$$

re, $v-\bar{v}$ being the relative velocity of the acther and wesoupe, /v has been written for the dragging cocfficient, as yot suppesed e unknown. Hence, to account for Airy's negative result, ix. to e (9) identical with (8), we have to write $(1-x) n n^{2}+1$, or

$$
\kappa \approx 1-\frac{1}{n^{2}},
$$

## Fresnel's formula.

ee, if necessary, for instance N. R. Campleill's, Woutron Rheitrient Theory
 e 28 , which are placed the wrong way ; correet alwy sme daslies wh p. 304 ead at the bottom of the page 'presence' insteal of 'presourc.' As regard u's experinent, amend the shocking anaclironism on p, 295 : 'Rizeau tried' -'to test the correctness of Airy's hypothesis'-1871).


Thus, Airy's negative result is perfectly accounted for by Fresnel's dragging coefficient, terms of the order of $10^{-s}$ being, of course, beyond the possibility of observation.
But Fresnel's formula found also, twenty years earlier, an immediate verification in Fizeau's optical interference-experiment with flowing water.* The arrangement of the apparatus which was used by Fizeau is seen at a glance from Fig. 5. Light from a narrow slit, $S$, after reflection from a plane parallel plate of glass, $A A$, is rendered parallel by a lens $L$ and separated into two pencils by apertures in a screen $E E$ placed in front of the tubes $T_{1}, T_{2}$ containing running water. The two pencils, after having traversed (towards the left hand) the respective columns of water, are focussed, by the lens $B$, upon a plane mirror $Z$, which interchanges their paths: the upper pencil returns towards $L$ by the tube $T_{2}$, the lower by $T_{1}$. On


Flt. 5.
emerging finally from the water, both pencils are brought, by $L$, to a focus behind the plate $A A$, at $S^{\prime}$ (and partly also at $S$ ). Here a system of interference fringes is produced which can be observed and measured in the usual way. Thus, each pencil traverses both tubes, $T_{1}$ and $T_{2}$, i.e. the same thickness of flowing water, say $l$. Moreover, the (originally) upper pencil is travelling always with, the other against the current. If, therefore, $v$ be the velocity of the water and $\kappa$ the dragging coefficient, the difference in light-time for the two pencils will be given by

$$
\Delta=l\left\{\begin{array}{c}
\mathrm{r} \\
\bar{c} / n-\kappa v \\
\bar{c} / n+\kappa v
\end{array}\right\},
$$

.where $n$ is the refractive index of water. Passing from stationary to flowing water, Fizeau observed a measurable displacement of the interference fringes, namely with $v=700 \mathrm{~cm} . / \mathrm{sec}$. ; and by reversing

[^14]the direction of the current of water the displacement of the fringes could be doubled. From the observed displacement it is easy to find the difference of times $\Delta$, and by equating it to the above expression of $\Delta$ to find the dragging coefficient $\kappa$ in terms of $l, n, z$, which can be measured. The result of Fizeau's experiment was that $\kappa$ is a fraction, sensibly less than unity. How much less, could not be ascertained with sufficient precision. Fizeau's experiment was therefore repeated in a form modified in several important points by Michelson and Morley * ( 1886 ), who found, for water (moving with the velocity of 800 cm . per second) at $18^{\circ} \mathrm{C}$., and for sodium light,
\[

$$
\begin{equation*}
\kappa=0.434 \pm 0.02 \tag{MM}
\end{equation*}
$$

\]

i.e. 'with a possible error of $\pm 0.02$.'

Now, $n$ being, in the case in question, equal to r.3335, Fresnel's formula gives

$$
\begin{equation*}
\kappa=\mathrm{r}-\frac{\mathrm{I}}{n^{2}}=0.438, \tag{Frsnl}
\end{equation*}
$$

a value agreeing very closely with Michelson and Morley's experimental result.

Thus, Fresnel's formula, deduced from what in our days may be deemed an assumption of naïve simplicity, proved to be in admirable conformity with experiment, like everything predicted by Fresnel in optics. His dragging coefficient has acquired a special importance in recent times, and every modern theory is proud to furnish his $\kappa$, which has become, in fact, one of the first requirements demanded from every theory of electrodynamics and optics of moving bodies which is being proposed. 'Agreeing with Fresnel' has become almost a synonym of 'agreeing with experience.'

Now Maxwell's and Hertz-Heaviside's equations for moving media, (Mx) and (HH), giving, as we have just seen, $\kappa \doteqdot \frac{1}{2}$ and $\kappa=\mathrm{r}$, or half and full drag, respectively, for any medium, be it as dense as water or glass or as rare as air, proved thereby to be in full disagreement with Fresnel, i.e. with experiment.
The first successful attempts to smooth out this discordance of ( Mx ) and (HH) from experiment, which-as has been mentionedmanifested itself also in the case of electromagnetic experiments properly so called, were made by H. A. Lorentz in 1892. The

* Michelson and Morley, American Jourrn. of Science, Vol. XXXI. p. 377 ; r886. See also A. A. Michelson's popular book, Light Waves and their Uses; Chicago 1907; p. 155.
theory proposed in a paper published in that year,* and which led with sufficient approximation to Fresncl's dragging coefficient, was then simplified and extended in 1895 , in a paper $\dagger$ which has since become classical.
Stokes' moving aether ( 8845 ) leading to serious difficulties, $\ddagger$
unic
scar
pres
Ma.
Lorentz decided in favour of Fresnel's inmovable, stationary atether, as the all-pervading electromagnetic medium.

Thus, Lorentz's theory, presently known widely as the Electron Theory, is, first of all, based on the assumption of a stationary, isotropic and homogencous acther. In calling it shortly 'stationary' (ruhend), Lorentz states expressly that to speak of the acther's 'absolute rest' would be pure nonsense, and that what he means is only that the several parts of the acther do not move relatively to one another (Essay; p. 4). In other words, Lorenta's aether is not deformed, it is subjected to no strain, and does not, consequently, execute any mechanical oscillations. And this being the case, it has, of course, no kind of elasticity, nor inertia or density. It is thus far less corporeal than Fresnel's acther. (One fails to see what properties, in fact, it still has left to it , hesides that of being a colourless seat (we camnot even say substratum) of the electromagnetice vectors $\mathbf{E}, \mathbf{M}$. And although Lorentz himself continues to tell us, in ryog, sis that he ' camot but regard the ether as endowed with a certain degree of substantiality,' yet, for the use he ever made of the atether, he might as well have called it an empty theatre of $\mathbf{E}, \mathbf{M}$, and their performances, or a purely geometrical system of reference, stationary with regard to the (or at least to some) 'fixed' stars. 'lhis acther, having been deprived of many of its precious propertics, was at any rate already so nearly non-substantial, that the first blow it had to sustain from modem research knocked it out of existence altogether,--as will be seen later. Still, substantial or not, for the theory of Lorentz we are now considering, it is something, namely its unique system of reference. So long, therefore, as it was thought that there is such an

[^15]gue system, Lorentz's all-pervading medium could continue its ty existence.
or this free aether, i.e. where it is not contaminated by the ence of ponderable matter, Lorentz assumes the exact validity of xzvell's equations, (4), i.e.
$$
\frac{\partial \mathbf{E}}{\partial t}=c \cdot \operatorname{curl} \mathbf{M} ; \quad \frac{i \mathbf{M}}{i t}=-i \cdot \operatorname{curl} \mathbf{E} ; \quad \operatorname{div} \mathbf{M} \cdots \mathbf{0},
$$
${ }^{1} \rho=\operatorname{div} \mathbf{E}=0$. (As to terminology, Lorent\% calls the above $\mathbf{E}$ dielectric displacement, and $\mathbf{M}$ the mas,netic forct.)
hen, to account for the optical and, more penerally, electro metic phenomena in moving ponderable matter, he has recours* sectro-atomism, an hypothesis alrendy employed (1882-1888) by se, Schuster, Arrhenius, Elster and (icitel, and others, and hater , by Helmholtz (1893) in his famous clectromagnetic the ory of ersion, and in various writings of Sir foseph Larmor. Acoorting Lorentz, matter by itself has no influences whatever on the stromagnetic phenomena: in this respect it behaves like the free eer. Only when and as far as matter is the satat of 'ions,' it -entz's, or electrons in modern terminology,* it mootities the atromagnetic field and its variations. In other words, Maxwells, ations, (4), are assumed to be strictly valicl not omly in the fere her, but also in all those portions of ponderable molecules in ch there is no charge, i.e. wherever $p=0$. Ant as to the yutestiom ather ponderable matter consists entirely of wectrical purtions arges) or not, Lorenta leaves it an open yuestiom, if 1 may iture an opinion, it was very wise of him not to have had Abraham's ambition to construct a purcly elertromagnetic 'eltbild,' as the Germans call it. (This remark will be under od better later on, when we shall see that, as lar ne we know, on the mass of the frec electrons, such ats the kuthorle raye of particles, may not be of purely electronagnetic origin.) "The part yed in Lorentz's theory by matter itself consists only in kerping electrons, or at least some of them, at or round certain placest, , restraining them from too wide excursions. Haxwell's equations, written above for the free aether, are nodified only where
$$
\operatorname{div} \mathbf{E} \boldsymbol{E} \neq 0
$$
'Electron' is due to Johnstone Stoney ( $\mathrm{IS91}$ ). The distinction maste now veen 'ions' and 'electrons' does not concern us here : levenites, it in generally
wn from a host of popular writings.

i.e. where there is, at the time being, some electric charge or electricity, and where, moreover, the electricity is moving.* The 'modification is the slightest imaginable,' to put it in Lorentz's own words (Electron Theory, p. 12). If p be the velocity of electricity at a point, relatively to the aether, i.e. relatively to that system of reference, $S$, in which the free-aether equations (4) are valid, thon the left-hand member of the first of thesc equations, or the displacement current, is supplemented by the convection current, per unit area, i.e. by $\rho \mathbf{p}$, while the second and third equations remain unchanged.

Thus, Lorentz's differential equations, assumed to be valid exactly or microscotically $\dagger$ throughout the whole space, are

$$
\left.\begin{array}{l}
\frac{\partial \mathbf{E}}{\partial t}+\rho \mathbf{p}=c . \operatorname{curl} \mathbf{M}, \text { where } \rho=\operatorname{div} \mathbf{E}  \tag{г.}\\
\frac{\partial \mathbf{M}}{\partial t}=-c \cdot \operatorname{curl} \mathbf{E} ; \quad \operatorname{div} \mathbf{M}=0 .
\end{array}\right\}
$$

These have been since generally called the fundamental equations of the electron theory. They contain, of course, the equations for the free aether as a particular case, namely for $\rho=0$.

An important supplement to the above system of equations consists in the formula for the ponderomotive force 'acting on the electrons and producing or modifying their motion,' which, guided by obvious analogies, Lorentz assumes to be, per unit volume,

$$
\begin{equation*}
\mathbf{P}=\rho\left[\mathbf{E}+\frac{\mathbf{I}}{c} V \mathbf{p} \mathbf{M}\right], \tag{ri.}
\end{equation*}
$$

or, per unit charge,

$$
\begin{equation*}
\mathbf{E}^{\prime}=\mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{VpM} . \tag{ro}
\end{equation*}
$$

This 'force' is supposed to be exerted by the aether on electrons or matter containing electrons. Vice versa, as Lorentz states it expressly, matter, whether containing electrons or not, exerts no action at all on the aether,--since the aether has already been supposed to undergo no deformations, etc. Of course, Lorentz's aether is massless as well. Lorentz tells us, with emphasis, not to

[^16]; in even the notion of a 'force on the acther.' It is true-me -that this is against Newton's third law (action reaction), , as far as I see, nothing compels us to clevate that proposition fundamental law of unlimited validity ' ( $E$ ssery, 1 1. 28). at there is no need to keep in mind all these, and similar, reas and verbal explanations,- ©specially as the absence of force te free acther is seen from (ni.) at a ghanes, by puting $p$
It is perfectly sufficient to state that the hasis of torenta's eory is entirely contained in the above (microsiopicall! valid) fuations (I.), (In),* all other things being ohtained from these fuations by more or less pure deduction, withont new hyper. eses. $\dagger$
otice, in passing, that (r.) is not a complete system in the sense de word explained in Chare I. For to trace the deetromagnetio ory, not only $\mathbf{E}_{0}, \mathbf{M}_{0}$ for $t=0$ and for the whole spare, hot atso o p for all values of $t$ must be given. In (f.) we have, (essmially, twin or equations of the first order for three veetors $\mathbf{E}, \mathbf{M}, \mathrm{p}$, and ther ula (ir.) does not complete the system, siner, on further reatareh, es not lead to an eçuation of the form ip/at $\Omega(E, M, p)_{1}$ b hat ne most favourable case to an integral e:gution entending wer tain interval of time, gencrally finite, but sometimes indefinitely onged. But this 'incompleteness' is men disadrantage in (t.), especially for the purpose of macernsoppie treathem, in whid isted Lorentz's main object of constructing these equmionts. ne equations assembled in (t.), which, together with the formula he ponderomotive foree, have been received inte the domain odern Relativity, as will be seen later, can be cusily ondensen a single quaternionic equation. lirme of all, put
\[

$$
\begin{equation*}
\mathbf{B}=\mathbf{M}-\iota \mathbf{I} \tag{11}
\end{equation*}
$$

\]

re $\iota=\sqrt{-1}$ ), and call it the electromagnetio biveotor. Alsw write, onvenience,

$$
15 u \%
$$

 rigid aether and deducel the equations in question from the prameptr of action. (Aether and Matter, (Cambribge, sgoo.)

 vectors or vector-fields.


Then, the first and third, and the second and fourth of (r.) coalesce respectively into the bivectorial equations

$$
\frac{\partial \mathbf{B}}{\partial l}+\operatorname{curl} \mathbf{B}=\frac{\mathbf{x}}{c} \rho \mathbf{p}
$$

and

$$
\operatorname{div} B=-t \rho ;
$$

or, in Familton's symbols,

$$
\begin{aligned}
\frac{\partial \mathbf{B}}{\partial l}+V \nabla \mathbf{B} & =\frac{I}{c} \rho \mathbf{p}, \\
\mathrm{~S} \nabla \mathbf{B} & =-(\nabla \mathbf{B})=-\operatorname{div} \mathbf{B}=\iota \rho .
\end{aligned}
$$

Acd up, and remember that the full quaternionic 'product' of the Familtonian $\nabla$ and of the bivector $\mathbf{B}$ is
then

$$
\nabla B=V \nabla B+S \nabla B ;
$$

$$
\frac{\partial \mathrm{B}}{\partial l}+\nabla \mathrm{B}=\rho\left(\iota+\frac{\mathbf{r}}{c} \mathbf{p}\right) .
$$

Next, introduce the operator

$$
\begin{equation*}
D=\frac{\partial}{\partial l}+\nabla \tag{13}
\end{equation*}
$$

which will turn out to be of fundamental importance for our subsequent relativistic considerations, and the quaternion

$$
\begin{equation*}
C=\rho\left(\imath+\frac{\mathbf{r}}{c} \mathbf{p}\right), \tag{14}
\end{equation*}
$$

which we may call the current-quaternion. Then the last equation becomes

$$
D \mathbf{B}=C
$$

(1. $a$

Thus, the four vectorial equations in (1.) coalesce into a singl, quaternionic equation (I. $a$ ), whose form will be most convenient fo relativistic electromagnetism. It is scarcely necessary to say tha what we have done here has nothing to do with Relativity. We ar not as yet so far. (1. a ) is simply a formal condensation of th fundamental electronic equations ( I .).

What we are mainly concerned with in the present chapter the macroscopic or average result of these equations and of the for formula (ir.). But before passing to consider Lorentz's macroscop equations, it will be good to dwell here a little upon the exact
microscopic formulae (r.), (ri.), and some of their immediate and most important consequences.

First, as regards the conservation of energy, multiply the first of (r.) by $\mathbf{E}$ and the third by $\mathbf{M}$, both times scalarly. Then, remembering that, by (II.), $\rho(\mathrm{Ep})=(\mathrm{Pp})$, and, by vector algebra,

$$
(\mathbf{E} \operatorname{curl} \mathbf{M})-(\mathbf{M} \operatorname{curl} \mathbf{E})=-\operatorname{div} \mathrm{VEM},
$$

the result will be

$$
\begin{equation*}
-\frac{\partial u}{\partial t}=(\mathbf{P p})+\operatorname{div} \mathfrak{p}, \tag{15}
\end{equation*}
$$

where

$$
\begin{equation*}
u=\frac{1}{2}\left(E^{2}+M^{2}\right) \tag{16}
\end{equation*}
$$

and
朋 $=c \mathrm{VEM}$.
Now, $(\mathbf{P p})$ is the activity of the ponderomotive force or the work done 'by the ether on the electrons' per unit time, and unit volume. Thus, by (1.5), the principle of conservation of energy will be satisfied for every portion of space, however small, if $u$ is interpreted as the density, and at the same time 7 as the flux, of electro-
 solenoidal distribution need not detain us here. 眈 is widely known as the poynting vector, in commemoration of the fact that this vector and the corresponding conception of the flow of energy were first formulated by Poynting (1884). Thus we see that the density and the flux of electromagnetic energy, given by (16) and (17), are in Lorentz's theory precisely as in Maxwell's and Hertz-Heaviside's theory.

Next, as regards the ponderomotive force $P$, in comparison with that of Maxwell as expressed by his electromagnetic stress, use the first and third of the fundamental equations (I.); then (II.) will become

$$
\mathbf{P}=\rho \mathbf{E}-\mathrm{VE} \operatorname{curl} \mathbf{E}-\mathrm{V} \mathbf{M} \operatorname{curl} \mathbf{M}-\frac{\mathbf{I}}{c} \mathrm{~V} \frac{\partial \mathbf{E}}{\partial t} \mathbf{M}-\frac{\mathbf{1}}{c} \mathrm{~V} \mathbf{E} \frac{\partial \mathbf{M}}{\partial t},
$$

or, introducing the Poynting vector,

$$
\begin{equation*}
\mathbf{P}=\rho \mathbf{E}-\mathrm{V} \mathbf{E} \operatorname{curl} \mathbf{E}-\mathrm{V} \mathbf{M} \operatorname{curl} \mathbf{M}-\frac{\mathrm{x}}{c^{\frac{2}{2}}} \frac{\partial \nmid}{\partial t} . \tag{18}
\end{equation*}
$$

This is the expression of Lorentz's force, equivalent, in virtue of (r.), to the original expression (i.). Now, Maxzeell's ponderomotive force, per unit volume, is given by

$$
\mathbf{P}_{\mathrm{Mx}}=\rho \mathbf{E}-\mathrm{V} \mathbf{E} \operatorname{curl} \mathbf{E}-\mathrm{V} \mathbf{M} \operatorname{curl} \mathbf{M} .
$$

This is the resultant of Maxwell's well-known electromagnetic stress i.e.

$$
\begin{gather*}
\mathbf{f}_{n}=u \mathbf{n}-\mathbf{E}(\mathbf{E n})-\mathbf{M n}(\mathbf{M n}),  \tag{20}\\
\mathbf{P}_{\mathrm{M} \mid \mathrm{x}}=-\mathbf{i} \operatorname{div} f_{1}-\mathbf{j} \operatorname{div} \mathrm{f}_{2}-\mathbf{k} \operatorname{div} f_{3}, \tag{2I}
\end{gather*}
$$

$f_{n}$ being the pressure* per unit area, on a surface element whose unit normal is $n$, and $f_{1}, f_{2}, f_{3}$ meaning the same things as $f_{n}$ for $\mathrm{n}=\mathrm{i}, \mathrm{j}, \mathrm{k}$ respectively. We do not stop here to show the equivalence of ( x 9 ) and (2I), for we shall have an opportunity to do so later. What concerns us here is the comparison of Lorentz's with Maxwell's ponderomotive force. From ( r 8 ) and ( 19 ) we see that the former is

Maxwell's force on the free wether, i.e. for $\beta=0$, is, by ( 19 ) and the system ( I. ), which in this case coincides with Maxwell's equations,
i.e.

$$
\begin{align*}
& P_{M x}=\frac{1}{c} V E M+\frac{1}{c} V E M, \\
& P_{M \times x}=\frac{1}{c} \frac{\partial y}{\partial t}, \text { for } p=0 . \tag{190}
\end{align*}
$$

'Thus, in a variable field, Maxwell's ponderomotive force on the free ather is, generally, different from zero. 'The supposed existence of such a force, which has been treated on various occasions by Heaviside, suggested to Helmholt\% the argument of his last paper, namely an investigation of the possible motions of the free aether. $\dagger$ On the other hand, Lorenta's force on the free acther is always wil, according to his fundamental formula (in); as has been already remarked, he forbids us even to talk about a force on the aether, since its elements are supposed once and for ever to be immovable. According to (22) the Maxwellian force on the aether is just compensated by Lorentz's supplementary term - $-\frac{1}{2}$, $1 \partial t$. In using the Maxwellian stress $f_{n}$ in his theory, Lorentz considers it, of course, as a system of 'merely fictitious tensions' (cf. Essay, p. 29). In

* Iressure proper heing counted pwsitive, and tension proper negative.
+11. v. Melmholty, folgerumgen aus Maxwells Theorie über die Bervegungen des reinen Aethers; Berl. Sitwher., July 5, 1893; Wied. Anh., Vol. LIII. p. 135, 1894.

Iaxwell's theory the ponderomotive actions observed in clectric: and agnetic fields were physically accounted for by the tensions and ressures of the aether. But Lorentz, in order to be consistent, voids considering the 'aether tensions' as something physical, since:' nese would mean forces exerted by the different parts of the ather a one another. Thus, the Maxwellian stress is to him but a conenient instrument for calculation.
Returning to the general case, $\rho \neq 0$, Lorentz's ponderomotive rce (ir.) may be written, by (22) and (21),

$$
P=-i \operatorname{div} f_{1}-j \operatorname{div}_{2}-\mathbf{k} \operatorname{div} f_{8}-\frac{1}{c^{2}} \frac{d}{\partial t}
$$

$$
\therefore \quad(3,3)
$$

thus consists of two parts, the first of which is deducible from the Laxwellian stress, while the sccond, forcign to Maxwell's theory, is ven by the negative time-rate of local change of the vecter 1 那 $/ c^{2}$. It this second term which always compensates the Máxwettion action a the pure aether.
Finally, to obtain Lorentz's resultant force

$$
\mathrm{II}=\int \mathrm{P} d \tau
$$

1 the whole system of electrons ( $\tau$ being any volune containing all e electrons), use the expression (23), and observe that

$$
\int \operatorname{div} f_{i} d \tau=\int\left(\operatorname{nf}_{i}\right) d r, \quad i=1, z, 3_{1}
$$

here n is the outward unit normal of the surface or earlosing the gion $\tau$. Also remember that

$$
\begin{equation*}
\mathrm{i}\left(\mathrm{f}_{2} \mathrm{n}\right)+\mathrm{j}\left(\mathrm{f}_{2} \mathrm{n}\right)+\mathbf{k}\left(\mathrm{f}_{8} \mathrm{n}\right) \cdots \mathrm{f}_{n}, \tag{24}
\end{equation*}
$$

ace the Maxwellian stress is irrotational or selfeconjugate, Then e result will be

$$
\begin{equation*}
\Pi=\int \mathrm{f}_{n} d r-\frac{d}{d t} \int \frac{\mathrm{x}}{\sigma^{2}} \nexists d d, \tag{25}
\end{equation*}
$$

being supposed fixed in the aether, i.e. relatively to the framework in which the fundamental equations are to be valid. Furmula (25) ates simply the same thing for the whole system, contained in $r$, ich is expressed by (23) for each of its elements. Of course, in ssing from (23) to (25), the continuity of the veetor $\boldsymbol{f}_{n}$ (or at hemst its components normal to surfaces of discontinuity, if there be auy) s.r.
has been tacitly assumed throughout $\tau$.* The last formula, again, may be written:

$$
\mathrm{\Pi}=\mathrm{\Pi}_{\mathrm{Mx}}-\frac{d}{d t} \int \frac{\mathrm{x}}{c^{2}} \text { 的 } d \tau,
$$

which needs no further explanation. Now, as the mathematicians say, let $\sigma$ expand to infinity, or at least so that, $E, M$ decreasing in the usual way as $\mathrm{I} / r^{2}$, the surface integral may vanish. Then

$$
\Pi_{M x}=0,
$$

while

$$
\begin{equation*}
\Pi=-\frac{d \mathbf{G}}{d t} \tag{26}
\end{equation*}
$$

where the vector $\mathbf{G}$ is defined by

$$
\begin{equation*}
\mathbf{G}=\frac{\mathrm{I}}{c^{2}} \int \mathfrak{\exists} d \tau, \tag{27}
\end{equation*}
$$

and is called the electromagnetic momentum.
Thus Maxwell's resultant force is strictly nil, satisfying Newton's third law (actio est par reactioni), while Lorentz's resultant force is generally different from zero, against the third law,-a result which has been already stated in a slightly different form. Thus Maxwell's theory, admitting an action on the pure aether, did, while Lorentz's theory, denying it, does not satisfy Newton's third law. But, as was observed by Lorentz himself, there is nothing to compel us to universalize that law of Newtonian mechanics. At first, Poincaré tried to use this as an argument against Lorentz's theory ; $\dagger$ but he soon gave it up. This was to be only one of a whole series of sacrifices, and not the greatest one, made by modern physicists.
Similarly, the resultant moment of the ponderomotive forces,

$$
\begin{equation*}
\Omega=\int \operatorname{VrP} d \tau \tag{28}
\end{equation*}
$$

where r is the vector drawn to any point of the field from a point $O$ fixed in the aether, or fixed relatively to $S$, may be easily put into the form

$$
\Omega=\int \operatorname{Vrf}_{n} d \sigma-\frac{d}{d t} \int \frac{\operatorname{Vr}^{2}}{c^{2}} d \tau
$$

[^17]Thus, for the whole space, and with the usual assumption as to the behaviour of $E, M$ at infinity,
and

$$
\Omega_{\mathrm{Mx}}=0,
$$

$$
\begin{equation*}
\Omega=-\frac{d \mathbf{H}}{\dot{d} t}, \tag{29}
\end{equation*}
$$

where

$$
\mathrm{H}=\frac{\mathrm{I}}{c^{2}} \int \operatorname{Vr} \beta d \tau
$$

is called the electromagnetic moment of momentum. Its analogy to the ordinary, mechanical, moment of momentum

$$
\Sigma m \mathrm{Vrv}
$$

is obvious. So is also the analogy of the above $\mathbf{G}$ with the ordinary momentum

$$
\Sigma m \mathrm{v}
$$

and the corresponding interpretation of (26) and (29). Both $\mathbf{G}$ and $\mathbf{H}$ are so constructed as if the aether contained (electromagnetic) momentum in each of its elements amounting to

$$
\begin{equation*}
\mathbf{g}=\frac{1}{c^{2}} \text { 刖 }=\frac{\mathrm{r}}{c} \mathrm{VEM} \tag{30}
\end{equation*}
$$

per unit volume.
So much as regards the chief consequences of the fundamental formulae (r.) and (II.).

Now for Lorentz's macroscopic equations. These are obtained from (r.), (II.) by averaging over 'physically infinitesimal' regions of space. Lorentz calls a length $l$ 'physically infinitesimal' (in distinction from what is called 'mathematically infinitesimal') if the values of any observable magnitude obtaining in two points distant $l$ from one another are sensibly equal to, i.e. indiscernible from, one another. Molecular, and, a fortiori, electronic, dimensions and mutual distances of molecules constituting a ponderable medium, are assumed to be small fractions of $l$. Let $\psi$ be any magnitude, scalar or vectorial. Round a point $P$ draw a sphere of physically infinitesimal radius ; let $\tau$ be the volume of this sphere. Then

$$
\frac{\mathrm{I}}{\tau} \int \psi d \tau
$$

is called the 'mean value of $\psi$ at $P$,' and is denoted by $\bar{\psi}$. If $\psi$ be any of the magnitudes involved in the fundamental (microscopic)
equations, as for instance $\rho$ or $\mathbf{M}$, then $\bar{\psi}$ is what is macroscopically observable.

We cannot reproduce here the details of the process of averaging based upon the above fundamental notion,* but shall simply write down the resulting macroscopic equations, limiting ourselves to the case of a perfectly transparent (i.e. non-conducting), non-magnetic ponderable medium, and leaving out of account dispersion. We must, however, explain first the meaning of the symbols involved in these equations.

Assuming that the molecules of the ponderable medium or body contain electrons, $\dagger$ to which belong certain positions of equilibrium within the individual molecules, Lorentz supposes their displacements from these positions, $\boldsymbol{q}$, and their velocities relative to the corresponding molecule,

$$
\dot{\mathbf{q}}=d \mathbf{q} / d t
$$

to be infinitesimal. In other words, he neglects the squares and products of $\mathbf{q}, \dot{\mathbf{q}}$, or any of their components in presence of their first powers. Notice that the only part played by the molecules of ponderable matter consists here in restraining the electrons, i.e. in keeping them near certain positions. For, as has already been remarked, one of Lorentz's fundamental assumptions is, that matter by itself, apart from electricity, behaves like the free aether, its presence having no influence whatever upon the electromagnetic field.

Let $e$ be the charge of an electron which has experienced the displacement $q$, as explained above. Then Lorentz brings in the notion of electrical moment, not unfamiliar to older theories, defining this vector to be, per unit volume, the average of eq, i.e.

## $\overline{e q}$.

Taking the sum of this and of the average of our above $\mathbf{E}$, Lorentz introduces the macroscopic vector

$$
\begin{equation*}
\mathfrak{G}=\overline{\mathrm{E}}+\overline{e q} \tag{3I}
\end{equation*}
$$

[^18]which he calls the dielectric polarization.* Thus, in the free aether © reduces to $\overline{\mathbf{E}}$, and generally $\mathbb{E}$ is what Maxwell called the dielectric displacement.

Next, the macroscopic magnetic force is defined to be the average of our above $\mathbf{M}$, i.e. $\overline{\mathbf{M}}$, instead of which, however, we shall write shortly M.

Finally, the macroscopic electric force is introduced, being defined as the average of $\overline{\mathbf{E}}^{\prime}$, i.e. of the ponderomotive force per unit charge, as given by the formula (土0). Instead of $\overline{E^{\prime}}$ we shall, again, write more conveniently $\mathbf{E}^{\prime}$. Thus Lorentz's macroscopic electric force will be

$$
\begin{equation*}
\mathbf{E}^{\prime}=\overline{\mathbf{E}}+\frac{\mathbf{x}}{c} \overline{\mathrm{Vp} \mathrm{\mathbf{M}}} \tag{2}
\end{equation*}
$$

Notice that here $\mathbf{p}$ means the resultant velocity of an electron, i.e. the vector sum of its velocity relatively to the molecule in question and of the velocity of the ponderable body as a whole, say $\boldsymbol{\nabla}$, 'relatively to the aether,' so that $\mathbf{p}=\dot{q}+\boldsymbol{\nabla}$.

With these meanings of the symbols, Lorentz's macroscopic equations for a transparent, non-magnetic, ponderable body, moving with constant $\dagger$ velocity $\nabla$ 'through the stagnant aether,' i.e. relatively to the framework $S$, are as follows (Essay, p. 76):

$$
\left.\begin{array}{rl}
\frac{\partial \mathbb{E}}{\partial t} & =c \cdot \operatorname{curl} \mathbf{M}^{\prime} ;  \tag{33}\\
\frac{\partial \mathbf{M}}{\partial t} & =-c \cdot \operatorname{curl} \mathbf{E}^{\prime} ; \\
\quad \operatorname{div} \mathbf{M}=0 \\
\mathbf{M}^{\prime} & =\mathbf{M}-\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{V} \mathbf{E}^{\prime} \\
K^{\prime} \mathbf{E}^{\prime} & =\mathbb{H}+\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{V} \mathbf{M} .
\end{array}\right\}
$$

Here the system of coordinates involved in div and curl, is rigidly attached to the ponderable body, thus sharing in its motion through the aether. But the time $t$ is the same as in the fundamental equations (I.); obviously, therefore, $\partial / \partial t$ is the time rate of change (for constant values of those coordinates, i.e.) at a fixed point of the body, not of the aether or of $S$.

* The above © $\mathbb{E}$ is Lorentz's 코.
$\dagger$ Constant in space and time, that is to say for a body having a uniform purely translational, rectilinear motion.

The second of (33) is an obvious expression of the (assumed) absence of macroscopic charge, i.e. of $\bar{\rho}=0$. In the more general case of a sensibly charged body we should have div $\mathbb{E}=\bar{\rho}$, where $\bar{\rho}$ is the observable density. As to $K$, appearing in the last of (33), it is a linear vector operator in crystalline, and a simple scalar coefficient in isotropic bodies, known as the 'dielectric constant' or permittivity, and depending in a complicated way on the distributional properties of the electrons. The numerical value of $K$ in an isotropic, and its principal values, $K_{1}, K_{2}, K_{3}$ in a crystalline body, are not constant, of course, but vary with the period $T$ of the incident light- or, generally, electromagnetic oscillations. However, to avoid unnecessary complication, we may think here of the simple case of homogeneous light, of a particular kind (colour). Then $K$, or $K_{1}, K_{2}, K_{3}$, are constants, whose numerical values are to be considered as deduced from the observable refractive properties of the body with regard to light of that particular kind. In case of isotropy we have to write $K=n^{2}$, if $n$ be the corresponding index of refraction.*

Notice that (33) contains, besides the solenoidal conditions for © $\mathbb{E}$ and $\mathbf{M}$, four vector equations for as many vectors,

$$
\mathfrak{G}, \mathbf{M}, \mathbf{E}^{\prime}, \mathbf{M}^{\prime},
$$

the velocity of motion $\boldsymbol{\nabla}$ being given. And since the differential equations are of the first order with regard to $t$, the electromagnetic history of the whole medium is determined by its initial state, say, by $\mathbb{C H}_{0}, \mathbf{M}_{0}$ given for $t=0$.

It must be kept in mind that, to obtain the system of equations (33) from the fundamental ones, Lorentz has consciously neglected not only various small terms concerning the minute influence of electrons, but also all terms of second order in $\beta$, or, to put it shortly, all $\beta^{2}$-terms, where

$$
\beta=\frac{v}{c} .
$$

This is especially true of the fifth of (33), which has been obtained from the more exact formula $\mathbf{M}^{\prime}=\mathbf{M}-\mathrm{V} \mathbf{\nabla} \overline{\mathbf{E}} / c$ by writing $\mathbf{E}^{\prime}$ instead of $\bar{E}$, and thus [cf. (32)] omitting $V \nabla \overline{\mathrm{VpM}} / c^{2}$, which is a $\beta^{2}$-term.

[^19]coefficient, $\kappa=\mathrm{r}-\frac{\mathrm{I}}{n^{2}}$. This, in fact, is a consequence of (33), when $\beta^{2}$-terms are neglected and when dispersion is not taken into account. For a dispersive medium that value of the index of refraction is to be taken which corresponds to the 'relative' period of oscillation, $T^{\prime}$ ',a concept to be explained further on. This gives a slight correction term, $-n^{-1} T \partial n / \partial T$ (Essay), p. IоI), where $n$ is the refractive index of the medium corresponding to the 'absolute' period $T$, i.e. the period of the oscillations emitted by the source, say, in Fizeau's experiment. Thus, Lorentz's formula is
\[

$$
\begin{equation*}
\kappa=\mathrm{I}-\frac{\mathrm{x}}{n^{2}}-\frac{\mathrm{I}}{n} T \frac{\partial n}{\partial T} . \tag{Lor}
\end{equation*}
$$

\]

For water, at $\mathrm{r} 8^{\circ} \mathrm{C}$., and for sodium light, this becomes

$$
\begin{equation*}
\kappa=0.45 \mathrm{I}, \tag{Lor}
\end{equation*}
$$

whereas Fresnel's value, and that obtained experimentally by Michelson and Morley, have been 0.438 and $0.434 \pm 0.02$ respectively. 'Thus Lorentz's dragging coefficient agrees with the experimental value (MM) quite as well as Fresnel's, especially if the 'possible error of $\pm 0.02$ ' be taken into account. In a word, Lorentz's equations give the right value of the dragging coefficient. And, from what has been said previously, it can be argued that these equations will also give correct results for all first order phenomena.

Next, putting $\nabla=0$, we see at once that (33) become

$$
\left.\begin{array}{rl}
\frac{\partial \mathbb{E}}{\partial t} & =c \cdot \operatorname{curl} \mathbf{M} ;  \tag{0}\\
& \operatorname{div} \mathbb{E}=0 \\
\frac{\partial \mathbf{M}}{\partial t} & =-c \cdot \operatorname{curl} \mathbf{E} ; \quad \operatorname{div} \mathbf{M}=0 \\
\mathfrak{E} & =K \mathbf{E}
\end{array}\right\}
$$

that is to say, identical with Maxwell's equations, for a stationary (non-magnetic) medium, (1), p. 24. Taking account of

* Proceeding, mutatis mutandis, similarly as in Note 3, concerning (HH). Another, more simple, method of obtaining the dragging coefficient is to apply Lorentz's 'theorem of corresponding states,' to be considered later.
magnetization-electrons, we would have, in the second and third equation, $\mathfrak{A l t}$ instead of $\mathbf{M}$, where $\mathfrak{A t t}=\mu \mathbf{M}, \mu$ being the permeability.

This is a very satisfactory result, for, as already mentioned, Maxwell's equations for stationary media, agreeing fully with experiment, have been able to stand even the severe criticism of the modern relativists, who have adopted them without the slightest modification whatever.
'Stationary' means, of course, in Lorentz's theory, fixed relatively to the aether.

In order to exhibit the properties of his equations, (33), in the general case of any constant $\mathbf{v}$, i.e. for a material medium having any uniform motion of rectilinear translation relative to the aether, Lorentz transforms these equations by introducing instead of the time $t$ a new variable of very remarkable properties. This, the socalled 'local time,' which was to become one of the most immediate forerunners of Einstein's relativistic theory, deserves a rather more extended treatment. It will occupy our attention in the next chapter.

## NOTES TO CHAPTER II.

Note 1 (to page 28). Let $\sigma$ be a surface of electromagnetic discontinuity of first order, for example ; that is to say, the vectors $\mathbf{E}, \mathbf{M}$ being themselves continuous across $\sigma$, let their space- and time-derivatives of first order be different in absolute value and direction on the two sides of the surface. Call one of its sides 1 , andl the other 2; draw the normal unit vector n from I towards 2 , and denote by $[a]$ the jump of any magnitude $\alpha$, i.e. the clifference $\alpha_{2}-a_{1}$. Then the so-called identical conditions, to be fulfilled in any case, are

$$
\begin{equation*}
[\operatorname{div} \mathbf{E}]=(\mathbf{n e}) ; \quad[\operatorname{curl} \mathbf{E}]=\operatorname{Vne} ; \tag{a}
\end{equation*}
$$

and the kinematical condition of compatibility, valid uncler the supposition that the surface is neither being split into two or more nor dissolved, is

$$
\begin{equation*}
\left[\frac{\partial \mathbf{E}}{\partial t}\right]=-\mathfrak{v e}, \tag{b}
\end{equation*}
$$

e being the same vector as in (a), characterizing the electrical discontinuity, and $\mathfrak{b}$ (an independent scalar) the velocity of propagation of $\sigma$, counted positively along $n$. Both $\mathfrak{e}$ and $\mathfrak{b}$ remain so far indeterminate, in numerical value and direction. Similarly, for the magnetic discontinuity,

$$
\begin{array}{cc}
{[\operatorname{div} \mathbf{M}]=(\mathbf{n m}), \quad[\operatorname{curl} \mathbf{M}]=\operatorname{Vnm},} & \left(a_{1}\right) \\
{\left[\frac{\partial \mathbf{M}}{\partial t}\right]=-\mathfrak{b m},} & \left(b_{1}\right) \tag{1}
\end{array}
$$

$m$ being a new vector and $\mathfrak{b}$ the same scalar as above, since the electric and magnetic discontinuities are supposed not to part from one another. (For the deduction of the above conditions see J. Hadamard's Leȩons sur la propagation des ondes et les équations de l'hydrodynannique, Paris, 1903, or, in vectorized form, my book on Vectorial Mechanics, London, Macmillan \& Co., r9r 3 ; also Annalen der Physik, Vol. XXVI., 1908, p. $75 \mathbf{1}$ and Vol. XXIX., 1909, p. 523.)

If $e, m$ are normal to $\sigma$, we have a longitudinal, and if tangential, a transversal discontinuity.

So far everything has been independent of any electromagnetic connections. Now use Maxwell's equations (4), with (41) ; since they are valid on both sides of $\sigma$, we have also

$$
\left[\frac{\partial \mathbf{E}}{\partial t}\right]=c[\operatorname{curl} \mathbf{M}], \text { etc. }
$$

and, using (a), (b) with their magnetic analogues,

$$
\left.\begin{array}{rlrl}
\frac{\mathfrak{V}}{c} \mathrm{e} & =\operatorname{Vmn} ; & & \frac{\mathfrak{V}}{c} \mathrm{~m}=V \mathrm{ne}  \tag{c}\\
(\mathrm{mn}) & =0 ; & & (\mathrm{en})=0 .
\end{array}\right\}
$$

Notice that if $\mathfrak{b}$ does not vanish, i.e. if there is propagation at all, the second pair of equations becomes superfluous, since it then follows identically from the first pair. Now, eliminating $m$ from the first pair of (c), we have

$$
\frac{\mathfrak{b}^{2}}{c^{2}} \mathrm{e}=\operatorname{VnVen}=\mathrm{e}-\mathbf{n}(\mathrm{en})
$$

n being a unit vector. But (en) $=0$; hence
and similarly

$$
\frac{\mathfrak{b}^{2}}{\mathfrak{c}^{2}} \mathrm{e}=\mathrm{e}
$$

$$
\frac{\mathfrak{b}^{2}}{c^{2}} \mathrm{~m}=\mathbf{m}
$$

Thus, if $e, m$ do not vanish, i.c. if there is at all a discontinuity,

$$
\begin{equation*}
\mathfrak{b}= \pm c \tag{d}
\end{equation*}
$$

that is to say, each element $d \sigma$ of the wave is propagated normally to itself with the velocity $c$. Q.E.D.

Notice that the sign of $\mathfrak{b}$, left undetermined in (d), due to the quadratic result of elimination, may be defined uniquely by means of the original pair of equations (c), which are linear in b. In fact, multiply the first scalarly by $e$ (or the second by $m$ ), then

$$
\mathfrak{b}=s(\mathrm{eVmn})=s(\mathrm{nVem}),
$$

where $s$ is a positive scalar, namely $c / e^{2}$. Thus, if $\mathbf{n}, \mathbf{e}, \mathrm{m}$ is a righthanded system, like the usual $\mathbf{i}, \mathbf{j}, \mathbf{k}$ then $\mathfrak{b}$ is positive, i.c. the sense of
propagation is that of $n$, and if $n, e, m$ is left-handed, then the propagation is along $-\mathbf{n}$. Thus, the sense of propagation coincides always with that of the vector

## Vem.

If $e$ points upwards and $m$ to the right, then the wave is propagated forwards. Notice the similarity with the sense of the flux of energy, or the Poynting vector, in relation to $\mathbf{E}, \mathbf{M}$,

$$
\text { 招 = } c \mathrm{VEM} .
$$

Finally, notice, in passing, that by the first pair of (c),

$$
c^{2}=m^{2}
$$

similarly to the known characteristic, $E^{2}=M^{2}$, of the usual 'pure' waves.
The above results may easily be extended to waves of discontinuity of any order.

Note 2 (to page 30). Take as a surface element the parallelogram constructed on two coinitial line elements $a, b$, composed always of the same particles, so that, $\mathbf{n}$ being its positive normal,

$$
\mathrm{n} d \sigma=\mathrm{Vab}
$$

Write, generally, $\boldsymbol{R}$ for $\mathbb{E}$ or $\mathfrak{f l}$. Then the induction through $d o$ will be given by the volume of the parallelopiped $\mathbf{R}, \mathrm{a}, \mathrm{b}$, i.e.

$$
(\mathbf{R n}) d \sigma=(\mathbf{R V a b})
$$

The current through $d \sigma$, say ( pn ) $d \sigma$, being the rate of change of this induction, is

$$
\begin{equation*}
(\mathbf{p n}) d \sigma=(\dot{\mathbf{R}} \mathbf{n}) d \sigma+(\mathbf{R} V \dot{\mathbf{a}} \mathbf{b})+(\mathbf{R V a} \dot{\mathbf{b}}) \tag{a}
\end{equation*}
$$

where the dots stand for individual variation. Thus

$$
\begin{equation*}
\dot{\mathbf{R}}=\frac{d \mathbf{R}}{d t}=\frac{\partial \mathbf{R}}{\partial t}+(\nabla \nabla) \mathbf{R} \tag{b}
\end{equation*}
$$

and [Vectorial Mechanics, Chap. V., formula (75)]

$$
\dot{\mathrm{a}}=\frac{d \mathrm{a}}{d t}=(\mathrm{a} \nabla) \nabla ; \quad \dot{\mathrm{b}}=(\mathrm{b} \nabla) \nabla
$$

Now, $\mathbf{i}, \mathbf{j}, \mathbf{k}$ being the usual right-handed system of mutually normal unit vectors, take a rectangular $d \sigma$, say

$$
\mathrm{a}=\mathbf{j} d y, \quad \mathbf{b}=\mathbf{k} d z
$$

and, consequently,
Then

$$
\mathbf{n}=\mathbf{i}, \quad d \sigma=d y . d z
$$

$$
\dot{\mathrm{a}}=d y \frac{\partial \mathrm{v}}{\partial y} ; \quad \dot{\mathrm{b}}=d z \frac{\partial \mathbf{v}}{\partial z}
$$

so that the sum of the last two terms in (a) will be

$$
\left(\mathbf{R} V \frac{\partial \mathbf{v}}{\partial y} \mathbf{k}\right) d \sigma+\left(\mathbf{R} V \mathbf{j} \frac{\partial \mathbf{v}}{\partial z}\right) d \sigma
$$

or, per unit area,

$$
R_{1}\left(\frac{\partial v_{1}}{\partial y}+\frac{\partial v_{3}}{\partial z}\right)-R_{2} \frac{\partial v_{1}}{\partial y}-R_{3} \frac{\partial v_{1}}{\partial z}=R_{1} \operatorname{div} \nabla-(\mathbf{R} \nabla) v_{1} ;
$$

hence, substituting (b) in the first term of (a) and remembering that

$$
\begin{gathered}
(\mathbf{p n})=(\mathbf{p} \mathbf{i})=p_{1}, \quad(\mathbf{R n})=R_{1}, \\
p_{1}=\frac{\partial R_{1}}{\partial t}+(\nabla \nabla) R_{1}+R_{1} \operatorname{div} \nabla-(\mathbf{R} \nabla) z_{1},
\end{gathered}
$$

with similar expressions for $p_{2}, p_{3}$ if $d \sigma$ be taken normal to $\mathbf{j}$ or $\mathbf{k}$ respectively. Thus the resultant current will be

$$
\mathbf{p}=\operatorname{current}(\mathbf{R})=\frac{\partial \mathbf{R}}{\partial t}+(\boldsymbol{\nabla} \nabla) \mathbf{R}-(\mathbf{R} \nabla) \mathbf{v}+\mathbf{R} \operatorname{div} \mathbf{v},
$$

or

$$
\begin{equation*}
\mathbf{p}=\operatorname{current}(\mathbf{R})=\frac{\partial \mathbf{R}}{\partial t}+\mathbf{v} \operatorname{div} \mathbf{R}+\operatorname{curl} \mathrm{VR} \mathbf{v} \tag{c}
\end{equation*}
$$

which is the required formula.
In the simplest case, considered on p. 33, in which the material medium moves as a whole with purely translational velocity $\nabla=v i$, we have to take only the first term of (a), so that in this case

$$
\begin{equation*}
\mathbf{p}=\frac{d \mathbf{R}}{d t}=\frac{\partial \mathbf{R}}{\partial t}+(\mathbf{v} \nabla) \mathbf{R}=\frac{\partial \mathbf{R}}{\partial t}+v \frac{\partial \mathbf{R}}{\partial x} . \tag{1}
\end{equation*}
$$

Note $\mathbf{3}$ (to page 34 ). Take $\mathbf{E}$, etc., proportional to an exponential function of the argument

$$
g(x-\mathfrak{b} t),
$$

where $g$ is an imaginary constant, as usual. Then

$$
\frac{\partial}{\partial t}=-g \mathfrak{b} ; \quad \nabla=\mathrm{i} \frac{\partial}{\partial x}=g \mathfrak{i},
$$

and, consequently, curl $=\mathrm{V} \nabla=g \mathrm{Vi}$. Introducing this in the equations ( Mx ), remembering that $\mathbf{\nabla}=\tau \mathrm{i}$ and omitting the common factor $g$, we obtain at once

$$
\begin{aligned}
-\frac{\mathfrak{b}}{c} \mathfrak{E} & =\text { ViM, } \\
\frac{\mathfrak{v}}{c} \mathfrak{E l t} & =\mathrm{Vi}\left\{E-\frac{v}{c} \mathrm{Vi} \mathfrak{A t}\right\} \\
& =\operatorname{ViE}-\frac{v}{c}\{\mathrm{i}(\mathfrak{A l i})-\mathfrak{y l t}\} ;
\end{aligned}
$$

but $\operatorname{div} \mathfrak{f l l}=0$ gives in the present case $(\mathfrak{G t i})=0$. Thus

$$
\begin{aligned}
\mathfrak{U} \mathfrak{E} & =c \mathrm{VMi} \\
(\mathfrak{b}-v)^{\prime} \mathfrak{A t} & =c \mathrm{ViE}
\end{aligned}
$$

and, the medium being isotropic,

$$
\begin{aligned}
(\mathrm{b}-v) \mathbf{M} & =\frac{c}{\mu} \mathrm{ViE} \\
\mathfrak{U E} & =\frac{c}{K} \mathrm{VMi}
\end{aligned}
$$

Eliminate $\mathbf{E}$, remembering that $(\mathbf{M i})=0$; then the result will be

$$
\mathfrak{b}(\mathfrak{b}-v) \mathbf{M}=\frac{c^{2}}{K \mu} \mathbf{M}=\mathfrak{b}^{\prime 2} \mathbf{M}
$$

where $\mathfrak{b}^{\prime}$ would be the velocity of propagation, if the medium were stationary in $S$. Thus

$$
\mathfrak{b}(\mathfrak{v}-v)=\mathfrak{v}^{\prime 2}
$$

and, the sense of propagation being that of $\nabla$,

$$
\mathfrak{v}=\frac{1}{2} v+\sqrt{\mathfrak{b}^{\prime 2}+v^{2} / 4}
$$

which is the required formula.
Note 4 (to page 39). To spare me trouble and to give the reader a sample of Fresnel's charming manner of exposition, I quote here simply the closing passages of his letter to Arago (loc. cit. pp. 633-636), in which he treats in a masterly manner the zouter-telescope experiment, both on the corpuscular and on the undulatory theory of light:
'Je terminerai cette lettre par une application de la même théorie à l'expérience proposée par Boscovich, consistant à observer le phénomène de l'aberration avec des lunettes remplies d'eau, ou d'un autre fluide beaucoup plus réfringent que l'air, pour s'assurer si la direction dans laquelle on aperçoit une étoile peut varier en raison du changement que le liquide apporte dans la marche de la lumière. Je remarquerai d'abord qu'il est inutile de compliquer cle l'aberration le résultat que l'on cherche, et qu'on peut aussi bien le déterminer en visant un objet terrestre qu'une étoile. Voici, ce me semble, la manière la plus simple et la plus commode de faire l'expérience.'
'Ayant fixé à la lunette même, on plutôt au microscope $F B D E$ [figure 2 of Fresnel's letter], le point de mire $M$, situé dans le prolongement de son axe optique $C A$, on dirigerait ce système perpendiculairement à l'écliptique, et, après avoir fait l'observation dans un sens, on le retour-

## BOSCOVICH'S EXPERTMENT

nerait bout pour bout, et l'on ferait l'olservation entsenseontraire. Si le mouvement terrestre déplaçait l'image du point $M$ parstrapport au fil de loculaire, on la verrait de cette manière jantôt à droite ettof du fil.'
 marqué, que le mouvement terrestre ne doit rieffohger a apparepres du phénomène. En effet, il résulte de ce moubcment quie ke rayen partant de $M$ doit prendre, pour passer par le centre de l'objectiryne direction $M A^{\prime}$ telle que l'espace $A A^{\prime}$ soit parcouru par le globe dans e. même intervalle de temps que la lumière emploie à parcourrid $A^{\prime}$, ou $M A$ (à cause de la petitesse de la vitesse de la terre relativement偮celle

de la lumière). Représentant par $v$ la vitesse de la lumière dans l'air, et par $t$ celle de la terre [i.e. our $c$ and $v$ respectively], on a donc:

$$
M A: A A^{\prime}:: v: t \text { ou } \frac{A A^{\prime}}{M A}=\frac{t}{v}
$$

c'est le sinus d'incidence. $v^{\prime}$ étant la vitesse de la lumière dans le milieu plus dense que contient la lunette [ $v^{\prime}$ is our $\left.c / n\right]$, le sinus de l'angle de réfraction $C^{\prime} A^{\prime} G$ sera égal à $\frac{t}{v^{\prime}}$; on aura donc $C^{\prime} G=A^{\prime} C^{\prime} \frac{t}{v^{\prime}}$; d'où l'on tire la proportion

$$
C^{\prime} G: A^{\prime} C^{\prime}:: t: v^{\prime} .
$$

Par conséquent le fil $C^{\prime}$ de l'oculaire placé dans l'axe optique de la lunette arrivera en $G$ en même temps que le rayon lumineux qui a passé par le centre de l'objectif.'

So far the corpuscular or emission theory. Again :
${ }^{\text {' La théorie des ondulations concluit au même résultat. Je suppose, }}$ pour plus de simplicité, que le microscope est dans le vide. $d$ et $d^{\prime}$ étant les vitesses de la lumière dans le vide et dans le milieu que contient la
lunette, on trouve pour le sinus de l'angle d'incidence $A M A^{\prime}, \frac{t}{d}$, et pour celui de l'angle de réfraction $C^{\prime} A G, \frac{t d^{\prime}}{d^{2}}$. Ainsi, indépendamment du déplacement des ondes dans le sens du mouvement terrestre,

$$
C^{\prime} G=A^{\prime} C^{\prime} \frac{t d^{\prime}}{d^{2}}
$$

Mais la vitesse avec laquelle ces ondes sont entraînées par la partie mobile du milieu dans lequel elles se propagent est égale à

$$
t\left(\frac{d^{2}-d^{\prime 2}}{d^{2}}\right)
$$

[i.c. in our notation $v\left(\mathrm{I}-\frac{\mathrm{I}}{n^{2}}\right)$ ]; donc leur déplacement total $G g$, pendant le temps qu'elles emploient à traverser la lunette, est égal à

$$
\frac{A^{\prime} C^{\prime}}{d^{\prime}} t\left(\frac{d^{2}-d^{\prime 2}}{d^{2}}\right)
$$

ainsi

$$
C^{\prime} g=A^{\prime} C^{\prime} \cdot t\left(\frac{d^{\prime}}{d^{\prime 2}}+\frac{d^{2}-d^{\prime 2}}{d^{\prime} d^{2}}\right)=A^{\prime} C^{\prime} \cdot t\left(\frac{d^{2}}{d^{\prime} d^{2 i}}\right)=A^{\prime} C^{\prime} \cdot \frac{t}{d^{\prime \prime}}
$$

On a donc la proportion $C^{\prime} g: A^{\prime} C^{\prime}:: t: d^{\prime}$; par conséquent l'image du point $M$ arrivera en $g$ en même temps que le fil du micromètre. Ainsi les apparences du phénomène doivent toujours rester les mêmes quel que soit le sens clans lequel on tourne cet instrument. Quoique cette expérience n'ait point encore été faite, je ne doute pas qu'elle ne confirmât cette conséquence, que l'on déduit également du'système de l'émission et de celui des onclulations.'

Note 5 (to page 42). Stokes' theory of aberration ('On the Aberration of Light,' Phil. Mag., Vol. XXVII., 1845, p. 9, reprinted in Math. and Phys. Papers, Vol. I. p. 134) was based on the assumption that the aether surrounding the earth is dragged by this planet in its annual motion, in such a way that the velocity of the aether relative to the earth is nil near its surforce, and, increasing gradually, becomes equal and opposite to the earth's orbital velocity at very considerable distances from our planet. It is obvious that this hypothesis led at once to a rigorous independence of purely terrestrial optical phenomena from the earth's annual motion. But in order to explain correctly astronomical aberration, Stokes had to assume that the aether's motion, betwreen the earth and the 'fixed' stars, is purely irrotationcal, which assumption could not be reconciled with the absence of sliding over the earth's surface, so long as the aether was regarded as incompressible. It is true that this difficulty, as has been shown by Planck, can be overcome by giving up the incompressibility, namely by supposing the aether to be condensed around the earth and the celestial bodies, as if it were subjected to
gravitation and behaved more or less like a gas. But the condensation around the earth, required to reduce the sliding to, say, one half per cent. of the earth's orbital velocity, would be something like $e^{11}$, i.e. corresponding to a density of the aether near the earth about 60,000 times as great as its density in celestial space. Now, it is certainly difficult to admit that the velocity of light is not to any sensible extent altered by this enormous condensation of the aether around the earth.

Particulars concerning the discussion of this most interesting subject will be found in Lorentz's book on Theory of Electrons (Chap. V.), and in his original paper on 'Stokes' Theory of Aberration in the Supposition of a Variable Density of the Aether,' Amsterdam Proceedings, 1898-1899, p. 443, reprinted in Abhandlungen üb. theor. Physik, Vol. I. p. 454.

## CHAPTER III.

THEOREM OF CORRESPONDING STATES. SECOND ORDER DIFFICULTIES. THE CONTRACTION HYPOTHESIS. LORENTZ'S GENERALIZED THEORY.

Let us return to Lorentz's macroscopic equations, for a material medium moving relatively to the aether with uniform velocity $\mathbf{v}$,

$$
\left.\begin{array}{rl}
\frac{\partial \mathbb{G}}{\partial t} & =c \cdot \operatorname{curl} \mathbf{M}^{\prime} ; \\
\frac{\partial \mathbf{M}}{\partial t} & =-c \cdot \operatorname{curl} \mathbf{E}^{\prime} ;  \tag{L}\\
\operatorname{div}^{\prime}=0 \\
\mathbf{M}^{\prime} & =\mathbf{M}-\frac{\mathbf{I}}{c} \mathbf{V} \mathbf{V} \mathbf{E}^{\prime} \\
K \mathbf{E}^{\prime} & =\mathbb{E}+\frac{\mathbf{I}}{c} \mathbf{V} \mathbf{V} \mathbf{M} .
\end{array}\right\}
$$

In the simplest case of a medium fixed in the aether, i.e. for $v=0$, these, as already noticed, become identical with Maxwell's equations for a stationary dielectric,

$$
\left.\begin{array}{cc}
\frac{\partial \mathfrak{E}}{\partial t}=c \cdot \operatorname{curl} \mathbf{M} ; & \operatorname{div} \mathfrak{E}=0  \tag{0}\\
\frac{\partial \mathbf{M}}{\partial t}=-c \cdot \operatorname{curl} \mathbf{E} ; & \operatorname{div} \mathbf{M}=0 \\
\mathfrak{C}=K \mathbf{E} .
\end{array}\right\}
$$

In order to exhibit the properties of the more general equations (L), Lorentz introduces instead of the 'universal time,' as he calls $t$, a new variable $t^{\prime}$, which will now be explained.

Let $O^{\prime}$ be a point fixed in the material body, chosen arbitrarily but once and for ever as the origin of coorclinates, $x^{\prime}, y^{\prime}, z^{\prime}$, measured
along axes rigidly attached to the body. From $O^{\prime}$ draw to any individual point of the body $P^{\prime}\left(x^{\prime}, y^{\prime}, z^{\prime}\right)$ the vector $\mathrm{r}^{\prime}$, so that the three Cartesian coordinates are condensed in

$$
\mathbf{r}^{\prime}=\mathbf{i} x^{\prime}+\mathbf{j} y^{\prime}+\mathbf{k} z^{\prime}
$$

Let us call the framework of reference rigidly attached to the body the system $S^{\prime}$. For comparison and to impress better upon your mind the meaning of $r^{\prime}$, take also an initial point $O$ fixed in the aether, i.e. relatively to the system $S$, and draw from $O$ to $P$ the vector $\mathbf{r}$, or in semi-Cartesian expansion, using the same unit vectors as above,*

$$
\mathbf{r}=\mathbf{i} x+\mathbf{j} y+\mathbf{k} z .
$$

If $O^{\prime}$ is taken to coincide with $O$ at the instant $t=0$, we have simply


Fig. 6.

Remember that the equations (L) hold for $t$ and $x^{\prime}, y^{\prime} z^{\prime}\left(\right.$ not $x, y^{\prime}, z$ ) as independent variables, or, more shortly, for

$$
\mathbf{r}^{\prime}, \quad t
$$

This fixes the meaning of curl, div and $\partial / \partial t$, as already mentioned in Chap. II. As regards the curls and divergences, they are, of course, the same in $x^{\prime}, y^{\prime}, z^{\prime}$ as in $x, y^{\prime}, z$.
*This is always possille, since the material body or medium moves relatively to $S$ in a purely translational manner.

Now, $\mathbf{r}^{\prime}$ being the above vector characterising any given point $P^{\prime}$ of the moving body or medium, the new variable $t^{\prime}$ is defined by

$$
\begin{equation*}
t^{\prime}=t-\frac{\mathrm{r}}{c^{2}}\left(\mathrm{r}^{\prime} \mathbf{v}\right) \tag{I}
\end{equation*}
$$

and is called the local time at $P^{\prime}$. Since the scalar product in the second term vanishes for $\mathbf{r}^{\prime} \perp \mathbf{v}$, the local time coincides with the 'universal' one at all points lying on the plane passing through $O^{\prime}$ and perpendicular to the direction of motion. But at all other places the new and the old time differ from one another, the local time being behind the 'universal' time in the anterior portion of the body, and the reverse being the case in its posterior portion (Fig. 6). In Cartesians, if $\mathbf{v}=\mathbf{i} v_{1}+\mathbf{j} v_{2}+\mathbf{k} v_{3}$, the local time is

$$
t^{\prime}=t-\left(x^{\prime} v_{1}+y^{\prime} v_{3}+z^{\prime} v_{3}\right) / c^{2}
$$

or if $\mathbf{i}$ be taken along the direction of motion, $t^{\prime}=t-x^{\prime} v / c^{2}$.
Notice that Lorenty's local time, as just defined, has nothing physical about it. It is merely an auxiliary mathematical quantity to be used instead of the 'universal' time $t$ in order to simplify the form of equations (L). It is constructed expressly for this purpose, and serves it excellently.

In fact, taking instead of $\mathbf{r}^{\prime}, t$ (or $x^{\prime}, y^{\prime}, z^{\prime}, t$ )

$$
\vec{r}^{\prime}, t^{\prime}
$$

as the new independent variables, and denoting the divergence and curl in terms of the new variables by

$$
\operatorname{div}^{\prime} \text { and curl', }
$$

we obtain, for example, by ( r ) and by the third of equations (L),

$$
\begin{aligned}
\operatorname{div} \mathbf{M} & =\operatorname{div}^{\prime} \mathbf{M}+\frac{\mathbf{I}}{c} \nabla \operatorname{curl} \mathbf{E}^{\prime} \\
& =\operatorname{div}^{\prime} \mathbf{M}-\frac{\mathbf{I}}{c} \operatorname{div} V \nabla \mathbf{E}^{\prime}
\end{aligned}
$$

since curl $\nabla=0$, by hypothesis. But for $V \nabla \mathbf{E}^{\prime}$, as for any vector normal to $\mathbf{\nabla}$, we have, obviously, div $=$ div'. Hence, by the fifth $^{\prime}$ of (L),

$$
\operatorname{div} \mathbf{M}=\operatorname{div}^{\prime}\left(\mathbf{M}-\frac{\mathbf{I}}{c} V \nabla \mathbf{E}^{\prime}\right)=\operatorname{div}^{\prime} \mathbf{M}^{\prime}
$$

Thus, the fourth of equations ( $L$ ), div $\mathbf{M}=0$, becomes, in the new variables, $\operatorname{div}^{\prime} \mathbf{M}^{\prime}=0$. Similarly, the second of ( $L$ ), $\operatorname{div} \mathbb{E}=0$, is transformed into $\operatorname{div}^{\prime} \mathbb{E}^{\prime}=0$, where $\mathbb{E}^{\prime}$ is a new vector defined by the formula

$$
\begin{equation*}
\mathbb{E}^{\prime}=\mathbb{E}+\frac{I}{C} V^{\prime} \mathbf{v M} . \tag{2}
\end{equation*}
$$

Using this new rector and the vector $\mathbf{M}^{\prime}$, defined by the fifth equation, the remaining equations (L) may be transformed, with equal ease, to the new variables.

The result is surprisingly simple. The system of Lorentz's equations ( $L$ ) for a moving medium takes with the new variables $\mathbf{r}^{\prime}, t^{\prime}\left(x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}\right)$ the form

$$
\left.\begin{array}{cc}
\frac{\partial \mathbb{E}^{\prime}}{\partial t^{\prime}}=c \cdot \operatorname{curl}^{\prime} \mathbf{M}^{\prime} ; & \operatorname{div}^{\prime} \mathbb{E}^{\prime}=0 \\
\frac{\partial \mathbf{M}^{\prime}}{\partial t^{\prime}}=-c \cdot \operatorname{curl}^{\prime} \mathbb{E}^{\prime} ; & \operatorname{div}^{\prime} \mathbf{M}^{\prime}=0 \\
\mathfrak{E}^{\prime}=K^{\prime} \mathbf{E}^{\prime},
\end{array}\right\}
$$

that is to say, precisely the same form as for a stationary medium, $\left(L_{0}\right)$, the only difference being that the electromagnetic vectors $\mathbf{E}$, $\mathfrak{E}, \mathbf{M}$ are replaced by their dashled correspondents, as are also the independent variables r , $t$.
This remarkable discovery, made by Lorentz, has played a most important rôle not only in his own theory, but also in the subsequent evolution of ideas concerning electromagnetism and optics. Undoubtedly, it may, to a great extent, be regarded as the germ of modern relativistic tendencies. It will therefore be worth our while to treat this subject at some length, and not only as an historical episode.
The above result may be put into the form of what has been called by Lorentz the Theorem of corresponding states:

If zve have for a stationary medium or system of bodies any solution (of Maxzeell's equations $\mathrm{L}_{0}$ ), in zekhich

$$
\mathbf{E}, \mathbb{E}, \mathbf{M}
$$

are certain functions of

$$
x, y, z, t,
$$

zve zoill obtain a solution for the same system of bodies moving zeith uniform translation-velocity $\mathbf{v}$, taking for

$$
\mathbf{E}^{\prime}, \mathbb{E}^{\prime}, \mathbf{M}^{\prime}
$$

exactly the same functions of the variables

$$
x^{\prime}, y^{\prime}, z^{\prime} a m d t^{\prime}=t-\frac{I}{c^{2}}\left(\mathbf{v r}^{\prime}\right) .
$$

In other words, and somewhat more shortly:
For each state in which $\mathbf{E}, \mathbb{E}, \mathbf{M}$ depend in a certain way on $x, y, z, t$ in the stationary system, there is a corresponding state in the moving system characterised by $\mathbf{E}^{\prime}, \mathbb{E}^{\prime}, \mathbf{M}^{\prime}$ which depend in the same way on $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$.

It will be useful to put here together the scattered definitions of the dashed vectors. These are, by (32), Chap. II.,* by (2) and by the fifth of equations ( L ),

$$
\left.\begin{array}{rl}
\mathbf{E}^{\prime} & =\mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{V} \mathbf{M}  \tag{3}\\
\mathfrak{E}^{\prime} & =\mathbb{E}+\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{V} \mathbf{M} \\
\mathbf{M}^{\prime} & =\mathbf{M}-\frac{I}{c} \nabla \mathbf{V} \mathbf{E}^{\prime} .
\end{array}\right\}
$$

As to the coordinate systems, notice that they are in both cases rigidly attached to the material medium or to the system of bodies in question, $x, y, z$ being fixed together with it in the aether, and $x^{\prime}, y^{\prime}, z^{\prime}$ sharing its motion through the aether.
'The above theorem of corresponding states has, of course, like the equations ( L ) themselves, the character of a first approximation only, terms of the order of $\beta^{2}=v^{2} / c^{2}$ having been neglected.
The broad and easy applicability of this beautiful theorem of Lorentz is obvious. It will be cnough to cuote here a few illustrative examples.

* Remembering that $\mathbf{M}$ itself is of the first order, so that

$$
\frac{\mathrm{I}}{c} \overline{\mathrm{VpM}} \div \frac{\mathrm{I}}{c} \overline{\mathrm{~V} \mathrm{VM}}=\frac{\mathrm{I}}{c} \mathrm{VV}_{\mathrm{V}},
$$

i.e. in the adopted short notation, $\frac{I}{c}$ VrM.

If, in the stationary medium or system $S$ of bodies, $\mathbf{E}, \mathbb{E}, \mathbf{M}$ are periodical functions of $t$, with period $T$, then, in the moving system $S^{\prime}$, the vectors $\mathbf{E}^{\prime}, \mathbb{E}^{\prime}, \mathbf{M}^{\prime}$ are periodical functions of the local time $t^{\prime}$, and consequently, at a point $P^{\prime}$ fixed in $S^{\prime}$, also of $t$, zeith the same relative period $T$. What Lorentz calls the relative period is the period of changes going on at a fixed point of the system $S^{\prime}$ moving relatively to the aether, i.e. for a constant $\mathbf{r}^{\prime}$, whereas the period of changes taking place at a point fixed in the aether, i.e. for a constant $\mathbf{r}$, is called the absolute period. Similarly, relative rays are distinguished from absolute rays, and so on. Thus, to luminous vibrations in $S$ of a given absolute period correspond luminous vibrations in $S^{\prime}$ of the same relative period.

If, in certain regions of the stationary system, $E=0$, etc., then also $E^{\prime}=0$, etc., in the corresponding regions of the moving system. Thus, to clarkness corresponds darkness. Also, limitations of beams in $S$ and $S^{\prime}$ correspond to one another. Luminous ray's in $S^{\prime}$, of relative period $T$, are refracted and reffected according to the same lazes as rays of (alsolute) period $T$ in $S$. The same is true of the distribution of dark and bright interference fringes, and consequently also of the concentration of light in a focus, by mirrors or lenses, this being a limiting case of diffraction.
But, although the lateral limitations of beams for corresponding states are the same, corresponding zvave normats in $S, S^{\prime}$ have generally different directions, this being again an immediate consequence of the theorem of corresponding states. In fact, if we have in $S$, say, plane waves whose normal is given by the unit vector $n$ and whose velocity of propagation is $\mathfrak{V}$, i.e. if $\mathbf{E}, \mathbb{E}, \mathbf{M}$ are proportional to a function of the argument

$$
(\mathrm{rn})-\mathrm{v} t,
$$

then, in the moving system, $\mathbf{E}^{\prime}$, etc., will be the same functions of the argument

$$
\begin{equation*}
\left(\mathbf{r}^{\prime} \mathbf{n}\right)-\mathfrak{b} t^{\prime}=\left(\mathbf{r}^{\prime} \mathbf{n}\right)+\frac{\mathfrak{b}}{c^{2}}\left(\mathbf{r}^{\prime} \mathbf{v}\right)-\mathfrak{b} t \tag{4}
\end{equation*}
$$

Consequently, the direction of the wave normal in the moving system will be given by that of the vector

$$
\begin{equation*}
\mathbf{N}^{\prime}=N^{\prime} \mathbf{n}^{\prime}=\mathbf{n}+\frac{\mathfrak{W}}{c^{\mathbf{z}}} \mathbf{v} \tag{5}
\end{equation*}
$$

Thus, unless $\mathbf{n} \| \mathbf{v}$, the directions of the wave normals in $S$ and $S^{\prime}$ are different. To state the same thing in Cartesians, the directioncosines of the wave normal in the moving system will be given by the proportions

$$
n_{1}^{\prime}: n_{2}^{\prime}: n_{3}^{\prime}=\left(n_{1}+\frac{\mathfrak{b}}{c^{2}} v_{1}\right):\left(n_{2}+\frac{\mathfrak{b}}{c^{2}} v_{2}\right):\left(n_{3}+\frac{\mathfrak{b}}{c^{2}} v_{3}\right) .
$$

In particular, for a vacuum or, very approximately, for air, in which case $\mathfrak{b}=c$,

$$
\begin{equation*}
N^{\prime}=n+\frac{1}{c} \nabla \tag{5a}
\end{equation*}
$$

or, in clumsy Cartesians,

$$
n_{1}^{\prime}: n_{2}^{\prime}: n_{3}^{\prime}=\left(n_{1}+\begin{array}{c}
v_{1} \\
c
\end{array}\right):\left(n_{2}+\frac{v_{2}}{c}\right):\left(n_{3}+\begin{array}{c}
v_{3} \\
c
\end{array}\right) .
$$

These formulae may, after a slight transformation, be applied at once to the case of astronomical aberration, the relative period being here that reduced according to I)oppler's law. Thus Lorentz obtains immediately the right results for air- and water-telescope aberration. (Cf. Essay, p. 89.)

To obtain the dragging coefficient it is enough to write the argument (4)

$$
\left(\mathbf{r}^{\prime} \mathbf{N}^{\prime}\right)-\mathfrak{v} t=N^{\prime}\left\{\left(\mathbf{r}^{\prime} \mathbf{n}^{\prime}\right)-\frac{\mathfrak{v}}{N^{i} t}\right\}
$$

Since here $\mathbf{n}^{\prime}$ is a unit vector, the velocity of propagation in $S^{\prime}$ is

$$
\mathfrak{v}^{\prime}=\frac{\mathfrak{b}}{\bar{V}^{\prime}}=\mathfrak{v}\left\{\mathrm{I}+\frac{\mathfrak{v}^{2}}{c^{2}} \beta^{2}+\frac{2 \mathfrak{v}}{c^{2}}(\mathrm{mn})\right\}^{-\frac{1}{2}},
$$

or, neglecting the term containing $\beta^{2}=(v / c)^{2}$, developing the square root and neglecting again the second and higher powers of ( $\mathbf{v n}$ ) $/ c_{\text {, }}$

$$
\begin{equation*}
\mathfrak{v}^{\prime}=\mathfrak{v}-\left(\frac{b}{c}\right)^{2}(\mathrm{~m}) . \tag{6}
\end{equation*}
$$

In particular, if the propagation is in the direction of motion or against it, as in Fizeau's experiment,

$$
v^{\prime}=\mathfrak{b}\left(\frac{v}{c}\right)^{2} v .
$$

Thus, the velocity of propagation relative to the aether will be

$$
\mathfrak{V} \pm\left\{I-\left(\frac{\mathfrak{b}}{\boldsymbol{c}}\right)^{2}\right\} v
$$

and the value of the dragging coefficient

$$
\kappa=\mathrm{I}-\left(\frac{\mathfrak{b}}{c^{2}}\right)=\mathrm{r}-\frac{\mathrm{r}}{v^{2}} .
$$

Here $v=c / \mathrm{b}$ is the refractive index of the medium, say water, corresponding to the relative period which is connected with the period $T$ of the emitted light by the formula

$$
T_{\mathrm{rel}}=\left(\mathrm{I} \pm \frac{v}{\mathfrak{b}^{\prime}}\right) T \doteqdot\left(\mathrm{I} \pm \frac{v}{\mathrm{~b}}\right) T
$$

second order terms being neglected. Thus, if $n$ be the refractive index for the period $T$,

$$
v=n \pm \frac{v}{b} T \frac{\partial n}{\partial T},
$$

whence Lorentz's formula for the dragging coefficient,

$$
\kappa=\mathrm{I}-\frac{\mathrm{r}}{n^{2}}-\frac{\mathrm{x}}{n} T \frac{\partial n}{\partial T}
$$

closely agreeing with experiment, as already mentioned in Chapter II.
For purely terrestrial experiments, in which not only the observer but also every part of his apparatus and the source of light are attached to the earth, the theorem of corresponding states leads to the following result:

The earth's motion has no first order influence whatever on any of such experiments.

The possibility of a second order influence remains, of course, in this stage of the research, an open question. For, as will be remembered, before arriving at the macroscopic equations (L), from which the theorem of corresponding states has been seen to follow, $\beta^{2}$-terms have been throughout neglected. In other words, that beautiful theorem, developed and illustrated by a series of most important examples in the fifth section of Lorentz's classical Essay, is but a first order approximation.

So far everything is quite satisfactory. But now, in the sixth, and last, section of Lorentz's Essay the difficulties begin.* In this section Lorentz investigates three problems, of which two concern the rotation of the plane of polarization and Fizeau's polarization experiments. But without dwelling on these, we shall pass straight on to the third one, namely to the famous interference experiment of Michelson and Morley. This second order or $\beta^{2}$-experiment, originally suggested by Maxwell, $\uparrow$ was performed by Michelson in 1881, and six years later repeated on a larger scale and with a higher degree of exactness by Michelson and Morley. $\ddagger$ A beam of luminous rays coming from the source $s$, after having been made parallel in the usual way, is divided by the semi-transparent

plane mirror (half-silvered plate) $a b$, which is inclined at an angle of $45^{\circ}$ to $s O A$, into a transmitted beam $O A$, and a reflected one $O B$. After having been reflected by the mirrors placed at $A$ and $B$ (at right angles to $O A, O B$, which directions are perpendicular to each other), the two beams of light return to the central mirror ; here a part of the first beam is reflected along $O C$ and a part of the second

* As is explicitly stated in the title: 'Alsschnitt VI.-Versuche, deren Frgel)nisse sich nicht ohne Weiteres erklären lassen.'
$\dagger$ See Note at the end of chapter.
$\ddagger$ A. A. Michelson, 'The relative motion of the earth and the luminiferous ether,' Amer. Journ. of Science, 3rd Ser. Vol. XXII., 188r. A. A. Michelson and E. W. Morley, Sill. Journ., and Ser. Vol. XXXI., 1886 ; Amer. Journ. of Science, 3rd Ser. Vol. XXXIV., 1887 ; Phil. Mas., 5 th Ser. Vol. XXIV., 1887. What is given above is but the usual rough scheme; details of the actual arrangement will be found in the original papers quoted and, to a certain extent, also in Michelson's popular book on Light Waves and their Uses, where a diagram of the actual apparatus is given (Fig. ro8):
beam is transmitted towards $C$, thus producing with one another a system of bright and dark interference fringes, which can be observed through a telescope placed on the line $O C$. To resume it shortly, the paths, taken relatively to the earth, of the two interfering beams of light are :

> sOAAOC and sOBBOC.

Let $O A$ (Fig. 7) be in the direction of the motion of the earth, and consequently also of the apparatus, source and all, with respect to the aether of Fresnel and Lorentz, and let $v$ be the velocity of this motion, i.e. the resultant of the earth's orbital velocit;, at the time being, and of the velocity of the solar system with respect to the 'fixed stars' or to those 'fixed' stars relatively to which the aether is supposed to be at rest. (Cf. Note 2.) On this assumption let us calculate the times taken by the two beams in travelling along their paths. Since the parts $s O$ and $O C$ are common to both, we have only to consider the intervals of time, say $T_{1}$ and $T_{2}$, taken to traverse

$$
O A A O \text { and } O B B O
$$

respectively, where the letters denote, of course, points attached to the apparatus.

Now, as has been already said in Chapter II., in connexion with Maxwell's equations for the 'free aether,' the velocity of light with respect to the aether is always equal $c=3 \cdot 10^{10} \mathrm{~cm} . \mathrm{sec}^{-1}$, quite independently of the motion of its source. This is no novel idea at all ; Fresnel himself considers it apparently as an obvious matter, when he says (in an early part of his letter, already mentioned) without any further explanations: 'car la vitesse avec laquelle se propagent les ondes est indépendante du mouvement du corps dont elles émanent.' Thus, according to both the classical and the more recent adherents of the aether, the velocity of light relative to the aether does not depend on the source's inotion; and on the wave-theory there is no reason why it should. Newton's corpuscular theory, revived in a more elaborate form in the writings of the late I)r. Ritz, need not detain us here.

Thus, the mirror $A$, receding from the waves on the part $O A$ of their journey, and the mirror $O$ moving toward them on their return from $A$ to $O$, we have

$$
T_{1}=\bar{O} \bar{A}_{l}\left\{\frac{1}{c-v}+\frac{1}{c+v}\right\}=\frac{2 c}{c^{2}-v^{2}} O A_{l},
$$

where the index ${ }_{l}$ is to remind us that $O A$ is 'longitudinal,' i.e. along the direction of motion. Putting $v / c=\beta$ and

$$
\begin{equation*}
\gamma=\frac{\mathrm{I}}{\sqrt{\mathrm{I}}-\overline{-\beta^{2}}} \tag{7}
\end{equation*}
$$

we may write shortly, without yet making any use of the smallness of $\beta^{2}$,

$$
\begin{equation*}
T_{1}=\frac{2}{c} \gamma^{2} \overline{A_{i}} \tag{8}
\end{equation*}
$$

To obtain $T_{2}$, the time for the second beam, we could say simply, after the manner of some authors, that the relative velocity of light, being the vector sum of the velocity $c$ parallel to $O B$ and of the velocity $v$ of the aether with respect to the apparatus, perpendicular to $O B$ and directed backwards, is equal $\left(c^{2}-v^{2}\right)^{\frac{1}{2}}$, so that

$$
T_{2}=2 \overline{O B}_{t}\left(c^{2}-v^{2}\right)^{-\frac{1}{2}}
$$

or

$$
\begin{equation*}
T_{2}=\frac{2}{c} \gamma \overline{O B_{l}} \tag{9}
\end{equation*}
$$


where the index ${ }_{t}$ is to remind us that $O B$ is 'transversal' or perpendicular to the direction of motion. But since this may not seem very satisfactory, we can support it by the following, equally frequent, reasoning which is but formally different from the above short statement. Contemplate for a moment Fig. 8, the paper on which it is drawn being now supposed to be stationary in the aether, and the apparatus moving past it from left to right. Let. the centre of the inclined mirror be at $O$ at the instant $t=0$, when the light leaves it, and at $O^{\prime \prime}$ at the instant $t=T_{2}$, when the light returns to it; let $B^{\prime}$ be the position of $B$ when the beam reaches it, and let $O^{\prime}$ be the simultaneous position of $O$. If it be granted
that the three distinct points of the aether, $O, O^{\prime}, O^{\prime \prime}$, are the consecutive positions of exactly the same point of the inclined mirror, that is to say, that the ray in question returns to exactly, or sensibly, the same point of the mirror from which it started, then $O B^{\prime} O^{\prime \prime}$ will be an isosceles* triangle, so that $O B^{\prime}=\frac{1}{2} c T_{2}$, and

$$
\frac{1}{4} c^{2} T_{2}^{2}=\frac{1}{4} v^{2} T_{2}^{2}+\overline{O B}_{t}^{2}
$$

This gives $T_{2}=2 \overline{O B}_{t}\left(c^{2}-v^{2}\right)^{-\frac{1}{2}}$, which is identical with (9).
By (8) and (9) we get for the time-difference of the two beams, by which the phenomenon of their interference is determined,

$$
\begin{equation*}
T_{1}-T_{2}=\frac{2}{c} \gamma\left\{\gamma \overline{O A}_{l}-{\overline{O B_{i}}}_{t}\right\} \tag{IO}
\end{equation*}
$$

Let us now turn round the whole apparatus through $90^{\circ}$, so that $O A$ becomes transversal, and $O B$ longitudinal. Then we shall have, using dashes to distinguish this case from the above one,

$$
T_{1}^{\prime}=\frac{2}{c} \gamma \overline{O A}_{t}, \quad T_{2}^{\prime}=\frac{2}{c} \gamma^{2} \overline{O B}_{l}
$$

so that the time-difference of the two beams will become

$$
T_{1}^{\prime}-T_{2}^{\prime}=\frac{2}{c} \gamma\left\{\overline{O A}_{t}-\gamma \overline{O B}_{l}\right\}
$$

If therefore the fixed-aether theory is true, such a rotation of the apparatus should produce a shift in the position of the interference fringes, corresponding to the change of the time-difference of the two beams, $\Delta=(10)-\left(10^{\prime}\right)$, i.e.

$$
\begin{equation*}
\Delta=\frac{2}{c} \gamma\left\{\gamma\left(\overline{O A}_{l}+\overline{O B}_{l}\right)-\left({\overline{O A_{l}}}_{l}+\overline{O B}_{t}\right)\right\} \tag{II}
\end{equation*}
$$

The indices ${ }_{l}$ and ${ }_{t}$, distinguishing between longitudinal and transversal orientation, have been introduced here (contrary to the historical order) only for the sake of subsequent discussions. To Michelson and Morley there was no question of distinguishing between the lengths of a segment in different orientations. To put

[^20]ourselves into agreement with their manner of treatment we have, therefore, to write simply
\[

$$
\begin{aligned}
& \overrightarrow{O A}_{l}=\overrightarrow{O A}_{t}=\overrightarrow{O A}, \\
& \overrightarrow{O B}_{l}=\overline{O B}_{t}=\overrightarrow{O B} .
\end{aligned}
$$
\]

' Co secure these equalities Michelson and Morley mounted the mirrors* and, in fact, the whole of the apparatus, on a heavy slab of stone mounted on a disc of wood which floated in a tank of mercury, so as to be able 'to rotate the apparatus without introducing strains.' In a word, they made the configuration of $O$, $A$, etc., 'rigid,' that is to say as rigid as a stone is. On this understanding, formula (ir) may be written

$$
\begin{equation*}
\Delta=\frac{2}{c} \gamma(\gamma-1) \cdot(\overline{O A}+\overline{O B}) . \tag{r2}
\end{equation*}
$$

As to the mutual relation of $\overline{O A}, \overline{O B}$, they were made 'nearly equal,' to suit the well-known requirements for producing neat interference fringes, in each of the two orientations of the apparatus. Moreover, since these lengths or distances enter in the formula only by their sum, their equality or non-equality is of no essential importance. We may therefore, without any more ado, write $\overline{O A}=\bar{O} \bar{B}=L$ or else call the sum of these lengths $2 L$. Then, as regards the factor depending on the velocity of motion, we have, by (7),

$$
\gamma(\gamma-1)=\left(\mathrm{I}-\beta^{2}\right)^{-1}-\left(\mathrm{I}-\beta^{2}\right)^{-\frac{1}{2}}
$$

or, up to quantities of the second order, i.e. neglecting $\beta^{4}$-terms, etc.,

$$
\mathrm{I}+\beta^{2}-\left(\mathrm{I}+{ }_{2}^{1} \beta^{2}\right)=\frac{1}{2} \beta^{2}
$$

Thus, the second-order effect to be expected on the stationaryaether theory would be determined by the change of the timedifference of the two beams

$$
\begin{equation*}
\Delta=\frac{2 \beta^{2}}{c} L \tag{r2a}
\end{equation*}
$$

If $T$ be the period of the light and $\lambda=c T$ the wave-length, the corresponding shift $s=\Delta / T$ of the interference bancls, measured as a fractional part of the distance of two neighbouring bands, would be given by

$$
\begin{equation*}
s=\beta^{\prime} \frac{2 L}{\lambda} . \tag{I3}
\end{equation*}
$$

* In the actual experiment not three but sixteen in number.

The length $2 L$, which in Michelson's original apparatus was too small, was in Michelson and Morley's experiment (1887) increased to about 22 metres, by multiple reflection from suitably plac̣ed mirrors. And since, for sodium light, $\lambda=5.89 .10^{-5} \mathrm{~cm}$., the tioio $2 L / \lambda$ had nearly the value $\cdot 37 \cdot 10^{s}$. As regards $\beta^{2}$, we should have, taking for $v$ simply the earth's orbital velocity, i.e. 30 kiloni, per second, $\beta^{2}=10^{-8}$. It is true that, at least in some of the experments, the rays of light, being horizontal, made a considerable ande with the earth's orbit, but on the other hand the motion of whe whe solar system exerted a favourable influence, so as to double the value of $\beta^{2}$ (as was already mentioned). So that to put $\beta^{2}$. equal $10^{-s}$ is certainly not to overestimate its value considerably. "Thus the shift should be on the stationary-aether theory, in rbiind figures,

$$
s=0.4 \text { of a fringe width. }
$$

In no case, however, did the actual displacement the fringes exceed .02, and probably it was less than or i.e. léss than $\frac{1}{40}$ th of the expected value. The experiment was repented. in 1905 by Morley and Miller* with considerably increased accuracy; and their result was that, if there is any fringe-shift of the kind expected, it is something like $s=.0076$ instead of I . 5 , i.e. not greater than one two-hundredth of the computed value. $\dagger$
'Thus, not nearly the expected second-order effect of the earth's motion relatively to the aether was observed. It seems, therefore, reasonable to say at least that, as far as we know, the above $\Delta$ is $n i t$.

In order to explain this negative result and to save, at the same time, the stationary-aether theory, Lorentz has had recourse to a peculiar hypothesis, constructed ad hoc, which occurred to him independently of Fitzgerald, who was the first to suggest $\mathrm{it} . \ddagger$ It is

[^21]now widely known under the name of the contraction hypothesis, and it consists in assuming that, in Lorentz's words, 'the dimensions of a solid body undergo slight changes, of the order $\beta^{3}$, when it moves through the ether,' namely a longitudinal contraction amounting to $\frac{1}{2} \beta^{2}$ per unit length or, more generalls; both a transversal and a longitudinal lengthening, $\epsilon$ and $\delta$, per unit length, such that $\epsilon-\delta={ }^{2} \beta^{2}$. This would amount for the whole earth to about 6.5 centimetres only:

To see at once that the negative result of the Michelson experimont is thus accounted for and to grasp as clearly as possible the nature of the hypothesis, let us return to the more general formula (ri) for $\Delta$, from which ( r 2 ) or ( $\mathrm{r} 2 a$ ) followed by identifying $\bar{O} A_{l}$ with $O \vec{A}_{t}$, and similarly $\overline{O \dot{B}_{l}}$ with $\overline{O B}_{t}$. Now, to simplify matters, assume $\bar{O} \bar{B}_{l}=O \bar{A}_{l}$ and $\bar{O} B_{t}=O \bar{A}_{t}$ (which, as we saw, is of no essential importance), but on the other hand distinguis/l betzueen $\bar{O} A_{l}$ and $\overline{O A} I_{l}$. Then formula ( r ), valid by the fixed-aether theory, will become

$$
\begin{equation*}
\Delta=\frac{4}{c} \gamma\left(\gamma O A_{1}-\widetilde{O A}_{t}\right) \tag{I4}
\end{equation*}
$$

and since $\Delta=0$, by experience, we have to write, in order to respect both that theory and experience,

$$
O A_{l}=\frac{1}{\gamma} \overline{O A_{t}}=\overline{O A}_{t}\left(\mathrm{r}-\beta^{2}\right)^{\frac{1}{2}}
$$

or, up to quantities of the second order,

$$
O \dot{A}_{l}: O A_{t}=1-\frac{1}{2}
$$

which is the Fitzserald-Lorentz hypothesis.
Notice that it would be a perfectly idle thing to cquarrel whether $\overline{O A}_{t}$ is shortened, while $O \bar{A}_{t}$ remains unchanged, by the earth's motion through the aether, or whether $O A_{t}$ alone is lengthened, or, finally, whether both are changed in suitable proportions. The only thing we are required by the aether theory and by experiment to do is to consider the ratio of the lengths of one and the same 'material'

[^22]segment $O A$, or shortly $L$, in those two orientations as being equal to $\mathbf{r}-\frac{1}{2} \beta^{2}$, or, more rigorously,
\[

$$
\begin{equation*}
L_{l}: L_{t}=\sqrt{\mathrm{r}-\beta^{2}} \tag{15}
\end{equation*}
$$

\]

This implies that for $\beta=0$, i.e. if the earth stopped moving through the aether, or nearly so, we should havc $L_{l}=L_{t}$, say, both equal to $L_{0}$. But it cannot inform us as to the ratio which either length bears to $L_{0}$, when the earth is moving through that medium ; moreover, such considerations are, thus far, physically meaningless.

At any rate, Lorentz soon decided in favour of a purely longitudinal contraction, which amounts to writing

$$
\begin{equation*}
L_{t}=L_{0} \quad \text { and } \quad L_{l}=\frac{L_{0}}{\gamma}=L_{0} \sqrt{1-\beta^{3}} . \tag{15a}
\end{equation*}
$$

In doing so he based himself on certain results obtained from the fundamental (microscopic) equations in an early part of his classical Essay, to be mentioned presently. That this, in fact, was his choice we see explicitly from the shape attributed by him to moving electrons. While Abraham's electron is and remains always a sphere, being rigid in the classical sense of the word, Lorentz's electron is a sphere of radius $R$, say, when at rest, and becomes flattened longitudinally, when in uniform motion, to a rotational ellipsoid of semiaxes

$$
\frac{I}{\gamma} R, R, R .
$$

Such an electron, of homogeneous surface- or volume-charge, is now generally known as the Lorentz electron. The history of its rivalry with the rigid one, and of its rather victorious issue from the contest, need not detain us here. It is, besides, sufficiently well known.

Lorentz's attitude towards the contraction hypothesis may be seen best from his own words, written in 1909 (Electron Theory', p. 196):
'The hypothesis certainly looks rather startling at first sight, but we can scarcely escape from it, so long as we persist in regarding the ether as immovable. We may, I think, even go so far as to say that, on this assumption, Michelson's experiment proves the chanses of dimension in question, and that the conclusion is no less legitimate than the inferences concerning the dilatation by heat or the changes of the refractive index that have been drawn in many other cases from the observed positions of interference bands.'

The obvious criticism of the above comparison may be left to the reader.

As regards the justification of the contraction hypothesis which to an unprepared mind certainly does 'look rather startling,' Lorentz observes in his original Essay' of 1895 (p. 124) that we are led precisely to the change of dimensions defined by ( $\mathrm{I}_{5}$ a) , if, disregarding the molecular motion, we assume that the attractive and repulsive forces acting on any molecule of a solid body which 'is left to itself' are in mutual cquilibrium, and if we apply to these molecular forces the same law which, by the fundamental equations, holds for electrostatic actions. It is truc, as Lorentz himself confesses, that 'there is, of course, no reason' for making the second of these assumptions. But those who entertain the hope of constructing an electromagnetic theory of matter will easily adhere to it. To obtain the law in question return to the fundamental electronic equations (r.), Chap. II., and introduce the so-called vector potential A and the scalar potential $\phi$, satisfying the differential equations

$$
\left.\begin{array}{l}
\left(\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}-\nabla^{2}\right) \phi=\rho  \tag{r6}\\
\left(\frac{1}{c} \frac{\partial t^{2}}{\partial^{2}}-\nabla^{2}\right) \mathbf{A}=\frac{1}{c} \rho \mathbf{p}
\end{array}\right\}
$$

and subject to the condition

$$
\begin{equation*}
\operatorname{div} \mathbf{A}+\frac{\mathrm{I}}{c} \frac{\partial \phi}{\partial t}=0 \tag{7}
\end{equation*}
$$

Then all of the equations (r.) will be satisfied by

$$
\left.\begin{array}{l}
\mathbf{E}=-\frac{\mathbf{I}}{c} \frac{\partial \mathbf{A}}{\partial t}-\nabla \phi  \tag{x8}\\
\mathbf{M}=\operatorname{curl} \mathbf{A},
\end{array}\right\}
$$

so that every slectromagnetic problcm is reduced to finding the potentials according to (16) and (17). Suppose, now, that a material body moves as a whole, relatively to the aether or to the system $S$, with uniform translational velocity $\mathbf{v}$, and that all the electrons it carries are at rest with respect to the body. Then the above $\mathbf{p}$ will have throughout the constant value $\mathbf{v}$, so that, by ( $\mathbf{r} 6$ ),

$$
\begin{equation*}
\mathbf{A}=\frac{\mathrm{r}}{c} \boldsymbol{\nabla} \phi . \tag{r9}
\end{equation*}
$$

Thus everything is made to depend on $\psi$ alene. 'Take the $x$-axis in $S$ along the direction of motion, so that $\mathbf{V}=z \mathbf{i}, \mathbf{A}=\mathrm{i} \beta\langle$, ancl suppose that the electromagnetic: field is invariable with respect to the material body. 'This assumption will be satisfied if $\psi$ is suppesed to depend only on the coordinates attached to the booly,

$$
\xi=x-\pi t, \quad \eta \cdot \quad \zeta,
$$

Thus we shall have
and the equation for $\psi$ will become
while the condition (17) will be satisfied idemtionly. Here

$$
\gamma^{\prime \prime \prime}=\left(1-\beta^{2}\right),
$$

as above. Again, by (18),

$$
\begin{aligned}
& \text { M. }{ }^{1}{ }^{1} \mathbf{V} \mathbf{V E}-\beta \backslash \mathbf{i E},
\end{aligned}
$$

whence the penderomotive force per mit charge, or loorente's electric forec, $\mathbf{E}+\beta \vee \mathrm{iM},(10)$, (haty. II., which we shall now denote: by fif (since the dashed E would be misleading),

$$
\begin{equation*}
\mathrm{A}=-\nabla\binom{\phi}{\gamma^{*}} \tag{21}
\end{equation*}
$$

 (here acting as the slopec), taken with respect to the aether or, which in our case is the same thing, with respect to the material body. Thus, the electric force is derived from a sealar potential $\phi / \gamma^{2}$, precisely as in ordinary bectrostatics, hy the way, $\phi / \gamma^{2}$ is called the convection potentiat. Notice that it is 捒, the electric, force, and not the 'diekentric displacenemt' $\mathbf{E}$, that has a sealar potential.
S.R, F

Now, supposing always $\beta^{2}<1$ and consequently $\gamma$ real, write

$$
\begin{equation*}
x^{\prime}=\gamma \xi, \quad y^{\prime}=\eta, \quad z^{\prime}=\zeta, \tag{22}
\end{equation*}
$$

and denote the corresponding Hamiltonian, i $\partial / \partial x^{\prime}+$ etc., by $\nabla^{\prime}$. Then (20) will become

$$
\begin{equation*}
\nabla^{\prime \prime 2} \phi=-\rho \tag{23}
\end{equation*}
$$

To adopt for the moment Lorentz's notation, call the moving material body or system of bodies the system $S_{1}$, and compare it with a system $S_{2}$ which is fixed in the aether and which is obtained from $S_{1}$ by stretching all its constituent bodies, together with the electrons, longitudinally in the ratio $\gamma: \mathrm{x}$, so that to any point $\xi, \eta, \zeta$ of $S_{1}$ corresponds the point $x^{\prime}, y^{\prime}, z^{\prime}$ of $S_{2}$, and so that corresponding volume-elements, $d \tau$ and $d \tau^{\prime}=\gamma d \tau$, contain equal charges. 'Then, $\rho$ and $\rho$ ' being the densities of electric charge at corresponding points,

$$
\rho^{\prime}=\frac{\mathrm{I}}{\gamma} \rho,
$$

and, by (23),

$$
\nabla^{\prime \prime} \phi=-\gamma \rho^{\prime}
$$

If then $\phi^{\prime}$ be the scalar, electrostatic, potential in $S_{2}$, so that

$$
\nabla^{\prime 2} \phi^{\prime}=-\rho^{\prime}
$$

we shall have

$$
\phi^{\prime}=\frac{\mathrm{I}}{\gamma} \phi,
$$

and consequently, instead of (2I), using (22),

$$
\mathrm{df}^{\prime \prime}=-\frac{\mathrm{I}}{\gamma} \nabla \phi^{\prime}=-\mathbf{i} \frac{\partial \phi^{\prime}}{\partial x^{\prime}}-\frac{\mathbf{r}}{\gamma}\left(\mathbf{j} \frac{\partial \phi^{\prime}}{\partial y^{\prime}}+\mathbf{k} \frac{\partial \phi^{\prime}}{\partial z^{\prime}}\right) .
$$

But the electric force in the stationary system $S_{2}$ is

$$
d f^{\prime}=-\nabla^{\prime} \phi^{\prime}=-\mathbf{i} \frac{\partial \phi^{\prime}}{\partial x^{\prime}}-\mathbf{j} \frac{\partial \phi^{\prime}}{\partial y^{\prime}}-\mathbf{k} \frac{\partial \phi^{\prime}}{\partial s^{\prime}} .
$$

Therefore, using the indices ${ }_{l}$ and ${ }_{l}$ to denote the longitudinal and the transversal components of the electric forces,
and since charges of corresponding elements are equal, exactly the same relations will hold between the ponderomotive forces acting on each electron in the moving system $S_{1}$ and on the corresponding electron in the stationary system $S_{2}$.

This is the 'law' alluded to. Now, suppose that it is obeyed by the molecular forces keeping together the parts of a moving solid which, disregarding its interior molecular and electronic motions, is to be taken for the system $S_{1}$. Then, if the molecular forces balance each other in the corresponding stationary body $S_{2}$, they will do so in the moving body $S_{1}$. But, by (22), $S_{1}$ is the body $S_{2}$ contracted longitudinally with preservation of its transversal dimensions, exactly as in ( $\mathrm{r} 5 a$ ), and the motion would produce this flattening 'by itself.' Whence Lorentz's justification of the contraction hypothesis.

Thus, the longitudinal contraction, though at first manifestly invented ad hoc, to account for the negative result of the Michelson experiment, found a kind of legitimate support by being brought into connexion with the fundamental assumptions of the clectron theory. But the cure of the disease has not bcen radical. In fact, the idea naturally suggested itself, that the Lorentz-Fitzgerald contraction, like an ordinary strain, might give rise to double refraction, of the order $\beta^{2}$, in solids or liquids, a property which should be directionally connected with the earth's motion round the sun. But here again the result of experiments has been sensibly negative. Lord Rayleigh's* experiments (r902) with licquids (water and carbon disulphide) as well as those with solids, with glass plates piled together, have given no trace of an effect of the expected kind. At least, if there was any effect on turning round the apparatus, it was less than $\frac{1}{100}$ th of that sought for. Rayleigh's experiment was then repeated (r904) by Brace $\dagger$ with considerably increased accuracy, and the result has again been negative: the relative retardation of the rays due to the supposed double refraction should be of the order $10^{-8}$, whereas, if existent at all, it was certainly less than $5 \cdot 10^{-11}$, in the case of glass, and even less than $7 \cdot 10^{-18}$, in the case of water.

To account for these obstinately negative results, and with a view to settle the matter once and for ever, Lorentz undertook what he

* Lord Rayleigh, Phil. Maag., Vol. IV. p. 678, 1902.
†D. B. Brace, Phil. MIag., Vol. VII. p. 317, 1904; Boltsmanh-fiestschrifi. p. $576,1907$.
thought a radical discussion of the whole subject, that is to say, of the electromagnetic phenomena in a uniformly moving system, not as hitherto for small values of $v$, but for $\alpha n y$ velocity of translation smaller than that of light, i.e. for any $\beta<\mathrm{x}$. Lorentz's ideas, laid down in a paper published in 1904,* are fully developed in his Columbia University Lectures, already quoted (p. 196 et seq.). His aim was now to reduce, 'at least as far as possible,' the electromagnetic equations for a moving system to the form of those that hold for a system at rest-always, of course, relatively to the aetherwithout neglecting either $\beta^{2}$ - or, in fact, terms of any order whatever.

It will be remembered that even in his first approximation, i.e. when neglecting $\beta^{2}$-terms, Lorentz employed the 'local time' $t^{\prime}=t-(\boldsymbol{\sigma}) / c^{2}$, or, measuring $x$ along the line of motion,

$$
\begin{equation*}
t^{\prime}=t-\frac{v}{c^{2}} x . \dagger \tag{a}
\end{equation*}
$$

'Then the necessity of accounting for the negative result of Michelson's interference experiment brought him to the contraction hypothesis, according to which the longitudinal dimensions of the moving system are reduced in the ratio $\mathrm{x}: \gamma^{-1}$, where $\gamma=\left(\mathrm{I}-\beta^{2}\right)^{-\frac{1}{2}}$, while the transversal ones remain unchanged. This contraction corresponds to $t=$ const., and consequently may easily be shown to be equivalent to transforming $x, y, z$, the coordinates of a point with respect to axes fixed in the aether, or the 'absolute' coordinates, into

$$
\begin{equation*}
x^{\prime}=\gamma(x-v t), \quad y^{\prime}=y^{\prime}, \quad z^{\prime}=z . \tag{b}
\end{equation*}
$$

It is true that the transformation (a) was as yet purely formal, and that the contraction, or (b), was introduced by Lorentz first ad hoc, but afterwards to be justified. But at anyrate, having already (a) and (b), Lorentz has been naturally led to investigate in a general way the consequences of introducing, instead of $x, y, z, t$,

[^23]new independent variables, called by him the effective coordinates and the effective time,
\[

\left.$$
\begin{array}{l}
x^{\prime}=\lambda \gamma(x-v t), \quad y^{\prime}=\lambda y^{\prime}, \quad z^{\prime}=\lambda z  \tag{25}\\
t^{\prime}=\lambda \gamma\left(t-\frac{v}{c^{2}} x\right),
\end{array}
$$\right\}
\]

where $\gamma$ is as above and $\lambda$ is a numerical coefficient of which Lorentz, provisionally, assumes only that it is a function of $v$ alone, whose value equals $I$ for $v=0$ and differs from I by an amount of the order $\beta^{2}$ for small values of the ratio $\beta=v / c^{*}$. Introducing the new variables (25) into the fundamental electronic equations, (r.), Chap. II., and defining new vectors $\mathbf{E}^{\prime}, \mathbf{M}^{\prime}$,
$\left.\begin{array}{lll}E_{1}^{\prime}=\lambda^{-2} E_{1}, & E_{2}^{\prime}=\gamma \lambda^{-2}\left(E_{2}-\beta M_{3}\right), & E_{3}^{\prime}=\gamma \lambda^{-2}\left(E_{3}+\beta M_{2}\right), \\ M_{1}^{\prime}=\lambda^{-2} M_{1}, & M_{2}^{\prime}=\gamma \lambda^{-2}\left(M_{2}+\beta E_{3}\right), & M_{8}^{\prime}=\gamma \lambda^{-2}\left(M_{3}-\beta E_{2}\right),\end{array}\right\}$
and also, instead of the relative velocity $\mathbf{p}-\mathbf{\nabla}$ of an electric particle, the vector

$$
\mathbf{p}^{\prime}=\gamma\left\{\mathbf{i} \gamma\left(p_{1}-v_{1}\right)+\mathbf{j}\left(p_{2}-v_{2}\right)+\mathbf{k}\left(p_{3}-v_{3}\right)\right\},
$$

i.e. with the above choice of axes, simply

$$
\begin{equation*}
\mathbf{p}^{\prime}=\gamma\left\{\mathrm{i} \gamma\left(p_{1}-v\right)+\mathrm{j} p_{2}+\mathrm{k} p_{\mathrm{g}}\right\}, \tag{27}
\end{equation*}
$$

and, instead of the density $\rho$,

$$
\begin{equation*}
\rho^{\prime}=\gamma^{\lambda^{\mathbf{3}} \rho}, \tag{28}
\end{equation*}
$$

Lorentz obtained again the equations (r.) with dashes,

$$
\partial \mathbf{E}^{\prime} \partial \partial t^{\prime}+\rho^{\prime} \mathbf{p}^{\prime}=c . \operatorname{curl}^{\prime} \mathbf{M}^{\prime}, \quad \text { etc. }
$$

but with the difference that $\operatorname{div} \mathbf{E}=\rho$ was replaced by

$$
\begin{equation*}
\operatorname{div}^{\prime} \mathbf{E}^{\prime}=\left[\mathrm{I}-\frac{\mathrm{I}}{c^{2}}\left(\mathbf{v p ^ { \prime }}\right)\right] \rho^{\prime}, \tag{29}
\end{equation*}
$$

* Columbia University Lectures, p. 196. The above $v, \gamma, \lambda$ stand for Lorentz's $w, k, l$ respectively. A transformation equivalent to (25) was previonsly applied by Voigt, as early as 1887 , to equations of the form $\frac{1}{c^{4}} \frac{j^{3}}{\partial t^{2}}-\nabla^{4}=0$; ' Ueber das
Doppler'sche Princip, Göttinger Nachrichten, 1887, p. 41. Lorent\% himself states (loc. cit., p. 198; 1909) that Voigt's paper had escaped his notice all these years, and adds: 'The idea of the transformation' (25) 'might therefore have been borrowed from Voigt, and the proof that it does not alter the form of the equations for the free ether is contained in his paper.'
not by $\operatorname{div}^{\prime} \boldsymbol{E}^{\prime}=\rho^{\prime}$. Thus, the fundamental equations for the free aether ( $\rho=\rho^{\prime}=0$ ) turned out to be rigorously invariant with respect to the transformation (25), which, especially for $\lambda=I$, has since been universally called the Lorentz transformation. The same invariance holds also in the general case, that is to say, in the presence of electric charges, but for the slight deviation given by (29).

Using this result, Lorentz generalized his Theorem of corresponding states for any velocity $v$ smaller than $c$, and succeeded in showing that the theorem thus extended not only accounts for the contraction required by the result of the Michelson experiment, but that it explains, among other things, why Lord Rayleigh and Brace failed to detect a double refraction due to the earth's orbital motion. A discussion of the formulae for the longitudinal and transversal masses of an electron, which need not detain us here,* led Lorentz to attribute to the coefficient $\lambda$ (his $l$ ) the value r , whereby the transformation formulae (25) and (26) were reduced to

$$
\left.\begin{array}{l}
x^{\prime}=\gamma(x-v t), \quad y^{\prime}=y, \quad z^{\prime}=z,  \tag{30}\\
t^{\prime}=\gamma\left(t-\frac{v}{c^{2}} x\right),
\end{array}\right\}
$$

and

$$
\left.\begin{array}{ll}
E_{1}^{\prime}=E_{1}, & E_{2}^{\prime}=\gamma\left(E_{2}-\beta M_{3}\right),  \tag{31}\\
M_{1}^{\prime}=M_{1}, & E_{2}^{\prime}=\gamma\left(E_{3}^{\prime}=\gamma\left(M_{2}+\beta E_{3}\right),\right. \\
M_{3}^{\prime}=\gamma\left(M_{3}-\beta E_{2}\right) .
\end{array}\right\}
$$

With this specialization, Lorentz's modified theory, which in its essence was built up in 1904, satisfied the requirements of selfconsistency and accounted for the negative results of all, second as well as first order, terrestrial experiments intended to show our planet's motion through the aether. In other words, by modifying and gradually extending his original theory, Lorentz obtained the desired physical equivalence of the 'moving' system $S^{\prime}$, with its effective coordinates and time $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$, and of a corresponding 'stationary' system with its absolute coordinates and time $x, y, z, t$.

But still one of the two systems $S$, $S^{\prime}$, namely $S$, was privileged, being regarded by Lorentz as 'fixed in the aether.' Their equivalence, as indicated persistently by such numerous experiments, was not placed as the basis of the theory, but followed as the result of long, laborious, and rather artificial constructions, intended to com-

[^24]pensate gradually the pretended play of the 'aether.' For, to repeat, Lorentz continued to assume this hypothetical medium of his classical Essay in his extended theory, dated 1904, and acheres to it even now, if we may judge from the last sentences of his American Lectures (p. 230). Not only is the aether for Lorentz a unique framework of reference, but he 'cannot but regard it as endowed with a certain degree of substantiality.' According to this standpoint, then, there certainly is such a thing as the aether, though every physical effect of the motion of ordinary, ponderable matter through it, being compensated by more or less intricate processes, remains undiscoverable for ever.
As' regards the above transformation of Lorentz, we may further notice here that Poincaré made, in 1906, an extensive use of its more general form (25) [Rend. del Circolo mat. di llalerno, Vol. XXI. p. 129] for the treatment of the dynamics of the electron and also of universal gravitation. Some of Poincare's results continue even now to be of considerable interest.
In the meantime, 1905, Einstein published his paper on 'the electrodynamics of moving bodies,'* which has since become classical, in which, aiming at a perfect reciprocity or equivalence of the above pair of systems, $S, S^{\prime}$, and denying any claims for primacy to either, he has investigated the whole problem from the bottom. Asking himself questions of such a fundamental nature, as what is to be understood by 'simultaneous' events in a pair of distant places, and dismissing altogether the idea of an acther, and in fact of any unique framework of reference, he has succeeded in giving a plausible support to, and at the same time a striking interpretation of, Lorentz's transformation formulae and the results of Lorentz's extended theory. Einstein's fundamental ideas on physical time and space, opening the way to modern Relativity, will occupy our attention in the next chapter.

[^25]

## NOTES TO CHAPTER III.

Note 1 (to page 72). It seems desirable to quote here after Lorentz (Abhandlungen iiber theor. Physik, Vol. I. p. 386, footnote) a passage from Maxwell's letter 'On a possible mode of detecting a motion of the solar system through the luminiferous ether,' published after his death in Proc. Roy. Soc., Vol. XXX. (1879-1880), p. 108 :
'In the terrestrial methods of determining the velocity of light, the light comes back along the same path again, so that the velocity of the earth with respect to the ether would alter the time of the double passage by a quantity depending on the square of the ratio of the earth's velocity to that of light, and this is quite too small to be observed.'

Note 2 (to page 73). Usually, at least in all text-books, it is simply said : 'Suppose that the aether remains at rest, and let $v=$ the velocity of the apparatus, i.e. of the earth in its orbit.' For this to be correct, the aether would have to be at rest with respect to our sun. But when astronomical aberration is in question, we are told that the aether is stationary with respect to the 'fixed stars,' say, with respect to the constellation ot Hercules, which, I hope, is 'fixed' enough. Now, as las incidentally been mentioned (p. 17), the sun or the whole solar system has a uniform velocity of something like 25 kilometres per second towards that constellation, which, being nearly equal in absolute value to the earth's orbital velocity ( 30 klm . per sec.), certainly cannot be neglected. Thus, the velocity ( $($ ) of Michelson's interferometer with respect to the aether would oscillate to and fro, in half-year intervals, between considerably distinct maximum- and minimum-values. According to Lorentz ('De Pinfluence du mouvement de la terre sur les phénomènes lumineux,' 1887 , reprinted in Abhandlungen, Vol. I.; see p. 388) the resultant of the earth's orbital and the solar system's velocity had at the time when Michelson was performing his experiment both a direction and an absolute value 'very favorable' to the effect sought for, even so much as to double the displacement of the fringes expected. I am not aware whether or no the defenders and the adversaries of the aether have discussed this circumstance with sufficient care. But at any rate it seemed worth noticing here. Of course, it is for the adherents of the aether (and not those of empty space) to tell us explicitly with respect to what celestial bodies, the sun, or Hercules or other groups of stars, the aether is to be stationary, if it be granted that the parts of that medium do not move relatively to each other. For these stars certainly move relatively to one another.

I cannot help remarking here that it is repugnant to me to think of an omnipresent rigid aether being once and for ever at rest relatively rather to one star than to another. For, this medium, unlike Stokes's aether, being non-deformable and not acted on by any forces whatever, none of the celestial bodies, be it ever so conspicuous in bulk or mass, can claim for itself this primacy of holding fast the aether. The bare idea
of action exerted upon the aether by material bodies being dismissed at the outset, there is nothing which could confer this distinctive privilege upon any one of them. But, then, I am quite aware that what 'is repugnant to think of' may not necessarily be wrong altogether. There are other reasons to be urged against the aether.

Note 3 (to page 75). Let a plane wave $\sigma$ (Fig. 9) proceed towards the inclined mirror (half-silvered plate) $O a$ in the direction of its motion, i.e. from left to right. Let $s O$, sma represent the incident wave normals, limiting a part of the beam of breadth $\overline{O m}=b$, and let $O X X$ be the normal to the mirror, so that $\theta=s O X$ is the angle of incidence. Let the wave reach the centre $O$ of the mirror at the instant $t=0$. Let $O_{1}$ and $a_{1}$ be the positions of the points $O$ and $a$ of the mirror (both taken in the plane

of the figure) at a later instant $t=\tau$, when the wave of disturbance reaches $a_{1}$, so that

$$
\overline{a c_{1}}=\overline{O O_{1}}=v \tau
$$

Draw round $O$ a circle with the radius

$$
\overline{m u a_{1}}=c \tau ;
$$

then the tangent to this circle, drawn from $\alpha_{1}$, will represent the reflected wave, and $O N$ will be the reflected wave normal. To obtain the angle of reflection, $\theta^{\prime}=X O N$, consider the triangles $O N a_{1}$ and $O m a_{1}$, having the side $O a_{1}$ in common and right angles at $m$ and at $N$. Since, moreover, their sides $O N$ and $a_{1} m$ are equal to one another, $\left.\bar{c}_{1} N=O\right) m=b$, so that the breadth of the beam remains unchanged by reflection, as for a stationary mirror, and

$$
\nless N O a_{1} \equiv \epsilon=\Varangle m a_{1} O=\frac{\pi}{2}-\theta-\zeta,
$$

where $\zeta=\nless a O a_{1}$. But $\theta^{\prime}=\pi / 2-\epsilon+\zeta$. Thus, the angle of reffection $\theta^{\prime}$ and the angle of incidence $\theta$ are connected by the relation

$$
\begin{equation*}
\theta^{\prime}=\theta+2 \zeta, \tag{A}
\end{equation*}
$$


where the angle $\zeta$ is determined by the given properties of the parallelogram $O a a_{1} O_{1}$. Writing

$$
\overline{O c}=\overline{O_{1} c_{1}}=l
$$

we have at once

$$
\overline{O c}_{1}^{2}=(v \tau)^{2}+l^{2}+2 v \tau l \cdot \sin \dot{\theta}
$$

and
whence

$$
v \tau: \bar{O}_{1}=\sin \zeta: \cos \theta
$$

$$
\frac{\cos ^{2} \theta}{\sin ^{2} \zeta}=1+\frac{l^{2}}{(\nu \tau)^{2}}+\frac{2 l}{v \tau} \sin \theta
$$

But $v \tau=v l \sin \theta /(c-v)$, or $/ / v \tau=\frac{1-\beta}{\beta \sin \theta}$, so that the required formula
for $\zeta$ is

$$
\begin{equation*}
2 \sin \zeta=\frac{\beta \sin (2 \theta)}{\sqrt{1-\beta(2-\beta) \cos ^{2} \theta}} \tag{B}
\end{equation*}
$$

(A) and (B) contain the rigorous solution of the problem, based, of course, on the assumption of a stationary aether:

In Michelson and Morley's experiment, as treated above (Fig. 8), $2 \theta=90^{\circ}$, so that (B) becomes

$$
\begin{equation*}
2 \sin \xi=\beta\left(\mathrm{r}-\beta+\frac{1}{2} \beta^{2}\right)^{-\frac{1}{2}} \tag{1}
\end{equation*}
$$

To connect Fig. 9 with Fig. 8, notice that, according to ( 1 ), the angle $B O B^{i}$ should be equal to $2 \zeta$. The approximate treatment given in connexion with Fig. 8 (p. 74) amounts to writing

$$
\begin{equation*}
\sin \left(B O B^{\prime}\right)=v: c=\beta \tag{c}
\end{equation*}
$$

Now, developing ( $\mathrm{B}_{1}$ ) and remembering that $\beta$ is a small fraction, we have, up to quantities of the second order,

$$
2 \sin \zeta=\beta+\frac{1}{2} \beta^{2}
$$

or, neglecting the third and higher powers of the small angle $\zeta$,

$$
\sin (2 \zeta)=\beta+\frac{1}{2} \beta^{2}
$$

But the term $\frac{1}{2} \beta^{3}$ appearing in this formula for the angle would give in the final formula for $T_{2}$ only terms of the order of $\beta^{3}$ and $\beta^{\ddagger}$. Thus, aiming at results which are correct only up to quantities of the second order, we may write the last formula

$$
\sin (2 \zeta)=\beta
$$

in agreement with (c). Our Huygens-construction shows then that the treatment adopted on page 74 is sufficiently correct for the purpose in question.

That treatment, which is given in all text-books (including also such valuable modern works as Laue's Relativitätsprinsip, r913) without any further remark, would be rigorously correct if $O$ were, say, a point
source of (spherical) waves spreading out in all directions, and not, as it actually is, one of the points of a mirror at which reflection of plane waves is taking place.
A different way of treating rigorously the above question will be found in Lorentz's paper entitled 'De linfluence du mouvement de la terre sur les phénomènes lumineux,' Arch. néerl., Vol. XXI. (1887), pp. 169-172 (reprinted in Abhandlungen iuber theor. Plysik, Vol. I. pp. 389-392) and partly also in his Columbia University Lectures, p. 194.
The discussion of our general formulae (A), (B) connecting the angle of reflection with that of incidence, for large values of $\beta$, may be left to the reader as a curious exercisc.

## CHAPTER IV.

## EINSTEIN'S DEFINITION OF SIMULTANEITY. THE PRINCIPLES OF RELATIVITY AND OF CONSTANT LIGHTVELOCITY. THE LORENTZ TRANSFORMATION.

We are now sufficiently prepared to grasp the meaning of Einstein's ideas* and to appreciate their relation to the work of his predecessors, especially of Lorentz.

In Chapter I. we have seen how it is possible to define the time as a physically measurable quantity fulfilling certain reasonable and fairly general requirements. Practically, it was the variable $t$ measured by the rotating earth as time-keeper or what, with a slight correction connected with tidal friction, has been called the 'kinetic time.' It has certainly not escaped the reader's notice that the requirements on which that choice was based had nothing absolute or necessary about them, being merely recommended by their simplicity and convenience. But this circumstance need not detain us here any further. Suppose we have secured a clock indicating, with a sufficient degree of precision, the kinetic time $t$. Suppose we keep that clock at a certain place $a$, relatively to a given space-framework of reference, say in a certain physical laboratory or astronomical observatory. Thus far we have tacitly assumed that the time $t$, measured by such a chronometer, is universal, if I may say so, i.e. that it is valid for all points of space, for all parts of any system, be it near to our clock or very far from it, be it at rest or moving with respect to it. It is very likely that nobody has ever

* As laid down in his paper of 1905, already quoted, and then (1907) developed by him more fully in a paper, 'Ueber das Relativitätsprinzip und die aus demselben gezogenen Folgerungen,' Jahrbuch der Radioaktivität und Elektronik, Vol. IV. p. 4 II . In what follows we shall refer principally to the former of these papers by quoting simply the original numbers of its pages.
asserted explicitly this universality and uniqueness of time, but everybody has certainly given to it his tacit consent, and would willingly endorse it if asked to do so. As far as we know, the first to question this pretended universality of time was Einstein.

Our clock, placed at $a$, indicates the time $t$, i.e. marks different time-instants and measures the intervals between them, to begin with, only at the place a, or nearly so. It is, to give it a short name, the time $t_{a}$. Suppose that some well-marked instant is chosen as the initial instant, $t_{a}=0$. Then, if any event is happening at $a$ or near $a$, we give to it that date or, as it were, label it with that number $t_{a}$ which is simultaneously shown by the index of our clock. We are exempted from defining what 'simultaneous' (as well as 'earlier' or 'later') means when applied to a pair of events occurring at the same place or near that place, as the passage of the index through a given division of the dial of our clock and the production of an electric spark closely to it.* But we do not know, beforehand, what we are to understand by saying that of two events occurring at places $a, b$ distant from one another the first occurs earlier or later than the second, or that both are simultaneous. The meaning of these words has to be defined. If the labelling of all possible kinds of events, occurring at distant points, fixed or moving relatively to one another, is to be of any use at all, we must establish the rules according to which we are going to label them with the $t$-numbers. And first of all we have to decide which of these events have to receive the same labels, i.e. we have to define simultaneity at distant points.

This notion is to be defined in terms of simultaneity at the same place, which alone is assumed to be known to us, and of some other things or processes which are actually realizable. In other words, distant simultaneity has to be reduced to local simultaneity by some physical process. Abstractly speaking, the choice of such a process is arbitrary, in very wide limits at least; but practically the choice will be reduced to such processes as are of possibly universal occurrence, and which are independent of the capricious peculiarities of different sorts of matter. Einstein has chosen for this purpose the propagation of light in vacuo. Gravitation being, chiefly due

[^26]to its alleged instantaneous action, out of question, this has been, in fact, the only possible choice. Moreover, it was not unprecedented in the history of physics and astronomy, and it suggested itself most obviously because the recent difficulties met with lay in the optical and, more generally, electromagnetic departments of physics.

To an unbiassed mind the question may present itself: Why label everything in the world with $t$-numbers at all? Such a question is not altogether unreasonable, and it may deserve some careful attention. But once we decide to attach a time-label to every event, we are forced to reduce in some kind of way distant simultaneity to local simultaneity (for pairs of points at rest or moving relatively to one another), and not to delucle ourselves with thinking that we know what 'universal simultaneity' means, or that it is, in fact, a self-consistent notion. To have initiated a critical analysis of the concept of simultaneity at all is certainly a great merit of Einstein's.

But let us leave aside these generalities and pass to the definition in question. We shall have to consider in the first place the simpler case of distant points $a$, $b$, etc., in relative rest, and then the somewhat intricate case of distant points belonging to systems which are uniformly moving with respect to one another.

Let $a$, $l$, etc., be points or 'places' fixed relatively to one another and with respect to a certain space-framework or system $S$, say, the system of the fixed stars.* Suppose we succeeded in manufacturing at the place $a$ a number of equal clocks, each measuring the same, say the 'kinetic,' time $t$ and set equally or synchronously, and that retaining one of them at $a$ we sent the others to $b$, etc., together with an equal number of observers who are to remain at those distant places with their clocks for ever. Then, to begin with, we should have as many 'times' as there are places in consideration, $t_{a}, t_{b}$, etc., valid, respectively, for the places $a, b$, etc., and for their nearest neighbourhoods. For, though all of these clocks were manufactured equally at $a$, we do not know whether they continue to be 'equal' or permanently synchronous, when one of them is

[^27]still kept at $a$, while the others are sent far away, to the places $b$, etc. More than this, we do not know what their being synchronous
or not, when far apart, means. We have yet to fix how we are going to test it. To invoke the preservation of rate of clocks of 'grood make' in spite of their being carried to distant places, on the title of the high precision of their mechanisms, would not help us out of the difficulty. For, supposing we also decided to assert such infallible and rigorous permanence, at different places within $S$, of the mechanical laws, necessarily involved, still we should have to verify whether the accessorial conditions of validity of those laws (and practically there would be a host of such conditions) are fulfilled at and round each place in question. To avoid this verification, which soon would prove to be a difficult task, we must have some means of testing in a clirect manner the synchronism of our distant clocks and, more generally, of correlating with one another the times $t_{a}, t_{b}$, etc., without being obliged to enter upon the properties and structure of the corresponding clock mechanisms.*
Now, the kind of test adopted by Einstein, and constituting at the same time the essence of his definition of distant simultancity, is as follows.

Let an observer stationed at $a$ send $a$ flash of light at the instant $t_{a}$ (as indicated by the $a$-clock) towards $b$, where it arrives at the instant $t_{b}$ (according to the $b$-clock). Let another observer send it back from $b$ without any delay, or let the flash be automatically reflected at $b$, towards $a$, where it returns at the instant $t_{a}{ }^{\prime}$. Then the $b$-clock is said, by definition, to be synohronous with the $a$-clock, if

$$
\begin{equation*}
t_{a}^{\prime}-t_{b}=t_{b}-t_{a} . \tag{1}
\end{equation*}
$$

This amounts to requiring, per definitionem, that 'the time' employed by light to pass from $a$ to $l$ should be equal to 'the time' employed to return from $b$ to $a$. Instead of (r) we may write, equivalently,

$$
\begin{equation*}
t_{u}=t_{a}+\frac{1}{2}\left(t_{a}^{\prime}-t_{u}\right)=\frac{1}{2}\left(t_{a}+t_{a}^{\prime}\right) . \tag{1a}
\end{equation*}
$$

Thus, the instant of arrival at $b$ is expressed by the arithmetic mean of the $a$-times of departure and return of the light-signal. Such

[^28]being the connexion of the $a$-time and of the $b$-time, the clock placed at $b$ is said to be synchronous with that placed at $a$.

This definition of synchronism is supposed to be self-consistent, for any number of clocks placed at different points of the system $S$, say, besides $a$ and $b$, at $c, d, e$, etc. To secure this consistency, Einstein makes, explicitly, the following two assumptions:
r. If the clock at $b$ is synchronous with that at $a$, then also the clock at $a$ is synchronous with that at $b$. In other words: clocksynchronism is reciprocal, for any pair of places taken in $S$.
2. If two clocks, placed at $a$ and $b$, are synchronous with a third clock, placed at $c$, they are also synchronous with one another. Or, more shortly, clock-synchronism is transitive throughout the system $S$.

This is the way that Einstein himself puts the matter. But it may easily be shown that the first of his assumptions will be fulfilled if we require that 'the time' employed by the light-signal to pass from $a$ to $b$ is alrvays the same. In fact, let us denote the $a$-time, taken generally, by $a$ instead of $t_{a}$, and similarly, let us write $b$ instead of the general variable $t_{b}$, and let us use the suffixes $a_{,} a_{2} r$ to denote the instants of departure, arrival and return. Then, if the $b$-clock is synchronous with the $a$-clock, we have, by definition, $b_{a}=\frac{1}{\Sigma}\left(a_{a l}+a_{r}\right)$, or

$$
b_{a}-a_{d}=a_{r}-b_{a}=a_{a}-b_{a},
$$

for the 'return' at $a$ may be equally well considered as an arrival at that place. Now, if at the instant $a_{a}$ the flash be sent again towards $b$, where it arrives at the instant $b_{r}$, we have, by our above requirement,

$$
b_{a}-a_{a}=b_{r}-a_{a}
$$

and, by the last equation,

$$
a_{a}-b_{a}=b_{r}-a_{a}
$$

But here $b_{a}$ is identical with the instant of departure $b_{d}^{\prime}$, and, consequently,

$$
a_{a}=\frac{1}{2}\left(b_{l a}+b_{r}\right),
$$

i.c. the clock placed at $a$ is synchronous with that placed at $b$. Q.E.D.

A similar treatment of assumption 2. may be left to the reader, who will find sufficient hints in Fig. ro. This assumption will be easily seen to imply that if a pair of flashes be sent out simultaneously
from $a$, one via $b, c$ and the other via $c, b$, they will both return simultaneously at $a$. More generally, the time elapsing between the instant of departure and that of return of the light-signal sent round alcoz will be equal to the time elapsing between departure and return of the signal sent round acba, and similarly for every other closed path in $S$, both times being measured by the clock placed at $a$. This form of the property attributed to the system $S$ is worthy or being especially insisted upon, as it implies only operations to be performed at one and the same spot. To state this property of the system $S$, the observer has not to move from his place.


Fig. 10.
Such then are the physical properties of this system of reference.
It is strange that Einstein, after having made explicitly the above assumptions r. and 2., considers it necessary to add (p. 894) that 'according to experience' the quantity

$$
2 \frac{\overline{a b}}{a_{r}-a_{d}}=c,
$$

or, in the notation of formula ( r ), the quantity

$$
\begin{equation*}
2 \frac{\overline{a b}}{t_{u}^{t}-t_{u}}=c, \tag{2}
\end{equation*}
$$

is to be taken as 'an universal constant (the velocity of light in empty space).' At any rate, if the last assumption is made, for any pair of points $a, b$ in $S$, once and for ever, then the above statements r . and 2 . are certainly superfluous. But considerations of this order need not detain us here any more.
S.R.

In this way the various times, $t_{a}, t_{b}$, etc., originally foreign to one another, are all connected so as to constitute one time only, valid for the whole system, which we may denote simply by $t$, calling it shortly the s-time.

There is, thus far, nothing essentially new in Einstein's procedure. It was more or less unconsciously applied since people began to measure the velocity of light, and even sound, nay, since they began to exchange with one another letters or messages of any kind. The novelty does not come in until the next stage, when the time-labelling is extended to different systems moving (uniformly) with respect to one another.
Let $S$ be as above, and let us consider other systems of reference, $S^{\prime}, S^{\prime \prime}$, and so on, each having with respect to $S$ a motion of uniform (rectilinear) translation. Having settled the matter for the system $S$, i.e. having established the $S$-time, $t$, let us similarly establish an $S^{\prime}$-time, $t^{\prime}$, an $S^{\prime \prime}$-time, $t^{\prime \prime}$, etc., and let us see how the times $t^{\prime}, t^{\prime \prime}$, etc., are connected with the time $t$ valid for $S$. It can reasonably be expected that these processes of (time-) labelling of events happening at different places, being undertaken from different standpoints, $S, S^{\prime}, S^{\prime \prime}$, etc., will generally not coincide with one another, e.g. that events obtaining identical t-labels may receive different $t^{\prime}$-labels, and so on. Such, in fact, will be the case; the labels of different sorts, dashed and non-dashed, though none is privileged in any way, will have to be carefully distinguished from one another. In a word, it will appear that, with the above definition of simultaneity, no universal, no unique time-labelling is possible.

It will be enough to consider explicitly, besides $S$, one other system only, say, $S^{\prime}$. Supposing that a consistent time-labelling of events occurring at different places of $S^{\prime}$ or an $S^{\prime}$-time, is possible, like the above $S$-time, the question is, how is this time $t^{\prime}$ to be connected with the time $t$ ? We shall see that the connexion sought for will involve also the coordinates defining the position of points

[^29]






















pulse or wave of discontinuity is a spherical surface, of centre $P$ and of radius
$$
r=c t,
$$
if $t$ is reckoned from the instant of emission. Again, if $P^{\prime}$ is a point fixed in $S^{\prime}$, and if the arbitrarily moving source emits a flash just when it is passing through $P^{\prime}$, then the wave, ats it appears to observers rigidly attached to $S^{\prime}$, will be a sphere whose centre is permanently situated at $P^{\prime}$ and whose radius at any instant of the $S^{\prime}$-time is
$$
r^{\prime}=c t^{\prime},
$$
if $t^{\prime}$ is reckoned from the instant of emission. Such is, in virtue of $I$, the meaning of the principle of 'constancy' of light-propagation in empty space. Of especial interest is the particular case, in which our source is fixed at a point $P^{\prime}$ of the system $S^{\prime}$, and therefore moving uniformly with respect to $S$. In this case the centre of the spherical wave will, to the $S^{\prime \prime}$-observers, bo permanently situated at the material particle playing the part of source, whereas for the $S$-observers the centre of the spherical wave, fixed in $S$, will detach itself from the material source, the source moving away from it with uniform velocity together with the whole system $S^{\prime \prime}$. This case will be made use of presently.

Let us now return to the first of the above principles, and let us remember how the time $t$, valid for the whole system $S$, has been defined. Since $S$ has been endowed with physical properties required for a consistent method of time-labelling of events occurring at its various points, the same propertics will, in virtue of I., hold also for $S^{\prime}$. Again, local clocks satisfying the requirements of convenience, e.g. the causality-maxim, being possible in $S$, such time-keepers are, by I., possible also for various stations taken in $S^{\prime}$. We can therefore consider first a time $t_{a}^{\prime}$, measured by a clock placed at a point $a^{\prime}$ in $S^{\prime}$, then distant clocks placed at $l^{\prime}$, etc., leaving the task of testing their synchronism to observers attached to the system $S^{\prime}$, and repeating in fact literally all that has been said before with regard to the system $S$. In this way we should obtain out of the originally local times a unique time $t^{\prime}$ applicable to the whole system $S^{\prime}$. Let us call the time thus constructed the $\mathrm{s}^{\prime}$-time.

The question now is, how are the $S^{\prime \prime}$-time and the $S$-time connected with one another (and, possibly, with other things, viz. lengths or distances as measured by the $S$ - and $S^{\prime}$-observers)?

The answer to this fundamental question nay be obtained, with the help of the two above principles, in a variety of ways. But for certain reasons the following way, though not the shortest, scems to me the most instructive to begin with.* It is, moreover, intimately connected with what has been said in the last chapter with regard to the Michelson experiment.
Let us imagine an $S^{\prime}$-observer having at his, clisposal a point-source of light at a place $P^{\prime}$ fixed in the system $S^{\prime}$. Let $A^{\prime}$ and $B^{\prime}$ be a pair of distant points also fixed in $S^{\prime}$, and such that the straight line $P^{\prime} A^{\prime}$ is in the direction of motion of $S^{\prime}$ relatively to $S$, and that ' $B^{\prime}$ is perpendicular to that direction (Fig. ri). As before, we shall call $P^{\prime} A^{\prime}$ longitudinal, and $P^{\prime} B^{\prime}$ transversal. Let $l^{\prime}$ be the

length' of the first of these segments or the 'distance' from $f^{\prime}$ o $A^{\prime}$, according to the estimation of the $S^{\prime}$-inhahitants, and similarly the length of the second segment. Suppose that our obscrver nds an instantaneous light-flash from $P^{\prime}$ towards $A^{\prime}$ and receives back at $P^{\prime}$ after the lapse $\theta^{\prime}$ of the $t^{\prime}$-time. Then, having assured imself by any means that an assistant stationed at $A^{\prime}$ sends him ack his signals without any delay, our observer will write

$$
\theta^{\prime}=\frac{2 l^{\prime}}{c} .
$$

Inder similar conditions, if he sends a flash towards $\dagger \not \beta^{\prime}$ and * Einstein's method of reasoning, as given in his original paper ( $\$ 3$, see alson otes at the end of this Chap.) may le mathomatically interesting, but dres mot em to be the fittest when a clear discussion of the physical aspect of the question aimed at.
$\dagger$ To avoid unnecessary difficulties as to hitting the receiving station, now $i^{\prime \prime}$ d now $A^{\prime}$, it will be best to imagine that our observer sends each time a full herical wave of discontinuity or a very thin spherical pulse. This will lie and especially convenient when we come next to consider the same processes on the $S$ standpoint.
receives it back after the interval $\tau^{\prime}$ of the $t^{\prime}$-time, he will put down the equation

$$
\tau^{\prime}=\frac{2 s^{\prime}}{c} .
$$

There is, in fact, by the above principles, no difference between longitudinal and transversal light signalling between stations fixed in $S^{\prime \prime}$, as observed by the inhabitants of this same system.

Let us now see how each of the above two processes will be described by an observer attached to the system $S$. Call the lengths or distances $P^{\prime} A^{\prime}, P^{\prime} B^{\prime}$, as estimated by the $S$-observer, $l$ and $s$ respectively. Each of thesc is obtained by ascertaining, with the help of an appropriate number of synchronous $t$-clocks, which are the points of the $S$-system, through which $P^{\prime}$ and $A^{\prime}$, or $P^{\prime}$ and $B^{\prime}$ pass simultaneously, and by measuring the mutual distances of these points by means of an $S$-standard rod. Similarly, $l^{\prime}$ and $s^{\prime}$ are to be considered as the distances $P^{\prime} A^{\prime}$ and $P^{\prime} B^{\prime}$ measured by standard rods which the $S^{\prime}$-observers are carrying with themselves. Notice that, by Principle I., $l^{\prime}$ and $s^{\prime}$, thus measured, will be the same whether the system $S^{\prime}$, together with its observers, clocks and measuring rods, is at rest with respect to $S$ or whether it moves uniformly with respect to that system, as it actually does. But $l, s$ are not necessarily equal to $l$ ', $s$ '. For although they are 'distances of the same pairs of material points,' the source and the receiving stations, they are not obtained by the same processes. Having thus explained the meaning of $l, s$, let us consider, from the S-standpoint, first the longitudinal and then the transversal signalling. The flash sent out by the luminous source will, according to Principle II., appear to the S-observers in both cases as a spherical wave expanding with the velocity $c$ and having its centre at that point $P_{0}$, fixed in $S$, through which the source has passed when emitting the flash. Now, if $v$ be the velocity of $S^{\prime}$ relative to $S$, the receiving station $A^{\prime}$ moves away from $P_{0}$ with the uniform velocity $v$. If, therefore, $\theta_{1}$ be the $S$-time required for the wave to expand from $P_{0}$ to $A^{\prime}$,

$$
c \theta_{1}=l+v \theta_{1}
$$

and

$$
\theta_{1}=\frac{l}{c-v} .
$$

In the same way, if $\theta_{2}$ be the $S$-time employed by the light
o return from the receiving station* to the sending station $P^{\prime}$,

$$
\theta_{2}=\frac{l}{c+v} .
$$

['hus, the $S$-time $\theta$ elapsing between the first appearance and the eappearance of a light flash at $A^{\prime}$, being the sum of $\theta_{1}$ and $\theta_{2}$, vill be given by

$$
\theta=\frac{2 l c}{c^{2}-v^{2}} .
$$

A similar reasoning applied to the case of transversal signalling, in which case the sphericity of the wave will be found particularly onvenient, will give us for the $S$-time elapsing between the appearnce of the first and second flash at $A^{\prime}$ the value

$$
\tau=2 \gamma \frac{s}{c}
$$

here $\gamma=\left(\mathrm{r}-\beta^{3}\right)^{-\frac{1}{2}}, \beta=v / c$.
Compare the last two formulae with the above ones for $\theta^{\prime}$ and $\tau^{\prime}$, nd denote the ratio $s / s^{\prime}$ by $a$. Then the result will be

$$
\begin{equation*}
\stackrel{\theta}{\theta^{\prime}}=\gamma^{2} \frac{l}{l^{\prime}} ; \quad \frac{\tau}{\tau^{\prime}}=\gamma a ; \quad \frac{s}{s^{\prime}}=a, \tag{3}
\end{equation*}
$$

here $a$ is a number which for $y=0$ becomes equal x , but is otherise an unknown function of the data of the problem.
Now, each of the two processes, i.e. the longitudinal and the ansversal signalling, may (by disregarding the receiving stations) e considered as phenomena consisting in a clouble appearance of flash at one and the same station, at the same individually disernible point $A^{\prime}$, fixed in $S^{\prime}$. Thus far we have, purposely, kept ese two processes separate. But now we can advantageously ombine them with one another. If the receiving stations were oosen so that $s^{\prime}=l^{\prime}$, then we should have, by the first pair of rmulae,

$$
\theta^{\prime}=\tau^{\prime}, \quad \text { say }=T^{\prime \prime}
$$

d if the two processes were started simultaneously, from the -point of view, they would also have ended simultaneously for e $S^{\prime}$-inhabitants. In other words, we would have, in $S^{\prime}$, a pair of

* This station $A^{\prime}$ (and similarly, in the case of transversal signalling, the station may be imagined to become an instantaneous point-source emitting a spherical ve at the moment when it is reached by the original wave.
simultaneous events followed by another pair of simultaneous events, all. of these occurring at the same place $A^{\prime}$. Let us now require (what, as far as I know, is tacitly assumed by most authors) that
> III. Events locally* simultaneous for an $S^{\prime \prime}$-olserver should also be simultaneous for the $S$-oliservers.

This amounts to supposing that there is a one-to-one correspondence between the $t$-labels and the $t^{t}$-labels to be applied to events occurring at any given place, i.c. for $f$ fived values of the coordinates $x^{\prime}, y^{\prime}, z^{\prime}$ in $S^{\prime}$. (The analogous one-to-one correspondence between $x^{\prime}, y^{\prime}, z^{\prime}$ and $x, y, z$ for $t^{\prime}=$ const. is tacitly assumed as a matter of course.) On the other hand, two events occurring at distinct places, being simultaneous in $S^{\prime}$, are generally non-simultaneous from the $S$-standpoint.

Now, in virtue of the requirement III., call it postulate or desideratum, or whatever you prefer, the above two simultaneous processes or phenomena occurring at $A^{\prime}$ will also begin and end simultaneously for the $S$-observers, so that

$$
\theta=\tau, \quad \text { say }=T,
$$

and

$$
\theta / \theta^{\prime}=\tau / \tau^{\prime}=T / T^{\prime} .
$$

Consequently, by the equations (3),

$$
\left.\begin{array}{r}
T / T^{\prime}=u \gamma=a\left(\mathrm{x}-\beta^{2}\right)^{-\frac{h}{t}}  \tag{4}\\
l / l^{\prime}=a \gamma^{-1} ; \quad s / s^{\prime}=a .
\end{array}\right\}
$$

These are the required comexions between durations and lengths, measured in $S$ and in $S^{\prime}$. They are based on the above assumptions I., II., III., the last of which is certainly the most obvious. The common coefficient $a$ is, thus far, indeterminate. If we are to endow (empty) space with homogeneity, as well as with isotropy, $\dagger$ and if it be granted that the relations between the $S$ - and $S^{\prime}$-measurements do not vary in time, the unknown coefficient $u$ can depend only upon $v=c \beta$. The only thing we thus far know about this

[^30]function is that it reduces to unity for $\beta=0$, when $S^{\prime}$ is at rest relatively to $S$, when, in fact, both systems cease to be discernible from one another. Thus
$$
u=u(\beta), \quad \alpha(0)=\mathrm{r} .
$$

Notice that for $\eta=0$ we have also $\gamma=\mathrm{r}$, so that in this casc $Z, l, s$ become, by (4), identical with $T^{\prime}, l^{\prime}, s^{\prime}$, as was to be expected.

To put the relations (4) in words as simply as possible, and to fix the ideas, let us assume for the moment $\alpha=1$. Then

$$
\begin{equation*}
\frac{T}{T^{\prime}}=\gamma ; \quad \frac{l}{l^{\prime}}=\frac{\mathrm{x}}{\gamma} ; \quad s=s^{\prime} \tag{+fa}
\end{equation*}
$$

Thus, a transversal bar sharing the motion of $S^{\prime}$ will have the same length from the standpoint of either of the two systems $S, S^{\prime} S^{\prime}$, while a bar of longitudinal orientation and of length $l^{\prime}$ in $S^{\prime}$ will, according to the estimation of the $S$-observers (with equal t-values for both terminals of the bar), be shortened to $l=l^{\prime} \sqrt{r}-\beta^{2} . ~ \Lambda ~ s o l i d ~ f i x e d ~$ in $S^{\prime}$, which for the inhabitants of that system is a sphere of radius $R$, will, according to the estimation of the $S$-observers, become a longitudinally flattened ellipsoid of semi-axes

$$
\frac{\mathrm{x}}{\gamma} R, R, K
$$

precisely as in the contraction hy'pothesis of Fitzgerald and Lormilk.. It is a slightly different thing to say, instead of this, that a bocly which for the $S$-observers is spherical while at rest in $S$ becomes flattened down to the above ellipsoid when set in motion with the translation-velocity $v$ relative to $S$. The clause hinted at is in connexion with the manner in which the body is set from rest to motion and cannot satisfactorily be dealt with at this stage of our considerations. Again, as regards the ratio of times, remember that $T^{\prime}$ is the $S^{\prime}$-duration of a phenomenon or process groing on at a place $P^{\prime}$ fixed in $S^{\prime}$, i.e. for constant $x^{\prime}, y^{\prime}, z^{\prime}$. This cluration or time-interval is then lengthened in the estimation of the $S$-observers - $T^{\prime}=\gamma T^{\prime}=T^{\prime} \sqrt{\mathrm{I}-\beta^{2}}$. We are assuming here, of course, that $\beta<\mathrm{r}$, so that $\gamma$ is real and greater than unity. Instead of a pair of flashes, as considered above, we may think of two consecutive indica tions of an $S^{\prime}$ clock preserved at $P^{\prime}$, and we may say that a cloc:k moving relatively to $S$ with the uniform velocity $v$ goes sloverer, in the
ratio $\sqrt{\mathrm{I}-\beta^{22}}: \mathrm{x}$, than 'the same' clock when at rest in $S$. This at
'The same' is taken to mean that the mechanism of the clock has undergone no changes due to its passing from rest to motion, except those which are implied by the fundamental relativistic principles themselves. This statement does by no means look satisfactory, but it can be made more rigorous and clear by returning to it after certain portions of relativistic physics have been worked out. 'The practically important question is, which are the physical systems we are going to consider as such clocks whose ' internal mechanism' is not subject to changes due to their merely passing from rest to uniform motion relatively, say, to the earth or the fixed stars. Now, as far as I know, the prevailing tendency is to consider as such physical systems the various atoms (or at least, if they are to serve us for thousands of years, those which are not sensibly radioactive) with their 'natural' periods of vibration, manifested in their characteristic spectrum lines.* The influence felt by such minute mechanisms in the presence of a strong magnetic field (Zeeman's effect) will not, of course, be forgotten. Who knows but that some remote future generations, to get rid of such physical influences, may choose to consider as 'invariable' the mechanism not of light emission but of radioactive disintegration of atoms. If such is to be the case, the formula $T=\gamma T^{\prime}$ will be interpreted by saying that the 'half-life' of radium, which is about 1760 years, is in the estimation of a terrestrial observer lengthened by a month or so, when flashing before him with something like one hundredth of the velocity of light.

We have already remarked in passing that two events occurring simultaneously in $S^{\prime}$ at places distant from one another will generally be non-simultaneous to the $S$-observers. 'This may be seen immediately by the principle of constant light-velocity, valid by I. for both $S$ and $S^{\prime}$. For let a spherical wave or a very thin pulse be started from our point-source placed at $P^{\prime}$. Then, if $l^{\prime}=s^{\prime}$, the arrivals of flashes at $A^{\prime}$ and $B^{\prime}$ will be a pair of events simultaneous
*Thus we read, in M. Laue's Relativilötsprinzip, second edition, 1913, p. 42 : ' In einem lewegten Wasserstoffatom (Kanalstrahlen) werden, zum Beispiel, die Licht emittierenden Eigenschwingungen geringere Frequenz haben, als in einem ruhenden.'

As regards the experimental side of the subject, see J. Laub's report in Jahrb. d. Rad. u. Elektronik, Vol. VII. p. 439.
, the $S^{\prime}$-observers. On the other hand, the $S$-time required for he wave to reach $A^{\prime}$ will be

$$
T_{P^{\prime} A^{\prime}}=\stackrel{l}{c-v},
$$

nd that to reach $B^{\prime}$

$$
T_{P^{\prime \prime} B^{\prime \prime}}=\gamma \frac{s}{c} .
$$

Jow, by (4a), and also by the more general formulac (4),

$$
l / s=\gamma^{-1} l^{\prime} / s^{\prime}=1 / \gamma,
$$

hence

$$
T_{P^{\prime} A^{\prime}}-T_{P^{\prime} B^{\prime} B^{\prime}}=\frac{\beta \gamma}{6} s
$$

hus, the pair of events in question will not be simulancous for oe $S$-observers. Instead of the two particular points $A^{\prime}, \beta^{\prime}$, the hole wave may be considered. Then it will be seen at once that ee sphere $r^{\prime}=$ const. with centre $P^{\prime}$ will, to the $S^{\prime \prime}$-(bservers, be the ous of points reached simultaneously by the wave, but not so to te $S$-observers. For to these the loci of simultaneously illuminated oints will be spheres centered at a point, $P_{0}^{\prime}$, fixed in $S$, from hich $P^{\prime}$ is continuously moving away.
Thus, the notion of distant simultancity, to call it again by this hort name, has no 'absolute' or universal meaning, but involves a pecification of one out of $\infty^{8}$ systems of reference. For such is re manifold of the vector-values of their relative volocity $\mathbf{v}$, its bsolute value $v$ amounting to one scalar, and its direction to two 1ore.
Let us now once more return to our formulae (4), with the view $f$ deducing from them the relations connecting the $S$-time and oordinates $t, x, y, z$ with the $S^{\prime}$-time and coordinates $t^{\prime}, x^{\prime}, y^{\prime}, z^{\prime}$. ake the $x^{\prime}$-axis coincident in direction and sense with the $x$-axis, oth concurrent with the vector $\mathbf{v}$ fixing the velocity of $S^{\prime \prime}$ relative ) $S$ (Fig. Ir),* and the axes of $y^{\prime}, z^{\prime}$, both transversal and perendicular to one another, parallel to and concurrent with the axes f $y, z$ respectively. Count both the $S^{\prime}$ - and the $S$-time from the
*In that figure the systems $S^{\prime}, S$ are represented as sliding along one another nly to avoid confusion in the drawing, but in reality they are to be imagined an iterpenetrating one another throughout the whole (three-dimensional) space.
instant at which the origins of the coordinates, $O^{\prime}$ and $O$, coincide with one another, i.e. assume $\cdot$.
as corresponding to $\because t=x=y=s=0$

$$
t^{\prime}=x^{\prime}=y^{\prime}=z^{\prime}=0
$$

which is a.pure convention. The axes of $y^{\prime}$ and $z^{\prime}$ will then coincide at that instant with the axes of $y$ and $z$. Let us fix our attention on any point $P^{\prime}\left(x^{\prime}, y^{\prime}, s^{\prime}\right)$ taken in $S^{\prime}$. Then by the second of formulae (4), in which we have to write $l=x-v t, l^{\prime}=x^{\prime}$,

$$
\begin{equation*}
x^{\prime}={ }_{\alpha}^{\gamma}(x-v t) \tag{5}
\end{equation*}
$$

and, by the third of those formulae,

$$
\begin{equation*}
y^{\prime}=\frac{1}{a} y^{\prime} ; \quad z^{\prime}=\frac{1}{\alpha} z \tag{6}
\end{equation*}
$$

To obtain $t$ as a function of $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$, notice first of all that events occurring at various points of a transversal plane ( $x^{\prime}=$ const.), being simultaneous in $S^{\prime}$, are also simultaneous with one another according to the $S$-point of view. For if $M^{\prime}, N^{\prime}$ be a pair of such points, and if $\overline{M^{\prime}} N^{\prime}=s^{\prime}$, then a wave started at their mid-point $C^{\prime}$ at the instant $t^{\prime}-\frac{1}{2} s^{\prime} / c$ will reach both $M^{\prime}$ and $N^{\prime}$ simultaneously, at the instant $t^{\prime}$. Again, from the $S$-standpoint, in our previous notation,

$$
T_{Q^{\prime} H^{\prime}}=\frac{\gamma s}{2 c}=T_{l^{\prime \prime} N^{\prime}}
$$

so that $M^{\prime}$ and $N^{\prime}$ will receive the signals at the same instant $t$. Thus, $t$ is independent of $y^{\prime}, z^{\prime}$, and consequently

$$
t=t\left(x^{\prime}, t^{\prime}\right)
$$

Next, take a longitudinal pair of points, say $P^{\prime}$ on the $x^{\prime}$-axis and the origin $O^{\prime}$. Call $x^{\prime}$ the abscissa of $P^{\prime}$. Imagine a wave started at the mid-point of $O^{\prime}$ and $P^{\prime}$ at the instant $t^{\prime}-\frac{1}{2} x^{\prime} / c$; then the wave will reach $O^{\prime}$ and $P^{\prime}$ at the same instant $t^{\prime}$, and, by Principle II. and by the second of formulae (4),

$$
t\left(x^{\prime}, t^{\prime}\right)-t\left(0, t^{\prime}\right)=\frac{\alpha}{\gamma} \frac{x^{\prime}}{2}\left(\frac{\mathrm{I}}{c-v}-\frac{\mathrm{r}}{c+v}\right)=\alpha \gamma \frac{v}{c^{2}} x .
$$

But, by the first of formulae (4) and by the aborg convention as to the origin of time-reckoning

Hence

$$
\begin{aligned}
& t\left(o, t^{\prime}\right)=\alpha \gamma t^{\prime} . \sin \\
& \left.t^{\prime}\right)=\alpha \gamma\left(t^{\prime}+\frac{v}{c^{2}} x^{\prime}\right)
\end{aligned}
$$

which is the required connexion. Substituting here and remembering that $\beta^{2}+1 / \gamma^{2}=1$, we shall obtain of $t, x$.

Thus, the complete set of formulae connecting the $S^{\prime}$ - with the S-time and coordinates will be

$$
\left.\begin{array}{c}
x^{\prime}=\frac{\gamma}{\alpha}(x-v t) ; \quad y^{\prime}=\frac{1}{\alpha} y ; \quad z^{\prime}=\frac{r}{\alpha} z  \tag{8}\\
t^{\prime}=\frac{\gamma}{\alpha}\left(t-\frac{v}{c^{2}} x\right) .
\end{array}\right\}
$$

Conversely, resolving these equations with respect to $t, x, y, z$, or simply copying (7) and using it to eliminate $t$ from the first equation,

$$
\left.\begin{array}{c}
x=\alpha \gamma\left(x^{\prime}+v t^{\prime}\right) ; \quad y=\alpha y^{\prime} ; \quad z=\alpha z^{\prime}  \tag{9}\\
t=\alpha \gamma\left(t^{\prime}+\frac{v}{c^{2}} x^{\prime}\right) .
\end{array}\right\}
$$

Notice that, disregarding $\alpha$, the set (9) follows from ( 8 ), and vice versca, by simply interchanging $x, y, z, t$ with $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$ and by writing $-v$ instead of $v$. Now $\nabla$ being the velocity of $S^{\prime}$ relative to $S,-\mathbf{v}$ will be the velocity of $S$ relative to $S^{\prime}$.* As to $c$, it is common to both systems, and $\gamma(v)=\gamma(-v)=\left(I-v^{2} / c^{2}\right)^{-\frac{1}{2}}$. Thus, there is reciprocity between the two systems of reference, except for the common arbitrary coefficient which is $\alpha^{-1}$ in the

[^31]first, and $a$ in the second set of formulac. As a matter of fact, there is a plyysical reciprocity anyhow, i.e. for any $\alpha=\alpha .(v)$, subjected to the condition $a(0)=\mathrm{I}$. For the conditions imposed upon the time-labellings in $S$ and in $S^{\prime}$, in order to make them self-consistent, will continue to be satisfied when all values of time and coordinates, in $S$ or in $S^{\prime}$, have been multiplied by a common factor; $a^{-1}$ in one, and $a$ in the other case may be thrown back upon the choice of the units of measurement. Thus, the choice of a being a matter of indifference, we may take $a=\mathrm{r}$. But, if not content with the physical, we require also a formal reciprocity, then we have to write
$$
a^{-1}=a, \quad \text { i.e. } a^{2}=\mathrm{I} .
$$

But $a(o)=\mathrm{r}$. Thus, if $\alpha(v)$ is to be continuous, $\alpha=+\mathrm{r}$.*
In this way we obtain the formulae of what is universally called the Lorentz transformation,

$$
\left.\begin{array}{l}
x^{\prime}=\gamma(x-v t) ; \quad y^{\prime}=y^{\prime} ; \quad s^{\prime}=z  \tag{10}\\
t^{\prime}=\gamma\left(t-\frac{v}{c^{i}} x\right), \quad \check{v}^{2} /
\end{array}\right\}
$$

already met with in Chap. III. But here, as can be judged from the whole line of reasoning, the meaning and the rôle of this transformation are essentially different from what they were in Lorentz's theory, based as it was on the assumption of a privileged system of referencc, the rether.

Let us write also the inverse transformation

$$
\left.\begin{array}{l}
x=\gamma\left(x^{\prime}+v t^{\prime}\right) ; \quad y=y^{\prime} ; \quad z=z^{\prime} \\
t=\gamma\left(t^{\prime}+\frac{z^{\prime}}{c^{2}} x^{\prime}\right) .
\end{array}\right\}
$$

The above postulate I., or the Principle of Relativity, may now be expressed in the concise and more definite form:

Ia. The lazes of physical phenomena, or rather their mathe-
matical expressions, are invariant zeith respect to the Lorentz
transformation. $\dagger$

[^32]That is to say, if a law $L$, valid in $S$, involves-besides other magnitudes $-x, y, z, t$ in a certain way, and if these are transformed according to (ro), then the resulting law $L^{\prime}$, valid in $S^{\prime}$, will involve $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$ in exactly the same way. Any system $S^{\prime}$, with its corresponding tetrad of independent variables, is as 'legitimate" as $S$. The choice of one out of $\infty^{3}$ systems of reference moving uniformly with respect to one another is a matter of indifference. As regards the behaviour of the 'other magnitudes' involved in the laws, any attempt to elucidate it by general remarks in this place would be useless. We shall come to unclerstand this point by and by when considering various applications of the above principle. And, with regard to the specification 'physical,' it has, of course, to be taken in the broadest sense of the word. 'The phenomena in question may as well be chemical or physiological (though, for the present, physiology is far from being prepared to receive a theory of such a high degree of accuracy). Instead of 'physical phenomena' the reader can, at any rate, put theoretically : any phenomena which are at all localizable in space and in time. But subtleties of this kind need not detain us here any further.

The principle of relativity excludes all such laws as are not invariant with respect to the Lorentz transformation. Thus, for instance, Newton's inverse square law of universal gravitation, or even his general laws of motion, cannot stand in their original form, but require some slight modifications, if they are to be brought into line with the principle in question. But there is certainly no need to multiply such negative examples; the reader can pick out at random as many cases as he wants, and he is sure never to hit a case which does not contradict the principle of relativity. Maxwell's equations for the 'free aether,' also with the supplementary term $\rho \mathbf{p}$, and for 'stationary' ponderable media, are, as has been already remarked, in an exceptional position. But these electromagnetic equations will occupy our special attention in later chapters.

Thus far we have had only one example of a 'lay' which is proclaimed to be rigorously valid, with reference to $S$, namely the law of light propagation, as enunciated in the principle of constant light-velocity.* Thus, the true office of II. is to fix a particular case of a physical law which is postulated rigorously to satisfy I.

* Notice that, in considering this law, we need not yet trouble about the electromagnetic, or any other, theory of light.

This law then has certainly to be invariant with respect to the Lorentz transformation. And since this transformation has been obtained by means of the law itself, applied both to $S$ and $S^{\prime}$, it can be foreseen without calculation that this law will prove to be invariant. In fact, this prevision may be verified at once. For the law in question states that if light be emitted at the instant $t=0$ by a point-source, placed at or just passing through a given point, which may be taken as the origin of the coordinates, $O$, then at any instant $t>0$ it reaches a spherical surface of radius $r=c t$ and centre $O$, i.e. such that, $x, y, z$ being the coordinates of that surface,

$$
\begin{equation*}
x^{2}+y^{2}+z^{2}-c^{2} t^{2}=0 \tag{II}
\end{equation*}
$$

Now, squaring the equations (ro) and adding up, we have, identically,

$$
\begin{equation*}
x^{\prime 2}+y^{\prime 2}+z^{\prime 2}-c^{2} t^{\prime 2}=x^{2}+y^{2}+z^{2}-c^{2} t^{2} \tag{I2}
\end{equation*}
$$

and consequently also

$$
\begin{equation*}
x^{\prime 2}+y^{\prime 2}+z^{\prime 2}-c^{\prime 2} t^{\prime 2}=0 . \tag{ri'}
\end{equation*}
$$

Thus, the law of light propagation, (ri), is invariant with respect to the Lorentz transformation. Remember that $O^{\prime}$ coincides with $O$ for $t=0$, when also $t^{\prime}=0$, and that, therefore, ( $\mathrm{Ir}^{\prime}$ ) expresses for $S^{\prime}$ precisely the same thing as (ir) for $S$. Notice, moreover, that the law under consideration would be invariant with any value of $a$ (not zero). For, then, we should have, by (8),

$$
x^{\prime 2}+y^{\prime 2}+z^{2}-c^{\prime \prime 2} t^{2}=a^{2}\left(x^{\prime 2}+y^{\prime 2}+z^{\prime 2}-c^{2} t^{\prime 2}\right),
$$

and what we require is not so much the invariance of the quadratic function $x^{2}+y^{2}+z^{2}-c^{2} t^{2}$ as that of the equation (ri). But having once decided, be it only for purely formal reasons, to take $\alpha=\mathrm{r}$, the property ( r 2 ), which will in the sequel be often referred to, is worth keeping in memory.

It may be expressed shortly by saying that

$$
x^{2}+y^{2}+z^{2}-c^{2} t^{2}, \text { or } r^{2}-c^{2} t^{2}
$$

is a relativistic invariant. Any function of this expression alone is, of course, again an invariant. But all of these count as one invariant. It is worth noticing that, on this understanding, there are among all functions of $x, y, z, t$ no other invariants than

$$
x^{2}+y^{2}+z^{2}-c^{2} t^{2}
$$

In what precedes we have used the integral form, (II), of (a particular case of) the law of propagation. We might as well have used its differential form,

$$
\begin{equation*}
\square \phi=0, \tag{13}
\end{equation*}
$$

where $\phi$ may be thought of as any one of the rectangular components of a 'light-vector,' and where

$$
\begin{equation*}
\square=\nabla^{2}-\frac{\mathrm{I}}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}=\frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}+\frac{\partial^{2}}{\partial z^{2}}-\frac{\mathrm{r}}{c^{2}} \partial^{2} \partial t^{2} \tag{1+4}
\end{equation*}
$$

is Cauchy's symbol, called also the Daiembertian. The physical meaning of this famous differential equation is (among other things) that any element of a wave of discontinuity is propagated normally to itself with the velocity $c$ (cf. Note 2). This then is the gencral law of which the previous is but a particular case, corresponding to a particular form of the wave. Now, by (ro),

$$
\begin{gather*}
\frac{\partial}{\partial x}=\gamma\left(\frac{\partial}{\partial x^{\prime}}-\beta \frac{\partial}{c \partial t^{\prime}}\right) ; \quad \frac{\partial}{\partial y^{\prime}}=\frac{\partial}{\partial y^{\prime}} ; \quad \partial z=\frac{\partial}{\partial z^{\prime}} ; \\
\frac{\partial}{c \partial t}=\gamma\left(\frac{\partial}{c d t^{\prime}}-\beta \frac{\partial}{\partial x^{\prime}}\right), \tag{15}
\end{gather*}
$$

so that
which proves the invariance of the differential law of the propagation of light in empty space. But since (13) involves further particulars not yet entered upon (embodied summarily in (d) concerning light, the reader is recommended to keep rather to the above integral form (Ir), until we come to consider the relativistic propertics of electromagnetic laws. Meanwhile he is asked to retain in memory solely that the Dalembertian is an invariant as good as $r^{2} \ldots c^{*} /$, although the latter is a magnitude and the former an operator.

Conversely, the Lorentz transformation may be obtained by postulating the invariance of the Dalembertian and by making some auxiliary assumptions (Note 3). But the above method of obtaining the transformation formulae seemed to me to be more suitable for bringing into prominence their physical meaning.

Basing ourselves upon the Principles I., II., and upon the obvious requirement III., we have obtained the formulace (4a) for the ratios of time-intervals and lengths as measured in $S$ and $S^{\prime}$. From these formulae the Lorentz transformation (ro), and its inverse

[^33](ro'), followed almost immediately. Now, it may be well to notice here how (4a) are to be obtained conversely from (ro), (ró). The third of ( $4 a$ ) is identical with $y=y^{\prime}, z=z^{\prime}$. To obtain the first of formulae ( $4 a$ ), remember that it was valid for a point (any point) fixed in $S^{\prime}$. Take therefore, in the last of (ro'), $x^{\prime}=$ const., and denote by $\Delta$ any increment. Then the result will be
$$
\Delta t=\gamma \Delta t^{\prime} .
$$

Similarly, remembering that the terminals of the segment $l$ are to be taken simultaneous in $S$, take, in the first of (ro), $t=$ const. ; then the result will be

$$
\Delta x=\frac{i^{\prime}}{\gamma} \Delta x^{\prime} .
$$

Now, these are precisely the relations stated by (4a). Notice that the constancy or variability of the transversal coordinates $y, z$ is a matter of indifference. As to the fact, mentioned on several occasions, that simultaneous events occurring at distant places in $S^{\prime}$ are generally not simultaneous in $S$, and vice versa, it is most immediately expressed in (ro), (ro') by the circumstance that $t$ contains $x^{\prime}$ besides $t^{\prime}$, and similarly, that $t^{\prime}$ contains $x$ besides $t$.

So long as $\eta<c$, or $\beta<\mathrm{r}$, the coefficient $\gamma$ is real and greater than unity, so that the duration of any process, local in $S^{\prime}$, is lengthened, to the $S$-observers, $\gamma$ times or in the ratio $\mathrm{I}:\left(\mathrm{r}-\beta^{2}\right)^{\frac{1}{2}}$, and any longitudinal segment $\Delta x^{\prime}$ is contracted to $\Delta x^{\prime}\left(\mathrm{r}-\beta^{2}\right)^{\frac{1}{2}}$. In the critical case of $\eta=c$, or $\beta=\mathrm{x}$, we have $\gamma=\infty$. Then any finite duration $\Delta t^{\prime}$ becomes infinite in $S$, and any finite distance $\Delta x^{\prime}$, as judged by the $S$-observers, dwindles down to nothing: the whole of $S^{\prime}$, with all the bodies sharing its motion, becomes a transversal flatland. Finally, for $\beta>\mathrm{I}$, i.e. when the velocity of $S^{\prime}$ relative to $S$ exceeds the velocity of light or when it becomes what may conveniently be called a hypervelooity,* $\gamma$ is purely imaginary and so also are $x, t$ for any real values of $x^{\prime}, t^{\prime}$. But, as far as I can see, this does not necessarily mean that motion with hypervelocity, of one body relative to another, is 'impossible.' It would, thus far, be enough to say simply that there is in this case no correlation in real terms between $S^{\prime}$ and $S$ to be obtained by light-signalling. Notice that, from the $S$-standpoint, any station $P^{\prime}$ can then succeed in sending light-signals only to points contained in a certain back-

* The Germans call it 'Ueberlichtgeschwindigkeit.'
cone, so that, according to that standpoint, no such station can ever receive back any of its signals, and that therefore the whole of our previous reasoning ceases to be applicable to the case in question. In what sense hypervelocities are, or by what reasons they may be required to be, 'impossible,' will be seen from the physical applications of the principle of relativity.
For the present, and for what follows, we shall simply assume

$$
v<c
$$

considering only now and then the limiting value $v=c$.
To touch the other extreme, let us suppose that $v$ is a very small fraction of $c$. Then, neglecting $\beta^{2}$-terms, and limiting ourselves to such values of $x$ as are not enormously great compared with ct, we obtain from (ro) the Newtonian transformation (Chap. I.)

$$
x^{\prime}=x-v t ; \quad y^{\prime}=y ; \quad z^{\prime}=z ; \quad t^{\prime}=t .
$$

Or, if we like, we can say also that, if $\infty$ is taken instead of $c$, the Lorentz transformation reduces to the Newtonian transformation. Just as the equations of classical or Newtonian mechanics were invariant with respect to the Newtonian transformation, so are the fundamental laws of optics and (as we shall see later) of electromagnetism invariant with respect to the Lorentz transformation. Let us call the principle associated with the former the classical principle of relativity, and that corresponding to the latter of these transformations the nere principle of relativity. Then it is obvious that we cannot have both, retaining the classical principle for our mechanics and using the new one for our clectromagnetism. For if $S$ be a particular system or space-framework of reference in which the laws of both classical mechanics and electromagnetism are valid, then, among all the systems moving with respect to it with uniform velocity, no other would have this property.* In other words, the system $S$ would be privileged, being the system for both classes of laws, whereas, according to the general principle of relativity, i.e. according to I. taken by itself (without yet touching II.), none of the manifold of $\infty^{8}$ systems moving uniformly with respect to one another is to be privileged, equal rights being claimed for all of them with regard to any physical phenomena. Thus, if we are

[^34]to construct a truly relativistic theory at all, we can have but one Principle of Relativity, that is to say, one at a time. (It may well happen that the next, or even the present, generation will have to give up the 'new' principle for a yet broader one.) Now, Hertz's and Heaviside's attempts to extend the classical principle of relativity to the domain of electromagnetism proved a complete failure. And since, for the time being, tertium non datur, the 'new' principle, involving the Lorentz transformation, has become the principle of relativity of modern physics. In this comnexion it must not be forgotten that electromagnetic and especially optical phenomena have been known all these years with a much higher degree of accuracy than the various instances of motion of material bodies. No wonder, therefore, that the physicist has so easily decided to mould his mechanics and thermodynamics according to a principle which sprang out from optical and, generally, electromagnetic ground. This is not to say, of course, that mechanical and all other phenomena must be 'ultimately' electromagnetic, i.e. that everything must be explained by, or reduced to, electromagnetism. The theory of relativity is not concerned at all with such reduction of one class of phenomena to another. It does not force upon us an electromagnetic view of the world any more than a mechanical view. Quite the contrary; it opens before us a wide field of possibilities of asserting that even the mass of a free electron, say a $\beta$-particle, must not be entirely electromagnetic.

Like the Newtonian transformations, the Lorentz transformations, generally with the inclusion of pure space-rotations,* constitute a group, that is to say, two of such transformations applied successively one after the other are equivalent to a single transformation, which is again a Lorentz transformation. In the case of the Newtonian transformation, if $\nabla_{1}$ be the velocity of $S^{\prime}$ rclative to $S$, and $\nabla_{2}$ the velocity of $S^{\prime \prime}$ relative to $S^{\prime}$, the vectorial parameter $\mathbf{V}$ of the resultant transformation is simply the sum of the parameters of the component transformations, i.e. $\mathbf{v}=\mathbf{v}_{1}+\mathbf{V}_{2}$. The parameter of the resultant Lorentz group is a more complicated function of the parameters of the component transformations, thus leading to a more complicated rule of physical addition of velocities, which will be given in the next two chapters. Only when the absolute values of $\mathrm{v}_{1}, \mathrm{v}_{2}$ are small compared with the critical velocity, does the familiar rule of composition of velocities reappear. Classical kine-
*This rescrvation will become clear in Chapter VI.
matic is but a limiting case of modern relativistic kincmatic. So are also most of the remaining branches of mechanics and generally of physics. For slow motion the reader will recognise throughout his good old friends in this new and strange land of relativistic connexions.

To close this somewhat lengthy chapter on the foundations of the theory of relativity, one short remark more. Einstein's results concerning electromagnetic and optical phenomena will be seen to agree in the main with those which have been obtainecl by Lorentz in his generalized theory, the chief difference being (to quote Lorentz's own words, Columbia University Lectures, p. 230) that Einstein simply postulates what Lorentz has deduced 'with some difficulty, and not altogether satisfactorily, from the fundamental equations of the electromagnetic field. By doing so, he may certainly take credit for making us see in the nogative result of experiments like those of Michelson, Rayleigh and Brace, not a fortuitous compensation of opposing effects, but the manifestation of a general and fundamental principle. . . . It would be unjust not to add that, besides the fascinating boldness of its starting point, Einstein has another marked advantage over minc. Whereas I have not been able to obtain for the equations referred to moving axes exactly the same form as for those which apply to a stationary system, Einstein has accomplished this by means of a system of new variables slightly different from those which I have introduced.' (As to these slight differences, cf. Note 86 to Lorentz's Lectures.)

We see from the above quotation that Lorentz himself aimed at an exact sameness of form of the laws of all, or at least of elestrom magnetic, phenomena for a pair of systems moving uniformly with respect to one another. Why then not postulate this sameness at once? But Lorentz had not the heart to abandon the atether which he confessedly 'cannot but regard as endowed with a certain degree of substantiality.'

## NOTES TO CHAPTER IV.

Note 1 (to page Iro). Einstein, Ann2. d. Physik, Vol. XVII., 1905, §3, obtains the formulae of transformation (10) in the following way:

Let, to use our notation, $x, y, z, t$ be the coordinates and the time in $S$, and $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$ those in $S^{\prime}$. Write

$$
\xi=x-v t ;
$$

then to a point fixed in $S^{\prime}$ corresponds a system of values of $\xi, y, z$ independent of $t$. To obtain $t^{\prime}$ as a function of $\xi, y, z$, Einstein considers a signal sent at the instant $t_{0}^{\prime}$ from the origin $O^{\prime}$ along the axis of $x^{\prime}$ towards the point $\xi$, where it arrives at the instant $t_{1}$, and, being reflected there, returns to $O^{\prime}$ at the instant $t_{2}^{\prime}$. Then, according to the definition of synchronism, (Ia), p. 95, which is to hold equally for $S^{\prime}$ as for $S$,

$$
t_{1}^{\prime}=\frac{1}{2}\left(t_{0}^{\prime}+t_{2}^{\prime}\right)
$$

i.e. filling in the arguments and applying the principle of constant light propagation,

$$
t^{\prime}(0,0,0, t)+t^{\prime}\left(0,0,0,\left[t+\frac{\xi}{c-v}+\frac{\xi}{c+v}\right]\right)=2 t^{\prime}\left(\xi, 0,0, t+\frac{\xi}{c-v}\right)
$$

whence, for an infinitesimal $\xi$,

$$
\frac{\partial t^{\prime}}{\partial \xi}+\frac{v}{c^{2}-v^{2}} \frac{\partial t^{\prime}}{\partial t}=0
$$

Applying the same reasoning to signals sent along the axes of $y$ or $z$, Einstein obtains

$$
\frac{\partial t^{\prime}}{\partial y}=0, \quad \frac{\partial t^{\prime}}{\partial z}=0,
$$

and, assuming $t^{\prime}$ to be a linear function of its arguments,

$$
t^{\prime}=\phi(v) \cdot\left(t-\frac{v \xi}{c^{2}-v^{2}}\right),
$$

where $\phi(v)$ is thus far an unknown function of $v$, and where $t^{\prime}=0$ has been put at $O^{\prime}$ for $t=0$.

Next, to obtain from the last equation $x^{\prime}, y^{\prime}, z^{\prime}$ in terms of $x, y, z, t$, Einstein writes the principle of constant light-propagation in $S^{\prime}$. A signal started at $O^{\prime}$ at the instant $t^{\prime}=0$ reaches at the instant $t^{\prime}$ a point of the positive $x^{\prime}$-axis, for which

$$
x^{\prime}=c t^{\prime}=\phi(v) \cdot c\left(t-\frac{v \xi}{c^{2}-v^{2}}\right) .
$$

But the same process, if considered from the $S$-standpoint, gives $\xi=t(c-v)$. Thus

$$
x^{\prime}=\phi(v) \frac{c^{3}}{c^{2}-v^{2}} \xi=\phi(v) \gamma^{2} \xi .
$$

Similarly

$$
y^{\prime}=c t^{\prime}=\phi(v) \cdot c\left(t-\frac{v}{c^{2}-v^{2}} \xi\right)
$$

where $t=y\left(c^{2}-v^{2}\right)^{-\frac{t}{2}}, \xi=0$. Thus

$$
\begin{aligned}
& y^{\prime}=\phi(v) \gamma y \\
& z^{\prime}=\phi(v) \gamma z .
\end{aligned}
$$

Consequently, writing again $\xi=x-v t$, and throwing the common factor $\gamma$ upon $\phi(\nu)$,

$$
\begin{gathered}
x^{\prime}=\phi(v) \cdot \gamma(x-v t), \quad y^{\prime}=\phi(v) \cdot y, \quad z^{\prime}=\phi(v) \cdot z \\
t^{\prime}=\phi(v \prime) \cdot \gamma\left(t-\frac{v}{c^{2}} x\right)
\end{gathered}
$$

These are identical with the formulae (8) of the present chapter, for $\phi(v)=1$. The way that Einstein obtains the particular value $\phi(v)=1$ (loc. cit. pp. 90r-902) need not detain us here. We know that the value of such a common coefficient is essentially, from the physical standpoint, a matter of indifference.

As to Laue (Das Relatizitäatsprinzip, 2nd edition, p. 38, etc.), his method of obtaining the Lorentz transformation consists in postulating the invariance of the 'wave-equation'

$$
\nabla^{2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}=0
$$

and in assuming linearity and symmetry round the axis of motion, i.e. in writing

$$
\left.\begin{array}{c}
x^{\prime}=\kappa(v) \cdot(x-v t), \quad y^{\prime}=\lambda(v) \cdot y, \quad z^{\prime}=\lambda(v) \cdot z  \tag{a}\\
t^{\prime}=\mu(v) \cdot t-v(v) \cdot x,
\end{array}\right\}
$$

where $\kappa, \lambda, \mu, \nu$ are functions of $\nu$ alone. These functions are then easily determined from the postulated invariance which Laue writes

$$
\begin{equation*}
\nabla^{\prime 2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{\prime 2}}=a\left\{\nabla^{2}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}\right\} \tag{b}
\end{equation*}
$$

where $\alpha$ is again an unknown function of $v$ alone. The value of $\lambda$ is easily shown to be equal to unity, by requiring reciprocity, i.e.

$$
y=\lambda(-v) \cdot y^{\prime}, \quad z=\lambda(-v) \cdot z^{\prime}
$$

and by remembering that 'for the $y$ - and $z$-directions it is exactly the same thing whether $S^{\prime}$ moves relatively to $S$ in the positive or in the negative sense of the $x$-axis,' so that $\lambda(v)=\lambda(-v)$. Thus $y^{\prime}=y, z^{\prime}=z$, and, by (b), $\kappa=\mu=\left(\mathrm{I}-\frac{z^{2}}{c^{2}}\right)^{-\frac{7}{2}}=\gamma, v=\frac{v}{c^{2}} \gamma$. Substituting these values in (a), Laue obtains the required formulae (ro). The discussion of Laue's method of obtaining for $\alpha$ the particular value I , rather than any other, is again left to the reacler.

Note 2 (to page II3). Let the function $\phi$, satisfying the equation $\square \phi=0$, be continuous, as well as its first derivatives $\partial \phi / \partial t, \partial \phi / \partial x$, etc., that is to say, let

$$
[\nabla \phi]=0, \quad\left[\frac{\partial \phi}{\partial t}\right]=0
$$

but let the clcrivatives of the second order; $\partial^{2} \phi / \partial t^{2}, \partial^{2} \phi / \partial x^{2}$, etc., experience a discontinuity at the surface $\sigma$. Then, $\mathrm{n}=\mathrm{i} n_{1}+\mathrm{j} n_{2}+\mathrm{k} n_{3}$ being the normal of any surface-element $d \sigma$, at the instant $t$, the identical conditions and the kinematic conclitions of compatibility, expressing that $\sigma$ is neither split into two or more surfaces, nor dissolved, at the next instant $t+d t$, are (cf. Ann. der Physik, Vol. XXIX., 1909, p. 524)

$$
\begin{gathered}
{\left[\begin{array}{l}
\partial^{2} \phi \\
\partial x^{2}
\end{array}\right]=n_{1}^{2} \lambda,}
\end{gathered} \begin{aligned}
& {\left[\frac{\partial^{2} \phi}{\partial y^{2}}\right]=n_{2}^{2} \lambda, \quad\left[\frac{\partial^{2} \phi}{\partial z^{2}}\right]=n_{3}^{2} \lambda} \\
& {\left[\frac{\partial^{2} \phi}{\partial t^{2}}\right]=v^{2} \lambda}
\end{aligned}
$$

where $\mathfrak{b}$ is the velocity of propagation, along $n$, and $\lambda$ a scalar characterizing the discontinuity. Now, $n$ being a unit vector,

$$
\left[\nabla^{2} \phi\right]=\lambda, \quad \text { and } \quad[\square \phi]=\lambda\left(\mathrm{r}-\frac{\mathfrak{b}^{2}}{c^{2}}\right)=0,
$$

whence $|\mathfrak{b}|=c$. Q.E.D.
In electromagnetism $\phi$ has in turn the meaning of the components of the electrical and the magnetic vector, and the sense of propagation, $\pm \mathrm{n}$, follows from the mutual relations of these two vectors.

Note 3 (to page 113). Postulate the invariance of the Dalembertian, i.e.
and assume

$$
\begin{gathered}
\square^{\prime}=\square \\
y^{\prime}=y, \quad z^{\prime}=z
\end{gathered}
$$

or make any set of plausible assumptions leading to this. Then

$$
\partial^{2} / \partial y^{\prime 2}=\partial^{2} / \partial y^{2}, \quad \partial^{2} / \partial z^{\prime 2}=\partial^{2} / \partial z^{2}
$$

and

$$
\frac{\partial^{2}}{\partial x^{\prime 2}}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{\prime 2}}=\frac{\partial^{2}}{\partial x^{2}}-\frac{1}{c^{3}} \frac{\partial^{2}}{\partial t^{2}} .
$$

Instead of $x, t$ introduce new independent variables

$$
\begin{aligned}
& \xi=x-c t, \\
& \eta=x+c t,
\end{aligned}
$$

and similarly, for the system $S^{\prime}$,

$$
\begin{aligned}
& \xi^{\prime}=x^{\prime}-c t^{\prime}, \\
& \eta^{\prime}=x^{\prime}+c t^{\prime} .
\end{aligned}
$$

Then the required invariance will assume the form

$$
\begin{equation*}
\frac{\partial^{2}}{\partial \xi^{\prime} \cdot \partial \eta^{\prime}}=\frac{\partial^{2}}{\partial \xi \cdot \partial \eta} \tag{a}
\end{equation*}
$$

Now, considering $\xi^{\prime}, \eta^{\prime}$ as functions of $\xi, \eta$, zvithout assuming their linearity, we have

$$
\frac{\partial}{\partial \xi}=\frac{\partial \xi^{\prime}}{\partial \xi} \cdot \frac{\hat{c}^{\prime}}{\partial \xi^{\prime}}+\frac{\partial \eta^{\prime}}{\partial \xi^{\prime}} \cdot \frac{\partial}{\partial \eta^{\prime}}
$$

and

$$
\frac{\partial^{2}}{\partial \xi \partial \eta}=\frac{\partial \xi^{\prime}}{\partial \xi} \frac{\partial \xi^{\prime}}{\partial \eta} \cdot \frac{\partial^{2}}{\partial \xi^{\prime 2}}+\frac{\partial \eta^{\prime}}{\partial \xi} \frac{\partial \eta^{\prime}}{\partial \eta} \cdot \frac{\partial^{2}}{\partial \eta^{\prime 2}}+\left(\frac{\partial \xi^{\prime}}{\partial \eta} \frac{\partial \eta^{\prime}}{\partial \xi^{\prime}}+\frac{\partial \xi^{\prime}}{\partial \xi^{\prime}} \partial \eta^{\prime}\right) \cdot \frac{\partial^{2}}{\partial \xi^{\prime} \partial \eta^{\prime}}
$$

$$
+\frac{\partial^{2} \xi^{\prime}}{\partial \xi^{\prime} \partial \eta} \cdot \frac{\partial}{\partial \xi^{\prime}}+\frac{\partial^{2} \eta^{\prime}}{\partial \xi \partial \eta} \cdot \frac{\partial}{\partial \eta^{\prime}}
$$

Thus, by (a),

$$
\begin{gathered}
\frac{\partial \xi^{\prime}}{\partial \xi^{\prime} \partial \eta}=0, \quad \frac{\partial \eta^{\prime} \eta^{\prime}}{\partial \xi^{\prime} \partial \eta}=0, \\
\frac{\partial \xi^{\prime}}{\partial \xi^{\prime}} \frac{\partial \xi^{\prime}}{\partial \eta}=0, \quad \frac{\partial \eta^{\prime} \partial \eta^{\prime}}{\partial \xi^{\prime} \partial \eta}=0, \\
\partial \xi^{\prime} \cdot \frac{\partial \eta}{\partial \xi^{\prime}}+\frac{\partial \xi^{\prime}}{\partial \xi^{\prime} \eta_{\eta}^{\prime}}=\mathrm{I} .
\end{gathered}
$$

To satisfy the third of these conditions, put
then the fifth will become

$$
\frac{\partial \xi^{\prime}}{\partial \eta}=0
$$

$$
\frac{\partial \xi^{\prime}}{\partial \xi} \frac{\partial \eta^{\prime}}{\partial \eta}=\mathrm{I}
$$

so that the only possibility of fulfilling the fourth condition consists in taking

Thus,

$$
\frac{\partial \eta^{\prime}}{\partial \xi}=0
$$

$$
\xi^{\prime}=\xi^{\prime}(\xi), \quad \eta^{\prime}=\eta^{\prime}(\eta)
$$

Hereby the first and second of the above conditions are identically satisfied, and the fifth becomes

$$
\begin{equation*}
\frac{d \xi^{\prime}}{d \xi} \cdot \frac{d \eta^{\prime}}{d \eta}=\mathrm{I} \tag{b}
\end{equation*}
$$

[An alternative solution would be $\partial \xi^{\prime} \partial \xi=0$, and $\partial \eta^{\prime} \partial \eta=0$, i.e. $\xi^{\prime}=\xi^{\prime}(\eta)$, $\eta^{\prime}=\eta^{\prime}(\xi)$, with $\left(d \xi^{\prime} / d \eta\right) .\left(d \eta^{\prime} / d \xi\right)=1$; but this may easily be shown to lead substantially to the same final result as the above one.] Now, for

$$
x=v t=c \beta t,
$$

we require $x^{\prime}=0$, i.e.

$$
\xi^{\prime}[\operatorname{ct}(\beta-\mathrm{I})]+\eta^{\prime}[c t(\beta+\mathrm{I})]=0,
$$

for every $t$; hence, differentiating with respect to $t$, and supposing $v$ constant,

$$
(\mathrm{r}-\beta)\left(\frac{d \xi^{\prime}}{d \xi}\right)=(\mathrm{r}+\beta)\left(\frac{d \eta^{\prime}}{d \eta}\right)
$$

and, by (b),

$$
\left(\frac{d \xi^{\prime}}{d \xi}\right)=\sqrt{\frac{I}{I}+\beta} \begin{aligned}
& \mathrm{I}-\beta^{\prime}
\end{aligned} \quad\left(\frac{d \eta^{\prime}}{d \eta}\right)=\sqrt{\frac{\mathrm{I}-\beta}{\mathrm{I}+\beta^{\prime}}}
$$

where both square roots are to be taken with the same sign, namely the positive (since $\xi^{\prime}=\xi$, etc., for $\beta=0$ ). Here (), in the differential coefficients, means ' for $x=v t$ '; but since $\xi^{\prime}, \eta$ ' depend only on $\xi, \eta$ respectively, these formulae are valid for any arguments. Hence, integrating, and remembering that for $x=t=0$, i.c. for $\xi=\eta=0$, we require $\xi^{\prime}=\eta^{\prime}=0$,

$$
\begin{equation*}
\xi^{\prime}=\sqrt{\frac{I+}{I-\beta}} \xi ; \quad \eta^{\prime}=\sqrt{\frac{I}{I}-\beta} \beta^{I} \eta \tag{c}
\end{equation*}
$$

This intermediate form is worth notice, since it shows at once that
i.e. $x^{\prime 2}-c^{2} t^{\prime 2}=x^{2}-c^{2} t^{2}$.

$$
\xi^{\prime} \eta^{\prime}=\xi \eta
$$

Introducing again the values of $\xi$, etc., in terms of $x$, etc., (c) are readily seen to be identical with the required formulae

$$
x^{\prime}=\gamma(x-v t), \quad t^{\prime}=\gamma\left(t-\frac{v}{c^{2}} x\right)
$$

## CHAPTER V. <br> VARIOUS REPRESENTATIONS OF THE LORENTZ TRANSFORMATION.

Passing now to consider the various expressions of the Lorentz transformation, which was seen to be fundamental for the whole theory of Relativity, let us first of all deprive the $x$-axis of its (formal) privilege and write (io), Chap. IV., symmetrically in $x, y$, $z$, or, using vectors, avoid splitting into Cartesians altogether. This is done in a moment. In fact, remembering that our axis of $x$ was longitudinal, and those of $y, z$ transversal, and calling r the vector drawn from $O$ to any point in $S$, and $r^{\prime}$ its $S^{\prime}$-correspondent, we can write the first of (IO),

$$
\left(\mathrm{r}^{\prime} \mathrm{i}\right)=\gamma[(\mathrm{ri})-, v t]
$$

where $\mathbf{i}$ is the unit of $\mathbf{v}$, similarly the second and third,

$$
\mathbf{r}^{\prime}-\left(\mathbf{r}^{\prime} \mathbf{i}\right) \mathbf{i}=\mathbf{r}-(\mathrm{r}) \mathrm{i},
$$

and, finally, the last of (10),

$$
t^{\prime}=\gamma\left[t-\underset{c^{2}}{v}(\mathrm{ri})\right]=\gamma\left[t-{ }_{c^{2}}^{\mathrm{r}}(\mathrm{r} \mathrm{v})\right] .
$$

To obtain the full vector $\mathbf{r}^{\prime}$ combine its transversal and longitudinal parts, and to get rid of the new letter $i$, write $(\mathrm{ri}) \mathrm{ima}(\mathbf{r v}) \mathbf{v} / v^{2}$. Thus, the concise vectorial form of the Lorentz transformation, exhibiting its independence of the choice of coordinate axes, will be

$$
\left.\begin{array}{c}
\mathbf{r}^{\prime}=\mathbf{r}+\left[\frac{\gamma-\mathrm{I}}{v^{2}}(\boldsymbol{\nabla} \mathbf{r})-\gamma t\right] \mathbf{\nabla}  \tag{I}\\
t^{\prime}=\gamma\left[t-\frac{1}{c^{2}}(\boldsymbol{\nabla})\right] .
\end{array}\right\}
$$

Here $\nabla$ is the velocity of $S^{\prime}$ relative to $S$, and $\gamma=\left(\mathrm{I}-\beta^{2}\right)^{-\frac{1}{2}}, \beta=v / c$, as before.

To suit the non-vectorial reader we may again split ( r ) into Cartesians. But in doing so, let us this time take any set of mutually perpendicular axes $x, y, z$, for $S$, which are also to be the axes of $x^{\prime}, y^{\prime}, z^{\prime}, x^{\prime \prime}, y^{\prime \prime}, z^{\prime \prime}$, etc., for all other systems $S^{\prime}, S^{\prime \prime}$, etc., moving uniformly with respect to one another. Call $v_{x}, v_{y}, v_{z}$ the components of $\nabla$ taken along these universal, but quite arbitrary, axes. Then, projecting the first of ( I ) upon these axes and rewriting the second of ( I ), the required symmetrical form will follow, viz.

$$
\left.\begin{array}{c}
x^{\prime}=x+\left[\frac{\gamma-1}{v^{2}}(\mathrm{rv})-\gamma t\right] v_{x} \\
y^{\prime}=y+\left[\frac{\gamma-\mathrm{x}}{v^{2}}(\mathrm{rv})-\gamma t\right] v_{y} \\
z^{\prime}=z+\left[\frac{\gamma-1}{v^{2}}(\mathrm{rv})-\gamma t\right] v_{z}  \tag{a}\\
t^{\prime}=\gamma\left[t-\frac{\mathrm{x}}{c^{2}}(\mathrm{rv})\right],
\end{array}\right\}
$$

where (rv) may be looked at as an abbreviation for $x v_{x}+y v_{y}+z v_{z}$. The inverse transformation is obtained by transferring the dashes from $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$ to $x, y^{\prime}, z, t$, and by changing the sign of $\mathrm{\nabla}$, that is of $v_{x}, v_{y}, v_{z}$.

On the other hand, to condense the vectorial form (I) still a little more, observe that $r$ enters into the first of ( r ) by the expression $\mathbf{r}+\frac{\gamma-\mathrm{r}}{\nu^{2}} \nabla(\mathbf{v r})$ only. Introduce therefore the linear vector operator

$$
\begin{equation*}
\epsilon=\mathrm{r}+\frac{\gamma-\mathrm{I}}{v^{2}} \nabla(\nabla . \tag{z}
\end{equation*}
$$

Then the Lorentz transformation will be expressed by

$$
\left.\begin{array}{l}
\mathbf{r}^{\prime}=\epsilon \mathbf{r}-\mathbf{v} \gamma t  \tag{xb}\\
t^{\prime}=\gamma\left[t-\frac{\mathrm{x}}{c^{2}}(\mathbf{r v})\right]
\end{array}\right\}
$$

Write again, for a moment, $\mathbf{v} / v=\dot{i}$, and let $\mathbf{j}, \mathbf{k}$ be a pair of unit vectors normal to one another and to $\nabla$. Then ( 2 ) may be written $\epsilon=\gamma i(i+r-i(i$, or, $r$ being the 'idemfactor,' i.e. $i(i+j(j+k(k$,

$$
\epsilon=\gamma i(i+j(j+\mathbf{k}(\mathbf{k} .
$$

This is called a dyadic.* Considered as an operator it is a symmetrical linear vector operator, so that if $\mathbf{A}, \mathbf{B}$ be any pair of vectors

$$
\begin{equation*}
(\mathbf{A} . \epsilon \mathbf{B})=(\mathbf{B} . \epsilon \mathbf{A}) \tag{3}
\end{equation*}
$$

But the operator $\epsilon$ may be described most immediately by calling it a longitudinal stretcher, since it stretches or magnifies $\gamma$ times any longitudinal vector, i.e. any vector parallel to $\nabla$, and leaves unchanged any transversal vector. According to the usual terminology, $\gamma$ would be the ratio of this stretcher.

Observe that $\nabla$ enters into $\epsilon$ through $\gamma$ only, i.e. quadratically. Thus, the inverse transformation will be

$$
\left.\begin{array}{l}
\mathbf{r}=\epsilon \mathbf{r}^{\prime}+\nabla \gamma t^{\prime} \\
t=\gamma\left[t^{\prime}+\frac{\mathrm{I}}{c}\left(\nabla \mathbf{r}^{\prime}\right)\right]
\end{array}\right\}
$$

The above form of the Lorentz transformation, involving (one vectorial parameter $\mathbf{v}$ or) three scalar parameters $v_{x}, v_{y}, v_{z}$, is especially useful when there are more than two systems, $S, S^{\prime}, S^{\prime \prime}$, to be considered, and when the velocity of $S^{\prime \prime}$ relative to $S^{\prime}$ is not parallel to that of $S^{\prime}$ relative to $S$.

But before proceeding further let us yet dwell a little more upon the properties of the sub-group contained in ( $\mathrm{r} b$ ), which involves one scalar parameter only, and which covers the particular case of parallel velocities. This case is especially interesting and instructive as illustrating a fundamental theorem of Lie's theory of groups of transformations $\dagger$ and as preparing the way for a subsequent form of the Lorentz transformation, adopted for illustrative purposes by Minkowski.

Measuring $x$, and $x^{\prime}$, along the direction of motion of $S^{\prime}$ relative to $S$, write again, as in the last chapter,

$$
x^{\prime}=\gamma(x-v t), \quad t^{\prime}=\gamma\left(t-\frac{v}{c^{2}} x\right), \quad y^{\prime}=y, \quad z^{\prime}=z, \ddagger
$$

[^35]and differentiate $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$ with respect to the parameter $v$. Then, denoting $d \gamma / d v$ by $\dot{\gamma}$,
$$
\frac{d x^{\prime}}{d v}=\frac{\dot{\gamma}}{\gamma} x^{\prime}-\gamma t ; \quad \frac{d t^{\prime}}{d v}=\dot{\gamma}\left(t-\frac{v}{c^{2}} x\right)-\frac{\gamma}{c^{2}} x
$$
and using the inverse transformation $x=\gamma\left(x^{\prime}+v t^{\prime}\right)$, etc.,
\[

$$
\begin{gather*}
\frac{d x^{\prime}}{d v}=\left(\frac{\dot{\gamma}}{\gamma}-\frac{\gamma^{2} v}{c^{2}}\right) x^{\prime}-\gamma^{2} t^{\prime} ; \quad \frac{d y^{\prime}}{d v}=0 ; \quad \frac{d z^{\prime}}{d v}=0,  \tag{4}\\
\frac{d t^{\prime}}{d v}=\left(\frac{\dot{\gamma}}{\gamma}-\frac{\gamma^{2} v}{c^{2}}\right) t^{\prime}-\frac{\gamma^{2}}{c^{2}} x^{\prime} . \tag{5}
\end{gather*}
$$
\]

To see that this is precisely the form corresponding to Lie's theorem, which, writing $a$ instead of $v$, and $x_{i}^{\prime}(i=1,2,3,4)$ for $x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}$, would be

$$
\begin{equation*}
\frac{d x_{i}^{\prime}}{d a}=\psi_{i}(a) \cdot \xi_{i}\left(x_{1}^{\prime}, x_{2}^{\prime}, x_{8}^{\prime}, x_{4}{ }^{\prime}\right), \tag{6}
\end{equation*}
$$

we have to remember only that $\gamma^{2}=\left(x-\beta^{2}\right)^{-1}, \beta=v / c$, so that

$$
\dot{\gamma}=\frac{x}{c} \beta \gamma^{3},
$$

and consequently

$$
\dot{\gamma} / \gamma-\gamma^{2} v / c^{2}=0
$$

relative to $S$, and $v_{2}$ that of $S^{\prime \prime}$ relative to $S^{\prime}\left(v_{2}\right.$ being taken from the $S^{\prime}$-point of view and $v_{1}$ from the $S$-standpoint), then we have

$$
x^{\prime}=\gamma_{1}\left(x-v_{1} t\right), \quad t^{\prime}=\gamma_{1}\left(t-\frac{v_{1}}{c^{2}} x\right), \quad y^{\prime}=y^{\prime}, \quad z^{\prime}=z
$$

and

$$
x^{\prime \prime}=\gamma_{2}\left(x^{\prime}-z_{2} z^{\prime}\right), \quad t^{\prime \prime}=\gamma_{2}\left(t^{\prime}-\frac{\gamma_{2}}{c^{2}} x^{\prime}\right), \quad y^{\prime \prime}=y^{\prime}, \cdot z^{\prime \prime}=z^{\prime}
$$

and substituting the first in the second, we obtain at once

$$
x^{\prime \prime}=\gamma(x-v t), \quad t^{\prime \prime}=\gamma\left(t-\frac{v}{c^{2}} x\right), \quad y^{\prime \prime}=y, \quad s^{\prime \prime}-z,
$$

which is again a Lorentz transformation like each of the above ones, namely with the parameter (velocity of $S^{\prime \prime}$ relative to $S$ )

$$
v=\frac{v_{1}+v_{2}}{\mathrm{I}+\frac{v_{1} v_{2}}{c^{2}}}
$$

This formula embodies the simplest case of Einstein's 'addition-theorem' of velocities, which will occupy our attention in the next chapter.
identically. Thus, the differential equations (4), (5), with the omission of the obvious $d y^{\prime} / d y=d z^{\prime} / d v=0$, become at once

$$
\begin{equation*}
\frac{d x^{\prime}}{d v}=-\gamma^{2} t^{\prime} ; \quad \frac{d t^{\prime}}{d v}=-\frac{\gamma^{2}}{c} x^{\prime}, \tag{7}
\end{equation*}
$$

or, writing

$$
\begin{equation*}
l^{\prime}=l c t^{\prime}, \quad \text { and similarly } \quad l=t c t, \tag{8}
\end{equation*}
$$

where $\iota=\sqrt{-x}$,

$$
\left.\begin{array}{l}
\frac{d x^{\prime}}{d v}=\iota \frac{\gamma^{2}}{c} l^{\prime}  \tag{9}\\
\frac{d l^{\prime}}{d v}=-\iota \frac{\gamma^{2}}{c} x^{\prime} .
\end{array}\right\}
$$

Here, the coefficient on the right side being in both equations the same known function of $v$, the idea easily suggests itself to introduce instead of $v$ the new parameter

$$
\omega=\frac{\iota}{c} \int_{0}^{v} \gamma^{3} d v=\iota \int_{0}^{\beta} \frac{d \beta}{\mathrm{I}-\beta^{2}},
$$

i.e.

$$
\begin{equation*}
\omega=\arctan (\iota \beta) . \tag{10}
\end{equation*}
$$

With this new variable the above equations become

$$
\begin{equation*}
\frac{d x^{\prime}}{d \omega}=l^{\prime} ; \quad \frac{d l^{\prime}}{d \omega}=-x^{\prime} . \tag{9a}
\end{equation*}
$$

Using the well-known general integral of these simple equations and remembering that for $\beta=0$ (i.e. for $\omega=0$ ) $x^{\prime}=x, l^{\prime}=l$, we obtain the remarkable expression of the Lorentz transformation:

$$
\left.\begin{array}{rl}
x^{\prime} & =x \cos \omega+l \sin \omega ; \quad y^{\prime}=y ; \quad z^{\prime}=z  \tag{II}\\
l^{\prime} & =l \cos \omega-x \sin \omega,
\end{array}\right\}
$$

which was first given by Minkowski, who made it his starting point.*
Thus, the Lorentz transformation may be described as a rotation, in the four-dimensional space $x, y, z, l$, through/a an imaginary angle $\omega$ in the plane $x, l$, or 'round the plane' $y, z$.
*H. Minkowski, 'Die Grundgleichungen fiur die clektromagnetischen Vorgange in bewegten Körpern,' Göttinger Nachrichten, 1907 ; reprinted in 'Zwei AbhandIungen uber die Grundgleichungen der Elektrodynamik,' Teulner, Leiprig, 19ro, p. 10.

That the transformation in question is a pure rotation, i.e. without change of 'length,' $\left(x^{2}+y^{2}+z^{2}+l^{2}\right)^{\frac{1}{3}}$, is best seen from ( $9 a$ ), which give at once

$$
\frac{d}{d \omega}\left(x^{\prime 2}+l^{\prime 2}\right)=0
$$

showing thus the invariance, already noticed, of $x^{2}+l^{2}$, and consequently also of $x^{2}+y^{2}+z^{2}+l^{2}$. Notice that the above rotation $\omega$ is an imaginary Euclidean rotation in $x, y, z, l$, or, which is the same thing, a real non-Euclidean (Lobatchewskyan) rotation in the space $x, y, z$, ct through an angle $\psi$ connected with $\omega$ by

$$
\begin{equation*}
\tan \omega=\iota \tan \psi \tag{12}
\end{equation*}
$$

We shall soon have an opportunity to return to this real angle, which, according to (ro), is defined by

$$
\begin{equation*}
\tan \psi=\beta \tag{13}
\end{equation*}
$$

Let again $\mathbf{v}_{1}$ be the velocity of $S^{\prime}$ relative to $S$, and $\mathbf{v}_{2}$ that of $S^{\prime \prime}$ relative to $S^{\prime}$, the former from the $S$ - and the latter from the $S^{\prime}$-point of view. Then, if $\mathbf{v}_{1}$ and $\mathbf{v}_{2}$ be parallel to and, say, concurrent with one another, the corresponding rotations are

$$
\omega_{1}=\arctan \left(\iota \beta_{1}\right)
$$

round a certain plane, in the four-climensional space $x, y, z, l$, and

$$
\omega_{2}=\arctan \left(\iota \beta_{2}\right)
$$

round the same plane. (In three dimensions the rotation is round an axis, or line, in four 'round a plane,' i.e. leaving fixed a whole plane instead of a line.) Thus, the resultant rotation, corresponding to the passage from the $S$ - to the $S^{\prime \prime}$-variables, will be

$$
\begin{equation*}
\omega=\omega_{1}+\omega_{2} \tag{14}
\end{equation*}
$$

Not the velocities themselves are added but the corresponding angles of rotation.

To verify the last formula, call $v=c \beta$ the resultant velocity, corresponding to $\omega$. Then

$$
\iota=\tan \omega=\tan \left(\omega_{1}+\omega_{2}\right)=\iota \frac{\beta_{1}+\beta_{2}}{\mathrm{I}+\beta_{1} \beta_{2}},
$$

or

$$
v=\frac{v_{1}+v_{2}}{1+\frac{v_{1} v_{2}}{c^{2}}}
$$

Now, this is but a particular case (cf. footnote on pp. ${ }^{125-6}$ ) of Einstein's general formula for the composition of velocities, to be fully considered later on.

Since the sub-group under consideration contains the identicial transformation, namely for $v=0$ or $\omega=0$, it must be possible, according to Lie's Theorem 6 (loc. cit. p. 49), to represent it as a 'group of translations,' i.e. by

$$
\phi_{1}^{\prime}=\phi_{1} ; \quad \phi_{2}{ }^{\prime}=\phi_{2} ; \quad \phi_{3}^{\prime}=\phi_{3} ; \quad \phi_{4}^{\prime}=\phi_{1}-\omega .
$$

In fact, by ( $9 a$ ) we have the simultaneous system

$$
d x^{\prime} l^{\prime}=-\frac{d l^{\prime}}{x^{\prime}}=d \omega ; \quad d y^{\prime}=d z^{\prime}=0
$$

with the initial conditions $x^{\prime}=x, y^{\prime}=y, z^{\prime}=z, l^{\prime}=l$, for $\omega=0$. Whence
and

$$
x^{\prime 2}+l^{\prime 2}=x^{2}+l^{2}=\phi_{1}^{2}, \text { say, }
$$

$$
\frac{d l^{\prime}}{\sqrt{\phi_{1}^{2}}-l^{\prime \prime 2}}=-d^{\prime} \omega .
$$

Thus, we have only to write

$$
\begin{equation*}
\phi_{1}=\left(x^{2}+l^{2}\right)^{\frac{1}{2}} ; \quad \phi_{2}=y ; \quad \phi_{3}=z ; \quad \phi_{4}=\arcsin \frac{l}{\sqrt{x} x^{2}+l^{2}} \tag{15}
\end{equation*}
$$

and the Lorentz transformation will assume the recpuired canonic form

$$
\begin{equation*}
\phi_{t}^{\prime}=\phi_{l}(t=1,2,3) ; \quad \phi_{4}^{\prime}=\phi_{4}-\omega . \tag{16}
\end{equation*}
$$

The interpretation of this simple result, and especially that of the meaning of $\phi_{4}$, is left to the reader.

We shall now pass to a remarkable and instructive graphic representation of the Lorentz transformation, due to Minkowski.*

Minkowski calls a space-point at an instant of time, i.e. the whole tetrad of values $x, y, z, t$, a world-point (Weltpunkt), and the fourdimensional manifold of all possible systems of values $x, j^{\prime}, z, t$ the world (die Welt). Thus, a point of the world represents a material, or, in Minkowski's terminology, a 'substantial' particle at a certain instant. Suppose that the particle can be recognized and watched
*H. Minkowski, 'Raum und Zeit,' lecture delivered during the meeting of the 'Naturforscherversammlung' at Cologne, rgo8, Thysik. Zeitschrift, Vol. X., p. Io4, 1909, reprinted, with a preface by A. Gutzurer, by B. G. Teulmer,
Leipzig and Berlin, 19og. s.R.
during its whole history. Then a one-dimensional continuum, contained in the four-dimensional world, may be constructed, whose element has the components

$$
d x, d y, d z, d t
$$

along the space- and time-axes, and which represents the history of the particle. This line, whose points may be uniquely referred to the parameter $t$, say, from $-\infty$ to $+\infty$, is called a world-ine (Weltlinie). Thus the whole world would consist of a maze of such world-lines, and the physical laws would find 'their most perfect expression in the mutual relations obtaining between these worldlines.' This, of course, can be only an ideal task, and in putting it before the eyes of physicists and mathematicians, Minkowski, no doubt, was very well aware how far we are from its accomplishment.

If instead of a particle or substantial point we have a body of finite space-extension, then drawing through each of its points a world-line, we shall obtain a tubular portion of the four-dimensional world, which may be called a world-tube. In his previous paper, of 1907,* Minkowski calls it a space-time flament. The utility of the conception of a space-time filament or tube in mechanical problems and those concerned with the motion of electrons is obvious.

The world-line of a particle will in general be curvilinear, e.g. for any non-uniform motion, whether the particle's path or orbit in ordinary space be curvilinear or its velocity be changing in absolute value. But if the particle is moving uniformly, with respect to a given system $S(x, y, z, t)$, then its world-line will be a straight line, which means only that the corresponding equations obtaining between the four variables will be linear. In particular, if the particle is at rest in $S$, then its world-line will coincide with the $t$-axis, this axis, as also the axes of $x, y, z$, being considered as straight lines in the four-dimensional world.

The complete representation cannot of course be given, either by a plane drawing or by a three-dimensional model. $\dagger$ But this is no serious objection against Minkowski's method. For, first of all, it

## * Grundgleichungen für die elektromagnetischen Vorgänge, p. 47.

$\dagger$ For a remarkable attempt to ohtain a geometrical image of Minkowski's world by means of systems of spheres see a paper by IF. E. Timerding in Jahresbericht der deutschen Math. Vereinigung, Vol. XXI. 1913, p. 274.
is very advantageous, especially for the trained geometer of our days, even merely to think and to speak about these relations in terms of four-dimensional geometry. And then we can help ourselves by taking various sections of the four-dimensional world, by constructing three-dimensional models ( $x, y, t$, or $y, z, t$, etc.) or, still better, plane drawings in $t$ and one of the space-axes.

It is such a graphic representation that we are offered in Minkowski's inspired lecture.

Let $B_{1} O B$ (Fig. 12) be the axis of $c t$, and $A_{1} O A$ that of $x$.* Draw the straight line $L_{1} O L$ bisecting the right angle $A O B$. This line would represent the world-line of a particle moving uniformly,


Frg. 12.
along the axis of $x$, with the velocity of light $c$. Now, according to one of the assumptions of the theory of relativity, the velocity $\geqslant$ of any particle is always smaller than $c$, or at least does not exceed $c$. Consequently no world-line will be steeper than, or even as steep as, $L_{1} O L$ or $N_{1} O N$. Every world-line passing through $O$, i.c. belonging to a particle for which $x=0$ at the instant $t=0$, is entirely confined to the region consisting of $L O N$ and $L_{1} O N_{1}$. For, to penetrate into $L O N_{1}$ or $N O L_{1}$, the particle would have to move, at least during a certain part of its wandering, with a hypervelocity.

* The plotting of $x$ and the time against one another has, of course, nothing novel about it. It is familiar to everybody from elementary text-books on mechanics by the name of a 'displacement curve.' But none the less its application to relativistic connexions has been a happy idea.

Let $O B^{\prime}$ be a world-line representing a particle in uniform motion with velocity $y=c \beta$. Then

$$
\tan \psi=\beta
$$

where $\psi$ is the angle $B O B^{\prime}$. Notice that our previous angle $\omega$, endowed with the remarkable additive property with regard to the composition of parallel velocitics, is connected with this real angle $B O B^{\prime}$ by $\tan \omega=\iota \tan \psi$. By what has been said above, the absolute value of the trigonometric tangent of this angle is smaller than unity,

$$
|\tan \psi|<\mathrm{r}
$$

Now, to obtain a representation of the Lorentz transformation from $S(x, t)$ to the system $S^{\prime}\left(x^{\prime}, t^{\prime}\right)$ attached to our uniformly moving particle, draw the hyperbola

$$
\begin{equation*}
x^{2}-c^{2} t^{2}=-1 \tag{r7}
\end{equation*}
$$

and the conjugate hyperbola

$$
\begin{equation*}
x^{2}-c^{2} t^{2}=\mathrm{I} \tag{r8}
\end{equation*}
$$

of which the previous $L_{1} O L, N_{1} O N$, given by

$$
\begin{equation*}
x^{2}-\epsilon^{2} t^{2}=0 \tag{19}
\end{equation*}
$$

is the common pair of asymptotes.
In order to represent the particle as being at rest, i.e. in order to pass from $S$ to $S^{\prime}$, take $O B^{\prime}$, insteacl of $O B$, as the new axis of time, that is to say of $c t^{\prime}$, and as the axis of $x^{\prime}$ a straight line $O A^{\prime}$, such that

$$
L O B^{\prime}=L O A^{\prime}
$$

or

$$
A O A^{\prime}=B O B^{\prime}=\psi
$$

and, instead of $O A$ and $O B$, the segments $O A^{\prime}$ and $O B^{\prime}$ as the units of $x^{\prime}$ and $c t^{\prime}$, as explained in Fig. 12. The obvious proof that this is equivalent to the Lorentz transformation $x^{\prime}=\gamma(x-v t)$, $t^{\prime}=\gamma\left(t-v x / c^{2}\right)$, is left to the reacler. Notice further that, by construction, $O A^{\prime}$ and $O B^{\prime}$ are conjugate semi-diameters of the hyperbola $x^{2}-c^{2} t^{2}=-1$, as were also $O A$ and $O B$.

Thus, the Lorentz transformation consists in passing from one to another pair of conjugate semi-diameters of the hyperbola $x^{2}-c^{2} t^{2}=-1$ and in taking their lengths as the new units for $x$ and $c t$.*

[^36]The new $x$ - and $t$-axes are obtained by turning each of the old ones, towards or away from the asymptote $O L$, through the angle

$$
\psi=\arctan \beta
$$

not exceeding $45^{\circ}$.
Since $x^{2}-c^{2} t^{2}$ is invariant with respect to the Lorent\% transformation, the asymptotes $L_{1} O L$ and $N_{1} O N$ and the hyperbolace are fixed, i.e. remain always the same no matter whether $a, t$ or $x^{\prime}, t^{\prime}$ or $x^{\prime \prime}, t^{\prime \prime}$, etc., are chosen as variables. The same property belongs of course to the whole system of hyperbolae

$$
x^{2}-c^{2} t^{2}=-\kappa^{2}
$$

and of the conjugate hyperbolae

$$
x^{2}-c^{2} t^{2}=\kappa^{2}
$$

where $\kappa$ is any real number. The asymptotes may be considercel as a particular, limiting case of these curves, corresponding $10 \kappa=0$.

The reader is recommencled to compare the case under con" sideration with that of an ordinary rotation of a planc, say $x, y$, in itself, when $x^{2}+y^{2}=\kappa^{2}$ is invariant, giving circles, instead of hyperbolae, as permanent paths of the points of the plance, and a single fixed point $\kappa=0$ instead of a pair of straight lincs. In connexion with this remark hyperbolic functions may convenicmely be introduced, to replace the ordinary sine and cosine. Writing

$$
\begin{equation*}
\tan \psi=\tanh a \tag{20}
\end{equation*}
$$

Fig. 12 will easily lead to the formulae

$$
\left.\begin{array}{l}
x^{\prime}=x \cosh a-c t \sinh a  \tag{21}\\
c t^{\prime}=c t \cosh a-x \sinh a,
\end{array}\right\}
$$

which agree with (II), since, by (20) and ( I 2 ),

$$
\begin{equation*}
\omega=\iota a \tag{22}
\end{equation*}
$$

Remember that, by the definition of the hyperbolic functions,

$$
\begin{aligned}
& \sinh a=-\iota \sin (\iota a) \\
& \cosh a=\cos (\iota a) .
\end{aligned}
$$

Notice that, the region of $O B^{\prime}$ being $L O N,^{*}$ a time-axis can bo

* For positive values of $t$, and $L_{1} O N_{1}$ for negative values of $t$, and similarly as regards $L O N_{1}$ when the $x$-axis is in question.
drawn from $O$ through any world-point situated in this region, that is to say, through any point for which

$$
\begin{equation*}
x^{2}-c^{2} t^{2}<0 \tag{23a}
\end{equation*}
$$

Similarly, the region of $O A^{\prime}$ being $L O N_{1}$, an $x$-axis can be drawn through any world-point for which

$$
\begin{equation*}
x^{2}-c^{2} t^{2}>0 \tag{23b}
\end{equation*}
$$

so that any of such world-points can be made simultaneous with $O$. This is an eminently characteristic feature of the new doctrine as distinguished from the old system of physics in which simultaneity was an absolute property of events, independent of our choice of a standpoint. It is plainly an immediate consequence of the reform, of the concept of simultaneity introduced by Einstein. Pairs of events are or are not simultaneous according to the choice of our standpoint, i.e. of one out of an infinity of legitimate systems $S, S^{\prime}$, etc., in exactly the same way as pairs of space-points have or have not equal values of $x$ and $y$ (or $y$ and $z$, or $z$ and $x$ ) according to our choice of the coordinate-planes. There is thus far an intrinsic similarity, a kind of coordinateness, between space and time, or as the Time Traveller, in a wonderful anticipation of Mr. Wells, puts it: 'There is no difference Zetzeeen Time and Space except that our consciousness moves along it.'*

The process of passing the time-axis through a world-point corresponding to a given particle, since it brings it to rest, is often referred to as transforming that particle to rest. $\dagger$ In view of the above property, a vector ('world-vector,' to be treated fully further on) in the plane $x, c t$, drawn from $O$ to any point of the region limited by $L O N$, or $L_{1} O N_{1}$, satisfying the condition (23a), may be called a time-like vector, and a vector drawn from $O$ towards any

[^37]point of the remaining region, $L O N_{1}$ or $L_{1} O N$, i.e. satisfying (23b), a space-like vector.* On the border between these two classes of vectors we have singular vectors, drawn from the origin to any point of the asymptotes, i.e. coinciding, in this bi-dimensional case, in fact, with parts of the asymptotes, and characterized by $x^{2}-c^{2} t^{2}=0$. For $t>0$ the world-point at the end of a singular vector represents a particle when it just receives a light-signal from $O$, that is to say, a signal started at $x=0$ at the instant $t=0$. Similarly, for $t<0$, the end-point of a singular vector represents a particle just at the instant when it sends a light-flash which arrives at $x=0$ at the instant $t=0$. Or, as Minkowski puts it, $L_{1} O N_{1}$ consists of all the worldpoints that send light towards $O$, and $L O N$ of all those that receive light from 0 .
Notice that $x= \pm c t$, if transformed, gives $x^{\prime}= \pm c t^{\prime}$, which follows from the invariance of $x^{2}-c^{2} t^{2}$ (together with the requirement $x^{\prime}=x$, $t^{\prime}=t$ for $v=0$ ), and is only a verification of the assumption, made at the outset, that the velocity of light in empty space is the same for all legitimate systems of reference. In this case both $x$ and $t$ are reduced by the Lorentz transformation in the same ratio. In fact, substituting $x=c t$ in $x^{\prime}=\gamma(x-v t), t^{\prime}=\gamma\left(t-v x / c^{2}\right)$, we obtain
$$
x^{\prime}: x=t^{\prime}: t=(\mathrm{x}-\beta)^{\frac{1}{2}}(\mathrm{x}+\beta)^{-\frac{1}{2}} .
$$

Thus far we have considered besides $t$ one independent variable only, the space coordinate $x$. Accordingly, any world-linc, traced in that bi-dimensional diagram, has been the representation of a particle, or, in the limiting case, of a flash of light travelling along a straight line, the $x$-axis. Now, bring in the coordinate $y$. Then the resulting three-dimensional diagram or model will be appropriate to represent the motion of a particle, or the propagation of light, in a plane, the plane of $x, y$. Return to Fig. 12, and imagine the axis of $y$ to be drawn through $O$ perpendicularly to the paper. To obtain the required representation, we have only to spin the two hyperbolac of Fig. 12 and their asymptotes round $B_{1} O B$ as axis. The two branches of the hyperbola ( 17 ) will generate a hyperboloid of revolution of two sheets

$$
\begin{equation*}
x^{2}+y^{2}-c^{2} t^{2}=-\mathrm{r}, \tag{24}
\end{equation*}
$$

and the two branches of the hyperbola (18), exchanging roles after

[^38]a rotation through $180^{\circ}$, will give rise to a hypertholoid of revolution of one sheet
\[

$$
\begin{equation*}
x^{2}+y^{2}-c^{2} t^{2}=\mathrm{I} \tag{25}
\end{equation*}
$$

\]

which will be cut by the $y$-axis in a pair of points, say, $C$ and $C_{1}$, one above and the other below the paper, while the asymptotic lines will generate a right cone

$$
\begin{equation*}
x^{2}+y^{2}+z^{2}-c^{2} t^{2}=0 \tag{26}
\end{equation*}
$$

the asymptotic cone of the hyperboloids. As regards this conic surface, let us distinguish its two parts $L_{1} O N_{1}$ and $L O N$ (revolved), corresponding to negative and positive times respectively, and let us call the first the fore-cone and the second the aft-cone of $O$.* The fore-cone consists of all world-points, out of those under consideration, which 'send light' towards $O$, and the aft-cone of all those which 'receive light' from $O$. Any vector drawn from $O$ to a world-point contained within the fore- or aft-cone will be a timelike vector, and vectors drawn from $O$ to any point of the remaining region of the world, outside the cones, will be space-like vectors.

Now, let $v$ be the ordinary vector-velocity of a particle in uniform motion, and let it have any direction whatever in the plane of $x, y$. Then the world-line of this particle will be a straight line passing through $O$ in the plane $v, O B$, and including with $O B$, the original time-axis, the angle

$$
\psi=\arctan \beta
$$

where $\beta=v / c$. To transform the particle to rest, take this worldline as the axis of $c t^{\prime}$, and to obtain at the same time the new coordinates $x^{\prime}, y^{\prime}$ turn the old plane $x y^{\prime}$ through the angle $\psi$ round an axis passing through $O$ and perpendicular to both v and $O B$. For the moment, call the coordinates measured in the $x y$-plane, along $\mathbf{v}$ and perpendicularly to $i t, \xi$ and $\eta$ respectively. Then the turning round of that plane from its original position ( $t=0$ ) will amount to writing

$$
c t=\xi \cdot \tan \psi=\beta \xi
$$

On the other hanc, we have

$$
x^{2}+y^{2}=\xi^{2}+\eta^{2}
$$

for any point of the plane $x y$, so that (25) will become

$$
\xi^{2}+\eta^{2}-c^{2} t^{2}=\mathrm{I}
$$

* Minkowski, 'Vorkegrel' and 'Nachkegel.'

The intersection of the new plane, $x^{\prime} y^{\prime}$, with the surface (25), will, therefore, be given by

$$
\left(\mathrm{I}-\beta^{2}\right) \xi^{2}+\eta^{2}=\mathrm{I} .
$$

Now, $\beta^{2}<\mathrm{I}$. Thus the $x^{\prime} y^{\prime}$ plane will cut the one-sheeted hyperboloid in an ellipse. To complete the Lorentz transformation we have only to take the semi-diameters of this ellipse as the nere units of length measured from the origin along any direction in the $x^{\prime} y^{\prime}$ plane. The major principal axis of this metric ellipse will be contained in the plane $\mathbf{v}, O B$, and the other axis will be normal to it. This ellipse of our graphical representation will, of course, in the new units of length, be a circle, i.e. $x^{\prime 2}+y^{\prime 2}=\mathrm{r}$. So also did the old plane of coordinates ( $x y^{\prime}$ ) cut the one-sheeted hyperboloid in a circle $x^{2}+y^{2}=\mathrm{I}$. This is seen at once to agree with the invariance of $x^{2}+y^{2}-c^{2} t^{2}$. We have generally

$$
x^{2}+y^{2}-c^{2} t^{2}=x^{\prime 2}+y^{\prime 2}-c^{2} t^{\prime 2}
$$

and since the sections under consideration are obtained by putting $t=0, t^{\prime}=0$ respectively, the $S$-circle

$$
x^{2}+y^{2}=\mathrm{I}
$$

has for its $S^{\prime}$-correspondent the circle

$$
x^{\prime 2}+y^{\prime 2}=\mathbf{I}
$$

The new unit of time, i.e. of $c t^{\prime}$, is again represented by the segment of the $c^{\prime}$-axis cut off by one of the sheets of the two-sheeted hyperboloid of revolution, i.e. by the semi-diameter conjugate to the plane $x^{\prime} y^{\prime}$. So also was the old time-axis, $O B$, conjugate to the old plane of coordinates $(x y)$, and the unit of $c t$ was the semidiameter $O B$.

To resume this three-dimensional graphic representation:
The Lorentz transformation consists in passing from one to another set composed of a time-like semi-diameter and the conjugate space-like semi-diameters of the hyperboloid

$$
x^{2}+y^{2}-c^{2} t^{2}=-\mathrm{x}
$$

and in taking the lengths of the new semi-diameters as the units for the time (ct') and for the space-coordinates; the units of length being thus given in each case by the semi-diameters of the ellipse cut out from the one-sheeted hyperboloid $x^{2}+y^{2}-c^{2} t^{2}=1$ by the plane of coordinates.

The new time-axis and the new coordinate-plane are obtained by turning each of the old ones, towards or away from the asymptotic cone round an axis passing through $O$ and perpendicular both to the old time-axis and to the velocity $\mathbf{v}$ of the new system with respect to the old one.

Having gone through all of this, we can now pass to the most general, four-dimensional case. Here, it is true, our imagery fails us. But we can still advantageously avail ourselves of the geometrical language as a guide to, and as a short expression of, the analytical process involved.
Instead of the hyperboloidic surfaces we have now the two-'sheeted' hyperboloidic space or, as we may conveniently call it, the double hyperbolotd

$$
\begin{equation*}
r^{2}-c^{2} t^{2} \equiv x^{2}+y^{2}+z^{2}-c^{2} t^{2}=-\mathrm{x} \tag{27}
\end{equation*}
$$

and its conjugate, the one-'sheeted' hyperbolodic space or the single hyperboloid

$$
\begin{equation*}
r^{2}-c^{2} t^{2} \equiv x^{2}+y^{2}+z^{2}-c^{2} t^{2}=\mathrm{I}, \tag{28}
\end{equation*}
$$

with their common asymptotic conic space

$$
\begin{equation*}
r^{2}-c^{2} t^{2} \equiv x^{2}+y^{2}+z^{2}-c^{2} t^{2}=0, \tag{29}
\end{equation*}
$$

consisting of the fore-cone $t<0$ and the aft-cone $t>0$, as before, with the only difference that these, like the hyperboloids, are now three-dimensional entities.

The $t$-axis cuts the double hyperboloid (27) in a pair of points, namely
and

$$
x=y=z=0, \quad c t=1
$$

$$
x=y=z=0, \quad c t=-1 .
$$

Take the first, contained in the positive 'sheet.' Call it $P$, so that $O P$, a semi-diameter of the hyperboloid (27), is the old timeaxis, and the length of this semi-diameter is the unit of ct. The space $t=0$, that is to say, the ordinary space-manifold $x, y, z$ is the three-space conjugate to the semi-diameter $O P$, just as the $x y$-plane, in the previous case, was conjugate to $O B$ (Fig. 12). Now, instead of $P$, take any other point $P^{\prime}$ of the positive sheet of (27), and consider $O P^{\prime}$ as the new time-axis and the length of this semidiameter as the unit of $c t^{\prime}$. Turn the $x y z$-space ( $t=0$ ) which cut the single hyperboloid (28) in a sphere,

$$
x^{2}+y^{2}+z^{2}=x,
$$

round the plane passing through $O$ and perpendicular to $v$, till this space, or pencil of semi-diameters, becomes conjugate to the semi-diameter $O P^{\prime}$. Then it will become the $x^{\prime} y^{\prime} z^{\prime}$-space. This space cuts the single hyperboloid in an ellipsoid (ellipsoidic surface). Take the semi-diameters of this ellipsoid as the new units of length measured from the origin along any direction in the $x^{\prime} y^{\prime} z^{\prime}$ 'space. Then the Lorentz transformation, from $S$ to $S^{\prime}$, will be completed, and the new metric surface which, from the $S$-point of view, is an ellipsoid of revolution will for the $S^{\prime}$-standpoint become a sphere,

$$
x^{\prime 2}+y^{\prime 2}+z^{\prime 2}=\mathrm{I}
$$

So also was the old metric surface, viewed from the old standpoint, a sphere of unit radius. Remember that $O P^{\prime}$ is time-like, i.e. contained within the four-dimensional region bounded by the threedimensional cone, but otherwise the choice of this axis as a time-axis is free. The possible positions of $p^{\prime}$ constitute a triple manifold, namely all the points of the positive sheet of (27). Thus, the systems $S^{\prime}\left(x^{\prime}, y^{\prime}, z^{\prime}, t^{\prime}\right)$ equally legitimate with $S$ are $\infty^{3}$, as has been repeatedly observed.*

To resume what has just been said with regard to the general, four-dimensional case:

The Lorentz transformation consists in passing from one (timelike) semi-diameter $O P$ and the pencil of conjugate (space-like) semi-diameters of the hyperboloid $r^{2}-c^{2} t^{2}=-1$ to another semidiameter $O P^{\prime}$ with its corresponding pencil of conjugate semidiameters, and in taking the lengths of the new semi-diameters as the units of time ( $\left(t^{\prime}\right)$ and of space-coordinates; the units of length being thus given in each case by the semi-diameters of the ellipsoid cut out from the hyperboloid $r^{2}-c^{2} t^{2}=x$ by the new space of coordinates.

The property of two lines $O P_{1}$ and $O P_{2}$ being conjugate may be expressed analytically by the equation

$$
\begin{equation*}
x_{1} x_{2}+y_{1} y_{2}+z_{1} z_{2}-c^{2} t_{1} t_{2}=0, \tag{30a}
\end{equation*}
$$

where $x_{1}, y_{1}, z_{1}, c t_{1}$ and $x_{2}, y_{2}, z_{2}, c t_{2}$ are the values of the four

[^39]variables defining the world-points $P_{1}$ and $P_{2}$ respectively, or, using the ordinary vectors $r_{1}$ and $r_{2}$,
\[

$$
\begin{equation*}
\left(\mathbf{r}_{1} \mathbf{r}_{2}\right)-c^{2} t_{1} t_{2}=0 \tag{30b}
\end{equation*}
$$

\]

or finally, writing $l=\iota c t$,

$$
\begin{equation*}
\left(r_{1} r_{2}\right)+l_{1} l_{2}=0 \tag{30}
\end{equation*}
$$

By an obvious analogy such lines $O P_{1}, O P_{2}$ are also called mutually perpendicular or normal lines in the world $x, y, z, l$. Notice that this property of a pair of lines is invariant with respect to the Lorentz transformation, i.e. that (30) is transformed into

$$
\left(\mathbf{r}_{1}^{\prime} \mathbf{r}_{2}^{\prime}\right)+l_{1}^{\prime} l_{2}^{\prime}=0
$$

In other words, conjugate diameters remain conjugate, independently of the choice of a reference-system. This is obvious, at least in two and three dimensions. More generally, for any pair of lines $O P_{1}, O P_{2}$,

$$
\begin{equation*}
\left(\mathrm{r}_{1}^{\prime} \mathrm{r}_{22}^{\prime}\right)+l_{1}^{\prime} l_{2}^{\prime}=\left(\mathrm{r}_{1} \mathrm{r}_{2}\right)+l_{1} l_{2}, \tag{3I}
\end{equation*}
$$

as the reader himself may prove, using for instance the form (1b) of the Lorentz transformation, and noticing that

$$
\left(\epsilon r_{1}, \epsilon r_{2}\right)-\gamma^{2} c^{-2}\left(r_{1} v\right)\left(r_{2} v\right)=\left(r_{1} r_{2}\right)
$$

identically. Thus, the invariance of orthogonality is but a particular case of the invariance of

$$
\left(\mathrm{r}_{1} \mathrm{r}_{2}\right)+l_{1} l_{2}
$$

We shall return to the last property later on.
Given the origin $O$ (and any world-point can be made the origin), the set of any four values of

$$
x, y, z, l
$$

or, more generally, of any four scalar magnitudes

$$
z v_{x}, z e_{y}, z w_{z}, s
$$

zehich are transformed like $x, y, z, l$ respectively, and of which the first three are real and the fourth purely imaginary, defines what is called a world-vector or space-time vector of the first kind* (Minkowski) or a four-vector (Sommerfeld).

[^40]Thus, if (30) is satisfied, the four-vectors $O P_{1}$ and $O P_{12}$ are said to be perpendicular to one another. Generally, if

$$
\begin{equation*}
\left(\mathrm{w}_{1} \mathrm{w}_{2}\right)+s_{1} s_{2}=0, \tag{32}
\end{equation*}
$$

then $\mathbf{w}_{1}, s_{1}$ and $\mathbf{w}_{2}, s_{2}$ form a pair of perpendicular four-vectors. Here $\mathrm{w}_{1}$ is the ordinary or three-vector whose components arc $z e_{1}, z v_{y}^{\prime}, z e_{i}^{\prime}$, and $w_{2}$ has a similar meaning, while ( $w_{1} w_{2}$ ) is, as before, the ordinary scalar product of $w_{1}, w_{2}$.
Any four-vector drawn from $O$ to a world-point contained within the asymptotic cone, i.e. such that $r^{2}-c^{2} t^{2}=r^{2}+l^{2}<0$ or, more generally, any four-vector $\mathrm{w}, s$, such that

$$
\begin{equation*}
z v^{2}+s^{2}<0, \tag{33t}
\end{equation*}
$$

is called, as in the two- and three-dimensional cases, a time-dike vector, while four-vectors satisfying the condition $r^{2}+l^{2}>0$ or, generally,

$$
\begin{equation*}
z v^{2}+s^{2}>0, \tag{33s}
\end{equation*}
$$

are called space-like vectors.
The reader will easily prove that if one of a pair of normal fourvectors is time-like, the other is space-like, or that, in other words, if one is contained within the asymptotic cone, the other is outside it.
Again, as in the above special cases, any vector drawn from () towards a point of the asymptotic cone, whether the fore- or alt-cone, is called a singular four-vector. The analytical expression of a singular vector is $r^{2}-c^{2} t^{2}=r^{2}+l^{2}=0$, or, generally,

$$
\begin{equation*}
z w^{2}+s^{2}=0 . \tag{34}
\end{equation*}
$$

Finally, as in the less-dimensional cases, the aft-cone may be said to consist of all world-points which 'receive light' from $O$, and the fore-cone of all those that 'send light' towards 0 .
Fig. 13, which is Fig. 12 redrawn with the omission of the arbitrary axes, and thus contains only what is 'absolute' or independent of the choice of such time- and spacc-axcs, may aid the reader in remembering the meaning of the various names employed in the above representation. This figure is drawn perspectively (in three dimensions, of course), so as to show that the hyperboloids (27) and (28) are hyperboloids of revolution, the former consisting of two disconnected 'sheets' and the latter of one 'sheet.' We may mention further that the world-region contained
within the fore-cone (left) was called by Minkowski this side of $O$ and that contained within the aft-cone that side of $O$. Every worldpoint of the first region is necessarily (independently of the selection of a reference-system) or essentially earlier, and every world-point of the second region is essentially later than $O$. Any point of the remaining, cyclical, region of the world, called the intermediate region, can be made simultaneous with or earlier or later than $O$ (i.e. can be given a value of $t=$ or $<$ or $>0$ ) by an appropriate choice of the time-axis, and is therefore essentially neither earlier nor later than $O$. This region is the domain of all space-like four-

vectors which can be drawn from $O$. Between the time-like and space-like classes of world-vectors are the singular vectors, composing the cones which are three-dimensional entities.

This partitioning of the world and the characteristic properties of the cones are obviously conditioned by the assumption that no particle, or at least, no legitimate system, can ever move (relatively to another one) with a velocity $v$ exceeding that of light in empty space. In classical physics there was no limit whatever to $v$. The Newtonian transformation follows from the Lorentz transformation by taking $\infty$ instead of $c$, or, figuratively, by widening both the cones till they coalesce with one another in a plane, squeezing out the space-like four-vectors and opening the whole world to the timelike vectors. Any straight line would, in the Newtonian world,
represent a possible uniform motion of a particle with respect to certain frames of reference.
So much as regards the geometric representation of the Lorentz transformation.
Now for its analytical expression and the methods of dealing with the world-vectors.
Minkowski, though availing himself now and then of the fourdimensional vector language and ideology, made a systematical and extensive use of Cayley's calculus of matrices.* Thus, the fundamental world-vector $\mathrm{r}, l$ and, more generally, any space-time vector of the first kind $\mathrm{w}, s$ is considered as a matrix of I row and 4 columns, say,

$$
\begin{equation*}
X=|x, y, z, l| \tag{35}
\end{equation*}
$$

and, in general,

$$
W=\left|v_{x}, w_{y}, w_{z}, s\right|
$$

The transformed world-vector $r^{\prime}, l^{\prime}$ will then be another matrix of $\mathrm{I} \times 4$ constituents,

$$
x^{\prime}=\left|x^{\prime}, y^{\prime}, z^{\prime}, l^{\prime}\right|
$$

which is obtained from $X$ by taking its 'product' into a certain matrix of $4 \times 4$ constituents,

$$
A=\left|\begin{array}{cccc}
a_{11}, & a_{12}, & a_{13}, & a_{14}  \tag{36}\\
a_{21}, & a_{22}, & a_{28}, & a_{24} \\
a_{31}, & a_{32}, & a_{38}, & a_{84} \\
a_{41}, & a_{42}, & a_{43}, & a_{44}
\end{array}\right|
$$

and which is written simply

$$
\begin{equation*}
X^{\prime}=X A \tag{37}
\end{equation*}
$$

Thus, the Lorentz transformation is expressed by the matrix $A$ taken as a postfactor of the world-vector to be transformed. This matrix is characterized by the condition that its determinant is +1 ,

$$
\begin{equation*}
\operatorname{det} A=\mathrm{r} \tag{8}
\end{equation*}
$$

and further that all of its constituents containing the index 4 once
*Cf. Minkowski's Grundgleichungen, already quoted, \&s ir et seq. Those readers who are not familiar with this branch of mathematics may consult: the Note at the end of this chapter, where the definition of different kinds of matrices and some rules of operating with them are given.
only are purely imaginary, while the remaining seven constituents are real and the right lowermost positive:

$$
\left.\begin{array}{cc}
a_{11}, & a_{12}, \ldots \\
a_{14}, & a_{33}, \\
a_{41}, & a_{34} \\
& a_{42}, \\
a_{43}
\end{array}\right\} \text { pureal } \begin{gathered}
\\
a_{44}>0 .
\end{gathered}
$$

The inverse transformation is represented by the reciprocal of $A$, which is at the same time the transposed of $A, A^{-1}=\bar{A}$, so that

$$
\begin{equation*}
\bar{A} A=\ddot{A} \bar{A}=\mathrm{r} \tag{39}
\end{equation*}
$$

It is this property that insures the invariance of $r^{2}+l^{2}$. Using $\bar{A}$ and $\ddot{X}$, we may write also, instead of (37),

$$
\dot{X}^{\prime}=\bar{A} \bar{X}
$$

The short formula (37) replaces

$$
x^{\prime}=\alpha_{11} x+\alpha_{21} y+\alpha_{31} z+a_{41} l,
$$

and three similar equations, with $2,3,4$ as second indices. If, in particular, the $x, y, z$-axes are taken along $\nabla$ and normal to it, and if $x^{\prime}, y^{\prime}, z^{\prime}$ are, as before, measured along the same directions, then, as we saw,

$$
\begin{aligned}
& x^{\prime}=\gamma(x+\iota \beta l) ; \quad y^{\prime}=0 ; \quad z^{\prime}=0 ; \\
& l^{\prime}=\gamma(l-\iota \beta x) .
\end{aligned}
$$

Hence, for this particular choice of coordinate-axes the matrix representing the Lorcntz transformation reduces to

$$
A=\left|\begin{array}{cccc}
\gamma, & \circ, & \circ, & -\iota \beta \gamma  \tag{40}\\
\circ & \mathrm{\imath}, & \circ, & \circ \\
\circ, & \circ, & \mathrm{r}, & \circ \\
\iota \beta \gamma, & 0, & \circ, & \gamma
\end{array}\right|
$$

The transposed matrix $\bar{A}$ which represents the inverse Lorentz transformation is obtained from this by simply changing the signof $\beta$, as it should be.

Writing, instead of $x, \ldots l$, the differentiators $\partial / \partial x, \ldots \partial / \partial l$, we obtain a matrix of $1 \times 4$ constituents, which Minkowski called lor, in honour of Lorentz,

$$
\begin{equation*}
\text { lor }=\left|\frac{\partial}{\partial x}, \quad \frac{\partial}{\partial y}, \quad \frac{\partial}{\partial z}, \quad \frac{\partial}{\partial l}\right| \cdots \tag{4x}
\end{equation*}
$$



This is the matrix-equivalent of our quaternionic differential pperator $D$, as defined by (I3), Chap. II. It can be easily verified that $\partial / \partial x, \ldots \partial / \partial l$ are transformed in exactly the same way as. $x, y, z, l$ respectively.* Thus, lor is covariant, or equally, traninsformed, with the matrix $X$ representing the standard world-vector, i.e.

$$
\begin{equation*}
\operatorname{lor}^{\prime}=\operatorname{lor} A \tag{42}
\end{equation*}
$$

Moreover, it has the same structure as $X$, its first three coinstituents (differentiators) being real and the fourth, $\partial / \partial l$, purely imaginary. Thus, lor, though an operator, behaves in every respect like a space-time vector of the first kind.

We cannot stop here to consider the matrix form of space-time vectors of the second kind and their analytical connexion with those of the first kind (although it could be done in a few lines), for the reader does not yet know their relativistic physical significance. Moreover, it is not our purpose to develop fully the matrix method of treating relativistic questions, since we shall avail ourselves chicfly of other methods. But one simple property of products of $W /$-matrices in connexion with the above remarks is worth mentioning here, namely that, if $W_{1}, W_{2}$ are matrices representing a pair of vectors of the first kind ( $w_{1}, s_{1} ; w_{2}, s_{2}$ ), the product

$$
\begin{equation*}
W_{1} \bar{W}_{2}=\left(\mathbf{w}_{1} \mathbf{w}_{2}\right)+s_{1} s_{2} \tag{43}
\end{equation*}
$$

is an invariant. For by (39), and by the associative property of products of matrices,

$$
W_{1}^{\prime} \bar{W}_{2}^{\prime}=W_{1} A \cdot \bar{A} \bar{W}_{2}=W_{1} \ddot{W}_{2}
$$

$$
\begin{aligned}
& \text { *Thus, for instance, measuring } x \text { along } \mathbf{v}, \\
& \text { we have } \\
& \qquad x^{\prime}=\gamma(x+\iota \beta l) ; y^{\prime}=y ; z^{\prime}=z ; l^{\prime}=\gamma(l-\iota \beta x), \\
& \text { whence } x=\gamma\left(x^{\prime}-\iota \beta l^{\prime}\right) \text {, etc., and } \\
& \quad \frac{\partial}{\partial x^{\prime}}=\gamma\left(\frac{\partial}{\partial x}+\iota \beta \frac{\partial}{\partial l}\right) ; \quad \frac{\partial}{\partial y^{\prime}}=\frac{\partial}{\partial y} ; \frac{\partial}{\partial z^{\prime}}=\frac{\partial}{\partial z} ; \frac{\partial}{\partial l^{\prime}}=\gamma\left(\frac{\partial}{\partial l}-\iota \beta^{\prime} \frac{\partial}{\partial x}\right) .
\end{aligned}
$$

Thus, the orthogonality of two four-vectors, which is an invariant property, is expressed by

Similarly,

$$
W_{1} \bar{W}_{2}=0
$$

$$
\operatorname{lor}^{\prime} \bar{W}^{\prime}=\operatorname{lor} \bar{W}
$$

or lor $\bar{W}$ is a relativistic invariant. Notice that, similarly to (43),

$$
\text { lor } \vec{W}=\frac{\partial w_{x}}{\partial x}+\frac{\partial z w_{y}}{\partial y}+\frac{\partial z w_{z}}{\partial z}+\frac{\partial s}{\partial l},
$$

or, using div in its ordinary sense,

$$
\begin{equation*}
\operatorname{lor} \bar{W}=\operatorname{div} \mathrm{w}+\frac{\partial s}{\partial l} . \tag{44}
\end{equation*}
$$

So much as regards Minkowski's matrix-form of the fundamental relativistic connexions.

Sommerfeld, whose aim was to elucidate Minkowski's ideas, replaced his language of matrices by a four-dimensional vector-algebra (and -analysis) which he developed in two very lucid papers,* and which is an obvious generalization of the familiar threedimensional calculus of vectors. Sommerfeld begins by drawing our attention to the well-known circumstance that in space of three dimensions there are two kinds of vectors to be distinguished, e.g. vectors of the 'first kind' or polar, and those of the 'second kind' or axial vectors. A vector of the first kind, such as a translation velocity, is a segment of a straight line having a certain direction (and sense) ; its components are the projections upon the coordinateaxes. On the other hand, a vector of the second kind, such as angular velocity, is represented by a plane surface of a certain area with a given sense of circulation round its circumference, and its components are the projections of that area upon the coordinateplanes. Consequently, the components of a vector of the first kind should be written with single indices, $v_{x}, v_{y}, v_{z}$; or $v_{1}, v_{2}, v_{3}$, while those of a vector of the second kind, as, for instance, rotational velocity $\omega$, with double indices, $\omega_{y z}, \omega_{z x}, \omega_{x y}$, or $\omega_{23}, \omega_{31}, \omega_{12}$. This discrimination, which in three dimensions is not very important (or at least ceases to be so when, instead of the plane area, a representative line-segment normal to it is introduced), becomes in
*A. Sommerfeld, 'Zur Relativitätstheorie. I. Vierdimensionale Vektoralgebra,' Ann. der Physik, Vol. XXXII., Igro, p. 749, and 'II. Vierdimensionale Vektoranalysis, ${ }^{\text {' }}$ Annz. der Phjisik, Vol. XXXIÏ., igio, p. 649.

Minkowski's four-dimensional world quite essential. For hereargues Sommerfeld-we have

$$
\begin{gathered}
\binom{4}{1}=\text { four coordinate axes, } \\
x, y, z, l, \\
\binom{4}{2}=s i x \text { coordinate planes, } \\
v z, z x, x y, x l, y l, z l,
\end{gathered}
$$

and

$$
\begin{gathered}
\binom{4}{3}=\text { four coordinate spaces, } \\
x y z, y z l, z x l, x y l .
\end{gathered}
$$

Accordingly, we have to distinguish in the 'world' between vectors of the first kind having four components, or four-vectors ; those of the second kind having six components, or six-vectors ; and, finally,
those of the third kind, which again have four components, and can be replaced by their 'supplements,' which are vectors of the first kind.

Consequently, vectors of both the first and the third kind are called by Sommerfeld, summarily, four-vectors.
This classification will be found useful for what is to occupy us later on. But meanwhile we are concerned only with space-timevectors of the first kind, which we shall simply call four-vectors.

The standard or typical example of such vectors is that drawn from the origin $O$ to any world-point. Call it $P$.* Then its components would be, according to Sommerfeld's general notation,

$$
P_{x}, \quad P_{y}, \quad P_{z}, \quad P_{l} .
$$

These, of which the first three are real and the last imaginary, are simply the previous

$$
x, \quad y^{\prime}, \quad z, \quad l .
$$

* Sommerfeld does not use any special type of print for his fuur-vectors, to distinguish them from six-vectors. A certain uniformity of notation was introduced later by Laue, loc. cit. But we shall not want very nuch of it for our subsequent purposes.

What Sommerfeld denotes by $|P|$ and calls the size of the vector $P$, or its length, i.e. the 'length' of the corresponding four-dimensional straight line, is the positive (or positive imaginary) value of

$$
\sqrt{x^{2}+y^{2}+z^{2}+l^{2}}=\sqrt{x^{2}+y^{2}+z^{2}-c^{2} t^{2}}
$$

or of $\sqrt{r^{2}-c^{2} t^{2}}$.
The length of this, and of every other, four-vector is invariant with respect to any Lorentz transformation. It is its only invariant.
Notice that the length, thus defined, of a four-vector may be either real, or purely imaginary, or nil, according as we have what was previonsly called a space-like, a time-like, or a singular vector.
If $A, B$ be any pair of four-vectors, the sum of the products of their corresponding components is called their scalar product, and is denoted by $(A B)$. Thus

$$
\begin{equation*}
(A B)=\left(A_{x} B_{w}+A_{y} B_{y}+A_{z} B_{z}+A_{l} B_{l}\right) . \tag{45}
\end{equation*}
$$

Guided by the analogy of ordinary vector-algebra, Sommerfeld defines then the direction-cosine of $A$ relative to $B$, or vice versa, by writing

$$
\begin{equation*}
\left(A B^{\prime}\right)=|A| \cdot|B| \cdot \cos (A, B) \tag{45a}
\end{equation*}
$$

Consequently, when

$$
(A B)=0
$$

the four-vectors $A, B$ are said to be perpendicular to one another This is identical with the previous definition of pairs of perpendicular vectors.

What Sommerfeld calls the 'vector product' of $A, B$ cannot here occupy our attention. For such a product is a (special) six-vector, which as yet is unfamiliar to us.

As to the Lorentz transformation itself, it appears in Sommerfeld's treatment as a rotation of the system of four axes. Let $P$ be any four-vector, and $P_{x}$, etc., its components along the old axes; then Sommerfeld defines the components of $P$ along the new axes by

$$
\Gamma_{w}^{\prime \prime}=P_{x \mathrm{x}} \cos \left(x^{\prime}, x\right)+P_{y} \cos \left(x^{\prime}, y^{\prime}\right)+P_{z} \cos \left(x^{\prime}, z\right)+P_{l} \cos \left(x^{\prime}, l\right), \quad(46)
$$

and by similar formulae for $P_{y}, P_{k^{\prime}}, P_{l^{\prime}}$. Here, the meanings of the cosines are defined by (45a). If the $x^{\prime}$-axis is space-like, then the first three cosines in (46) are real, while $\cos \left(x^{\prime}, l\right)$ is purely imaginary, like $P_{l}$, so that $P_{a d^{\prime}}$ is real. Similarly, the $y^{\prime}$ - and $z^{\prime}$-axes being space-like, $P_{y^{\prime}}, P_{z^{\prime}}$ will be real. And the $l^{\prime}$-axis being time-
like, $P_{l^{\prime}}$ will be purely imaginary. In order to show that the projection- or component-formulae (46), etc., are identical with those of the Lorentz transformation, Sommerfeld considers the particular case of rotation round the $y z$-plane, i.e. in the $x$-plane, when

$$
\begin{aligned}
& \cos \left(x^{\prime}, x\right)=\cos \left(l^{\prime}, l\right), \text { say }=\cos \omega, \\
& \cos \left(x^{\prime}, l\right)=-\cos \left(l^{\prime}, x\right)=\sqrt{\mathrm{r}-\cos ^{2}(\omega)=\sin \omega,} \\
& \cos \left(y^{\prime}, y\right)=\cos \left(z^{\prime}, z\right)=\mathrm{x},
\end{aligned}
$$

while all other cosines vanish. Here we have, obviously, $\cos \omega>\mathrm{I}$, so that the angle $\omega$, as well as its sine and tangent, are purely imaginary, and the absolute value of the latter is $<\mathrm{x}$. Consequently we can write

$$
\tan \omega=\iota \beta, \quad \cos \omega=\left(\mathrm{r}-\beta^{2}\right)^{-\frac{1}{2}}=\gamma, \quad \sin \omega=\iota \beta \gamma,
$$

so that (46), etc., are at once reduced to the formulae (ri), on p. 127 , with the same meaning of $\omega$, provided that the new system of spaceaxes ( $x^{\prime} y^{\prime} z^{\prime}$ ) moves relatively to the old one ( $x y z$ ) with the uniform velocity $v=c \beta$ along the $x$-axis. There is in fact no difference whatever between Sommerfeld's and Minkowski's method of representing the relativistic transformation.
It is true that the systematic use of the four-dimensional vector language may offer some advantages, when compared with that based on the use of matrices. But, on the other hand, there are rather important arguments which may be brought forward in defence of the matrix-method. Thus, for instance, Sommerfeld's 'scalar product,' say $(A B)$, is the same thing as Minkowski's product of the corresponding matrices, $W_{1} \bar{W}_{2}$, (43). But whereas the invariance of $W_{1} \bar{W}_{2}$ is seen at a glance, viz. by writing, in virtue of the fundamental formula (39),

$$
W_{1} A \cdot \bar{A} \bar{W}_{2}=W_{1} W_{2}
$$

the invariance of $(A B)$ cannot be proved without splitting the fourvectors into their components and multiplying out expressions like (46) and adding them up. For Sommerfeld's only definition of 'scalar product'. (45) is of such a character. It is essentially Cartesian, not vectorial. Of course, we know that, in threedimensional space, the scalar product of a pair of vectors can be, and generally is, defined without any reference to axes, so that its invariance with respect to space rotations requires no proof. But
this does not by itself enable us to see the invariance of $(A B)$, when we are asked to pass into the four-dimensional world, where our imagery fails us. Similar remarks could be made with respect to other points of Sommerfeld's method of treatment. But discussions of this kind need not detain us here any further.*
In the sequel we shall not avoid either of these two methods of analytic expression. In fact, we shall now and then profitably employ matrices as well as world-vectors. But principally we shall avail ourselves of the language of Hamilton's quaternions, the utility of which for relativistic purposes I have endeavoured to show in two papers. $\dagger$ I may notice that Minkowski himself (Grundgleichungen, p. 28, footnote) despised Hamilton's calculus of quaternions as 'too narrow and clumsy for the purpose' in question. But, notwithstanding that, I am still under the impression that quaternions are admirably suitable for most, if not for all, relativistic needs. We had a sample of the conciseness of Hamilton's language in Chapter II., when we saw how easily the four vector equations of the electron theory are condensed into a single quaternionic equation, viz. $D \mathbf{B}=C$. But in advocating here the cause of quaternions I am doing so not only because they furnish us very short formulae and simplify their handling. Quite independently of this, the quaternion seems to me intrinsically better adapted than the world-vector to 'express that 'union' of time and space which was (too strongly, perhaps) emphasized by Minkowski. For, although there is a certain union between the two, which manifests itself when we pass from one system to another, there is no total fusion. In each system, out of the four scalars $x, y, z, l$, the first three are more intimately bound to one another than any of them to the last one. The first three are artificial components of a vector, $\mathbf{r}$, which certainly is a more immediate entity than each of them. Now, in a four-vector, as well as in a matrix, $x, y, z, l$ are, as it.

[^41]were, on entirely equal footing with one another, being the four 'components' of the former, or the four 'constituents' of the latter.* On the other hand, a quaternion $q$ has a distinct vector part, V. $q$ or simply $\mathrm{V} q$, and a scalar part, $\mathrm{S} q$, and none of the components of the former can be confounded with the latter. Now, the position of a particle is determined by a vector (in its ordinary sense), and its date by a scalar. What then more natural than to take the first as the V and to embody the second in the S of a quaternion? We could insist upon loosely juxtaposing both entities, and write simply
$$
\mathbf{r}, \quad l .
$$

But, if instead of the comma the plus sign is used, we have just enough of 'union' to express the relativistic standpoint, and yet enough distinction not to amalgamate time and space entirely.

Let us therefore combine the position vector $\mathbf{r}$ of a particle with its date, $l=l c t$, into a quaternion,

$$
\begin{equation*}
q=l+\mathbf{r}, \tag{47}
\end{equation*}
$$

which, if it needed a name of its own, we might call the positionquaternion. Those who are particularly fascinated by the worldconcept can consider this 'position' to be the 'position in the world.' But, in fact, the above provisional name is simply an abbreviation for 'position-date quaternion.'

The conjugate of $q$, i.e. Hamilton's $\mathrm{K} q$, will be denoted by $q_{c}$. Thus,

$$
\begin{equation*}
q_{0}=l-\mathrm{r} . \tag{47c}
\end{equation*}
$$

The reader must not be afraid of quaternions. If he is familiar with the elements of ordinary vector-algebra, the following short remarks will enable him to understand thoroughly all of our subsequent calculations.

1. Without returning to Hamilton's original expression of a quaternion as the 'ratio' or the quotient of two vectors, he can conveniently define it from the outset as the sum of a scalar and a vector, using for the latter heavy type. Thus

$$
a=\sigma+\mathrm{A}
$$

will be a quaternion, whose scalar part is $\sigma$, and whose vector part is A,

$$
\mathrm{S} a=\sigma, \quad \mathrm{V} a=\mathbf{A} .
$$

*It is true, that the fourth, $l$, is imaginary, while the first three are real, but this does not seem to me to emphasize the distinction sufficiently.
2. The conjugatc $a_{c}$ of the quaternion $a$ is defined, as above, by
i.e. by $a_{c}=\mathrm{S} a-\mathrm{V} a$.

$$
a_{c}=\sigma-\mathrm{A},
$$

3. Two quaternions $a, b$ are said to be cqual if both their scalars and their vectors are equal to one another. Thus,
means the same thing as

$$
a=b
$$

$$
\mathrm{V} a=\mathrm{V} b \text { and } \mathrm{S} a=\mathrm{S} b .
$$

4. Quaternions are added to one another by adding separately their scalars and their vectors. Thus
means the same $\quad c=a+b$

$$
\mathrm{S} c=\mathrm{S} a+\mathrm{S} b, \quad \mathrm{~V} c=\mathrm{V} a+\mathrm{V} b
$$

Now, since the addition of scalars and the addition of vectors are both commutative, the commutative property belongs also to the sum of quaternions,

$$
b+a=a+b
$$

And for the same reason the rssociative law holds for the sum of any number of quaternions. Thus

$$
a+[b+c]=[a+b]+c,
$$

so that both sides may be simply written $a+b+c$.
5. Subtroction of quaternions, and the change of the sign of a quaternion are at once reduced to the same operations applied to scalars and vectors. Thus, if $a=\sigma+\mathrm{A}$,

$$
-a=-\sigma-\mathbf{A}
$$

Also, by 4, $a-b=-b+a$.
6. Two quaternions, $a=\sigma+\mathbf{A}$ and $b=\tau+\mathrm{B}$, are multiplied by the formula

$$
a b=\sigma \tau+\tau \mathrm{A}+\sigma \mathrm{B}+\mathrm{AB}
$$

where the first three terms require no further explanation, and the last is defined to be a quaternion

$$
A B=V A B+S A B
$$

such that VAB is identical with the 'vector product' and SAB is the negative 'scalar product,' both supposed to be known from ordinary vector algebra. Thus, in our usual notation,

$$
A B=V A B-(A B)
$$

The minus sign is introduced to suit the whole of Hamilton's calculus; I do not think there is any trouble in doing so. Ultimately, the product $a b$ of a pair of quaternions is given by

$$
\begin{aligned}
& \mathrm{S} a b=\sigma \tau-(\mathrm{AB}), \\
& \mathrm{V} a b=\tau \mathrm{A}+\sigma \mathrm{B}+\mathrm{VAB} .
\end{aligned}
$$

Thus, $a b$, and similarly, the product of any number of quaternions, is again a quaternion, with uniquely determinate vector and scalar parts.

Both ( AB ) and VAB being clistributive, quaternion multiplication is distributive, i.e.

$$
\begin{aligned}
& a[b+c]=a b+a c, \\
& {[b+c] c=b c a+c a .}
\end{aligned}
$$

It can easily be shown that it is also crssociative, i.e. that

$$
a \cdot b c=a b \cdot c
$$

so that both sides may be simply written robc. The sance thing is true of the product of any number of quaternions. It is chiefly this associative property which makes Hamilton's calculus so powerful.

From the above formulae we see that

$$
S b a=S a b,
$$

because ( $\mathbf{A B}$ ), like $\sigma \tau$, is commutative. On the other hand we have, generally,
because VBA $=-\mathrm{VAB}$.
Thus, multiplication of quaternions is, generally, not commutative,

$$
b a \neq a b .
$$

It becomes commutative only when VAB vanishes, i.e. when $\mathbf{A} \| \mathbf{B}$, or $\mathrm{V} a \| \mathrm{V} b$. This is, for instance, the case for a pair of conjugate quaternions, and, consequently, we have, for any quaternion $\pi$,

$$
a a_{\mathrm{o}}=a_{\mathrm{c}} r .
$$

7. Writing again $a=\sigma+A$, we have, by 6 ,

$$
a a_{\mathrm{c}}=a_{0} a=\sigma^{2}+\mathrm{A}^{2},
$$

where $A^{2}=(\mathrm{AA})$. Thus $a a_{0}$ is always a pure scalar. Its syuare root is called the tensor of the quaternion $a$, and is denoted by $\mathrm{T} a$,

$$
\mathrm{T} a=\left(\sigma^{2}+A^{2}\right)^{\frac{1}{2}}
$$

If it is real, the positive value of the root is taken, and if purely imaginary, the positive imaginary value of the root is taken. (In cases of complex values, when ambiguity of $T$ might arise, special explanations

will be given.) But the chief thing is to keep in mind the formula for the square of the tensor,

$$
(\mathrm{T} \alpha)^{2}=\alpha \alpha_{\mathrm{c}}=\alpha_{\mathrm{c}} \alpha,
$$

which is called the norm of the quaternion $a$.
Let a be the unit of A , so that $(\mathrm{aa})=\mathrm{I}$ and $\mathrm{A}=A \mathrm{a}$. (There is, I hope, no danger of confounding the quaternion $a$ with the absolute value of a, which is r .) Then the quaternion $a$ can be written

$$
a=\mathrm{T} c \cdot[\cos \alpha+\mathbf{a} \sin \alpha]
$$

or, by an obvious analogy,

$$
a=\Gamma a \cdot e^{\alpha a}, *
$$

where $c$ is the basis of natural logarithms. The factor of $T a$, which is a quaternion of unit tensor or a unit quaternion, is called the versor of the quaternion $a$, and is denoted by U $a$, so that $a=\mathrm{T} a$. U $a$. The unit vector $a$ is called the axis of the quaternion $a$, and the angle $\alpha$, which can be real or imaginary, is called the angle of the quaternion $a$.

Thus, conjugate quaternions may be described as quaternions having equal tensors and equal angles, but opposite axes.

Notice that if (as in the case of our above $q$ ) $\sigma$ is imaginary and $\mathbf{A}$ real, or vice versa, the tensor of $a$ may vanish, though $a$ is not simply 'zero,' but a definite quaternion having a certain axis and a certain angle. Such a quaternion was called by Hamilton a nulllifier, and by Cayley a nullitat. In our physical applications we shall not avail ourselves of either of these names, but shall adopt for such quaternions the name singular, already used for the corresponding world-vectors.
8. The following rule, which will be often required, can easily be proved:

The conjugate of the product of any number of quaternions is the product of their conjugates in the reversed order.

Thus, if
then

$$
\begin{aligned}
& m=a b, \\
& m_{0}=b_{c} a_{0} .
\end{aligned}
$$

9. Finally, as regards division by quaternions, it may be entirely reduced to multiplication by what are called their reciprocals.

The reciprocal of a quaternion $a$ is again a certain quaternion, which is denoted by $a^{-1}$ and which is defined by the equation

$$
a^{-1} a=1
$$

* But this analogy cannot be pushed so far as to write in the expression of a product of two quaternions $a, b$

$$
e^{a \mathrm{a}+\beta b}
$$

and to invert the order of addends in the exponent. For, unless $a \| b$, the product $a b$ is not commutative.

Multiply both sides by $a_{c}$ as a postfactor. Then $a^{-1} a a_{c}=a_{c}$, and, by 7,

$$
a^{-1}=\frac{a_{c}}{(\mathrm{~T} a)^{2}}=\frac{\mathrm{I}}{\mathrm{~T} a} \cdot \mathrm{U} a_{c}=\frac{\mathrm{I}}{\mathrm{~T} a} \cdot e^{-\alpha \mathrm{a}} .
$$

Thus, the reciprocal of a quaternion is its conjugate divided by its norm. In other words, the reciprocal of $a$ has the reciprocal tensor, the opposite axis and the same angle as $\alpha$.

Consequently, we can also write

$$
a a^{-1}=\mathrm{r}
$$

Thus, if we have an equation

$$
a m=b
$$

and we wish to isolate $m$, we have only to multiply both sides by $a^{-1}$ as prefactor, obtaining

$$
\begin{array}{ll}
\text { pretactor, odtaining } & m=a^{-1} b . \\
\text { Similarly, if } & n c z=b, \\
\text { we shall have } & n=b a^{-1} .
\end{array}
$$

Notice, in particular, that the reciprocal of a unit quaternion is at the same time its conjugate.
10. The differentiation of quaternions, with respect to time or position in space, does not require any explanations. The definition of 'curl' and 'div' being supposed known from usual vector-analysis, it will be enough to remember here what was said already in Chap. II., namely that $\nabla$, the vector part of our $D$, when applied to a vector A, gives

$$
\nabla \mathrm{A}=\mathrm{V}_{\nabla \mathrm{A}}+S_{\nabla \mathrm{A}}=V_{\nabla \mathrm{A}}-(\nabla \mathrm{A})
$$

(as in 4 , because $\nabla$ apart from its differentiating properties is to be treated as an ordinary vector), or ultimately

$$
\nabla \mathrm{A}=\operatorname{curl} \mathbf{A}-\operatorname{div} \mathbf{A} .
$$

For all of our purposes we shall hardly want more than is given in the above ten sections,-which will in the sequel be shortly referred to as 'Quat. 1, 2, etc.'

Returning to our position-quaternion $q$, let us write its $S^{\prime}$-correspondent, or the transformed quaternion

$$
q^{\prime}=l^{\prime}+\mathbf{r}^{\prime}
$$

Since $l=\iota c t$, i.e. $t=-\downarrow l / c$, we have, by ( $1 b$ ), p. 124, and denoting now the unit of $\mathbf{v}$ by $\mathbf{u}$,

$$
\left.\begin{array}{l}
l^{\prime}=\gamma[l-\iota \beta(\mathbf{u r})]  \tag{48}\\
\mathbf{r}^{\prime}=\epsilon \mathbf{r}+\iota \beta \gamma \mathbf{u} .
\end{array}\right\}
$$

Here, it will be remembered, $\epsilon$ is the longitudinal stretcher, whose developed form is, by (2),

$$
\epsilon=\mathrm{r}+(\gamma-\mathrm{r}) \mathrm{u}(\mathrm{u} .
$$

Now, such being the scalar and the vector parts of $q^{\prime}$ in terms of those of $q$, we can easily find a quaternion $Q$ such that

$$
\begin{equation*}
q^{\prime}=Q q Q . * \tag{49}
\end{equation*}
$$

First of all, since we know that $l^{2}+r^{2}$ is an invariant, or that $\mathrm{T} q^{\prime}=\mathrm{T} q$, we can at once take for $Q$ a unit quaternion,

$$
Q=\cos \theta+a \cdot \sin \theta
$$

Thus, we have only to find the angle and the axis of $Q$ in terms of $\beta$ and $u$. Now, developing the triple product in (49), we obtain easily, by Quat. 6,

$$
\begin{aligned}
& l^{\prime} \equiv S Q_{q} Q=\cos (2 \theta) \cdot l-\sin (2 \theta) \cdot(\mathrm{ar}), \\
& \mathrm{r}^{\prime} \equiv \mathrm{V} Q_{q} Q=\mathrm{r}-2 \sin ^{2} \theta \cdot \mathrm{a}(\mathrm{ar})+\sin (2 \theta) \cdot l \mathrm{a},
\end{aligned}
$$

whence, comparing with (48),

$$
a=u ; \quad \cos (2 \theta)=\gamma ; \quad \sin (2 \theta)=\iota \beta \gamma,
$$

and $r-2 \sin ^{2} \theta \cdot a\left(a=\epsilon\right.$, i.e. $2 \sin ^{2} \theta=r-\gamma$, which is identical with the third of the above equations, and this, again, says the same thing as the second. Thus, all conditions are satisfied at once, and we have ultimately

$$
\mathrm{a}=\mathrm{u} \quad \text { and } \quad \theta=\frac{1}{2} \arctan (\iota \beta)=\frac{1}{2} \omega,
$$

where $\omega$ is the (imaginary) angle of rotation, as previously defined. [Cf. (10), p. 127.] To resume:

The position-quaternion $q$ is transformed by the operator

$$
Q[] Q
$$

the vacant place being destined for the operand.
The axis of the unit quaternion $Q$ is $\mathfrak{u}$, the unit of $\mathbf{v}$, and its angle is half that of Minkozeski's imaginary angle of rotation, i.e.

$$
\begin{equation*}
Q=\cos \frac{\omega}{2}+\mathbf{u} \cdot \sin \frac{\omega}{2}=e^{\frac{\omega}{2}} \mathbf{u} . \tag{50}
\end{equation*}
$$

* As regards the reason why particularly this form, involving a quaternionic prefactor and postfactor, is sought for, see my paper in Phil. Mag., Vol. XXIII., quoted before, where I gave references going back to Cayley's original discovery (1854).

Another form of this quaternion is

$$
\begin{equation*}
Q=\left(\frac{I+\gamma}{2}\right)^{\frac{1}{2}}+\mathbf{u} \cdot\left(\frac{I-\gamma}{2}\right)^{\frac{1}{2}} \tag{50a}
\end{equation*}
$$

Observe that, $\gamma$ being $>\mathrm{I}$, the vector part of $Q$ is imaginary, while its scalar part is real.
Since $Q$ is a unit quaternion, we have $Q^{-1}=Q_{c}$, or

$$
Q Q_{c}=Q_{c} Q=I
$$

a property which we shall constantly use. Thus, to obtain from (49) the inverse transformation, multiply both sides by $Q_{c}$ as a post- and a prefactor. Then the result will be

$$
\begin{equation*}
q=Q_{c} q^{\prime} Q_{c} \tag{49a}
\end{equation*}
$$

as it should be, since $Q_{0}$ is obtained from $Q$ by a reversal of $u$ or $\nabla$. Again, to see once more, or to verify, the invariance of

$$
\begin{equation*}
(\mathrm{T} q)^{2}=q q_{c}=r^{2}+l^{2}=r^{2}-c^{2} t^{2} \tag{51}
\end{equation*}
$$

take the conjugate of (49), which, by Quat. 8, is

$$
q^{\prime}=Q_{0} q_{0} Q_{c} .
$$

Now, by the same formula (49), and by the associative law,

$$
q^{\prime} q_{c}^{\prime}=Q q Q Q_{c} q_{c} Q_{c}=Q q q_{c} Q_{c} .
$$

But, since $q q_{c}$ is a scalar, it may be written before the $Q$, or if you wish, after the $Q_{c}$, so that

$$
q^{\prime} q_{0}^{\prime}=q q_{0} Q Q_{c}=q q_{0}
$$

(2.E.D.

We shall see later on, when we come to consider products of two, or more, of such quaternions, that they are transformed with equal ease.

Consecutive transformations assume the following simple form. Let $\nabla_{1}=v_{1} \mathbf{u}_{1}$ be the velocity of $S^{\prime}$ relative to $S$, and $\nabla_{2}=\tau_{1} \mathbf{u}_{2}$ the velocity of $S^{\prime \prime}$ relative to $S^{\prime}$, and let $Q_{1}, Q_{2}$ be the corresponcing transforming quaternions, i.e.

$$
Q_{1}=e^{\frac{\omega_{1}}{2} u_{1}}, \quad Q_{2}=e^{\frac{\omega_{2}}{2} u_{2}}
$$

Then
and

$$
q^{\prime}=Q_{1} q Q_{1}
$$

$$
q^{\prime \prime}=Q_{2} q^{\prime} Q_{2}=Q_{2} Q_{1} q Q_{1} Q_{2}
$$

so that the compound transformer is

$$
Q_{2} Q_{1}[] Q_{1} Q_{2}
$$

In general, for non-parallel axes $\mathrm{u}_{1}, \mathrm{u}_{2}$,

$$
Q_{2} Q_{1} \neq Q_{1} Q_{2}
$$

so that the compound transformer has not the form $Q[] Q$. This is but the quaternionic expression of the fact, to be considered fully in the following chapter, that a pair of consecutive three-parametric Lorentz transformations, (48), does not generally give again such a transformation, but is equivalent to (48) combined with a pure rotation in ordinary three-dimensional space. In other words, the transformations (48) do not constitute a group. But, as we saw before, they contain sub-groups, namely for parallel velocities. Then, and only then, $Q_{2} Q_{1}$ becomes equal to $Q_{1} Q_{2}$, and the compound transformer assumes the form $Q[\quad] Q$. Suppose, for instance, that the velocities $\mathbf{v}_{1}, \nabla_{2}$, being parallel, are also concurrent with one another, i.e. that

Then

$$
\mathbf{u}_{1}=\mathbf{u}_{2}=\mathbf{u}
$$

$$
Q=e^{\frac{\omega}{2}} \mathbf{u}=e^{\frac{1}{2}\left(\omega_{1}+\omega_{2}\right) \mathbf{u}},
$$

so that the previous formula for the composition of parallel velocities, $\omega=\omega_{1}+\omega_{2}$, follows from the quaternionic form immediately.

Imitating the name 'world-vector,' we could now call $q$, or $q_{c}$, the standard of world-quaternions. But the more modest name physical quaternion will do as well. Also, to begin with, no further specification of the 'kind' is needed. But it may be convenient to have a pair of short symbols, in order to compare any quaternions with respect to their relativistic behaviour. By writing

$$
X \sim q,
$$

we shall understand that the quaternion $X$ is covariant or, equally transformed, with $q$, i.e. that

$$
X^{\prime}=Q X Q
$$

without taking into account the structure of $X$. And if $X$ has also the structure of $q$, that is to say, if it has a purely imaginary scalar and a real vector,* then we shall write

$$
x \simeq q
$$

The latter will then be equivalent to saying that $X$ is a physical quaternion, viz. covariant with $q$. This being the case, the conjugate of $X$ will, of course, be also a physical quaternion, e.g.

$$
X_{c} \simeq q_{c} .
$$

The same notation we shall extend to quaternionic operators. Thus, as we saw, $\partial / \partial l$ and $\nabla$, the scalar and the vector parts of the operator $D$, are transformed like $l, \mathbf{r}$, the scalar and the vector parts of the position-quaternion, i.e.

$$
\begin{equation*}
D^{\prime}=Q D Q \tag{52}
\end{equation*}
$$

and similarly, $D_{c}=\partial / \partial l-\nabla$ being the conjugate operator,

$$
\begin{equation*}
D_{c}^{\prime}=Q_{c} D_{c} Q_{0} \tag{52c}
\end{equation*}
$$

But $D$ has also the same structure as $q$. Consequently, apart from its differentiating properties, $D$ behaves as a genuine physical quaternion, or

$$
D \simeq q .
$$

Analogously to Minkowski's classification of four-vectors, we may call any physical quaternion $X$ a space-1ike, or a time-iike, or finally a singular quaternion, according as its norm, $(\mathrm{T} X)^{2}=X X_{0}$, is positive, or negative, or zero.
But it does not seem desirable to dwell any longer upon the formal side of the subject until our stock of materials has. been somewhat enlarged. For as yet we have only one physical quaternion, namely $q$.

[^42]
## NOTE TO CHAPTER $V$.

(To page 143.) A matrix is any rectangular array of magnitudes or, more generally, of symbols cither of magnitude or of operation, each of which has its assigned place, i.c. belongs to a given row and a given column. Thus

$$
A=\left|\begin{array}{cccc}
a_{11}, & a_{12}, & \ldots & a_{1 n} \\
a_{21}, & a_{22}, & \ldots & a_{2 n} \\
\vdots & \vdots & & \vdots \\
a_{m 1}, & a_{m 2}, & \ldots & a_{m n}
\end{array}\right|
$$

is a matrix of $m$ roous and $n$ columns. The first index of any constituent $a_{t k}$ denotes the row, and the second the column to which it belongs.

The matrix whose rows are the columns of $A$ is called the transposed of $A$, and is denoted by $\bar{A}$. Thus, $A$ being as above,

$$
\dot{A}=\left|\begin{array}{cccc}
a_{11}, & a_{21}, & \ldots & a_{m 1} \\
a_{12}, & a_{2 n}, & \ldots & a_{m 2} \\
\vdots & \vdots & & \vdots \\
a_{1 n}, & a_{2 n}, & \ldots & a_{m n}
\end{array}\right|
$$

To specify the number of rows and columns of a matrix we may conveniently attach to its symbol a pair of indices. Thus, $A$ will be $A_{m n}$, and similarly $\mathcal{A}=\bar{A}_{m n}$. Or we may say, equivalently, that $A$ is a matrix of $m \times n$ constituents, and $\bar{A}$ a matrix of $n \times m$ constituents.

If we have a pair of matrices $A=A_{m n}$ and $B=B_{m n}$, then the matrix $C=C_{m n}$, whose constituents are sums of the corresponding constituents of $A, B$ (i.c. $\left.c_{i \kappa}=a_{l \kappa}+b_{\iota k}\right)$, is written

$$
C=A+B
$$

If, in particular, $B=A$, the result of acldition is written $2 A$, and so on. (ienerally, if $a$ be any number (or symbol of operation) and $A$ any matrix, then the matrix $C$, whose constituents are $c_{L K}=\alpha u_{L \kappa}$, is called the product of $\alpha$ into $A$, and is denoted by $a A$.

If the matrix $B$ has as muny rows as $A$ luas columns, i.e. if

$$
A=A_{m n}, \quad B=B_{n p}
$$

(where $p$ may be equal to or different from $m$ ), then the matrix $C$, of which any constituent $c_{l K}$ is equal to the sum of the products of the constituents of the $\iota$ th row of $A$ into those of the $\kappa$ th column of $B$, is called the product of $A$ into $B$, and is written

$$
C=A B
$$

Thus, if $A$ is as above, and if

$$
B=\left|\begin{array}{ccc}
b_{11}, & b_{12}, & \ldots b_{1 p} \\
b_{21}, & b_{22}, & \ldots b_{2 p} \\
\vdots & \vdots \\
b_{n 1}, & b_{n 2}, & \ldots b_{n p}
\end{array}\right|,
$$

then
where

$$
C=C_{m 1}=A_{m n} B_{n 1}=\left|\begin{array}{cccc}
c_{11}, & c_{12}, & \ldots & c_{11} \\
c_{21}, & c_{22}, & \ldots & c_{22} \\
\vdots & \vdots & & \vdots \\
c_{n 1}, & c_{m 2}, & \ldots & c_{m 1}
\end{array}\right|
$$

$$
\begin{aligned}
& c_{11}=a_{11} b_{11}+a_{12} b_{21}+\ldots+a_{1 n} b_{n 1} \\
& c_{12}=a_{11} b_{12}+a_{12} b_{22}+\ldots+a_{1 n} b_{n 2}
\end{aligned}
$$

and so on, generally

$$
c_{\mathrm{\iota K}}=a_{\mathrm{L} 1} b_{1 \kappa}+a_{\mathrm{L} 2} b_{2 \kappa}+\ldots+a_{\mathrm{tn}} b_{n \kappa}
$$

Notice that, if $p \neq m$, the expression $B A$ would be meaningless. But, since $\bar{B}=\bar{B}_{m}$ and $\bar{A}=\bar{A}_{n m}$, we can have the product $\bar{B} \bar{A}$, which will be a matrix of $p \times m$ constituents. This, as can easily be seen, will be the transposed of $A B$, i.e.

$$
\overline{A B}=\bar{B} \bar{A} .
$$

Compare this property with Quat. 8. p. 154.
Since $A B=C_{m p}$, it can be multiplied into a third matrix $\eta=I I_{1 m}$, thus giving rise to $A B . D$, which will be a matrix of $m \times q$ constituents. It can be proved that for such products the associative lezv holds (supposing, of course, that the constituents themselves, which generally can be operators, obey this law), i.c.

$$
A B \cdot D=A \cdot B D .
$$

Hence, both sides may be simply written $A B D$. The same property belongs to the product of any number of matrices. Thus

$$
A_{m n} B_{n p} D_{p q} \ldots M_{x y} N_{p z z}=R_{n n z}
$$

will be a definite matrix of $m \times z$ constituents, independent of the grouping of the factors. Notice the analogy with quaternionic products.
Let each of the constituents of the principal diagonal (from left uppermost to right lowermost) of a square matrix $U=U_{m a}$ be equal r , i.e.

$$
u_{11}=u_{22}=\ldots=u_{n n}=1
$$

and let each of its remaining constituents $u^{\text {*K }}$ be zero. Then, if $M T$ be any matrix of $n$ rows,
s.r.
$U M=M$.
L

In view of this property, $U$ is called a unit-matrix, and may be simply denoted by I.

Now, let $M$ be any square matrix. Then the determinant formed of its constituents is called the determinant of $M$, and is shortly written det $M$. Suppose that det $M$ does not vanish. Then there exists a definite matrix which, multiplied into $M$, gives a unit matrix or simply 'unity.' This matrix is called the reciprocal of $M$, and is denoted by $M^{-1}$. The above definition is written shortly

$$
M^{-1} M=\mathrm{I}
$$

where $I$ stands for $U_{n n}$. The reciprocal is, of course, as $M$ itself, a square matrix of $n \times n$ constituents.

Other particulars concerning matrices will be given incidentally, as the need arises in the subject under consideration.

## CHAPTER VI.

## COMPOSITION OF VELOCITIES AND THE LORENTZ GROUP.

Consider a particle moving about in an arbitrary manner in the system $S^{\prime}$, which in its turn moves with uniform velocity $\mathbf{v}$ relatively to the system $S$. Let $\mathbf{p}^{\prime}$ be the instantaneous velocity of the particlc from the point of view of the $S^{\prime}$-observers, i.e. let at the instant $t^{\prime}$

$$
\frac{d \mathbf{r}^{\prime}}{d t^{\prime}}=\mathrm{p}^{\prime} .
$$

What is the velocity $\mathbf{p}$ of this particle from the $S$-standpoint, at the instant $t$ corresponding to $t^{\prime}$ ?
To answer this simple but very fundamental question of relativistic: kinematics, use the form ( $\mathrm{I} b$ ), Chap. V., of the Lorentz transformation. Then its inverse will be, as in ( $\left(b^{\prime}\right)$,

$$
\begin{aligned}
& \mathbf{r}=\epsilon \mathbf{r}^{\prime}+\gamma \mathbf{v} t^{\prime}, \\
& t=\gamma\left[t^{\prime}+\frac{1}{c^{2}}\left(\mathbf{v \mathbf { r } ^ { \prime }}\right)\right],
\end{aligned}
$$

and, since $d \epsilon \mathbf{r}^{\prime}=\epsilon d \mathrm{r}^{\prime}$ and $d\left(\mathrm{vr}^{\prime}\right)=\left(\mathbf{v} d \mathbf{r}^{\prime}\right)$,

$$
\mathrm{p}=\frac{d \mathbf{r}}{d t}=\frac{\epsilon d \mathbf{r}^{\prime}+\gamma \mathbf{v} d t^{\prime}}{\gamma\left[d t^{\prime}+\frac{\mathrm{T}}{c^{2}}\left(\mathbf{v} d \mathbf{r}^{\prime}\right)\right]} .
$$

Divide the numerator and the denominator on the right by $d t^{\prime}$, and remember the meaning of $\mathbf{p}^{\prime}$. Then the required velocity will follow at once under the simple form

$$
\begin{equation*}
\mathbf{p}=\frac{\gamma \boldsymbol{\nabla}+\epsilon \mathrm{p}^{\prime}}{\gamma\left[\mathbf{x}+\frac{\mathrm{I}}{c^{2}}\left(\boldsymbol{\mathrm { p }} \mathbf{p}^{\prime}\right)\right]} \tag{t}
\end{equation*}
$$

This is the vectorial expression of Einstein's famous Addition Theorem.*
As before, $\gamma=\left(x-\beta^{2}\right)^{-\frac{1}{2}}, \beta=v / c$, and $\epsilon$ is the longitudinal stretcher of ratio $\gamma$. Thus, in Cartesians, with $x$ measured along $\nabla$, ( $\mathrm{I} a$ ) will become

But having explained this for the non-vectorial reader, we shall henceforth use the above short formula.
By writing $\mathbf{p}^{\prime}, \mathbf{p}$ we wished to emphasize that the latter is the $S$-correspondent of the Cormer. But we may as well look at $\mathbf{p}$ as the resullant of $\mathbf{v}$ and $\mathbf{p}^{\prime}$, keeping in mind that the first of these component velocities is taken relatively to one system, $S$, and the second relatively to another $\dagger$ system $S^{\prime \prime}$. Then it may be more convenient to write for the velocitics to be compounded $\mathbf{v}_{1}, \boldsymbol{\nabla}_{2}$ (instead of $\boldsymbol{\nabla}, \mathbf{p}^{\prime}$ ), and for the resultant velocity $\boldsymbol{\nabla}$ (instead of $\mathbf{p}$ ). Thus, attaching the correspondent index to $\gamma$ and $\epsilon$, we shall write

$$
\begin{equation*}
\mathbf{v}=\frac{\gamma_{1} \boldsymbol{v}_{1}+\epsilon_{1} \boldsymbol{v}_{2}}{\gamma_{1}\left[I+\frac{I}{c^{2}}\left(\boldsymbol{v}_{1} \boldsymbol{v}_{2}\right)\right]} . \tag{I}
\end{equation*}
$$

Notice that the resultant is, in general, anon-symmetrical function of the two component velocitics. It is important to know which of these comes first, and which next. In Newtonian or classical kinematics the resultant is simply $\mathbf{v}_{1}$ plus $\nabla_{2}$ and at the same time $\mathbf{v}_{2}$ plus $\mathbf{v}_{1}$. Here the case is different. We may still speak of 'addition,' as a non-pedantic synonym of composition of velocities, but to avoid confusion we should employ instead of the ordinary + another symbol, say $\#$, and write the above $\mathbf{v}$, as given by ( I ),

$$
\boldsymbol{v}_{1} \# \boldsymbol{v}_{2} .
$$

[^43]Then the resultant of $\mathbf{V}_{2}$ and $\nabla_{1}$ (i.e. the $S$-velocity of a particle moving with velocity $\mathbf{v}_{1}$ relative to $S^{\prime}$, which in its turn moves with velocity $\nabla_{2}$ relative to $S$ ) would be

$$
\begin{equation*}
\mathbf{v}_{2} \# . \boldsymbol{v}_{1}=-\frac{\gamma_{2} \mathbf{v}_{2}+\epsilon_{2} \mathbf{v}_{1}}{\gamma_{2}\left[\mathrm{I}+\frac{\mathrm{I}}{c^{2}}\left(\mathbf{v}_{1} \mathbf{v}_{2}\right)\right]} \tag{2}
\end{equation*}
$$

where $\epsilon_{2}$ is a stretcher acting along $\boldsymbol{\nabla}_{2}$, of ratio $\gamma_{2}$.
In short, the relativistic composition of velocities is, generally speaking, non-commutative.

But it is interesting and, in view of what has to come later, uscful to notice that the two vectors ( I ), ( 2 ), though differing in direction, are identical in their absolute magnitude. To see this, we have only to prove that the squares of the two vectors
and

$$
a=\boldsymbol{v}_{1}+\frac{I}{\gamma_{1}} \epsilon_{1} \mathbf{v}_{2}
$$

$$
\mathrm{b}=\mathbf{v}_{\mathbf{2}}+\frac{\underline{r}}{\gamma_{2}} \epsilon_{2} \mathbf{v}_{1}
$$

are equal to one another. Now, by the elementary rules of vector algebra,

$$
a^{2}=v_{1}^{2}+\frac{2}{\gamma_{1}}\left(\nabla_{1} \cdot \epsilon_{1} \nabla_{2}\right)+\frac{I}{\gamma_{1}^{2}}\left(\epsilon_{1} \nabla_{2}\right)^{2}
$$

and, since $\epsilon_{1}$ is a symmetrical vector-operator,

$$
\left(\boldsymbol{v}_{1} \cdot \epsilon_{1} \nabla_{2}\right)=\left(\epsilon_{1} \nabla_{1} \cdot \nabla_{2}\right)=\gamma_{1}\left(\nabla_{1} \nabla_{2}\right)
$$

Again, denoting by $\theta$ the angle between $\nabla_{l}$ and $\nabla_{2}$,

$$
\left(\frac{I}{\gamma_{1}} \epsilon_{1} \nabla_{2}\right)^{2}=v_{2}^{2}\left[\cos ^{2} \theta+\frac{I}{\gamma_{1}^{2}} \sin ^{2} \theta\right]=v_{2}^{2}\left[I-\beta_{1}^{2} \sin ^{2} \theta\right] .
$$

## Hence

$$
a^{2}=v_{1}^{2}+v_{2}^{2}+2\left(\mathbf{v}_{1} \mathbf{v}_{2}\right)-\frac{1}{c^{2}} v_{1}^{2} v_{2}^{2} \sin ^{2} \theta=\left(\mathbf{v}_{1}+\mathbf{v}_{2}\right)^{2}-\frac{\mathrm{r}}{c^{2}}\left(V \mathbf{v}_{1} \mathbf{v}_{2}\right)^{2},
$$

and this, being a symmetrical function of $\mathbf{v}_{1}, \mathbf{v}_{2}$, is at the same time the value of $b^{2}$. Q.E.D.

Thus we have the general property of relativistic composition of velocities:

$$
\begin{equation*}
\left(\mathbf{v}_{1} \# \mathbf{v}_{2}\right)^{2}=\left(\mathbf{v}_{2} \# \mathbf{v}_{1}\right)^{2} . \tag{3}
\end{equation*}
$$

The common value of these scalars is, by ( r ) and by the formula just found for $a^{2}$,

$$
=\begin{gather*}
\left(\mathbf{v}_{1}+\nabla_{12}\right)^{2} \cdot \frac{1}{c^{2}}\left(V_{v_{2}} \mathbf{v}_{2}\right)^{2} \\
\left.1.1+\frac{1}{c^{2}}\left(\mathbf{v}_{1} \mathbf{v}_{2}\right)\right]^{2} \tag{4}
\end{gather*}
$$

This is linustin's fimmus formula for the square of the resultant Mocity, written vectorially.

Hefore masing to give a few examples and a certain very remarkalble geometric: representation of the Addition Theorem ( r , let us aproach the question of composition of velocities from anowher side, cos. by considering a pair of consecutive Lorentz tramsformations.
lart aguin $v_{1}$ be the velocity of $S^{\prime}$ relative to $S$, but instead of 'ur particle take a thirel system $S^{\prime \prime \prime}$ moving relatively to $S^{\prime}$ with the Wheity $v_{n}$, the former velocity being taken from the $S$-standpoint and the hatter from the $S^{\prime \prime}-p$ oint of view, both being now uniform tranhlational veloxitics. Lat $\gamma_{1}, \epsilon_{1}$ and $\gamma_{2}$, $\epsilon_{2}$ be the corresponding moanings of $\gamma$ e. 'Then, $t$ ', $r$ ' loing the time and the space-vector in $S^{\prime \prime}$, mad $t^{\prime \prime}, x^{\prime \prime}$ the time and the space-vector in $S^{\prime \prime}$, we shall have, by ( 10 ), (laty. V.,

$$
\begin{equation*}
\mathbf{r}^{\prime} \quad \epsilon_{1} \mathbf{r} \cdots \mathbf{v}_{1} \gamma_{1} t ; \quad t^{\prime}=\gamma_{1}\left[t-c_{c^{2}}^{\mathrm{r}}\left(\mathbf{v}_{1} \mathbf{r}\right)\right] \tag{1}
\end{equation*}
$$

athl

$$
\begin{equation*}
r^{\prime \prime} \quad \epsilon_{a} r^{\prime} \cdots \nabla_{a} \gamma_{a} t^{\prime} ; t^{\prime \prime}=\gamma_{2}\left[t^{\prime}-\frac{1}{c^{2}}\left(\boldsymbol{v}_{3} r^{\prime}\right)\right] . \tag{2}
\end{equation*}
$$

Intruluce the values ( $5_{1}$ ) of $r^{\prime}$ and $t^{\prime}$ into ( $5_{2}$ ), and remember that, it being a symmetrical vector operator, $\left(\boldsymbol{\nabla}_{2}, \epsilon_{1} r\right)=\left(\epsilon_{1} \nabla_{2}, r\right)$. Then the result will low

The ferent, transformations hitherto considered, of which ( $5_{1}$ ) and (5.) are individual cases, involve three scalar parameters ( $P_{0}, q_{q}, \theta_{0}$ ) or une vectorial parameter $\mathbf{v}$. Let us therefore denote any ome: of these transformations by $L(v)$. Thus, the two above compenent transformations will $\ln \ell\left(\boldsymbol{v}_{1}\right), L\left(\boldsymbol{v}_{2}\right)$, and their resultant,
i.e. the first followed by the second, or the transformation (6), may be written $L\left(\boldsymbol{v}_{2}\right) L\left(\mathbf{v}_{1}\right)$.

We know that any $L(\mathbf{v})$ leaves invariant the quadratic expression

$$
x^{2}+y^{2}+z^{2}+l^{2},
$$

and can therefore be considered as a rotation in the four-dimensional world. But it is not the most general rotation, since it does not include the rotation round the time-axis, i.e. a rotation of the spaceframework, or an equivalent rotation of the three-dimensional vectors. If any transformation $\mathcal{L}(\mathbf{v})$ is followed by such a rotation of $\mathbf{r}^{\prime}$, which does not change the value of $r^{\prime 2}=x^{\prime 2}+y^{\prime 2}+z^{\prime 2}$, then the: above quadratic expression will, obviously, continue to be an invariant. Let $\Omega$ be a purely rotating operator, or what (iibls** called a 'versor,' i.e. such a linear vector operator that, for any vector $\mathbf{R}$,

$$
(\Omega \mathbf{R})^{2}=R^{2} .
$$

Then the amplified or, as it is sometimes called, the general Lorentz transformation will be given by

$$
\begin{align*}
& \mathbf{r}^{\prime}=\Omega[\epsilon \mathrm{r}-\mathrm{v} \gamma t] \\
& t^{\prime}=\gamma\left[t-\frac{\mathrm{I}}{c^{2}}(\mathrm{vr})\right] .
\end{align*}
$$

Since $\Omega$ involves three scalar data, viz. one for its angle and $\mathfrak{t w o}$ for its axis, $L(\mathbf{v}, \Omega)$ will be a six-parametric transformation. Thas, the above symbol $L(v)$ of the special Lorentz transformation stands for $L(\mathbf{v}, \mathrm{r})$. Notice that the scalar product of two vectors, e.g. (vr), is not changed at all by a pure space-rotation. This is the reasm that $\Omega$ does not enter into the expression for $t^{\prime}$, and would not enter into it even if the rotation preceded the special Lorentz transformation.

Let us now return to our $L\left(\boldsymbol{v}_{2}\right) L\left(\boldsymbol{v}_{1}\right)$, as given by the formulae ( 6 ).
We have seen in the last chapter that, if the velocitics $\nabla_{1}$ and $\nabla_{3}$ are parallel to one another, the resultant transformation is again it special Lorentz transformation, i.e.

$$
L\left(\mathbf{v}_{2}\right) L\left(\mathbf{v}_{1}\right)=L(\mathbf{v}),
$$

where $\boldsymbol{\nabla}\left\|\boldsymbol{v}_{1}\right\| \boldsymbol{\nabla}_{2}$. Now, it can easily be shown that this is the calst ouly for $\mathbf{v}_{1} \| \mathbf{v}_{\mathbf{2}}$.
*J. Willard Gibbs, Scientific Papers, Vol. II. p. 64.

In fact, suppose that (6) is an $L(\boldsymbol{v})$, that is to say, suppose there is a vector $\boldsymbol{\nabla}$ (with the corresponding $\gamma$ and $\epsilon$ ), such th

$$
\mathbf{r}^{\prime \prime}=\epsilon \mathbf{r}-\mathbf{\nabla} \gamma t ; \quad t^{\prime \prime}=\gamma\left[t-\frac{\mathbf{I}}{c^{2}}(\mathbf{\nabla} \mathbf{r})\right] .
$$

Then, remembering that this has to coincide with (6) for eve (as well as for every $t$ ) and taking, for instance, $r=0$, you obtain, from the first of (6),

$$
\gamma \mathbf{\nabla}=\gamma_{1}\left[\epsilon_{2} \nabla_{1}+\gamma_{2} \nabla_{2}\right],
$$

and at the same time, from the second of (6),

$$
\gamma \mathbf{\nabla}=\gamma_{2}\left[\epsilon_{1} \nabla_{2}+\gamma_{1} \nabla_{1}\right],
$$

and, consequently,

$$
\gamma_{1}\left[\epsilon_{2} \nabla_{1}+\gamma_{2} \nabla_{2}\right]=\gamma_{2}\left[\epsilon_{1} \nabla_{2}+\gamma_{1} \nabla_{1}\right] .
$$

Now, this equation cannot be satisfied unless $\boldsymbol{\nabla}_{1}$ and $\boldsymbol{\nabla}_{2}$ are pare To see this, call $l_{1}$ and $n_{1}$ the parts of $\nabla_{1}$ taken along and no to $\boldsymbol{v}_{2}$, and similarly $\mathbf{l}_{2}$ and $\mathrm{n}_{2}$ the parts of $\mathbf{v}_{2}$ taken along and no to $\nabla_{1}$, and write $\nabla_{1}=l_{1}+n_{1}, \nabla_{2}=l_{2}+n_{2}$. Then, remembering $\epsilon_{1}, \epsilon_{2}$ are longitudinal stretchers, the above equation will assume form

$$
\gamma_{1}\left[\gamma_{2} l_{1}+n_{1}+\gamma_{2} l_{2}+\gamma_{2} n_{2}\right]=\gamma_{2}\left[\gamma_{1} 1_{2}+n_{2}+\gamma_{1} l_{1}+\gamma_{1} n_{1}\right]
$$

or

$$
\gamma_{1}\left(r-\gamma_{2}\right) \mathbf{n}_{1}=\gamma_{2}\left(r-\gamma_{1}\right) \mathbf{n}_{2} .
$$

Hence, either $\gamma_{1}=\gamma_{2}=\mathrm{r}$, which corresponds to the trivial $v_{1}=v_{2}=0$, or $\mathrm{n}_{1} \| \mathrm{n}_{2}$, and consequently also $\mathrm{\nabla}_{1} \| \mathrm{\nabla}_{2}$. Q.E.D.

Thus, if $\boldsymbol{\nabla}_{1}$ and $\boldsymbol{\nabla}_{2}$ are not parallel to one another, the resu transformation (6) is not an $L(\mathbf{v})$. In other words, the class of transformations $L(v)$ does not constitute a group, although it con one-parametric subgroups, each ranging over parallel velocities.

But the six-parametric transformations $L(\mathbf{v}, \Omega)$ do constitu group, i.e.

$$
\mathcal{L}\left(\mathbf{v}_{2}, \Omega_{2}\right) L\left(\mathbf{v}_{1}, \Omega_{1}\right)=L(\mathbf{\nabla}, \Omega)
$$

for any pair of velocities and any pair of versors, and hence particular, also for $\Omega_{1}=\mathrm{r}, \Omega_{2}=\mathrm{r}$, as in our case. For non-par velocities, then, our $L\left(\mathbf{v}_{2}\right) L\left(\mathbf{v}_{1}\right)$ is not again an $L(\mathbf{v})$, but it i

## RESULTANT TRANSFORMATION

$L(\nabla, \Omega)$ with a certain space-rotation,* to be determined. In the formulae (6) are of the form

$$
\begin{aligned}
\mathbf{r}^{\prime \prime} & =\Omega[\epsilon \mathbf{r}-\mathbf{\nabla} \gamma t]=\Omega \epsilon r-\gamma t \Omega \mathbf{v} \\
t^{\prime \prime} & =\gamma\left[t-\frac{\mathrm{I}}{c^{2}}(\mathbf{\nabla r})\right]
\end{aligned}
$$

where $\Omega \neq \mathrm{I}$.
A comparison with (6) will give us the four equations

$$
\begin{aligned}
& \Omega \epsilon=\epsilon_{2} \epsilon_{1}+\frac{\gamma_{1} \gamma_{2}}{c^{2}} \nabla_{2}\left(\nabla_{1}\right. \\
& \gamma=\gamma_{1} \gamma_{2}\left[\mathrm{I}+\frac{\mathrm{r}}{c^{2}}\left(\nabla_{1} \nabla_{2}\right)\right] \\
& \Omega \boldsymbol{v}=\frac{\gamma_{1}}{\gamma}\left[\epsilon_{2} \nabla_{1}+\gamma_{2} \mathbf{\nabla}_{2}\right] \\
& \boldsymbol{\nabla}=\frac{\gamma_{2}}{\gamma}\left[\epsilon_{1} \nabla_{2}+\gamma_{1} \nabla_{1}\right] .
\end{aligned}
$$

From (b), (d) we have at once the resultant velocity, of $S^{\prime \prime} \mathrm{r}$ to $S$,

$$
\mathbf{v}=\mathbf{v}_{1} \# \boldsymbol{v}_{2,2}=\frac{\gamma_{1} \boldsymbol{v}_{1}+\epsilon_{1} \boldsymbol{v}_{2}}{\gamma_{1}\left[\mathrm{r}+\frac{\mathrm{r}}{c^{2}}\left(\boldsymbol{\nabla}_{1} \boldsymbol{v}_{2}\right)\right]},
$$

identical with ( r ), which was obtained by differentiation. verification that $\gamma$, as given by $(b)$, is equal to $\left(\mathrm{r}-v^{2} / c^{2}\right)^{-\frac{1}{2}}$, to the reader. Again, the right side of $(c)$ is what $\nabla$ becor permutation of ${ }_{1,2}$, so that

$$
\Omega \mathbf{v}=\Omega\left[\mathbf{v}_{1} \# \mathbf{v}_{2}\right]=\mathbf{v}_{2} \# \mathbf{v}_{1},
$$

and this agrees with the nature of the operator $\Omega$. For, shown explicitly, the tensors of the two resultant velocities are cf. (3). Thus, $\Omega$ turns $\mathbf{v}_{1} \# \mathbf{\nabla}_{2}$ into $\nabla_{2} \# \nabla_{1}$. The equati

* In four-dimensional language the case under consideration may be e as follows. Call $t$ the time-axis in Minkowski's world. Then $L\left(\nabla_{1}\right)$
of course, does not by itself suffice for a complete det the operator, for it states the result of its application vector $\boldsymbol{\nabla}$ only. But we have still ( $a$ ), which is valid for as operand, i.e.

$$
\Omega \epsilon r=\epsilon_{2} \epsilon_{1} r+\frac{\gamma_{1} \gamma_{2}}{c^{2}} \nabla_{2}\left(\nabla_{1} r\right) .
$$

As to $\epsilon$, the reader may verify that none of the above fo is contradicted by assuming it to be a longitudinal str sponding to $\mathbf{\nabla}$, i.e. by writing, for any $\mathbf{r}$,

$$
\epsilon r=r+\frac{\gamma-\mathrm{I}}{v^{2}} \nabla(\nabla r) .
$$

Then $\Omega$ will be determined by ( $a$ ). In fact, take for r normal to the plane $\boldsymbol{\nabla}_{1}, \boldsymbol{\nabla}_{2}$, and consequently norma (which is always coplanar with $\boldsymbol{\nabla}_{1}, \boldsymbol{v}_{\mathbf{2}}$ ). Then ( $\boldsymbol{v}_{1} \mathbf{n}$ ) a vanish, and $\epsilon \mathbf{n}=\mathbf{n}$, so that (a) will become

$$
\Omega \mathrm{n}=\epsilon_{2} \epsilon_{1} \mathrm{n},
$$

and since $\epsilon_{1}, \epsilon_{2}$ are longitudinal stretchers and n is no axes of both,

$$
\Omega \mathrm{n}=\mathrm{n}
$$

Thus, the axis of rotation, or simply the axis of $\Omega$, is nc plane $\nabla_{1}, \nabla_{2}$, while the angle of rotation is given by outstanding determination of the sense of rotation is reader. To resume:

The general or six-parametric Lorentz transformatic constitute a group, but the special or three-parametric tions $L(\mathbf{V}, \mathrm{r})$ or $L(\mathbf{v})$ do not constitute a group, though the subgroups for parallel velocities. The successive of two special Lorentz transformations with non-para $\mathrm{v}_{1}, \mathrm{v}_{2}$ gives alzoys an $\mathcal{L}(\mathbf{v}, \Omega)$, that is to say, it is equ special Lorentz transformation followed by a pure sz round an axis normal to $\nabla_{1}$ and $\nabla_{2}$, which turns $\nabla=v$ $\mathbf{v}_{2} \# \mathbf{v}_{1}$, -the former of these vectors being given

## PARALLEL VELOCITIES

axes are not parallel. But this subject need not further us here.

We have touched the six-parametric Lorentz group on elucidate the question of successive transformations, as intir connected with the composition of velocities. But hencefor shall hardly need it any more. In fact, our previous transforn $\mathcal{L}(v)$, without any rotation of the space-framework, will be sufficient for all physical purposes.

Having got through this, let us return to the 'Addition The of velocities, ( x ), with the purpose of illustrating its meaning few remarks and some simple examples.

In the first place, if both $\nabla_{1}$ and $\nabla_{2}$ are small as compared wi velocity of light, then, if magnitudes of second order are negl (r) reduces at once to

$$
\mathbf{v}=\mathbf{v}_{1}+\mathbf{v}_{2}=\mathbf{v}_{2}+\mathrm{v}_{1},
$$

which is the Newtonian or classical formula for the composit velocities.

Next, consider the simplest case of parallel velocities. $\epsilon_{1} \mathbf{v}_{2}=\gamma_{1} \mathbf{V}_{2}$, and, as in Chap. V.,

$$
\nabla=\frac{\nabla_{1}+\nabla_{2}}{I+\left(\nabla_{1} \nabla_{2}\right) / c^{2}},
$$

or, counting the resultant velocity positively along $\mathbf{v}_{1}$,

$$
v=\frac{v_{1} \pm v_{2}}{1 \pm v_{1} v_{2} / c^{2}}
$$

according as $\boldsymbol{\nabla}_{2}$ is concurrent with or against $\boldsymbol{\nabla}_{\mathbf{1}}$. It will be e to consider the former case, for which

$$
v=\frac{v_{1}+v_{2}}{r+v_{1} v_{2} / c^{2}} .
$$

Let both $v_{1}$ and $v_{2}$ be smaller than $c$, say, $v_{1}=c-m, v_{2}=$ where $m, n$ are positive and smaller than $c$. Then
obtained by the accumulation of any number of veloc than $c$. This property, proved here for concurrent ve be expected to hold a fortiori for velocities of any dire rigorous proof, to be based upon the general formula the reader as a useful and interesting exercise. .

Again, if one of the compounded velocities, say $v_{1}$ then, by ( 9 ),

$$
v=\frac{c+v_{2}}{1+\tau_{2} / c}=c,
$$

i.e. the resultant of $c$ and of any other parallel velocity whether it is smaller or equal to or even greater thar the velocity of light c.* This result becomes obvious, remembered that in the present case the system $S^{\prime}$ flatland, perpendicular to the direction of motion, and the former $p^{\prime}$ is the velocity of our particle relative whole path of the particle appears to the $S$-observers point of that flatland, so that, for these observers, might as well be fixed in $S^{\prime}$.

The following is one of the most beautiful applications that were made in the early times of the doctrine.

To emphasize better the meaning of the various vel again, for the moment, $p, v^{\prime}, f^{\prime}$ instead of $\tau^{\prime}, v_{1}, v_{2}$,

$$
p=\frac{v+p^{\prime}}{1+z^{\prime} p^{\prime} \mid c^{\prime}}
$$

Now, this can be put into the form

$$
p=p^{\prime}+\kappa v
$$

where $\kappa$, expressing the fraction of $v$, which is added to rigorously by

$$
\kappa=\frac{1-p^{\prime} 2 / c^{2}}{1+v p^{\prime} / c^{2}},
$$

and approximately, for moderate values of $p^{\prime} / c$ and of $v / c$, by

Here $p^{\prime}$ is the velocity, as observed in $S^{\prime}$, of what we have hithertc called a 'material particle.' But in doing so, we have assumed only that it is something that can be recognized and watched in its changing position. Its being 'material' or not, mattered, in fact but little. We might as well have spoken from the beginning o any comparatively permanent complex of sense-data, distinctly localizable in the $S$ - and $S$ '-spaces. 'Thus, if $p^{\prime}$ be the velocity o propagation or transfer of anything that can be watched,* from the $S^{\prime}$-standpoint, and if $\tau^{\prime}$ be the velocity of $S^{\prime}$ relative to $S$, then $p$ as given by $(9 a)$, will be the corresponding velocity of propagatior or transfer, from the S-point of view, and the above $\kappa$ will be the dragring coefficient of $S^{\prime}$ (if it be empty excent for the framework) or, as the case may be, of the bodies or media carried along with $S^{\prime}$ If, for example, $S^{\prime}$ is attached to a column of air blowing uniformly past an observer resting on earth $(S)$, and if $p^{\prime}$ be the velocity o sound relative to $S^{\prime}$ (and consequently, by the principle of relativity also the velocity of sound as would be obtained by our $S$-observe in quiet air), then ( rr ) will be the dragging coefficient of air fo sound. In this case $f^{\prime} / \mathrm{c}$ is of the order $3 \cdot 3 \cdot 10^{4} / 3 \cdot 10^{10} \doteqdot 10^{-6}$ so that $\kappa$ differs from unity by little more than one millionth, anc we have a sensibly (though not rigorously) full draig of sound by air. Similarly, for light $\dagger$ propagated along a column of flowins water, as in Fizeau's experiment, if $p^{\prime}$ be its velocity relative to the water and taken from the $S^{\prime}$-stand point (and hence also the velocity of light in stationary water from the standpoint of an ordinary o S-observer), formula (ir) will express the drag of light by water

[^44]$\dagger$ In this case we can imagine an irregular train of light waves or a solitar wave or a sufficiently thin electromagnetic sheet which can be watched, at leas theoretically. And if we wish we can reduce this case to that of the motio

## THE THEORY OF RELATIVITY

The only difference is that in this case the value of $p^{\prime} / c$ exceedingly small as for sound and air, and this is why of considerable physical importance. For water in or ditions $p^{\prime} / c$ is as great as $3 / 4$, and it approaches unity nearly for optically 'rarer' media. Generally, if $n$ be sponding index of refraction, we have $p^{\prime} / c=\mathrm{r} / n$, so tha at once

$$
\kappa=\mathrm{I}-\frac{\mathrm{r}}{n^{2}},
$$

and this is the famous dragosing coefficient of Fresnel, whic so much of our attention in the early part of this $v$ which was found to be in such good agreement with

Thus, Fresnel's formula, which on the ground of t theory appeared as the outcome of a rather complica minute particles, follows here as a simple conseque fundamental theorem of relativistic kinematics, quite in of any theory of the structure of matter.

- Notice that the above is but an approximate value of $t$ coefficient, and that its rigorous value would be, by (r

$$
\kappa=\begin{aligned}
& 1-1 / n^{2} \\
& 1+\beta / n
\end{aligned},
$$

where $\beta=v / c$. But for the present Fresnel's formula, the technical difficulties of the measurements, is more than accurate. Remember that in Fizeau's experiment, as rep improved form by Michelson and Morley (p. 4r), the flowing with a velocity of 8 metres per second, so tha the order $\mathrm{ro}^{-8}$, while the observed value of the dra trusted to hardly more than two decimal figures. I do what possibilities lie in canal rays. At any rate the e discrimination between ( I 2 ) and the Fresncl formula is reserved for the future.

As a further example of composition of velocitics, let the case of perpendicular components. Returning once notation adopted in the general formula ( r ), we have in

## PERPENDICULAR COMPONENTS

Similarly, the resultant of $\mathbf{V}_{\mathbf{2}}$ followed by $\mathbf{v}_{\mathbf{1}}$ will be

$$
\nabla_{21}=\nabla_{2} \# \nabla_{1}=\nabla_{2}+\frac{\nabla_{1}}{\gamma_{2}}=\nabla_{2}+\nabla_{1} \sqrt{I-\beta_{2} 2^{2}} .
$$

In Fig. 14, in which $O A N B$ is a rectangle, the former of t ] vectors is given, in absolute value and direction, by $O C$, and latter by $O D$, while the diagonal $O N$ represents the Newto resultant. As was already remarked, the absolute values of

relativistic resultants $\boldsymbol{\nabla}_{12}, \boldsymbol{\nabla}_{21}$ are equal to one another, the sq of each being in the present case given by

$$
v^{2}=v_{1}^{2}+v_{2}^{2}-\frac{1}{c^{2}} v_{1}^{2} v_{2}^{2},
$$

instead of which we may conveniently write

$$
\beta^{2}=\beta_{1}{ }^{2}+\beta_{2}{ }^{2}-\beta_{1}{ }^{2} \beta_{2}{ }^{2},
$$

or also, as a particular case of (b), p. 169 ,

$$
\gamma=\gamma_{1} \gamma_{2}
$$

To obtain the angle $\zeta=C O D$ enclosed by the two resultants, their scalar product and divide it by $v^{2}$. The result will be

$$
\cos \zeta=\frac{\gamma_{1} \beta_{1}^{2}+\gamma_{2} \beta_{2}^{2}}{\gamma \beta^{2}} .
$$

Thus, for $v_{1}=v_{2}$ equal $\frac{1}{5}, \frac{1}{2}, \frac{9}{10}$ of the velocity of light, the an would be, in round figures, $\mathrm{I}^{\circ}, 8^{\circ}, 43^{\circ}$ respectively, or more accur
while it is moved relatively to the paper ( $S$ ) horizont velocity $v_{1}$, the point of a pencil is led along the vertic the velocity $v_{2}$ relative to the ruler $\left(S^{\prime}\right)$, then the pen the line $O C$, e.g. the segment $\overline{O C}$ in unit time ( $S$-tim other hand, if the ruler is moved vertically with velocit pencil is led along its horizontal edge with velocity $v_{1}$, the pencil will draw the line $O D$. According to classica the line drawn would be in both cases the diagonal of $t$ Notice that from the paper-standpoint the velocities pounded are: in the first case $O A$ and $A C$ (not $A N$ ) second case $O B$ and $B D($ not $B N)$. In the old kine was no question of discriminating between the pap ruler-standpoints.

So much to explain the true meaning of $\mathbf{v}_{1} \# \mathbf{\nabla}_{2}$, as from $\nabla_{2} \# \nabla_{1}$.

The space in the ordinary sense of the word, or positions being assumed Euclidean in both the old a theory, the space representative of velocities, or wh called the kinematic space, is again the Euclidean spac kinematics, but non-Euclidean in relativistic kinematic to represent the resultant $\mathbf{V}_{12}$ on the same Euclidean p with the component velocities, we had to cut off from $C N$, and similarly, in constructing $\nabla_{21}$ we had to cut the piece $D N$. If we want to obtain the resultant construction without cutting off anything from the seg senting the component velocities or any functions of e velocities alone, then we have to use a non-Euclideal Lobatchewsky's and Bolyai's space of constant negati or, as it is appropriately called, a hyperbolic space.*

In short, the relativistic kinematic space is a nyperb

[^45]
## COMPOSITION OF VELOCITIES

To see this, take, for simplicity, the above case of $\nabla_{1} \perp$ Denote the angle contained between $\mathbf{v}_{1}$ and the resultant $\boldsymbol{\nabla}=\mathbf{v}$ i.e. the angle $A O C$ of Fig. r4, by $\theta_{2}$. Then, by (13),

$$
\tan \theta_{2}=\frac{v_{2}}{v_{1} \gamma_{1}}=\frac{\beta_{2}}{\beta_{1} \gamma_{1}}
$$

and, by (16),

$$
\gamma=\gamma_{1} \gamma_{2} .
$$

Now, instead of the absolute value of each of the velocit introduce the corresponding imaginary angle $\omega$,

$$
\omega=\arctan (\iota \beta)
$$

as defined by (ro), Chap. V. Then $\gamma=\cos \omega, \beta \gamma=-\iota \sin \omega$, the above pair of formulae will become

$$
\begin{aligned}
& \cos \omega=\cos \omega_{1} \cdot \cos \omega_{2}, \\
& \tan \theta_{2}=\frac{\tan \omega_{2}}{\sin \omega_{1}},
\end{aligned}
$$

and these are the known formulae of spherical trigonometry fo right-angled triangle, whose sides and hypothenuse are $\omega_{1}, \omega_{2}$ and whose angle opposite to $\omega_{2}$ is $\theta_{2}$, the only difference being $t$ here all the sides are imaginary. This is the property remarked Sommerfeld (cf. last footnote).

Next, to get rid of the imaginary sides, introduce, for each veloc instead of $\omega$ the real angle $a$, as defined by (20), Chap. V., such th

$$
\tanh \alpha=\beta=v / c .
$$

Then, as was previously noticed, $\omega=\iota a$, and, since

$$
\sin (\iota a)=\iota \sinh a, \quad \cos (\iota a)=\cosh a,
$$

the above formulae become at once

$$
\left.\begin{array}{rl}
\cosh a & =\cosh \alpha_{1} \cdot \cosh \alpha_{2} \\
\tan \theta_{2} & =\frac{\tanh \alpha_{2}}{\sinh a_{1}}
\end{array}\right\}
$$

Now, these are exactly the formulae for a right-angled triangle Lobatchewskyan or hyperbolic space.* Thus, if $a_{1}$ and $\alpha_{2}$ (Fig.
are segments of geodesics or shortest lines in hyperb representing the component velocities, the shortest line $a$, the triangle, will represent the resultant velocity, as regard and inclination, $\theta_{2}$. The same property may be proved general, i.e. for component velocities including with on any angle. Here it will be enough to give the ler


Fig. 15


Fite. $x$.

Denoting by $\pi-\theta$ the angle $\mathbf{v}_{1}, \mathbf{v}_{\mathbf{n}}$, so that $\theta$ itself is opposite to $a$ (Fig. 16), we have

$$
\frac{\mathrm{I}}{i^{2}}\left(\mathbf{v}_{1} \mathbf{v}_{2}\right)=-\beta_{1} \beta_{2} \cos \theta,
$$

so that our previous formula (i),

$$
\gamma=\gamma_{1} \gamma_{2}\left[r+\frac{r}{e^{2}}\left(\nabla_{1} \nabla_{2}\right)\right],
$$

becomes at once

$$
\cosh a=\cosh \alpha_{1} \cdot \cosh \alpha_{2}-\sinh a_{1} \cdot \sinh \alpha_{2} \cdot \cos \theta
$$

The determination of the angle $\theta_{2}$, by means of the gene ( x ), is left to the reader.

Notice that, as long as we are concerned only with tw and their resultant, we have no need of three-dimensional space. What we want then is a Lobatchewskyan plane c of constant negative curvature. Now this may be easil of any size in Euclidean space. Models of such a surfe as a pseudosphere, which is a surface of revolution,* belo the outfit of many mathematical class-rooms. Our last must be imagined to be drawn on a pseudosphere (whic has nothing more imaginary about it than the page on $v$
commutativity and all the remaining properties of the addition of velocities. In this way the relativistic rules of the composition of velocities could be made accessible even to all those who do not like to think of hyperbolic, and other non-Euclidean, spaces.

It has been proposed by Ir. Robb* to call our above $a$, as defined by (r8), that is

$$
\begin{equation*}
a=\operatorname{arctanh} \frac{\pi}{c} \tag{2I}
\end{equation*}
$$

the rapidity, corresponding to the velocity 7 . It seems a very convenient name for the purpose. Using it, we may briefly restate the above result as follows:

Any two rapidities are compounded by the triangle-rule in hyperbolic space.

Whence also: the resultant of any number of rapidities arranged in a chain in hyperbolic space, is the geodesic or the straight line of that space, drawn from the beginning to the end of the chain.


Fici. x 7.
Notice that if rapidity is to involve 'direction' as well as size or absolute value, it has to be considered as a vector localized in its ozen line, i.e. in a Lobatchewskyan straight line or shortest line upon our pseudosphere. In connexion with this we have only the triangle-rule, and not the parallelogram-rule, as in Newtonian kinematics. There are no parallelograms in hyperbolic space or upon a pseudosphere, any more than upon a sphere. To express that direction is involved, we may write for the rapidities $a_{1}, a_{2}$, etc., and use the ordinary sign + for their addition, keeping in mind that each of these rapidity-vectors can be shifted only along its own line, and, consequently, that their addition is non-commutative, unless $a_{1}$, $a_{2}$ are on the same line. Thus, the rapidity $a_{1}+a_{2}$ (Fig. 17) is $A B$, while $a_{21}+a_{1}$ is $C D$, which, though of the same length, is on a different line.

* Alfred A. Robl, Optical Geometry of Motion, Cambridge, W. Ileffer \& Sons, 1911.

Remembering that $\tanh a=\left(e^{a}-e^{-a}\right) /\left(e^{a}+e^{-a}\right)$, we can of (2 I),

$$
a=\frac{1}{2} \log \frac{1+\beta}{1-\beta}=\beta+\frac{1}{83} \beta^{3}+\frac{1}{5} \beta^{5}+\ldots
$$

For small values of $\beta$ we have, up to quantities of the $a \doteqdot \beta=v / c$, so that for small velocities the corresponc are small fractions, of the order of $\beta$, and the Lo triangle becomes a Euclidean triangle, as in classica It seems worth mentioning that to unit rapidity corres velocity, amounting to $\frac{3}{4}$ of the velocity of light; mo we have

$$
\beta=\cdot 7616 \text { for } a=1 .
$$



Fic. 18.
From (21a) we see most immediately that to the vel itself corresponds an infinite rapidity,

$$
a=\infty \text { for } \beta=x \text {. }
$$

Now, if two sides of a pseudospherical triangle are fir side is also finite. Thus, our previous statement, that of any velocities smaller than that of light is again sme velocity of light, is reduced to an obvious property triangles.

To close the discussion of this beautiful subject, bu more. Lobatchewsky's II $(a)$, the angle of parallelism for

## COMPOSITION OF VELOCITIES

Thus, equations (r9) can be written, in terms of ordinary trigo metric functions of the respective angles of parallelism,

$$
\left.\begin{array}{rl}
\sin \Pi(a) & =\sin \Pi\left(a_{1}\right) \cdot \sin \Pi\left(a_{2}\right) \\
\tan \theta_{2} & =\tan \Pi\left(a_{1}\right) \cdot \cos \Pi\left(a_{2}\right),
\end{array}\right\}
$$

which is the original form of Lobatchewsky's own formulae, fo right-angled triangle. Similarly, the general formula (20), will beco

$$
\sin \Pi(a)=\frac{\sin \Pi\left(a_{1}\right) \cdot \sin \Pi\left(a_{2}\right)}{1-\cos \Pi\left(a_{1}\right) \cdot \cos \Pi \Gamma\left(a_{2}\right) \cdot \cos \theta}
$$

which is Lobatchewsky's fundamental formula. The unit of len here adopted is that employed by Lobatchewsky, i.e. that ler whose negative reciprocal square is the curvature of the representa hyperbolic space, or the curvature of the pseudosphere upon wh the triangles are to be drawn. Thus, if we take for that purp a pseudosphere of curvature $-1 / 100 \mathrm{~cm}^{2}$., a segment of its geod ro centimetres long will correspond to the rapidity $a=1$, and sequently will represent the velocity 76 c which is a little ab the velocity of light in water.

Instead of (18), we shall now have, by the second of (22), omitting the unnecessary argument,

$$
\cos \mathrm{I}={ }_{c}^{n}-\beta
$$

For very small values of $\beta$ the angle of parallelism II is nearl right angle, as in a Euclidean plane. 'Thus, for the earth's orl motion $\beta=10^{-4}$ and $\Pi=89^{\circ} 59^{\prime} 39^{\prime \prime} \cdot 4$, so that the departure fi Euclid amounts only to $20^{\prime \prime} \cdot 6$. But if we turn to swift electr as observed in kathode rays and $\beta$-rays of radionctive substan the angle of parallelism is very considerably reduced. For $\beta=$ and $\cdot 95$ (Kaufmann observed cven $\cdot 99$ and more) I find $\mathrm{IL}=25^{\circ}$ and $18^{\circ} \mathrm{I} 2^{\prime}$ respectively. At the limit, for light-velocity, the angl parallelism would vanish altogether.

## CHAPTER VII.

## PHYSICAL QUATERNIONS. DYNAMICS OF A

The importance of the study of world-vectors quaternions for relativistic investigations is obvious form of the laws of physical phenomena is to be pr Lorentz transformation, they can involve besides the coordinates, and, of course, besides any invaria sets of magnitudes which, caeteris paribus, bear legitimate systems the same relations to its time and in any other of such systems. Therefore, physical whatever mathematical form we may choose for tet tudes transformed like $l, x, y, z$ and of sets of magn from them) constitute, as it were, the building 1 modern relativist. And what is most important to $k$ that he cannot use any other material. For if he did, $h$ to infringe against the fundamental principle of the wl

To try to describe in a few abstract sentences th material is procured and how it is used, would be The reader will see it best from particular cases.

As yet we had, properly speaking, only one physi which we made the standard of such quaternions, $e$. quaternion

$$
q=l+\mathbf{r}=c c t+\mathbf{r} .
$$

This was transformed into $q^{\prime}$ by the operator $Q[$ quaternion $X$ was transformed into $X^{\prime}$ by the sam wrote $X \sim q$, and if it had also, like $q$, an imagina real vector, we wrote $X \simeq q$, and called $X$ a phys Such was our definition given in Chap. V., entirel that of a four-vector.

## VELOCITY-QUATERNION

Now let us look for other physical quaternions. An indefi number of such can be obtained at once from $q$ itself.

In fact, let $q$ belong, say, to a material particle at a given insta of its history. Let the particle move about in an arbitrary man and let p be its instantaneous velocity in $S$. Then its posit quaternion at the instant $t+d t$ will be $q+d_{l}$, and this as well will certainly be a physical quaternion. And since $Q[] Q$ distributive (or since the Lorentz transformation is linear and ho geneous), the difference of these two quaternions, i.e.

$$
d q=d l+d \mathbf{r}=[\iota c+\mathbf{p}] d t
$$

will again be a physical quaternion, $\simeq q$. Therefore, as we kr from Chap. V., its tensor

$$
\mathrm{T} d y=t d t \sqrt{c^{2}-p^{2}}
$$

will be an invariant. Divide it by $c$; then

$$
d \tau=d t \sqrt{\mathrm{I}-\frac{p^{2}}{c^{2}}=\frac{d t}{\gamma_{p}}}
$$

will again be an invariant. Its value will be real, provided $p$ is not greater than $c$. And since $d q$ is a physical quaternion, shall have also

$$
Y=\frac{d q}{d \tau} \simeq q,
$$

that is, $Y$ will again be a physical quaternion. Let us call it velocity-quaternion of the particle in question. Its developed form

$$
Y=\gamma_{p}[\iota c+\mathbf{p}],
$$

where $\mathbf{p}$ is the ordinary vector-velocity of the particle, justify the above name.

The plain meaning of our result is that $Y^{\prime}=Q Y Q$, i.e. that

$$
\iota c \gamma_{p} \text { and } \mathbf{p} \gamma_{p}
$$

Using this, the reader will obtain at once the addition th velocities, identical with ( $\mathrm{r} a$ ), Chap. VI., along with the

$$
\gamma_{p}=\gamma_{v} \gamma_{p^{\prime}}\left[\mathrm{I}+\frac{\mathrm{x}}{c^{2}}\left(\nabla \mathrm{p}^{\prime}\right)\right]
$$

(identical with (b), p. ェ69), which is a consequence of that Thus, the relativistic rule for the composition of velocities i in the statement that $Y$ is a physical quaternion.

The infinitesimal scalar $d \tau$, as defined by (z), deserve attention. For $p=0$ it reduces to $d t$, the element of $S$-time, but is, in general, smaller than dt. It has the a of being an invariant, which $d t$ is not. In other words, of $d \tau$ is independent of the choice of our standpoint, bei for all legitimate systems. It belongs to the particle. I property will obviously hold for

$$
\tau=\int d \tau=\int \begin{aligned}
& d t \\
& \gamma_{p}
\end{aligned},
$$

where the integral is taken along any portion of the particle' or along any segment of its world-line, from an arbitrar initial point to the variable end-point. The parameter $\tau$, thus may be called, after Minkowski, the proper time* of the The velocity $\mathbf{p}$ of the particle, entering into cach clement its square, may, in general, vary from instant to instant, as both absolute value and direction. If the particle is fixed proper time is the ordinary time $t$ of the system $S$. An particle moves uniformly in $S$, we can imagine a system $S^{\prime \prime}$ it will be at rest. And then the proper time of the par become the ordinary time of that system.

The velocity-quaternion may now be described as the d of the position-quaternion with respect to the proper tim particle. It will often be convenient to use the dot differentiation. Thus, $Y=\dot{q}$.

The name corresponding to $Y$ in the language of four-dim algebra would be four-velocity, $\dagger$ and its matrix-form would be

## VELOCITY AND ACCELERATION

Remember that $d \tau$, as originally defined, was simply the of $d q$ divided by $c c$. The tensor of the velocity-quaterni therefore,

$$
\mathrm{T} Y=\iota c .
$$

We know, from Chap. V., that the tensor of every physical quate is an invariant. In the present case this knowledge doe furnish us anything new. For $c$ is, by the fundamental assum of the theory, a universal constant. The norm of $Y$ being nes namely equal to $-c^{2}$, the velocity-quaternion is always tin In Minkowski's language we should say that the four-velo along the world-line of the particle in question.

Since $Y$ is a physical quaternion and $\tau$ is an invariant,

$$
Z=\frac{d Y}{d \tau}=\ddot{q}
$$

will again be a physical quaternion which, for obvious reasons be called the acceleration-quaternion. So also will $d^{3} q / d \tau^{3}$, c physical quaternions, each $\simeq q$, and obviously also $d^{3} q_{c} / d \tau^{3}$, each $\simeq q_{c}$. But of all these derivatives of $q$ we shall hardly more than the first two, containing the velocity and the accele

Let $Y_{c}^{r}$ be the conjugate of $Y$. Then, by Quat. 7, we can for its norm the product $Y Y_{c}$ or also $S Y Y_{c}$, and consequ instead of (4),

$$
Y Y_{c}=-c^{2} .
$$

Differentiating this with respect to $\tau$, we have

$$
Z Y_{c}+Y Z_{c}=0,
$$

or also

$$
S Z Y_{c}=0,
$$

which says preciscly the same thing as ( $6 a$ ).* Such then relation which holds always between the acceleration- an velocity-quaternion of a particle. Using the developed form $q$ we should have, correspondingly,

$$
\begin{aligned}
& (\dot{\mathbf{r}} \dot{\mathbf{r}})+\dot{l}^{2}=-c^{2} \\
& (\ddot{\mathrm{r}} \ddot{\mathrm{r}})+\dot{l} \ddot{l}=0,
\end{aligned}
$$

[^46]or, in a still more developed form,
\[

$$
\begin{aligned}
& \dot{x}^{2}+\dot{y}^{2}+\dot{z}^{2}+\dot{l}^{2}=-c^{2} \\
& \dot{x} \ddot{x}+\dot{y} \ddot{y}+\dot{z} \ddot{z}+\dot{l} \dddot{z}=0 .
\end{aligned}
$$
\]

In four-dimensional language, as explained in Chap. V formula would read: The four-acceleration is alzeay's nor four-velocity and, consequently, to the world-line of the p famous statement of Minkowski. This cardinal property its short quaternionic expression in (6). Observe that th of that equation is the same thing as Sommerfeld's scald of the corresponding four-vectors. But the invarianc expressions is seen more immediately on the quaternion In fact, remembering that $Q Q_{c}=Q_{c} Q=r$, we have, by

$$
\mathrm{S} Z^{\prime} Y_{c}^{\prime}=\mathrm{S} Q Z Q Q_{c} Y_{c} Q_{c}=\mathrm{S} Q Z Y_{c} Q_{c}=\mathrm{S} Q_{c} Q Z Y_{c}=\mathrm{S} \%
$$

Next, as regards the transformational properties of the ac These are entirely determined by saying that $Z=c \ddot{t}+\ddot{\mathrm{r}}$ is quaternion. For this means simply that $\ddot{t} \ddot{\mathrm{r}}$ are transfo $t$, r. If, therefore, $S^{\prime}$ be a system moving relatively to uniform velocity v , we have, according to ( $\left.\mathrm{r}^{\prime} b\right)$, Chap. $V$

$$
\left.\begin{array}{l}
\ddot{\mathrm{r}}=\epsilon_{\nabla} \ddot{\mathrm{r}}^{\prime}+\nabla \gamma_{v} \ddot{t}^{\prime} \\
\ddot{t}=\gamma_{v}\left[\ddot{t^{\prime}}+\frac{I}{c^{2}}(\dot{\mathrm{vi}})\right],
\end{array}\right\}
$$

where the subscripts are to remind us that $\gamma, \epsilon$ are to be the velocity $\nabla$. The dots denote, on both sides, differenti respect to the same variable $\tau$. For, as the reader alrea $d \tau^{\prime}=d \tau$. There is no difficulty in developing these for thus finding the ordinary acceleration

$$
\mathrm{a}=\frac{d \mathbf{p}}{d t}
$$

in terms of $\mathbf{a}^{\prime}=d \mathbf{p}^{\prime} / d t^{\prime}$ and $\mathbf{p}^{\prime}$ (or vice versa), for any pair of systems $S, S^{\prime}$ picked out at random. But this would worth the trouble.

To see the plain kinematical meaning of the second with respect to $\tau$ and hence of the whole acceleration-c we have to place ourselves at a standpoint which bears th

## REST-ACCELERATION

possible relation to the moving particle itself. Let us then for $S^{\prime}$ that particular system of reference with respect to wh the particle is instantaneously at rest. In other words, let $S^{\prime}$ b system whose uniform velocity v , relative to $S$, is equal in size direction to the instantaneous velocity of the particle, i.e. to value of p at a given instant of its history. Then, at that inst $p^{\prime}=0$ and $\gamma^{\prime}=\gamma\left(p^{\prime}\right)=1$. Now, we had, generally, $d t / d \tau=\gamma$ Therefore,

$$
\ddot{t^{\prime}}=\gamma^{\prime} \frac{d \gamma^{\prime}}{d t^{\prime}}=\frac{d \gamma^{\prime}}{d t^{\prime}}=\frac{\mathrm{I}}{c^{2}} \gamma^{\prime 3} p^{\prime} \frac{d p^{\prime}}{d t^{\prime}}=0,
$$

or $\ddot{l}^{\prime}=0$, as might have been expected, and in a similar way,

$$
\dddot{\mathbf{r}}^{\prime}=\frac{d \gamma^{\prime} \mathbf{p}^{\prime}}{d t^{\prime}}=\frac{d \mathbf{p}^{\prime}}{d t^{\prime}}=\mathrm{a}^{\prime},
$$

so that $Z^{\prime}=\ddot{Z}^{\prime}+\ddot{i}^{\prime}$, the acceleration-c puatemion relative to $S^{\prime \prime}$, for instant in question, is simply

$$
Z^{\prime}=\mathrm{a}^{\prime}
$$

i.e. equal to the ordinary acceleration of the particle with res to $S^{\prime}$. Since $S^{\prime}$ is that particular system of reference in which particle is instantancously at rest, it may be called the rest-sy: and the corresponding $a^{\prime}$ the rest-acceleration of the particle.*

Thus, the scalar part of the acceleration-quaternion $Z^{\prime}$ vani identically, and its vector part is equal to the rest-accelerati Consequently, $\mathrm{T} \not /{ }^{\prime}=a$ '. And since the tensor of every phy quaternion is an invariant, we have also, for any legitimate syster

$$
\mathrm{T} \%=a^{\prime} .
$$

In words, the tensor of the acceleration-guaternion is equal to absolute value of the rest-acceleration of the particle. It acqu thus an immediate kinematical meaning. At the same time form (7), in which we have now to write $\mathbf{v}=\mathrm{p}$, give us, for the syste which in a certain sense is an unnatural system of reference,
and $\ddot{t}=c^{-2} \gamma\left(\mathrm{pa}^{4}\right)$, sorthat the whole acceleration-quat written: : wa $_{1}$

$$
Z=\frac{d Y}{d \tau}=\frac{\imath}{c} \gamma\left(\mathrm{pa}^{\prime}\right)+\epsilon \mathrm{a}^{\prime}
$$

Here $_{j} \cdot \mathbf{p}$ is the velocity of the particle relative to the stretcher $\epsilon=\epsilon_{\mathfrak{p}}$, of ratio $\gamma_{p}$, acts along the instanta) of $p$ or tangentially to the path of the particle. Thus if the tangent to the path of the particle, drawn in $t$ its motion, be our instantaneous $x$-axis,

$$
\ddot{x}=\gamma a_{x}^{\prime}, \quad \ddot{j}=a_{y}^{\prime}, \quad \ddot{z}=a_{z}^{\prime}
$$

and $c \ddot{t}=\beta \gamma a_{x}{ }^{\prime}$. If the $y$-axis be taken in the oscul the path, then $\ddot{z}=0$. Since we already know, by (

$$
\ddot{x}^{2}+\dot{y}^{2}+\ddot{z}^{2}-c^{2} \ddot{t}^{2}=a^{\prime 2},
$$

the formula for $\ddot{t}$ becomes superfluous.
Finally, to express $\ddot{\mathbf{r}}=d^{\ddot{ }} \mathbf{r} / d \tau^{2}$ in terms of the ordir tion $\mathrm{a}=d \mathbf{p} / d t$, remember once more that $d t / d \tau=\gamma$. definition of $\gamma$,

$$
\frac{d \gamma}{d t}=\frac{\mathrm{x}}{c^{2}} \gamma^{3} \phi \frac{d p}{d t}={ }_{c^{3}}^{\mathrm{r}} \gamma^{3}(\mathrm{pa}),
$$

the result will be

$$
\ddot{\mathrm{r}}=\gamma \frac{d \gamma \mathbf{p}}{d t}=\gamma^{2}\left[a+\frac{\mathrm{r}}{\gamma^{2}} \gamma^{\mathrm{g}} \mathbf{p}(\mathbf{p a})\right]=\gamma^{2}\left[a+\beta^{2} \gamma^{2} \mathbf{u}\right.
$$

where $\mathfrak{u}$ is the unit of $\mathbf{p}$. Now, $\mathrm{I}+\beta^{2} \gamma^{2}=\gamma^{2}$, identica the bracketed expression is the vector sum of the lo of a magnified $\gamma^{2}$ times and of its unaltered trans simply the result of a double application of the stre ultimately,

$$
\ddot{\mathbf{r}}=\frac{d^{2} \mathrm{r}}{d \tau^{2}}=\gamma^{2} \epsilon^{2} \mathrm{a},
$$

whence also, by (10),

$$
\gamma^{2} \in a=a^{\prime},
$$

giving the conncxion between a and the rest-accele Cartesians, with the above choice of axes, for the ld the transversal components of $\ddot{\mathrm{r}}$,
and

$$
\ddot{x}=\gamma^{4} a_{x}, \quad \ddot{y}=\gamma^{2} a_{y}, \quad \ddot{z}=\gamma^{2} c z_{z},
$$

$$
\gamma^{3} a_{x}=a_{x}^{\prime}, \quad \gamma^{2} a_{y}=a_{y}^{\prime}, \quad \gamma^{2} a_{z}=a_{z}^{\prime}
$$

## HYPERBOLIC MOTION

By (r3) we have also, writing $p / c=\beta$,

$$
a \cdot \gamma^{3} \sqrt{ } \mathrm{x}-\beta^{2} \sin ^{2}(\mathbf{p}, \mathbf{a})=a^{\prime},
$$

which is merely a developed form of (9). In fact, the right-h side of ( I 4 ) is seen, by ( Ir ), to be identical with $T Z$.

The simplest case of motion of a particle occurs wher permanently nil, and consequently also $a=0$. This is, as in class kinematics, the trivial case of uniform rectilinear motion. Such mot preserves its character in all legitimate systems. In fact, owing to linearity of the Lorentz transformation, any motion which is uniform rectilinear with respect to one of these systems will be so relati to any other of them. A straight world-line will remain strai 'The next simplest kind of motion, which also preserves its chara in all such systems of reference, occurs when the non-vanishing acceleration is constant in size and direction, i.e. when $d \mathrm{a}^{\prime} / d \tau^{\prime}=0$, hence also $d a^{\prime} / d t=0$. Then, by ( r 3 ), the vector $\gamma^{2} \in a$ is cons in $S$, that is to say, independent of $t$. But since the axis of stretcher $\epsilon$, or the $a$-axis in ( $13 a$ ), instead of being fixed, is at instant tangential to the path of the particle, which may be curvilin it does not follow that even the direction of the acceleration a be constant in S. 'Thus, the general case of such motion, whic the counterpart of the uniformly accelerated or parabolic motion classical kinematics, would still be fairly complicated. The simp sub-case, which also will show best the characteristic propertie this kind of motion, occurs, of course, when the particle moves a straight line. Let this be our $x$-axis. Then, by ( $\mathrm{r}_{3}$ ),

$$
\gamma^{3} a=\gamma^{3} \frac{d t}{d t}=a^{\prime},
$$

or

$$
\left.a^{x^{\prime}} d t=\gamma^{3} d \beta=\frac{d \gamma}{\beta}=\frac{1}{2}\left(\gamma^{2}-1\right)^{2}\right),
$$

whence, counting the time $t$ from the instant at which $p=0$,

$$
p=\frac{d x}{d t}=a^{\prime} t \cdot\left[1+\binom{a^{\prime} t}{c}^{2}\right]^{-\frac{1}{2}} .
$$

## THE THEORY OF RELATIVITY

Thus, as long as $a^{\prime} t$ is small in comparison with the ve light (whether before or after the instant when the particl rest in $S$ ), we have, approximately, $p \doteqdot a^{\prime} t$, and $x=\frac{1}{2} a^{\prime} t$ as in the Galilean free fall. But after a sufficiently long neglected terms begin to make themselves sensible, and the of the particle, instead of increasing beyond all limit asymptotically to the velocity of light. In fact, we ha (15), for any given $a^{\prime}, p=\mp c$ for $t=\mp \infty$.

Integrating once more, and choosing the origin of $x$ for $t=0, x=x_{0}=c^{2} / a^{\prime}$, we obtain

$$
x^{2}-c^{2} t^{2}=\left(\frac{c^{2}}{a^{\prime}}\right)^{2} .
$$

Thus, the world-line of our particle, in rectilinear moti constant rest-acceleration, is an equilateral hyperbola (Fig.


Fig. rg.
length of whose semiaxes is equal $\frac{c^{2}}{a^{\prime}}$. This motion has th been called by Born, who was the first to study it, hyperbolic The asymptotes $A O, O A$ correspond, as in a previous figure velocity $c$, directed towards and away from the origin. The arrives from $x=\infty$ with light-velocity, moves up the $x$-ax ever-diminishing velocity towards $P$, the vertex of the repres hyperbola, where its velocity is nil. Then it turns and move the $x$-axis with increasing velocity, which again tends asympt

## HYPERBOLIC MOTION

more sudden is the passage of the particle's velocity from $-c$ throt zero to $+c$. Taking, instead of $S$, another system of reference which moves uniformly along the $x$-axis, and whose origin coinci with $O$ at the instant $t^{\prime \prime}=t=0$, we shall have again equation ( for the new variables. For, $x^{2}-c^{2} t^{2}$ is an invariant, and so is acceleration $a^{\prime}$, by its very definition. It is this we meant by say that the considered kind of motion preserves its character in differ systems of reference, -a property which is not shared by Galilean uniformly accelerated motion to which would correspo a parabola as world-line. Remember that in classical kinema there was no question of discriminating between the ordin $S$-acceleration $a$ and the rest-acceleration $a^{\prime}$.

We may mention here that the hyperbolic motion is particule interesting in connexion with the theory of the relativistic 'rig body. But its chief importance lies herein that any varia motion can be closer approximated by it than by uniform moti In other words, any curved world-line can be brought into clo contact with a hyperbola than with a straight line. There is every point $P$ of such a world-line a hyperbola of closest cont with the world-line, which plays the part of the familiar cir of curvature and which was called by Minkowski the hyperl of curvature. If $O$ be the centre of this hyperbola (whose ver is at $P$ ), then the four-acceleration will be given by the wo vector drawn in the direction $O P$ and having the absol value $c^{\Downarrow} / \overline{O P}$, or $\frac{c^{2}}{c^{2}}$. In fact, as we have just seen, the expression is simply equal to $a^{\prime}$, and this again was seen to identical with $T \not Z$, or with the size of the four-acceleration wh was always normal to the world-line. Remembering, on the ot hand, that 'TY=cc, or that the square of the four-velocity is ec to $-c^{2}$, the reader will at once perceive the perfect analogy betw the above property of $c^{2} / O P$ and the familiar formula: nor acceleration $=$ square of velocity divided by radius of curvature. will also be noticed that, the square of the four-vclocity be negative, the four-acceleration is directed away from the centre

Such then are the properties of the velocity- and th quaternion. These being simply the (first and seco of the position-quaternion $q$ of a particle with respec time, our above considerations had a purely kinema Although we have spoken of $q$ as defining the position a 'particle,' the latter could mean anything which car at all and watched in its varying position. Of cour be possible, the 'particle' must have some or other cl its own. But these must not necessarily be quantitativ to say nothing of their being constant in time or equ standpoints or systems of reference. The moving th might have no such characteristic at all.

But let us suppose there is a certain magnitude that there is, more especially, a scalar coefficient attached to the particle and fulfilling the latter invoriant with respect to the Lorentz transformation. $m$, without yet giving it any name. 'Then $m y, m Y_{F}$ and so on, will all be physical quaternions, and, con of them may be employed, along with other physic for relativistic purposes, i.e to write down laws of particle. Such laws would be admissible, in tha word, that they would not infringe against the princi But this does not imply, of course, that they will Nature. If such laws or equations are to be of a physicist, and if they do not happen to cover an enti ground, they have to coincide, roughly at least, an circumstances, with what is otherwise known to hold In the present case we shall require that the relati of motion should coincide, approximately for smal rigorously, when referred to the rest-system, with N law of motion.

Keeping this in mind, let us see what are the c assuming, as the equation of motion of our particle

$$
\frac{d m Y}{d \tau}=X .
$$

## DYNAMICS OF A PARTICLE

the quaternion $X$, and implies that it has an imaginary scalar a real vector; the coefficient $m$ being supposed real.

By its construction, equation ( I 7 ) will preserve its form in legitimate systems of reference.

Remembering that $d t / d \tau=\gamma$, write, instead of ( 17 ),

$$
\frac{d m Y}{d t}=\stackrel{\mathrm{r}}{\gamma} X,
$$

and denote the imaginary scalar of $\gamma^{-1} X$ by $\iota v$ and its real ve by $\mathbf{N}$, i.e. put

$$
\frac{\mathrm{I}}{\gamma} X=\iota \nu+\mathrm{N} .
$$

Then ( r 7 ) will split into the vector and the scalar equations

$$
\left.\begin{array}{l}
d \\
d t \\
d \gamma \mathbf{p}=\mathbf{N} \\
d \\
d t \\
d \gamma \gamma=v
\end{array}\right\}
$$

where $\mathrm{p}=d \mathrm{r} / d t$ is the ordinary velocity of the particle relative t and $\gamma=\left(x-p^{2} / c^{2}\right)^{-\frac{1}{2}}$.

Written for the rest-system, which we shall again denote by the first of these equations becomes at once

$$
m \mathbf{a}^{\prime}=m \frac{d \mathbf{p}^{\prime}}{d t^{\prime}}=\mathbf{N}^{\prime}
$$

i.e. identical with the classical equation of motion of a particl mass $m$ under the action of the impressed force $\mathrm{N}^{\prime}$. Thus, above requirement is fulfilled. In view of this property, coefficient $m$ is called the rest-mass of the particle.* The ordi force, $\mathbf{N}^{\prime}$ in the rest-system and, gencrally, $\mathbf{N}$ in any legitir system $S$, is called the 'Newtonian force' in distinction from the vector part of $X$, which is the 'Minkowskian force.' For rea:

[^47]which will appear when we come to consider the actions of the electromagnetic field, we shall have former and not the latter as the force acting upon The second of (r7a) becomes, for the rest-syste
$$
\frac{d m}{d t^{\prime}}=\frac{v^{\prime}}{c} .
$$

As will be seen in Chapter IX., there are reasor that even the rest-mass may vary with time. Ir in general, be the case when the internal state of th during its motion. But, to simplify matters, let the particle's internal state is kept constant. Then will be constant in time. This implies $\nu^{\prime}=0$, so quaternion $X$ will be reduced, for the rest-system,

$$
X^{\prime}=\mathbf{N}^{\prime}
$$

and we shall have, for any legitimate system $S$,

$$
\mathrm{T} X=\mathrm{T} X^{\prime}=N^{\prime}
$$

where $N^{\prime}$ is the absolute value of the (Newtonian) fo from the standpoint of the rest-system.

With this supposition of a constont rest-mass, equati

$$
m \frac{d Y}{d \tau}=m \%=X .
$$

Now, by (6), $\mathrm{S} \% Y_{\mathrm{c}}=0$, and consequently also

$$
S X Y_{0}=0,
$$

or, in developed form, by (3a) and ( x 8 ),

$$
(N p)=c v .
$$

Hence, by the second of ( $17 a$ ), which is simply of (20),

$$
(\mathrm{N} \mathbf{p})=\frac{d}{d t}\left(m c^{2} \gamma\right)
$$

Thus, $(\mathbf{N p})$ being the activity of the forec $\mathbf{N}$, th the quaternionic equation (20) expresses the fri

## DYNAMICS OF A PARTICLE

(principle of Vis-viva), giving for the kinetic energy of the the value

$$
T=m c^{2}(\gamma+\text { const } .) .
$$

If we require that, for $p=0$ (i.e. for $\gamma=1$ ), $T_{0}=0$, we have const. $=-\mathrm{I}$. Ultimately, therefore, the kinetic energy of the moving with the velocity p relative to $S$, becomes

$$
T=m c^{2}(\gamma-1)=m c^{2}\left[\left(1-\beta^{2}\right)^{-\frac{1}{2}}-\mathrm{r}\right],
$$

or, developed in a series,

$$
T=\frac{1}{2} m p^{2}\left(\mathrm{r}+\frac{3}{4} \beta^{2}+\frac{5}{8} \beta^{4}+\ldots\right) .
$$

For small velocities, this reduces sensibly to the first tern which is the classical value of the kinetic energy, since in tl the rest-mass becomes sensibly identical with its $S$-value.

The above expression of the kinetic energy was first $g$ Einstein's fundamental paper of 1905 . An alternative, rem: form of (23), due to Minkowski, is

$$
T=m c^{2} \frac{d t-d \tau}{d \tau}
$$

and reads as follows : the kinetic energy of a particle, as es from the $S$-standpoint, is the product of its rest-mass by the of the light-velocity and by the proportionate gain of the with respect to the particle's proper time.

Let us now consider the vector part of the quaternionic e of motion, or the first of $(17 a)$. This, which holds also for a rest-mass, may be read in the usual way: rate of change of mor $=$ force. Then the momentum of the particle, of rest-mass $m$,

$$
\mathbf{G}=m \gamma \mathbf{p}=\sqrt{\sqrt{\mathrm{I}}-\mathrm{p}^{2} / /^{2}} \mathbf{p}
$$

Thus, to obtain the momentum we have to multiply the o velocity p of the particle by $m \gamma$, and not by $m$. Some auth therefore, $m \gamma$ the 'ordinary mass' of the particle. But w rather to avoid so many different names. It is quite suffic know that $m$, the rest-mass, enters in a certain way into the exp of momentum, and in a certain way into that of kinetic The momentum-quaternion, which is always a physical quaterni simply be $m Y$.

Next, to see the propertics of $m$ with respect acceleration $\mathrm{a}=d^{2} \mathrm{r} / d t^{2}$, return once more to th constant $m$, and write the first of ( $17 a$ )

$$
m \frac{d \dot{\mathbf{r}}}{d t}=\frac{m}{\dot{\gamma}} \ddot{\mathbf{r}}=\mathbf{N} .
$$

Then, by (12),

$$
m \gamma \epsilon^{2} a=\mathbf{N},
$$

where, it will be remembered, é is a streteher tangentially to the path. Thus, the force, though in the osculating plane, will, in general, differ in d acceleration. Instead of the old 'mass,' which wa factor converting the acceleration a into the force the linear vector-operator

$$
m \gamma \cdot \epsilon^{2} .
$$

Or, splitting the acceleration into its tangential longitudinal and transversal) components, $a_{u}, a_{u}$,

$$
m \gamma^{3} \cdot \alpha_{x}=N_{r r}, \quad m \gamma \cdot a_{y}=N_{y} .
$$

This result is expressed by saying that the particle h mass

$$
m_{l}=m \gamma^{3}=\begin{gathered}
\\
\sqrt{ }\left(\mathrm{I}-\beta^{3}\right)^{3}
\end{gathered}
$$

and the transversal mass

$$
m_{t}=m \gamma{ }^{m} \sqrt{1}-\beta^{3} .
$$

For vanishing velocities both of these masses beco the rest-mass of the particle. With increasing wh tudinal mass increases more rapidly than the trans $p=c$ both would become infinite. So also would th of the particle increase beyond all limits when the is approached.

It is worth noticing here that the above $m_{l}$ and, velocity of motion in exactly the same way as the the transversal electromagnetic masses of a Lorent;

## DYNAMICS OF A PARTICLE

formula of Lorentz for the transversal electromagnetic mass is fairly well verified by experiments on electrons constituting the $\beta$ In the early stage of such experimental research other elect formulae coincided equally well with the observed facts. It been argued therefore that the zehole mass of the electron is of $p$ electromagnetic origin. Now, the above relativistic formulae, $g$ the required dependence on velocity, have nothing electromag about them. If, therefore, the doctrine of relativity is accepted part of the observed mass of the electron may be attributed non-electromagnetic origin. To obtain this we have only to to the electron, instead of the usual $10^{-13} \mathrm{~cm}$., a correspond greater radius, reducing thus its electromagnetic mass. Remer that what is given by observation is the total mass and the charge of an electron, while its dimensions remain free, in very limits at least. But this subject cannot profitably be discussed any further.

The longitudinal and the transversal masses of a particle, de as the quotients of the corresponding components of force acceleration, may also be written, by (24), in terms of the abs value $G$ of the momentum,

$$
m_{l}=\frac{G}{p}, \quad m_{l}=\frac{d G}{d p} .
$$

The first of these is simply (24) itself, and to see the truth of second, we have only to remember that $d \gamma / d p=\gamma^{3} \phi / c^{2} . *$ formulae (27) would even continue to be true if we had in expression of the momentum, instead of the factor $m \gamma$, any function of $\beta$ alone, as the reader may easily prove for himse

Let us once more return to the first of equations ( $\mathrm{r} 7 a$ ), w may be written

$$
\frac{d \mathbf{G}}{d t}=\mathbf{N} .
$$

Multiply it on both sides vectorially by $r$. Remember that momentum coincides in direction with the velocity $\mathrm{p}=d \mathrm{r} / d$ that $\mathrm{VpG}=0$. Then the result will be

In words: The rate of change of the moment of m in absolute value and direction, to the moment force, both moments being taken about $O$, the oris the relativistic equivalent of what is known in clas: the principle of areas. The above moment of mome of the rest-mass $m$, and $\mathbf{r}, \dot{\mathbf{r}}$,

$$
m \vee r G=m \vee r \frac{d \mathbf{r}}{d \tau} .
$$

In particular, if the moment VrN is permanem impressed force is central, we have the equivalent of conservation of areas, that is, $m$ being again supp

$$
\operatorname{Vr} \frac{d \mathbf{r}}{d \tau}=\mathbf{A},
$$

where the vector $\mathbf{A}$ is constant both in size and in to the frame-work of reference $S$. In this case the a plane, normal to $\mathbf{A}$, as it would also accordi mechanics. But there is the following clifference. usual polar coordinates $r, \theta$, we have, by the las

$$
r^{2} \frac{d \theta}{d \tau}=A,
$$

that is to say, equal areas swept out by the radiu intervals of the proper time of the particle, and $n$ Using the time $t$ of the observers fixed in $S$ we

$$
r^{2} d \theta=A \sqrt{1}-f^{2} / c^{2}
$$

and this is variable, unless the particle happens to along its orbit. Such then is the relativistic modifi second law, valid for any eentral forces. For slo back, of course, to the ordinary conservation of

Leaving, for the present, any further dynamical ( close this chapter by developing some simple and of certain combinations of physical quaternions, inc particular meaning. These will be found useful

## PHYSICAL QUATERNIONS

quaternions, obtained by their addition and multiplication, ca used for relativistic purposes, that is to say, for writing equations which will satisfy the principle of relativity?

We need not dwell upon the sum $a+b+d+\ldots$ (or $a_{c}+b_{c}+d_{c}+$ which is again a physical quaternion, in the original sense of word, and as such, is relativistically available. But having menti the sum at all, it may be good to observe that a sum of antiva quaternions,* as, for example,

$$
a+b_{c},
$$

cannot be used. For not only is this sum not covariant with $q$; with $q_{c}$, but, when subjected to the Lorentz transformation, split, the two addends being torn asunder, thus

$$
a^{\prime}+b_{c}^{\prime}=Q a Q+Q_{c} b_{c} Q_{c}
$$

In other words, such a sum is not transferred as a whole from legitimate system of reference to another.

Now for the product of physical quaternions. Begin with simplest case of two factors. Leave aside $a b$ which is split in act of transformation, thus

$$
a^{\prime} b^{\prime}=Q a Q^{2} b Q,
$$

and pass straight on to the product of antivariant factors, say,

$$
H=a_{c} b .
$$

Pass from the system $S$ to any other legitimate system $S^{\prime}$. $H^{\prime}=Q_{c} a_{c} Q_{c} \cdot Q^{b} Q$, whence, by the associative property, and membering that $Q_{c} Q=1$,

$$
H^{\prime}=Q_{0} H Q
$$

Thus, the new quaternion $H$, though it is neither covariant with standard $q$ nor covariant with $q_{c}$, is transformed as a zehole ( posed of constituents already admitted) and can therefore be for relativistic purposes. A moment's reflection will convince reader that such a procedure will not infringe against the prin of relativity. And the meaning of these abstract remarks
become plainer when we come, in the next cha a concrete law involving a magnitude which, in pas: is transformed exactly as the above quaternion $H$. us look for some further properties of that quate

Consider $H_{c}$, the conjugate of $H$. This will be, rule of the conjugate of a product, ( uat. 8,

$$
H_{\mathrm{c}}=l_{\mathrm{c}} c r .
$$

Now, transforming this, we get $H_{c}^{\prime}=Q_{c} b_{c} Q_{c} Q_{a} Q$, same way as above,*

$$
H_{c}^{\prime}=Q_{0} H H_{c} Q
$$

Thus we see that

$$
Q_{a}[] Q
$$

is the relativistic transformer of both $H$ and its a hence also of their sum and of their difference scalar and of the vector parts of the quaternion $h$ and L ,

$$
s=\mathrm{S} / t, \quad \mathrm{~L}=\mathrm{V} / H .
$$

Now, $s$ being a scalar, we have simply

$$
s^{\prime}=Q_{a} s Q=s Q_{a} Q=s
$$

i.e. $s$ is an invariant, as was proved before. Thu of $a_{c} b$ need not detain us any further.

What we really need for the subsequent physical the vector part of this quaternion. 'This is trans

$$
\mathrm{I}^{\prime}=Q_{\mathrm{o}} \mathrm{~L}(,
$$

and since $Q, Q_{c}$ are unit quaternions, the tensor of 'TL' = 'TL, which may also be written, more conv

$$
I^{\prime 2}=I^{2 \prime} .
$$

[^48]
## PHYSICAL BIVECTORS

These being the transformational properties of the vector $\mathbf{L}=$ let us see what is its structure.

Since both $a$ and $b$ have the structure of $q$, the standa physical quaternions, write

$$
\begin{gathered}
a=\iota \alpha+\mathbf{A} ; \quad \therefore a_{c}=\iota \alpha-\mathbf{A} \\
b=\iota \beta+\mathbf{B},
\end{gathered}
$$

where $\alpha, \beta$ are real scalars and $\mathbf{A}, \mathbf{B}$ ordinary, i.e. real vectors.

$$
\mathbf{L}=\mathbf{L}_{1}-\iota \mathbf{L}_{2},
$$

where $L_{1}$ and $L_{L_{2}}$ are the real vectors

$$
\mathbf{L}_{1}=V B \mathbf{B A}, \quad \mathbf{L}_{2}=\beta \mathbf{A}-\alpha \mathbf{B} .
$$

Thus, $L$ is a complex vector or a bivector,-called so, sir consists of two ordinary vectors. We had, in Chap. II., a sa of such a magnitude in the electromagnetic bivector. The co invariant, (34), of $L$ splits into its tze re real invariants,

$$
L_{1}{ }^{2}-L_{2}^{2} \quad \text { and } \quad\left(\mathrm{L}_{1} \mathrm{~L}_{2}\right) .
$$

The second of these invariants vanishes, since, by (36), perpendicular to $\mathrm{L}_{2}$. This being the case, $\mathrm{L}=\mathrm{V} a_{c} b$ is a bivector (and is equivalent to Sommerfeld's 'special six-vec In order to obtain the general bivector, whose two real v are mutually independent, we have only to add to the abo another, appropriate, special bivector having the same transf tional properties. For this purpose we can take the special bi $\mathrm{L}^{(s)}$, the supplement of L , defined by $\mathrm{L}^{(s)}=\mathrm{V} a_{c}^{(s)} l^{(s)}$, where $a^{(s)}$, $b$ pair of physical quaternions, such that

$$
\mathrm{S} a^{(s)} a_{\mathrm{c}}=\mathrm{S} a^{(s)} b_{c}=\mathrm{S} b^{(s)} a_{c}=\mathrm{S} b^{(s)} b_{c}=0 .
$$

But particulars concerning the choice of a sufficiently $g$ supplement, as this is, need not detain us here.

Henceforth we shall denote by $\mathbf{L}$ the general bivector, obtainable. And we shall call it, where it will be needed fo sake of distinction, a left-handed bivector, owing to the posit
general right-handed bivector, $\mathbf{R}$, consisting of vectors $\mathbf{R}_{1}, \mathbf{R}_{2}$. This will be transformed by $Q[$

$$
\mathbf{R}^{\prime}=Q \mathbf{R} Q_{c}
$$

and will, therefore, have again the two real inve

$$
R_{1}^{2}-R_{2}^{2} \quad \text { and } \quad\left(\mathbf{R}_{1} \mathbf{R}_{2}\right) .
$$

Both $L$ and $R$ can be used, with equal convenic purposes, and will be found useful for the tre magnetic questions.

To illustrate the above properties by a simple ki take, as the generating factors, the velocity- anc quaternions of a particle. Then

$$
\mathrm{L}=\mathrm{V} Y_{c} Z=-\mathrm{V} \ddot{\mathrm{r}} \ddot{\mathrm{r}}+c c(\ddot{t} \ddot{\mathrm{r}}-\ddot{t} \ddot{\mathrm{r}}),
$$

i.e., after a slight calculation, in terms of the ordir acceleration $\mathbf{a}$,

$$
\mathrm{L}_{1}=\gamma^{3} \operatorname{Vap}, \quad \mathrm{~L}_{2}=-c \gamma^{3} \mathrm{a} .
$$

Thus, besides $\left(\mathrm{L}_{1} \mathrm{~L}_{2}\right)$ which vanishes, obvious invariant ( $L_{1}{ }^{2}-L_{2}{ }^{2}$ and, therefore, also)

$$
\frac{1}{c} \sqrt{L_{2}^{2}-L_{1}^{2}}=a \gamma^{3} \sqrt{r-\beta^{2} \sin ^{2}(\mathbf{p},}
$$

and this invariant has a simple kinematical me identical with the absolute value of the rest-acc particle, as given by (I4).

Returning to our general topic, let us consider number of left-handed bivectors. Then we shall $s$ transforming it, all the internal $Q$ 's and $Q_{c}$ 's, as one another $\left(Q Q_{c}=I\right)$, and what is left is or beginning and the $Q$ at the end of the whole cha single $I$. In other words, the vector part of the pro of left-handed bivectors is again a left-handed bi we see, by (38), that the vector part of the produ bivectors is again a right-handed bivector. But wo physical application of such products.

## PHYSICAL QUATERNIONS

Notice, therefore, that if $a$ be any physical quaternion cove with $q$ (not necessarily that already involved in $\mathbf{L}$ or $\mathbf{R}$ ), the pr $a \mathrm{~L}$ will transform into

$$
a^{\prime} \mathbf{L}^{\prime}=Q a Q Q_{c} \mathbf{L} Q=Q a \mathbf{L} Q
$$

that is to say, $a \mathrm{~L}$ will agrain be covariant zerith $q$. So also wi be covariant with $q$. And similarly will $\mathbf{L} a_{c}$ and $a_{c} \mathrm{R}$ be cov with $q_{c}$. In short symbols,

$$
\begin{gathered}
a \mathbf{L} \sim \mathbf{R} a \sim q \\
\mathbf{L} a_{c} \sim a_{c} \mathbf{R} \sim q_{c}
\end{gathered}
$$

Each of these products can be, and is, in fact, used for relat purposes. As regards their structure, they are biquaternio Hamilton's (not in Clifford's) sense of the word, that is t quaternions, of which both the scalar and the vector parts are plex.* But, as we shall see in the next chapter, any one of biquaternions can be split into a pair of our original pl quaternions, each $\simeq q$ or $\simeq q_{c}$ in the case of (40) or (40a) r tively. In this way we fall back to the quaternions conside the outset.

Thus, the product of any number of antivariant physica ternions

$$
\ldots a b_{c} d e_{c} \ldots
$$

will furnish us (after the rejection of the invariant scalar bivector L or R , which is transformed by $Q_{c}[] Q$ and $Q[$ respectively, or again, biquaternions consisting of pairs of p physical quaternions, which are transformed by $Q[] Q$, $Q_{c}[] Q_{c}$.

And, as was already remarked, products of covariant factor: as $a b$, are out of question.

As concerns the operation of division by a physical quat we know that it is reduced to multiplication by its reciprocal. it will be enough to observe that the reciprocal of a $p$ quaternion is again a physical quaternion. For we have
and the tensor $\mathrm{T} a$ is a relativistic invariant. Noti are mutually antivariant.

Finally, notice that any one of the above factors by the quaternionic differential operator

$$
D=\frac{\partial}{\partial l}+\nabla \simeq q,
$$

or by its conjugate $D_{c}$, which is $\simeq q_{c}$. Thus, fo quaternion $\Phi \simeq q$ be a function of time and the $\mathrm{V} D_{c} \Phi$ will be a left-handed bivector ; and so also right-handed bivector. For, independently of the power, these operators behave with respect to th formation exactly as any of our primary quaternic

## CHAPTER VIII.

## FUNDAMENTAL ELECTROMAGNETIC EQUATIONS

Is this chapter we shall consider, from the relativistic standp the fundamental, or 'microscopic,' equations of the electron th and their consequences. These equations, written in their ord vector form, are, as under (r.) and (ir.), Chapter II.,

$$
\left.\begin{array}{l}
\frac{\partial \mathbf{E}}{\partial t}+\rho \mathbf{p}=c \cdot \operatorname{curl} \mathbf{M} ; \quad \rho=\operatorname{div} \mathbf{E} \\
\partial \mathbf{M} \\
\partial t=-c \cdot \operatorname{curl} \mathbf{E} ; \quad \operatorname{div} \mathbf{M}=0
\end{array}\right\}
$$

and

$$
\mathbf{P}=\rho\left[\mathbf{E}+{ }_{c}^{\mathrm{I}} \mathrm{VpM}\right] \equiv \rho \text { 我. }
$$

Here, $\mathbf{p}$ is the velocity of a charge-clement with respect to framework $S$, for which, to begin with, the equations are supp to be rigorously valid; $\mathbf{P}$ is the ponderomotive force, per volume, and df the ponderomotive force por unit charge, of electric force.

First of all, we have to ask whether these equations satisfy principle of relativity, that is to say, whether they preserve form when we pass from the system $S(t, x, y, z)$ to another sy $S^{\prime}\left(t^{\prime}, x^{\prime}, y^{\prime}, z^{\prime}\right)$ moving with uniform velocity relatively to $S$. if the answer be, as it is in fact, in the affirmative, what ar connexions between $\mathbf{E}^{\prime}, \mathbf{M}^{\prime}$, the dielectric displacement and

Lorentz transformation $x=\gamma_{v}\left(x^{\prime}+v t^{\prime}\right)$, etc., and expressing of $\mathbf{p}^{\prime}$ by means of his addition theorem of velocities, sho variance of the form of these equations, and finally gathe the terms which in the transformed equations play the pa field-vectors.* But the shortest method to obtain these to write the four equations (1.) in their condensed qu form,

$$
D \mathrm{~B}=C,
$$

as given in Chap. II., and to test the constituents of this with regard to their relativistic qualities.

Here, it will be remembered, $\mathbf{B}=\mathbf{M}-\iota \mathbf{E}$, while

$$
C=\rho\left[\iota+\frac{\mathrm{I}}{c} \mathbf{p}\right],
$$

or, in terms of the velocity-quaternion, (3a), Chap. VII.,

$$
C=\frac{\rho}{c \gamma_{p}} Y,
$$

where $\gamma_{p}=\left(\mathrm{r}-p^{2} / c^{2}\right)^{-\frac{1}{2}}$.
Keeping this in mind, consider the equation (r). already that the differentiator $D$ behaves exactly as a quaternion, viz. that $D \simeq q$. The only thing, therefore, require, is to find the nature of the current-quaternion $C$

Now, the electric charge de of any individual portion of a is a relativistic invariant, i.e. if $d S$ be the volume of tha and $d S^{\prime}$ its $S^{\prime}$-correspondent, then

$$
\rho d S=\rho^{\prime} d S^{\prime} .
$$

In fact, taking the divergence of the first of (r.), we hav

$$
\circ=\frac{\partial \rho}{\partial t}+\operatorname{div}(\rho \mathbf{p})=\frac{\partial \rho}{\partial t}+(\mathbf{p} \nabla) \rho+\rho \operatorname{div} \mathbf{p},
$$

which is known as 'the equation of continuity, or, denot the rate of individual change, as on p. 3 r ,

## CHARGE AND CURRENT

whence, multiplying by $d S$ and observing that $\frac{d}{d t}(d S)=d S$. di

$$
\frac{d}{d t}(\rho d S)=\frac{d}{d t}(d e)=0
$$

Thus, the charge, as estimated from the $S$-standpoint, is invar in time, notwithstanding the motion and deformation of the vol element we are watching. This being the case, we can imagin charge first fixed in $S$ and then set it into motion, bringing it by by to the velocity $\mathbf{v}$, when it will be at rest in $S^{\prime}$. Claiming, t fore, in the name of the principle of relativity, the same rights $f$ as for $S$, we shall have $d e^{\prime}=d e$. (If the reader does not like kind of proof, he can simply postulate the invariance of charge verify a posteriori, after having obtained $\mathbf{E}^{\prime}$ in terms of $\mathbf{E}, \mathbf{M}$, this postulate is fulfilled.)

On the other hand, remembering that volumes are transfo in the same way as longitudinal dimensions, and denoting for moment by $d S_{0}$ the rest-volume of the element considered shall have

$$
d S=d S_{0} \sqrt{\mathrm{~J}-p^{2} / c^{2}} \quad \text { and } \quad d S^{\prime}=d S_{0} \sqrt{\mathrm{r}-p^{\prime 2} / c^{2}}
$$

or
Therefore, by (3),

$$
\gamma_{p} d S=\gamma_{p^{\prime}} d S^{\prime}
$$

$$
\frac{\rho}{\gamma_{p}}=\frac{\rho^{\prime}}{\gamma_{p^{\prime}}},
$$

that is to say, $\rho / \gamma_{p}$, the coefficient of $Y$ in $(2 a)$, is an invaria
Now, as we know from the last chapter, $Y$ is a physical quater Therefore, $C$, the current-quaternion, as it was already calle Chapter II., is again a physical quaternion, like the standard

$$
C \simeq q,
$$

as'well as $D \simeq q$.
This proves the invariance of the form of the equation (1 of the equations (r.), with respect to the Lorentz transformation gives at the same time the connexion between $\mathbf{B}^{\prime}$ and $\mathbf{B}$.
and inserting $Q Q_{c}=1$ between $D$ and $\mathbf{B}$,

$$
\begin{aligned}
D^{\prime} Q_{c} \mathbf{B} Q & =C^{\prime}, \\
D^{\prime} \mathbf{B}^{\prime} & =C^{\prime},
\end{aligned}
$$

i.e.
where $\mathbf{B}^{\prime}=Q_{c} \mathbf{B} Q$.* Thus, $\mathbf{B}$, the electromagnetic bivect left-handed bivector.

Or, to obtain this bivector in its typical form $\mathrm{V} a_{c} b$, proceed as follows. Operate on both sides of ( I ) with $D_{c}$.

$$
D_{c} D \mathbf{B}=D_{c} C .
$$

But $D_{c} D$ is an invariant. This, therefore, is already the form. We need not even put the prefix V before $D_{\mathrm{c}}$ $\mathrm{S} D_{c} C=0$, as we shall see when we next return to the last e

Thus, B is a left-handed bivector, having the same struct the same transformational properties as our L of the last Henceforth we can consider it as the standard of physical in the same way as $q$ has been the standard of physical quat

It will be found convenient for subsequent work to write thr $\mathbf{I}$ (instead of our previous $\mathbf{B}$ ) for the electromagnetic bivecto

$$
\mathbf{L}=\mathbf{M}-\iota \mathbf{E} .
$$

The quaternionic equivalent of the electromagnetic differential e (土.) zeill now be

$$
D \mathrm{~L}=C,
$$

and the transformation formula of the electromagnetic bivector

$$
\mathrm{I}^{\prime}=Q_{c} \mathrm{~L} Q .
$$

The invariance of the formula (ir.) for the ponderomoti will, with equal ease, be proved later on. Meanwhile let $u$ attention upon (5).

As already pointed out in the last chapter, $Q$ and unit quaternions, the square of the electromagnetic bivect invariant, i.e.

$$
\mathrm{L}^{\prime 2}=\mathrm{L}^{2} .
$$

## ELECTROMAGNETIC BIVECTOR

Now, by (4),

$$
-\mathbf{L}^{2}=M^{2}-E^{2}-2 \iota(\mathbf{E M}),
$$

and similarly for $\mathbf{L}^{\prime \prime}$. Thus we have the two real invariant

$$
\frac{1}{2}\left(M^{2}-E^{2}\right) \quad \text { and } \quad(E M) .
$$

The first of these invariants, the difference of the densities magnetic and the electric energies, is the electromagnetic Lagr function per unit volume.* The second invariant, the scalar p of $\mathbf{E}$ and $\mathbf{M}$, has no particular name of its own. Notice tha is called a pure electromagnetic wave is defined by $M^{2}=Z$ $(E M)=0$. In words: energy half electric and half magneti $\mathbf{E}$ and $\mathbf{M}$ perpendicular to one another. Using the electroma bivector we can characterize pure waves more shortly by $\mathrm{L}^{2}=1$ At the same time we see that a wave which is pure from the $S$ point is equally pure from the $S^{\prime}$-point of view. In short, pu least in this domain of relations, is an invariant property. B only by the way.

Next, to develop (5) into its vectorial form, remember that, $b$ Chap. V.,

$$
Q=\cos \frac{\omega}{2}+u \cdot \sin \frac{\omega}{2},
$$

where $\mathbf{u}$ is the unit of $\mathbf{v}$, the velocity of $S^{\prime}$ relative to $S$, and $\mathbf{w}$ is the imaginary angle previously defined. Multiply out th side of (5). Then

$$
\mathbf{I}^{\prime}=(\mathrm{I}-\cos \omega) \cdot \mathbf{u}(\mathrm{uL})+\cos \omega \cdot \mathrm{I}+\sin \omega \cdot V L u .
$$

From this intermediate form we can easily see that $\mathrm{L}^{\prime}$ is ob from L by turning it about u , the axis of the guaternion $Q$, ther the doutble of the angle of that quaternion. Such then is the of the operator $Q_{c}[\quad] Q$. This is only a particular insta a theorem of the calculus of quaternions, given by Ha himself. $\dagger$

[^49]But let us write the last formula in terms of $\gamma$, which is a abbreviation for $\gamma_{v}=\left(\mathrm{x}-\beta^{2}\right)^{-\frac{1}{2}}, \beta=\tau^{\prime} / c$. Remembering that $\cos \omega=$ and $\sin \omega=\iota \beta \gamma$, we have

$$
\mathrm{L}^{\prime}=(\mathrm{r}-\gamma) \mathfrak{u}(\mathrm{uL})+\gamma \mathrm{L}+\frac{\iota}{c} \gamma \mathrm{VLv},
$$

or, employing again the longitudinal stretcher $\epsilon$, of ratio $\gamma$,

$$
\begin{equation*}
\mathbf{L}^{\prime}=\gamma\left[\frac{\mathrm{I}}{\epsilon} \mathbf{L}+\frac{\imath}{c} \mathrm{~V} \mathbf{L} \mathbf{v}\right], \tag{7}
\end{equation*}
$$

and splitting into the real and the imaginary parts, according to (4),

$$
\left.\begin{array}{l}
\mathbf{E}^{\prime}=\gamma\left[\frac{\mathrm{I}}{\epsilon} \mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{VVM}\right] \\
\mathbf{M}^{\prime}=\gamma\left[\frac{\mathrm{I}}{\epsilon} \mathbf{M}-\frac{\mathrm{I}}{c} \mathrm{VVE}\right] .
\end{array}\right\}
$$

Or, finally, in Cartesian expansion, using $1,2,3$ for the rectangula components of the vectors taken along the direction of motion an perpendicular to it (right-handed system of axes),

$$
\left.\begin{array}{lll}
E_{1}^{\prime}=E_{1}, & E_{2}^{\prime}=\gamma\left(E_{2}-\beta M_{3}\right), & E_{3}^{\prime}=\gamma\left(E_{3}+\beta M_{2}\right)  \tag{7}\\
M_{1}^{\prime}=M_{1}, & M_{2}^{\prime}=\gamma\left(M_{2}+\beta E_{3}\right), & M_{3}^{\prime}=\gamma\left(M_{3}-\beta E_{2}\right) .
\end{array}\right\}
$$

These are the relativistic formulae for the transformation of th electric and the magnetic vectors, as obtained by Einstein. The agree entirely with those given by Lorentz in his modificed theor (see p. 86). Notice that, in passing from the $S$ - to the $S^{\prime}$-standpoin the longitudinal components of $\mathbf{E}, \mathbf{M}$ remain unchanged, while th changes brought about in their transversal components involve th vector products VV M and VEv and the coefficient $\gamma$.

Multiplying both sides of (5) by $Q$ as a prefactor and by $Q_{c}$ as postfactor, we have at once

$$
\mathrm{L}=Q \mathrm{~L}^{\prime} Q_{c} .
$$

But $Q_{c}$ follows from $Q$, and zice versa, by a mere change of the sig of $\mathbf{v}$. Thus, the inverse transformation, giving $\mathbf{E}, \mathbf{M}$ in terms o

## CONVECTIVE FIELDS

as the reader may also prove by solving $(7 b)$. This shows more that none of the systems of reference is privileged.

The invariance of electric charge, used at the outset, can no directly verified by differentiation of the transformed electric $v$ or of its components.*

The applicability of the above formulae of transformatic obvious. For, if we know a solution of the electrömagnetic $\delta$ ential equations for one of the legitimate systems of reference can deduce from it at once the solution for any other of systems. Now, the problem of integration may be much easie one of these systems than for any other, owing to some parti simplicity of the conditions as stated from the standpoint of former system. Whence the advantage of the method. $\dagger$

The simplest solution of the electromagnetic equations i electrostatic field corresponding to a given distribution of ch (electrons), which are all fixed with respect to a legitimate fr work, say $S^{\prime}$. The $S$-correspondent of this will be the ele magnetic field accompanying a system of electrons in un? translational motion, with velocity v relative to $S$, or what is cal convective field. 'The framework $S^{\prime}$ will be the rest-system belor permanently to these charges. It will be good, before procee further with our general subject, to consider this example at length.

Let us suppose, therefore, that we have in $S^{\prime}$ a purely electro: field, so that $\mathbf{E}^{\prime}=-\nabla^{\prime} \phi^{\prime}$, where $\phi^{\prime}$ is the scalar potential of the distribution of charge, while $\mathbf{M}^{\prime}=0$. Then, remembering tha inverse of the first of $(7 a)$ is

$$
\mathbf{E}=\gamma\left[\frac{\mathbf{I}}{\epsilon} \mathbf{E}^{\prime}-\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{V M}^{\prime}\right]
$$

we shall have, from the $S$-point of view,

$$
\mathbf{E}=\gamma \epsilon^{-1} \mathbf{E}^{\prime},
$$

i.e., in Cartesians,

$$
E_{1}=E_{1}^{\prime}, \quad E_{2}=\gamma E_{2}^{\prime}, \quad E_{3}=\gamma E_{3}^{\prime} .
$$

The second of ( $7 a$ ) gives us at once $\mathbf{M}$ in terms of $\mathbf{E}$,

$$
\mathbf{M}={ }_{c}^{\epsilon} \mathrm{VvE}=\frac{\mathrm{I}}{{ }_{c}} \mathrm{~V} \mathbf{V E},
$$

since the stretcher $\epsilon$ acts along $\mathbf{v}$, while the vector produc to V .

Thus we have for the most general convective field, acc any system of charges which moves as a whole with $t$ l translational velocity v relative to $S$,

$$
\left.\begin{array}{r}
\mathbf{E}=\gamma \epsilon^{-1} \mathbf{E}^{\prime} \\
\mathbf{M}=\frac{r}{c} V \mathbf{V} \mathbf{E} .
\end{array}\right\}
$$

Here $\mathbf{E}^{\prime}=-\nabla^{\prime} \phi^{\prime}$, the scalar function $\phi^{\prime}$ being the electrostati of the given distribution of charge fixed in $S^{\prime}$. The therefore reduced to finding, for each particular case of $d$ the scalar potential $\phi^{\prime}$. Observe that this is the poten while $\mathbf{E}$ has no such potential. Notice, further, that the lines, due to the motion of charges, are everywhere norm $\mathbf{E}$ and the direction of motion. And since $\mathbf{E}^{\prime}$ is coplanar the magnetic lines are also at right angles to $\mathbf{E}^{\prime}$.

The gradient or slope $\nabla^{\prime} \phi^{\prime}$ can casily be replaced by fact, measuring $x$ along the direction of motion, so that $x^{\prime}=$ and remembering that, by assumption, 'o $\delta \phi^{\prime} / \partial t^{\prime}=0$, we have i.e.

$$
\frac{\partial \phi^{\prime}}{\partial x}=\gamma \frac{\partial \phi^{\prime}}{\partial x^{\prime}}, \quad \frac{\partial \phi^{\prime}}{\partial y}=\frac{\partial \phi^{\prime}}{\partial y^{\prime}}, \quad \partial \phi^{\prime}=\partial \psi^{\prime},
$$

$$
\epsilon \nabla^{\prime} \phi^{\prime}=\nabla \phi^{\prime},
$$

so that the first of (8) can be written

$$
\mathbf{E}=-\gamma \epsilon^{-2} \nabla \phi^{\prime} .
$$

Thus, the displacement $\mathbf{E}$, as already remarked, has potential. But the electric force aff, or the ponderomotive unit of charge carried along with $S^{\prime \prime}$, has such a potenti as in Lorentz's treatment, given in Chapter III. p. 8 r .

## CONVECTIVE FIELDS

or

$$
\text { 㮽 }=\gamma^{-2} \epsilon^{2} \mathrm{E} \text {, }
$$

and by our last formula,

$$
\mathrm{dt}=-\nabla\left(\frac{\phi^{\prime}}{\gamma}\right) .
$$

Thus $\phi^{\prime} / \gamma$ is the scalar potential of the electric force. This is convection potential of Chap. III., the above equation being ident with formula (2I) of that chapter, in which $\phi$ was $\gamma \phi^{\prime}$. The sa result may be deduced more directly from the transformatio properties of the ponderomotive force, to be developed later o

Since $\gamma$ is constant throughout $S^{\prime}$, the surfaces of constant vection potential and those of constant $\phi^{\prime}$ overlap. We see, theref that the lines of electric force ff (but not those of displacement cut perpendicularly the surfaces of constant electrostatic poter of the rest-system, $\phi^{\prime}=$ const. The electric force and displacem of that system are identical, of course, i.e. $\mathbb{f f}^{\prime}=\mathbf{E}^{\prime}$.
' $o$ illustrate the general formulae (8) of the convective fi suppose that the distribution of electric charge in $S^{\prime}$ is in homogene concentric spherical sheets round $O^{\prime}$, the origin of the coordin or the origin of the vectors $r^{\prime}$. Then $\phi^{\prime}$, and consequently also will be functions of $r^{\prime}$ alone, and the lines of displacement in will be straight and radial or, say,

$$
\mathbf{E}^{\prime}=f\left(r^{\prime}\right) \cdot \mathbf{r}^{\prime}
$$

where $f$ is a scalar function of its argument. By the fundame formulae of transformation, $\mathbf{r}^{\prime}=\epsilon \mathbf{r}-\mathrm{v} \gamma t$. Now, since the whole fi together with the charges, moves past $S$ without being deformec is enough to consider it at one single instant. Let this be instant $t=0$, when $O^{\prime}$ coincides with $O$, the origin of the $S$-coordin or of all vectors r . Then

$$
\text { and, by (8), } \left.\begin{array}{l}
\mathbf{r}^{\prime}=\epsilon \mathbf{r}, \\
\mathbf{E}=\gamma f\left(r^{\prime}\right) \cdot \mathbf{r} \\
\mathbf{M}=\frac{\mathrm{r}}{-} \gamma f\left(r^{\prime}\right) \cdot \mathrm{Vvr},
\end{array}\right\}
$$

The whole electromagnetic field is symmetrical, of course, roun this longitudinal axis. Since $\mathbf{r}=\epsilon^{-1} \mathbf{r}^{\prime}$, or

$$
x=\frac{x^{\prime}}{\gamma}=x^{\prime} \sqrt{1-}-\beta^{2}, \quad y=y^{\prime}, \quad z=z^{\prime},
$$

the spheres $r^{\prime}=$ const. become, in $S$, oblate ellipsoids of revolution as in the FitzGerald-Lorentz contraction, i.e. having

$$
\begin{aligned}
& r^{\prime} \\
& \gamma^{\prime}
\end{aligned} r^{\prime}, r^{\prime}
$$

for their semiaxes. These are known as Hecviside ellipsoids. Suc then will be the surfaces of constant convection potential, and th lines of electric force (dyt), cutting these ellipsoids at right angle will be parabolic arcs, contained in the meridian planes.

If $s=\left(y^{2}+z^{2}\right)^{\frac{1}{2}}$ be the distance of a point from the axis symmetry, we have

$$
r^{\prime}=\sqrt{\gamma^{2} x^{2}+s^{2}}
$$

or also, denoting by $\theta$ the angle contained between r and the axis,

$$
r^{\prime}=\gamma r \sqrt{1-\beta^{2} \sin ^{2} \theta} .
$$

This is to be substituted in each particular case for the argumer of the given function $f$ in ( I 0 ).

Take, as the simplest case of the above kind, a single sphere homogeneous surface-charge, or a Lorentz electron. Call its res radius $R$ and its total charge $e$ (which, as we know, is the sam thing as $e^{\prime}$ ). Then $E^{\prime}=0$ inside the sphere $r^{\prime}=R$, and consequentl also $E=0$ inside the oblate ellipsoid $\gamma^{2} \alpha^{2}+s^{2}=R^{2}$, while at th surface of and outside the electron*
and therefore

$$
\mathbf{E}^{\prime}=\frac{e \mathrm{r}^{\prime}}{4 \pi r^{\prime 3}}
$$

$$
\mathbf{E}=\frac{e \gamma}{4 \pi r^{\prime 3}} \mathbf{r}, \quad r^{\prime} \equiv R
$$

that is, by (II),

$$
\mathbf{E}=\frac{e}{4 \pi r^{2}} \cdot \frac{1-\beta^{2}}{\left(1-\beta^{2} \sin ^{2} \theta\right)^{\frac{3}{2}}} \mathbf{r}
$$

## ELECTROMAGNETIC MASSES

tional to the square of the distance from the centre of the elect The unit tubes of displacement, though everywhere radial, crowded towards the equator, and the more so, the greater velocity of motion. At any given distance $r$, the density of tubes at the equator is greater than that at the poles $(\theta=0$ o in the ratio $E_{\pi / 2}: E_{0}=\mathrm{I}:\left(\mathrm{I}-\beta^{2}\right)^{\frac{3}{2}}$.

From the above, widely known, formulae the longitudinal and transversal electromagnetic masses of the electron may be ea deduced in the usual way. The flux of energy or the Poyn vector being

$$
\text { 挧 }=c \mathrm{VEM}=\mathrm{VEVVE}=E^{2} \mathrm{~V}-(\mathbf{E v}) \mathbf{E} \text {, }
$$

we have for the electromagnetic momentum, per unit volume, (30), Chap. II.,

$$
\mathbf{g}=\frac{v}{c^{2}}\left[E^{2} \mathbf{u}-E_{1} \mathbf{E}\right],
$$

where $u$ is the unit of $\nabla$ and $E_{1}$ the longitudinal component of Integrating through the whole field (from $r^{\prime}=R$ till $r^{\prime}=\infty$ ) taking advantage of its axial symmetry, we obtain, for the t electromagnetic momentum,*

$$
\mathbf{G}=\frac{e^{2}}{6 \pi c^{2} R} \gamma \mathbf{v},
$$

whence the longitudinal electromagnetic mass $m_{l}$ of the electron the transversal one, $m_{t}$, defined by $m_{l}=d G / d v, m_{t}=G / v$ :

$$
m_{l}=m_{0} \gamma^{3}, \quad m_{t}=m_{0} \gamma,
$$

where

$$
m_{0}=\frac{e^{2}}{6 \pi c^{2} R} .
$$

These are the well-known formulae of Lorentz, as mentior previously. They are valid for an electron of homogeneous surfa charge. In the case of volume-charge, we should obtain for electromagnetic momentum $\frac{6}{3}$ of the above value, so that (14) wo continue to hold with $m_{0}$ equal to $\frac{6}{5}$ of the above,

The electromagnetic momentum can, in either case, be $v$

$$
\mathbf{G}=m_{0} \gamma \mathbf{v} .
$$

Thus, $m_{0}$, the electromagnetic rest-mass, plays the same $p a$ rest-mass, of any origin, in the relativistic dynamics of a Cf. (24), Chap. VII.

Having for the present sufficiently illustrated the transfo properties of the electromagnetic bivector, let us now retul general subject.

Consider again the equation

$$
D \mathbf{L}=C
$$

embodying in itself the whole of the electronic differential (r.), and showing at the same time their invariance. Ope both of its sides with $D_{c}$. Then

$$
D_{c} D \mathbf{L}=D_{c} C
$$

But $D_{c} D$ is the Dalembertian,

$$
D_{c} D=\frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}+\frac{\partial^{2}}{\partial z^{2}}-\frac{1}{c^{2}} \frac{\partial^{2}}{\partial t^{2}}=\square
$$

and this is a purely scalar operator ; that is to say, if app scalar it gives a scalar, and if applied to a vector it give vector. Now, $\mathbf{L}$ is scalarless. Therefore

$$
\mathrm{S} D_{c} C=0
$$

This is the equation of continuity. In fact, its developec by (2),

$$
\frac{\partial \rho}{\partial t}+\operatorname{div}(\rho \mathbf{p})=0 .
$$

But this only by the way.
Next, introduce an auxiliary quaternion $\Phi$, satisfying ential equation

$$
\square \Phi=-C
$$

and the supplementary condition

## POTENTIAL-QUATERNION

Now, $D_{c} D=\square$, being the norm of $D \simeq q$, will be an inva as was already remarked on p. 113. Therefore, by (18), $\Phi$ be a physical quaternion, having an imaginary scalar and a vector. Write it, therefore,

$$
\Phi=\iota \phi+\mathbf{A} \simeq q,
$$

and call it the potential-quaternion, since the whole electromag bivector is derived from it by simple differentiation. The sponding world-vector is called the four-potential.

The scalar part of $\Phi$ is $\iota$ times the usual scalar potential, ar vector part is the vector potential. In fact, splitting (20) int real and the imaginary parts, we obtain at once

$$
\begin{aligned}
\mathbf{M} & =V \nabla \mathbf{A}=\operatorname{curl} \mathbf{A}, \\
\mathbf{E} & =-\nabla \phi-\frac{\mathrm{I}}{\mathrm{c}} \frac{\partial \mathbf{A}}{\partial t},
\end{aligned}
$$

while the condition (rg) becomes -

$$
\frac{r}{c} \frac{\partial \phi}{\partial t}+\operatorname{div} \mathbf{A}=0
$$

and these are the familiar formulac of the electron theor employed incidentally in Chapter III., p. 8o. The differ equation (18) splits, of course, into the familiar pair of equat

$$
\square \phi=-\rho ; \quad \square \mathbf{A}=-\frac{\mathrm{I}}{c} \rho \mathbf{p}
$$

identical with (x6), Chap. III.
According to (2r), $\phi$ and $\mathbf{A}$ are transformed as $c t$ and $\mathbf{r}$. for instance, if we have in $S^{\prime \prime}$ a purely electrostatic field, i.e. if then, for the convective field, as estimated from the $S$-standpoir

$$
\phi=\gamma \phi^{\prime}, \quad \gamma=\gamma_{v}
$$

as mentioned above, and

$$
\mathbf{A}=\frac{\mathrm{r}}{c} \boldsymbol{\nabla} \gamma \psi^{\prime}=\frac{\mathrm{r}}{c} \boldsymbol{\nabla} \phi,
$$

as in (19), Chap. III.
So much as regards the potential-quaternion and its relati
which can be called the complementary electromagnetic Then we would have obtained as the condensed equivaler fundamental equations (r.), instead of and in exactly the $s$ as (I. $a$ ),

$$
D_{c} \mathrm{R}=C_{c},
$$

where $C_{c}$ is the conjugate current-quaternion $\rho(\imath-\mathbf{p} / \mathbf{c})$. on both sides of this equation with $D$. Then the result $\square \mathrm{R}=D C_{c}$. And since the Dalembertian is an invariant, at once that R is a right-handed physical bivector,* i.e. tha

$$
\mathbf{R}^{\prime}=Q \mathbf{R} Q_{c} .
$$

Henceforth $\mathbf{R}$ can be considered as the standard of all such $b$ just as $\boldsymbol{L}$ became the standard of the left-handed ones. O the differential equation ( $1 . b$ ) is invariant with respect to the transformation, i.e.

$$
D_{c}^{\prime} \mathbf{R}^{\prime}=C_{c}^{\prime} .
$$

(1.a) and (1.b) differ, of course, only formally from one each, when split, gives the four electromagnetic differential e (r.). Thus, as far as the equations of the field and all th sequences are concerned, we do not need both $L$ and require only one of them at a time.

For some other purposes, however, the simultaneous use bivectors will prove to be very advantageous.

Their symbols, being the initials of 'left' and 'right,' are so as to remind the reader of their transformational pr In connexion with these, $\boldsymbol{L}$ can admit a physical quaterı variant with $q$, only on its left as neighbour, and $\mathbf{R}$ onl right. And vice versa, if the neighbour is covariant with

Now for the outstanding proof of the invariance of th mental formula (II.) for the ponderomotive force. To obtain t we have only to write that formula in terms of legitimate re magnitudes.

If we multiply our left-handed electromagnetic bivector left side, by any physical quaternion $\sim q$, then, as in (40), Ch

[^50]
## PONDEROMOTIVE FORCE

the resulting product will again be transformed like $q$. N current-quaternion $C$ being precisely such a quaternion, consi full product

## $C \mathrm{~L}$.

This then will again be transformed by $Q[\quad] Q$. Develop it and (4). Then the result will be

$$
C \mathbf{L}=F+\iota F_{m},
$$

where

$$
F=\rho\left[\frac{\iota}{c}(\mathbf{p} \mathbf{E})+\mathbf{E}+\frac{\mathrm{r}}{c} \mathrm{~V} \mathbf{p M}\right],
$$

and $F_{m}$, the magnetic analogue* of this,

$$
F_{m}=\rho\left[\frac{\iota}{c}(\mathbf{p} \mathbf{M})+\mathbf{M}-\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{p E}\right] .
$$

Now, the vector part of $F$ is exactly P , the ponderomotive fc unit volume, as given by (II.), and the scalar part of $F$ is $\iota$ the activity of this force. 'Thus,

$$
F={ }_{c}^{l}(\mathrm{Pp})+\mathrm{P} .
$$

Observe that the whole product $C L$, though covariant standard $q$, has not the structure of $q$, since it is a full biqua in the Hamiltonian sense of the word. But $F$, and its n analogue, have each the structure of q, i.e. a real vector imaginary scalar.

Similarly, the complementary $\mathbf{R}$ being a right-handed multiply it on the right side by $C$. Then the product $\mathbf{R} C$ w be transformed by $Q[\quad] Q$. I)evelop it. Then, by (2) al

$$
\mathbf{R} C=--F+\imath F_{n},
$$

with precisely the same meanings of $F$ and $F_{m}$ as above again is a full biquaternion.

Now, since both biquaternions, $C \mathrm{~L}$ and $\mathrm{R} C$ are transfor $Q[\quad] Q$, this will also be the relativistic transformer of their of their difference. Ieave alone the sum, which would
physically uninteresting $F_{m}$, and take half the difference and (24a). This will give

$$
F=\frac{1}{2}[C \mathbf{L}-\mathrm{R} C] .
$$

Thus, we see that $F$ taken by itself (as well as $F_{n n}$ ) is with $q$. And since $F$ has also the structure of $q$, it is a quaternion, and may as such be called the force-quaternion volume. It has a dynamic vector, the ponderomotive force volume, and an energetic scalar, proportional to the a that force.

At the same time we have obtained for $F$ the expressi and we know that the vector part of this is equal to $\mathbf{P}$ by (in.). Now, (ri. $a$ ) transforms into

$$
F^{\prime}=Q F Q=\frac{1}{2}\left[C^{\prime} \mathbf{L}^{\prime}-R^{\prime} C^{\prime}\right]
$$

and the vector part of this quaternion is again

$$
\mathbf{P}^{\prime}=\rho^{\prime}\left[\mathbf{E}^{\prime}+\frac{\mathrm{r}}{c} \mathrm{Vp}^{\prime} \mathbf{M}^{\prime}\right]
$$

which proves explicitly the invariance of the formula (ir.) wit to the Lorentz transformation.

Thus, the whole of 'the fundamental equations for the as (r.) and (ir.) are called, satisfy rigorously the principle of and it was for this reason possible to incorporate them e the new doctrine.

By (25) we have, identically,

$$
\mathrm{SFF} C_{c}=0,
$$

and therefore also, by ( $2 a$ ),

$$
\mathrm{S} F Y_{c}=0 .
$$

In four-dimensional language we should say that the f equivalent to the quaternion $F$, is perpendicular to t current, and consequently also to the world-line of the el electric charge acted upon. We met with this proper treating the dynamics of a particle moving under the act force of any nature whatever. See (2r), Chapter VII.

Remember that what is, in our present case of electro

## PONDEROMOTIVE FORCES

The latter is not the vector part of a physical quaternion. Bi the other hand, we know that

$$
\gamma_{p} \times \text { volume }
$$

is an invariant. Therefore

$$
\gamma_{p} \times \text { vol. } \times F \simeq q,
$$

that is to say, $\gamma_{p}$ times the force-quaternion calculated for any pc of electricity is again a physical quaternion. Such then is the formational property of ponderomotive forces due to an el magnetic field.

Now, if one of these forces is in equilibrium with a for any other origin, from the standpoint of the system $S^{\prime}$ (sc the particle acted upon is at rest with respect to that system), these two forces have also to balance each other when estir from the standpoint of any other legitimate system $S$. For rela to $S$, the particle in question will move uniformly. Hence tl quirement, that ponderomotive forces of any origin shall be transf in exactly the same way as those of electromagnetic origin,* i.e. sc

$$
\gamma_{p}\left[\text { total force }+\frac{\iota}{c} \text { times its activity }\right] \simeq \mathrm{r}+\iota c t .
$$

Here 'total force' means the force acting upon a particle velocity relative to $S$ is p , or upon a body of any dimensions its parts happen to have the same velocity.

Now, what in Chap. VII. has been called the Newtonian $\mathbf{N}$, satisfies exactly this relativistic requirement. In fact, accc to the formula ( $\mathbf{r} 8$ ) of that chapter (where $\gamma$ stands for $\gamma_{p}$ ),

$$
\gamma_{p}(\mathbf{N}+\iota v)=X
$$

is a physical quatemion, and, as we have seen, $\nu=\frac{\mathrm{r}}{c}(\mathbf{N p})$. precisely for this reason that the Newtonian force, not the kowskian, has been considered as the force, and the magi $m c^{2}(\gamma-1)$, whose rate of change has been equal to ( Np ), (kinetic) energy of the particle.

This procedure of transferring the transformational propertie: certain physical magnitudes to others of the same kind

## THE THEORY OF RELATIVITY

After this short digression of a general nature, let us our electromagnetic topic.

The formula (ir. a), obtained above for the force-quat has nothing to do with the differential equations (r.) of th magnetic field. It is only another form of the original for for the ponderomotive force. Now, use those differential in their quaternionic condensation (r. $a$ ), that is to say, $C=D \mathrm{~L}$. Then the double of the force-quaternion will b

$$
{ }_{2} F=D \mathbf{L} \cdot \mathbf{L}-\mathbf{R} \cdot D \mathbf{L},
$$

where the dot stands for a separator, stopping the diffe action of $D$. This formula, when subjected to a sligh somewhat peculiar change, will prove to be very conve further application. The peculiarity of the formal chang to, consists in this, that it requires us to give up an Hitherto, in conformity with the general convention, we ha used the differential operator $D$ as a 'prefactor,' i.e. actin only, just as an ordinary scalar differentiator, such as $\partial / \partial t$ Now, the position of a scalar being a matter of indiff would be utterly useless and extravagant to write $\partial / \partial t$, for behind the scalar or vector function to be differentiated; expressions would mean just the same as $\frac{\partial s}{\partial t}$ or $\frac{\partial v}{\partial t}$. But tl different when the differentiator has the nature of a vect Hamiltonian $\nabla$, or of a quaternion, as $D$. Since the mult of vectors, and more generally of quaternions, is non-com we obviously deprive ourselves of possible advantages if the rôle of quaternionic differential (or other) operators prefactors. Henceforth, therefore, we shall use $D$ as an acting both forzeard and backzeard,* i.e. as both a prefac postfactor, and we shall, for instance, write

$$
\mathbf{R}[D] \mathbf{L}=\mathbf{R} D \cdot \mathbf{L}+\mathbf{R} . D \mathbf{L}
$$

where the dots stop $D$ 's differentiating power, and where the (which could also be omitted) are used for better empha

## PONDEROMOTIVE FORCE

bilateral action of the enclosed operator. The only thing to be explained in this symbolism is the meaning of $\boldsymbol{R} D$, which is unus inasmuch as the operator $D$ follows the operand. Now, if $D$ were ordinary quaternion, that is a quaternionic magnitude, with $s, \nabla$ its scalar and vector parts, we should have, by elementary rule

$$
\mathbf{R} D=\mathbf{R} s+V \mathbf{R v}-(\mathbf{R v})=s \mathbf{R}-\mathrm{V} \mathbf{v} \mathbf{R}-(\mathbf{v R}) .
$$

Writing therefore $\partial / \partial l$ instead of $s$ and $\nabla$ instead of $\nabla$, the pl meaning of $\mathrm{R} D$ will be

$$
\mathbf{R} D=\frac{\partial \mathbf{R}}{\partial l}-V \nabla \mathbf{R}-(\nabla \mathbf{R})=\frac{\partial \mathbf{R}}{\partial l}-\operatorname{curl} \mathbf{R}-\operatorname{div} \mathbf{R} .
$$

This settles the question. Notice that $D \mathbf{R}$ could not be used relativistic purposes, since $\mathbf{R}$ is right-handed.

Now, to see the utility of $\mathbf{R} D$, return to ( $\mathrm{I} . b$ ), by which $D_{c} \mathbf{R}=$ Take the conjugate of each side, and remember that $\mathbf{R}_{c}=$ Then, by the rule of conjugate of a product,

$$
-\mathrm{R} D=C
$$

and consequently, by (1. $a$ ),

$$
D \mathbf{L}=-\mathbf{R} D
$$

and, substituting this in (28),

$$
2 F=-\mathbf{R} D \cdot \mathbf{L}-\mathbf{R} . D \mathbf{L}=-\mathbf{R}[D] \mathbf{L} .
$$

In this way we obtain the required short expression for the fo quaternion, in terms of the electromagnetic bivectors,

$$
F=-\frac{1}{2} \mathrm{R}[D] \mathrm{L} .
$$

Thus, $\mathbf{R}[] \mathbf{L}$, when applied to $D$, or more correctly, when expo to the bilateral differentiating action of $D$, gives the force-quatern We shall see in the next chapter that the same operator $\mathbf{R}[$ when applied to an ordinary vector, e.g. the normal of a surfe element, will give us the corresponding stress, and, when app to a scalar, the density and the flux of electromagnetic energy

As regards the matrix-equivalents of our bivectors and quaterni equations, it seemed preferable, for the sake of avoiding any poss confusion, not to insert them in the text of this chapter. Some of tl

## NOTES ON CHAPTER VIII.

Note 1 (to page 206). Take first the case $\rho=0$, that is to sa the equations (r.) outside the charges. Measure $x$ along $\mathbf{v}$, of $S^{\prime}$ relative to $S$. Then

$$
\frac{\partial}{c \partial t}=\gamma \frac{\partial}{c \partial t^{\prime}}-\beta \gamma \frac{\partial}{\partial x^{\prime}}, \frac{\partial}{\partial x}=\gamma \frac{\partial}{\partial x^{\prime}}-\beta \gamma \frac{\partial}{c \partial t^{\prime}}, \frac{\partial}{\partial y}=\frac{\partial}{\partial y^{\prime}}, \frac{\partial}{\partial z}=
$$

and the equations

$$
\frac{\mathrm{I}}{\mathrm{c}} \frac{\partial E_{1}}{\partial t}=\frac{\partial M_{3}}{\partial y}-\frac{\partial M_{2}}{\partial z}
$$

and

$$
\operatorname{div} \mathbf{E} \equiv \frac{\partial E_{1}}{\partial x}+\frac{\partial E_{2}}{\partial y}+\frac{\partial E_{3}}{\partial z}=0
$$

will be transformed into

$$
\gamma \frac{\partial E_{1}}{c \partial t^{\prime}}-\beta \gamma \frac{\partial E_{1}}{\partial x^{\prime}}=\frac{\partial M_{3}}{\partial y^{\prime}}-\frac{\partial M_{2}}{\partial z^{\prime}}
$$

and

$$
\gamma \frac{\partial E_{1}}{\partial x^{\prime}}-\beta \gamma \frac{\partial E_{1}}{c \partial t^{\prime}}=-\frac{\partial E_{2}}{\partial y^{\prime}}-\frac{\partial E_{3}}{\partial z^{\prime}} .
$$

Take the sum of the first and $\beta$ times the second of these Then the result will be

$$
\frac{\mathrm{r}}{c} \frac{\partial E_{1}}{\partial t^{\prime}}=\gamma \frac{\partial}{\partial y^{\prime}}\left(M M_{3}-\beta E_{2}\right)-\gamma \frac{\partial}{\partial z^{\prime}}\left(M M_{2}+\beta E_{3}\right) .
$$

Thus the form of the equation (a) reappears. Treat similarly t ing of the equations contained in (I.). Then the whole of thes with $\rho=0$, will reappear in dashed letters, thus:

$$
\frac{\mathrm{I}}{\mathrm{c}} \frac{\partial E_{1}^{\prime}}{\partial t^{\prime}}=\frac{\partial M_{3}^{\prime}}{\partial y^{\prime}}-\frac{\partial M_{2}^{\prime}}{\partial z^{\prime}}, \text { etc. }
$$

where

$$
\begin{array}{lll}
E_{1}^{\prime}=\psi(v) \cdot E_{1}, & E_{2}^{\prime}=\psi(v) \cdot \gamma\left(E_{2}-\beta M_{3}\right), & E_{3}^{\prime}=\psi(\gamma) \cdot \gamma(E \\
M_{1}^{\prime}=\psi(\tau) \cdot M_{1}, & M_{2}^{\prime}=\psi(\tau) \cdot \gamma\left(M I_{2}+\beta E_{3}\right), & M_{3}^{\prime}=\psi(\tau) \cdot \gamma(M
\end{array}
$$

the common factor $\psi(v)$ being thus far an indeterminate fun which for $z=0$ reduces to unity. But solving the last six equ respect to the non-dashed components and claiming mutually 6 for the two systems, $S$ and $S^{\prime}$, we obtain at once

$$
\psi(v) \cdot \psi(-z)=1,
$$

## LAGRANGIAN FUNCTION

and these are the required formulae of transformation, identical wi of this chapter.

Next, pass to the general case of $\operatorname{div} \mathbf{E}=\rho \neq 0$. Bring in the 0 terms $\rho p_{1}$, etc., the components of $\rho \mathbf{p}$, and, by means of the a theorem of velocities, express $\mathbf{p}$ in terms of $\mathbf{p}^{\prime}$ and $\boldsymbol{\nabla}$. Then the $\mathbf{w}$ the general equations collected under (1.) will reappear in dashed thus :
where

$$
\frac{\mathrm{I}}{c} \frac{\partial E_{1}^{\prime}}{\partial t^{\prime}}+\rho^{\prime} p_{1}^{\prime}=\frac{\partial M_{3}^{\prime}}{\partial y^{\prime}}-\frac{\partial M_{2}^{\prime}}{\partial z^{\prime}}, \text { etc. }
$$

$$
\rho^{\prime}=\operatorname{div}^{\prime} \mathbf{E}^{\prime}=\gamma\left(\mathrm{I}-\frac{z \not \mathscr{1}_{1}}{c^{2}}\right) \rho,
$$

or

$$
\rho^{\prime}=\gamma\left[\mathrm{r}-\frac{\mathrm{I}}{c^{2}}(\mathrm{\nabla p})\right] \rho,
$$

and where the components of $\mathbf{E}^{\prime}, \mathbf{M}^{\prime}$ are still connected with tl $\mathbf{E}, \mathbf{M}$ by the above formulae (b). The details of calculation, sim those for $\rho=0$, may be left as an exercise for the reader. By wor out fully he will convince himself best of the advantages of shortne simplicity offered by the quaternionic method employed for the purposes in the text of the chapter.

Note 2 (to page 209). The difference of the magnetic energy $C$ the electric energy $U_{e}$,

$$
L=U_{m}-U_{e}=\frac{1}{2} \int\left(M I^{2}-E^{2}\right) d S,
$$

has been called the Lagrangian function, because it has been rer that the fundamental electronic equations, (r.) and (ir.), can be con into a single variation-formula having the structure of Han Principle (or the principle ' of least action'), $\delta \int_{t_{1}}^{t_{2}} \ldots=0$, in which pr that difference of the two kinds of energy appears (along with possible terms) under the sign of integration. This result is hardl than a purely formal condensation of the original equations since some authors have attributed to it an exaggerated mechan dynamical significance, it may be well to give here a short sketcl bare result and of the method by which it is usually obtained.

Consider a region of space, bounded by the surface $\sigma$, fixed system $S$ in which the equations (I.) and (II.) hold. Let $\rho=0$ at the points of the surface $\sigma$, whose choice is otherwise arbitrary. space region, whose volume-elements will be denoted by $d S$, cont
the charge of each element of matter unchanged. With this ass and since $\rho=\operatorname{div} \mathbf{E}$, the distribution of the infinitesimal vector

$$
\delta^{\prime} \mathbf{C} \equiv \delta \mathbf{E}+\rho \delta \mathbf{r}
$$

will be solenoidal, i.e. such that $\operatorname{div}\left(\delta^{\prime} \mathbf{C}\right)=0$. Let $W$ be the infir virtual work of the ponderomotive forces of electromagnetic oris i.e. by (II.),

$$
W=\int(\mathbf{P} \delta \mathbf{r}) d S=\int \rho\left(\delta \mathbf{r}\left[\mathbf{E}+\frac{\mathbf{I}}{c} \mathrm{~V} \mathbf{p} \mathbf{M}\right]\right) d S .
$$

Then, by the differential electronic equations ( I. ), and after a easy calculation (the details of which, together with the literatur subject, will be found in Lorentz's article in the Encyklop. der matischen Wissenschaften, Vol. V2, pp. 167 et seq. ; Leipsic, 1904

$$
W=\delta\left(U_{m}-U_{e}\right)-\frac{d}{d t}\left(\delta^{\prime} U_{m}\right)-\int(\mathbf{X n}) d \sigma
$$

where $\mathbf{X}$ is the infinitesimal vector

$$
\mathbf{X}=\mathrm{V} \mathbf{A} \delta \mathbf{M}-\mathrm{VA} \frac{\partial \delta^{\prime} \mathbf{M}}{\partial t}+V \mathbf{E} \delta^{\prime} \mathbf{M}
$$

A being the usual vector potential, so that $\mathbf{M}=\operatorname{curl} \mathbf{A}$. The s denotes the variation which would correspond to a change of electric current

$$
\mathbf{C} \equiv \frac{\partial \mathbf{E}}{\partial t}+\rho \mathbf{p}=c . \operatorname{curl} \mathbf{M}
$$

by $\delta^{\prime} \mathbf{C}$, the elements of matter being kept fixed. This am defining $\delta^{\prime} \mathbf{M}$ by $c$. curl $\delta^{\prime} \mathbf{M}=\delta^{\prime} \mathbf{C}$, so that

$$
\begin{aligned}
\delta^{\prime} U_{m} & =\int\left(\mathbf{M} \delta^{\prime} \mathbf{M}\right) d S=\int\left(\delta^{\prime} \mathbf{M} \cdot \operatorname{curl} \mathbf{A}\right) d S \\
& =\int\left(\mathbf{A} \operatorname{curl} \delta^{\prime} \mathbf{M}\right) d S+\int\left(\mathbf{n V A} \delta^{\prime} \mathbf{M}\right) d \sigma \\
& =\frac{1}{c} \int\left(\mathbf{A} \delta^{\prime} \mathbf{C}\right) d S+\int\left(\mathbf{n} V \mathbf{A} \delta^{\prime} \mathbf{M}\right) d \sigma .
\end{aligned}
$$

Such then is the value of the variation appearing in the seco of (c). But this only by the way.

Now, let $\sigma$ expand indefinitely. Then, in virtue of the usual ass as to the behaviour of the field 'at infinity,' the surface integral is vanish, and

## LAGRANGIAN FUNCTION

principle applied to the ordinary, non-electromagnetic masses $\bar{\pi}$ system (if there be any such masses),

$$
W=\Sigma \bar{m}\left(\frac{d^{2} \mathbf{r}}{d t^{2}} \delta \mathbf{r}\right)+\delta V=\frac{d}{d t} \Sigma \bar{m}\left(\frac{d \mathbf{r}}{d t} \delta \mathbf{r}\right)+\delta V-\delta T .
$$

Any relativistic amendment of d'Alembert's principle is here disre of course. Combining the last two equations, integrating from $t=t_{2}$, and assuming that $\delta \mathbf{r}$ and $\delta \mathbf{E}$ vanish at these limiting time we obtain, finally,

$$
\delta \int_{t_{1}}^{t_{2}}\left(U_{m}-U_{e}+T-V\right) d t=0,
$$

that is to say, Hamilton's Principle, in which to the ordinary energy the magnetic energy $U_{m}$ and to the potential energy the energy $U_{e}$ is added. In particular, if the whole energy is electrom as in Abraham's theory, we have simply

$$
\delta \int_{t_{1}}^{t_{2}} L d t=\delta \int_{t_{1}}^{t_{2}}\left(U_{m}-U_{e}\right) d t=0
$$

The more general equation (c) corresponds to the broader vi by Lorentz.

Thus, $L=U_{m}-U_{e}$ plays the rôle of the Lagrangian function versely, assuming $\partial E / \partial t+\rho \mathbf{p}=c$. curl $\mathbf{M}$, with $\rho=\operatorname{div} \mathbf{E}$, and d the remaining fundamental electronic equations, i.e.

$$
\partial \mathbf{M} / \partial t=-c \cdot \operatorname{curl} \mathbf{E} \quad \text { and } \quad \mathbf{P}=\rho\left[\mathbf{E}+\frac{\mathrm{r}}{c} \mathrm{~V} \mathbf{p} \mathbf{M}\right]
$$

can be deduced from (e). For slowly varying motion of the e formula (d) gives at once the ponderomotive forces of electror origin, corresponding to any set of configurational parameter well-known Lagrangian form.

Remember that what is invariant with respect to the Loren formation is the Lagrangian function per unit volume, i.e. $\frac{1}{2}(\Lambda)$ But since $\gamma_{p} d S$ and $d t / \gamma_{p}$, and consequently also $d S$. $d t$ are invar element of 'action'

$$
L_{c} d t=\left(U_{m}-U_{e}\right) d t
$$

is an intuariant. And so also is the whole 'action' $\int_{t_{1}}^{t_{2}} L d t$ invar respect to the Lorentz transformation. It may be noticed here is only a particular instance of a general theorem of relativistic $d$ obtained by Planck.
Note 3 (to page 2r1). Differentiating $E_{1}^{\prime}=E_{1}, E_{2}^{\prime}=\gamma\left(E_{2}-\right.$

## THE THEORY OF RELATIVITY

Now, by the addition theorem of velocities (see Chap. VI., and formula (b), p. 169),

$$
\gamma_{p}=\gamma_{v} \gamma_{p^{\prime}}\left[\mathrm{I}+\frac{\mathrm{I}}{c^{2}}\left(\boldsymbol{\nabla \mathbf { p } ^ { \prime }}\right)\right]
$$

whence, by inversion,

$$
\gamma_{p^{\prime}}=\gamma_{v} \gamma_{p}\left[\mathrm{I}-\frac{\mathrm{I}}{c^{2}}(\mathrm{vp})\right] .
$$

Thus $\rho^{\prime} / \gamma_{p^{\prime}}=\rho / \gamma_{p}$, and since $\gamma_{p^{\prime}} d S^{\prime}=\gamma_{p} d S$,

$$
\rho^{\prime} d S^{\prime}=\rho d S
$$

which is the required verification of the invariance of electrical $c$
Note 4 (to page 215). Using the formula obtained for $\mathbf{g}$ on F have, for the electromagnetic momentum of the whole field,

$$
\mathbf{G}=\int \mathbf{g} d S=\frac{v^{2}}{c^{2}} \int\left[E^{2} \mathbf{u}-E_{1} \mathbf{E}\right] d S,
$$

where $\mathbf{u}$ is the unit of $\boldsymbol{v}$ and $E_{1}$ the longitudiual component of $\mathbf{E}$. the transversal part of $\mathbf{E}$, the bracketed terms may be written

$$
\left(E^{2}-E_{1}^{2}\right) \mathbf{u}-E_{1} \mathbf{E}_{t},
$$

and since the field is, in the case under consideration, syr round $u$, the transversal terms cancel one another in the proc tegration, so that

$$
\mathbf{G}=\frac{\boldsymbol{\nabla}}{c^{2}} \int\left(E^{2}-E_{1}^{2}\right) d S=G \mathbf{u} .
$$

For a Lorentz electron of homogeneous surface-charge,

$$
\mathbf{E}=\frac{e \gamma}{4 \pi r^{\prime 3}} \mathbf{r}, \quad r^{\prime} \geqq R,
$$

and $E=0$ inside the electron. Writing, therefore, $r^{2}-x^{2}=s^{2}$, w $\epsilon$

$$
G=\frac{v}{c^{2}}\left(\frac{e \gamma}{4 \pi}\right)^{2} \int \frac{s^{2}}{\gamma^{\prime 6}} d S,
$$

where the integral is to be taken throughout the $S$-space lyin the ellipsoid $r^{\prime}=\left(\gamma^{2} x^{2}+s^{2}\right)^{\frac{3}{2}}=R$. But since this ellipsoid is, $S^{\prime}$-standpoint, a sphere of radius $R$, it is easier, of course, to the integration in the $S^{\prime}$-space. Thus, remembering that $d S=d S^{\prime} / \gamma$ (or that the functional determinant of $x, y, z$ with r

## PHYSICAL BIVECTORS

so that

$$
G=\frac{e^{2} \gamma}{6 \pi c^{2} R}{ }^{2 \prime}
$$

and

$$
\mathbf{G}=G \mathbf{u}=\frac{e^{2} \gamma}{6 \pi c^{2} h^{2}} \mathbf{\nabla}
$$

which is the required formula.
Note 5 (to page 218 ). Let $\mathbf{A}, \mathbf{B}$ be a pair of real vectors and $\mathbf{A}$ left-handed physical bivector, i.e. such that

$$
\mathbf{A}^{\prime}-\iota \mathbf{B}^{\prime}=Q_{c}[\mathbf{A}-\iota \mathbf{B}] Q=Q_{c} \mathbf{A} Q-\iota Q_{c} \mathbf{B} Q
$$

This splits into
and

$$
\left.\begin{array}{l}
\mathbf{A}^{\prime}=\text { re. } Q_{c} \mathbf{A} Q-\iota . \text { imag. } Q_{c} \mathbf{B} Q \\
\iota \mathbf{B}^{\prime}=\iota . \text { re. } Q_{c} \mathbf{B} Q-\text { imag. } Q_{c} \mathbf{A} Q
\end{array}\right\}
$$

where re. and imag. stand for 'real pairt of' and 'imaginary pa Now, since $Q$ has a real vector and an imaginary scalar, and since the conjugate of $Q$, it is obvious that

$$
\begin{gathered}
\text { re. } Q_{c} \mathbf{A} Q=\text { re. } Q \mathbf{A} Q_{c}, \\
\text { imag. } Q_{c} \mathbf{A} Q=-\operatorname{imag} \cdot Q \mathbf{A} Q_{c},
\end{gathered}
$$

and similarly for $\mathbf{B}$. Therefore, by (a),

$$
\begin{aligned}
\mathbf{A}^{\prime}+\iota \mathbf{B}^{\prime} & =\text { re. } Q \mathbf{A} Q_{c}+\iota \cdot \text { re. } Q \mathbf{B} Q_{c}+\text { imag. } Q \mathbf{A} Q_{c}+\iota \cdot \text { imag. } Q \mathbf{B} Q_{c} \\
& =Q \mathbf{A} Q_{c}+\iota \mathbf{B} Q_{c}=Q[\mathbf{A}+\iota \mathbf{B}] Q_{c},
\end{aligned}
$$

that is to say, $\mathbf{A}+\iota \mathbf{B}$ is a right-handed bivector. Q.E.I.
Note 6 (to page 223). Our physical bivector is equivalent to Minko space-time vector of the second kind and to Sommerfeld's six-2 Minkowski represents this world-vector by an 'alternating' matrix

$$
h=\left|\begin{array}{llll}
0, & h_{12}, & h_{13}, & h_{14} \\
h_{21}, & 0, & h_{23}, & h_{24} \\
h_{31}, & h_{32}, & 0, & h_{34} \\
h_{41}, & h_{42}, & h_{43}, & 0
\end{array}\right| \quad\left(h_{\kappa c}=-h_{L \kappa}\right)
$$

subjected to the condition that

$$
h^{\prime}=\bar{A} h A
$$

The matrix $h$ is built up of six independent constituents (not counting the diagonal which is always the same). Out of these six constituents three, not containing the index 4 , are real, and the remaining three imaginary :

$$
\begin{aligned}
& h_{23}, h_{31}, h_{12} \text { real, } \\
& h_{14}, h_{24}, h_{34} \text { imaginary. }
\end{aligned}
$$

Along with $h$, Minkowski uses the corresponding 'dual' matrix which he denotes by $h^{*}$, and which is again an alternating matrix, e.g.

$$
h^{*}=\left|\begin{array}{llll}
0, & h_{34}, & h_{42}, & h_{23} \\
h_{43}, & 0, & h_{14}, & h_{31} \\
h_{24}, & h_{41}, & 0, & h_{12} \\
h_{32}, & h_{13}, & h_{21}, & 0
\end{array}\right| .
$$

This is transformed like $h$. The product of both matrices,

$$
\begin{equation*}
h^{*} h=h_{32} h_{14}+h_{13} h_{24}+h_{21} h_{34}, \tag{a}
\end{equation*}
$$

which is also the square root of det $h$, and

$$
\begin{equation*}
h_{23}{ }^{2}+h_{31}{ }^{2}+h_{12}{ }^{2}+h_{14}{ }^{2}+h_{24}{ }^{2}+h_{34}{ }^{2} \tag{b}
\end{equation*}
$$

are invariant with respect to the Lorentz transformation. Both of these invariants are contained in the square of our physical bivector.

Let, in particular,

$$
\left.\begin{array}{ll}
h_{23}=M_{1}, & h_{31}=M_{2},  \tag{c}\\
h_{14}=-\iota E_{1}, & h_{24}=M_{3} \\
=-\iota E_{2}, & h_{34}=-\iota E_{3} .
\end{array}\right\}
$$

Then the matrix $h$ will correspond to the electromagnetic bivector $\mathbf{L}=\mathbf{M}-\iota \mathbf{E}$. (In Sommerfeld's four-dimensional language we should say that the magnetic components are projections of the six-vector $h$ upon the planes $y z, z x, x y$, and $-\iota$ times the electric components the projections of $h$ upon the planes $x l, y l, z l$.) With this particular meaning of $h$ the matrix form of the electronic differential equations (I.) consists of the equations

$$
\left.\begin{array}{l}
\text { lor } h=-s  \tag{d}\\
\operatorname{lor} h^{*}=0,
\end{array}\right\}
$$

the former embodying the first pair and the latter the second pair of the equations (r.). Here $s$ is the current-matrix,

$$
s=\rho\left|\frac{p_{1}}{c}, \frac{p_{2}}{c}, \frac{p_{3}}{c}, \iota\right|
$$

## MINKOWSKI'S FORMULAE

Both of these are contained in $\mathbf{L}^{2}$. The ponderomotive force $\mathbf{P}$, (iI.), its activity are given by the matrix -sh. In fact, taking the produ $s$ into $h$, by the rule given in the Note to Chap. V., we obtain

$$
\frac{1}{\rho} s h=\left\lvert\,-\frac{p_{2}}{c} M_{3}+\frac{p_{3}}{c} M_{2}-E_{1}\right., \text { etc., } \left.\quad-\iota \frac{p_{1}}{c} E_{1}-\iota \frac{p_{2}}{c} E_{2}-\iota \frac{p_{3}}{c} E_{3} \right\rvert\,,
$$

i.e.

$$
-s / l=\left|P_{1}, P_{2}, P_{3}, \frac{\iota}{c}(\mathbf{P p})\right|
$$

Since $s^{\prime}=s A$, as in (37), p. 143, and $k^{\prime}=\bar{A} h A$, we have $s^{\prime} h^{\prime}=\operatorname{sh} A$, sho that the four-dimensional force, per unit volume, is indeed a world-v of the first kind. Its quaternionic equivalent is $F=\frac{l}{c}(\mathbf{P p})+\mathbf{P}$, the f quaternion of this chapter. The expression $\mathbf{R} C-C \mathbf{L}$ in formula takes the place of the matrix $2 s h$.

## CHAPTER IX.

## ELECTROMAGNETIC STRESS, ENERGY AND MOMENTUM. EXTENSION TO GENERAL DYNAMICS.

Is the preceding chapter we have seen that the fundamental electronic equations are invariant with respect to the Lorentz transformation, and we have obtained for the force-quaternion per unit volume, i.e. for

$$
\begin{equation*}
F={ }_{c}^{l}(\mathbf{P p})+\mathbf{P} \tag{r}
\end{equation*}
$$

the short formula (II.b), p. 223,

$$
\begin{equation*}
F=-\frac{1}{2} \mathbf{R}[D] \mathbf{L} \tag{2}
\end{equation*}
$$

Here $D$ is intended to operate on both $R$ and $L$, and the only office of the brackets is to remind us of this bilateral differentiation.

We shall now deduce from this formula the electromagnetic stress $f_{n}$ together with the density and the flux of energy. All these magnitudes have already been treated in Chap. II. But now, in virtue of (2), they will appear in a form which will disclose at once their transformational properties.

Take first the scalar part of (2). This gives, by (1), and since SRL $=-(R L)$,

$$
\frac{t}{c}(\mathbf{P p})=\frac{\mathrm{r}}{2} \frac{\partial}{\partial l}(\mathbf{R L})-\frac{1}{2} \operatorname{div} \mathbf{V R L}
$$

or

$$
\begin{equation*}
(\mathbf{P p})=-\frac{\partial u}{\partial t}-\operatorname{div} \tag{3}
\end{equation*}
$$

where

$$
\left.\begin{array}{rl}
u & =\frac{1}{2}(\mathrm{RL})  \tag{4}\\
\\
\forall H & ={ }_{2}^{c} \mathrm{VLR} .
\end{array}\right\}
$$

## STRESS, ENERGY, ETC.

Remembering the meaning of $\mathbf{L}$ and $\mathbf{R}$, the reader will see at that these are identical with the familiar formulae $u=\frac{1}{2}\left(E^{2}+\right.$ $z_{p}=c$ VEM. But the above form will better answer our purp

Thus, the scalar part of the equation (2) expresses the conserv of energy, giving the flux of energy or the Poynting vector 据, $^{\prime}$ with $u$, the density of electromagnetic energy. Both of these also be condensed into the full product,

$$
\frac{1}{2} \mathrm{RL}=-u+\frac{\imath}{c}
$$

It is hardly necessary to say that this is not a physical quatern But the formula recommends itself by its shortness.

Next, consider the vector part of (2). This is, by ( I ), the deromotive force,

$$
\mathbf{P}=-\frac{1}{2} \partial l V R L-\frac{1}{2} V R[\nabla] L
$$

or, by the second of (4),

$$
\mathbf{P}=-\frac{\mathrm{r}}{\mathrm{c}^{2}} \frac{\partial \| \mathrm{J}}{\partial t}-\frac{\mathrm{r}}{2} \operatorname{VR}[\nabla] \mathrm{L} .
$$

Writing $\nabla=\mathrm{i} \partial / \partial \mathrm{x}+\mathrm{j} \partial / \partial y+\mathrm{k} \partial / \partial z$, and remembering that bo and $L$ are to be differentiated, we have

$$
V \boldsymbol{R}[\nabla] \mathbf{L}=\frac{\partial}{\partial x} V \operatorname{RiL}+\frac{\partial}{\partial j} V \operatorname{RjL}+\frac{\partial}{\partial z} V R k L .
$$

On the other hand, if $f$ is a stress-operator, i.e. if

$$
f \mathbf{n}=\mathrm{f}_{n}
$$

is the pressure, per unit area, on a surface element whose normal is n , and if we write in particular, as on p. 48,

$$
f \mathbf{i}=\mathbf{f}_{1}, \quad f \mathbf{j}=\mathbf{f}_{2}, \quad f \mathbf{k}=\mathbf{f}_{3},
$$

then the corresponding resultant force per unit volume will

$$
\mathbf{i}\left(\frac{\partial f_{11}}{\partial x}+\frac{\partial f_{21}}{\partial y}+\frac{\partial f_{31}}{\partial z}\right)+\mathbf{j}\left(\frac{\partial f_{12}}{\partial x}+\frac{\partial f_{22}}{\partial y}+\frac{\partial f_{32}}{\partial z}\right)+\mathbf{k}\left(\frac{\partial f_{13}}{\partial x}+\frac{\partial f_{23}}{\partial y}+\frac{\partial f^{2}}{\partial}\right.
$$

which is exactly of the form of $(a)$. We see, therefore, that

$$
\mathbf{P}=-\frac{\partial \mathbf{g}}{\partial t}-\frac{\partial f \mathrm{i}}{\partial x}-\frac{\partial f \mathbf{j}}{\partial y}-\frac{\partial f \mathbf{k}}{\partial z},
$$

where $\mathbf{g}=\mathrm{d} / \mathrm{c}^{2}$, as on $\mathrm{p} .5^{\mathrm{I}}$, and where, for $\mathrm{i}, \mathbf{j}, \mathbf{k}$, and hence also for any unit vector $\mathbf{n}$,

$$
f \mathrm{n}=\mathrm{f}_{n}=\frac{1}{2} \mathrm{VRn} \mathbf{L} .
$$

This is the required formula for the stress. Multiplying out the right-hand side, the reader will easily obtain

$$
\begin{aligned}
\mathbf{f}_{n} & =\frac{1}{2}(\mathbf{R L}) \mathbf{n}-\frac{1}{2} \mathbf{R}(\mathbf{L n})-\frac{1}{2} \mathbf{L}(\mathbf{R n}) \\
& =u \mathbf{n}-\mathbf{E}(\mathbf{E n})-\mathbf{M}(\mathbf{M n}),
\end{aligned}
$$

which is the Maxwellian stress, (20), p. 48. But the above form obtained directly from (2), is more appropriate for our purposes Again, since the stress is irrotational, or since $f$ is a symmetrica operator, we have $f \mathrm{i}=\mathrm{i} f$, etc., so that we may write, in the las formula for $\mathbf{P}$,

$$
\frac{\partial f \mathbf{i}}{\partial x}+\frac{\partial f \mathbf{j}}{\partial y}+\frac{\partial f \mathbf{k}}{\partial z}=\nabla f
$$

where $f$ is to be considered as a dyadic. (See Note 1.) Had w used this short form at the beginning, we might have obtained th above formula for $f$ even more directly.

Thus, the vector part of the equation (2) gives for the pondero motive force the expression

$$
\mathbf{P}=-\frac{\partial \mathbf{g}}{\partial t}-\nabla f
$$

where g , the electromagnetic momentum per unit volume, and $f \mathrm{n}$, th stress for any orientation of n , are determined by

$$
\mathrm{g}=\frac{\text { 挸 }}{c^{2}}=\frac{l}{2 c} \mathrm{VLR}
$$

and

$$
f n=\frac{1}{2} V R n L .
$$

## STRESS，ENERGY，ETC．

Now，the scalar part of this ternary product is

$$
\operatorname{SRnL}=S R V n L=-(R V n L)=(n V R L)
$$

so that，by（4），

$$
\frac{1}{2} \mathrm{SRnI}=\frac{l}{c}(\text { 羽 }) .
$$

Consequently，the full product will be

$$
\frac{1}{2} \mathrm{RnL}=\frac{l}{c}(\nmid(\mathrm{n})+f \mathrm{n} .
$$

It will be convenient to combine this with（4a）into one forn Let $\sigma$ be a real，but otherwise arbitrary scalar，and let us i duce for the moment the auxiliary quaternion

$$
K=\iota \sigma+\mathrm{n} .
$$

Adding $\sigma \sigma$ times $(4 a)$ to $(7 a)$ ，we have

$$
\frac{1}{2} \mathrm{R} k \mathrm{~L}={ }_{c}^{l}[(\mathrm{f} \mathrm{nn})-c \sigma u]+f \mathrm{n}-\frac{\sigma}{c} \text { 羽. }
$$

This is valid for any $k$ ，that is，for any direction of $\mathbf{n}$ and fo value of $\sigma$ ．

Since（2）transforms into itself，i．e．into $F^{\prime}=-\frac{1}{2} \mathrm{R}^{\prime}\left[D^{\prime}\right] \mathrm{L}^{\prime}$ for legitimate system $S^{\prime \prime}$ ，the same thing is true of the equatio energy（3）and of the formula for the ponderomotive force Both are invariant with respect to the Lorentz transformation． we have，in $S^{\prime \prime}$ ，

$$
\left.\mathbf{P}^{\prime} \mathbf{p}^{\prime}=-\frac{\partial u^{\prime}}{\partial t^{\prime}}-\operatorname{div}^{\prime} \sharp\right]^{\prime}
$$

and

$$
\mathbf{P}^{\prime}=-\frac{\partial \mathbf{g}^{\prime}}{\partial t^{\prime}}-\nabla^{\prime} f^{\prime}
$$

where $\mathrm{g}^{\prime}=\underline{j^{\prime}} / c^{2}$ and where $\exists ⿰ 习 习^{\prime}, u^{\prime}, f^{\prime}$ are determined by the pre formulae，i．e．also by（8）with dashed letters．Remember that the stress－operator in $S^{\prime}$ ，so that if $\mathbf{n}^{\prime}$ is a unit vector，$f^{\prime} \mathbf{n}^{\prime}=$ the pressure on a unit area whose normal is $\mathbf{n}^{\prime}$ ．

What are the connexions between $\left\{刀^{\prime}, u^{\prime}, f^{\prime}\right.$ on the one sid鲥，$u, f$ on the other side？To answer this question，return $t$
$\mathbf{n}^{\prime}$ being the unit of $\mathbf{N}^{\prime}$ ．Then $\mathbf{R} k \mathbf{L}$ will also be a physical qua ternion，$\simeq q$ ．Denoting，therefore，by（8）the right side of th equation thus numbered，and by（ $8^{\prime}$ ）the same expression wit dashes，we have

$$
\frac{1}{2} \mathrm{R}^{\prime} h^{\prime} \mathrm{L}^{\prime}=\left(8^{\prime}\right)=Q(8) Q .
$$

Writing down $Q(8) Q$ and equating its scalar and vector parts t the scalar and vector parts of $\left(8^{\prime}\right)$ ，we obtain the two relations
in which $\mathbf{V}$ is the velocity of $S^{\prime}$ relative to $S$ and $\epsilon$ our previou longitudinal stretcher of ratio $\gamma=\left(\mathrm{I}-\nu^{2} / c^{2}\right)^{-\frac{1}{2}}$ ．Now，since these relations hold for any value of $\sigma$ ，take first $\sigma=0$ ，and then $\sigma=1$ and remember that，by（ 9 ），

$$
\begin{gathered}
\sigma_{0}^{\prime}=-\frac{\gamma}{c}(\nabla \mathbf{n}), \quad \mathbf{N}_{0}^{\prime}=\epsilon \mathfrak{n}, \\
\sigma_{1}^{\prime}-\sigma_{0}^{\prime}=\gamma, \quad \mathbf{N}_{1}^{\prime}-\mathbf{N}_{0}^{\prime}=-\frac{\gamma}{c} \nabla .
\end{gathered}
$$

Then of the four relations，obtained in this way，one，containing the $n$－component of $f(f v$ ，will turn out to be a consequence o the three others．

These three relations，after a simple rearrangement of terms，anc without Cartesian splitting，give us the required relativistic trans formation of the density and the flux of electromagnetic energy and of the stress in the short form

$$
\begin{align*}
& \frac{\mathrm{I}}{\gamma^{2}} u=u^{\prime}+\frac{2}{c^{2}}\left(\boldsymbol{y}^{\prime} \boldsymbol{v}\right)+\frac{\mathrm{I}}{c^{2}}\left(\mathbf{v} f^{\prime} \mathbf{v}\right) \tag{10}
\end{align*}
$$

$$
\begin{aligned}
& \left.\underline{\mathrm{I}} f=f^{\prime}+\underline{\gamma}_{\left[H 月^{\prime}\right.}+\ell^{\prime} \mathbf{v}\right] / \mathbf{v}+\underline{\mathbf{v}}\left(\epsilon_{\epsilon} 月^{\prime}\right.
\end{aligned}
$$

## TRANSFORMATION OF STRESS, ETC.

vector n as operand, and closing the parentheses, we obtain corresponding pressure $\mathrm{f}_{n}=f \mathrm{n}$, thus:

Remember that $\epsilon$ is a symmetrical operator, so that $\left(\epsilon \exists^{\prime} \cdot \mathbf{n}\right)=\left(\right.$ 朋 $^{\prime}$.
To obtain the stress in its more familiar form, take the $u$ system of normal unit vectors, $\mathbf{i}$ along and $\mathbf{j}, \mathbf{k}$ at right angles to direction of motion. Write in turn $\mathbf{n}=\mathbf{i}, \mathbf{j}, \mathbf{k}$, and remember $\epsilon \mathrm{i}=\gamma \mathbf{i}, \epsilon \mathrm{j}=\mathbf{j}, \epsilon \mathbf{k}=\mathbf{k}$. Then

$$
\begin{aligned}
& { }_{\boldsymbol{\epsilon}}^{\mathrm{I}} \mathrm{f}_{3}=\mathrm{f}_{3}{ }^{\prime}+\frac{\mathrm{V}}{\mathrm{c}^{2}} \mathrm{pl}_{3}{ }^{\prime} .
\end{aligned}
$$

Splitting each of the stress vectors $f_{1}$, etc., into its three rectang components along the same set of axes, we obtain nine st formulae which contract to six, since $f_{12}{ }^{\prime}=f_{21}{ }^{\prime}$, etc., and $f_{12}=f_{21}$, Treating similarly the first two of the equations ( x 0 ), we have the transformation of stress and of flux and density of energy the Cartesian formulae, which were first given by Laue,

The transformation formula of $\mathbf{g}$, the electromagnetic momentum unit volume, which is simply the energy flux divided by $c^{2}$, will be
and momentum，as estimated from the $S$－standpoint，are up of the stress，energy and momentum or energy flux cor to the $S^{\prime}$－point of view．This entanglement of the vari tudes，which in classical physics led an independent ex characteristic of the theory of relativity．It is a conse the way in which time and space are involved in the fu Lorentz transformation．

In deducing the formulae（ ro ）of transformation of associated magnitudes，we have used their expressions i the electromagnetic bivectors，as condensed in（8）．Our doing so was to show the properties of the simple opera But，as a matter of fact，these formulae hold quite indepe the particular，electromagnetic meaning of $f, u$ and $\mathbf{g}$ or $⿰ 习 习 ⿱ 一 𫝀 口 㐄 ~$ are valid in virtue of（3）and（5）alone（with 眈 $=c^{2} \mathbf{g}$ ），tha for stresses etc．of any origin，electromagnetic or not，pr the corresponding ponderomotive force，per unit volume，and can be represented in the form

$$
\begin{gathered}
\mathbf{P}=-\nabla f-\frac{\partial \mathbf{g}}{\partial t} \\
(\mathbf{P p})=-\frac{\partial u}{\partial t}-c^{2} \cdot \operatorname{div} \mathbf{g} .
\end{gathered}
$$

The proof of this statement is most simply obtained by method，which in this case is superior to the quaternioni course，each method has advantages for certain purposes consider the symmetrical matrix

$$
S=\left|\begin{array}{cccc}
f_{11}, & f_{12}, & f_{13}, & \iota c_{g} g_{1} \\
f_{21}, & f_{22}, & f_{23}, & \iota c_{g 2}^{g_{2}} \\
f_{31}, & f_{32}, & f_{33}, & \iota c_{3} \\
\iota c_{g 1}, & \iota c_{22}, & \iota c_{3}, & -\psi
\end{array}\right|
$$

in which $f_{\iota \kappa}=f_{\kappa \iota}$＊Multiply it by，or operate upon it matrix lor $=\left|\frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z}, \frac{\partial}{\partial l}\right|$ ，according to the rule gi Note to Chapter V．Then the result will be

## STRESS-ENERGY MATRIX

the constituents of the matrix on the right side being exactly given by (A) and (B). This matrix is the equivalent of the phy quaternion $F=\mathbf{P}+\frac{\iota}{c}(\mathbf{P p})$. We can therefore use for it the letter $F$. Thus, the last equation can be written

$$
F=-\operatorname{lor} \mathfrak{S}
$$

To write this for the force-matrix is exactly the same thin to postulate ( A ) and ( B ) for the force and its activity.

Let me observe here that the matrix (ri)* can be written siderably shorter, thus:

$$
\mathfrak{S}=\left|\begin{array}{ll}
f, & \iota \mathrm{~g} \\
\iota c g, & -u,
\end{array}\right| .
$$

Here one constituent is a linear operator, or, say, a dy $\left.\left.\mathbf{f}=\mathbf{i}) \mathbf{f}_{1}+\mathbf{j}\right) \mathbf{f}_{2}+\mathbf{k}\right) \mathrm{f}_{3}$, two other constituents are vectors, and the fc a scalar. But this heterogeneity of the various constituents of one the same matrix need not alarm us. It seems even to harm fully with the original intention of the creation of Cayley, wished to see his instrument of multiple algebra treated as bro as possible. The only requirement is that the array should rectangular. Using the abbreviated form ( $\mathrm{r} a$ ), we have, of co to use lor, correspondingly, as the matrix of $\mathrm{r} \times 2$ constitu $\nabla$ to be applied scalarly, and $\partial / \partial l$. In this way we obtain

$$
-\operatorname{lor} \mathscr{S}=-\left|\nabla f+\frac{\partial g}{\partial t}, \quad{ }_{c}\left(\operatorname{div} c^{2} \mathrm{~g}+\frac{\partial u}{\partial t}\right)\right|=F
$$

at once, instead of writing first so many scalar terms and gathering them together.

But let us return to our subject. We know already that, wha the nature of the ponderomotive force, $F$ is a physical quaternio the matrix $l^{\prime}$ is transformed as $|\mathbf{r}, l|$. And the same thing is of lor. Thus, if $A$ be the fundamental transforming matrix, p. 143, we have

$$
F^{\prime}=F A, \operatorname{lor}^{\prime}=\operatorname{lor} A
$$

and therefore, by ( r 2 ),

$$
\operatorname{lor} A \mathbb{S}^{\prime}=\operatorname{lor} \mathfrak{S} A
$$

Now, substituting here for $A$ the matrix (40), p. i44, re that the transposed matrix $\bar{A}$ is obtained from $A$ by a m of the sign of $\beta$, and multiplying out the right side, the easily convince himself of the identity of ( 13 ) with the trar formulae ( $\mathrm{r} \circ a$ ), in which $1 \boldsymbol{1}=c^{2} \mathrm{~g}$. This proves the above F which may be restated as follows:

If we make with regard to any ponderomotive assumptions (A) and (B), or, which is the same assumption

$$
F=-\operatorname{lor} \mathfrak{S}
$$

then the corresponding pressures, etc., are transformed to ( $\mathrm{r} \circ a$ ) or ( r 0 ), with 7 羽 $=c^{2} \mathrm{~g}$.

It is, of course, an entirely different question whe assumptions are to be considered as universally vali Assumption (в) is the expression of the principle of cons enersy together with the concepts of its localization anc (A) leads to the principle of conservation of momentum, is a strong tendency among the relativists to retain bot principles of classical physics. Thus, M. Abraham involving both principles, in his paper on the electrod ponderable bodies,* and appeals to this equation ev theory of gravitation, which does not satisfy the p relativity, while Laue makes of it the basis of the genera of continuous bodies. On the other hand, according to 1 electrodynamics of moving ponderable bodies, the pon force and its activity are expressed by that part of the $v$ $F=-$ lor $\mathfrak{S} \dagger$, which is normal to the four-velocity $Y$, matrix

$$
F+\frac{\mathrm{I}}{c^{2}} Y \bar{F} Y
$$

or, which is the same thing, by the physical quaternion

$$
\frac{1}{2}\left[F+\frac{\mathrm{I}}{\hat{c}^{2}} Y F_{c} Y\right]
$$

## TRANSFORMATION OF STRESS，ETC．

and not by $F=-$ lor $\mathfrak{S}$ itself．Now，it is true that Abraham Laue＇s device recommends itself by its simplicity in the ca the general mechanics of ponderable continua；but，on the hand，Minkowski＇s device seems to offer advantages for a relat theory of gravitation．In fact，a pair of such theories，both sati rigorously the principle of relativity，have been recently pro by Nordström，＊in one of which the four－dimensional force the form lor $\mathbb{S}$ ，while in the other it is given by the part of a world－vector perpendicular to $Y$ ．Now，the latter of these th is physically simpler，inasmuch as it leads to a rest－mass pendent of the gravitation potential，while the former requir rest－mass to become an exponential function of this potentia

Certainly，then，the principle of relativity does not compel attribute to the forms（A）and（B）of ponderomotive force a activity an universal validity．But it is at any rate interest see the consequences of making the assumptions（A），（в）a accepting，therefore，the formulae（ro）also for pressures，mom and energy of non－electromagnetic nature in any material me Once the reader knows expressly the conditions of their va there is no danger in doing so．

We shall therefore proceed to give here some consequences formulae（ro）．

Let the system of reference $S^{\prime}$ be such that there is no $f$ energy with respect to it，i．e．such that $\exists ⿰ 习 习^{\prime}=0$ ，and therefor $g^{\prime}=0$ ．This will，under a restriction，be the case when $S^{\prime}$ rest－system either of the whole material body，if all its parts ha same velocity relative to $S$ ，or，more generally，of its volume－el under consideration．We may retain in both cases the sym which will then generally denote the velocity of an element body with respect to $S$ ．The restriction hinted at consists obv in supposing that in $u^{\prime}$ are contained only such kinds of energy not flow through the element in question，e．g．energy of deformation，energy stored up in the atoms，heat for the of uniform temperature，and－as Laue adds－＇possibly also new kinds of energy，yet undiscovered．＇But to these the e

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## THE THEORY OF RELATIVITY

or a magnetostatic field, we can include in $u^{\prime}$ the density corresponding energy, combining at the same time the or the magnetic stress with the mechanical one.

Keeping this in mind, and writing $c^{2}$ g for 据, we obtain, density of energy and of momentum and for the stress, as es from the $S$-point of view, the formulae ( r 0 ), considerably simp

$$
\left.\begin{array}{l}
u=\gamma^{2}\left[u^{\prime}+\frac{\mathbf{r}}{c^{2}}\left(\mathbf{v} f^{\prime} \mathbf{v}\right)\right] \\
\mathbf{g}=\frac{\gamma^{2}}{c^{2}}\left[u^{\prime} \mathbf{v}+\gamma^{-1} \epsilon f^{\prime} \mathbf{v}\right] \\
f=\epsilon f^{\prime} \epsilon+\frac{\gamma^{2} u^{\prime}}{c^{2}} \mathbf{v}(\mathbf{v} .
\end{array}\right\}
$$

In Cartesians, with axes taken along the velocity $\mathbf{v}$ of a parti at right angles to it, these formulae are, as (roa) without the fluxes of energy,

$$
\begin{aligned}
& f_{11}=\gamma^{2}\left(f_{11}{ }^{\prime}+\beta^{2} u^{\prime}\right) ; \quad f_{22}=f_{22}{ }^{\prime} ; \quad f_{33}=f_{33}{ }^{\prime} \\
& f_{23}=f_{23} ; \quad f_{31}=\gamma f_{31^{\prime}} ; \quad f_{12}=\gamma f_{12}{ }^{\prime} \\
& g_{1}=\frac{\gamma^{2} v}{c^{2}}\left(u^{\prime}+f_{11}{ }^{\prime}\right) ; g_{2}=\frac{\gamma v}{c^{2}} f_{12}{ }^{\prime} ; \quad g_{3}=\frac{\gamma v}{c^{2}} f_{13}{ }^{\prime} \\
& u=\gamma^{2}\left(u^{\prime}+\beta^{2} f_{11}{ }^{\prime}\right) .
\end{aligned}
$$

We may notice in passing that the sum of the diagonal cons of the matrix $\mathfrak{S}$, i.e.

$$
f_{11}+f_{22}+f_{33}-u,
$$

is always an invariant. With the above choice of axes, $v$ also separately, by ( $\mathrm{r} 4 a$ ) or by ( $\mathrm{r} \circ \mathrm{a}$ ),

$$
f_{11}-u=f_{11}^{\prime}-u^{\prime} \text { and } f_{22}=f_{22}^{\prime}, f_{33}=f_{33}^{\prime} .
$$

The invariant ( $\mathrm{r}_{5}$ ) vanishes in the case of purely electron Maxwellian stress. But for mechanical stresses its value general differ from zero.

## RELATIVE STRESS

the name of relative stress. According to Laue, we have to $p$ every dynamically complete system, $F=0$, and to write, theref the head of dynamics of continuous bodies, the equation*

$$
\operatorname{lor} \mathbb{S}=0
$$

This amounts to putting $\mathbf{P}=0$ in (A) and (B), so that

$$
\frac{\partial u}{\partial t}+c^{2} \operatorname{div} \mathbf{g}=0, \quad \frac{\partial \mathbf{g}}{\partial t}=-\nabla f .
$$

At the same time it is assumed that the resultant force acting any individual portion of the body in question is given by

$$
\mathbf{N}=\frac{d \mathbf{G}}{d t},
$$

where $\mathbf{G}=\int \mathbf{g} d S$, the integral being taken throughout the vol that portion. If, therefore, $d S$ is an individual volume-e of the body, the relative stress which we shall denote by symbol of an operator, $\dagger$ will, according to the familiar defi be given by

$$
\frac{d}{d t}(\mathbf{g} d S)=-\mathbf{i}\left(\frac{\partial p_{11}}{\partial x}+\frac{\partial p_{21}}{\partial y}+\frac{\partial p_{31}}{\partial z}\right) d S-\mathbf{j} \ldots
$$

or

$$
\frac{d}{d t}(\mathbf{g} d S)=-\left[\frac{\partial \mathbf{p}_{1}}{\partial x}+\frac{\partial \mathbf{p}_{2}}{\partial y}+\frac{\partial \mathbf{p}_{3}}{\partial z}\right] d S
$$

where $\mathrm{p}_{1}=p \mathrm{i}$, etc., and where $\frac{d}{d t}$ is the individual rate of c On the other hand, the meaning of the absolute stress $f$ is gi the second of ( $\mathrm{r} 6 a$ ) or, in expanded form, by

$$
\frac{\partial \mathbf{g}}{\partial t}=-\frac{\partial \mathbf{f}_{1}}{\partial x}-\frac{\partial \mathbf{f}_{2}}{\partial y}-\frac{\partial \mathbf{f}_{3}}{\partial z},
$$

where $\frac{\partial}{\partial t}$ is the local rate of variation, corresponding to cc

## THE THEORY OF RELATIVITY

values of $x, y, z$. Now, we have, for any orientation of the syste of rectangular axes,

$$
\begin{aligned}
\frac{d}{d t}(\mathbf{g} d S) & =\left[\mathbf{g} \operatorname{div} \mathbf{v}+\frac{\partial \mathbf{g}}{\partial t}+(\mathbf{v} \nabla) \mathbf{g}\right] d S \\
& =\left[\frac{\partial \mathbf{g}}{\partial t}+\frac{\partial}{\partial x}\left(\mathbf{g} v_{1}\right)+\frac{\partial}{\partial y}\left(\mathbf{g} v_{2}\right)+\frac{\partial}{\partial z}\left(\mathbf{g} v_{3}\right)\right] d S,
\end{aligned}
$$

and therefore, comparing (a) with (b),

$$
p_{1}=f_{1}-g v_{1}, \quad p_{2}=f_{2}-g v_{2}, \quad p_{3}=f_{3}-g v_{3}
$$

i.e. for any direction of $n$,

$$
\mathbf{p}_{n}=\mathbf{f}_{n}-\mathbf{g}(\mathrm{vn})
$$

Omitting the operand $n$, we may write this result, in terms of th stress-operators themselves,

$$
p=f-g(v
$$

This is the required connexion between the relative stress $p$ and $t$ absolute stress $f$. Notice that, $f$ being symmetrical or self-conjugat $p$ is in general non-symmetrical, since $g$ may differ in directic from $\boldsymbol{v}$. Thus, for instance, $p_{12}=f_{12}-g_{2} v_{1}$, while $p_{21}=f_{12}-g_{1}$, Only when $\mathbf{g} \| \mathbf{v}$ does the relative stress become self-conjugate.

Let us now return to (14). Remember that, for the rest-syster $p^{\prime}=f^{\prime}$, write down $g(v$ by the second of those formulae, ar subtract it from the third one. 'Then the terms containing $u$ ' w cancel one another, and the result will be

$$
p=\epsilon p^{\prime} \epsilon-\frac{\gamma}{c^{2}} \epsilon \rho^{\prime} \mathbf{v}(\boldsymbol{v},
$$

or, if $i$ be the unit of $\mathbf{v}$,

$$
p=\epsilon p^{\prime} \epsilon-\beta^{2} \gamma . \epsilon f^{\prime} \mathrm{i}(\mathrm{i} .
$$

Such then is the transformation formula of the relative stress. In reader will find no difficulty in splitting ( 19 ) into nine Cartesia equations for $p_{11}, p_{12}$, etc., especially as this procedure has be illustrated a moment ago by the passage from (14) to ( $14(z)$. It interesting to remark that $p$ depends only upon $f^{\prime}$ and the motion the element in question, but not upon $u^{\prime}$, the density of energ

## ENERGY AND MOMENTUM

pressure (i.e. to a pressure which is purely normal and equal all directions of $\mathbf{n}$ ), either uniform or varying from point to $p$ Then the stress-operator $p^{\prime}$ degenerates into an ordinary scalar, pressure in the more familiar sense of the word.* In this cas can be written before the stretching operator, so that (19) g at once

$$
p: p^{\prime}=\epsilon^{2}-\beta^{2} \gamma \epsilon \mathrm{i}\left(\mathrm{i}=\epsilon^{2}-\beta^{2} \gamma^{2} \mathbf{i}(\mathbf{i} .\right.
$$

Now, $\epsilon^{2}=\gamma^{2} \mathrm{i}\left(\mathrm{i}+\mathrm{j}\left(\mathrm{j}+\mathrm{k}\left(\mathbf{k}\right.\right.\right.$, and $\gamma^{2}-\beta^{2} \gamma^{2}=\mathrm{r}$, so that the right sid the last formula is, in Gibbs' terminology, an idemfactor, $\mathbf{i}(\mathbf{i}+\mathbf{j}(\mathbf{j}+\mathbf{k}$ leaving unchanged any operand $n$ whatever. The result, there is that

$$
p=p^{\prime},
$$

or that isotropic pressure is a relativistic invariant. This result first obtained by Planck $\dagger$ from thermodynamical considerations a by the principle of relativity, then by Sommerfeld $\ddagger$ from what believed to be a purely geometric enunciation of the behaviou four-dimensional vectors and their projection, and, finally, by L whose method has been here adopted. The reader will find it w his while to compare the latter with the two former methods, is for that purpose referred to the papers of Planck and Somme just quoted.

So much as regards the stress and its transformation. N consider $u$ and $\mathbf{g}$, the densities of energy and of momentum which the first pair of (r4) hold. In these formulae we have to substitute the identity $f^{\prime}=p^{\prime}$. Thus, taking i along the direc of motion of the given element of the body, we have in gen that is to say, for any elastic stress $p^{\prime}$,

$$
u=\gamma^{2}\left[u^{\prime}+\beta^{2} p_{11}{ }^{\prime}\right]
$$

and

$$
\mathbf{g}=\frac{\gamma^{2}}{c^{2}}\left[u^{\prime} \mathbf{v}+\gamma^{-1} \epsilon p^{\prime} v\right],
$$


Let $d S^{\prime}$ be the rest-volume of an element of the body, and sequently $d S=d S_{0}^{\prime} / \gamma$ its $S$-volume. Then we shall have for

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energy of that individual element, as estimated from the $S$-point of view,

$$
u d S=\gamma\left(u^{\prime}+\beta^{2} p_{11}^{\prime}\right) d S^{\prime} .
$$

To obtain the whole energy $U$, this is to be integrated throughout the body. Generally speaking, there will be no simple relation between $U$ and $U^{\prime}$. For, even if $u^{\prime}$ and the stress were constant throughout the body, the value of $\beta$ and also the direction of $v$ may change from point to point. And if but one particle of the body moves with varying velocity, then the velocity will also, as a rule, vary from particle to particle. Let us suppose, however, that this heterogeneity of the inner state $\left(u^{\prime}, p^{\prime}\right)$ and of the motion of the body can be neglected. Then, if $V$ and $V^{\prime}$ be the volumes of the whole body from the two standpoints, its total energy, as estimated by the $S$-observers, will be

$$
\begin{equation*}
U=\gamma\left(U^{\prime}+\beta^{2} p_{11}^{\prime} V^{\prime}\right) . \tag{20a}
\end{equation*}
$$

We shall return to this formula presently, in order to compare the difference $U-U^{\prime}$ with the expression of kinetic energy given, for the simplest particular case, in Chapter VII.
Treating similarly the equation (2I), and making the same assump tion of homogeneity, or considering the whole body as a particle, we have, for its total momentum,

$$
\begin{equation*}
\mathbf{G}=\frac{\gamma}{c^{2}}\left[U^{\prime}+V^{\prime} \cdot \gamma^{-1} \epsilon \rho^{\prime}\right] \mathbf{v} . \tag{2Ia}
\end{equation*}
$$

We have seen in Chap. VII., formula (24), that, according to Minkowski's dynamics of a particle, the momentum of the particle would be simply $\gamma m$ times its velocity, where $m$, the rest-mass o the particle, is an ordinary scalar magnitude. Thus, according to that manner of treatment, the momentum would always coincide in direction with the velocity. This isotropic behaviour of the rest-mass appears now as the simplest particular case of formula (2 $1 a$ ), which holds for a particle conceived as the limit of an extended body We can still write

$$
\mathbf{G}=\gamma m \mathbf{v},
$$

but now $m$, instead of being a simple scalar, will be a linear vecto

## INERTIA OF ENERGY

The first part of $m$ is an ordinary scalar, namely

$$
U^{\prime} / c^{2}
$$

This is the expression of the famous inertia of energy which, a consequence of the principle of relativity, has been enunciated Einstein.* If a body gains or loses $n$ ergs of energy, say, in form of heat, then we have to look for an increase or diminut of its rest-mass by $\frac{n}{9} 10^{-20}$ grams. The second part of $m$ is due the stress. Since ${ }^{\prime \prime}$ is, in general, an operator, this part of $m$ also be an operator. It will be remembered that $p^{\prime}$, being ident with the original $f^{\prime}$, is self-conjugate. The stress, therefore, have three mutually perpendicular principal axes. Let these represented by the unit vectors $a, b, c$, each of which can be tal of course, in both its positive and negative sense. And let denote the corresponding principal pressures, which are ordir scalars, by $p_{a}{ }^{\prime}, p_{c}{ }^{\prime}, p_{c}{ }^{\prime}$. Then, if v is along a, for instance, we s . have

$$
\frac{\mathrm{r}}{\gamma} \epsilon p^{\prime} \mathrm{a}=\frac{\mathrm{I}}{\gamma} \epsilon \mathrm{a} \cdot p_{a}^{\prime}=\mathrm{a} \cdot p_{a}^{\prime},
$$

since $\epsilon a=\gamma \mathrm{a}$. Similarly, if the body happens to move along b o Thus, the principal axes of the mass-operator $m$ coincide zeith principal axes of the stress. $\dagger$ The corresponding principal val of the rest-mass are

$$
\left.\begin{array}{l}
m_{a}=\frac{\mathrm{I}}{c^{2}}\left(U^{\prime}+V^{\prime} p_{a}^{\prime}\right) \\
m_{b}=\frac{\mathrm{I}}{c^{2}}\left(U^{\prime}+V^{\prime} p_{b}^{\prime}\right) \\
m_{c}=\frac{\mathrm{I}}{c^{2}}\left(U^{\prime}+V^{\prime} p_{c}^{\prime}\right) \cdot
\end{array}\right\}
$$

* Cr. Einstein's papers in Ann. der Physik, Vol. XVIII., 1905, p. Vol. XX., 1906, p. 627, lut especially ' Ueher die vom Relativitätsprinzip g derte Trägbeit der Energie,' ibid., Vol. XXIII., 1907, p. 371. Independe of the principle of relativity, the inertia of energy, in the case of radiation, app in a valuable paper of K. v. Mosengeil, Ann. der Physik, Vol. XXII., r p. 867. The history of this concept can, of course, be traced a long way far lack. Its origin can be looked for in Maxwell's pressure of light, and in

The momentum is parallel to the velocity of the body when anc only when it happens to move along one of its principal stress-axes
Notice that, by what has been said, this anisotropy would be a property of the rest-mass itself. When, therefore, we pass to conside the acceleration of such a body, or particle, in relation to the moving force, according to the equation of motion

$$
\frac{d}{d t} \gamma m \mathbf{v}=\mathbf{N},
$$

we can no longer express the inertial behaviour of the body in terms of a 'longitudinal' and a 'transversal' mass, as in Chapter VII. The axial symmetry produced round $\boldsymbol{v}$ in that comparatively simple case was due to the assumption of a scalar rest-mass. The case now before us is much more complicated. Even if the inner state o the body is supposed to remain invariable, a full description o acceleration in connexion with force requires a linear vector operator involving six scalar inertial coefficients. The dynamics of trans lational motion of such a body is, obviously, entangled with the dynamics of its rotations. Unlike classical mechanics, these twe kinds of motion cannot, rigorously speaking, be treated separately It can be shown, by considering the moment of momentum, that to maintain such a body in uniform rectilinear motion, a certain couplc is required. Only when the constant vector-velocity v of the body coincides in direction with one of its principal stress-axes, would the moment of this couple vanish. Again, suppose that there is no impressed resultant force, i.e. that $\mathbf{N}=0$. Then the momentum wil be constant in both size and direction relative to $S$, say, equal $\mathbf{C}$, anc

$$
\gamma \boldsymbol{\gamma}=m^{-1} \mathbf{C} .
$$

If, therefore, the body rotates together with its stress-axes, the motion of translation will not be uniform and even not rectilinear Notwithstanding the absence of a resultant S-force the body may move with varying velocity relative to the framework $S$. . nd it will do so if, for instance, its initial velocity does not coincide ir direction with one of the principal stress-axes and if the couple mentioned above is not applied. But we cannot dwell any longel

## ANISOTROPY OF INERTIA

any departure from isotropy. On the other hand, it must $k$ confessed that no phenomena of this kind have been sought f expressly and that direct comparisons of inert masses (i.e. apa from gravity) could not easily be made more accurate than to or in ten or hundred thousand parts. One thing, at any rate, seen certain : If the above formulae are accepted, we cannot reasonab hope to produce observable anisotropy of mass by artificial pressur or tensions in any lump of matter. For, according to (22a), hundre of atmospheres appropriately applied would produce a departu from isotropy of mass amounting only to $10^{2} \cdot 10^{6} \mathrm{c}^{-2} \doteqdot 10^{-13}$ of gram per cubic centimetre. But for all that we know there mig be anisotropy of inertia in natural crystals, corresponding to son enormous 'latent stresses.' And to embody such stresses into seems no less, and no more, legitimate than to condense in $U^{\prime}$ much 'latent energy' as is necessary to account for the observab mass of a body. But, apart from any theory, experiments crystals seem worth trying, whether to reveal some traces anisotropic inertia or to push it below a numerically defini limit.*

Of course, if it is assumed that the stresses represented by $p^{\prime}$ ar under all circumstances, only of the order of manifest tensions ar pressures known as such from experience, then the influence the differences $p_{a}{ }^{\prime}-p_{v}{ }^{\prime}$, etc., upon inertia will be far too small to ever detected. But if so, then there will be also no sensib contribution of stress to inertia at all. Such, in fact, is the prevaili opinion.

According to this opinion the stress-term in (22), (2 I) and, slow motion, a fortiori in (20), where it appears with the coefficient $A$ can be omitted for all ordinary material bodies. But the case

[^51]
## THE THEORY OF RELATIVITY

different, of course, when the energy and the stress are electromagnetic, when the 'body' becomes simply a region c containing an electromagnetic field. Under these circumstan part played by $p^{\prime}$ is no longer negligible, unless we wish to the whole mass $m$, and therefore also the whole momentum. not only then are the pressures or tensions $p_{a}^{\prime}$, etc., of th order as the density $u^{\prime}$ of electromagnetic energy, but some c can even wholly annul the contribution of energy to mass for instance, the field in $S^{\prime}$ be a homogeneous electrostat $\mathbf{E}^{\prime}=$ const., such as is contained between the plates (discs) of vacuum-condenser, far enough from the edges of the plates. $u^{\prime}=\frac{1}{2} E^{\prime 2}$, and if a be taken along the axis of the condenser o the Faraday tubes, $p_{a}{ }^{\prime}$, being a tension proper, is equal to while $p_{b}^{\prime}, p_{c}^{\prime}$, being pressures proper, are each equal to $\frac{1}{2} E^{\prime 2}$. fore, by (22a),

$$
m_{b}=m_{c}=\frac{2 U^{\prime}}{c^{2}}=\frac{E^{\prime 2} V^{\prime}}{c^{2}},
$$

while

$$
m_{a}=0 .
$$

Thus the condenser, apart from the plates, has equal rest-ma all transversal directions, while its longitudinal principal res vanishes altogether. If it is moved along the tubes it 1 momentum. This property, which holds separately for each element of a Faraday tube, harmonizes with Sir J. J. Tho well-known representation. The tubes may be straight, as above case, or curved and of varying section. The only cor being that there shall be no flux of energy in $S^{\prime}$, we can co apply the above reasoning to any electrostatic ficld. Summ the contributions due to the elements of infinitesimal filament appropriate consideration of their directions), the mass-oper the whole field can be found. If the field is radial and symm round a point $O^{\prime}$, as in the case of the Lorentz electron, the operator $m$ degenerates into an ordinary scalar, the rest-mass electron, or rather of its whole field. The reader is recommen prove this in detail, and to compare the result to be thus ob

Let us now once more return to stresses and energies of ar origin. In the simplest case of hydrostatic or isotropic pressur whatever its order of magnitude, our above $p^{\prime}$ degenerates into ordinary scalar, so that, in ( $2 \mathrm{I} a), \gamma^{-1} \epsilon p^{\prime} \mathrm{v}=\gamma^{-1} \epsilon \mathrm{~V} \cdot p^{\prime}=\mathrm{v} \cdot p^{\prime}$, while, (20a), $p_{11}{ }^{\prime}=p^{\prime}$, and therefore

$$
\left.\begin{array}{l}
U=\gamma\left(U^{\prime}+\beta^{2} p^{\prime} V^{\prime}\right) \\
\mathbf{G}=\frac{\gamma}{c^{2}}\left(U^{\prime}+p^{\prime} V^{\prime}\right) \mathbf{v} .
\end{array}\right\}
$$

These are Planck's formulae (loc. cit.). Since isotropic pressure is invariant and $V=V^{\prime} / \gamma$, we have also

$$
\chi=U+p V=\gamma\left(U^{\prime}+p^{\prime} V^{\prime}\right)=\gamma \chi^{\prime},
$$

where $\chi^{\prime}$, the rest-value of $\chi$, is Gibbs' 'heat function for consta pressure' or enthalpy.* The momentum is now in the direction motion. The mass-operator (22) degenerates into

$$
m=\frac{U^{\prime}+p^{\prime} V^{\prime}}{c^{2}}=\frac{\chi^{\prime}}{c^{2}},
$$

the scalar rest-mass.
Thus, in the case of isotropic stress, the inertial behaviour of $t$ body, or particle, is characterized by a simple scalar, as in Chap. V] But still the rest-mass will in general vary in time, inasmuch as $t$ inner state of the particle ( $U^{\prime}, p^{\prime}, V^{\prime}$ ) may undergo changes duri its motion. If this is the case, e.g. if the enthalpy of the parti varies, then $\mathrm{SX} Y_{c}$ does not vanish, or, in other words, the M kowskian four-force $X$ is no longer perpendicular to the particl world-line. In fact, instead of equation (20), p. r94, we now ha

$$
m \frac{d Y}{d \tau}+Y \frac{d m}{d \tau}=X
$$

suspended condenser due to the earth's orbital motion was sought for a somewhat thorough exposition of this subject would be beyond the limits a purposes of the present volume, and the interested reader must therefore
and consequently, since $t^{\prime}$ can be written for the proper tir

$$
\mathrm{S} X Y_{c}=Y Y_{c} \frac{d m}{d t^{\prime}}=-c^{2} \frac{d m}{d t^{\prime}},
$$

or, by (26),

$$
\mathrm{S} X Y_{c}=-\frac{d \chi^{\prime}}{d t^{\prime}} .
$$

This proves the statement. Developing the left-hand side, b ( $\mathrm{r}_{7}$ ), p. 193, we have, in terms of the Newtonian force $\mathbf{N}$ a velocity $v$ of the particle,

$$
(\mathbf{N v})=\frac{d}{d t}\left(m c^{2} \gamma\right)-\frac{\mathrm{r}}{\gamma^{2}} \frac{d \chi^{\prime}}{d t^{\prime}}, *
$$

or also, by (25) and (26),

$$
(\mathrm{Nv})+\frac{1}{\gamma^{2}} \frac{d \chi^{\prime}}{d t^{\prime}}=\frac{d \chi}{d t} .
$$

This is now, instead of (22), p. r94, the equation of energ
To see its meaning, consider the particular case of cc pressure, or what may be called isopiestic motion. Then, if the heat communicated to the particle per unit $t^{\prime}$-time,

$$
\frac{d \chi^{\prime}}{d t^{\prime}}=\frac{d U^{\prime}}{d t^{\prime}}+p^{\prime} \frac{d V^{\prime}}{d t^{\prime}}=h^{\prime}
$$

the heat supply being estimated from the point of view system $S^{\prime}$ in which the particle is instantaneously at rest. sequently,

$$
(\mathrm{Nv})+\frac{\mathrm{I}}{\gamma^{2}} h^{\prime}=\frac{d U}{d t}+p \frac{d V}{d t} .
$$

The first term on the right is the rate of increase of the total of the particle, the second term gives the work done per un by the particle in expanding, while ( $\mathbf{N v}$ ) is the activity impressed force, everything being estimated from the $S_{\text {-point o }}$ If, therefore, (28) is to express the conservation of energy in as $\left(28^{\prime}\right)$ does with respect to $S^{\prime}$, we have to write for $h$, the heat supply as estimated from the $S$-point of view, $\dagger$

$$
h=\frac{h^{\prime}}{\gamma^{2}} .
$$

## HEAT SUPPLY

And, since $d t=\gamma d t^{\prime}$, we have to require that the relativistic cor nexion between corresponding infinitesimal amounts of heat supplie or withdrawn shall be

$$
\delta H=\frac{\mathrm{x}}{\gamma} \delta H^{\prime}
$$

This transformation formula agrees entirely with what follows fron Planck's thermodynamical investigation. In fact,* one of Professo Planck's most fundamental results is that entropy is invariant witl respect to the Lorentz transformation,

$$
\eta=\eta^{\prime},
$$

and another of his results states that temperature is transformed lik volume,

$$
\theta=\frac{\mathrm{r}}{\gamma} \theta^{\prime} .
$$

Now, the temperature being here defined in the well-known thermo dynamical way, we have, for reversible heat supply, $\delta H^{\prime}=\theta^{\prime} d \eta$ and on the other hand (granting that a process reversible is $S^{\prime \prime}$ is also reversible from the $S$-standpoint), $\delta H=\theta d \eta$, whenc $\delta H=\delta H^{\prime} / \gamma$.

But, instead of recurring to temperature and the second law o thermodynamics, the transformation formulae (29) and (30) cal equally well be considered as consequences of the principle of con servation of energy combined with (28), which in its turn is consequence of the equation of motion (23) and of the relativisti behaviour of momentum. Whatever the logical order of exposition the important thing to notice is that the several properties are con sistent with one another.

Before leaving the discussion of variable rest-mass, only one more remark. It has been shown in Chap. VIII. that the electro magnetic ponderomotive force per unit volume plus $1 / 6$ times it activity is a physical quaternion. In agreement with this the tota force $\mathbf{N}$ of Chap. VII. had the property that $\gamma[\iota(\mathbf{N v}) / c+\mathbf{N}]$ was physical quaternion. Both of these were particular instances of
and consequently, since $t^{\prime}$ can be written for the prop

$$
\mathrm{S} X Y_{c}=Y Y_{c} \frac{d m}{d t^{\prime}}=-c^{\frac{d}{2}} \frac{d m}{d t^{\prime}}
$$

or, by (26),

$$
\mathrm{S} X Y_{c}=-\frac{d X^{\prime}}{d t^{\prime}}
$$

This proves the statement. Developing the left-hand $s$ (r7a), p. r93, we have, in terms of the Newtonian forc velocity v of the particle,

$$
(\mathbf{N v})=\frac{d}{d t}\left(m c^{2} \gamma\right)-\frac{1}{\gamma^{2}} \frac{d \chi^{\prime}}{d t^{\prime}}, *
$$

or also, by (25) and (26),

$$
(\mathrm{Nv})+\frac{\mathbf{1}}{\gamma^{2}} \frac{d \chi^{\prime}}{d t^{\prime}}=\frac{d \chi}{d t}
$$

This is now, instead of (22), p. 194, the equation of
To see its meaning, consider the particular case pressure, or what may be called isopiestic motion. Tl the heat communicated to the particle per unit $t^{\prime}$-time

$$
\frac{d \chi^{\prime}}{d t^{\prime}}=\frac{d U^{\prime}}{d t^{\prime}}+p^{\prime} \frac{d V^{\prime}}{d t^{\prime}}=h^{\prime}
$$

the heat supply being estimated from the point of system $S^{\prime}$ in which the particle is instantaneously at sequently,

$$
(\mathrm{Nv})+\frac{1}{\gamma^{2}} h^{\prime}=\frac{d U}{d t}+力 \frac{d V}{d t}
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Before leaving the discussion of variable rest-mass, only one more remark. It has been shown in Chap. VIII. that the electro magnetic ponderomotive force per unit volume plus $t / c$ times it activity is a physical quaternion. In agreement with this the tota force $\mathbf{N}$ of Chap. VII. had the property that $\gamma[\iota(\mathbf{N} v) / c+\mathbf{N}]$ was physical quaternion. Both of these were particular instances of
is varying, then the above expression is no longer a quaternion. What now continues to be such a quaternion is

$$
X=\frac{d}{d \tau} m Y=\gamma\left[\frac{d m \gamma v}{d t}+c \frac{d \gamma m}{d t}\right]
$$

or

$$
X=\gamma\left[\frac{\iota}{c} \frac{d}{d t}\left(m c^{2} \gamma\right)+\mathbf{N}\right]
$$

In Chap. VII. we had simply $\frac{d}{d t}\left(m c^{2} \gamma\right)=(\mathrm{Nv})$, while now w instead of this, the equation (27). Thus, in general, for any of the body,

$$
X=\gamma\left\{\frac{c}{c}\left[(\mathbf{N v})+\frac{\mathrm{r}}{\gamma^{2}} \frac{d \chi^{\prime}}{d t^{\prime}}\right]+\mathbf{N}\right\}
$$

is a physical quaternion, and more especially, for isopiestic mo

$$
X=\gamma\left\{\frac{\iota}{c}[(\mathbf{N v})+h]+\mathbf{N}\right\} \simeq q .
$$

In all such cases, therefore, we have to add to the activity impressed force the amount of heat supplied to the body F time. This property will reappear, in the next chapter, in cor with Joule's heat in electrical conductors.

If the enthalpy $\chi^{\prime}$, and therefore also the rest-mass of the kept constant, we fall back to the simple case treated in Chapt The activity then becomes, by (27),

$$
(\mathrm{Nv})=\frac{d}{d t}\left(m c^{2} \gamma\right)
$$

identical with (22), p. 194. Using the form (27a), we m write, equivalently,

$$
(\mathbf{N v})=\frac{d \chi}{d t}=\frac{d U}{d t}+p \frac{d V}{d t},
$$

which reads : Work done upon the body $=$ increase of its ener work done by the body in expanding. The corresponding co

$$
\chi^{\prime}=U^{\prime}+f^{\prime} V^{\prime}=\text { const. }
$$

## CONSTANT REST-MASS

of the moving body may be invariable, i.e. $U^{\prime}, p^{\prime}$ as well as $V^{\prime} \mathrm{m}$ be kept constant. But even then the work done by expansion do not disappear from ( $32 a$ ) unless the motion is uniform. For, wi constant $V^{\prime}$, we have

$$
\frac{d V}{d t}=V^{\prime} \frac{d \gamma^{-1}}{d t},
$$

which expresses the varying FitzGerald-Lorentz contraction. B whatever the way in which $\chi^{\prime}$ is kept constant, we have the san equation of motion as in Chap. VII.,

$$
m \frac{d \gamma \mathbf{v}}{d t}=\mathbf{N},
$$

and consequently the longitudinal and the transversal masses retu to their rights, being again given by

$$
m_{l}=m \gamma^{3}, \quad m_{t}=m \gamma,
$$

where $m$ has now the explicit meaning

$$
m=\frac{x^{\prime}}{c^{2}}=\frac{U^{\prime}+p^{\prime} V^{\prime}}{c^{2}}
$$

Finally, if the pressure $p^{\prime}$, and therefore also $p$, is assumed to vanis the equation of energy becomes

$$
(\mathrm{Nv})=\frac{d U}{d t},
$$

and the constancy of the rest-mass

$$
m=\frac{U^{\prime}}{c^{2}}
$$

means constancy of the particle's store of energy. In this case tl difference between the energies $U$ and $U^{\prime}$ can be looked at entirely due to the motion of the particle and called its kinet energy relative to $S$. The value of the kinetic energy thus defin is identical with that given on p. 195. In fact, the first of (2 becomes now $U=\gamma U^{\prime}$, so that

$$
\begin{aligned}
U-U^{\prime} & =(\gamma-\mathrm{r}) U^{\prime}=m c^{2}(\gamma-\mathrm{r}) \\
& =\frac{1}{2} m v^{2}\left(\mathrm{r}+\frac{3}{4} \beta^{2}+\frac{5}{8}{ }^{4}+\ldots\right) .
\end{aligned}
$$

It may be useful to illustrate here the mass formula (33) numerical examples. Thus, taking $2 \cdot \mathrm{I}$ gram calories for called the solar constant (energy received from the sun pe1 per $\mathrm{cm}^{2}$. at the earth's mean distance), we have for the su radiation per minute

$$
4 \pi\left(\mathrm{I} \cdot 5 \cdot 10^{13}\right)^{2} \cdot 2 \cdot 1 \cdot 4 \cdot 2 \cdot 10^{7} \mathrm{ergs},
$$

so that the diminution of the sun's mass due to radiation $w$ per minute, $2 \cdot 8 \cdot 10^{14}$ grams, and per year

$$
\delta m=\mathrm{r} \cdot 5 \cdot \mathrm{r} \mathrm{o}^{20} \text { grams. }
$$

This seems at first a prodigious loss; but the sun's ma. $2.10^{33} \mathrm{gr}$., the proportionate loss per year,

$$
\frac{\delta m}{m}=\frac{3}{4} \cdot 10^{-19}
$$

is quite insignificant. Next, take the example adduced by A mixture of 2 gr . of hydrogen and r 6 gr . of oxygen dev the act of producing water, at ordinary pressure and teml $2.9 .10^{12}$ ergs of heat; the corresponding diminution of mass to $3.2 .10^{-9}$ gr., and the proportionate loss due to this reaction,

$$
\frac{\delta m}{m}=2 \cdot 10^{-10}
$$

would again be far too small to be observed. Numbers o order would result for other instances of chemical In short, the 'latent energy' which (if we neglect the con due to stress) is to account for mass does not manifest any one of those processes in which atoms are implied as We are thus driven back to the interior of the old atom, and have to look for that energy in the disintegr atoms known in connexion with radioactive phenomena. if we are to judge from their observed heat-effect o amounts of energy developed in such processes exceed im all those liberated in ordinary chemical reactions, and Planck seems to see in radioactivity a kind of verific the energetic theory of inertia. Now, it is true that th
quote Planck's own example, one gram-atom of radium would los through its heat production ( 30240 gram calories per hour) or miligr. of its mass per year; the proportionate loss, therefor amounting to

$$
\delta m / m=\cdot 5 \cdot 10^{-7} \text { per annum }
$$

is again too small to be observed.
We may add that the latter $\delta m$ is small even when compare with the mass disintegrated during the same interval of time. F this amounts, per 225 gr . of radium present, and per year, $m_{\text {dis }}=9.10^{-2}$ gr., so that, in round figures,

$$
\delta m / m_{\mathrm{dis}}=10^{-4} .
$$

The mass of the disintegrated parent substance reappears sensib undiminished in the masses of the descendants.

Thus, even radioactive phenomena reveal to us practica nothing of the assumed latent energy $c^{2} m$. Its bulk remains latent as anything ever was. It must, therefore, be confessed th the energetic theory of rest-mass, attractive and promising as it m seem, has for the time being the character of a purely formal red tion of one concept to another. Nobody doubts, of course, that $t$ chemical atoms are themselves exceedingly complicated systems, a that there are therefore many ways left of throwing the chief sto of latent energy upon a host of ultra-atomic entities, electrons what not. If so, then some spontaneous disintegration, affecting atomic structure even more profoundly than that which in our d is associated with the name of radioactivity, may induce the gates those copious stores to open to the human eye. But as yet we h not the least knowledge of such phenomena. It is for this reas we have said that it is equally legitimate to assume latent stres along with the manifest ones in the mass formulae as to assu latent energies. Both are originally defined only by their variatic in time and in space respectively. And, for the present, $b$ would have a purely formal character.

The above mechanical, and partly thermodynamical, subje have been treated at some length because of their affinity the fundamental electroma netic equations for vacuum. Returr

## NOTES TO CHAPTER IX.

Note $\mathbf{1}$ (to page 234). Let $\mathbf{i}, \mathbf{j}, \mathbf{k}$ be the antecedents, and $\mathbf{f}_{1}$, consequents* of the stress-dyadic $f$. Thus, if $f$, as in equation (5) applied as a post-factor,

$$
\left.\left.f=\mathrm{i}) \mathrm{f}_{1}+\mathrm{j}\right) \mathrm{f}_{\mathbf{2}}+\mathrm{k}\right) \mathrm{f}_{3} .
$$

This means that $\mathrm{if}=(\mathrm{ii}) \mathrm{f}_{1}=\mathrm{f}_{1}$, etc., and in general,

$$
\begin{aligned}
\mathrm{n} f & =(\mathrm{ni}) \mathrm{f}_{1}+(\mathrm{nj}) \mathrm{f}_{2}+(\mathrm{nk}) \mathrm{f}_{3} \\
& =n_{1} \mathrm{f}_{1}+n_{2} \mathrm{f}_{2}+n_{3} \mathrm{f}_{3},
\end{aligned}
$$

which is equal to $\mathbf{f}_{n}$, as it should be. Similarly, writing instead Hamiltonian $\nabla$,

$$
\begin{aligned}
\nabla f & =(\nabla \mathrm{i}) \mathbf{f}_{1}+(\nabla \mathrm{j}) \mathbf{f}_{2}+(\nabla \mathbf{k}) \mathbf{f}_{3} \\
& =\frac{\partial \mathbf{f}_{1}}{\partial x}+\frac{\partial \mathbf{f}_{2}}{\partial y}+\frac{\partial \mathbf{f}_{3}}{\partial z} .
\end{aligned}
$$

Using the notation of Gibbs, Scientific P'apers, Vol. II. p. 76, w write $f=\mathrm{if}_{1}+\mathrm{jf}_{2}+\mathrm{kf}_{3}$, and

$$
\frac{\partial \mathbf{f}_{1}}{\partial x}+\frac{\partial \mathbf{f}_{2}}{\partial y}+\frac{\partial \mathbf{f}_{3}}{\partial z}=\nabla \cdot f,
$$

where the dot means scalar application of $\nabla$. But since, in our prescription of applying $\nabla$ scalarly is already given by the open par in the dyadic (a), we do not require the dot or any other symbol multiplication.
Note 2 (to page 238). Let $h$, as in Note 6 to Chapter VIII., kowski's alternating matrix equivalent to the electromagnetic i.e. let, according to (c), p. 230,

$$
h=\left|\begin{array}{rrrr}
\circ, & M_{3}, & -M M_{2}, & 1 E_{1} \\
-M_{3}, & \circ, & M_{1}, & \cdots\left(E_{2}\right. \\
M_{2}, & -M_{1}, & 0, & -1 E_{3} \\
1 E_{1}, & \iota E_{2}, & \iota E_{3}, & 0
\end{array}\right| .
$$

Multiply it into itself. Then the first constituent of the first ro resulting matrix $h h$ will be

$$
(h h)_{11}=-M_{2}^{2}-M_{3}^{2}+F_{1}^{2}=-f_{11}+\lambda,
$$

where $f_{11}$ is the corresponding component of the Maxwellian str $\lambda=\frac{1}{2}\left(M^{2}-E^{2}\right)$ the electromagnetic Lagrangian function per unit Similarly,

## STRESS-ENERGY MATRIX

where $u$ is the density of electromagnetic energy and $\mathbf{g}$ that of momentur as throughout the chapter. Thus

$$
-(h h)=\left|\begin{array}{rrrr}
f_{11}+\lambda, & f_{12}, & f_{13}, & \operatorname{lcg} g_{1} \\
f_{21}, & f_{22}+\lambda, & f_{23}, & \operatorname{lcg} 2 \\
f_{31}, & f_{32}, & f_{33}+\lambda, & \operatorname{cog} 3 \\
\operatorname{cog}, & \operatorname{cog} 2, & \operatorname{cog}, & -u+\lambda
\end{array}\right|=\mathfrak{s}+\lambda,
$$

where $\mathbb{S}$ is the matrix defined by (ri), p. 238, and $\lambda$ is written $f$ $\lambda$ times the unit matrix of $4 \times 4$ constituents. The required connexion therefore,

$$
S=-h h-\lambda .
$$

It will be remembered that $\lambda$ is one of the invariants of $h$. And, $\sin$ $h^{\prime}=\bar{A} h A$, the last equation gives at once

$$
\mathbb{S}^{\prime}=\bar{A} \mathfrak{S} A
$$

in agreement with (13), p. 239. On the quaternionic scheme we har instead of (b), the operator $\mathbf{R}[] \mathbf{L}$ of an analogous and somewhat simpl structure.

## CHAPTER X.

## MINKOWSKIAN ELECTROMAGNETIC EQUATIONS PONDERABLE MEDIA.

In Minkowski's notes, which after his death were worked out M. Born,* the electromagnetic equations for moving bodies, sati rigorously the principle of relativity, are deduced in a very ing way from the fundamental equations of the electron theory. since the electronic equations were previously known to be inv with respect to the Lorentz transformation, and gave to the rel his first standard magnitudes, such a deduction was certainl desirable and interesting. In fact, it occupied Minkowski's th vividly during his last days. But in his own paper of 190 peatedly quoted, Minkowski adopts a purely phenomeno method, and deduces the equations for moving bodies, now rally associated with his name, from Maxwell's equations for stat media by subjecting them to a Lorentz transformation.

In the present chapter we shall avail ourselves of the latter $m$ only, which, apart from other considerations, recommends its its mathematical simplicity. Readers, and especially those desire to see the electron theory made the foundation of all $e$ magnetic science, are referred to Dr. Born's paper just quoted, the resulting equations $\dagger$ are wholly identical with Minko original equations to be given presently.

[^52]We shall retain here the notation adopted in Chapter II., wher Maxwell's equations for a perfect insulator are collected under (3 p. 26. In the more general case of a conducting body, we have supplement the displacement current by the conduction curren The latter, reckoned per unit area, we shall denote by $\mathrm{I}^{\prime}$, and th electrical conductivity by $\sigma$. Thus, Maxwell's equations, written $f$ the system $S^{\prime}$, in which the ponderable body is at rest, will consist the two groups:

$$
\left.\begin{array}{l}
\frac{\partial \mathfrak{E}^{\prime}}{\partial t^{\prime}}+\mathbf{I}^{\prime}=c \cdot \operatorname{curl}^{\prime} \mathbf{M}^{\prime} ; \operatorname{div}^{\prime} \mathfrak{\mathbb { E } ^ { \prime }}=\rho^{\prime} \\
\frac{\partial \mathfrak{Z l t}}{\partial t^{\prime}}=-c \cdot \operatorname{curl}^{\prime} \mathbf{E}^{\prime} ; \operatorname{div}^{\prime} \mathfrak{G} \mathfrak{t}^{\prime}=0,
\end{array}\right\}
$$

independent of the properties of the particular body, and

$$
\mathfrak{G} \mathbb{E}^{\prime}=K \mathbf{E}^{\prime}, \quad \mathfrak{H} \mathfrak{H}^{\prime}=\mu \mathbf{M}^{\prime}, \quad \mathbf{I}^{\prime}=\sigma \mathbf{E}^{\prime},
$$

containing its specific 'constants.' These, the permittivity, induc vity (or permeability) and conductivity, which hereafter will play t] part of invariants,* may be either simple scalars (more generall linear vector operators) if dispersion is disregarded, or otherwi compound differential operators. In the latter case (in which prac cally $K$ alone is concerned) the operator $K$ is to be express constructed so as to be invariant. Thus it may consist of derivatio of any order with respect to the proper time of the body.

In what follows we shall limit ourselves to isotropic media, so th $K, \mu, \sigma$ will have at any rate a scalar character, being either scal magnitudes or scalar operators involving differentiations.

Let now $S$ be another system of reference (say, the earth), relati to which our ponderable medium, together with its rest-system moves with a uniform velocity v. Assuming the rigorous validity Maxwell's equations ( $\mathrm{r}^{\prime}$ ) and ( $2^{\prime}$ ) in $S^{\prime}$, and subjecting them to t appropriate Lorentz transformation, we shall obtain two groups equations for the $S$-standpoint. Call them (x) and (2). Wh properties are we to require from ( r ) and (2) in the name of t
principle of relativity? In the previous case of vacuum, there was nothing to be carried along with the observers, a timate systems, $S, S^{\prime}, S^{\prime \prime}, \ldots$ were wholly equivalent to one ar and the relativistic requirement was simply invariance or preser of form of the equations. The case before us is different. ponderable dielectric, with its specific properties, is at rest system at a time, and moves relatively to all other systems. rest-system, in our concrete case $S^{\prime}$, is an uniquely privileged work. If, in other concrete cases, the body were fixed in $S$ or and so on, we should have to require the non-dashed or the $d$ dashed equations to be of the same form as the above ( $r^{\prime}$ ) an But, $S$ being a system, relative to which the body does (uniformly), we have to require only that the groups of equ ( 1 ) and (2), which might both contain the velocity $\nabla$,* shor invariant with respect to the Lorentz transformation by me which we pass from $S$ to any other legitimate system. If this rt ment were not fulfilled, Maxwell's equations could not be us relativistic purposes at all. But, as a matter of fact, they stan test completely.

It seemed advisable to dwell a little upon these explana firstly, to avoid possible misunderstanding, and secondly, becau procedure and the test here exemplified are of general impor They are the same in every other case in which the rela equations to be constructed concern any phenomena in pond bodies.

In order to obtain the two groups of equations, numbered ir cipation ( I ) and (2), and to see at the same time their invarianc

$$
\mathbf{L}^{\prime}=\mathfrak{A l} \mathfrak{A l}^{\prime}-\iota \mathbf{E}^{\prime}, \quad \mathbf{R}^{\prime}=\mathfrak{y} \mathfrak{t l}^{\prime}+\iota \mathbf{E}^{\prime}
$$

and

$$
\text { 焎 }^{\prime}=\mathbf{M}^{\prime}-\iota \mathbb{E}^{\prime}, \quad \text { 退 }{ }^{\prime}=\mathbf{M}^{\prime}+\iota \mathbb{E}^{\prime},
$$

and similarly for the non-dashed letters. Further, introduc quaternion

$$
C^{\prime}=\iota \rho^{\prime}+\mathbf{I}_{c}^{\prime},
$$

# PONDERABLE MEDIA 

:our equations ( $\mathrm{r}^{\prime}$ ) will assume the quaternionic form

$$
\left.\begin{array}{c}
D^{\prime} \mathbb{R O}^{\prime}-\mathrm{FR}^{\prime} D^{\prime}=2 C^{\prime} \\
D^{\prime} \mathbf{L}^{\prime}+\mathbf{R}^{\prime} D^{\prime}=0 .
\end{array}\right\}
$$

dentical with the first and the second pairs of ( $\mathrm{I}^{\prime}$ )
? [ $] Q$ be our usual transformer from $S$ to $S^{\prime}$, and there= the inverse transformer. Apply the latter to each of $\mathrm{s}\left(\mathrm{r}^{\prime} a\right.$ ) and insert $Q_{0} Q=\mathrm{x}$ between $D^{\prime}$ and $\mathbb{C}^{\prime}{ }^{\prime}$ (or $L^{\prime}$ ), between $\mathbb{R}^{\prime}$ (or $\mathbf{R}^{\prime}$ ) and $D^{\prime}$, in very much the same way - Then the result will be

$$
D Q^{\mathbb{N P}^{\prime}} Q_{c}-Q_{c} \mathrm{ZR}^{\prime} Q D=2 Q_{c} C^{\prime} Q,
$$

$r$ for the second equation, i.e.

$$
\begin{align*}
D & \text { 靱 } D \tag{xa}
\end{align*}=2^{C} C,
$$

 id $C=Q_{c} C^{\prime} Q_{c}$. Conversely,

$$
\mathbb{C N}_{\mathbb{C}^{\prime}}=Q_{c} \mathbb{C C} Q, \text { etc. }, \quad C^{\prime}=Q C Q .
$$

$C=\iota \rho+\mathrm{I} / c$ is a physical quaternion, $\mathbb{C}$ and $\mathbf{L}$ are leftical bivectors, and 7R and $\mathbf{R}$ right-handed ones.* $\quad C$ may he (macroscopic) current-quaternion, while the electrovectors need no special names.
$S$-equations ( $\mathrm{r} a$ ) are precisely of the same form as those, e rest-system. And so they will be also for every other $r$ stem of reference. The velocity of the body does not, rr into these differential equations at all. We can now reir quaternionic form ( $\mathrm{x} a$ ) to the vectorial one, and shall the required first group of equations:

$$
\begin{equation*}
\frac{\partial \mathscr{H}}{\partial t}+\mathbf{I}=c . \operatorname{curl} \mathbf{M}, \text { etc. }, \tag{x}
\end{equation*}
$$

in ( $\mathrm{I}^{\prime}$ ) without the dashes. At the same time we have

This property finds its immediate expression in the above nionic form ( $\mathrm{r} a$ ).*

Moreover, the stated transformational properties of the magnetic bivectors and of the current-quaternion lead at onc second group of equations for the moving body, to be deduc the Maxwellian connexions ( $2^{\prime}$ ). In fact, since both $\mathbf{L}=\mathfrak{y l}$ $\mathbb{C}=\mathbf{M}$ - $\iota \mathbb{C}$ are left-handed bivectors, $\dagger$ we have in exactly th way as on p. 210 , writing again $\epsilon$ for the longitudinal stre ratio $\gamma=\left(\mathrm{r}-\boldsymbol{v}^{2} / c^{2}\right)^{-\frac{1}{2}}$,

$$
\begin{aligned}
& \mathbf{E}=\gamma\left[\frac{\mathrm{x}}{\epsilon} \mathbf{E}^{\prime}-\frac{\mathrm{I}}{c} \nabla \nabla \mathbb{C} \mathfrak{f l}^{\prime}\right] ; \mathbb{E}=\gamma\left[\frac{\mathrm{I}}{\epsilon} \mathbb{E}^{\prime}-\frac{\mathrm{I}}{c} \mathrm{~V} \boldsymbol{\nabla} \mathbf{M}^{\prime}\right],
\end{aligned}
$$

whence, by the first and second of the connexions ( $2^{\prime}$ ), at an easy rearrangement of terms,

$$
\begin{aligned}
& \mathbb{E}+\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{M} \mathbf{M}=K\left[\mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{VV}\{\mathbb{A}]\right. \\
& \mathfrak{y t}-\frac{\mathrm{I}}{c} \mathrm{VVE}=\mu\left[\mathbf{M}-\frac{\mathrm{I}}{c} \mathrm{~V} \mathbb{E}\right] .
\end{aligned}
$$

Both of these relations, involving the substantial properties medium, contain its velocity. Again, since $C=\iota \rho+\mathbf{I} / c$ is a quaternion, we have, by ( $\mathrm{r}^{\prime} b$ ) of Chap. V., p. 125 ,

$$
\mathbf{I}=\boldsymbol{\epsilon} \mathbf{I}^{\prime}+\gamma \rho^{\prime} \mathbf{\nabla}
$$

and

$$
\rho^{\prime}=\gamma\left[\rho-\frac{\mathbf{I}}{c^{2}}(\mathbf{I} \mathbf{v})\right],
$$

whence, by the last of ( $2^{\prime}$ ),

$$
\mathbf{I}=\sigma \epsilon \mathbf{E}^{\prime}+\gamma^{2}\left[\rho-\frac{\mathbf{x}}{c^{2}}(\mathbf{I} \mathbf{\nabla})\right] \mathbf{\nabla}
$$

[^53]$$
\operatorname{lor} h=-s, \quad \text { lor } H^{*}=0,
$$

## PONDERABLE MEDIA

But $\frac{1}{\gamma} \mathbf{E}^{\prime}=\frac{\mathrm{I}}{\epsilon} \mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{Vv} \nmid \mathfrak{f l}$. Hence, after a slight rearrangemen terms,

$$
\mathbf{I}-\rho \mathbf{\nabla}=\mathfrak{E}=\sigma \gamma \frac{\mathrm{I}}{\epsilon^{2}}\left[\mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{VV} \mathfrak{d l}\right] .
$$

Thus I appears as the sum of the convection current $\rho \mathrm{v}$ and conduction current, for which we have written $\mathfrak{F}$, the latter b proportional to the conductivity.*

Using the convenient abbreviations

$$
\begin{aligned}
& \mathbf{E}^{\times}=\mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{\nabla} \text { 象 }, \quad \mathbb{E}^{\times}=\mathbb{E}+\frac{\mathrm{I}}{c} \mathrm{~V} \mathbf{\nabla M}
\end{aligned}
$$

and gathering together the above results, we obtain the req second group of equations, valid from the standpoint of system $S$,

$$
\left.\begin{array}{c}
\mathbb{E}^{\times}=K \mathbf{E}^{\times}, \quad \mathscr{E l}^{\times}=\mu \mathbf{M}^{\times} \\
\mathbf{I}-\rho \boldsymbol{V}=\mathbb{E}=\sigma \gamma \frac{\mathrm{I}}{\epsilon^{2}} \mathbf{E}^{\times} .
\end{array}\right\}
$$

These three connexions involve the velocity of the ponde medium relative to that system. It remains only to prove they are invariant with respect to the Lorentz transform Now, introducing the velocity-quaternion

$$
Y=\gamma[\iota c+\mathbf{v}],
$$

we have, identically,

$$
\begin{aligned}
& \eta \equiv \frac{\mathbf{I}}{2 c}[Y \mathbf{L}-\mathbf{R} Y]=\gamma\left[\frac{\iota}{c}\left(\mathbf{E}^{\times} \mathbf{V}\right)+\mathbf{E}^{\times}\right], \\
& \frac{\mathrm{r}}{2 a c}[Y \mathbf{L}+\mathbf{R} Y]=\gamma\left[\frac{L}{c}\left(\mathfrak{G l}^{\times} \mathbf{v}\right)+\mathfrak{G l} \mathrm{l}^{\times}\right], \\
& \frac{x}{2 c}[Y \mathbb{C}-\mathbb{Z} Y]=\gamma\left[\frac{\iota}{c}\left(\mathbb{E}^{\times} \times v\right)+\mathfrak{E}^{\times}\right] \text {, } \\
& \zeta \equiv \frac{I}{2 c \iota}[Y \mathbb{C}+\mathbb{Z} Y]=\gamma\left[\frac{\iota}{c}\left(\mathbf{M}^{\times} \mathbf{V}\right)+\mathbf{M}^{\times}\right],
\end{aligned}
$$

and each of these expressions* is a physical quatern Moreover, starting from the current-quaternion $C$ and its $C_{c}$, we easily obtain the identical equation

$$
\frac{c}{2}\left[C+\frac{\mathbf{I}}{c^{2}} Y C_{c} Y\right]=\epsilon^{2} \mathbf{I}-\rho \gamma^{2} \mathbf{v}+\frac{\iota}{c}\left[(\mathbf{I} \mathbf{v})-\rho v^{2}\right] \gamma^{2},
$$

of which the left-hand side is, obviously, again a physical qu So also is its right-hand side, which, by the third of (2), to $\sigma \eta$. Using, therefore, the above identities we can whole of (2), in terms of physical quaternions alone,

$$
\left.\begin{array}{rl}
Y \text { 暂 } Y & =K[Y \mathbf{L}-\mathbf{R} Y] \\
Y \mathrm{~L}+\mathbf{R} Y & =\mu[Y \mathbb{C}+\mathbb{Z} Y] \\
C+\frac{\mathrm{r}}{c^{2}} Y C_{o} Y & ={ }_{\sigma_{2}^{2}}^{\sigma}[Y \mathrm{~L}-\mathbf{R} Y]
\end{array}\right\}
$$

This proves the invariance of the relations (2) with ro the Lorentz transformation. $\dagger$ Thus the whole of equa and (2) satisfy the principle of relativity. Q.E.D.

It is worth noticing here that the world-vector corre to the quaternion

$$
\frac{1}{2}\left[C+\frac{\mathrm{r}}{c^{2}} Y C_{c} Y\right]
$$

is the part of the four-current $C$ normal to the four-vel Generally, for any pair of physical quaternions $a, b$, the e

$$
\frac{1}{2} a-\frac{b a_{c} b}{2(\mathrm{~T} b)^{2}}
$$

represents that part of the four-vector corresponcling to is normal to the four-vector $b$ (Note 1). The above $s$ is deduced from this, remembering that $\mathrm{T} Y=c c$.

[^54]In the course of the above calculations we came across the formula $\rho^{\prime} / \gamma=\rho-(\mathrm{Iv}) / c^{2}$. Its inversion will be

$$
\rho=\gamma\left[\rho^{\prime}+\frac{\mathrm{I}}{c^{2}}\left(\mathbf{I}^{\prime} \mathbf{v}\right)\right] .
$$

Substituting here $\left(\mathbf{I}^{\prime} \boldsymbol{v}\right)=\gamma(\mathbf{I} \boldsymbol{v})-\gamma \rho \nu^{2}$ and remembering that $\mathrm{I}=\mathbb{£}+\rho \mathbf{v}$, we obtain the interesting relation

$$
\begin{equation*}
\rho=\gamma \rho^{\prime}+\frac{\gamma^{2}}{c^{2}}(£ \mathbf{v}) \tag{3}
\end{equation*}
$$

about which a few words will be said later on. To resume the above results:

The equations for a moving isotropic* conducting dielectric, obtained from Maxwell's equations for stationary media, are invariant with respect to the Lorentz transformation. They consist $I^{\circ}$ of a set of differential equations not containing the velocity of motion at all, and $2^{\circ}$ of a set of relations concerning the substantial properties of the medium and involving its velocity $\nabla$ relative to the observing system. The quaternionic form of these two sets of equations is given in ( $\mathrm{I} a$ ) and ( $2 a$ ), where $\mathbb{C}, \mathrm{L}$ are left-handed and $\mathfrak{K}, \boldsymbol{R}$ right-handed physical bivectors, and $C$ a physical quaternion, $\simeq q$. The vector form of the first set is

$$
\left.\begin{array}{cc}
\frac{\partial \mathbb{E}}{\partial t}+\mathbf{I}=c \cdot \operatorname{curl} \mathbf{M} ; & \operatorname{div} \mathbb{E}=\rho  \tag{I}\\
\frac{\partial \mathfrak{n l}}{\partial t}=-c \cdot \operatorname{curl} \mathbf{E} ; & \operatorname{div} \mathfrak{G t}=0
\end{array}\right\}
$$

and that of the second set

$$
\left.\begin{array}{l}
\mathbb{E}^{\times}=K \mathbf{E}^{\times}, \quad\left\{\mathfrak{Z}=\mu \mathbb{M}^{\times}\right.  \tag{2}\\
\mathrm{I}-\rho \boldsymbol{\nabla}=\mathfrak{E}=\sigma \gamma \epsilon^{-2} \mathbf{E}^{\times},
\end{array}\right\}
$$

where $\mathbf{E}^{\times}$stands for $\mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{~V} \boldsymbol{f}$ fll, etc., as in (A), and $K, \mu, \sigma$ for the permittivity, inductivity and conductivity of the body, as originally defined from the standpoint of the rest-system.

These are Minkozeski's equations. They were first giv fundamental paper of 1907 , in both their vectorial and $r$ forms already quoted. We may notice here that Minkowsl assumed that Maxwell's equations ( $\mathrm{r}^{\prime}$ ) and ( $\mathrm{a}^{\prime}$ ) are vali corresponding instantaneous rest-system $S^{\prime}$ ) at each poir material body, whatever the state of motion around that F as if the whole body were fixed in $S^{\prime}$. It is this that he 'first axiom' (loc. cit., §8). Such being Minkowski's start he asserts, consequently, the validity of the resulting equa and (2) for each element of a material medium moving in an manner with respect to the framework $S$, in short, for $\mathbf{v}$ both space and time. His only restriction is that $v<c$. not unlikely that the first set of Minkowski's equations o such a general validity. (Notice that these are, properly two equations for five vectors, otherwise yet unconnected.) case is different when the first set is supplemented by the For, apart from other reasons, if we pass to $K=\mu=\mathrm{I}$ and whole of equations (r), (2) reduce, as will be seen pres the vacuum-equations, and the acceptance of the latter fo works whose relative motion is variable, would require a reconstruction of the principle of relativity underlying tl theory. Retaining, therefore, this principle, we can consi kowski's equations as rigorously valid only for uniform Accordingly our $\mathbf{\nabla}$ has been treated from the outset as a vector and $Y$ as a constant velocity-quaternion belonging body as a whole. Of course, as an approximation of $m$ sufficient accuracy, the equations ( 1 ) and (2) can well be velocities experiencing all such time- and space-variation practically realizable. Thus, for instance, they can safely b to bodies kept rotating, as in the case of Wilson's experim unequal FitzGerald-Lorentz contraction and the ensuing st its influence upon $K$, etc., being of the order of $\beta^{2}$.

The comparison of the equations ( I ), (2) with those 0 Heaviside, Lorentz and Cohn, none of which satisfy $r$ the principle of relativity, must be left to the reader given at sufficient length in Minkowski's pa er. As

## VACUUM AS LIMITING CASE

invariant with respect to the Newtonian, and not to the Lore transformation.

Let us now stop a while at Minkowski's equations in order to le some of their properties.

In the first place, if $K=\mu=\mathrm{I}$ and $\sigma=0$, then Minkowski's equati reduce at once to the fundamental or the vacuum-equations. In fa in this limiting case we have, by the third of (2), I= $\mathbf{V}$, and if $\rho^{\prime}=$ also $\rho=0$, by (3). Again, by the first and second of (2), $\mathbb{E}^{\times}=$ and $\mathfrak{y l}{ }^{\times}=\mathbf{M}^{\times}$, i.e.

$$
\begin{aligned}
\mathfrak{E}-\mathbf{E} & =\frac{\mathbf{I}}{c} V \nabla[\mathfrak{G t}-\mathbf{M}], \\
\mathfrak{G t}-\mathbf{M} & =-\frac{\mathrm{I}}{c} V \nabla[\mathfrak{E}-\mathbf{E}],
\end{aligned}
$$

whence, by elimination,

$$
\mathfrak{E}-\mathbf{E}=\beta^{2}(\mathbb{H}-\mathbf{E}),
$$

and since $\beta \neq \mathrm{I}, \mathfrak{E}=\mathbf{E}$, and similarly, $\mathfrak{f t}=\mathbf{M}$. Q.E.D. The sa result may be obtained from the quaternionic form (2a). In present case $\mathbb{Z}=\mathrm{L}$ becomes identical with the electromagn bivector of the preceding chapters. And since at the same ti $\mathfrak{q} \mathbb{R}=\mathbf{R}$, the sum of the equations ( $\mathrm{x} a$ ) gives at once $D \mathbf{I}=$ Properly speaking, to obtain $K=\mu=\mathrm{r}, \sigma=0$, we have (on the elec atomistic doctrine) to consider a region outside the electrons, ol least outside electronic assemblages crowded within atomic regic Then $\rho=0, D \mathbf{L}=0$, and here the macroscopic bivector coinci with our previous microscopic $\mathbf{L}$. Thus the announced reduct becomes complete.

As regards the meaning of the vector I, we have already remar that it is the sum of the convection- and the conduction-current. virtue of the properties of the stretcher $\epsilon$, the longitudinal compon of the latter current will be

$$
\mathfrak{F}_{1}=\frac{\sigma}{\gamma} E_{1}^{\times}=\frac{\sigma}{\gamma} E_{1},
$$

and the transversal ones

$$
¥_{2}=\sigma \gamma E_{2}^{\times}, \quad ¥_{3}=\sigma \gamma E_{3}^{\times} .
$$

way of splitting the conduction current into factors. But sinc far, the only requirement is that 'resistance' should reduce to $v=0$, we may equally well give the name of 'electromotive to the line-integral of the vector $\mathbf{E}^{\times}$itself; then we shall ha specific resistance-operator $\epsilon^{2} / \sigma \gamma$, instead of an ordinary If second-order terms are neglected, the distinction disappears. conduction current may then be written, with more than su approximation,

$$
\mathfrak{E} \doteqdot \sigma \mathbf{E}^{\times} .
$$

We will not stop here to discuss the nomenclature propo various authors for $\mathbf{E}^{\times}$and its magnetic companion. It advisable to leave them for the time being without any names.

The integral properties of $\mathbf{E}^{\times}$and $\mathbf{M}^{\times}$, in relation to $\mathfrak{y t l}$, et at once be put into a form with which the reader has $b$ familiar in Chapter II. In fact, by (A) and (I), we have

$$
\begin{aligned}
& -c \cdot \operatorname{curl} \mathbf{E}^{\times}=-c \cdot \operatorname{curl} \mathbf{E}-\operatorname{curl} \operatorname{VV}\left\{\begin{array}{l}
\text { flt }
\end{array}\right.
\end{aligned}
$$

and this is precisely what in Note 2 to Chap. II. has been calle
current (気tl).

That is to say, if $d \sigma$ be a surface element composed always same particles of the body, and $\mathbf{n}$ the normal of $d \sigma$, we have

$$
c\left(\mathbf{n} \cdot \operatorname{curl} \mathbf{E}^{\times}\right)=-\frac{d}{d t}(\sqrt[f t l n]{d \sigma})
$$

Similarly,

$$
c\left(\mathbf{n} \cdot \operatorname{curl} \mathbf{M}^{\times}\right)=\frac{d}{d t}(\mathbb{f} \mathrm{f} \mathrm{n} d \sigma)+(\mathrm{£n}) .
$$

Recurring, therefore, to Stokes' theorem, we have for any sur which together with its bounding circuit $s$ is carried along zu body,

$$
\int_{(0)}\left(\mathbf{M}^{\times} d \mathbf{s}\right)=\frac{\mathrm{I}}{c} \int(\mathfrak{£ n}) d \sigma+\frac{\mathrm{I}}{c} \frac{d}{d t} \int(\mathbb{E} \mathbf{n}) d \sigma,
$$

## BOUNDARY CONDITIONS

and to p. 30 , the reader will find this integral form of equation most suitable for a direct comparison of Hertz's theory wit that of Minkowski. Instead of Hertz's $\mathbf{E}, \mathbf{M}$ we have here $\mathbf{E}$ $\mathbf{M}^{\times}$, and instead of his $\mathscr{H}=K \mathbf{E}$, etc., the Minkowskian relation (2), involving the velocity of the medium relative to the observin system.

Applying (4) to a pair of surfaces bounded by one and the sam circuit $s$, as on p. 25 , we obtain the familiar equation,

$$
\frac{d e}{d t}=-\int(\mathrm{\Psi n}) d \sigma,
$$

where $e$ is the total charge of any portion of the medium enclose completely by the surface $\sigma$, whose outward normal is n. If th bounding surface is entirely composed of lines of conduction-curren then the charge remains constant. The same result follows, course, from the first pair of the differential equations ( r ), wit $\mathrm{I}=\rho \nabla+\mathfrak{F}$. And since these are independent of the Minkowskia connexions, involving the substantial properties of the medium there is no wonder that the equation of continuity reappears in i familiar form.

The above equations (4) and (5) lead at once to a pair of what ar usually called the boundary conditions. The other pair follow directly from $\operatorname{div} \mathbb{C}=\rho$ and $\operatorname{div} \mathfrak{G}\{t=0$. In fact, let $\Sigma$ be, in Had mard's phraseology, a stationary surface of discontinuity,* $i$. permanently affecting the same material particles, such as the surfac of contact of two different media. And let us require that $£$ and th individual time-rate of change of $\mathbb{E}$ and $\mathfrak{f l t}$ should be finite. Th condition, to be fulfilled at any point of $\Sigma$ and elsewhere, is necessar to prevent $\mathscr{E}$, $\mathfrak{A t}$ mounting up to infinite values at any point of th medium. $\dagger$ Under these assumptions apply (4) and (5), in the usu way, to an infinitesimal rectangle, with its shorter sides normal to Then the result will be that the tangential components of $\mathbf{E}^{\times}$and $\mathbf{M}$ must be continuous. The two remaining conditions are as in th older theory. They follow at once from the divergence-formula and require the normal component of $\mathfrak{y l}$ to be continuous, and th

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 THE THEORY OF RELATIVITYjump of the normal component of $\mathbb{E}$ to be equal to the density of charge. Thus, if there is no such charge, we h following boundary conditions:
( $\mathfrak{G l t n}$ ) and ( $\mathbb{E} n$ n) continuous,

where $\mathbf{n}$ is normal to the boundary. The latter pair of exp gives the tangential parts of the vectors, i.e. in both $s$ direction.

Next, as regards the formula (3) for the density of charge, a consequence of the nature of $C$ as a physical quaternion. first, that there is no conductivity. Then

$$
\rho=\gamma \rho^{\prime},
$$

just as for the microscopic density of charge, whence, for any of the body,

$$
e=\int \rho d S=\int \rho^{\prime} d S^{\prime}=e^{\prime}
$$

which means relativistic invariance of macroscopic charge property then continues to hold for a moving body, provided is a perfect insulator.

On the other hand, suppose that the body is conductive, there is no rest-charge $\left(\rho^{\prime}=0\right)$. Then there will be for the $S$-o an apparent charge of density

$$
\rho=\frac{\gamma^{2}}{2^{2}}(\Psi \nabla) .
$$

The history of this conduction charge, or compensation ch it previously has been called, can be traced back as far in which year it was deduced by Budde (Wied. Ann., p. 553) from Clausius' fundamental law of electrodynamics. whose formula differed from the above one by containi instead of $\gamma^{2}$, was able to defend Clausius' law from a seriou by showing that this charge accounted for the non-existen action between a current circuit and a charged body sh
two ways leading to one and the same result will be found usefu and the electronic interpretation of a formula which here appea as a relativistic consequence of Maxwell's equations will not $b$ lacking in interest. But even apart from electro-atomistic col cepts the reader will not fail to see that if the densities of positi electricity, flowing one way, and negative flowing the other wa cancel one another for an observer attached to the conductir body, then the corresponding values $\rho_{+}$and $\rho_{-}$, as estimated fro any other ( $S$-) point of view, will in general not annul themselve They will do so only when the current has no longitudinal con ponent. There is no difficulty in working out the quantitati details of such a reasoning, and thus re-obtaining the above formul

Next, as regards the dragging of waves. We know already fro Chapter VI. that, whatever the value $\mathfrak{b}^{\prime}$ of the velocity of propagatic in the rest-system, its $S$-value $\mathfrak{b}$ will follow by the addition theore of velocities, and will give, therefore, the Fresnelian coefficier And that Einstein's theorem is in fact applicable to the prese case, can be concluded from the manner in which the equatio ( 1 ), (2) have been obtained from those, ( $\left.\mathrm{I}^{\prime}\right)$, $\left(2^{\prime}\right)$, holding in Thus we know beforehand that Minkowski's equations will le to the correct Fresnelian value of the dragging coefficient. Ar this expectation is readily confirmed on performing the expli calculation. Cf. Note 2.

Finally, let us remark that Minkowski's electromagnetic equatio account fully for the well-known results of Rowland's, Wilson Röntgen's and Eichenwald's experiments. We cannot enter he upon the corresponding details, and must confine ourselves short indications concerning each of these famous experimen The magnetic effect of the convection current, first proved expe mentally by Rowland, and confirmed by other physicists,* directly expressed by the term $\rho \nabla$, which together with $t$ conduction current makes up I, and thus equally with that curre contributes to the magnetic field. It is scarcely necessary to $s$ that the Rowland effect was equally well expressed by the Her Heaviside equations. The result of Wilson's experiments on $t$

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## THE THEORY OF RELATIVITY

electric effect of rotating a dielectric between the connected of a condenser in a magnetic field $M$ consisted in each plates being found charged to a surface-density

$$
\begin{equation*}
\left(K^{\prime}-\mathrm{r}\right) \beta M \tag{II}
\end{equation*}
$$

of opposite signs.* In the theoretical treatment of the pr uniform translation (of each element) can, with sufficient acc be substituted for the actual spin, and the state being sup stationary (and $\sigma=0, \rho=0$ ), Minkowski's differential equations r to $\operatorname{curl} \mathbf{E}=0$, etc. Using these, with the appropriate bour conditions, and the first pair of (2), Einstein and Laub $\dagger$ de for the surface-density in question, the value

$$
(K \mu-\mathrm{r}) \beta M,
$$

with the correct sign for each plate. The authors observe Lorentz's theory would give, instead of this,

$$
(K-1) \beta \mu M .
$$

Since in Wilson's case $\mu$ was $=\mathrm{r}$, both of these theoretical for coincide with his experimental result. If a dielectric of consid inductivity were available, experiment would readily decide in of the former or the latter theory. As to Hertr-Heaviside's tl it would give for the Wilson-effect

$$
K \mu \beta M,
$$

i.e. practically $K \beta M$, which is equally contradicted by Wilson' by Blondlot's results. This disagreement, even in the case first-order effect, might have been expected, in view of the fac Hertz-Heaviside's equations give a full instead of a Fresnelian Lastly, as regards the experiments on the magnetic effect of $m$ polarized dielectrics, which were first carried out by Rö and more recently with increased accuracy by Eichenwald, * be enough to write down the expression of what is generally

* H. A. Wilson, Phil. T'rans., Vol. CCIV. A, r904, p. r2r. Wilson's p result agrees with the absence of any such effect stated previously by R. Bli Comptes rendus, Vol. CNXXIII. 1901, D. 778 , in the case of $a \dot{r}$ as die
the Röntgen-current. In fact, if we limit ourselves to homogeneous media, the experimental results may be concisely stated by saying that the observed value of the Röntgen-current is

$$
\begin{equation*}
(K-\mathrm{r}) \mathrm{curl} \text { VEv. } \tag{Exper.}
\end{equation*}
$$

Now, according to the Hertz-Heaviside equations (p. 3r), this current would be

$$
K^{\prime} . \operatorname{curl} \text { VEv, }
$$

(HH)
so that the disagreement is exactly of the same kind as for the Wilson-effect. On the other hand, Minkowski's equations, with $\mu=\mathrm{r}$, give for the Röntgen-current the rigorous value

$$
\begin{equation*}
\operatorname{curl} \mathrm{V}[\mathfrak{E}-\mathbf{E}] \mathbf{v} \tag{Mnk}
\end{equation*}
$$

where, by the first of (2) and by (A),

$$
\mathfrak{E}-\mathbf{E}=(K-\mathrm{I}) \mathbf{E}+{ }_{c}^{\mathrm{I}} \mathrm{VV}\left[K_{\mathfrak{G}} \mathfrak{f l}-\mathbf{M}\right] .
$$

Thus the first-order term of the Minkowskian expression represents correctly the observed facts. The second-order terms are, of course, for the time being far too small to be detected. The Minkowskian value of the Röntgen-current follows also from a later form of Lorentz's equations deduced (1902) from the electron theory.* In what consists the violation done by these last equations to the principle of relativity may be seen from Minkowski's paper. There the reader will find also the appropriate coordination of the fieldvectors involved in the various theories.

So much as regards the electromagnetic equations for moving bodies, contained in (1) and (2). Now for the dynamical part of the subject. Before proceeding to a relativistic construction of the formulae for the ponderomotive force and the associated physical magnitudes, some preliminary remarks seem indispensable. These will concern the requirements to be postulated in addition to those dictated by the principle of relativity itself. The choice of such supplementary requirements or postulates is free, within fairly wide limits. We shall select those which seem to offer the advantage of possible simplicity and which will lead to results but slightly different

Let $\mathbf{P}$ be the ponderomotive force due to the electromagnetic fie per unit volume of the medium, and, therefore, ( Pv ) its activi Further, let $J$ be Joule's heat, or the Joulean waste, per unit time al unit volume, and $F$ the force-quaternion, i.e., according to what b been said in the last chapter,

$$
F=\frac{\mathrm{r}}{c}[(\mathrm{Pv})+J]+\mathrm{P} .
$$

Let $u$ be the density of electromagnetic energy, $\mathbf{g}$ that of elect magnetic momentum, and finally $f$ and $\ddagger$ the ('absolute,' $n$ relative) stress-operator and flux of energy, as defined in the ust way with respect to the observing system $S$. With this meaning the symbols, let our requirements be as follows:
I. F, a physical quaternion,

$$
F \equiv_{i}^{b}[(\mathrm{Pv})+/]+\mathrm{P} \simeq q
$$

$2^{\circ}$. Principle of momentum, to call it by its usual short name, th is to say,

$$
\mathbf{P}=-\nabla f-\frac{\partial \mathbf{g}}{\partial t},
$$

where $\nabla f$ stands for $\partial \mathrm{f}_{1} / \partial x+\partial \mathrm{f}_{2} / / y+\partial \mathrm{f}_{3} / \partial z$.
$3^{\circ}$. Principle of conservation of enersy, i.e.

$$
(\mathbf{P v})+J=-\frac{\partial u}{\partial t}-\operatorname{div} \| J,
$$

where 扔 has, thus fur, nothing to do zeith the momentum.
It is needless to add that, besides fulfilling these explicit requi ments, the resulting formulac have to agree with experience, far as it goes, and to reduce, for $K^{*}=\mu=1, r=0$, to the previo vacuum-formulae, as, in fact, they will.

We have seen in the preceding chapter that there is at the prese time a strong tendency to universalize the simple relation of equal holding between $g$ and $\mathfrak{j} / c^{a}$ in the ideal limiting case of a vacuum.

[^55]But, as far as I can see, there is nothing to compel us to such a generalization. If it is assumed that the matrix embodying the stress, momentum, etc., should be symmetrical, then, of course, the equality under consideration follows from $(\beta)$ and $(\gamma)$. But nothing prevents us from abandoning, at least in the case of ponderable media, that assumption of symmetry.* We shall see that in doing so we need not even give up the formulae (14) or (14a) of Chap. IX., which have led to so many far-reaching consequences. These formulae will continue to hold within wide limits, although the more general formula (ro) of that chapter will have to be modified. Thus, there will still be 'inertia of energy,' with its manifold corollaries.

So much to justify the abandoning of the assumption of universal proportionality of momentum and energy-flux.

Returning to our above requirements, let us, first of all, observe that, with the given meaning of $F$, assumptions $(\beta)$ and $(\gamma)$ may be condensed into

$$
\begin{equation*}
F=-\operatorname{lor} \mathbb{S} \tag{ıо}
\end{equation*}
$$

where

$$
S=\left|\begin{array}{cc}
f, & \frac{\iota}{c} \nmid \boldsymbol{j u}  \tag{Ir}\\
\iota g, & -u
\end{array}\right|
$$

or, written out fully,

$$
S=\left|\begin{array}{cccc}
f_{11}, & f_{12}, & f_{13}, & \frac{\iota}{c}-\exists_{1}  \tag{IIa}\\
f_{21}, & f_{22}, & f_{23}, & \frac{\iota}{c}-1_{2} \\
f_{31}, & f_{32}, & f_{33}, & \frac{\iota}{c}-1_{3} \\
\iota c_{3}, & \iota g_{2}, & \iota c_{3}, & -u
\end{array}\right| .
$$

* In Sommerfeld's four-dimensional algebra (loc. cit.), the symmetrical worldtensor, corresponding to such a matrix, is generated by what he calls 'a


## THE THEORY OF RELATIVITY

Here, in general, $f \iota \neq f \kappa \iota$, so that the matrix lacks symmet altogether.

Next, to satisfy $(\alpha)$, we have to write, for any pair of legitima frameworks of reference $S$ and $S^{\prime}$, as on p. 239,

$$
S=A S^{\prime} \cdot \bar{I}
$$

where $A, A$ are as before. This fixes the transformational properti of the stress, momentum, etc., quite inclependently of the elect magnetic expressions they will hereafter receive. Developing ( $x$ we have the following table of Cartesian formulae, which ta the place of (roa), p. 237, and which, though not needed our electrodynamical investigation, are here given because of the bearing upon the subjects treated in the preceding chapter:

$$
\begin{aligned}
& f_{11}=\gamma^{2}\left[f_{11}{ }^{\prime}+\beta^{2} u^{\prime}+\frac{\beta}{C}\left(\mathfrak{m p}_{1}{ }^{\prime}+\alpha^{2} \xi_{1}{ }^{\prime}\right)\right] ; f_{222}=f_{32}{ }^{\prime} ; f_{33}=f_{33}{ }^{\prime} \\
& f_{23}=f_{23}{ }^{\prime} ; f_{31}=\gamma\left(f_{31}{ }^{\prime}+\frac{\beta}{6} \text { 羽 }{ }_{3}{ }^{\prime}\right) ; f_{12}=\gamma\left(f_{12}{ }^{\prime}+\tau^{\prime} g_{2}{ }^{\prime}\right) \\
& f_{32}=f_{32}{ }^{\prime} ; f_{13}=\gamma\left(f_{13}{ }^{\prime}+\tau g_{3}{ }^{\prime}\right) ; \quad f_{21}=\gamma\left(f_{21}{ }^{\prime}+\frac{\beta}{c}{ }_{c}{ }^{\prime}{ }^{\prime}\right)
\end{aligned}
$$

$$
\begin{aligned}
& u=\gamma^{2}\left[u^{\prime}+\beta^{2} f_{11}^{\prime}+{ }_{c}^{\beta}\left(\eta_{1}^{\prime}+c^{2} s_{1} r^{\prime}\right)\right] .
\end{aligned}
$$

Here the $x$-axis is taken along $\mathbf{v}$, the velocity of.$\delta^{\prime \prime}$ relative to (The reader can condense these formulat into a more convenic shape by using vectors and the stretcher c.) If there is, from $S^{\prime}$-point of view, no flux of encrgy and no momentum, then $u$ a the stress-components become as in ( $14 a$ ) of Chap. IX.; obtain also the same $S$-momentum as before, i.e.

$$
g_{1}=\frac{\gamma^{2} v}{c^{2}}\left(u^{\prime}+f_{11}^{\prime}\right), \quad g_{2}=\frac{\gamma^{v}}{c^{2}} f_{12}^{\prime}, \quad g_{3}=\frac{\gamma^{\prime} v^{\prime}}{c^{\prime}} f_{13}^{\prime}
$$

whereas

## PONDEROMOTIVE FORCE

seen to hold for any electromagnetic field, if $S^{\prime}$ is attached to th ponderable medium ; and the condition of vanishing $g^{\prime}$ and $\exists^{\prime}$ wi be satisfied in the case of a purely electrostatic, or a purel magnetostatic field.

With $f_{c k}{ }^{\prime}=f_{k l}^{\prime}$ alone, we have, from ( $12 a$ ), the interesting relatio

$$
\text { 羽 }-c^{2} \mathbf{g}=\frac{\gamma}{\epsilon}\left[\text { 羽 }-c^{2} \mathbf{g}^{\prime}\right] \text {, }
$$

which will hold for any electromagnetic field, provided that $S^{\prime}$ the rest-system of the ponderable medium.

But let us return to our chief subject. After what has bee said we could either employ the form (ro) of the force-quaternior and would then have to prove that $\mathbb{S}$ is transformed accordin to ( 12 ), or we can proceed by satisfying our three requirement in their original forms $(\beta),(\gamma)$, and $(\alpha)$. The two ways are wholl equivalent to one another. Minkowski chooses the former: $h$ constructs $\$$ in a manner that ensures by itself the validity (12), subjects it to the operation lor, and develops the resultin four-vector.* We shall take the latter way, which the reader ma find easier to follow. Thus, we shall first construct $F$ so that should be a physical quaternion, and then find the correspondin expressions for the energy, stress, etc., according to $(\beta),(\gamma)$, aided of course, by the electromagnetic equations (r), (2).

The first step to be taken is suggested by analogy with th construction of the fundamental electronic force-expression (c p. 220). We know that

$$
C=\iota \rho+\frac{\mathbf{I} .}{c} \simeq q,
$$

and that $\mathbf{L}=\mathbf{f t l}-\mathfrak{E}$ is a left-handed bivector. Therefore, $C \mathbf{L} \sim$ Similarly, $\mathbf{R}=\mathfrak{y t}+c \mathbf{E}$ being a right-handed bivector, we hav $\mathbf{R} C \sim q$. The difference of both products has also the structu of $q$, and thus is again a physical quaternion, and can be used far as $(x)$ is concerned. Try, therefore, to satisfy the remainir requirements of the problem by putting
be easily supplemented by another physical quaternion involving the variations of $K, \mu$. Develop the right-hand side of (r4a). Then the vector part will give the ponderomotive force,

$$
\begin{equation*}
\mathbf{P}=\rho \mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{VI} \mathfrak{A l}, \tag{}
\end{equation*}
$$

and the scalar part will lead to

$$
\begin{equation*}
(\mathbf{P v})+J=(\mathbf{E I}) . \tag{I6a}
\end{equation*}
$$

Eliminate (EI) from these two equations and remember that $\mathbf{I}=\rho \mathbf{\nabla}+£$. Then the result will be
giving for Joule's heat the expression

$$
\begin{equation*}
J=\left(\mathrm{EE}^{\times}\right) . \tag{17}
\end{equation*}
$$

Thus far $(\beta)$ and $(\gamma)$ have not yet been employed. Now take account of these conditions, beginning with the latter. This gives, by ( $16 a$ ),

$$
-(\mathbf{E I})=\frac{\partial u}{\partial t}+\operatorname{div} \text { 解. }
$$

Now, by the electromagnetic differential equations ( r ),

$$
-(\mathbf{E I})=\frac{\mathrm{I}}{2} \frac{\partial}{\partial t}(\mathbf{E} \mathbb{E}+\mathbf{M} \mathfrak{G} \mathfrak{f l})+c \cdot \operatorname{div} \text { VEM. }
$$

Thus ( $\gamma$ ), the principle of conservation of energy, is satisfied if the density of electromagnetic energy is taken to be

$$
\begin{equation*}
u=\frac{1}{2}\left(E \mathbb{E}+\mathrm{M}_{\mathfrak{C}} \mathfrak{l l}\right), \tag{18}
\end{equation*}
$$

and the flux of energy, from the standpoint of the observing system,*

$$
\begin{equation*}
\text { 羽 = } c \mathrm{VEM} \text {. } \tag{19}
\end{equation*}
$$

The addition of an arbitrary sourceless (solenoidal) flux, as well as of an invariable $u$-term, would be irrelevant.

Lastly, to represent the ponderomotive force in the form required

Using the second pair of equations ( r ), and writing, for the moment,

$$
\mathbf{A}=\mathbf{E} \operatorname{div} \mathfrak{E}-\mathrm{V} \mathbb{E} \operatorname{curl} \mathbf{E}+\mathbf{M} \operatorname{div} \mathfrak{f l}-\mathrm{V} \mathfrak{C} \mathfrak{f l} \operatorname{curl} \mathbf{M} \text {, }
$$

we have

$$
\mathbf{P}=\mathbf{A}-\frac{\mathbf{I}}{c} \frac{\partial}{\partial t} \mathrm{~V} \mathbb{E} \mathbb{G} \mathbb{A l} .
$$

This gives, first of all, for the electromagnetic momentum per unit volume,

$$
\begin{equation*}
\mathrm{g}=\frac{\mathrm{I}}{c} \mathrm{~V} \mathbb{E} \mathbb{E} \mathbb{E} \mathbb{C l}, \tag{20}
\end{equation*}
$$

and what remains to be shown is that the vector sum $\mathbf{A}$, familiar from the Maxwellian theory, is of the form $-\nabla f$. Now, this is exactly the case, provided that $K$ and $\mu$, involved in (2), are constant throughout the medium. In fact, take for the electromagnetic stress the familiar expression

$$
\begin{equation*}
\mathbf{f}_{n}=u \mathbf{n}-\mathbf{E}(\mathbb{E} \mathrm{n})-\mathbf{M}(\mathfrak{f l n}), \tag{2I}
\end{equation*}
$$

where $u$ is as in (18). Then, remembering that $\nabla f$ is used as shorthand for $\partial \mathbf{f}_{1} / \partial x+\partial \mathrm{f}_{2} / \partial y+\partial \mathbf{f}_{3} / \partial z$,

$$
\begin{aligned}
-\nabla f & =-\nabla u+\frac{\partial}{\partial x}[(\mathbb{F} \mathbf{i}) \mathbf{E}+(\mathfrak{A l i}) \mathbf{M}]+\frac{\partial}{\partial y} \cdots+\frac{\partial}{\partial z} \cdots \\
& =\mathbf{E} \operatorname{div} \mathfrak{E}+\mathbf{M} \operatorname{div} \mathfrak{f l}-\nabla u+(\mathfrak{E} . \nabla) \mathbf{E}+(\mathfrak{G l l} . \nabla) \mathbf{M} .
\end{aligned}
$$

On the other hand, we have

$$
\mathrm{V} \mathbb{E} \operatorname{curl} \mathbf{E}=\mathrm{V} \mathbb{E} \cdot \mathrm{~V} \nabla \mathbf{E}=\nabla(\mathbf{E} \cdot \mathbb{E})-(\mathbb{E} . \nabla) \mathbf{E}
$$

(where the dot stops $\nabla$ 's differentiating action), and a similar expression for the last term of $\mathbf{A}$. Thus,

$$
-\nabla f=\mathbf{A}+\mathbf{N},
$$

where

$$
\mathbf{N}=\nabla(\mathbf{E} \cdot \mathfrak{E}+\mathbf{M} \cdot \mathfrak{f l})-\nabla u,
$$

i.e.

$$
\mathbf{N}=\frac{1}{2} \nabla(\mathbf{E} \cdot \mathfrak{E}+\mathbf{M} \cdot \mathfrak{y l}-\mathbb{E} \cdot \mathbf{E}-\mathfrak{f l} . \mathbf{M}) .
$$

To prevent a possible misunderstanding, we may add that this is a vector whose components are
paper,* the reader will verify that the above vector is identi

$$
\mathbf{N}=-\frac{1}{2}(\mathrm{~T} \eta)^{2} \cdot \nabla K-\frac{1}{2}(\mathrm{~T} \zeta)^{2} \cdot \nabla \mu,
$$

where the quaternions $\eta$ and $\zeta$ are as in (B), p. 265. In the homogeneity, therefore, $\mathbf{N}$ vanishes, and we have $\mathbf{A}=-\nabla f$, the condition $(\beta)$ is satisfied, with the above stress and mon by taking ( $\mathrm{r} 5 a$ ) for the ponderomotive force, that is ( $14 a$ ) force-quaternion.

In the more general case of a heteroseneous medium we ha to supplement our original $\mathbf{P}$ by the vector $\mathbf{N}$, and consequ add to our original $F$ the quaternion

$$
-\frac{1}{2}(\mathrm{~T} \eta)^{2} \cdot D K-\frac{1}{2}(\mathrm{~T} \zeta)^{2} \cdot D \mu
$$

which, like that $F$ itself, is $\simeq q$, since $\mathrm{T} \eta$ and $\mathrm{T} \zeta$, being the of physical quaternions, are invariant with respect to the transformation.

Thus we shall have, as a generalization of ( $14 a$ ),

$$
F=\frac{1}{2}[C \mathrm{~L}-\mathrm{R} C]-\frac{1}{2}(\mathrm{~T} \eta)^{2} \cdot D K-\frac{1}{2}(\mathrm{~T} \zeta)^{2} \cdot D \mu,
$$

which splits into

$$
\mathrm{P}=\rho \mathbf{E}+\frac{\mathrm{I}}{c} \mathrm{VI} \begin{cases}\mathrm{tl} \\ -\frac{1}{2} & (\mathrm{~T} \eta)^{2} \cdot \nabla K-\frac{1}{2}(\mathrm{~T} \zeta)^{2} \cdot \nabla \mu\end{cases}
$$

and

$$
(\mathrm{Pv})+J=(\mathbf{E I})+\frac{1}{2}(\mathrm{~T} \eta)^{2} \frac{\partial K}{\partial t}+\frac{1}{2}(\mathrm{~T} \zeta)^{2} \frac{\partial \mu}{\partial t} .
$$

All requirements being now satisfied, with the above va density and flux of energy, and of stress and momentum, th thing to be still revised on account of the heterogeneity medium is the Joulean waste. Now, proceeding as beff obtain at once, from (15) and (16),

$$
J=\left(\mathfrak{\Psi E}^{\times}\right)+\frac{1}{2}(\mathrm{~T} \eta)^{2} \frac{d K}{d t}+\frac{1}{2}(\mathrm{~T} \zeta)^{2} \frac{d \mu}{d t},
$$

where

$$
\frac{d K}{d t}=\frac{\partial K}{\partial t}+(\nabla \nabla) K=\frac{\mathrm{r}}{\gamma} \frac{\partial K}{\partial t^{\prime}},
$$

the previous value, ( I 7 ), of Joule's heat. Under these circumstances we have also

$$
\nabla K=\epsilon \nabla^{\prime} K, \quad \nabla \mu=\epsilon \nabla^{\prime} \mu,
$$

which values can be substituted in the last two terms of ( 15 ). To resume:

The force-quaternion

$$
\begin{equation*}
F=\frac{1}{2}[C \mathrm{~L}-\mathrm{R} C]-\frac{1}{2}(\mathrm{~T} \eta)^{2} \cdot D K-\frac{1}{2}(\mathrm{~T} \zeta)^{2} \cdot D \mu, \tag{I4}
\end{equation*}
$$

being a physical quaternion, satisfies the fundamental relativistic recuirements and, at the same time, the so-called principle of mumentum,

$$
P=-\nabla f-\partial g / \partial t,
$$

and the principle of conservation of energy,

$$
(\mathrm{Pv})+J+\frac{\partial u}{\partial t}+\operatorname{div} \nexists \mathrm{p}=0 .
$$

It gives for the ponderomotive force, per unit volume of an isotropic medium,

$$
\begin{equation*}
\mathbf{P}=\rho \mathbf{E}+{ }_{c}^{\mathrm{r}} \mathrm{VI} \mathrm{Ift}-\frac{1}{2}(T Y)^{2} \cdot \nabla K-\frac{1}{2}(T \zeta)^{2} \cdot \nabla \mu, \tag{15}
\end{equation*}
$$

and for the foulectn zeraste (with $\partial K / \partial t^{\prime}=\partial \mu / \partial t^{\prime}=0$ ):

$$
\begin{equation*}
J=\left(\mathbf{玉 E}^{\times}\right), \tag{17}
\end{equation*}
$$

where $\mathbf{X}=\mathbf{I}-\boldsymbol{\rho}$ is the conduction current. The corresponding auxiliary magnitudes are as follows:

The density of electromagnetic energy

$$
\begin{equation*}
u=\frac{1}{2}(\mathbf{E} \mathbb{E}+\mathbf{M} \mathfrak{M t}), \tag{I8}
\end{equation*}
$$

the flux of energy

$$
\begin{equation*}
\eta=c \mathrm{VEM}, \tag{I9}
\end{equation*}
$$

the density of electromagnetic momentum

$$
\begin{equation*}
\mathrm{g}={ }_{c}^{\mathrm{I}} \mathrm{~V} \mathbb{E} \mathrm{ft}, \tag{20}
\end{equation*}
$$

and, finally, the electromagnetic stress

Our $F$ is the quaternionic equivalent of Minkowski's four-ve (loc. cit., § I4, for constant $\nabla$, of course), and so also our components, etc., are identical with the sixteen constitue Minkowski's matrix $-S$. The difference between the theor proposed and that given by Minkowski, is this :-What Min considers as the ponderomotive force is the vector, not of but, of

$$
F_{\mathrm{Mnk}}=\frac{1}{2}\left[F+\frac{\mathrm{r}}{c^{2}} Y F_{c} Y\right],
$$

i.e. of that part of the four-vector $F$, which is normal to th velocity $Y$. Thus Minkowski's ponderomotive force is the form $(\beta)$, though it becomes so in the rest-system. The why Minkowski proposed for the four-force the above $F$, instead of the whole $F$, is to be sought for in his dyr according to which the 'moving' four-force had always normal to the particle's world-line. This corresponded assumption of a constant rest-mass. But in general, as ha explained in Chapter IX., the four-force does not necessari that relation to the world-line, and it will not do so wh there is heat supply or heat generation. Now, this being the case with an electric conductor, we had to abando Minkowskian condition of orthogonality. And, in connexio this, Joule's heat has, from the beginning, been embodied in force-quaternion along with the activity of the force, such cedure being directly suggested by the dynamical conside of Chapter IX. If there is no conductivity, and therefore a Joulean waste, then the four-force represented by (14) is, i normal to the world-line of the body, i.e.

$$
S Y_{c} F=0, \quad \text { for } \quad \sigma=0
$$

as is proved in Note 4 at the end of the chapter. But in ducting body this property does not hold.

We will content ourselves here with the above modificat Minkowskian electrodynamics, which, besides fulfilling the requirements, is also in complete agreement with what is

## PONDEROMOTIVE FORCE, ETC.

convection-current, contained in VIftl/c). But, since nobody ever observed such an action of the magnetic field, this can $h$ be considered as a serious objection. In order to obtain desired term, Einstein and Laub recurred to the electron th and since in doing so they thought necessary to limit thems to the case of stationary bodies, their electrodynamics is c particular interest from the relativistic point of view. Abral ponderomotive force contains, in addition to (15), several terms, and among these the displacement-current term. His el dynamics of moving bodies,* as has already been mentione based upon the assumption of the particular relation $\mathbf{g}=$ borrowed from the vacuum-equations. The desire to retain relation throughout the theory makes Abraham's formulae siderably more complicated than the above ones. His expre for the Joulean waste, subject to the same conditions for is the same as above, but those for the density and flux of en and consequently also for the momentum, are less simple.

The set of electrodynamical formulae given above is best ch terized by saying that, while satisfying the principle of rela and the principle of conservation of energy, it gives for the system the ponderomotive force

$$
\mathbf{P}^{\prime}=-\nabla^{\prime} f_{\mathrm{Mxw}}^{\prime}-\frac{\mathrm{I}}{c} \frac{\partial}{\partial t^{\prime}} V \mathcal{E}^{\prime} \mathfrak{g l t}^{\prime} .
$$

In fact, $f$, as given by (2I), reduces in $S^{\prime}$ to the Maxwellian s operator,

$$
f^{\prime}=f_{\mathrm{Mxw}}^{\prime}=\frac{1}{2}\left(K E^{\prime 2}+\mu M M^{\prime 2}\right)-K \mathbf{E}^{\prime}\left(\mathbf{E}^{\prime}-\mu \mathbf{M}^{\prime}\left(\mathbf{M}^{\prime} .\right.\right.
$$

It will be remembered that Maxwell's stress, taken by itself, give, in absence of electric charge and of ponderable matter ponderomotive force

$$
\frac{\mathrm{r}}{c} \frac{\partial}{\partial t^{\prime}} \mathrm{VE}^{\prime} \mathbf{M}^{\prime},
$$

and this 'force on the free aether' is just balanced by the se term in (23). See p. 48. With the exception of this obvi desirable supplementary term everything is as in Maxwell's el dynamics of stationary bodies. Thus, (15), the developed
where $\mathfrak{F}^{\prime}=I^{\prime}$ is the conduction current, and each of the four has its old familiar form and meaning. Again, by ( $\mathrm{r}_{7}$ ) ar the Joulean waste takes its usual form,

$$
J^{\prime}=\left(\mathrm{I}^{\prime} \mathbf{E}^{\prime}\right)=\sigma E^{\prime 2}
$$

while ( x 8 ) and (19) give at once the Maxwellian density of magnetic energy, and the familiar Poynting vector for th of energy,

$$
u^{\prime}=\frac{1}{2}\left(K E^{\prime 2}+\mu M^{\prime 2}\right) ; \quad \text { 㡖 }=c \mathrm{VE}^{\prime} \mathbf{M}^{\prime} .
$$

The transformation formula for Joule's heat is easily ob In fact, since $F$, defined by ( 9 ), is a physical quaternion, and

$$
F^{\prime}=\frac{\iota}{c} J^{\prime}+\mathbf{P}^{\prime}
$$

we have at once, writing $P_{1}$ for the longitudinal compon the force,

$$
J+v P_{1}=\gamma\left(J^{\prime}+v P_{1}^{\prime}\right)
$$

and

$$
P_{1}=\gamma\left(P_{1}^{\prime}+v J^{\prime} c^{2}\right),
$$

whence, by subtraction,

$$
J=\gamma\left(\mathrm{r}-\beta^{2}\right) J^{\prime}=\frac{\mathrm{I}}{\gamma} J^{\prime}
$$

Consequently, if $d S^{\prime}$ be any volume-element of the bod $d S$ its correspondent,

$$
J d S=\frac{\mathrm{I}}{\gamma^{2}} J^{\prime} d S^{\prime}
$$

in complete agreement with (29), Chap. IX.
The electromagnetic momentum bears, in the rest-sys simple relation to the energy flux. In fact, by (20),

$$
\mathbf{g}^{\prime}=\frac{K \mu}{c} \mathrm{VE}^{\prime} \mathbf{M}^{\prime}=\frac{c}{\mathfrak{b}^{\prime 2}} \mathrm{VE}^{\prime} \mathbf{M}^{\prime}
$$

where, dispersion being disregarded, $\mathfrak{b}^{\prime}$ is the velocity of prop of disturbances, as estimated by the $S^{\prime}$-observers. Hence, of Planck's relation, we have

$$
\mathbf{g}^{\prime}=\frac{\mathrm{I}}{\mathfrak{v}^{\prime 2}} \mathfrak{z}
$$

so that, in a stationary ponderable medium, $\mathfrak{b}^{\prime}$ takes the pla And since $\mathfrak{b}^{\prime}$ plays in such a medium just the same part

## ENERGY FLUX AND MOMENTUM

critical velocity in empty space, it seems quite natural that should replace the relation which holds good in the absenc matter. The stress in.$S^{\prime}$ being self-conjugate, our previous equ (r3) can be applied, so that, in general,

$$
\text { 羽 }-c^{2} \mathbf{g}=\left(\mathrm{r}-n^{\prime 2}\right) \gamma \epsilon^{-1} \boldsymbol{q}^{\prime} \text {, }
$$

where $n^{\prime}$ is the refractive index of the medium. If, therefor differs at all from unity, we have $\nexists \neq c^{2} \mathrm{~g}$, unless there is in rest-system no Poynting flux.

Finally, notice that if $K=\mu=\mathrm{r}$, and $\sigma=0$, the ponderom force ( $\mathrm{r}_{5}$ ) coincides with that of the electron theory. And same thing is true of the above expressions for stress, density flux of energy, and momentum. So also were the vacuum-equa contained, as a limiting case, in Minkowski's electromagnetic $\dot{d}$ ential equations for moving bodies.

NOTES TO CHAPTER X.
Note 1 (to page 266). Let the quaternions $a$ and $b$ represent a p four-vectors. Then the component of the four-vector $a$ taken the four-vector $b$ (cf. p. 148) will be represented by

$$
(\mathrm{T} b)^{-1} \cdot \mathrm{~S} a_{c} b
$$

and, consequently, the part of $a$ normal to $b$, in both size and dire by

$$
a_{n}=a-\frac{b \mathrm{~S} a_{c} b}{(\mathrm{~T} b)^{2}} .
$$

Now, $\mathrm{S} a_{c} b=\frac{1}{2}\left[b_{c} a+a_{c} b\right]$, and $b b_{c}=(\mathrm{T} b)^{2}$. Hence

$$
a_{n}=\frac{1}{2}\left[a-\frac{b a_{c} b}{(\mathrm{~T} b)^{2}}\right],
$$

which is the required expression.
Note 2 (to page 273). It will be enough to consider here the ca plane waves, propagated along $\mathbf{v}$, in a non-conducting medium, car no charge, so that $\mathrm{I}=0$.

As in a previous Note (p. 59), take $\mathbf{E}$, $\mathbb{E}$, etc., proportional
discontinuity). Then the equations ( 1 ) and (2) will give, by (A) and since $\mathbf{V}=c \beta \mathbf{i}$,

$$
\begin{aligned}
& -\frac{\mathfrak{v}}{c} \mathscr{E}=\operatorname{ViM}, \frac{\mathfrak{b}}{c} \mathfrak{A f}=\mathrm{ViE}, \\
& \mathfrak{E}+\beta \operatorname{ViM}=K[\mathbf{E}+\beta \mathrm{Vi} \mathfrak{A H}], \\
& \mathfrak{d i t}-\beta \mathrm{ViE}=\mu[\mathbf{M}-\beta \mathrm{Vi} \mathbb{E}],
\end{aligned}
$$

the solenoidal conditions $(\mathbb{E} i)=(\mathfrak{f t i})=0$ being already satisfied Next, introduce (a) into (b), (c); then

$$
\begin{aligned}
& \mathfrak{E}\left[\mathrm{I}-\beta_{\frac{\mathfrak{V}}{\mathrm{V}}}^{\mathrm{C}}\right]=K^{\prime}\left[\mathrm{I}-\beta_{\overline{\mathfrak{b}}}^{c}\right] \mathbf{E}, \\
& \mathfrak{A t t}\left[\mathrm{I}-\beta \frac{\mathfrak{v}}{\mathfrak{c}}\right]=\mu\left[\mathrm{I}-\beta_{\mathrm{b}}^{\mathrm{c}}\right] \mathbf{M},
\end{aligned}
$$

showing that $\mathbf{E}$ and $\mathbf{M}$ are again transversal. Use the latter in (a), eliminate either $\mathbf{E}$ or $\mathbf{M}$, and remember that

Then the result will be

$$
\mathfrak{v}^{\prime}=c / \sqrt{K \mu} .
$$

$$
\frac{\mathfrak{b}-v}{1-v \mathfrak{b} / c^{2}}=\mathfrak{b}^{\prime},
$$

whence

$$
\mathfrak{b}=\frac{\mathfrak{b}^{\prime}+v}{\mathfrak{I}+v \mathfrak{b}^{\prime} / \mathfrak{c}^{2}} .
$$

Thus $\mathfrak{b}$ is obtained from $\mathfrak{b}^{\prime}$ and $v$ by Einstein's addition the velocities, and this, as we saw on p. 172, gives the Fresnelia for the dragging coefficient.
Note 3 (to page 279). Let $h$ and $H$, as on p . 264, be the alt matrices equivalent to the electromagnetic bivectors ${ }^{2}$ and $L$ tively, i.e.
and

$$
\begin{array}{lll}
h_{23}=M_{1}, & h_{31}=M_{2}, & h_{12}=M_{3} \\
h_{14}=-\iota \mathbb{E}_{1}, & h_{24}=-\iota \mathbb{E}_{2}, & h_{34}=-\iota \mathbb{E}_{3}
\end{array}
$$

$$
\begin{array}{lll}
H_{23}=\mathfrak{A l}_{1}, & H_{31}=\mathfrak{A t}_{2}, & H_{12}=\mathfrak{H t}_{3} \\
H_{14}=-\iota E_{1}, & H_{24}=-\iota E_{2}, & H_{34}=-\iota E_{3} .
\end{array}
$$

Both of these matrices reduce, for $K=\mu=\mathrm{r}$, to the matrix $h$ o to Chap. IX.

Minkowski begins by constructing the product of $h$ into $H$. each of the factors is transformed by $\bar{A}() A$, the same will of their product, which will be a matrix of $4 \times 4$ constituents similarly as on p. 259 , the reader will find

## STRESS-ENERGY MATRIX

or what in the rest-system becomes 'the Lagrangian function.' (It m: be mentioned, for the sake of comparison with Minkowski's paper, th our $h, H, \lambda, \mathcal{S}, F$ are his $f, F, L,-S, \mathrm{~K}$ respectively.) Similar $h^{*}$ and $H^{*}$ being the dual matrices,

$$
-H^{*} h^{*}=-\mathfrak{S}+\lambda .
$$

By (a) and (c), $\lambda$ is an invariant, and $\mathfrak{S}$ is transformed by $\bar{A}()$ so that

$$
F=-\operatorname{lor} S
$$

is a genuine four-vector (or physical quaternion). The latter become by ( $a$ ), (c) and (b),

$$
F=\operatorname{lor} h \cdot H-\operatorname{lor} H^{*} \cdot h^{*}+\mathrm{N},
$$

where the dots act as separators and N is the four-vector written quaternionic form under (c), p. 282.
Next, using in (d) the differential equations of the field, lor $h=-$ lor $H^{*}=0$, Minkowski obtains

$$
F=s H+\mathrm{N},
$$

where $s$ is the current-matrix, represented in this chapter by tl quaternion C. Minkowski's ponderomotive four-force is the part of normal to the four-velocity. Our force-quaternion (14) is the quaternion equivalent of the whole four-vector (e).

Notice that (b) can be written, in terms of the electromagnetic bivector

$$
\lambda=-\frac{1}{2} S L T C,
$$

whence the invariance of $\lambda$ is seen immediately.
Note 4 (to page 284). For a non-conducting medium, the curre quaternion becomes

$$
C=\rho\left[\iota+\frac{\mathrm{I}}{c} \nabla\right]=\frac{\rho}{c \gamma} Y,
$$

and therefore, the force-quaternion (14),

$$
F=\frac{\rho}{\gamma} \eta-\frac{1}{2}(T \eta)^{2} \cdot D K-\frac{1}{2}(T \zeta)^{2} \cdot D \mu,
$$

where $2 c \eta=Y \mathbf{L}-\mathbf{R} Y$, as on p. 265. Hence,

$$
S Y_{c \eta}=0 .
$$

Again, as regards the second term of $F$,

$$
{ }_{\gamma}^{\mathrm{I}} \mathrm{~S} Y_{c} D K=\frac{\partial K}{\partial t}+(\nabla \nabla) K=\frac{d K}{d t}=0,
$$

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[^0]:    * A clear and beautiful statement of the fundamental imprortance of the Copernican idea is to be found in ['. I'ainleve's article 'Meranigue' in the coller" tive volume De la Methode dans les Siciences, edited hy Hmile lorel, (Tariv, F. Alcan, 19ro.)
    S.R.

[^1]:    * Which gave so much trouble to the late Sir G. II. Darwin and his brother int their attempts to measure directly the gravitational action of the meen, as deseribed in Sir G. H. Darwin's attractive popular hook, The Tides and Xindred 'henoment in the Solar System, London, 1898 (German edition by A. [ockels, enharyed : Teubner, Leipzig, 1911).
    $\dagger$ With the exception of those of the type of Foucault's pendulum experiments, performed with the special purpose 'of showing the earth's rotation.' In more recent times the pendulum could be successfully replaced by a gyroscope, as originally suggested, and tried, by Foucault himself.

[^2]:    *The earth as the centre of the universe, with the 'crystal spheres,' with the starw stuck to them, spinning round the rarth, still formed part of the teachings of the Ionian sehool of philoswphers founded hy Thales (lerrn atheut 640 b.c.). The first to suggest the rohation of the earth round its axis and its motion round the sun enems to have been l'ythagoras, one of "Thales' diseiples, thought it has heen later unjustly attributed to Ihilolaus, one of l'ythagroras' disciples (born about 450 日. ©.).

    - Although I do net chaim to give here anything like a history of astronomy, it may tre worth memioning that the I'ythagurems alreuly taught that the planets and comets were circling round the sun. But al any rate the I'tolemaean geocentrie mystem reigned universally from the second till the fifternth century, the only serious objection against its complexity having leeen raised in the thirteenth century by Aphonnt, X., king of Castile, the author of the astronomical 'Tables' asociated with his name (published during 1248-1252).

[^3]:    * We speak of it in the singular, instead of infinite plural, only for the sake of shortness. For, as is well known, if 2 , say the 'fixed' stars, be such a symem, then any other system $\Sigma^{\prime}$ having relatively to $\Sigma^{\prime}$ any notion of uniform (reetilinear) translation is equally good for all purposes.
    $\dagger$ This point is expressly insisted upon and successfully apphied (o didactic purposes in Professor A. M. Worthington's Dynamios of Rotalion, sixth edition, new impression 1910; Longmans, Green \& Co., London.
    \$See, for example, the author's Vectorial Mechurnics, Chap. LIL. ; Macmillan \& Co., London, 1913.

[^4]:    * Thus it is manifest that the science of mechanics does not describe the motion of bodies in its quantitative dependence upon 'time, flowing at a constant rate' (Newton), but literally gives only sets of simultaneous states of motion of the various bodies, the time-keeper itself being included. What is besides contained in these sets or successions is a non-quantitative element, namely, of what is vaguely called 'before' and 'after.'

[^5]:    * Matter and Motion, p. 20, first paragraph of Art. xix. ; see also p. 21, lines 7-I i. ''Small' only so as not to be obliged to consider the different temperatures of its various parts.
    $\ddagger$ Or practically so, at least.

[^6]:    * Observe that $n$ mechanical 'degrees of freedom' amount to $2 n$ degrees of freedom in the sense here adopted.
    $\dagger$ See Note 1 at the end of the chapter.

[^7]:    - Not necessarily larger in volume; for often we introduce new parameters by going deeper iuto the original system itself, sometimes as deep as the molecular, atomic or even sul-atomic structure, say, of a piece of matter; or being originally concerned with the thermic history only, we supplement the temperature by the pressure, volume, electric potential, and so on.

[^8]:    *Thus we read in Painlevés article (loc. cit. page 91): 'La durée d'une ondulation lumineuse correspondant à une radiation déterminée (ou quelque durée deduite d'un phénomène electrique constant) sera vraisemblablement la prochaine unité de temps.' This idea seems to be suggested first by Maxwell; the corresponding wave-length would at the same time be the standard of length, when the platinum 'metre stalon' will be given up. Thus it may happen that the 'kinetic length ' (i.e, that based on our notion of a 'rigid' body) will be sacrificed for the benefit of an optical or 'electromagnetic length' in the same way as the 'kinetic time' may be replaced by an 'electromagnetic time.'

[^9]:    * (Quoted after Painlevé, loc. cit. page 117 .

[^10]:    * As yet we have no need to touch upon the suljeet of combtutityer media.

[^11]:    * By 'uniform' we mean homogencous or constant in space and invariahle in time, $c$ lieing constant with respeet to leoth.

[^12]:    *II. A. Wilson, Ihil. Trans., A. Vol. CCIV. p. 121; 19ro.-...W. C. Roentgen, Berl. Sitsber., 1885 ; Wiedem. Ann., Vol. XXXV. 1888, and Vol. XL. 1890.A. Eichenwald, Ann. der I'hysik, Vol. XI. xgoz.

[^13]:    * This result was obtained by J. J. Thomson. See Heaviside's Electromagnetic Theory, Vol. III, 8471 et seq., where some interosting remarks regarding this and allied subjects may be found.

[^14]:    *H. Fizeau, Comples rendus, Vol. XXXIII., I85r ; Annales de Chimie, Vol. LVII., 1859.

[^15]:    * II. A. Lorentz, La theoric dedtromagndique de .Maxzeell at son application aux corps mouvants; Leiden, I.. J. Mrill, 1892 (also in Alch. mederl., Vol. XXV.).
    † II. A. Lorenty, Versuch einer Theorie der electrischen und ontischen Eirscheinungen atn beweglen Körpern; Leiden, E. J. Brill, 1895. This paper will be shortly referred to as 'Essay' ( $=$ L"orsuch $)$.
    $\$$ See Note $s$ at the end of this chapter.
    \& Lorentz, The Theory of Electrons, etc., Lectures delivered in Columbia University, 1906; Leip\%ig, Teubner, 1909 ; p. 230.

[^16]:    *This, of course, implies the possibilty of our following an individual portion or element of charge in its motion,-a subtle point (due to circuital indeterminateness, etc.), which, however, need not detain us here.
    $\dagger$ To be contrasted afterwards with his macroscopic (or average) equations.

[^17]:    * The treatment of possible exceptions to this assumption, as electromagnetic surfaces of discontinuity or waves properly so called [which exceptions seem to be overlooked by the leading electronists, who claim for (25) general validity], need not detain us here.
    $\dagger$ H. Poincaré, Arch. Néerland., Vol. V.; 1900.

[^18]:    * See Sections II. and IV. of Lorentz's Essay, or his article in Encykl. a. math. Wiss., Vol. V. ${ }_{2}$, pp. 200 et seq.; Leipzig, 1904.
    $\dagger$ Viz. 'polarization-electrons,' and leaving out of account circling or 'mag. netization-' and free or 'conduction-electrons.'

[^19]:    *As to dispersion, which need not detain us here, it can be accounted for in the well-known way by attributing to the body (or to its molecules) one or more internal, 'natural periods,' and, to introduce these, plenty of opportunities are offered by the hypothesis of the electronic structure of molecules and atoms.

[^20]:    * That the above assumption is satisfied with a sufficient degrec of accuracy may be seen from Note 3 at the end of the chapter, where the corresponding Huygens construction is worked out.

[^21]:    -     * E. W. Morley and D. C. Miller, Phil. Mag., Vol. VIII. p. 753, 1904; Phil. Mag., Vol. IX. p. 680, 1905.
    $\dagger$ As to various objections raised against the correctness of the interference experiment by Sutherland, Litroth and Kohl, and their refutation by Lodge, Lorentz, Debye and Laue, see the 'Literaturihersicht' in J. Laub's report 'Ueber die experimentellen Grundlagen des Relativitätsprinzips,' Jahrbuch der Radioaktivitüt und Elektronik, Vol. VII. p. 405, 1910.
    $\ddagger$ Cf. Lorentz's Essay, p. 122 (1895), where reference is made to a paper of his, dated 1892-93. As regards Fitzgerald, we read in The Efher of Space by Sir Oliver Lodge (London, 1909, p. 65), referring to that hypothesis: 'It

[^22]:    was first suggested by the late I'rofessor G. F'. Fitygerald, of Trinity College, Dublin, while sitting in my study at Liverpool and discussing the matter with me. The suggestion bore the impress of truth from the first.' Happy are those who are gifted with that immediate fecling for 'iruth.'

[^23]:    * H. A. Lorentz, 'Electromagnetic phenomena in a system moving with any velocity smaller than that of light,' Pror. Almsterdam Acad., Vol. VI. p. 809 ; 1904.

    HHere, according to the original definition of 'local time,' p. 66, we should have rigorously (instead of the coordinate $x$, measured in the fixed framework) $x-v t$, so that $t^{\prime}=\left(1+\beta^{2}\right) t-\frac{v}{c^{2}} x$. But, since at that stage $\beta^{2}$.terms were neglected, we could write simply $x$ instead of $x-v t$. The symbols $x^{\prime}$, etc., in what follows are not to be confounded with the $x^{\prime}$, etc., of page 66 .

[^24]:    * See Columbia University Lectures, pp. 211-212.

[^25]:    *A. Einstein, Annal. der Phlysik, Vol. XVII. p. 89r; 1905.

[^26]:    * We need not stop here to consider such apparatus as Siemens' 'sparkchronometer,' in which the visible marks corresponding to pairs of events are brought very close to one another, and which enable the modern physicist to fix with a high degree of precision their time-relations.

[^27]:    * In his paper (p. 892) Einstein begins with taking, for the purpose of his definition of simultancity, that 'system of coordinates in which Newton's mechanical equations are valid.' But it seems advisable not to appeal at the outset, and in connexion with such a fundamental definition, to Newtonian mechanics, especially as it requires, according to the relativistic view itself, some essential, though numerically slight, modifications. On the other hand, the physical specification of what has been called alove the system $S$ will appear presently without recourse to any theory of mechanics.

[^28]:    *We may notice in this connexion that Einstein's specification (p. 893): 'eine Uhr $[$ at $b]$ von genau derselben Beschaffenheit wie die in $A[a]$ befindliche' is unnecessary and, to a certain extent, misleading.

[^29]:    * Which Einstein himself, in order to have a convenient name, provisionally, calls 'the stationary system.'

[^30]:    *i.e. occurring nt one and the same place.
    $\dagger$ Both properties having been already attributed to it physically, i.e. as regards propagation of light, by II.

[^31]:    * In fact, what we call the velocity of $S$ relative to $S^{\prime}$ is the vector whose components are the derivatives of $x^{\prime}, y^{\prime}, z^{\prime}$ with respect to $t^{\prime}$, for constant $x, y, z$, that is to say, by (8),

    $$
    \frac{d x^{\prime}}{d l^{\prime}}=-v, \quad \frac{d y^{\prime}}{d l^{\prime}}=0, \quad \frac{d z^{\prime}}{d l^{\prime}}=0,
    $$

    and this is the vector - $\boldsymbol{v}$. In exactly the same way, the velocity of $S^{\prime}$ relative to $S$ is the vector whose components are the derivatives of $x, y, z$ with respect to $t$, for constant $x^{\prime}, y^{\prime}, z^{\prime}$, i.e., again by (8),

    $$
    \frac{d x}{d t}=v, \quad \frac{d y}{d t}=0, \quad \frac{d z}{d t}=0 .
    $$

[^32]:    * With regard to Einstein's own treatment of this subject, and also that adopted in Laue's book, see Note 1 at the end of the present chapter.
    + Some authors employ in this connexion the mathematically sanctioned term covariant, instead of invariant. But it will be convenient to reserve 'covariant' for another use, namely to denote that two groups of magnitudes are equally transformed.

[^33]:    S.R.

[^34]:    *Supposing, of course, that the inhabitants of cach system avail themselves only of one set of coordinates and time.

[^35]:    * Cf. for instance my Vestorial Mechanics, London, Macmillan \& Co., 1913, p. 97. The dots used there as separators are here replaced by (. Thus $\epsilon \mathrm{r}$ means $\boldsymbol{\gamma} \mathbf{i}(\mathbf{i r})+\mathbf{j}(\mathbf{j} \mathbf{r})+\mathbf{k}(\mathbf{k r})=\gamma \mathbf{i} x+\mathrm{j} y+\mathbf{k} z$.
    †Theorem 3 in Vol. I. of S. Lie's Theorie der Transformationsgruppen, Leipzig, 1888, p. 33. See also the whole of 'Kapitel 3. Eingliedrige Gruppen und infinitesimale Transformationen,' Ybiden, p. 45.
    $\ddagger$ That these transformations form a group, and that therefore Lie's theorem must be applicable to them, is easily seen. In fact, if $v_{1}$ is the velocity of $S^{\prime}$

[^36]:    * Ifere, as before, $x$, that is to say $x$ for $S$ as well as the new $x^{\prime}$ for $S^{\prime}$, is the coordinate measured along $\mathbf{v}$, the velocity of $S^{\prime \prime}$ with respect to $S$.

[^37]:    * H. G. Wells: The Time Machine, 1898 (Tauchnitz edition), p. 13. It is interesting to remark that even the forms used by Minkowski to express these ideas, as 'Three-dimensional geometry becoming a chapter of the four-dimensional physics,' are anticipated in Mr. Wells' fantastic novel. Here is another sample (loc. cit. p. 14), illustrative of what is now called a world-tube: 'For instance, here is a portrait [or, say, a statue] of a man at eight years old, another at fifteen, another at seventeen, another at twenty-one, and so on. All these are evidently sections, as it were, Three-Dimensional representations of his FourDimensioned being, which is a fixed and unalterable thing.' Thus, Mr. Wells seems to perceive clearly the absoluteness, as it were, of the world-tube and the relativity of its various sections.
    $\dagger$ In German, 'Auf Ruhe transformieren.'

[^38]:    * If we are to translate thus the names used by Minkowski: seitartiger and rcaunartiger Vektor respectively.

[^39]:    * The simple turning round of $x, y, x$, leaving $x^{2}+y^{8}+x^{2}$ invariant, leeing always left out of account. Having once assumed the isotropy of space, we have, speaking physically, no need to consider such rotations. And with regard to their mathematical role, see Chap. VI.

[^40]:    * To be distinguished, later on, from those ' of the second kind' or 'six-vectors.'

[^41]:    * Nor can we enter here upon a paper of E. B. Witson and G. N. Lewis, Proc. Amer. Acad. of Arts and Sciences, Vol. XLVIII., Nov. 1912, p. 389, in which an attempt is made to work out the four-dimensional vector-algebra and -analysis, ab ovo, starting from a number of quasi-geometric postulates.
    +Phil. Mrag., Vol. XXIII., 1912, p. 790, and Vol. XXV., 1913, p. 135 ; also Bull. of the Societas Scientiartum Varsaviensis, Vol. IV. fasc. 9, communicated in November, 191r. I wish to mention here that Dr. G. F. C. Searle has drawn my attention to a paper of Prof. Conway, Proc. Irish Acad., Vol. XXIX. Section A, March IgII, in which some of my results are arrived at. Particulars of comparison are left to the reader.

[^42]:    * If the reverse is the case, then $L X$ will have the structure of $q$.

[^43]:    *'Additionshecrem der (jeseliwindigkeiten,' Amn. d. Phys., Vol. XVII., 1905; \&5.
    HIf both were taken with respect to the same system, then their resultant would, of course, he simply equal to their vector sum. But this is hardly worth mentioning. For all cases of composition of velocitics, which have any physical interest, are of the type considered alove, tis. imply component velocities referred to a chain of different systems: $\Lambda \mathrm{n}$ object $B$ moves in a given way relatively to $A$, a third object $C$ moves relatively to $B$, and so on; find the motion of the last relative to the first.

[^44]:    * 'Propagation,' as here defined, does not necessarily involve any materia medium as the 'substratum' of the thing to be recognized and watched in it migrations, the only requirement being the possibility of its being watched so Thus, we may have 'propagation' of a clistortion along a rope, or of sound wave in air, or of electromagnetic 'disturbances' through empty space as well a through glass or water. The process of deteeting and watehing the waves o disturbances may be immediate in some, and very indireet in other cases, but thi does not bring in any essential differences.

[^45]:    * This was first pointed out explicitly by V. Varićak, Phys. Zeits 1910, pp. 93, 287, 586; cf. also Jahresbericht der deutschen Mat. Vol. XXI. 1912, p. 103, where all his contributions to the subjec But it must be noticed that materially the discovery was made previo Sommerfeld (Verh. deutsch. Phys. Ges., XI. p. 577), when he I relativistic formulae for the composition of velocities are ' $n 0$ long

[^46]:    * In fact, the reader will find at once that, for any pair of quaternions $a$, $a b_{c}+b a_{c}=2 S a b_{c}=2 S a_{c} b$.

[^47]:    * Lorentz, Phys. Zeitschrift, Vol. XI. 1910, calls $m$ the 'Minkowskian n and $\mathrm{V} X=\gamma \mathbf{N}$ the 'Minkowskian force,' since ( $\mathbf{1} 7$ ), with constant $m$, is equivalent of the four equations of motion sriven by Minkowski: Grundoleichut

[^48]:    * Here, $H I_{c}^{\prime}$ is an abbreviation for ( $\left.H_{c}\right)^{\prime}$, the transforn taking the conjugate of the transformed quateruion, (3r) $\left.\left(H^{\prime}\right)_{o}=\Omega_{c} H_{0}\right)$, so that $\left(H H_{c}\right)^{\prime}=\left(H^{\prime}\right)_{c}$, and hoith sides may, simply $H H_{c}^{\prime}$.
    t Remember that the square of the tensor, or the norm is $X X_{6}$. Now, in our case, $L$ being a scalaresss quatern $\mathbf{I}_{c}=-\mathrm{L}$, so that its norm is simply $-\mathrm{L}^{2}$. If L were an , we could write (instead of $-L^{3}$ ) $l^{2}$, the square of its siz But since L is a complex vector, or a bizector, the above n $\mathrm{L}^{2}$ is a scalar, of course, c.g. a complex sealar, as will be need not put the prefix $S$ before il, since VLL is always $n$ definition of vector product.

[^49]:    * The properties of this function, belonging to the elements of the

[^50]:    * This property of $\mathbf{R}=\mathbf{M}+\iota \mathbf{E}$ may also be deduced directly from that of

[^51]:    * In connexion with this sulject, Prof. A. W. Porter of University Colle London, draws my attention to experiments made by Poynting and Gra who tested for anisotropy of gravitation between two quartz spheres ( $P h$ Trans., 192, 1899, A. p. 245; cf. also Poynting and Thomson's Text-Book Physics, Properties of Matter, London, 1909, p. 48). Their results showed th this anisotropy could not amount in one case to more than one part in 2800, a in another case to more than one part in 16000 . On the other hand, proportional between mass and gravitation, furst tested by Newton in his endulum experime

[^52]:    * Fortschritte der math. Wiss. in Monographien, edited by O. Blum Heft I, Teubner, igio, p. 58.

[^53]:    * Or in Minkowski's matrix form. This consists of the two equations

[^54]:    * Of which the first and the last, denoted for subsequent reference are the quaternionic equivalents of Minkowski's world-vectors of the $\Phi$ and $\Psi$, called by him elektrische Ruh-Kraft and magnctische respectively. Cf. his Grundgleichungen, pp. 33-34-

[^55]:    *This tendency was initiated by l'anck's paper (l'hys. Ztitwheift, Vol. 1908, p. 828) on the principle of action and reaction. M. Abraham uses

