



ASTRONOMY LIBRARY

TREATISES
ON
PHYSICAL ASTRONOMY,
LIGHT AND SOUND.

CONTRIBUTED TO THE ENCYCLOPÆDIA METROPOLITANA.

BY
SIR JOHN F. W. HERSCHEL, BART., M.A., F.R.S.

ST JOHN'S COLLEGE, CAMBRIDGE.

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TREATISES

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PHYSICAL ASTRONOMY.

Astronomy THE object of the philosopher in the investigation of nature is to arrange and classify facts and phenomena, with a view to trace the agency of their remote, or, at least, their proximate causes, and ascend as high as the imperfection of human means of observation, and the limited powers of the human intellect will allow us in the scale of generalization.

To beings endowed with more perfect faculties, and more comprehensive intelligence than ourselves, much of that complication we observe in natural phenomena would disappear: many effects, which seem to us independent of each other, and linked by no natural connection, would be in their eyes collateral results of one and the same principle: much, that to us seems fortuitous would to them appear pre-disposed and regularly arranged. The laws of nature would at once be reduced in number and enlarged in extent; and that higher order of generalization which would consist in classifying together laws of the same kind, and referring them to others yet more universal, would exercise *their* power and constitute *their* science. That such would be the case with more perfect beings, our own experience, limited as it is, amply shews. The man who has learnt to regard the fall of a leaf and the precession of the equinoxes as results equally certain and unavoidable of a law capable of being stated in three lines, and understood by a child of ten years old, has made already a considerable step in this way—the patient exercise of his natural reason has stood him in the stead of a sharper intellect; and secrets which an angel might penetrate perhaps at a glance, become revealed to man by the slow, yet sure, effects of persevering thought.

The progress of modern science has done more than the keenest metaphysical reasoning, and has given us the most convincing proofs of the agency of one general and intelligent cause throughout the whole system of nature. When we see on all sides phenomena grouping themselves under laws intelligible and simply expressed, which are themselves subordinate to others, yet more simple and extensive; when we see every anomaly which threatened destruction to a theory, becoming, in the progress of our knowledge, its firmest support; every inequality disappearing when viewed from a higher level; every exception proving a rule of greater generality; all, in short, conveying more and more towards order and simplicity the more severely we scrutinize it; it is impossible not to allow that that last great step, which unites all the phenomena of the universe under one general head, and refers them to one all-pervading agency—however inconceivably remote, and surpassing probably the utmost limits of the human intellect to comprehend, if explained, would still be but the continuation and final completion of a

chain of reasoning whose first links we hold within our grasp—the consummation of a process actually begun—the termination of a career into which we are fairly entered. **Physical Astronomy.**

It is difficult to avoid such contemplations at the outset of an essay on physical astronomy; they crowd upon us; and in rejecting them we should reject the noblest use of the sublimest of sciences. For scarcely in any are the phenomena presented by nature more various and more complicated; in none is the generalization so complete, the final result so simple, or the object more imposing.

From what has just been said, it may be gathered that the object of physical astronomy is to reduce, under general laws, the motions and phenomena of the heavenly bodies, and investigate their causes; to trace the history of what has already happened in our own system, and to ascertain what changes the causes demonstrably in action (unless interfered with by others we have no knowledge of,) will superinduce in the course of ages, and thus to appreciate the stability of the present order of creation. In a more limited and practical point of view, the physical astronomer is called on to furnish formulæ deduced from theory for determining the state of the system at any assigned instant; and adapted for the purposes of the observer, so as to serve as a basis for the construction of tables; and to descend, by the application of his general principles, to those more refined inequalities which, owing to their minuteness, or the length of their periods, would escape or mislead the observer unassisted by theory.

There is one feature in physical astronomy which renders it remarkable among the sciences, and has been the chief, if not the only, source of the perfection it has attained. It is this—that the fundamental law embracing all the minutiae of the phenomena so far as we yet know them, presents itself at once, on the consideration of broad features and general facts, deduced by observations even of a rude and imperfect kind, in such a form as to require no modification, extension, or addition when applied in minute detail. In other sciences, when an induction of a moderate extent has led us to the knowledge of a law which we conceive to be general, the further progress of our inquiries frequently obliges us either to limit its extent or modify its expression. To those who are familiar with the history of chemistry, instances of this will present themselves at every turn. In physical optics, the general representation of all the series of polarised tints and the colours of natural bodies by a certain universal scale—the Cartesian law of refraction when applied to the extraordinary ray in crystallized media, and even to the ordinary, if the reports of some recent experiments are to be relied on—together with innumerable other laws, simple, natural,

Astronomy. and resting on extensive inductions, have all been either overset, extended, or materially modified by the progress of the science.

In physical astronomy, however, when taken in that limited acceptation, which restricts it to the explanation of the planetary motions, our first conclusion is our last. The law on which all its phenomena depend, flows naturally and easily from the simplest among them, as presented by the rudest observation; and, in point of fact, such has really been the order of investigation in this science. The rude supposition of the uniform revolution of the moon in a circle about the earth as a centre, led Newton at once to the true law of gravity, as extending from the earth to its companion. The uniform circular motions of the planets about the sun, in times following the progression assigned by observation in Kepler's rule, confirmed the law, and extended its influence to the boundaries of our system. Every thing more refined than this—the elliptic motions of the planets and satellites—their mutual perturbations—the slow changes of their orbits and motions, denominated secular variations—the deviation of their figures from the spherical form—the oscillatory motions of their axes, which produce nutation and the precession of the equinoxes—the theory of the tides, both of the ocean and the atmosphere, have all in succession been so many trials for life and death in which this law has been, as it were, pitted against nature; trials, whose event no human foresight could predict, and where it was impossible even to conjecture what modifications it might be found to need. Even at this moment, if among the innumerable inequalities of the lunar or planetary motions any one, however small, should be discovered decidedly not explicable on the hypothesis of a force varying as the inverse square of the distance, that hypothesis must be modified till it accounts for it. It is hardly necessary to add, however, that in the present state of science, this is a case not to be contemplated.

Still, these are refinements. The deviations of the planetary orbits from circles are small, their deviations from ellipses excessively minute: the lunar orbit alone presents results of perturbation so large as to strike us at once with the appearance of irreconcilable anomalies, but it is only by a refinement of calculation that we can trace them all to the laws of gravity; but the motions of comets put the truth and generality of this law to a severe and rude test, by giving it a trial under the greatest varieties of distance, position, and velocity of motion, and instancing its influence on matter of a rarity almost spiritual, and differing so utterly from that of which our planet consists, as scarcely to authorize the admittance of any property in common.

The above observations have been made in conformity with the general language of natural philosophers, and the customary acceptation of the term physical astronomy, and are, no doubt, strictly applicable when we confine ourselves to the celestial phenomena of our own immediate system, and the motions of those larger masses of matter of which planets and their satellites consist. The cautious philosopher however will still regard it as worthy inquiry, whether, at enormous distances, like those of the fixed stars, or at such comparatively microscopic intervals as those we are ordinarily conversant with on the surface of

our planet, the rigorous law of a force as the inverse square of the distance may not suffer some modification. An emanation, like light, traversing in succession every part of space, may be conceived to go on without acceleration or retardation, without loss or change, to the remotest regions; but an active and immediate intercourse carried on between points infinitely distant, is not only incapable of demonstration, but, could it be proved, must, I suppose, be referred to direct spiritual agency. At the same time, it is worthy attention how strict and indissoluble a bond gravity establishes between remote objects—to see this in its real light, we must compare it with the most effectual of our ordinary means of transmitting power. If the earth and sun were connected by a rod of cast iron, in one piece, an impulse or pull, however violent and sudden, applied at the sun, would not begin to be felt at the earth till after a lapse of eighteen months from the moment of its communication, while a change in the sun's attraction, such as might arise from a sudden alteration of its figure or density, would demonstrably* affect our planet in an instant of time many thousand times less than the least interval perceptible to our senses.

The subsistence of sidereal clusters, in which the compression or crowding of the stars is carried to the extent we have instances of in many parts of the heavens, seems hardly compatible with a gravitating force, unopposed by some principle of conservation, unless we suppose them in a state rapidly verging to a catastrophe. On the other hand, with regard to small distances, we have no distinct proof, that within a few inches, or even miles, from a material point, the law of gravity may not begin to deviate appreciably from the Newtonian law. The experiments of Maskelyne and Cavendish, which may perhaps be adduced as supporting its rigorous application, are far too gross, and differ too widely in their results, to be cited in so delicate a matter, besides which, their results, as applied to such an inquiry, are affected with an unknown element, the mean density of the earth. At much closer intervals we are certain of the existence of attractive and repulsive forces following a widely different law; and by what imperceptible gradations these shade into that of gravity, or whether they are to be regarded as distinct from it in their nature and origin, is a point whose consideration seems reserved for a much higher state of science than we can boast of having yet attained.

But it is quite sufficient for the purposes of physical astronomy to know, that as far as the motions of the great masses of matter connected with our system either in the heavens or on our globe are concerned, observation and theory present no difference capable of being made an objection to the strict expression of Newton's law, and we shall therefore wave all further discussion of the subject, and proceed to the object of the present essay, in which the reader will be presumed acquainted with the general facts of astronomy, with the principles of mechanical philosophy, and so much of analysis, of the differential and integral calculus, plane and spherical trigonometry, as shall render it unnecessary for us to interrupt the general chain of our reasoning to demonstrate such theorems, &c. as we shall have occasion to call to our aid.

* Laplace, *Système du Monde*, p. 286.

ON THE CIRCULAR AND ELLIPTIC MOTIONS OF THE PLANETS AND SATELLITES.

By observing the places of the sun, moon, and planets in the heavens at different times, and measuring their angular diameters, we have learnt, that, provided certain excessively small inequalities, to be hereafter considered, are disregarded, their motions are all compatible with the supposition of the planets revolving about the sun in elliptic orbits of small eccentricity, and having the sun's centre in one focus; each orbit lying wholly in a fixed plane peculiar to itself, and slightly inclined to the ecliptic, or plane, in which the sun appears to revolve about the earth. We learn, moreover, that if the apparent motion of the sun be transferred in a contrary direction to the earth, and the sun be supposed at rest in space, the motion so assigned to the earth will be such as to include it in the expression of the same law. The supposition then, of the sun at rest, and the earth in motion, being agreeable to this general analogy, and supported by incontrovertible arguments drawn from the great magnitude of the former in comparison with the latter body, is assumed as a demonstrated truth. Observation moreover assisted, it is true, by calculation, but independent of all theory, (i.e. of all reasoning from causes,) has taught us the truth of the following remarkable laws, in which also the earth is included among the planets.

The areas described about the sun's centre by the radius vector of any one of the planetary orbits (or the line drawn from the sun to the place of the planet,) are proportional to the times of their description.

Fig. 1.

Let S be the sun, and APP' part of the orbit of any planet. Then, A being assumed as a point of departure, the area ASP is to the area ASP' as the time of the planet's describing the arc AP of its orbit is to the time of its describing AP'.

The squares of the periodic times of different planets, (or of the times of a complete revolution of each about the sun,) are as the cubes of their mean distances from the sun, or of the greater semiaxes of their respective ellipses.

The periodic time of the earth is 365·2564, and that of Mars 686·9796. The greater semiaxis of the earth's orbit being 1, that of Mars is 1·523693; and we may easily satisfy ourselves, by executing the computation, that $(365·2564)^2 : (686·9796)^2 :: 1^3 : (1·523693)^3$.

These three laws, viz. 1st. The elliptic motion of the planets about the sun as a focus: 2dly. The proportionality of the areas to the times: and 3dly. The law of the periodic times, which have immortalized the name of Kepler, and whose discovery, and the manner of it, afford at once matter of humiliation and triumph to the human intellect, were all deduced immediately from observation, as insulated, and, for aught their discoverer knew, unaccountable facts. It shall now be our business to demonstrate their mutual dependence, and to shew how the general law of attraction may be derived from them most simply.

The analogy observed between the motions of the other planets and of the earth, affords a reasonable presumption of their being masses of matter subject

to the same mechanical laws of rest, impulse, and resistance, as that of which our own planet consists. Moreover, from what we know of the constitution of our own atmosphere, and its rapid diminution of density as we recede from the earth, we have every reason to believe, that the immense space in which their revolutions are performed, is either completely void, or at least free from any material substance capable of sensibly resisting or impeding their motions, or preventing any external impulse they may receive from acting on them with its full effect. Setting out with these assumptions (the strict truth of which will be best tried by the conclusions they will lead to,) it is obvious that as the planets, instead of moving continually forward in straight lines, as masses of inert matter would do if projected in space and left to themselves, are, in fact, constantly deviating from this rectilinear progression—they must be under the perpetual influence of some agency external to themselves, which, (by the second law of motion,) can be no other than that of a mechanical force acting in a direction inclined to that in which they move at any instant.

The enormous distance at which the planets are from the sun, and their own minuteness, compared with it, permit us at present to regard them as points; and it will be shewn hereafter that this supposition introduced here merely for simplification, is strictly legitimate. Let us then consider the motion of a material point perpetually deflected from a straight line by the action of an external force; and to this end let us conceive the curve OPQ described by the planet to be replaced by a polygon of an infinite number of sides OP, PQ, &c. and setting out from O, let it describe the chord OP in the first instant of time dt . Fig. 2.

In an equal subsequent instant it would, if left to itself, go on describing PR equal to OP, and in the same straight line. But since we have regarded the curve as replaced by an elementary polygon, we must (on the principles of the differential calculus) conceive the deflecting force to act by interrupted impulses at the angles of that polygon. Let the first impulse therefore be conceived to take place at P. Then, since the material point P, in virtue of the motion inherent in it at P, would have described PR in the instant dt ; but, in virtue of that motion, combined with the new motion it receives at P, does actually describe PQ in the same time, that new motion (by the composition of motions,) must be such as, *alone*, would carry it from P over a space Pv equal and parallel to RQ; and as the change of motion takes place in the direction in which the moving force acts, Pv must be the direction of the deflecting force. Prolong Pv indefinitely, and take in it any point S; join OS, SQ, SR. Then, since OP = PR, the area OSP = SPR = SPQ, because QR is parallel to PS.

A force may be conceived to *tend* to any point in the line of its direction. We see, therefore, that any point to which the force acting at P tends is characterised by this remarkable property, that the areas described about it in equal evanescent instants on either side of P are equal. This property belongs to every point in the line PvS, but (as is obvious) to no point situated out of that line; and any point possessing this property may be regarded (at least for that moment) as a point of tendency, or centre of the force acting on the body at P.

Now, as we have seen, it is matter of observation that each planet describes areas proportional to the times, and consequently equal areas in equal infinitely small times before and after any given instant, about a fixed point in the system coincident with the sun's centre. This point, therefore, possesses at all times and in all positions of the planets, the property above demonstrated to belong to a point situated in the direction of the deflecting force; or, in other words, the forces deflecting the planets in their orbits are invariably directed to the centre of the sun.

The moon, (neglecting periodical inequalities,) describes about the centre of the earth, and the satellites of Jupiter and Saturn about their respective primaries, areas proportional to the times of their description. The forces, therefore, which deflect them in their orbits, are directed (small causes of inequality being neglected) to the centres of the earth, of Jupiter, and of Saturn respectively.

Having ascertained the directions of the forces which deflect the planets from their rectilinear paths, and retain them in their orbits, we come now to estimate their intensity, and investigate the laws of their action. In order to this, we shall find it more simple to abandon the supposition of the interrupted impulsive action of the deflecting force, and consider the body P as deflected from the *tangent* PR, and describing not the chord PQ, but the infinitesimal arc, the force being supposed to act during the whole time dt . This time, however, being infinitely small, the force may be regarded as constant; and since the angle PSQ between its first and last directions is also evanescent, it must be considered as acting constantly in a direction parallel to PS or QR. If then we take \bar{F} to represent the force at P, and $g = 32^{ft} \cdot 1908$ (or double the space through which a heavy body falls in the first second at the earth's surface,) and suppose unity to represent the force of gravity, we shall have by mechanics,

$$RQ = \frac{1}{2} g \cdot F dt^2$$

and*

$$F = \frac{2QR}{g \cdot dt^2}.$$

If, therefore, we know by observation the nature of the orbit, and the velocity of the body at any point P, we may thence calculate the magnitude of the deflection QR produced in any very minute time dt : and thus the intensity of the deflecting force will become known. Let us, for instance, take the case of the moon; and, supposing her orbit a circle, with the earth in the centre, let us inquire the actual magnitude of QR, the deflection from the tangent produced in some extremely minute portion of time as $1''$, by the force retaining it in its orbit.

Call the mean radius of the moon's orbit R, her period (in seconds of mean time) T; then will her velocity (being equal to the circumference of her orbit, divided by the number of seconds in the time of one revolution,) be represented by $\frac{2\pi R}{T}$, where $\pi = 3.14159$, &c.; and this is the actual length of the arc described in $1''$. Now, since (neglecting higher powers of the arc than the square)

* QR here represents the deflection from the *tangent*, and is only half the length of QR in the last figure, which represents the deflection from the preceding chord prolonged.

$$\text{versed sine} = \frac{(\text{arc})^2}{\text{diameter}};$$

and since QR or Pv is in this case equal to the versed sine of PQ, we must have,

$$QR = \frac{1}{2R} \times \left(\frac{2\pi R}{T} \right)^2 = \frac{2\pi^2 R}{T^2};$$

and, finally, (since $dt = 1''$)

$$F = \frac{4\pi^2}{g} \times \frac{R}{T^2};$$

and this is the general expression for the force perpetually urging a body to the centre of a circle. To reduce it to numbers in the case of the moon, we have $R = 238783^m = 39165700 \times g$ (since $g = 32^{ft} \cdot 1908$)

$$\text{and } \frac{R}{g} = 39165700$$

moreover,

$$T = 27^d \cdot 32167 = 2360592''.$$

Calculating from which data, we find

$$F = 0.00026394 = \frac{1}{3522}.$$

So that the force by which the moon is retained in its orbit, is about 3522 times feebler than that of gravity at the earth's surface.

When we observe a tendency in all bodies at the earth's surface to approach or fall towards its centre, and if hindered from approaching, still to *press* towards that point, we express these phenomena, by saying, that they are attracted towards the earth. At all moderate elevations above its surface, and in the same geographical situation this tendency seems invariable; but at great elevations, the delicate indications of modern instruments will detect a decrease in its energy;* and indeed the gradual enfeebling of attraction towards any body by an increase of distance, is not only a natural supposition of itself, but is borne out by the strong analogy of magnetic and electrical attractions. At vast elevations then, like that of the moon, there is reasonable ground to expect a considerable diminution of attraction; and if the force by which the moon be retained in her orbit be nothing more than this same attraction modified by the remoteness of the two bodies, we see that an increase of the distance to about 60 times the earth's radius from its centre, is sufficient to weaken it more than 3500 times. As the distance then from the earth's centre increases, the attraction diminishes in a much more rapid progression. What the exact nature of this progression is, we must satisfy ourselves by other phenomena; but even from the rude calculation already made, (in which every correction has been neglected) we may perceive, that a law of decrease as the squares of the distances (the next in simplicity to the distances themselves,) has a *prima facie* probability. In fact, $60^2 : 1^2 :: 3600 : 1$.

Having only one attendant satellite, however, we

* At an elevation of a mile above the surface of the earth, the intensity of gravity is diminished $\frac{1}{1977.291}$; and a pendulum clock, beating seconds at the level of the sea, would lose 21.898 seconds a day at this altitude, a quantity not to be overlooked. Any traveller having leisure, and the proper apparatus, might try the experiment in the barrack on Mont Cenis, or at the Hospice of St. Bernard.

Astronomy. have no means of obtaining any further verification of such a law, in this way; but if we regard the earth as well as the other planets, as so many satellites of the sun, we have here ample room to satisfy ourselves, having a progression of no less than eleven distances, from that of Mercury to that of Uranus, on which to ground the assumption of a law. And here we have the advantage of dispensing altogether with numerical calculation, and substituting in its stead the third law of Kepler: for if we call R and r the radii of the circles described by any two planets round the sun, T and t their periodic times, F and f the forces retaining them in their orbits, we have (as a result of observation,)

$$\frac{T^2}{t^2} = \frac{R^3}{r^3},$$

which, combined with the equations,

$$F = \frac{4\pi^2}{g} \times \frac{R}{T^2}, \quad f = \frac{4\pi^2}{g} \times \frac{r}{t^2}$$

which give

$$\frac{F}{f} = \frac{R}{r} \cdot \frac{t^2}{T^2},$$

we shall find

$$\frac{F}{f} = \frac{r^2}{R^2}.$$

Thus, then we encounter the same rate of diminution in the attractive tendencies of each of the planets towards the sun, which the lunar motions had given reason to surmise in the case of the earth and moon; but in the case now under consideration, the verification is much more satisfactory, and the numerical coincidences, when the calculations are gone through, complete; the third law of Kepler, on which the whole is founded, being almost rigorously exact.

We may now, with great confidence, presume the inverse proportion of the squares of the distances to be the law of variation of that force which retains the bodies of our system in their orbits; but previous to assuming its generality, it will be right to compare the force retaining the moon in its path round the earth, and that deflecting the earth in its orbit about the sun. Calling R , r , the respective radii of the earth's and moon's orbits, and T , t , their periodic times, we have still

$$\frac{F}{f} = \frac{R}{r} \cdot \left(\frac{t}{T}\right)^2.$$

Now we have

$$\frac{r}{R} = \frac{60 \cdot 23799}{23405}; \quad \frac{T}{t} = \frac{365 \cdot 25638}{27 \cdot 32167}$$

So that, executing the calculation,

$$\frac{F}{f} = 2 \cdot 17399.$$

The sun then, although more than 380 times the distance of the moon, exerts a force of more than double the intensity on the earth compared with the earth's attraction on the moon. At equal distances, then, the forces exerted by the sun and earth would be in the ratio of

$$2 \cdot 17399 \times \left(\frac{R}{r}\right)^2 : 1 \text{ or } 328196 : 1.$$

This enormous difference in the attractive energies

of the two bodies, must evidently be owing to some equally striking difference in the bodies themselves; and when we consider the immense magnitude of the sun (in comparison with our planet) we shall not be at a loss to what cause to assign it. Whatever be the cause of attraction, we may fairly conclude that, if it be the result of a force inherent in matter, two equal and similar bodies (i. e. each containing the same quantity of attracting matter,) placed close together, will each attract a third placed at a distance, with equal forces; and both together, with double the force of either separately, and pursuing the same idea—that 328196 such bodies as the earth, if placed close together, and forming one mass in the place of the sun, would attract as the sun actually does: in other words, that the sun only attracts other bodies with more energy than the earth, by reason of its being a greater body, and containing a greater quantity of attracting *gravitating* matter.

By such reasonings we are led to assume, as a general law, that similar and equal particles of matter, however situated in space, attract, or tend to each other with a force directly proportional to their masses or quantities of gravitating substance, and inversely proportional to the squares of their distances from each other; and having arrived at this law by the steps described, we must now proceed to verify its rigorous exactness, by applying it in succession to the phenomena as presented by nature in our system, which will be the object of the following sections.

SECTION II.

On the attractions of spherical bodies

The earth, sun, and planets, as well as their satellites, being shewn by observation to be spherical bodies of great magnitude, it becomes necessary to examine, *in limine*, whether the law of attraction above stated be compatible with this fact—in other words, whether from a knowledge that the gross attractions of the whole masses follow that law of decrease, we can argue that the attraction of each elementary molecule follows the same.

Let BDCE be the attracting sphere, m the body Fig. 3. attracted, which at present we will suppose to be a single particle, taking its mass as unity. Suppose $mBAC$ the axis of the sphere passing through the molecule m , and let M , M' be two equal and similar molecules, similarly, but oppositely situated with respect to the axis. Each of these molecules will attract m with a force represented by $\frac{M}{(Mm)^2}$, but the directions of their attractions not coinciding, we must resolve them into others, whose effects may directly assist or counteract each other. Draw MPM' (which will of course be perpendicular to mAC), and if we take $Mm = f$, and $MP = \rho$, we shall have $\frac{M}{f^2}$ to represent the force of M on m , which, reduced to the directions mC and PM , give the partial forces $\frac{M}{f^2} \times \frac{mP}{mM}$ and $\frac{M}{f^2} \times \frac{MP}{mM}$, that is, (if we call AP , x , and AM , a)

$$\frac{M(a-x)}{f^3} \text{ and } \frac{M \cdot \rho}{f^3}.$$

Physical Astronomy.

Astronomy. The partial forces of M' are represented by the same quantities, but the latter of them, acting in the direction PM' destroys the partial force of M in the opposite direction PM , while the former conspires with the corresponding force of M , and doubles its effect, so that we have

$$\frac{2(a-x) \cdot M}{f^3}$$

for the attraction of this pair of molecules in the direction mA ; and the sum of all such pairs throughout the sphere being found by the ordinary rules of the integral calculus will express the whole attraction of the sphere. The simplest way will be to regard the molecule M as a parallelopiped included, first, between two consecutive positions of the plane DE perpendicular to the axis Am , separated from each other by the interval dx ; 2dly. Between two cylindrical surfaces, having for their bases the circle MM' , whose radius is $PM (= \rho)$ and the same circle in its consecutive position, when its radius varies from ρ to $\rho + d\rho$; and 3dly. Between two consecutive positions of the plane PMm , assumed during its rotation about Pm as an axis. In virtue of this, if we put the angle $FPM = \theta$ the dimensions of the molecule in a direction perpendicular to PM will be $\rho d\theta$, and its dimensions in other two directions being respectively dx and $d\rho$, we have

$$M = \rho d\rho \cdot d\theta \cdot dx;$$

so that the whole attraction (A) will be expressed by the triple integral

$$\iiint \frac{2(a-x) \cdot \rho d\rho \cdot d\theta \cdot dx}{f^3}.$$

The variables ρ , θ , and x , are here independent; and it is therefore indifferent with which we begin, we will commence with θ , because f being $= \sqrt{(a-x)^2 + \rho^2}$ is independent on θ . Thus we have

$$A = \iint \frac{2(a-x) \rho d\rho \cdot dx}{f^3} (\theta + \text{Const.})$$

This integral must be extended only from $\theta = 0$ to $\theta = \pi$, or over only *half* the circumference of the circle MM' , otherwise the attraction of each molecule M, M' (having been grouped in pairs,) would be repeated twice over. Then we have

$$A = \iint \frac{2\pi(a-x) dx \cdot \rho d\rho}{\{(a-x)^2 + \rho^2\}^{\frac{3}{2}}}.$$

If we now perform the integration relative to ρ regarding x as constant, we get

$$A = \int 2\pi(a-x) dx \cdot \left\{ \text{Const.} - \frac{1}{\sqrt{(a-x)^2 + \rho^2}} \right\}$$

But the integral in this case is to be extended from $\rho = 0$ to $\rho = PD = \sqrt{r^2 - x^2}$, r being the radius of the sphere, so that it becomes

$$\begin{aligned} A &= \int 2\pi(a-x) dx \cdot \left\{ \frac{1}{a-x} - \frac{1}{\sqrt{(a^2 + r^2) - 2ax}} \right\} \\ &= 2\pi x + \pi \int \frac{(a-x) d \cdot (a-x)}{\sqrt{2a(a-x) - (a^2 - r^2)}} \\ &= \text{Const.} + 2\pi x + \frac{2\pi}{3} \left\{ 2a^2 - r^2 - ax \right\} \cdot \frac{\sqrt{a^2 + r^2 - 2ax}}{a^2} \end{aligned}$$

This integral must be extended from $x = AC = -r$ to $x = AB = +r$, when it finally becomes, after all reductions,

$$A = \frac{4}{3} \cdot \frac{\pi r^3}{a^2}.$$

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Now, $\frac{4}{3} \pi r^3$ represents the mass of the sphere, which being called S , we have

$$A = \frac{S}{a^2},$$

an equation which shews, that the attraction of the sphere is expressed by the whole mass, divided by the square of the distance of its centre from the attracted molecule, and is therefore precisely the same as if the whole sphere were condensed into its centre.

The hypothesis, then, which refers the observed attractions of the great masses composing our system to the effect of the mutual attraction of their ultimate molecules, varying according to the same law, has nothing in it incompatible with mathematical reasoning; but it is a very remarkable coincidence that this should so happen, as the only mathematical laws of attraction which would lead to a similar conclusion, are that of nature, and that in which the force is directly proportional to the distance, or one resulting from the combination of these two laws.

Let us next examine the case when the attracted body is also a sphere of sensible magnitude. S and s being the two spheres, and a the distance of their centres, it has been shewn that S will attract every molecule of s with the same force as if it were condensed into its centre. Now the mass of a single molecule having been regarded as unity, S , the mass of the first sphere will be proportional to, and represent, the number of molecules it consists of. The attraction then of s on S will be the same as if the latter sphere were removed, and in its centre a single molecule placed, endowed with an attractive energy S times as great as that of any molecule, such as S actually consists of. But the attraction of s on one molecule of the last named kind has been shewn to be $\frac{s}{a^2}$, therefore its attraction on S being S times as

forcible, will therefore be represented by $\frac{S \times s}{a^2}$

This expression represents the absolute force with which the two spheres tend to each other, or the number of pounds, grains, or other units, which must be opposed to it in order to hinder their approach. This in mechanical language, is called the *moving force*; and we therefore see, that the moving force with which two homogeneous spheres attract each other, is as the product of their masses directly, and the square of the distances of their centres inversely.

The moving force then is, of course, the same on each sphere; and in consequence of the equality of action and re-action, (which always refers to *moving force*,) it ought to be so. Were the spheres allowed to approach each other, however their velocities would obviously be different, the greater moving slower than the less. In fact, the accelerating force on any body being equal to the moving force divided by the mass moved, we have

$$\text{accelerating force on } S = \frac{s}{a^2}.$$

$$\text{accelerating force on } s = \frac{S}{a^2}.$$

Suppose now the two spheres at liberty in space, and

Astronomy. moving, in consequence of their mutual attraction and any projectile force. If to both of them we apply an accelerating force $\frac{s}{a^2}$ equal to that exerted by s , but in an opposite direction, or towards S , the sphere S will be urged by forces destroying each other, and will therefore either remain at rest, or move uniformly in a right line: but the sphere s will now be urged by an accelerating force equal to $\frac{S+s}{a^2}$. Now, it is shown in mechanics, that the application of a common accelerating force to all the bodies of a system, does not alter their relative motions. Hence, if we refer the motions of our spheres, not to a fixed point in space, but to the centre of one of them S ; or take that centre as the origin of our co-ordinates, we must then put $\frac{S+s}{a^2}$ for the accelerating force animating the other.*

Let us now consider the attractions of spheres not homogeneous, but composed of concentric strata varying in density according to any law of the distance from their centres. There is every reason to suppose this the actual constitution of the sun and planets; and it therefore becomes necessary to examine this case. Now, any stratum of infinitesimal thickness dr , may be regarded as the difference of two spheres s , and $s+ds$, homogeneous, and of the same density as the stratum, their radii being r and $r+dr$. The attraction of s on a molecule equal to 1 placed at a distance a , is $\frac{S}{a^2}$, and that of $s+ds$ is $\frac{s+ds}{a^2}$: consequently the attraction of the stratum is $\frac{s+ds}{a^2} - \frac{s}{a^2} = \frac{ds}{a^2}$; and therefore the same as if the stratum were collected in its centre. As this is true of every stratum separately, whatever be its density, and their attractions do not interfere, it will be true of all together; so that, whether the sphere be homogeneous, or composed of concentric layers, or strata of different density, the same property still holds good; and all we have demonstrated in the case of homogeneous spheres, remains true in this.

SECTION III.

Theory of elliptic motion.

We are now enabled to enter on the general theory of the planetary motions, but we will still confine ourselves to a case of comparative simplicity. The vast mass of the central body of our system, compared with those which circulate round it, permits us to regard their motions as influenced by it alone, and to neglect, in a first approximation, all the minute effects arising from the mutual attractions of the planets and satellites on each other. The case then we propose to consider in this section is that of the sun and a single planet, or a primary and one of its satellites.

* If, however, we suppose the sphere S forcibly retained in its place by some external agency *not* acting on s , the case will be different; and $\frac{S}{a^2}$ will continue to represent the accelerating force on S .

Physical Astronomy. Let M represent the mass of the sun or central body, and m that of the planet; and, fixing the origin of the co-ordinates in the centre of M , let x, y, z , represent the co-ordinates of m . Also let r be the radius vector, or line joining the centres of the two bodies; so that $r^2 = x^2 + y^2 + z^2$, and t the time (in seconds of mean solar time,) elapsed since any fixed epoch.

The accelerating force of M on m is $\frac{M}{r^2}$, and that of Fig. 4.

m on M $\frac{m}{r^2}$; and since we regard M as fixed, the latter quantity must be added to the force animating m , according to the observation made in the last section; so that the relative accelerating force acting on m in the direction of the radius vector mM will be $\frac{M+m}{r^2}$ which being resolved into forces in the directions mP, PQ, QM , of the three co-ordinates, becomes multiplied by $\frac{x}{r}, \frac{y}{r},$ and $\frac{z}{r}$, respectively, and produces the partial forces

$$P = \frac{(M+m)x}{r^3}, Q = \frac{(M+m)y}{r^3}, R = \frac{(M+m)z}{r^3}.$$

The effect of an accelerating force P acting, during an instant of time dt , on a body in a direction parallel to any given axis, that of the x , is to produce a variation in its velocity in that direction, which is to the variation gravity on the earth's surface would produce in the same time, as the force P is to the accelerating force of gravity which we will represent by unity. Now gravity producing the variation $g \cdot dt$ in that time, the variation produced by the force P , will be $P \cdot g dt$; or, if instead of taking one foot, as we have hitherto done, for the unit of linear measure, we take g ($= 32^{\text{ft}} \cdot 1908$) for our standard unit, simply $P dt$. But to this the negative sign should be prefixed, as the force P tends to diminish the co-ordinate x . Again, the velocity in the direction x being $\frac{dx}{dt}$, its variation

is $d \frac{dx}{dt}$; we have therefore $d \frac{dx}{dt} = -P dt$; or, sup-

$$\text{posing} \quad R = \frac{M+m}{r^3}$$

and writing for P its value Rx

$$d \frac{dx}{dt} + Rx = 0; \quad (1)$$

similarly,

$$d \frac{dy}{dt} + Ry = 0; \quad (2)$$

and

$$d \frac{dz}{dt} + Rz = 0; \quad (3)$$

These equations contain the whole theory of the planetary motions, neglecting their mutual perturbations; and if, instead of supposing $R = \frac{M+m}{r^3}$, we

suppose it equal to $\frac{\phi(r)}{r}$, the same equations will express the motion of a point m about a centre of force M , attracting it with a force represented by any function $\phi(r)$ of the distance Mm .

If we eliminate R from the two first of these

Astronomy. equations, by multiplying the first by y , and the second by $-x$, and adding, we get

$$y d \frac{dx}{dt} - x d \frac{dy}{dt} = 0$$

that is,
$$\frac{y d^2 x - x d^2 y}{dt} = \frac{(y dx - x dy) d^2 t}{dt^2}$$

or
$$\frac{y d^2 x - x d^2 y}{y dx - x dy} = \frac{d^2 t}{dt}$$

Each member of this equation being a complete differential, because $y d^2 x - x d^2 y = d \cdot (y dx - x dy)$, we get by integration

$$y dx - x dy = h dt; \quad (4)$$

and similarly,

$$z dx - x dz = h' dt; \quad (5)$$

$$z dy - y dz = h'' dt; \quad (6)$$

If, now, we multiply the first of these three equations (in which h , h' , and h'' , represent arbitrary constant quantities) by z , the second by $-y$, and the third by x , and add all together, we get

$$h z - h' y + h'' x = 0; \quad (7)$$

which is the equation of a plane passing through the origin of the co-ordinates. Consequently, the curve described by the body is one of simple curvature, and its plane passes through the centre of attraction.

For simplicity, let us suppose the plane of the orbit to be coincident with that of the x and y , we have then $z = 0$, and our equations are reduced to

$$d \frac{dx}{dt} + R x dt = 0$$

$$d \frac{dy}{dt} + R y dt = 0$$

$$y dx - x dy = h dt$$

Fig. 5. The area of the elementary sector $M m m'$, described by the radius vector $M m$ in the instant dt , is equal to

$$\begin{aligned} & M m m' P' - M m' P' \\ &= M m P - M m' P' + P m m' P' \\ &= -d \cdot \frac{xy}{2} + y dx = \frac{y dx - x dy}{2} \end{aligned}$$

If, then, we call A the area described by the radius vector since the commencement of the time t , we have $y dx - x dy = 2 dA$, and consequently

$$2 dA = h dt; \quad A = \frac{h}{2} t; \quad (8)$$

this equation expresses the proportionality between the areas and times in Kepler's law; and since the process by which it is deduced, is independent of any particular value of R , R having been eliminated to obtain it, the analytical demonstration here given applies generally for all possible laws of central force.

In order next to investigate the nature of the curve described, we must eliminate t , which will be easiest done after a transformation of the co-ordinates. Let then the angle $AM m$ described by the radius vector, since the origin of the time t , be called θ , and we have $m\mu = r d\theta$, and the elementary sector $m M m' = \frac{r \times r d\theta}{2}$, or $dA \frac{1}{2} r^2 d\theta$, so that

$$r^2 d\theta = h dt; \quad (9)$$

Now, since $r^2 = x^2 + y^2$, we have

$$r dr = x dx + y dy$$

$$r d^2 r + d r^2 = (x d^2 x + y d^2 y) + (dx^2 + dy^2)$$

so that (putting $ds = m m' = \sqrt{dx^2 + dy^2}$)

$$x d^2 x + y d^2 y = r d^2 r + d r^2 - ds^2.$$

But $ds^2 - d r^2 = m m'^2 - m' \mu^2 = m \mu^2 = r^2 d\theta^2$

whence, $x d^2 x + y d^2 y = r d^2 r - r^2 d\theta^2$.

This premised, since our equations (1) and (2) give

$$0 = d^2 x - \frac{dx}{dt} d^2 t + R x \cdot dt^2$$

$$0 = d^2 y - \frac{dy}{dt} d^2 t + R y \cdot dt^2$$

if we multiply the first of them by x , and the second by y , and add, we have

$$0 = (x d^2 x + y d^2 y) - \frac{x dx + y dy}{dt} d^2 t + R r^2 dt^2 \quad (10)$$

but if we take the logarithmic differential of the equation $h dt = r^2 d\theta$, and suppose $d\theta$ constant, (which we are at liberty to do, having as yet taken no differential constant,) we get

$$\frac{d^2 t}{dt} = 2 \cdot \frac{dr}{r}.$$

So that, making this substitution for $\frac{d^2 t}{dt}$, and for

$x d^2 x + y d^2 y$, writing its value $r d^2 r - r^2 d\theta^2$, and $r dr$ for $x dx + y dy$, we get

$$0 = r d^2 r - r^2 d\theta^2 - 2 dr^2 + R r^2 dt^2,$$

in which we have put for R its value $\frac{M+m}{r^3}$, and

for dt , $\frac{r^2 d\theta}{h}$, it becomes

$$0 = r d^2 r - r^2 d\theta^2 - 2 dr^2 + \frac{M+m}{h^2} \cdot r^3 d\theta^2$$

Put $r = \frac{1}{u}$, and since $dr = -\frac{du}{u^2}$, and $d^2 r = -\frac{d^2 u}{u^2} + \frac{2 du^2}{u^3}$, it becomes

$$\frac{d^2 u}{u^2} + \frac{2 du^2}{u^3} + u - \frac{M+m}{h^2} = 0; \quad (11)$$

This equation (being the simplest case of an equation of the second order of a linear form,) is immediately integrable, and gives

$$u = f \cos(\theta + g) + \frac{M+m}{h^2}; \quad (12)$$

whence,

$$r = \frac{\frac{h^2}{M+m}}{1 + \frac{h^2 f}{M+m} \cdot \cos(\theta + g)}; \quad (13)$$

f and g being two arbitrary constants.

In any conic section, if we call a the semiaxis major, $a(1-e^2)$ the semiparameter, and $\theta + g$ the angle included between the radius r and the vertex nearest the focus from which r is supposed to take its origin, we have

$$r = \frac{a(1-e^2)}{1 + e \cdot \cos(\theta + g)} \quad (14)$$

Consequently we see that the curve described by the body must necessarily be a conic section, having the body M in the focus, and the relation between the arbitrary constants f , h , and the axis and semipara-

Astronomy meter of this conic section will be (if we call p the semiparameter, or put $p = a(1 - e^2)$)

$$f = \frac{e}{a(1 - e^2)} = \frac{e}{p}; \quad (15)$$

$$h = \sqrt{a(1 - e^2)(M + m)} = \sqrt{p(M + m)}; \quad (16)$$

If we suppose the angle θ to commence from the vertex nearest the focus, or from the *nearer apside*, or *perihelion* of the orbit, we have $g = 0$, and

$$r = \frac{a(1 - e^2)}{1 + e \cos \theta} = \frac{p}{1 + e \cos \theta}; \quad (17)$$

If the value of e be less than unity, the conic section described is an ellipse, if equal and a infinite, a parabola, if greater, and a negative, an hyperbola.

But to complete the theory of the planetary motions, it is necessary to know, not only the nature of the orbit, in general, but also whereabouts in it the body will be at any moment assigned. For this purpose we must obtain a finite equation involving t , and either r or θ , or some functions of them. Now, the equation (9) gives $h dt$, or

$$\sqrt{a(1 - e^2)(M + m)} \times dt = r^2 d\theta$$

and substituting for r , its value in equation (17)

$$dt = \frac{a^{\frac{3}{2}}(1 - e^2)^{\frac{3}{2}}}{\sqrt{M + m}} \cdot \frac{d\theta}{(1 + e \cos \theta)^2} \quad (18)$$

and

$$t + C = \frac{a^{\frac{3}{2}}(1 - e^2)^{\frac{3}{2}}}{\sqrt{M + m}} \cdot \int \frac{d\theta}{(1 + e \cos \theta)^2}$$

To integrate this, take another variable v , such, that

$$\cos \theta = \frac{\cos v - e}{1 - e \cos v}; \quad (19)$$

whence,

$$\sin \theta = \frac{\sqrt{1 - e^2} \sin v}{1 - e \cos v}$$

$$d \cos \theta = -dv \cdot \frac{(1 - e^2) \sin v}{(1 - e \cos v)^2}$$

$$d\theta = -\frac{d \cos \theta}{\sin \theta} = \frac{dv \cdot \sqrt{1 - e^2}}{1 - e \cos v}$$

$$1 + e \cos \theta = \frac{1 - e^2}{1 - e \cos v}$$

and substituting these expressions in the value of t above given, it will be found to reduce itself to the following very simple form:

$$t + C = \frac{a^{\frac{3}{2}}}{\sqrt{M + m}} \int dv (1 - e \cos v)$$

so that, making

$$\frac{a^{\frac{3}{2}}}{\sqrt{M + m}} = \frac{1}{n}; \quad (20)$$

and taking the integral to commence, when $v = 0$, or $\theta = 0$, that is, from the instant of the body leaving the lower apside,

$$nt = v - e \sin v; \quad (21)$$

This equation fixes the relation between t and v ; but that between v and θ may be expressed more conveniently for the purposes of calculation than by equation (19), as follows. By the equation last mentioned, we get

$$1 - \cos \theta = \frac{(1 + e)(1 - \cos v)}{1 - e \cos v}$$

$$1 + \cos \theta = \frac{(1 - e)(1 + \cos v)}{1 - e \cos v}$$

whence,

$$\frac{1 - \cos \theta}{1 + \cos \theta} = \frac{1 + e}{1 - e} \cdot \frac{1 - \cos v}{1 + \cos v}$$

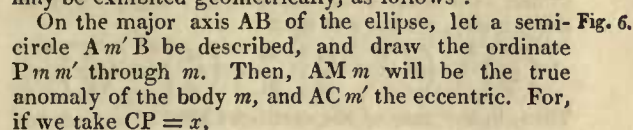
or,

$$\tan \frac{1}{2} \theta = \sqrt{\frac{1 + e}{1 - e}} \cdot \tan \frac{1}{2} v; \quad (22)$$

Finally, we obtain immediately the relation between r and v , by substituting, in equation (17) for $\cos \theta$, its value in (19), when we find

$$r = a(1 - e \cos v); \quad (23)$$

These equations comprise the whole theory of the motions of bodies in conic sections. Equation (17) exhibits the relation between r and θ , or the polar equation to the curve; while (21), (22), and (23), express relations between the time, t , the *true anomaly* (as it is called) θ , and the *radius vector*, r , respectively, and an auxiliary angle v , to which the name of the *eccentric anomaly* has been given. The quantity nt , when reduced into angular measure, by multiplying it by 180° , and dividing it by $\pi = 3.14159$, &c. is called also the *mean anomaly*. The eccentric anomaly may be exhibited geometrically, as follows:

On the major axis AB of the ellipse, let a semi- circle Am'B be described, and draw the ordinate Pm'm' through m . Then, AM m will be the true anomaly of the body m , and AC m' the eccentric. For, if we take CP = x ,

$$\cos ACm' = -\frac{CP}{Cm'} = -\frac{x}{a}$$

$$\cos AMm = -\frac{MP}{Mm} = -\frac{ae + x}{a + ex}$$

because, by the property of the ellipse, $Mm = a + ex$. Hence, we have

$$\begin{aligned} \frac{1 - \cos AMm}{1 + \cos AMm} &= \frac{(1 + e)(a + x)}{(1 - e)(a - x)} \\ &= \frac{1 + e}{1 - e} \times \frac{1 + \frac{x}{a}}{1 - \frac{x}{a}} \\ &= \frac{1 + e}{1 - e} \cdot \frac{1 - \cos ACm'}{1 + \cos ACm'} \end{aligned}$$

and, consequently,

$$\tan \frac{1}{2} AMm = \frac{1 + e}{1 - e} \times \tan \frac{1}{2} ACm'$$

Let T represent the time of one entire revolution in the ellipse, or the periodic time, then, as θ increases from 0 to 360° (or 2π) t increases from 0 to 360 (2π) also. Consequently, by (21) we get

$$nT = 2\pi; \quad T = \frac{2\pi}{n}$$

or

$$T = \frac{2\pi \cdot a^{\frac{3}{2}}}{\sqrt{M + m}}; \quad (24)$$

The periodic times, then, of several bodies revolving about the same central body, are, in the sesquuplicate ratio of the major axes of their orbits, (or mean distances from the central body,) directly, and in the subduplicate ratio of the sum of the masses of the revolving and central body inversely. The mass of the sun being enormously great compared with

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Astronomy. those of the planets, we may neglect m in comparison with M ; and M being the same for the whole system, we have

$$T \propto a^{\frac{3}{2}}; \quad T^2 \propto a^3$$

which is no other than Kepler's third law.

The periodic times of the planets being very exactly known, we might expect to find in equation (24) the means of ascertaining the masses of the planets, supposing that of the sun, and any one of them, known. Thus, if m, m' be the masses of two planets, and T, T' , their periodical times, we have

$$\left(\frac{T'}{T}\right)^2 = \left(\frac{a'}{a}\right)^3 \times \frac{M+m}{M+m'} \quad (25)$$

On applying calculation, however, all we learn from this relation is, that the resulting masses are so small, as to be incapable of accurate determination by this method; their values, as deduced from it, being materially affected by the small uncertainties still prevailing as to the lengths of the periods, and by the mutual perturbations of the sun and planets. There is a case, however, where it may be used with advantage, viz. in that of a planet accompanied by a satellite. If we call M the sun, and m the planet, and neglect the mass of the latter in comparison with the former, and that of the satellite in comparison with the primary, we have at once

$$\left(\frac{T'}{T}\right)^2 = \left(\frac{a'}{a}\right)^3 \times \frac{M}{m}; \quad \frac{m}{M} = \left(\frac{a'}{a}\right)^3 \times \left(\frac{T}{T'}\right)^2$$

Thus, in the case of the earth, we have

$$\frac{a'}{a} = \frac{60 \cdot 23799}{23405}; \quad \frac{T}{T'} = \frac{365 \cdot 25638}{27 \cdot 32167}$$

So that, by executing the computations, we find

$$\frac{m}{M} = 0 \cdot 00000304697 = \frac{1}{328196}$$

To find v in terms of t , or to calculate the eccentric (and thence the true) anomaly at any given instant, we must resolve the transcendental equation,

$$nt = v - e \cdot \sin v$$

This can only be done in a series; and fortunately, in the case of the planets, e is so small, that a series ascending by powers of e will converge sufficiently. Now we have

$$v = nt + e \cdot \sin v$$

and, since e is small, and $\sin v$ necessarily less than 1, nt itself expresses the value of v within a limit less than e , and is therefore a first approximation. Again, if in $\sin v$ we write its value for v or $(nt + e \cdot \sin v)$, we get

$$v = nt + e \cdot \sin (nt + e \cdot \sin v)$$

So that,

$$v = nt + e \cdot \sin nt$$

is an approximation carried one step farther, or to the first power of e . Let this be again substituted for e , and we have

$$v = nt + e \cdot \sin \{nt + e \cdot \sin nt\}$$

But (neglecting the squares and higher powers of e) we have

$$\begin{aligned} \sin \{nt + e \cdot \sin nt\} &= \sin nt + e \cdot \sin nt \cdot \cos nt \\ &= \sin nt + \frac{e}{2} \sin 2nt \end{aligned}$$

So that we find

$$v = nt + e \cdot \sin nt + \frac{e^2}{2} \sin 2nt \quad (26)$$

If we again repeat the process, we get an approxima-

tion, pushed to the third power of e , and so on, as far as we please.

For numerical calculation, however, the equation $v = nt + e \cdot \sin v$ furnishes the readiest solution; as we have only to reduce e into seconds, (taking $57^\circ 17' 44'' \cdot 8$ for the arc equal to radius or 1, or adding the logarithm $5 \cdot 3144251$ to the logarithm of e , which gives at once that of the number of seconds e is equal to,) and assuming nt (the given mean anomaly) for a first approximation, correct it successively, as in the following example.

Required Jupiter's eccentric anomaly corresponding to 53° of mean anomaly.

Here $nt = 53^\circ$, and in Jupiter's orbit we have

$$e = 0 \cdot 048077 \quad \log . 8 \cdot 6819374$$

$$5 \cdot 3144251$$

$$e = 9916'' \cdot 6; \quad \log . 3 \cdot 9963625$$

$$\text{Take } v = 53^\circ \dots \log \sin 53^\circ \quad 9 \cdot 9023486$$

$$e \cdot \sin v = 2^\circ 12' 0'' = 7919'' \cdot 8; \quad 3 \cdot 8987111$$

$$nt = 53$$

$v = 55 \quad 12 \quad 0$ Corrected value, with which resuming the process.

$$\log \sin v = 9 \cdot 9144221$$

$$e \dots 3 \cdot 9963625$$

$$2^\circ 15' 43'' = 8143'' \cdot 0 \dots \log \dots 3 \cdot 9107846$$

$$53^\circ \dots = nt$$

$55 \quad 15 \quad 43 \dots$ Second corrected value.

Another repetition of the very same process gives $v = 55^\circ 15' 49'' \cdot 1$, which is true within $0'' \cdot 2$.

But even this process, simple as it is, becomes tedious for the orbit of Mercury, and those of the new planets, Pallas and Juno, in which the value of e is not very small; and here we must have recourse to the well known method of trial and error, which may be applied in this case as follows:—Having assumed by estimation a value of v , (neglecting minutes and seconds,) by noticing whether the term $e \cdot \sin v$ is additive or subtractive, and increasing or diminishing nt accordingly; calculate the value of $v - e \cdot \sin v$ for that and the next subsequent degree; and let the values so found be called V , and V' : should either of these be exactly equal to the proposed value of nt , the corresponding value of v will be the truth; but as this will probably never happen, we have only to say,

$$V' - V : V' - nt :: -3600'' : x = -\frac{3600''}{V' - V} \times (V' - nt)$$

which correction being applied with its proper sign to the latter of the two assumed values of v , will give an approximate value. Let the value of $v - e \cdot \sin v$ be again computed with this value of v , and call the result V'' . This will always be found very nearly equal to nt ; but if not exactly so, the correction

$$-\frac{3600}{V'' - V} (V'' - nt)$$

must be computed and applied to the new value, and so on.

For instance: Let $nt = 332^\circ 28' 65''$, and $e = 50600'' = 14^\circ 3' 20''$. Here $\sin nt$ is negative; so that $e \cdot \sin nt$ is subtractive, and v must be less than nt . Take then for the two values of v , 325° and 326° respectively, and compute as follows:—

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$$\begin{aligned}\sin 325^\circ &= 9.7585913 \text{ (neg)} \\ \log e'' &= 4.7041513\end{aligned}$$

$$\begin{aligned}\sin 326^\circ &= 9.7475617 \text{ (neg)} \\ 4.7041513\end{aligned}$$

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$$\begin{aligned}e \cdot \sin 325^\circ &= -29023''.0 = -8^\circ 3' 43''.0 \\ \text{Hence } V &= 333^\circ 3' 43''.0 \\ V' - V &= 2872''.2\end{aligned}$$

$$\begin{aligned}4.4627426 \text{ (neg)} \\ e \cdot \sin 326^\circ &= -28295''.2 = -7^\circ 51' 35''.2 \\ V' &= 333^\circ 51' 35''.2 \\ V' - nt &= -4960.2\end{aligned}$$

$$\begin{aligned}\log 3600 &= 3.5563025 \\ \log 2872.2 &= 3.4582147\end{aligned}$$

$$\begin{aligned}0.0980878 \\ \log 4960.2 &= 3.6954992\end{aligned}$$

$$\begin{aligned}\log 6217.1 &= 3.7935870 \\ v &= 326^\circ - 6217''.1 = 324^\circ 16' 23''\end{aligned}$$

Taking this for a new value of v , we find, by another repetition of the process, Finally, if we resolve the equation

$$V'' = 332^\circ 28' 49''.6; \quad V'' - nt = -5''.4$$

$$\text{and, } -\frac{3600}{V' - V} \times (V'' - nt) = +6''.768$$

$$\text{so that } v = 324^\circ 16' 29''.768$$

which is true to the hundredth of a second; and this case is nearly the worst that can occur in the theory of the planets.

If we substitute the value of v given in equation (26) in the expression for r , $r = a(1 - e \cdot \cos v)$; (23), and develop in powers of e , we obtain r in a series of powers of e , and cosines of nt , and its multiples,

$$r = a \left\{ 1 - e \cdot \cos nt + \frac{e^2}{2} (1 - \cos 2nt) - e^3 \times \&c. \right\};$$

We shall content ourselves here with merely setting down the formulæ, which are as follows:—

$$v = nt + e \cdot \sin nt + \frac{e^2}{1 \cdot 2 \cdot 2} \cdot 2 \cdot \sin 2nt$$

$$+ \frac{e^3}{1 \cdot 2 \cdot 3 \cdot 2^2} \left\{ 3^2 \cdot \sin 3nt - 3 \cdot \sin nt \right\} \quad (27)$$

$$+ \frac{e^4}{1 \cdot 2 \cdot 3 \cdot 4 \cdot 2^3} \left\{ 4^3 \cdot \sin 4nt - 4 \cdot 2^3 \cdot \sin 2nt \right\}$$

$$+ \frac{e^5}{1 \cdot 2 \cdot 3 \cdot 4 \cdot 5 \cdot 2^4} \left\{ 5^4 \cdot \sin 5nt - 5 \cdot 3^4 \cdot \sin 3nt + \frac{5 \cdot 4}{1 \cdot 2} \cdot \sin nt \right\}$$

+ &c.;

$$\frac{r}{a} = 1 + \frac{e^2}{2} - e \cdot \cos nt - \frac{e^2}{2} \cos 2nt$$

$$- \frac{e^3}{2 \cdot 4} \left\{ 3 \cdot \cos 3nt - 3 \cdot \cos nt \right\} \quad (28)$$

$$- \frac{e^4}{2 \cdot 4 \cdot 6} \left\{ 4^2 \cdot \cos 3nt - 4 \cdot 2^2 \cdot \cos 2nt \right\}$$

$$- \frac{e^5}{2 \cdot 4 \cdot 6 \cdot 8} \left\{ 5^3 \cdot \cos 5nt - 5 \cdot 3^3 \cdot \cos 3nt + \frac{5 \cdot 4}{1 \cdot 2} \cdot \cos nt \right\}$$

- &c.;

$$\theta = v + 2e \cdot \sin v + \frac{2e^2}{2} \cdot \sin 2v + \frac{2e^3}{3} \cdot \sin 3v + \&c.; \quad (29)$$

where $e = \frac{e}{1 + \sqrt{1 - e^2}}$ and is, of course, a fraction smaller than e . Also, if we neglect powers of e higher than the fifth,

$$\begin{aligned}\theta &= nt + \left\{ 2e - \frac{1}{4}e^3 + \frac{5}{96}e^5 \right\} \cdot \sin nt + \left\{ \frac{5}{4}e^2 - \frac{11}{24}e^4 \right\} \sin 2nt \\ &+ \left\{ \frac{13}{12}e^3 - \frac{43}{64}e^5 \right\} \sin 3nt + \frac{103}{96}e^4 \cdot \sin 4nt + \frac{1097}{960}e^5 \cdot \sin 5nt \end{aligned} \quad (30)$$

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SECTION IV.

On the velocities of the planetary motions, and the determination a priori of the elements of their orbits.

The angular velocity of a body is measured by the angle which it appears to describe in any very small time to the eye of a spectator. In fact, if we call θ the angle, and t the time, we have

$$\text{Angular velocity about the sun} = \frac{d\theta}{dt}$$

Now, by our 9th equation, we have $r^2 d\theta = h dt$, and, consequently, $\frac{d\theta}{dt} = \frac{h}{r^2}$; that is to say, the angular velocity in any orbit is inversely as the square of its distance from the centre; and this law is general for all central forces.

The paracentric velocity is the approach to, or recess from, the centre; and is measured by $\frac{dr}{dt}$.

Now, $\frac{dr}{dt} = \frac{dr}{d\theta} \cdot \frac{d\theta}{dt}$, but

$$\frac{dr}{d\theta} = \frac{a(1-e^2) \cdot e \sin \theta}{(1+e \cos \theta)^2} = \frac{e \sin \theta}{a(1-e^2) \cdot r^2}$$

and, consequently,

$$\frac{dr}{dt} = \frac{he \sin \theta}{a(1-e^2)r^4} = \sqrt{\frac{M+m}{a(1-e^2)}} \cdot \frac{e \sin \theta}{r^4}; \quad (31)$$

To complete our knowledge of the body's motion, we must inquire its linear velocity at any instant.

To this end, if we call V the velocity, $V = \frac{ds}{dt}$, ds being the element of the curve. Now, first, we must remark, that if we write for dt its value $\frac{y dx - x dy}{h}$,

$$\text{this gives} \quad V = h \cdot \frac{ds}{y dx - x dy}; \quad (32)$$

But $\frac{y dx - x dy}{ds}$ expresses the length of a perpendicular dropped, or a tangent to the curve from the origin of the co-ordinates. Thus we see, that the velocity is inversely as the perpendicular so let fall, and directly as the quantity h , or the area described in a given time.

Moreover, we have $ds^2 = dr^2 + r^2 d\theta^2$; so that,

writing for dr its equal $-\frac{du}{u^2}$, or $-r^2 du$,

$$\left(\frac{ds}{dt}\right)^2 = r^2 \left(\frac{d\theta}{dt}\right)^2 \left\{ 1 + r^2 \cdot \left(\frac{du}{d\theta}\right)^2 \right\}$$

Now, by reason of the equation $r^2 d\theta = h dt$, we have

$$r^2 \left(\frac{d\theta}{dt}\right)^2 = \frac{h^2}{r^2}$$

Again, if we differential the equation (12), we find

$$\begin{aligned} \left(\frac{du}{d\theta}\right)^2 &= f^2 \cdot \sin^2(\theta + g) \\ &= f^2 - f^2 \cdot \cos^2(\theta + g) \\ &= f^2 - \left(u - \frac{M+m}{h^2}\right)^2, \quad (\text{by equation 12}) \\ &= \frac{e^2}{p^2} - \left(u - \frac{1}{p}\right)^2, \quad (\text{by equations 15, 16}) \\ &= \frac{e^2 - 1}{p^2} + \frac{2}{pr} - \frac{1}{r^2} \end{aligned}$$

writing for u its value $\frac{1}{r}$. Consequently, we find

$$\begin{aligned} 1 + r^2 \left(\frac{du}{d\theta}\right)^2 &= \left(\frac{2}{pr} - \frac{1-e^2}{p^2}\right) \cdot r^2 \\ &= \frac{r^2}{p} \left(\frac{2}{r} - \frac{1}{a}\right) \end{aligned}$$

because $p = a(1-e^2)$. Hence, we finally obtain

$$\left(\frac{ds}{dt}\right)^2 = \frac{h^2}{p} \left(\frac{2}{r} - \frac{1}{a}\right)$$

$$\text{or} \quad V^2 = (M+m) \cdot \left(\frac{2}{r} - \frac{1}{a}\right); \quad (33)$$

On the determination of the elements of the planetary orbits, *a priori*, there is no occasion to enter into any very extensive discussion, in a practical point of view. Since, however, it is a subject generally touched upon in astronomical works, and is not without its interest, when we consider what changes in our own system may have taken place, or may yet take place from the action of violent causes, we shall devote a part of our space to its consideration.

Let us suppose, then, a body of a given mass, to be launched in space from a given point, with a velocity and direction also given; and to be attracted by another body, whose mass is also given with a force inversely as the square of the distance: it is required to determine the form, magnitude, and position of the conic section it will describe.

The plane passing through the attracting body and the primitive direction of projection will, of course, be that in which the orbit will lie, there being no force to draw the body out of this plane. Taking then the central body for the origin of two co-ordinates, x and y , lying in this plane, and retaining all the other denominations of the foregoing pages, and considering, as the unit of velocity, that with which a body would describe a space equal to 32.1908 feet in one second, we shall have, by (29),

$$v^2 = (M+m) \left(\frac{2}{r} - \frac{1}{a}\right)$$

and, consequently,

$$\frac{1}{a} = \frac{2}{r} - \frac{v^2}{M+m} \quad (34)$$

Now, by hypothesis, the masses of the central and revolving body are given; and the distance from the centre, as well as the velocity with which the latter is projected. If, then, we suppose the quantities in this equation to correspond to the point of projection, r is this distance, and V, M, m , are known; so that we have at once the value of a , the semiaxis major of the orbit.

This result is a remarkable one. It shews us, that the major axis of the orbit is independent of the angle of inclination to the radius at which the original projection takes place: in other words, that any number of bodies (of equal, or of exceedingly small magnitudes compared with the central one,) launched from one point with equal velocities, but in any different directions, will all describe orbits having equal major axes. Another result, not less curious, follows from this—that they will all describe conic sections of the same nature, that is, all ellipses, or all parabolas, or all hyperbolas—for the nature of the conic section depends only on the algebraic sign of its major axis.

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ellipse. If equal, a will be infinite ($\frac{1}{a} = 0$), and the orbit will be a parabola; but if greater, a is negative, and the orbit will be an hyperbola

In the two latter cases, the bodies will never return to their original point of departure, but in the former they will do so: and since the periodic time depends only on the major axes and masses, if their masses be either all equal, or all extremely small, they will all return in the same time, and a collision will take place; after which it is impossible to say what will happen. If, on the other hand, there exists any sensible proportion between the revolving and central bodies, and any considerable inequality in their masses, their periods will be unequal, and each may perform its orbit undisturbed.

This is supposing their mutual attractions to be neglected. In fact, however, at the instant of their departure, these may be incomparably greater than that of the central body, and will then materially change their velocities, unless the latter be so great as speedily to carry them beyond the sphere of their mutual influence. If a small portion of the earth, for instance, were suddenly projected from its surface, the attraction of the earth on it would, at the moment of its departure, exceed that of the sun in the ratio of 3522 : 2.17399, or upwards of 1600 : 1. So that, in the first instants of its motion, it would move as if influenced by the earth alone. But this effect would diminish rapidly; by the time the projectile had reached the distance of the moon, the sun's action would already have a preponderance (as we have seen in section I,) in the ratio of 2.17399 : 1, and it would depend entirely on the relative velocity of projection, whether such a space could be described in a time small enough to escape the influence of the earth or not.

It has been a matter of some speculation, whether the small planets between Mars and Jupiter may not have had their origin in the destruction, by violence, of some larger mass once revolving in the situation they now occupy. The very considerable approximation of their periodic times, which, in the case of Ceres and Pallas is singularly near (within $\frac{1}{100}$ of the whole period,) and the equally remarkable fact of the mutual intersections of their orbits falling all in the same part of the heavens, (in a general way,) have given rise to this surmise; and it has even been conjectured, that an explosive rupture of a former planet may have scattered its fragments far and wide over our system, and produced these singular bodies. There is no limit to conjecture; but if any such event have taken place, we are forced to conclude, that the mass of the ruptured planet must have been very small, or the fragments must have collapsed by their mutual attraction; or, at least, their velocities would have been so materially modified by it, as to obliterate all traces of their

once having had a common velocity about the sun. **Physical Astronomy** The smallness of the ruptured mass renders the supposition of an explosion less revolting; and we know, at least from observation, that the fragments (if such) are extremely minute.

If the orbit be a circle, we have $r = a$, and $\frac{2}{r} - \frac{1}{a} = \frac{1}{r}$; so that, for the velocity in a circle,

we have $V^2 = \frac{M+m}{r}$, $V = \frac{\sqrt{M+m}}{\sqrt{r}}$; that is to say,

the velocity in different circles is in the sub-duplicate ratio of the sum of the masses, or the absolute force, as it is sometimes called, directly, and of the radii inversely. Moreover, if we denote by V the velocity

in a circle of the radius r , or $V = \sqrt{\frac{M+m}{r}}$, we have

$$\frac{v^2}{V^2} = \frac{\frac{2}{r} - \frac{1}{a}}{\frac{1}{r}} = 2 - \frac{r}{a} = \frac{2a-r}{a}$$

or $v^2 : V^2 :: 2a - r : a$.

Now, if APM be an ellipse, S, H, the foci, AM = $2a$ and SP = r , we have HP = $2a - r$; so that

$$v : V :: \sqrt{HP} : \sqrt{AR}$$

by which property the velocity in a conic section may be immediately compared with that in a circle at the same distance.

Hence, when SP = AC, or at the extremities of the conjugate axis, the velocity is equal to that in a circle.

In the parabola, we have $\frac{v^2}{V^2} = 2$, or $v = V \cdot \sqrt{2}$, so that in this curve the velocity bears a constant ratio to that in a circle at the same distance, $\sqrt{2} : 1$.

In the hyperbola, HP increases without limit, and the velocity bears continually a greater and greater ratio to that in a circle.

As the velocity depends only on the distance and major axis of the conic section, and not at all on its form, we may conceive the conjugate axis so diminished that the conic section shall pass into a straight line. In this case, the extremity of the axis will coincide with the focus, and the velocity at any distance r will be that acquired by falling from a distance $2a$ from the centre to the distance r . The expression for this velocity is therefore still the same with that in the conic section. Hence, "The velocity in a conic section at any point is that which would be acquired by falling freely towards the centre, from a distance equal to the longer axis, to that point." In the parabola the longer axis is infinite, and the velocity at any point is, therefore, that acquired by falling from an infinite distance. In the hyperbola the axis is negative, and even an infinite fall is not sufficient to give a body all the velocity requisite for the description of this curve.

We have then the following expressions:—

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Velocity in a conic section, semiaxis = a , distance = r ; $v = \sqrt{(M+m) \left(\frac{2}{r} - \frac{1}{a} \right)}$

$$\left. \begin{aligned} \text{--- circle, whose radius is } r, \dots\dots\dots V &= \sqrt{\frac{M+m}{r}} \\ \text{--- parabola at any distance } r, \dots\dots\dots v &= \sqrt{\frac{2(M+m)}{r}} \end{aligned} \right\} \quad (35)$$

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Let us now consider the effect on the form of the conic section and on the position of its major axis, arising from a change not in the velocity, but in the *angle of projection*, i. e. the angle made with the radius vector by the direction in which the body is projected.

Fig. 7.

APM being any conic section, and S, H, its foci, SY, SZ, perpendiculars on a tangent at P, the angle of projection SPy, which we will call A, is equal to HPZ, and therefore

$$SY = SP \cdot \sin A, \text{ and } HZ = HP \cdot \sin A.$$

consequently,

$$SY \times HZ = SP \cdot HP \cdot \sin^2 A;$$

but, by the property of the conic sections, $SY \times HZ = CD^2$; and therefore $CD^2 = SP \cdot HP \cdot \sin^2 A$. Now, we have $SP = r$, $HP = 2a - r$, $CD^2 = a^2 (1 - e^2)$;

$$\text{hence, } a^2 (1 - e^2) = r (2a - r) \cdot \sin^2 A$$

$$\text{and } a (1 - e^2) = r \times \left(2 - \frac{r}{a} \right) \cdot \sin^2 A.$$

But $a (1 - e^2)$ is the semiparameter of the conic section. Moreover, calling v the velocity in the curve, and V that in a circle; and denoting by n the ratio $\left(\frac{v}{V} \right)^2$ we have already seen, that $\left(2 - \frac{r}{a} \right) = n^2$.

Hence, we have

$$a (1 - e^2) = r \times n^2 \cdot \sin^2 A; \quad (36)$$

and therefore is given, when the angle and distance of projection are given, and also the velocity. When the angle of projection varies, other circumstances remaining, we see hence that the parameter varies as the square of its sine.

The eccentricity is easily found; for we have

$$e = \sqrt{1 - \frac{r}{a} \left(2 - \frac{r}{a} \right) \cdot \sin^2 A}$$

$$\begin{aligned} \text{Now } \frac{r}{a} &= 2 - \frac{r v^2}{M + m}; \text{ but since } v^2 = n^2 V^2 \\ &= n^2 \cdot \frac{M + m}{r}, \text{ we have } \frac{r v^2}{M + m} = n^2, \text{ and } \frac{r}{a} = 2 - n^2; \end{aligned}$$

$$2 - \frac{r}{a} = n^2. \text{ so that we get by substitution}$$

$$\begin{aligned} e &= \sqrt{1 - n^2 (2 - n^2) \cdot \sin^2 A} \\ &= \sqrt{\cos^2 A + \sin^2 A (1 - n^2)^2}; \end{aligned} \quad (37)$$

We see, therefore, that the ratio e of the eccentricity to the semiaxis, or the figure of the ellipse, or hyperbola, depends solely on the angle of projection and the ratio of the velocity of projection to that in a circle at the same distance; and if this latter ratio remain the same, the distance may be varied to any extent without changing the figure of the conic section.

It only remains to determine its position, or the angle made by the greater axis (or line of apsides) with the distance SP.

Now, if we call ASP, θ , we have

$$r = \frac{a (1 - e^2)}{1 + e \cdot \cos \theta};$$

$$\cos \theta = \frac{1}{e} \times \left\{ \frac{a (1 - e^2)}{r} - 1 \right\}$$

in which, substituting for e and $a (1 - e^2)$, their values before found, we shall obtain

$$\cos \theta = \frac{n^2 \cdot \sin^2 A - 1}{\sqrt{\cos^2 A + (1 - n^2)^2 \cdot \sin^2 A}} \quad (38)$$

on which value we may make the same remark as on that of e , and in both which it will be recollected,

$$\text{that } n^2 = \frac{r v^2}{M + m}.$$

SECTION V.

On certain peculiar cases of the celestial motion; viz. when the orbit is of very great, or very small eccentricity. Of circular and parabolic motion.

The circumstances of circular motion are too simple to need much consideration. The velocity is uniform and equal to that which would be acquired by falling down half the radius with the force at the circumference continued uniformly. This is evident, if we recollect that the velocity in a circle, by what

has been just proved, is $\sqrt{\frac{M+m}{r}}$, but that acquired by the action of a constant force F , acting through a space $\frac{1}{2} r$, is given by the equation $v^2 = 2 F s = F r$ in the present case $= \frac{(M+m)r}{r^2} = \frac{M+m}{r}$. The periodical time is equal to the circumference divided by this constant velocity, or to $\frac{2\pi r^{\frac{3}{2}}}{\sqrt{M+m}}$, as we have before shewn by a more general reasoning.

In a conic section of so small an eccentricity that its square may be neglected, it may be worth while to recapitulate some of the chief formulæ, developed to the first power. We have then

$$\left. \begin{aligned} r &= a (1 - e \cdot \cos \theta) \\ &= a (1 - e \cdot \cos v) \\ &= a (1 - e \cdot \cos nt) \end{aligned} \right\}; \quad (39)$$

$$\left. \begin{aligned} \theta &= v - e \cdot \sin v \\ \theta &= nt + 2e \cdot \sin nt \end{aligned} \right\}; \quad (40)$$

$$v = nt + e \cdot \sin nt; \quad (41)$$

$$nt = \theta - 2e \cdot \sin \theta; \quad (42)$$

Let us consider the motion in a parabola, and since in this case we have $a = \infty$, $e = 1$, and $p = a (1 - e^2)$, if we call the perihelion distance D , we have $D = \frac{1}{2} p$, and $p = 2 D$. Then will the equations (17) and (18) become

$$r = \frac{2 D}{1 + \cos \theta} = \frac{D}{\left(\cos \frac{\theta}{2} \right)^2}; \quad (43)$$

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$$dt = \frac{(2D)^{\frac{1}{2}}}{\sqrt{M+m}} \cdot \int \frac{d\theta}{(1 + \cos \theta)^{\frac{3}{2}}}$$

To integrate the latter, put $\tan \frac{\theta}{2} = x$, and we have

$$\begin{aligned} \tan \theta &= \frac{2x}{1-x^2} \\ \cos \theta &= \frac{1-x^2}{1+x^2} \\ d\theta &= \frac{2dx}{1+x^2} \end{aligned}$$

which substituted, give

$$\begin{aligned} dt &= \frac{(2D)^{\frac{1}{2}}}{\sqrt{M+m}} \times \left(\frac{dx}{2} + \frac{x^2 dx}{2} \right) \\ t &= \frac{D^{\frac{1}{2}} \cdot \sqrt{2}}{\sqrt{M+m}} \left(x + \frac{x^3}{3} \right) \end{aligned}$$

or,

$$t = \frac{D^{\frac{1}{2}} \cdot \sqrt{2}}{\sqrt{M+m}} \times \left\{ \tan \frac{\theta}{2} + \frac{1}{3} \left(\tan \frac{\theta}{2} \right)^3 \right\}; \quad (44)$$

This equation gives $\tan \frac{\theta}{2}$, and consequently θ , by

the resolution of a cubic equation of the form

$$x^3 + 3x = A$$

where $A = \frac{3t \cdot \sqrt{M+m}}{D^{\frac{1}{2}} \cdot \sqrt{2}}$. This cubic, it is easily

shewn, has but one real root; and since the co-efficient of x is an absolute number, and therefore x a function of A only, its root may be found in any proposed case by a table of single entry. In fact, suppose we have formed such a table, containing the values of x (or, for greater convenience, of $2 \times$ arc ($\tan = x$) or θ) for every value of t on the supposition of $D = 1$, then will this table serve for all cases, and for every other value of D ; for we have only to multiply t by $D^{-\frac{1}{2}}$; and calling the product T , look out in the table for the value of θ corresponding to the time T . A comet, describing a parabola, whose perihelion distance is 1, will describe 90° of anomaly from the perihelion in about 109 days. Hence, the use of the *table of a comet of 109 days* given in works on astronomy, Lalande's *Astron.* 2d edit. vol. iii. p. 335; Delambre's *Astronomy*, vol. iii. p. 434, from which we extract it as subjoined.

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Table of a comet of 109 days.

Days	True anomaly.	Days	True anomaly.	Days	True anomaly.	Days	True anomaly.	Days	True anomaly.	Days	True anomaly.
0	0° 0' 0"0	35	43°58'34"8	70	71°51'23"2	105	88°20'42"9	140	98°56'22"5	175	106°20'14"2
1	1 23 37.4	36	44 59 57.5	71	72 27 4.4	106	88 42 42.6	141	99 11 12.9	176	106 30 59.4
2	2 47 11.9	37	46 0 26.6	72	73 2 13.4	107	89 4 26.0	142	99 25 54.3	177	106 41 39.4
3	4 10 40.4	38	47 0 2.2	73	73 36 51.0	108	89 25 53.4	143	99 40 26.9	178	106 52 14.1
4	5 34 0.1	39	47 58 44.8	74	74 10 57.8	109	89 47 5.1	144	99 54 50.9	179	107 2 43.5
5	6 57 8.0	40	48 56 34.5	75	74 44 34.3	110	90 8 1.3	145	100 9 6.4	180	107 13 7.7
6	8 20 1.3	41	49 53 32.0	76	75 17 41.2	111	90 28 42.4	146	100 23 13.4	181	107 23 26.9
7	9 42 37.2	42	50 49 37.6	77	75 50 18.9	112	90 49 8.7	147	100 37 12.2	182	107 33 41.1
8	11 4 53.0	43	51 44 51.8	78	76 22 28.1	113	91 9 20.3	148	100 51 2.8	183	107 43 50.2
9	12 26 45.9	44	52 39 15.2	79	76 54 9.3	114	91 29 17.6	149	101 4 45.4	184	107 53 54.6
10	13 48 13.4	45	53 32 48.2	80	77 25 23.1	115	91 49 0.7	150	101 18 20.1	185	108 3 54.1
11	15 9 13.1	46	54 25 31.6	81	77 56 10.1	116	92 8 30.0	151	101 31 47.1	186	108 13 48.8
12	16 29 42.5	47	55 17 25.8	82	78 26 30.6	117	92 27 45.7	152	101 45 6.5	187	108 23 38.8
13	17 49 39.4	48	56 8 31.5	83	78 56 25.3	118	92 46 48.0	153	101 58 18.2	188	108 33 24.2
14	19 9 1.5	49	56 58 49.4	84	79 25 54.6	119	93 5 37.3	154	102 11 22.6	189	108 43 5.0
15	20 27 46.8	50	57 48 20.1	85	79 54 59.1	120	93 24 13.6	155	102 24 19.6	190	108 52 41.3
16	21 45 53.4	51	58 37 4.3	86	80 23 39.1	121	93 42 37.2	156	102 37 9.4	191	109 2 13.1
17	23 3 19.4	52	59 25 2.7	87	80 51 55.4	122	94 0 48.4	157	102 49 52.1	192	109 11 40.4
18	24 20 3.1	53	60 12 15.9	88	81 19 48.1	123	94 18 47.3	158	103 2 27.8	193	109 21 3.5
19	25 36 2.9	54	60 58 44.8	89	81 47 17.9	124	94 36 34.2	159	103 14 56.6	194	109 30 22.2
20	26 51 17.3	55	61 44 29.9	90	82 14 25.2	125	94 54 9.2	160	103 27 18.5	195	109 39 36.7
21	28 5 45.1	56	62 29 32.1	91	82 41 10.3	126	95 11 32.7	161	103 39 33.8	196	109 48 46.9
22	29 19 25.0	57	63 13 52.0	92	83 7 33.7	127	95 28 44.7	162	103 51 42.4	197	109 57 53.1
23	30 32 15.8	58	63 57 30.3	93	83 33 35.9	128	95 45 45.4	163	104 3 44.6	198	110 6 55.1
24	31 44 16.7	59	64 40 27.7	94	83 59 17.2	129	96 2 35.2	164	104 15 40.3	199	110 15 53.1
25	32 55 26.7	60	65 22 45.0	95	84 24 38.0	130	96 19 14.0	165	104 27 29.7	200	110 24 47.1
26	34 5 45.2	61	66 4 22.8	96	84 49 38.7	131	96 35 42.2	166	104 39 12.8		
27	35 15 11.4	62	66 45 22.0	97	85 14 19.8	132	96 51 59.8	167	104 50 49.8		
28	36 23 44.8	63	67 25 43.1	98	85 38 41.6	133	97 8 7.1	168	105 2 20.7		
29	37 31 24.9	64	68 5 26.9	99	86 2 44.4	134	97 24 4.2	169	105 13 45.6		
30	38 38 11.5	65	68 44 34.0	100	86 26 28.7	135	97 39 51.3	170	105 25 4.7		
31	39 44 4.2	66	69 23 5.1	101	86 49 54.7	136	97 55 28.6	171	105 36 17.9		
32	40 49 2.9	67	70 1 1.0	102	87 13 2.8	137	98 10 56.2	172	105 47 25.4		
33	41 53 7.6	68	70 38 22.1	103	87 35 53.3	138	98 26 14.3	173	105 58 27.2		
34	42 56 18.2	69	71 15 9.3	104	87 58 26.5	139	98 41 23.0	174	106 9 23.4		
35	43 58 34.8	70	71 51 23.2	105	88 20 42.9	140	98 56 22.5	175	106 20 14.2		

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2dly. Let us now consider the circumstances of the motion of a body in an ellipse of great eccentricity, or approaching very near to a parabola. This is neces-

sary to complete the theory of such comets as are known, or may be suspected, to revolve in very elongated ellipses. Physical Astronomy.

Since e , the ratio of the eccentricity, approaches very near to unity, $1 - e$ must be a very small quantity; and if we put $\beta = \frac{1-e}{1+e}$, we shall have $e = \frac{1-\beta}{1+\beta}$, $1 - e = \frac{2\beta}{1+\beta}$, $1 + e = \frac{2}{1+\beta}$, $(1 - e^2) = \frac{4\beta}{(1+\beta)^2}$; now, if D = the perihelion distance, we have $D = a(1 - e)$, consequently

$$r = \frac{D(1+e)}{2 + e \cos \theta} = \frac{2D}{(1+\beta) + (1-\beta) \cos \theta} = \frac{D}{\frac{1 + \cos \theta}{2} + \beta \cdot \frac{1 - \cos \theta}{2}}$$

$$= \frac{D}{1 + \beta \cdot \frac{1 - \cos \theta}{1 + \cos \theta}} = \frac{D}{\left(\cos \frac{\theta}{2}\right)^2 \left(1 + \beta \cdot \left(\tan \frac{\theta}{2}\right)^2\right)}$$

and developing this in powers of β , we obtain

$$r = \frac{D}{\left(\cos \frac{\theta}{2}\right)^2} \left\{ 1 - \beta \cdot \left(\tan \frac{\theta}{2}\right)^2 + \beta^2 \cdot \left(\tan \frac{\theta}{2}\right)^4 - \&c. \right\}; \quad (45)$$

which is a very simple expression for the radius vector in the elongated ellipse.

To find the time, we must make the same substitution in the expression,

$$t + C = \frac{a^{\frac{3}{2}}(1 - e^2)^{\frac{3}{2}}}{\sqrt{M + m}} \cdot \int \frac{d\theta}{(1 + e \cos \theta)^2}$$

Now $\left\{ a \cdot (1 - e^2) \right\}^{\frac{3}{2}} = \left(\frac{2D}{1+\beta} \right)^{\frac{3}{2}}$ and since

$$\begin{aligned} (1 + e \cos \theta)^2 &= \frac{1}{(1+\beta)^2} \{ 1 + \beta + (1-\beta) \cos \theta \}^2 \\ &= \frac{4}{(1+\beta)^2} \left\{ \frac{1 + \cos \theta}{2} + \beta \cdot \frac{1 - \cos \theta}{2} \right\}^2 \\ &= \frac{4 \cdot \left(\cos \frac{\theta}{2}\right)^2 \left\{ 1 + \beta \cdot \left(\tan \frac{\theta}{2}\right)^2 \right\}^2}{(1+\beta)^2} \end{aligned}$$

if we put $\tan \frac{\theta}{2} = T$, we shall have $\frac{\frac{1}{2} d\theta}{\left(\cos \frac{\theta}{2}\right)^2} = dT$; and the integral will become $\frac{(1+\beta)^2}{2} \cdot \int \frac{dT}{(1 + \beta \cdot T^2)^2}$

whence we get after all substitution,

$$t + C = \sqrt{\frac{2(1+\beta)}{M+m}} \cdot D^{\frac{3}{2}} \cdot \int \frac{dT(1+T^2)}{(1+\beta \cdot T^2)^2}$$

and the part of this under the integral sign being developed in powers of β , and the whole integrated, we get

$$t + C = D^{\frac{3}{2}} \cdot \sqrt{\frac{2(1+\beta)}{M+m}} \times \left\{ \left(T + \frac{T^3}{3} \right) - \frac{2}{1} \beta \left(\frac{T^3}{3} + \frac{T^5}{5} \right) + \frac{2 \cdot 3}{1 \cdot 2} \beta^2 \left(\frac{T^5}{5} + \frac{T^7}{7} \right) - \&c. \right\} \quad (46)$$

By the aid of this series, we may express very readily the difference between the true anomaly in a very eccentric ellipse, and that in a parabola of equal perihelion distance, for, T still expressing the tangent of half the former true anomaly, and τ that of the latter, if we develop according to powers of β , pursuing the process only as far as β^3 , we get

$$t + C = \frac{D^{\frac{3}{2}} \sqrt{2}}{\sqrt{M+m}} \left\{ \left(T + \frac{T^3}{3} \right) + \beta \left(\frac{T}{2} - \frac{T^3}{2} - \frac{2T^5}{5} \right) - \beta^2 \left(\frac{T}{8} + \frac{3T^3}{8} - \frac{2T^5}{5} - \frac{3T^7}{7} \right) \right\}; \quad (47)$$

Now, the time being the same both in the parabola and ellipse, this must equal $\frac{D^{\frac{3}{2}} \sqrt{2}}{\sqrt{M+m}} \left\{ \tau + \frac{\tau^3}{3} \right\}$, and

therefore the series within the brackets on both sides must be equal.

Let

$$\begin{aligned} \tau + \frac{\tau^3}{3} &= a; \quad \frac{\tau}{2} - \frac{\tau^3}{2} - \frac{2\tau^5}{5} = b; \quad -\frac{\tau}{8} - \frac{3\tau^3}{8} + \frac{2\tau^5}{5} + \frac{3\tau^7}{7} = c \\ T + \frac{T^3}{3} &= A, \quad \frac{T}{2} - \frac{T^3}{2} - \frac{2T^5}{5} = B; \quad -\frac{T}{8} - \frac{3T^3}{8} + \frac{2T^5}{5} + \frac{3T^7}{7} = C \end{aligned}$$

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$$A + B\beta + C\beta^2 = a;$$

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and the first member of this equation, though apparently a function of β , cannot be really so, as the second does not contain it. Hence, if we call Δ the difference between the elliptic and parabolic anomaly,

or suppose the former anomaly $\theta + \Delta$ the latter θ , since $T = \tan \frac{\theta + \Delta}{2}$, and $\tau = \tan \frac{\theta}{2}$, we may express A, B, and C, thus:—

$$A = a + \frac{da}{d\theta} \cdot \frac{\Delta}{1} + \frac{d^2a}{d\theta^2} \cdot \frac{\Delta^2}{1 \cdot 2} + \&c.$$

$$B = b + \frac{db}{d\theta} \cdot \frac{\Delta}{1} + \frac{d^2b}{d\theta^2} \cdot \frac{\Delta^2}{1 \cdot 2} + \&c.$$

$$C = c + \frac{dc}{d\theta} \cdot \frac{\Delta}{1} + \frac{d^2c}{d\theta^2} \cdot \frac{\Delta^2}{1 \cdot 2} + \&c.$$

but Δ itself is required to be expressed in terms of β and its powers; and, consequently we may suppose $\Delta = P\beta + Q\beta^2 + \&c.$

If then we make these substitutions, our equation becomes

$$\begin{aligned} a = a + P \frac{da}{d\theta} \beta + \left(Q \frac{da}{d\theta} + \frac{P^2}{2} \cdot \frac{d^2a}{d\theta^2} \right) \beta^2 + \&c. \\ + b \beta + P \frac{db}{d\theta} \cdot \beta^2 + \&c. \\ + c \beta^2 + \&c. \end{aligned}$$

whence we obtain by equating the co-efficients of β to zero,

$$0 = P \frac{da}{d\theta} + b; \quad 0 = Q \frac{da}{d\theta} + \frac{P^2}{2} \cdot \frac{d^2a}{d\theta^2} + P \cdot \frac{db}{d\theta} + c; \quad \&c.$$

Now we have $\frac{da}{d\theta} = \frac{da}{d\tau} \cdot \frac{d\tau}{d\theta}$, and since $a = \tau + \frac{\tau^3}{3}$, $\frac{da}{d\tau} = 1 + \tau^2$, moreover $\frac{d\tau}{d\theta} = \frac{1 + \tau^2}{2}$, whence $\frac{da}{d\theta} = \frac{(1 + \tau^2)^2}{2}$, and therefore $P = \frac{1}{(1 + \tau^2)^2} \left\{ -\tau + \tau^3 + \frac{4}{5} \tau^5 \right\}$

Again,

$$\begin{aligned} \frac{d^2a}{d\theta^2} &= \frac{d}{d\theta} \cdot \frac{da}{d\theta} = 2\tau(1 + \tau^2) \cdot \frac{d\tau}{d\theta} = \tau(1 + \tau^2)^2 \\ \frac{db}{d\theta} &= \frac{d}{d\theta} \left\{ \frac{\tau}{2} - \frac{\tau^3}{2} - \frac{2}{5} \tau^5 \right\} = \left\{ \frac{1}{2} - \frac{3\tau^2}{2} - 2\tau^4 \right\} \cdot \frac{1 + \tau^2}{2} \end{aligned}$$

whence, after all reductions, we obtain

$$Q = \frac{\frac{3}{4} \tau - \frac{5}{4} \tau^3 + \frac{1}{20} \tau^5 + \frac{167}{140} \tau^7 + \frac{24}{35} \tau^9 + \frac{18}{175} \tau^{11}}{(1 + \tau^2)^4}$$

and, finally,

$$\Delta = \frac{-\tau + \tau^3 + \frac{4}{5} \tau^5}{(1 + \tau^2)^2} \cdot \beta + \frac{\frac{3}{4} \tau - \frac{5}{4} \tau^3 + \frac{1}{20} \tau^5 + \frac{167}{140} \tau^7 + \frac{24}{35} \tau^9 + \frac{18}{175} \tau^{11}}{(1 + \tau^2)^4} \cdot \beta^2; \quad (48)$$

SECTION VI.

On the determination of the planetary orbits *a posteriori*, or from observation.

The determination of the elements of the orbit of a planet, or comet, is justly regarded as one of the most difficult problems of astronomy. By the *elements* of the orbit are meant those constant data which determine the position, form, and magnitude of the conic section described, and the body's place in it at a given epoch, and are six in number, viz.

The inclination of its plane to the ecliptic,

The longitude of its ascending node,

The position of its axis, or the longitude of the perihelion,

The length of its axis,

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The eccentricity, or the length of the semiparameter, or the perihelion distance,

The moment of passing the perihelion, or the heliocentric longitude at any assigned time,

which correspond to, and are functions of the six arbitrary constants introduced by the integration of the equations (1), (2), and (3). If we knew, and could subject to calculation, the causes which originally determined the motions of the heavenly bodies, we could assign, *a priori*, by the means already pointed out, the values of these constants from the conditions which subsisted at the commencement of their motions; but as these are, and must ever remain, unknown, we can only discover the values in question, *a posteriori*, by observation of the apparent places of each body at different times.

Astronomy. In this research, the great difficulties arise from our not observing from a fixed station in the centre of our system. To an observer stationed in the sun, the elements would offer themselves with comparative ease.

But the motion and eccentric position of the earth render the whole affair infinitely more complicated; and to analyse the problem, it becomes necessary to take into consideration the relations between the heliocentric and geocentric places of a heavenly body.

To this effect, let S be the sun, E the earth, P the planet, or comet, projected into p upon the ecliptic by the perpendicular Pp. Let S γ be the line of equinoxes, parallel to which draw E γ . Then will

$L = \gamma SE$ represent the earth's heliocentric longitude.
 $l = \gamma Sp$ — the planet's heliocentric longitude.
 $\lambda = \gamma Ep$ — the planet's geocentric longitude.
 $b = PSp$ — the planet's heliocentric latitude.
 $\beta = PEp$ — its geocentric latitude.

R = the earth's distance from the sun.

r = the planet's distance from the sun.

ρ = the planet's distance from the earth.

d, δ = the respective *curtate*, or projected distances of the planet from the sun and earth.

It is obvious, then, that we shall have the following relations:—

1st. $Pp = r \cdot \sin b$, also the same $Pp = \rho \cdot \sin \beta$, so that,

$$r \cdot \sin b = \rho \cdot \sin \beta; \quad (49)$$

$$2dly. \begin{cases} d = r \cdot \cos b \\ \delta = \rho \cdot \cos \beta \end{cases} \quad (50) \quad (51)$$

3dly. Since the angle PSE of the triangle PSE is equal to $l - L$, and its sides are R, d , and δ

$$\delta^2 = R^2 + d^2 - 2Rd \cos(l - L); \quad (52)$$

Lastly producing SE to C, the angle PEC = $pE\gamma - CS\gamma = \lambda - L$, but the angle $pEc = 180^\circ - pES$, consequently the equation

$$d^2 = R^2 + \delta^2 - 2R\delta \cdot \cos pES$$

gives

$$d^2 = R^2 + \delta^2 + 2R\delta \cdot \cos(\lambda - L); \quad (53)$$

Thus then we have five distinct and independent relations among the ten quantities $L, l, \lambda, \&c.$ Now, it is evident, that if R and L be given, (or the place of E fixed) the geocentric latitude and longitude, β and λ , will be determined when the place of P is, but this is not determined unless l, d , and b , are so. But when all the five quantities L, R, l, d, b , are given, the five equations above deduced suffice to determine the others, and therefore express all the relations subsisting between the heliocentric and geocentric places. The last may, however, be simplified by substituting for δ^2 its value given in (52), when the whole becomes divisible by $2R$, and gives

$$o = R + \delta \cdot \cos(\lambda - L) - d \cdot \cos(l - L); \quad (54)$$

Another equation (included, of course, in the foregoing,) may also be obtained if we consider, that in the triangle PSE, we have

$$r^2 = R^2 + \rho^2 - 2R\rho \cdot \cos PES$$

but $\cos PES = \cos pES \times \cos PEp$ by spherical trigonometry, the plane PEp being at right angles to

pES. Hence, we get (writing $\frac{\rho'}{\cos \beta}$), or $\rho' \cdot \sec \beta$ for

$$\rho, \text{ and } -\cos(\lambda - L) \text{ for } \cos pES$$

$$r^2 = R^2 + \delta^2 \cdot \sec \beta^2 + 2R\delta \cdot \cos(\lambda - L); \quad (55)$$

In like manner, eliminating ρ from (49), it becomes

$$r \cdot \sin b = \delta \cdot \tan \beta; \quad (56)$$

and putting in (54) for d , its value $r \cdot \cos b$,

$$o = R + \delta \cdot \cos(\lambda - L) - r \cdot \cos b \cdot \cos(l - L) \quad (57)$$

This equation may also be derived independently, if we drop the perpendicular p q , for we then have

$$Sq = SP \cdot \cos PSp \cdot \cos pSq$$

$$= r \cdot \cos b \cdot \cos(l - L)$$

and also

$$Sq = SE + Eq =$$

$$= R + \delta \cdot \cos(\lambda - L)$$

which equated, give the equation in question. We may also derive another, which will be useful in the sequel, by equating two values of the perpendicular itself.

$$pq = Ep \cdot \sin pEq$$

$$pq = Sp \cdot \sin pSq$$

which equated, give

$$r \cdot \cos b \cdot \sin(l - L) = \delta \cdot \sin(\lambda - L); \quad (58)$$

Thus we have reduced the five equations to four, 55, 56, 57, 58, which contain only the radius vector of the earth, that of the planet, and its curtate distance from the earth, and the heliocentric and geocentric latitudes and longitudes of the two bodies.

The problem of determining the elements of P's orbit from observation requires these relations to be combined with the conditions of P's motion and of the earth's. Now, if we call Ω the longitude of P's ascending node, and i the inclination of its orbit, we have, by spherical trigonometry, for the heliocentric latitude.

$$\tan b = \tan i \cdot \sin(l - \Omega) \quad (59)$$

Moreover, if we call π the longitude of the perihelion on the orbit, θ the true anomaly, and ψ the angle ΩSA , or the distance of the node from the perihelion on the orbit, (see fig. 9) we shall have

$$\cos \Omega a = \tan \Omega Sa \cdot \cot \Omega A$$

$$\text{or } \tan(\pi - \Omega) = \cos i \cdot \tan \psi \quad (60)$$

and also

$$\sin b = \sin i \cdot \sin(\psi + \theta) \quad (61)$$

these seven equations, of which the four first, 55, 56, 57, 58, express the relations between the heliocentric and geocentric places of a body any how situated, the other three, 59, 60, 61, merely express the condition of every point in the orbit lying in a plane passing through the sun. These must be combined with the dynamical results peculiar to the planetary motions, viz. 1st. The proportionality of the areas to the times; and, 2dly. The nature of the orbit, regarded as an ellipse or parabola of unknown position, form, and magnitude; all which are included in the three equations, 21, 22, 23, or others equivalent to them. Thus we have a system of nine equations, representing, in general, the relations between the time elapsed since a given epoch, and the observed geocentric latitude and longitude.

These equations involve, as constant quantities, the elements of the orbit, i, Ω, π, a, e, E , while the variable ones, being either given by observation, or expressed in functions of others that are so, the whole system of equations may be regarded as expressing a known (though a very complicated) relation among the elements; and as each observation affords a similar

Fig. 8.

Fig. 9.

Astronomy. relation, it is evident that the complete determination of the latter cannot require more than six observations. In fact, however, from the peculiar nature of some of the equations, fewer will suffice, half that number only being necessary.

The complication of the relations in question precludes all idea of a direct solution of the problem, and to apply them to any particular case indirect and approximative ones have been invented, but with every assistance from such simplification; and, after all the force of analysis has been exercised upon it, it still remains a very difficult problem, and one which our limits will by no means permit us to enter upon in its full extent. In fact, the case of an elliptical orbit to be determined is one of very rare occurrence; and we shall therefore content ourselves with pointing out the course to be pursued to obtain the readiest approximation to the elements of parabolic motion. The frequent visits of comets to our system render this an important case. These singular bodies, with one or two exceptions, have always been found to describe parabolas, or ellipses of such extreme eccentricity, as to be undistinguishable from parabolas within considerable distances on either side their perihelion.

Preparatory to this research, we must premise the following remarkable theorem.

In a parabola, if r, r' denote any two distances from the focus, k the chord of the arc they include, and t the time of a body describing that arc, and g a certain constant the same for all parabolas, we shall have

$$gt = (r + r' + k)^{\frac{3}{2}} - (r + r' - k)^{\frac{3}{2}}; \quad (62)$$

To shew this, let θ and θ' be the true anomalies corresponding to the radii r, r' , and t being the time of describing the intermediate arc, and ϵ the time elapsed since the perihelion passage to the commencement of the time t , we have, by (44)

$$\begin{aligned} \epsilon &= q \left\{ \tan \frac{\theta}{2} + \frac{1}{3} \left(\tan \frac{\theta}{2} \right)^3 \right\} \\ t + \epsilon &= q \left\{ \tan \frac{\theta'}{2} + \frac{1}{3} \left(\tan \frac{\theta'}{2} \right)^3 \right\} \end{aligned}$$

where $q = \frac{D^3 \sqrt{2}}{\sqrt{M+m}}$. Consequently, we have

$$\begin{aligned} t &= q \cdot \left\{ \left(\tan \frac{\theta'}{2} - \tan \frac{\theta}{2} \right) \right. \\ &\quad \left. + \frac{1}{3} \left(\left(\tan \frac{\theta'}{2} \right)^3 - \left(\tan \frac{\theta}{2} \right)^3 \right) \right\} \end{aligned}$$

Now, by the property of the parabola, we have also

$$r = \frac{D}{\left(\cos \frac{\theta}{2} \right)^2}, \quad r' = \frac{D}{\left(\cos \frac{\theta'}{2} \right)^2}$$

And if we put $\frac{\theta' - \theta}{2} = \phi$, we shall have

$$\sqrt{\frac{r}{r'}} = \frac{r}{\sqrt{r r'}} = \frac{\cos \frac{1}{2} \theta'}{\cos \frac{1}{2} \theta} = \frac{\cos \left(\frac{1}{2} \theta + \phi \right)}{\cos \frac{1}{2} \theta}$$

$$\text{or, } \frac{r}{\sqrt{r r'}} = \cos \phi - \tan \frac{1}{2} \theta \cdot \sin \phi$$

whence we get

$$\tan \frac{\theta}{2} = \operatorname{cosec} \phi \left\{ \cos \phi - \frac{r}{\sqrt{r r'}} \right\} \quad (63)$$

Similarly, since $\frac{1}{2} \theta = \frac{1}{2} \theta' - \phi$, we have,

$$\sqrt{\frac{r'}{r}} \text{ or } \frac{r'}{\sqrt{r r'}} = \frac{\cos \left(\frac{1}{2} \theta' - \phi \right)}{\cos \frac{1}{2} \theta'}$$

and therefore

$$\tan \frac{\theta'}{2} = -\operatorname{cosec} \phi \left\{ \cos \phi - \frac{r'}{\sqrt{r r'}} \right\}$$

but k being the side of a triangle opposite to the angle $(\theta' - \theta)$ or 2ϕ and r and r' the sides adjacent, we have, by trigonometry,

$$\cos \phi = \frac{\sqrt{(r+r')^2 - k^2}}{2 \sqrt{r r'}}$$

$$\sin \phi = \frac{\sqrt{k^2 - (r-r')^2}}{2 \sqrt{r r'}}$$

or, if we put

$$\sqrt{(r+r')^2 - k^2} = R \text{ and } \sqrt{k^2 - (r-r')^2} = S$$

$$R = 2 \sqrt{r r'} \cdot \cos \phi \text{ and } S = 2 \sqrt{r r'} \cdot \sin \phi$$

$$\cos \phi = \frac{R}{2 \sqrt{r r'}} \quad \sin \phi = \frac{S}{2 \sqrt{r r'}}$$

Hence, we obtain

$$\tan \frac{\theta}{2} = \frac{R-2r}{S}; \quad \tan \frac{\theta'}{2} = -\frac{R-2r'}{S}$$

$$\tan \frac{\theta'}{2} - \tan \frac{\theta}{2} = \frac{2(r+r'-R)}{S}$$

also

$$\begin{aligned} \left(\cos \frac{\theta}{2} \right)^2 &= \frac{1}{1 + \left(\tan \frac{\theta}{2} \right)^2} = \\ &= \frac{S^2}{(R^2 + S^2) - 4rR + 4r^2} \end{aligned}$$

but $R^2 + S^2 = (r+r')^2 - (r'-r)^2 = 4rr'$, so that we get

$$\left(\cos \frac{\theta}{2} \right)^2 = \frac{1}{4r} \cdot \frac{S^2}{r+r'-R}$$

and, in consequence,

$$D = r \cdot \left(\cos \frac{\theta}{2} \right)^2$$

$$= \frac{1}{4} \cdot \frac{S^2}{r+r'-R}$$

Now $(r+r'-R)(r+r'+R) = (r+r')^2 - R^2 = k^2$; so that

$$D = \frac{S^2}{4k^2} \cdot (r+r'+R)$$

and

$$D \left\{ \tan \frac{\theta'}{2} - \tan \frac{\theta}{2} \right\} = \frac{S}{2}$$

consequently

$$\frac{D^{\frac{3}{2}} \cdot \sqrt{2}}{\sqrt{M+m}} \left\{ \tan \frac{\theta'}{2} - \tan \frac{\theta}{2} \right\} = \frac{S^2 \cdot \sqrt{2}(r+r'+R)}{4k \cdot \sqrt{M+m}};$$

Now, if we put T and T' for $\tan \frac{\theta}{2}$ and $\tan \frac{\theta'}{2}$, we have

$$T'^3 - T^3 = (T' - T) \{ T'^2 + TT' + T^2 \}$$

and therefore

$$t = \frac{D^{\frac{3}{2}} \sqrt{2}}{\sqrt{M+m}} (T' - T) \left\{ 1 + \frac{T^2 + TT' + T'^2}{3} \right\}$$

The part without the brackets we have already considered. Let us next examine that within. Now, if in $T^2 + TT' + T'^2$, we substitute for T and T' , their

Astronomy. values $\frac{R-2r}{S}$ and $\frac{-R+2r'}{S}$, it will become

$$\frac{R^2 - 2R(r+r') + 4(r^2 - rr' + r'^2)}{S^2}$$

and substituting this, the quantity within the brackets becomes

$$\frac{3S^2 + R^2 + 4(r^2 - rr' + r'^2) - 2R(r+r')}{3S^2}$$

But $R^2 + S^2 = 4rr'$, therefore $3S^2 + R^2 = 12rr' - 2R^2 = 2k^2 - 2r^2 + 8rr' - 2r'^2$. So that our expression reduces itself to

$$\begin{aligned} & \frac{2}{3S^2} \{ (r+r')^2 + k^2 - (r+r') \cdot R \} \\ &= \frac{2}{3S^2} \{ k^2 + (r+r')(r+r'-R) \} \end{aligned}$$

and therefore we have, finally,

$$t = \frac{\sqrt{2(r+r'+R)}}{6k \cdot \sqrt{M+m}} \{ k^2 + (r+r')(r+r'-R) \}$$

but $\sqrt{r+r'+R} \cdot (r+r'-R) =$
 $= \sqrt{(r+r'+R) \cdot (r+r'-R)^2}$
 $= k \cdot \sqrt{r+r'-R}$

because $(r+r'+R)(r+r'-R) = k^2$. So that substituting, we get

$$= \frac{k \cdot \sqrt{r+r'+R} + (r+r') \cdot \sqrt{r+r'-R}}{6 \cdot \sqrt{M+m}} \cdot \sqrt{2}$$

Now $R = \sqrt{(r+r')^2 - k^2}$, so that the two radicals in the above expression are of the form

$$\sqrt{a \pm \sqrt{a^2 - k^2}}$$

putting a for $r+r'$. Hence the usual rule for extracting the root of a binomial surd applies to them, the root being, as it is shewn in all books of algebra,

$$\frac{\sqrt{a+k} \pm \sqrt{a-k}}{\sqrt{2}}$$

first and second; secondly, from the first and third; and lastly, from the second and third; we shall get the three following

$$0 = x(y''z' - y'z'') + x'(yz'' - y''z) + x''(y'z - yz') \quad (65, 1)$$

$$0 = y(z''x' - z'x'') + y'(zx'' - xz') + y''(xz' - zx'') \quad (65, 2)$$

$$0 = z(x''y' - x'y'') + z'(xy'' - x''y) + z''(x'y - xy'') \quad (65, 3)$$

Now $zy' - z'y$ represents the double area of the projection of the plane triangle included between the radii, r, r' drawn to the place of the comet at the first and second observations, and the chord k of the parabolic arc joining their extremities on the plane of the z and y ; similarly, $zy'' - z''y$ represents double the triangle formed by the projections of the radii r, r'' and the chord k' joining their extremities on the same plane; and if we denote by k'' the chord joining the extremities of the radii r', r'' , the double of the projection of the triangle included between r', r'' , and k'' will be represented by $z'y'' - z''y'$. If, therefore, we call the surfaces of these triangles respectively

$$S, S', S''$$

Since the projection of each is equal to the surface multiplied by the cosine of its inclination to the plane of the z, y , which is the same for all, we shall have the

Consequently, we have

$$t = \frac{1}{6\sqrt{M+m}} \left\{ k(\sqrt{r+r'+k} + \sqrt{r+r'-k}) + (r+r')(\sqrt{r+r'+k} - \sqrt{r+r'-k}) \right\}$$

or,

$$6\sqrt{M+m} \times t = (r+r'+k)^{\frac{1}{2}} - (r+r'-k)^{\frac{1}{2}}; \quad (62)$$

which agrees with the proposition as announced, if we take $g = 6\sqrt{M+m}$.

The relation just proved to subsist between the time of describing any parabolic arc, and the rectilinear distances of its two extremities from each other, and from the sun, is very useful, as it enables us from any three of these quantities given, to find the fourth without knowing either the perihelion distance, or the position of the perihelion with respect to the arc described. In an analytical point of view, it is certainly remarkable for the length and complexity of the transformations by which it is obtained from the general principles of parabolic motion. It appears to admit of no simple demonstration, and that above given, tedious as it may seem, can hardly be replaced by one materially shorter.

We shall now proceed to show how this relation may be rendered available in computing the elements of a comet's orbit.

Let x, y, z , be the co-ordinates of the comet's place at the epoch of a first observation, x', y', z' , their values after any time t and x'', y'', z'' , their values after any other times t' has elapsed. Let X, Y, Z, X', Y', Z' , and X'', Y'', Z'' , be the corresponding co-ordinates of the earth's place. (If the plane of the ecliptic be chosen for that of the X and Y , we shall have $Z = 0, Z' = 0, Z'' = 0$.) Also let

$$t' - t = t''$$

then will t'' be the time elapsed between the second and third observations.

Since the comet moves in a plane passing through the sun, we have

$$z = ax + by \quad (64, 1)$$

$$z' = ax' + by' \quad (64, 2)$$

$$z'' = ax'' + by'' \quad (64, 3)$$

from which, if we eliminate a and b ; first, from the

following equation instead of (65,1),

$$xS'' - x'S' + x''S = 0 \quad (66, 1)$$

and similarly,

$$yS'' - y'S' + y''S = 0 \quad (66, 2)$$

$$zS'' - z'S' + z''S = 0 \quad (66, 3)$$

Now, if we retain the denominations of the foregoing pages, and assume the line of the equinoxes $S\gamma$ for the axis of the x , we have

$$x = S\pi = Se + e\pi$$

$$= R \cdot \cos L + \delta \cdot \cos \lambda; \quad (67, 1)$$

$$x' = R' \cdot \cos L' + \delta' \cdot \cos \lambda'; \quad (67, 2)$$

$$x'' = R'' \cdot \cos L'' + \delta'' \cdot \cos \lambda''; \quad (67, 3)$$

also

$$y = Ee + pq = R \cdot \sin L + \delta \cdot \sin \lambda; \quad (67, 4)$$

$$y' = R' \cdot \sin L' + \delta' \cdot \sin \lambda'; \quad (67, 5)$$

$$y'' = R'' \cdot \sin L'' + \delta'' \cdot \sin \lambda''; \quad (67, 6)$$

Astronomy. we have also

$$z = \delta \cdot \tan \beta \quad (67, 7)$$

$$z' = \delta' \cdot \tan \beta' \quad (67, 8)$$

$$z'' = \delta'' \cdot \tan \beta'' \quad (67, 9)$$

and by the substitution of these values we shall obtain equations, in which, instead of the co-ordinates x, y, z , &c. the curvate distances $\delta, \delta', \delta''$, will be involved, viz.

$$\left. \begin{aligned} o &= S'' (\delta \cdot \cos \lambda + R \cdot \cos L) \\ &- S' (\delta' \cdot \cos \lambda' + R' \cdot \cos L') \\ &+ S (\delta'' \cdot \cos \lambda'' + R'' \cdot \cos L'') \end{aligned} \right\}; \quad (68, 1)$$

$$\left. \begin{aligned} o &= S'' (\delta \cdot \sin \lambda + R \cdot \sin L) \\ &- S' (\delta' \cdot \sin \lambda' + R' \cdot \sin L') \\ &+ S (\delta'' \cdot \sin \lambda'' + R'' \cdot \sin L'') \end{aligned} \right\} \quad (68, 2)$$

$$o = S'' \cdot \delta \cdot \tan \beta - S' \cdot \delta' \cdot \tan \beta' + S \cdot \delta'' \cdot \tan \beta''; \quad (68, 3)$$

R, L, λ , and β , and their accented values, are all

quantities given by the solar tables, or by observation at the three assigned instants, so that the only unknown quantities contained in these equations are the three curvate distances, $\delta, \delta', \delta''$, and the areas S, S', S'' ; and by elimination, any one of the three former may be expressed in terms of the latter. Let this process be executed then in succession for each of the quantities $\delta, \delta', \delta''$, and we shall have the following results:

$$(a \cdot \delta + AR) \cdot S'' - A'R' \cdot S' + A''R'' \cdot S = 0; \quad (69, 1)$$

$$BR \cdot S'' - (a \cdot \delta' + B'R') S' + B''R'' S = 0; \quad (69, 2)$$

$$CR \cdot S'' - C'R' \cdot S' + (a \cdot \delta'' - C''R'') S = 0; \quad (69, 3)$$

where, for brevity's sake, we have made the following substitutions—

$$\left. \begin{aligned} a &= \tan \beta \cdot \sin (\lambda'' - \lambda') \\ &- \tan \beta' \cdot \sin (\lambda'' - \lambda) \\ &+ \tan \beta'' \cdot \sin (\lambda' - \lambda) \end{aligned} \right\}; \quad (70, 1)$$

$$A = \tan \beta \cdot \sin (L - \lambda'') - \tan \beta'' \cdot \sin (L - \lambda'); \quad (70, 2)$$

$$A' = \tan \beta' \cdot \sin (L' - \lambda'') - \tan \beta'' \cdot \sin (L' - \lambda'); \quad (70, 3)$$

$$A'' = \tan \beta' \cdot \sin (L'' - \lambda'') - \tan \beta'' \cdot \sin (L'' - \lambda'); \quad (70, 4)$$

$$B = \tan \beta'' \cdot \sin (L - \lambda) - \tan \beta \cdot \sin (L - \lambda''); \quad (70, 5)$$

$$B' = \tan \beta'' \cdot \sin (L' - \lambda) - \tan \beta \cdot \sin (L' - \lambda''); \quad (70, 6)$$

$$B'' = \tan \beta'' \cdot \sin (L'' - \lambda) - \tan \beta \cdot \sin (L'' - \lambda''); \quad (70, 7)$$

$$C = \tan \beta \cdot \sin (L - \lambda') - \tan \beta' \cdot \sin (L - \lambda); \quad (70, 8)$$

$$C' = \tan \beta \cdot \sin (L' - \lambda') - \tan \beta' \cdot \sin (L' - \lambda); \quad (70, 9)$$

$$C'' = \tan \beta \cdot \sin (L'' - \lambda') - \tan \beta' \cdot \sin (L'' - \lambda); \quad (70, 10)$$

In these equations, the right hand members are composed entirely of quantities given either by observation, or by the solar tables; and their calculation, though long, is greatly facilitated by the regularity of their composition. The values of a, A, B, C , &c. obtained, we have, by (69, 1) (69, 2)

$$\delta = - \frac{ARS'' - A'R'S' + A''R''S}{a S''}$$

$$\delta' = + \frac{BRS'' - B'R'S' + B''R''S}{a S'}$$

whence we obtain

$$\frac{\delta}{\delta'} = - \frac{S'}{S''} \cdot \frac{A'}{B'} \quad (71)$$

$$+ \frac{(B'A - BA') \cdot RS'' + (B'A'' - B''A') R''S}{B' \{BRS'' - B'R'S' + B''R''S\}} \times \frac{S'}{S''}$$

So far, we have nothing but rigorous equations, and it does not immediately appear how these can become serviceable in the question before us; but if we consider attentively the composition of the numerator of the fraction which forms the second term of $\frac{\delta}{\delta'}$ here stated, we shall find that the equation last arrived at, affords

a means of obtaining a first approximation. Let us consider, first, the term $AB' - BA'$. If we substitute in this the values of A, B, A', B' , in equations (70, 2, 3, 5, 6,) execute all the multiplications, strike out such terms as mutually destroy each other, and then reduce as much as possible, we shall find for its value

$$\tan \beta'' \left\{ \begin{aligned} &\tan \beta \cdot \{ \sin (L - \lambda') \cdot \sin (L' - \lambda'') - \sin (L' - \lambda') \cdot \sin (L - \lambda'') \} \\ &+ \tan \beta' \cdot \{ \sin (L - \lambda'') \cdot \sin (L' - \lambda) - \sin (L' - \lambda'') \cdot \sin (L - \lambda) \} \\ &+ \tan \beta' \cdot \{ \sin (L - \lambda) \cdot \sin (L' - \lambda') - \sin (L' - \lambda) \cdot \sin (L - \lambda') \} \end{aligned} \right\}$$

Now, by trigonometry, we have, (since $\sin A \cdot \sin B = \frac{1}{2} \{ \cos (A - B) - \cos (A + B) \}$)

$$\sin (L - \lambda) \cdot \sin (L' - \lambda') = \frac{1}{2} \cos \{ (L - L') - (\lambda - \lambda') \} - \frac{1}{2} \cos \{ (L + L') - (\lambda + \lambda') \}$$

$$\sin (L' - \lambda) \cdot \sin (L - \lambda') = \frac{1}{2} \cos \{ (L - L') + (\lambda - \lambda') \} - \frac{1}{2} \cos \{ (L + L') - (\lambda + \lambda') \}$$

and subtracting, we find

$$\begin{aligned} \sin (L - \lambda) \cdot \sin (L' - \lambda') - \sin (L' - \lambda) \cdot \sin (L - \lambda') &= \\ &= \frac{1}{2} \cos \{ (L - L') - (\lambda - \lambda') \} - \frac{1}{2} \cos \{ (L - L') + (\lambda - \lambda') \} \end{aligned}$$

which another application of the same trigonometrical formula already used converts back again into

$$\sin (L - L') \cdot \sin (\lambda - \lambda')$$

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Astronomy. a similar transformation applies to the co-efficients of $\tan \beta$ and $\tan \beta'$; so that our expression becomes

$$AB' - BA' = \tan \beta'' \times \left\{ \begin{array}{l} \sin (L - L') \cdot \sin (\lambda' - \lambda'') \cdot \tan \beta \\ + \sin (L - L') \cdot \sin (\lambda'' - \lambda) \cdot \tan \beta' \\ + \sin (L - L') \cdot \sin (\lambda - \lambda') \cdot \tan \beta'' \end{array} \right.$$

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or

$$AB' - BA' = \sin (L - L') \cdot \tan \beta'' \cdot \left\{ \begin{array}{l} \tan \beta \cdot \sin (\lambda' - \lambda'') \\ + \tan \beta' \cdot \sin (\lambda'' - \lambda) \\ + \tan \beta'' \cdot \sin (\lambda - \lambda') \end{array} \right\}$$

or, if we put $\gamma = \tan \beta \cdot \sin (\lambda' - \lambda'') + \tan \beta' \cdot \sin (\lambda'' - \lambda) + \tan \beta'' \cdot \sin (\lambda - \lambda')$; (72)

$$AB' - BA' = \gamma \cdot \sin (L - L') \cdot \tan \beta'' \quad (73, 1)$$

Now, it is evident that γ is a symmetrical function, and is not changed by putting at the same time β' for β , and λ' for λ , and reciprocally; so that, by pursuing a process of reduction exactly similar, we should obtain

$$AB'' - BA'' = \gamma \cdot \sin (L - L'') \cdot \tan \beta' \quad (73, 2)$$

$$A'B'' - B'A'' = \gamma \cdot \sin (L' - L'') \cdot \tan \beta'' \quad (73, 3)$$

and were we to pursue similar processes, combining A and C, B and C, we should get

$$CA' - AC' = \gamma \cdot \sin (L - L') \cdot \tan \beta \quad (73, 4)$$

$$CA'' - AC'' = \gamma \cdot \sin (L - L'') \cdot \tan \beta' \quad (73, 5)$$

$$C'A' - A'C' = \gamma \cdot \sin (L' - L') \cdot \tan \beta' \quad (73, 6)$$

and

$$BC' - CB' = \gamma \cdot \sin (L - L') \cdot \tan \beta \quad (73, 7)$$

$$BC'' - CB'' = \gamma \cdot \sin (L - L'') \cdot \tan \beta \quad (73, 8)$$

$$B'C'' - C'B'' = \gamma \cdot \sin (L' - L'') \cdot \tan \beta \quad (73, 9)$$

Thus the numerator of the fraction in $\frac{\delta}{\delta'}$ above referred to, becomes

$$\gamma \cdot \tan \beta'' \cdot \{RS'' \cdot \sin (L - L') - R''S \cdot \sin (L' - L'')\}; \quad (74)$$

Now, let us suppose the three observations made at small intervals of time. The supposition almost always holds good in the case of comets, which usually excite sufficient interest about the time of their first appearance, to induce astronomers to observe them as frequently as possible; so that observations separated by an interval of a few days at most may generally be obtained for calculation. On this supposition, then, the variations $\lambda - \lambda'$, $\lambda' - \lambda''$ and $\lambda'' - \lambda$ of the observed geocentric longitudes may be regarded as very small quantities of the first order; so that γ is also a very small quantity of the first order, especially as β (the observed geocentric latitude) is usually below 45° ; so that $\tan \beta$, $\tan \beta'$, &c. are less than unity. Moreover, $L - L'$, $L' - L''$, &c. which represent the sun's motion in the intervals between the observations, are also very small quantities of the first order; so that the expression (74) is, on these accounts, a very small quantity of the second order.

Fig. 10.

But besides this, since S, S'' represent the rectilinear triangles SPP' and SP'P'' described about the sun in the times $t' - t$ and $t'' - t'$ (which for the moment we will call θ and $n\theta$) they will be nearly proportional to those times, being extremely near to equality with the sectors SPnP', SP'n'P'', so that we have very nearly $S'' = n \cdot S$. Again, the earth's orbit being nearly a circle, and its motion in it nearly uniform, we have

$$R = R'' \text{ and } (L' - L'') = n \cdot (L - L')$$

whence $(L - L')$ being very small, $\sin (L' - L'') = n \cdot \sin (L - L')$. If these suppositions were rigorously exact, we should have

$$R \cdot S'' \cdot \sin (L - L') = nRS \cdot \sin (L - L') \\ = R''S \cdot \sin (L' - L'')$$

and the quantity within the brackets of (74) would be

rigorously zero; and in all cases we see that it must be a very small quantity of the second order at least, so that the whole expression (74) must be a very small quantity of the third order; and that, whether the intervals between the consecutive observations approach to equality or not.*

Let us next consider the denominator of the fraction, and retaining the substitution $n = \frac{\theta''}{\theta} = \frac{t'' - t'}{t' - t}$

it becomes, on the same hypotheses as before

$$B' \cdot RS \{nB - B' (1 + n) + B''\}$$

because $S' = S + S''$ nearly. We have, therefore, first to consider the values of $B'' - B'$ and of $B - B'$. Now, if in $B'' - B'$ we substitute the values of B' , B'' , and reduce as far as possible, we shall find

* Mr. Littrow, from whose excellent and most useful work on theoretical and practical astronomy (*Theoretische und Practische Astronomie*, von J. J. Littrow, director der Sternwarte und professor der astron. an der K. K. universität in Wien.—Wien, 1821, 2 vols. 8vo.) I have taken this exposé of the determination of a comet's orbit and the example which follows, has merely remarked

that the quantity $\frac{\delta}{\delta'}$ is equal to $-\frac{A'}{B'} \times \frac{S}{S''} + \text{a remainder}$, and

"this remainder is of the order $AB' - A'B$, and therefore may be neglected in a first approximation, when the intervals are small and nearly equal," (vol. ii. p. 129.) This is not, however, a satisfactory view of the subject, and is, in fact, slurring over a very considerable difficulty and evading one of the most difficult and delicate points in the theory of comets. I have therefore judged it better to enter, though somewhat more at length than is absolutely consonant with the nature of this work into this part of the subject, rather than leave a doubt on the mind of the student, or a mystery attached to a fundamental proposition. Meanwhile I am happy to have to call the attention of the lovers of astronomy to the interesting work alluded to, of which a translation by any competent hand would be a real accession to our stock of elementary works.

$$\underbrace{\text{Astronomy.}} \quad B'' - B' = 2 \tan \beta \cdot \tan \beta'' \cdot \sin \frac{L' - L''}{2} \cdot \left\{ \cot \beta \cdot \cos \left(\frac{L' + L''}{2} - \lambda \right) - \cot \beta'' \cdot \cos \left(\frac{L' + L''}{2} - \lambda'' \right) \right\} \quad \underbrace{\text{Physical Astronomy.}}$$

and the quantity within the brackets being the difference between two values of the same function at a small interval of time is of the first order, as is also $\sin \frac{L' - L''}{2}$, so that $B'' - B'$ is of the second order. Similarly, we have

$$B - B' = -2 \cdot \tan \beta \cdot \tan \beta'' \sin \frac{L - L'}{2} \cdot \left\{ \cot \beta \cdot \cos \left(\frac{L + L'}{2} - \lambda \right) - \cot \beta'' \cdot \cos \left(\frac{L + L'}{2} - \lambda'' \right) \right\}$$

Now, $\sin \frac{L' - L''}{2} = n \cdot \sin \frac{L - L'}{2}$, and therefore we have

$$n(B - B') + (B'' - B') = 2n \cdot \tan \beta \cdot \tan \beta'' \cdot \sin \left(\frac{L - L'}{2} \right) \cdot \left\{ \begin{aligned} &\cot \beta \left(\cos \left(\frac{L'}{2} - \lambda + \frac{L''}{2} \right) - \cos \left(\frac{L'}{2} - \lambda + \frac{L}{2} \right) \right) \\ &- \cot \beta'' \left(\cos \left(\frac{L'}{2} - \lambda'' + \frac{L''}{2} \right) - \cos \left(\frac{L'}{2} - \lambda'' + \frac{L}{2} \right) \right) \end{aligned} \right\}$$

Let this be reduced as much as possible; and, putting M for $\frac{L + 2L' + L''}{4}$, and $(1 + n) \cdot \sin \left(\frac{L - L'}{4} \right)$ for $\sin \left(\frac{L - L'}{4} \right)$ it will become

$$\begin{aligned} R \cdot S \times 2n(1 + n)^2 \cdot \tan \beta^2 \cdot \tan \beta''^2 \cdot \sin \frac{L - L'}{2} \cdot \sin \frac{L - L'}{4} \times \\ \times \{ \sin(L' - \lambda) \cdot \cot \beta - \sin(L' - \lambda'') \cdot \cot \beta'' \} \times \\ \times \{ \sin(M - \lambda) \cdot \cot \beta - \sin(M - \lambda'') \cdot \cot \beta'' \} \end{aligned}$$

This is evidently a quantity of the fifth order, the two last factors of it being each the difference of consecutive values.

Granting then the hypotheses above employed, viz.
—1. The strict circularity and uniformity of the earth's motion.—2. The proportionality of the rectilinear, instead of the parabolic, sectors to the times we see that the numerator of the fraction vanishes absolutely, while the denominator on the same hypotheses does not strictly vanish, but only reduces itself to a quantity of the fifth order. Hence, we are entitled (at least in making a first approximation to the elements) to neglect the fraction, and suppose

$$\frac{\delta}{\delta'} = -\frac{S'}{S''} \cdot \frac{A'}{B'}; \quad (75)$$

and similarly

$$\frac{\delta}{\delta''} = +\frac{S}{S''} \cdot \frac{A'}{C'}; \quad (76)$$

and

$$\frac{\delta'}{\delta''} = -\frac{B'}{C'} \cdot \frac{S}{S''}; \quad (77)$$

It is true, from the minuteness of the denominator of the fraction in (71) the deviation from truth of the hypotheses assumed becomes *magnified*. But these hypotheses themselves are so very nearly correct, that unless in extremely unfavourable cases the equations (75) and (76) may safely be employed; and it will be observed too, that *independent of both hypotheses*, the numerator would be of the fourth order at least, for if we throw it into the form

$$\begin{aligned} &\gamma \cdot \tan \beta'' \cdot SS'' \cdot \left\{ R \cdot \frac{\sin(L - L')}{S} - R'' \cdot \frac{\sin(L' - L'')}{S''} \right\} \\ &= \gamma \cdot \tan \beta'' \cdot SS'' \cdot \left\{ R \cdot \frac{\sin(L - L')}{S} - R' \cdot \frac{\sin(L' - L'')}{S''} \right\} \\ &\quad + \gamma \cdot \tan \beta'' \cdot SS'' \cdot (R' - R'') \cdot \frac{\sin(L' - L'')}{S''} \end{aligned}$$

the second term is evidently so; and since the factor of the first within the brackets is the difference of

consecutive values, and therefore a very small quantity,—the first is also.

To apply these equations to the determination of a comet's elements, we will first suppose the sectors $PnSP'$, &c. equal to the plane triangles SPP' , &c. or at least in the same ratio with them. This amounts

to substituting $\frac{t' - t}{t'' - t'}$ or $\frac{\theta'}{\theta''}$ for $\frac{S'}{S''}$; when we obtain, instead of the equations (69, 1, 2, 3,)

$$(a\delta + AR) \times \theta'' - A'R' \cdot \theta' + A''R'' \theta = 0 \quad (78,1)$$

$$BR \theta'' - (a\delta' + B'R') \theta' + B''R'' \theta = 0 \quad (78,2)$$

$$CR \theta'' - C'R' \theta' + (a\delta'' - C'R'') \theta = 0 \quad (78,3)$$

which give the values of $\delta, \delta', \delta''$.

Now we get, by the equation (55)

$$r'^2 = R'^2 + \delta'^2 \cdot \sec \beta'^2 + 2R'\delta' \cdot \cos(L' - \lambda'); \quad (79)$$

whence the value of r' is obtained.

r' once obtained, the values of $\delta, \delta', \delta''$, &c. and r' may be corrected by the following process, let τ represent the time between the second observation and any assigned instant. If then x, y, z , represent the co-ordinates at the first observation, and x'', y'', z'' , at the third, we have

$$x = x' - \frac{\theta}{1} \cdot \frac{dx'}{d\tau} + \frac{\theta^2}{1 \cdot 2} \cdot \frac{d^2x'}{d\tau^2} - \&c. \quad (80,1)$$

$$x'' = x' + \frac{\theta''}{1} \cdot \frac{dx'}{d\tau} + \frac{\theta''^2}{1 \cdot 2} \cdot \frac{d^2x'}{d\tau^2} - \&c. \quad (80,2)$$

$$y = y' - \frac{\theta}{1} \cdot \frac{dy'}{d\tau} + \frac{\theta^2}{1 \cdot 2} \cdot \frac{d^2y'}{d\tau^2} - \&c. \quad (80,3)$$

$$y'' = y' + \frac{\theta''}{1} \cdot \frac{dy'}{d\tau} + \frac{\theta''^2}{1 \cdot 2} \cdot \frac{d^2y'}{d\tau^2} - \&c. \quad (80,4)$$

and similar expressions for z, z'' . Now, assuming τ as the independent variable, or $d\tau$ constant, we have, by (1), (2), (3),

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$$\left. \begin{aligned} \frac{d^2 x'}{d\tau^2} &= -(M+m) \cdot \frac{x'}{r^3} \\ \frac{d^2 y'}{d\tau^2} &= -(M+m) \cdot \frac{y'}{r^3} \\ \frac{d^2 z'}{d\tau^2} &= -(M+m) \cdot \frac{z'}{r^3} \end{aligned} \right\} \quad (81, 1, 2, 3)$$

The numerical value of $M+m$ is easily found, since it is the same for the whole solar system except in so far as the masses of the planets and comets denoted by m differ, and these bear so minute a proportion to the mass M of the sun, that they may be neglected: let τ denote the period of any one plane, as the earth, and a the semi-major axis of its orbit, then we shall have, by (24)

$$\frac{T^2}{a^3} = \frac{4\pi^2}{M+m}$$

so that $\frac{\tau^2}{a^3}$ is a quantity very nearly constant for the whole planetary system; and the period of the earth being 365.256384, and the semiaxis of its orbit 1, its value for the earth is $\frac{(365.256384)^2}{1^3}$, and the same quantity expresses its value for all the other planets and comets. Hence we have

$$M+m = \frac{4 \times (3.141592)^2}{(365.256384)^2} = 0.0002959122;$$

and $\log(M+m) = 6.4711628$

To correct the values of δ , δ' , &c. we must correct the supposition on which they were obtained, of the proportionality of the rectilinear triangles instead of the parabolic sectors to the times, or the values of S , S' , S'' . Now, if c be the inclination of the plane of the orbit to the plane of the x and y , we have

$$2S \cdot \cos c = yx' - xy'$$

$$2S' \cdot \cos c = y'x'' - x'y''$$

$$2S'' \cdot \cos c = yx''' - xy'''$$

In the first of these, for y and x write their values as given in (80, 1) and (80, 3), and we find

$$\begin{aligned} 2S \cdot \cos c &= \frac{\theta}{1} \cdot \frac{y'dx' - x'dy'}{d\tau} \\ &\quad - \frac{\theta^2}{1.2} \cdot \frac{y'd^2x' - x'd^2y'}{d\tau^2} + \\ &\quad + \frac{\theta^3}{1.2.3} \cdot \frac{y'd^3x' - x'd^3y'}{d\tau^3} - \&c. \end{aligned}$$

But if the values of $\frac{d^2 x'}{d\tau^2}$ and $\frac{d^2 y'}{d\tau^2}$ be written in the second term of this, it will vanish. Again, if we differentiate the equations (81, 1) and (81, 2) we get

$$\begin{aligned} \frac{d^3 x'}{d\tau^3} &= -(M+m) \cdot \left(\frac{1}{r^3} \cdot \frac{dx'}{d\tau} - \frac{3x'}{r^4} \cdot \frac{dr'}{d\tau} \right) \\ \frac{d^3 y'}{d\tau^3} &= -(M+m) \cdot \left(\frac{1}{r^3} \cdot \frac{dy'}{d\tau} - \frac{3y'}{r^4} \cdot \frac{dr'}{d\tau} \right) \end{aligned}$$

so that, by substitution, we obtain

$$\frac{y'd^3x' - x'd^3y'}{d\tau^3} = -\frac{M+m}{r^3} \cdot \frac{y'dx' - x'dy'}{d\tau}$$

If we neglect the higher powers of θ than the

third, and put, for the present, p for $\frac{x'dy' - y'dx'}{d\tau}$ Physical Astronomy

we shall have

$$2S \cdot \cos c = p\theta \left\{ 1 - \frac{M+m}{r^3} \theta^2 \right\}; \quad (82, 1)$$

and similarly,

$$2S' \cdot \cos c = p\theta' \left\{ 1 - \frac{M+m}{r'^3} \theta'^2 \right\}; \quad (82, 2)$$

$$2S'' \cdot \cos c = p\theta'' \left\{ 1 - \frac{M+m}{r''^3} \theta''^2 \right\}; \quad (82, 3)$$

If we now employ the value of r' as above approximately determined, and substitute it in the right hand members of these equations, we shall obtain, dividing one by the other the ratios of S , S' , S'' , much more accurately than on the original hypothesis, for we have

$$\frac{S}{S''} = \frac{\theta}{\theta''} \times \frac{1 - \frac{M+m}{r^3} \theta^2}{1 - \frac{M+m}{r'^3} \theta'^2} \quad (83, 1)$$

$$\frac{S'}{S''} = \frac{\theta'}{\theta''} \times \frac{1 - \frac{M+m}{r'^3} \theta'^2}{1 - \frac{M+m}{r''^3} \theta''^2} \quad (83, 2)$$

Now, in the equations for determining δ , δ' , &c. (69, 1, 2, 3) it is only the ratio of S , S' , S'' , that are required, and thus we are enabled to correct the values of their quantities, and thence again that of r' . To carry the process to a greater degree of precision, however, the series expressing the value of $2S \cdot \cos c$ must be continued further. Without, however, going through the process as here set down we may content ourselves, after computing first values of δ' and r' by the equations (78, 2) and (79) with calculating the quantity

$$\begin{aligned} &\frac{M+m}{r^3} \left\{ B'R' \cdot \frac{\theta}{\theta'} (\theta'^2 - \theta^2) \right. \\ &\quad \left. - BR \cdot \frac{\theta''}{\theta'} (\theta''^2 - \theta'^2) \right\}; \quad (84) \end{aligned}$$

which must be applied as a correction to the first found value of δ' with its proper sign.

The values of r , r' , r'' , δ , δ' , δ'' , so obtained, the remaining unknown quantities may be found as follows:

The equation (58)

$$r \cdot \cos b \cdot \sin(l-L) = \delta \cdot \sin(\lambda-L)$$

divided by the equation (38)

$$r \cdot \cos b \cdot \cos(l-L) = R + \delta \cdot \cos(\lambda-L)$$

member for member, gives

$$\tan(l-L) = \frac{\delta \cdot \sin(\lambda-L)}{R + \delta \cdot \cos(\lambda-L)} \quad (85)$$

the right hand member consists only of known quantities, so that the value of $l-L$ is easily found, and thus the heliocentric longitude becomes known at each of the three observations. Again, we have by the equation (56)

$$r \cdot \sin b = \delta \cdot \tan \beta$$

and by (58)

$$r \cdot \cos b = \delta \cdot \frac{\sin(\lambda-L)}{\sin(l-L)}$$

Astronomy. which, divided member for member, give

$$\tan b = \tan \beta \cdot \frac{\sin(l-L)}{\sin(\lambda-L)}; \quad (86)$$

and thus the heliocentric latitudes may be computed.

The values of l and b known, the same equations afford a value of r , viz.

$$r = \delta \cdot \frac{\tan \beta}{\sin b}$$

whence r, r', r'' may be found; and if these values agree with those before determined, it will be a proof of the correctness of the previous work. If l' and l'' be greater than l , the comet's motion is direct; if less, retrograde.

The heliocentric longitudes and latitudes thus obtained at each of the three observations, we easily obtain the inclination and place of the node. In fact, (taking the first and last observations, to embrace a greater arc of the great circle) our equation (40) gives

$$\begin{aligned} \tan b &= \tan i \cdot \sin(l - \Omega) \\ \tan b'' &= \tan i \cdot \sin(l'' - \Omega) \end{aligned}$$

so that we have

$$\frac{\sin(l - \Omega)}{\sin(l'' - \Omega)} = \frac{\tan b}{\tan b''}$$

which gives

$$\begin{aligned} \tan b'' (\sin l \cdot \cos \Omega - \cos l \cdot \sin \Omega) \\ = \tan b (\sin l'' \cdot \cos \Omega - \cos l'' \cdot \sin \Omega) \end{aligned}$$

$$\begin{aligned} \text{or } \sin \Omega \{ \tan b \cdot \cos l'' - \cos l \cdot \tan b'' \} = \\ = \cos \Omega \{ \sin l \cdot \tan b'' - \sin l'' \cdot \tan b \} \end{aligned}$$

and consequently

$$\tan \Omega = \frac{\tan b \cdot \cos l'' - \tan b'' \cdot \cos l}{\sin l \cdot \tan b'' - \sin l'' \cdot \tan b}; \quad (87)$$

and

$$\tan i = \frac{\tan b}{\sin(l - \Omega)} \quad (88)$$

If the motion of the comet be *direct*, the value of Ω given by the above equation, will be the longitude of the *ascending* node; but if *retrograde*, it is evidently that of the *descending*; and to get that of the ascending node, 180° must be added. The equation (61) or the corresponding one,

$$\tan(l - \Omega) = \cos i \cdot \tan(\psi + \theta) \quad (89)$$

gives the values of $\psi + \theta$ the longitudes on the orbit reckoned from the node, and these being obtained, $(\psi + \theta, \psi + \theta', \text{ and } \psi + \theta'')$ * the angle $\theta'' - \theta$ between the first and last places of the comet at the sun becomes known.

If we denote this by ϕ , we have at once, by the equation

$$\tan \frac{\theta}{2} = \operatorname{cosec} \phi \left\{ \cos \phi - \sqrt{\frac{r}{r'}} \right\} \quad (90)$$

demonstrated in the foregoing pages (p. 665) the value of θ , or the distance (on the orbit) of the perihelion from the place of the first observation; so that, subtracting the value found, from $\psi + \theta$ before determined, we get ψ , and thence π , the longitude of the perihelion on the ecliptic, by the equation (60)

It only remains to find the perihelion distance and time of perihelion passage. The equation of the parabola

$$D = r \cdot \left(\cos \frac{\theta}{2} \right)^2$$

gives the former immediately, r and θ being known, and the latter is found as follows. In the table of a comet of 109 days subjoined, seek the number of days (by interpolation) corresponding to the anomaly θ , and call it M , then, by equation (44) we have, (the perihelion distance of such a comet being unity)

$$N = \sqrt{\frac{2}{M+m}} \times \left\{ \tan \frac{\theta}{2} + \frac{1}{3} \cdot \tan \left(\frac{\theta}{2} \right)^3 \right\}$$

and consequently, if T be the time of the perihelion passage *before* the first observation,

$$T = N \times D^{\frac{1}{2}}$$

If we institute a similar operation for the time T'' of the perihelion passage *before* the third observation, we shall find

$$T'' = N'' \cdot D^{\frac{1}{2}}$$

Now, we ought to have $T'' - T =$ the interval between the first and last observations. If this be found on trial to be the case, it is a proof of the correctness of the work. The perihelion passage will be before or after the first observation (in the case of direct motion) according as the longitude π of the perihelion is less or greater than the heliocentric longitude l at the first observation; the reverse is the case in retrograde motion.

The chief difficulty in the solution of the problem of a comet's orbit, consists in the determination of the distances from the sun and earth, at the moments of the several observations. Our equations above derived afford a great variety of means for accomplishing this, among which we shall only select one, proposed by Dr. Olbers, which in many cases has particular advantages. It depends on the property already demonstrated of the chords of the parabolic arcs.

Let k represent the chord of the arc included between the two extreme places, then we have

$$k^2 = (x'' - x)^2 + (y'' - y)^2 + (z'' - z)^2$$

$$\text{or } k^2 = r^2 + r''^2 - 2(xx'' + yy'' + zz'')$$

In this equation, let the values of x, y, z, x'', y'', z'' , in (67, 1, 2, 3, 4, &c.) in terms of R and δ , &c. be substituted; and if we put

$$m = \frac{C'}{A'} \times \frac{\theta''}{\theta}$$

so that, by (76) $\delta'' = m \delta$

we shall have

$$\begin{aligned} k^2 &= r^2 + r''^2 \\ &- 2m\delta^2 \cdot \{ \cos(\lambda - \lambda'') + \tan \beta \cdot \tan \beta'' \} \\ &- 2mR\delta \cdot \cos(\lambda'' - L) \\ &- 2R''\delta \cdot \cos(\lambda - L'') \\ &- 2RR'' \cdot \cos(L - L'') \end{aligned} \quad (91)$$

This equation combined with the equation (63)

$$6 \cdot \sqrt{M+m} \cdot \theta' = (r + r'' + k)^{\frac{3}{2}} - (r + r'' - k)^{\frac{3}{2}}$$

and the two equations

$$r^2 = R^2 + \delta^2 \cdot \sec^2 \beta + 2R\delta \cdot \cos(\lambda - L)$$

$$r''^2 = R''^2 + m^2 \delta^2 \cdot \sec^2 \beta'' + 2mR''\delta \cdot \cos(\lambda'' - L)$$

gives four equations for determining the four unknown quantities. To resolve them, we must use the well known *rule of false*, one of the simplest and most

* The reader will not confound θ here with the same letter before used to denote the time.

Astronomy. widely useful rules mathematical science affords. Assume a value for δ , as near the truth as can be conjectured, compute m and thence $m\delta$, or δ'' ; and having used these to obtain k , r , and r'' , substitute these in the expression for $g\theta'$. Call the result A . Assume then a second value for δ , which call Δ , and executing the same process, call the result B . Then we shall have

$$B - A : g\theta' - A :: \Delta - \delta : x - \delta$$

where x is the corrected value of δ , and hence we get

Mean Time, Paris.	h	'	''		°	'	''		°	'	''
1779. Aug. 30.....	11	9	42,	$\lambda = 125$	48	39.3		$\beta = 41$	53	52.2	
Sep. 2.....	10	36	8,	$\lambda' = 132$	53	48.5		$\beta' = 45$	54	48.1	
Sep. 4.....	10	7	51,	$\lambda'' = 138$	56	31.2		$\beta'' = 48$	32	27.8	

$$\begin{aligned} L &= 337 \ 29 \ 8.7 & R &= 1.0087218 \\ L' &= 340 \ 22 \ 26.9 & R' &= 1.0079991 \\ L'' &= 342 \ 17 \ 47.8 & R'' &= 1.0074854 \end{aligned}$$

The intervals therefore between the first and second, first and third, and second and third observations are respectively* $\theta = 2.976690$ days, $\theta' = 4^d.957049$, $\theta'' = 1^d.980350$

Consequently, we have by the equations expressing the values of A' and C' reduced into numbers

$$\frac{C'}{A'} = 1.18408; \quad m = \frac{\delta''}{\delta} = 0.787752$$

and substituting this value of m , and the values of λ , λ' , β , β' , R , R'' , we get the following equations, noticing that $g = 0.103212$

$$r^2 = 1.01752 - 1.71693 \cdot \delta + 1.80493 \cdot \delta^2; \quad r'^2 = 1.01503 - 1.45724 \cdot \delta + 1.41564 \cdot \delta^2$$

$$k^2 = 0.00716 - 0.04739 \cdot \delta + 0.08627 \cdot \delta^2; \quad 135.84219 = \left(\frac{r + r'' + k}{2} \right)^{\frac{1}{2}} - \left(\frac{r + r'' - k}{2} \right)^{\frac{1}{2}}$$

These equations are very nearly satisfied by taking $\delta = 0.71469$, whence we find

$$r = 0.8440; \quad r'' = 0.8346; \quad k = 0.1317; \quad \delta'' = m\delta = 0.562998$$

whence we next deduce the heliocentric longitudes and latitudes at the extreme observations from equations (85) and (86)

$$\begin{aligned} l &= 20^\circ \ 37' \ 40''.8 \quad \text{and} \quad b = 49^\circ \ 26' \ 23''.1 \\ l' &= 6 \ 45 \ 39.7 \quad b' = 49 \ 46 \ 46.9 \end{aligned}$$

Since l'' is less than l , the comet's motion is retrograde. Moreover, we have by equations (87) and (88)

$$\varpi = 100^\circ \ 51' \ 53''.4; \quad i = 49^\circ \ 51' \ 7''.9$$

we obtain also, by following the process pointed out in equations (89) and (90) $\theta = 12^\circ \ 39' \ 35''.6$, $\pi = 4^\circ \ 32' \ 8''.2$ for the longitude of the perihelion, and $D = 0.830761$ for the perihelion distance. Consequently the time between the first observation and the perihelion passage is

$$\begin{aligned} \frac{D^{\frac{2}{3}} \cdot \sqrt{2}}{\sqrt{M + m}} \cdot \left\{ \tan \frac{\theta}{2} + \frac{1}{3} \left(\tan \frac{\theta}{2} \right)^3 \right\} &= + 6.97118 = \\ &= 6^d \ 23^h \ 18' \ 30'' \\ \text{30th Aug. 11} \quad 9 \ 42 \end{aligned}$$

$$10 \ 28 \ 12 \ \text{Sept. 6, 1799} = \text{time of perihelion passage.}$$

Of the two solutions given above, the latter only supposes the orbit to be determined a parabola; the former may be applied to the determination of elliptic elements, but the latter part of the process from equation (70) must be varied accordingly; but as the various methods of determining and correcting elliptic as well as parabolic elements would lead us a great deal beyond our proposed limits, we must content ourselves with referring to Mr. Gauss's *Theoria Motus Corporum Coelestium*, in which the whole subject is treated with the utmost generality. The reader may

$$x = \delta + (\Delta - \delta) \cdot \frac{g\theta' - A}{B - A}$$

and substituting this again, and employing that one of the former values of δ which gave a result nearest to $g\theta'$ for a corresponding value, we shall obtain a yet nearer approximation, and so on.

It remains to exemplify the method of computation by an instance. Let us therefore take the following observations of the comet of 1779, taken at Paris with a transit and a repeating circle.

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* The reader is requested to take the numbers in this example upon Mr. Littrow's authority, from whose work the example is taken.

PART II. SECTION I.

OF THE PERTURBATION OF THE ELLIPTIC MOTIONS OF THE HEAVENLY BODIES, ARISING FROM THEIR MUTUAL ATTRACTION

IN the preceding investigations we have supposed only two attracting bodies to exist in space; and on this supposition have succeeded in representing all the phenomena of their motions in finite equations, by the aid of which their relative situations may be assigned at any past or future moment. The resulting formulæ reduced into tables, and compared with actual observation, manifest an agreement sufficiently complete to leave no doubt of the correctness of the principles from which they have been deduced, at least, when observations separated only by moderately long intervals of time are compared together. Yet, on descending to a more rigorous nicety, and especially on comparing together observations embracing a very long series of years, it is found that the results of the calculations, founded on the assumption of elliptic motion according to the laws above demonstrated, do not represent the observations perfectly. Minute irregularities are still detected, and the planets are found sometimes a little in advance, and sometimes a little falling short, sometimes a little to the right or left, above or below, their calculated places in their orbits. Moreover, if the elements of the orbits themselves, as deduced from modern observations, be compared with those similarly deduced from ancient ones, they will be found not to correspond exactly, but to differ by small variations, which are however greater, as the observations themselves compared are more distant in their dates. In a word, though the elliptic theory agrees very nearly with observation; yet, to make it tally rigorously with them, it is necessary to introduce modifications of this kind.

1st. The ellipse in which each planet moves, must be conceived to change its eccentricity and position in space, by exceedingly slow gradations—so slow indeed, as to be insensible in a single revolution of the planet, and only discoverable by a comparison, such as we have described, of its nature in past and present ages.

2dly. The planet itself is not always found exactly in its place in the ellipse *even* when so varied; nor indeed is it always found in the exact periphery of the ellipse at all. But if we conceive an imaginary point to describe this ellipse according to the rigorous laws of elliptic motion, the real planet will never be very far distant from it, but will oscillate or revolve round it like a secondary about its primary in an orbit of extremely small dimensions, yet according to laws of a very complicated nature, too complicated indeed for observation alone to unravel.

These variations are much more sensible in the orbit and motions of the moon, where, in fact, they amount to very considerable quantities. If the moon, for instance, at its greatest northern latitude be observed to pass over a certain star, it should continue to do so each revolution, if its motion were strictly conformable to the elliptic hypothesis. So far, however, is this from being the case, that it will be observed to deviate visibly southward every lunation; and after a lapse of nine years and a half, its path, in the same part of the heavens as to longitude, will pass not less than 10° south of the star; after which it will again advance northward, and in nine years and a half more will once more pass over, or at least very near the star. Deviations like these must have some positive and decided cause; as much so as the elliptic motion itself; and if they are to be accounted for on the theory of universal attraction, it is manifest that the same mechanical principles applied in the same way to the case of several bodies abandoned in free space to their mutual attractions, ought to lead us to their explanation.

But the mathematical difficulties to be encountered in this research, are of a much higher order than in the case of two bodies only. There we had no difficulty in integrating the differential equations of the problem. Here, on the other hand, the equations are too complicated to allow of their integrals being exhibited otherwise than in series; and even then we have no means of ascertaining their laws. Fortunately, however, this is not necessary; the enormous preponderance of the sun's mass in our system being such, that the fractions, representing those of each of the other planets, are small enough to allow of their squares and products being neglected without any fear of inducing appreciable errors.

It is a well known theorem, which extends to all the applications of mathematical reasoning to natural phenomena, that when several causes of motion act together, if their effects taken singly are of such an order of minuteness, that their squares and products may be disregarded, then their joint effect will be the sum of the effects which would be produced by each acting alone. This principle allows us to regard the disturbance or perturbation of any one planet produced by the joint action of all the others, as the sum of the effects each would produce separately; so that we may simplify the investigation by leaving out of consideration all but a single disturbing planet, and applying the analytical formulæ so investigated to each in succession, we shall thus obtain expressions for the perturbations produced by each; which, added together, give the total perturbation.

Investigation of the forces exerted by one body to disturb the orbit of another revolving about a common central body, and of the differential equations of their motions.

Taking unity for the mass of the central body, and its centre for the origin of the co-ordinates, let m and m' represent the masses of two revolving bodies, and suppose x, y, z, r to denote the three co-ordinates and radius vector of the body m , whose perturbations we would investigate; and x', y', z', r' , the corresponding quantities relative to the disturbing body m' . Moreover, let λ represent the distance of the two bodies, m, m' ,

Astronomy. on each other. Then will the several attractive forces exerted by the bodies of the system on each other be Physical Astronomy. as follows :—

1st. The sun (or central body) attracts m and m' with forces represented respectively by $\frac{1}{r^2}$ and $\frac{1}{r'^2}$.

2dly. m attracts the sun with a force $\frac{m}{r^2}$, and it attracts m' with a force $\frac{m}{\lambda^2}$.

3dly. m' attracts the sun with the force $\frac{m'}{r'^2}$, and it attracts m with the force $\frac{m'}{\lambda^2}$.

Fig. 11

These are the forces exerted in the directions of the lines joining the several bodies in the system; but to estimate their effects, they must be reduced to the directions AS, BS, CS, of the three co-ordinates. This done, we shall find

For the forces acting on the sun,

$$\left. \begin{aligned} \text{Force of } m &= -\frac{m x}{r^3} \\ \text{Force of } m' &= -\frac{m' x'}{r'^3} \end{aligned} \right\} \text{ in the direction AS}$$

and similarly,

$$\begin{aligned} -\frac{m y}{r^3} \text{ and } -\frac{m' y'}{r'^3} &\text{ in the direction BS} \\ -\frac{m z}{r^3} \text{ and } -\frac{m' z'}{r'^3} &\text{ in the direction CS.} \end{aligned}$$

So that the three forces acting on the sun, are respectively

$$\begin{aligned} &-\left(\frac{m x}{r^3} + \frac{m' x'}{r'^3}\right) \\ &-\left(\frac{m y}{r^3} + \frac{m' y'}{r'^3}\right) \\ &-\left(\frac{m z}{r^3} + \frac{m' z'}{r'^3}\right) \end{aligned}$$

Again, the forces acting on m are

$$\begin{aligned} \text{In the direction AS} \dots\dots\dots &\left\{ \begin{aligned} \text{Force of the sun} &= +\frac{x}{r^3} \\ \text{Force of } m' \dots\dots &= -\frac{m' (x' - x)}{\lambda^3} \end{aligned} \right. \\ \text{In the direction BS} \dots\dots\dots &\left\{ \begin{aligned} \text{Force of the sun} &= +\frac{y}{r^3} \\ \text{Force of } m' \dots\dots &= -\frac{m' (y' - y)}{\lambda^3} \end{aligned} \right. \\ \text{In the direction CS} \dots\dots\dots &\left\{ \begin{aligned} \text{Force of the sun} &= +\frac{z}{r^3} \\ \text{Force of } m' \dots\dots &= -\frac{m' (z' - z)}{\lambda^3} \end{aligned} \right. \end{aligned}$$

So that the aggregate forces acting on m in these directions are

$$\begin{aligned} &+\frac{x}{r^3} - m' \cdot \frac{x' - x}{\lambda^3} \\ &+\frac{y}{r^3} - m' \cdot \frac{y' - y}{\lambda^3} \\ &+\frac{z}{r^3} - m' \cdot \frac{z' - z}{\lambda^3} \end{aligned}$$

and in the same manner, the forces acting on m' will be found to be respectively

$$\begin{aligned} &+\frac{x'}{r'^3} - m \cdot \frac{x - x'}{\lambda^3} \\ &+\frac{y'}{r'^3} - m \cdot \frac{y - y'}{\lambda^3} \\ &+\frac{z'}{r'^3} - m \cdot \frac{z - z'}{\lambda^3} \end{aligned}$$

Astronomy. From these three sets of forces we might determine separately the motions of the sun, of m , and of m' ; but as we have chosen to suppose the sun at rest, and fix the origin of our co-ordinates in its centre, we must transfer to both m and m' in a contrary direction the forces which act upon it. This done, the forces animating m in the respective directions AS, BS, CS, will become

$$\begin{aligned} \frac{(1+m) \cdot x}{r^3} + m' \left\{ \frac{x'}{r'^3} - \frac{x'}{\lambda^3} \right\} \\ \frac{(1+m) \cdot y}{r^3} + m' \left\{ \frac{y'}{r'^3} - \frac{y'}{\lambda^3} \right\} \\ \frac{(1+m) \cdot z}{r^3} + m' \left\{ \frac{z'}{r'^3} - \frac{z'}{\lambda^3} \right\} \end{aligned}$$

and if, in these expressions, we exchange the accented letters for the corresponding unaccented ones, and *vice versa*, (λ only excepted) we shall evidently have the forces acting on m' .

The first terms of these expressions, as will be observed, are the same as if the mass m' of the disturbing planet were nothing, or as if there were no such planet; the other parts, those multiplied by m' , express the *disturbing forces*, or the effect of the attraction of m' to derange the orbit of m from its elliptical form. These, it will also be observed, consist each of two parts; the one expressing the action of m' on the sun, and the other its action on m with an opposite sign. It is only therefore in virtue of the difference of the attractions of m' on the sun, and on m that the orbit of the latter is deranged.

These forces are susceptible of a very simple mode of expression; and the equations of m 's motion in consequence admit of a very compendious form, if we consider, that owing to the form of the function λ , which is equal to $\sqrt{(x'-x)^2 + (y'-y)^2 + (z'-z)^2}$, we have

$$\left. \begin{aligned} \frac{x'-x}{\lambda^3} &= \frac{d}{dx} \cdot \frac{1}{\lambda} \\ \frac{y'-y}{\lambda^3} &= \frac{d}{dy} \cdot \frac{1}{\lambda} \\ \frac{z'-z}{\lambda^3} &= \frac{d}{dz} \cdot \frac{1}{\lambda} \end{aligned} \right\} \dots\dots\dots (92)$$

and that moreover, if we put U for the function $\frac{x x' + y y' + z z'}{r^3}$, we shall have

$$\frac{x'}{r^3} = \frac{dU}{dx}; \quad \frac{y'}{r^3} = \frac{dU}{dy}; \quad \frac{z'}{r^3} = \frac{dU}{dz}$$

so that if we assume

$$\begin{aligned} \Omega &= U - \frac{1}{\lambda} \\ &= \frac{x x' + y y' + z z'}{(x'^2 + y'^2 + z'^2)^{\frac{3}{2}}} - \frac{1}{\sqrt{(x'-x)^2 + (y'-y)^2 + (z'-z)^2}} \end{aligned} \quad (93)$$

the disturbing forces will be represented by

$$m' \cdot \frac{d\Omega}{dx}; \quad m' \cdot \frac{d\Omega}{dy}; \quad m' \cdot \frac{d\Omega}{dz}$$

and the equations representing the motion of m , will be (if we put $1+m=\mu$) and suppose dt the element of the time invariable.

$$\left. \begin{aligned} 0 &= \frac{d^2 x}{dt^2} + \mu \cdot \frac{x}{r^3} + m' \cdot \frac{d\Omega}{dx} \\ 0 &= \frac{d^2 y}{dt^2} + \mu \cdot \frac{y}{r^3} + m' \cdot \frac{d\Omega}{dy} \\ 0 &= \frac{d^2 z}{dt^2} + \mu \cdot \frac{z}{r^3} + m' \cdot \frac{d\Omega}{dz} \end{aligned} \right\} \dots\dots\dots (94)$$

Similarly, if we put $U' = \frac{x x' + y y' + z z'}{r^3}$ and take $\mu' = 1+m'$, and

$$\Omega' = U' - \frac{1}{\lambda} \quad (95)$$

the motion of m will be represented by the equations

$$\left. \begin{aligned} 0 &= \frac{d^2 x'}{dt^2} + \mu' \cdot \frac{x'}{r'^3} + m \cdot \frac{d\Omega'}{dx'} \\ 0 &= \frac{d^2 y'}{dt^2} + \mu' \cdot \frac{y'}{r'^3} + m \cdot \frac{d\Omega'}{dy'} \\ 0 &= \frac{d^2 z'}{dt^2} + \mu' \cdot \frac{z'}{r'^3} + m \cdot \frac{d\Omega'}{dz'} \end{aligned} \right\} \dots\dots\dots (96)$$

Astronomy. The joint integration of these six equations would determine the motions of both m and m' , but their rigorous integration being impracticable in the present state of analysis, we are driven to have recourse to methods of approximation, of which the principle may be explained as follows:— **Physical Astronomy.**

Suppose $A = 0$ to be any equation, or system of equations, susceptible of rigorous integration, and let the process to be followed for this purpose, consist in deducing in succession from $A = 0$, by any analytical artifices, or transformations, the equations, or systems of equations,

$$B = 0, \quad C = 0, \dots\dots\dots K = 0$$

the last of which, $K = 0$, is integrable by known methods. Now, suppose, instead of $A = 0$, we had proposed the equation

$$A + m' \cdot a = 0$$

where $m' \cdot a$ is a minute term of the order of the disturbing forces, m' being an extremely small constant quantity, and a any assigned function of the co-ordinates. If then we pursue with this equation the very same process by which $B = 0$ was derived from $A = 0$, it is clear that we shall obtain an equation which can differ only from $B = 0$ by a quantity which vanishes when $m' = 0$, and which therefore must be of the form $m' \times b$. So that instead of $B = 0$, we have

$$B + m' \times b = 0$$

where b is some explicit function of the co-ordinates, their differential co-efficients, and of m' , and either developed, or at least developable in a series of ascending powers of m' . In like manner we may deduce equations

$$C + m' \cdot c = 0; \dots\dots\dots K + m' \cdot k = 0$$

in the place of the equations $C = 0, \dots\dots\dots K = 0$

In order then to obtain a final equation for the solution of the problem of three bodies, we have only to follow out any system of processes and transformations which, in the case of two only, would prove successful in reducing the differential equations to an integrable form.

This may be accomplished in a great variety of ways, as the equations of undisturbed motion are integrable by a great many different artifices, besides those which we have employed in the former part of this essay; and it is therefore necessary to select among them such as lead to final equations best adapted to the nature of the case under consideration. Now two courses have been adopted by geometers; the former adapted to the theory of the planetary, and the latter to that of the lunar perturbations.

In the theory of the planets, the disturbing force is so extremely minute, that its square and higher powers may be neglected with safety. This simplification being permitted, it becomes practicable (as we shall presently see) so to conduct the investigation as to make the time, or mean longitude of the disturbed planet our independent variable, and thus to express the perturbations at once in functions of the time, or of angles proportional to it, a simplification of the utmost moment in the construction of tables. In the more complicated theory of the moon, in which the part of the perturbations depending on higher powers of the disturbing force than the first is very conspicuous, it is no longer permitted to neglect them, and the necessity of preserving, as far as possible, the rigorous expressions of the forces, &c. at least in the differential equations, obliges us to employ, not the mean, but the true longitude of the moon for our independent variable, as by this means we are enabled to arrive at final equations perfectly rigorous, and can thus estimate the influence which the neglect of small quantities is capable of producing in their integration, with less likelihood of being misled.

The theory of the moon then differs entirely from that of the planets in its treatment. The general principles of approximation, however, are the same in both. Both theories, as we shall see, lead to final equations of the form

$$\frac{d^2 u}{dt^2} + n^2 u + m' \cdot k = 0 \quad (97)$$

where n^2 is constant, and k an explicit function of the several co-ordinates, distances, and angles, of the problem, and m' a very small quantity (which, in the lunar theory, also enters into the composition of k), of the order of the disturbing forces. To approximate then to the value of u , we first suppose $m = 0$, and we get a value of u corresponding to $m' = 0$, and which we call its *elliptic value*.

2dly. We deduce from this the *elliptic values* of all the variables which enter into the composition of the term $m' k$, in terms of the independent variable t , whether t represent the mean, or the true longitude. These being substituted in the last term, it will become an explicit function of the independent variable.

3dly. If we now again integrate the differential equation so prepared, the value of u will consist of two parts; the first will be the same as before, viz. the elliptic value, and the second will be a correction which must be of the order of disturbing forces, and will express the perturbation of u with a degree of precision corresponding to their first power.

If the same process of substitution be repeated, and the equation again integrated, another set of terms will be added to the value of u , which carry the approximation a step farther, or to the squares of the disturbing forces; and were the same process continued to infinity, the series of terms so obtained would be a rigorous analytical expression for u .

The final equation (97) is of the form so well known in analysis, under the name of linear differential equations, and (as we have observed) almost all the equations on which the planetary motions depend being of this nature, it will be right to premise some few points relative to their theory, to which to refer hereafter.

It is demonstrated in all works in the differential and integral calculus, and will be so in our article on that subject, that any linear differential equation of the second order is integrable, provided we can find two, or even one particular value capable of satisfying it when deprived of its last term. Thus, if

$$\frac{d^2 u}{dt^2} + M \cdot \frac{du}{dt} + N u + \Pi = 0 \quad (98)$$

Astronomy. be any such equation, M and N being functions of t , and if u' and u'' be any two functions of t which satisfy Physical Astronomy the equation

$$\frac{d^2 u}{dt^2} + M \cdot \frac{du}{dt} + N \cdot u = 0$$

then the complete integral will be

$$C' u' + C'' u'' - u' \int d \frac{u''}{u'} \int \frac{\Pi dt^2}{u' d \frac{u''}{u'}} \quad (99)$$

Now, the last term of this, being integrated by parts, becomes

$$\begin{aligned} u' \left\{ \frac{u''}{u'} \int \frac{\Pi dt^2}{u' d \frac{u''}{u'}} - \int \frac{\Pi dt^2 \times u''}{u'^2 d \frac{u''}{u'}} \right\} \\ = u'' \int \frac{\Pi dt^2}{u' d \frac{u''}{u'}} + u' \int \frac{\Pi dt^2}{u'' d \frac{u'}{u''}} \end{aligned}$$

Consequently, the complete integral of the equation (98) will be

$$u = C' u' + C'' u'' - \left\{ u' \int \frac{\Pi dt^2}{u'' d \frac{u'}{u''}} + u'' \int \frac{\Pi dt^2}{u' d \frac{u''}{u'}} \right\} \quad (100)$$

Suppose we have the equation

$$\frac{d^2 u}{dt^2} + n^2 u + \Pi = 0 \quad (101)$$

then u' and u'' , the particular integrals of $\frac{d^2 u}{dt^2} + n^2 u = 0$ offer themselves readily, being no other than $\sin nt$ and $\cos nt$; and if these be substituted in the general expression above given, we get at once

$$\begin{aligned} u = C' \cdot \cos nt + C'' \cdot \sin nt \\ + \frac{1}{n} \left\{ \cos nt \int \Pi dt \cdot \sin nt - \sin nt \int \Pi dt \cos nt \right\}; \end{aligned} \quad (102)$$

and if $n = 1$, or in the case of

$$\frac{d^2 u}{dt^2} + u + \Pi = 0 \quad (103)$$

the integral is

$$\begin{aligned} u = C' \cdot \cos t + C'' \cdot \sin t \\ + \cos t \cdot \int \Pi dt \cdot \sin t - \sin t \int \Pi dt \cdot \cos t \end{aligned} \quad (104)$$

These values of u are rigorous, whatever be the value of Π , and independent of any approximation, as it is easy for the reader to satisfy himself by substituting them in the differential equations from which they were deduced, when the whole will be found to vanish, independent of any particular value of Π , or any supposition made as to its magnitude. The equation (102) may however be obtained perhaps easier as follows.

Let $\frac{d^2 u}{dt^2} + n^2 u = -\Pi$ be multiplied by $dt^2 \cdot \cos nt$ and it becomes

$$d^2 u \cdot \cos nt + n^2 u dt^2 \cdot \cos nt = - \int \Pi dt^2 \cdot \cos nt$$

and integrating,

$$d u \cdot \cos nt + n u dt \cdot \sin nt = - \int \Pi dt^2 \cdot \cos nt$$

If we again multiply this by $\frac{1}{\cos nt^2}$ we get

$$\begin{aligned} \frac{du}{\cos nt} + n u dt \cdot \frac{\sin nt}{(\cos nt)^2} &= - \frac{dt}{(\cos nt)^2} \int \Pi dt \cdot \cos nt \\ &= - \frac{d \cdot \tan nt}{n} \int \Pi dt \cdot \cos nt \end{aligned}$$

and again integrating,

$$\begin{aligned} \frac{u}{\cos nt} &= - \frac{1}{n} \int d \cdot \tan nt \int \Pi dt \cdot \cos nt \\ &= \frac{1}{n} \left\{ - \tan nt \int \Pi dt \cdot \cos nt + \int \Pi dt \cdot \sin nt \right\} \end{aligned}$$

by integrating by parts. Hence we have

$$u = \frac{1}{n} \left\{ \cos nt \int \Pi dt \cdot \sin nt - \sin nt \int \Pi dt \cdot \cos nt \right\} \quad (105)$$

Astronomy. or, expressing the arbitrary constants by C' and C''

$$u = C' \cdot \cos nt + C'' \cdot \sin nt + \frac{1}{n} \left\{ \cos nt \int \Pi dt \cdot \sin nt - \sin nt \int \Pi dt \cdot \cos nt \right\}$$

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as before.

If Π be an explicit function of t , this value of u is always assignable, at least by quadratures, and, whenever the integrations can be executed, in finite terms. In the lunar and planetary theories, Π is always reducible to a series of sines or cosines of the form $\frac{\cos}{\sin} (At + B)$. Let us therefore consider this case more closely.

Now any term of Π , such as $a \times \sin (At + B)$ will introduce into u the term

$$\frac{a}{n} \left\{ \cos nt \int \sin nt \cdot \sin (At + B) dt - \sin nt \int \cos nt \cdot \sin (At + B) dt \right\}$$

But we have

$$\sin nt \cdot \sin (At + B) = \frac{1}{2} \{ \cos (\overline{A - n} \cdot t + B) - \cos (\overline{A + n} \cdot t + B) \}$$

whence

$$\int dt \cdot \sin nt \cdot \sin (At + B) = \frac{\sin (\overline{A - n} \cdot t + B)}{2(A - n)} - \frac{\sin (\overline{A + n} \cdot t + B)}{2(A + n)}$$

and similarly,

$$\int dt \cdot \cos nt \cdot \sin (At + B) = - \frac{\cos (\overline{A + n} \cdot t + B)}{2(A + n)} - \frac{\cos (\overline{A - n} \cdot t + B)}{2(A - n)}$$

So that the term introduced into u will become

$$\begin{aligned} & a \cdot \frac{\cos nt \cdot \sin (\overline{A - n} \cdot t + B) + \sin nt \cdot \cos (\overline{A - n} \cdot t + B)}{2n(A - n)} \\ & - a \cdot \frac{\cos nt \cdot \sin (\overline{A + n} \cdot t + B) - \sin nt \cdot \cos (\overline{A + n} \cdot t + B)}{2n(A + n)} \\ & = a \cdot \sin (At + B) \left\{ \frac{1}{2n(A - n)} - \frac{1}{2n(A + n)} \right\} \\ & = \frac{a \cdot \sin (At + B)}{A^2 - n^2} \end{aligned}$$

Similarly, if $a \cdot \cos (At + B)$ were any term of Π , the corresponding term in u would be $\frac{a \cdot \cos (At + B)}{A^2 - n^2}$

Π therefore consisting of a series of terms, such as

$$\Pi = a \cdot \frac{\cos}{\sin} (At + B) + a' \cdot \frac{\cos}{\sin} (A't + B') + \&c. \quad (106)$$

we shall have

$$u = \frac{a}{A^2 - n^2} \cdot \frac{\cos}{\sin} (At + B) + \frac{a'}{A'^2 - n^2} \cdot \frac{\cos}{\sin} (A't + B') + \&c. \quad (107)$$

$$+ C \cdot \cos nt + C' \cdot \sin nt$$

C and C' being two arbitrary constants.

Terms of the form $a \cdot \frac{\cos}{\sin} (At + B)$ being of perpetual occurrence in physical astronomy, it is necessary to designate them and their several parts by names. The whole term is called an equation, or inequality: the part $(At + B)$ within the sign \sin or \cos is called the *argument*; and the co-efficient a , the maximum. The *period* of the inequality, or the time (in units of time such as t consists of) which it occupies in passing through all its gradations of magnitude and sign, is equal to $\frac{360}{A}$, A being expressed in degrees. Hence, the period of an inequality is longer or shorter, according as the co-efficient of the time in its argument is less or greater.

The arguments of all the inequalities in u then are the same as those in Π , one remarkable case only excepted, in which $A = n$; for, in this case, $A^2 - n^2 = 0$; and the term having $(At + B)$ for its argument, changes its form. In fact, since the constants C and C' are arbitrary, we may change them into $C - \frac{a \cdot \sin B}{A^2 - n^2}$ and $C' - \frac{a \cdot \cos B}{A^2 - n^2}$ respectively, in which case $C \cdot \cos nt + C' \cdot \sin nt$ will be changed to

$C \cdot \cos nt + C' \cdot \sin nt - \frac{a}{A^2 - n^2} \cdot \sin (At + B)$. Thus u will contain the terms

$$\frac{a}{A^2 - n^2} \cdot \sin (At + B) - \frac{a}{A^2 - n^2} \cdot \sin (nt + B)$$

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But when $A = n$, $a \cdot \frac{\sin (A t + B) - \sin (n t + B)}{A^2 - n^2} = \frac{0}{0}$ and differentiating numerator and denominator with respect to A , it becomes simply (on making $A = n$)

$$\frac{a t}{2} \cdot \cos (n t + B) \quad (108)$$

Similarly, if Π contained the term $\cos (n t + B)$, this would introduce into u the term

$$-\frac{a t}{2} \cdot \sin (n t + B) \quad (108)$$

Terms of this kind form an exception to the law of periodicity observed by all the rest, having t disengaged from the sign \sin or \cos . If t represent either the time, or the mean or true longitude, of the disturbed body, they will represent inequalities, whose maxima go on continually increasing without limit. Such inequalities, if they really had an existence in our system, must end in its destruction, or at least in the total subversion of its present state; but we shall see hereafter, that when they do occur, they have their origin, not in the nature of the differential equations, but in the imperfection of our analysis, and in the inadequate representation of the perturbations, and are to be got rid of, or rather included in more general expressions, of a periodical nature, by a more refined investigation than that which led us to them. The nature of this difficulty will be easily understood from the following reasoning. Suppose that a term, such as $a \sin (A t + B)$ should exist in the value of u , in which A being extremely minute, the period of the inequality denoted by it would be of great length; then, whatever might be the value of the co-efficient a , the inequality would still be always confined within certain limits, and after many ages would return to its former state. Suppose now that our peculiar mode of arriving at the value of u , led us to this term, not in its real analytical form $a \cdot \sin (A t + B)$, but by the way of its development in powers of t , $a + \beta t + \gamma t^2 + \&c.$; and that, not at once, but piecemeal, as it were; a first approximation giving us only the term a , a second adding the term βt , and so on. If we stopped here, it is obvious that we should mistake the nature of this inequality, and that a really periodical function, from the effect of an imperfect approximation, would appear under the form of one not periodical.

This is what actually takes place in the theory of the problem of three bodies. These terms in the value of u , when they occur, are not superfluous; they are essential to its expression, but they lead us to erroneous conclusions as to the stability of our system and the general laws of its perturbations, unless we keep in view that they are only parts of series; the principal parts, it is true, when we confine ourselves to intervals of moderate length, but which cease to be so after the lapse of very long times, the rest of the series acquiring ultimately the preponderance, and compensating the want of periodicity of its first terms.

SECTION II.

General theory of the planetary perturbations depending on their mutual configurations.

Our first object in the theory of the planets is to transform the differential equations of the disturbed orbit, so as to obtain final equations in which the radius vector, and true longitude, or those parts of them arising from the action of the disturbing forces, shall be expressed in terms of the time; and to reduce them to the general form of the linear equation of the second order, whose theory we have just considered. To this end,

Let the equations (94) of m 's motion be respectively multiplied by x, y, z , and added together, and we get

$$0 = \frac{d x d^2 x + d y d^2 y + d z d^2 z}{d t^2} + \mu \cdot \frac{x d x + y d y + z d z}{r^3} + m' \cdot \left\{ \frac{d \Omega}{d x} d x + \frac{d \Omega}{d y} d y + \frac{d \Omega}{d z} d z \right\}$$

The portion within the brackets of the last term is the differential of Ω taken on a supposition of the co-ordinates of m only varying. Let this be represented by the Roman character d , so that

$$d \Omega = \frac{d \Omega}{d x} d x + \frac{d \Omega}{d y} d y + \frac{d \Omega}{d z} d z$$

bearing this in mind, and that $d \Omega$ is only an abbreviated expression for this function, we have by integration

$$0 = \frac{d x^2 + d y^2 + d z^2}{d t^2} - \frac{2 \mu}{r} + \frac{\mu}{a} + 2 m' \int d \Omega; \quad (109)$$

in which we must be careful not to confound $\int d \Omega$ with $\int d \Omega$ or Ω , $d \Omega$ being only an incomplete differential, and the characteristic \int denoting an integration relative to t , supposes the variation of the co-ordinates of the disturbing as well as the disturbed body.

Again, if we multiply the same equations (94) by x, y, z , respectively, we shall get

$$0 = \frac{x d^2 x + y d^2 y + z d^2 z}{d t^2} + \frac{\mu}{r} + m' \left\{ x \frac{d \Omega}{d x} + y \frac{d \Omega}{d y} + z \frac{d \Omega}{d z} \right\}.$$

Astronomy If we add this to the former, observing that

$$d x^2 + d y^2 + d z^2 + x a^2 x + y d^2 y + z d^2 z = \frac{1}{2} d^2 (x^2 + y^2 + z^2) = \frac{d^2 \cdot r^2}{2}$$

we get

$$o = \frac{1}{2} \frac{d^2 r^2}{d t^2} - \frac{\mu}{r} + \frac{\mu}{a} + 2 m' \int d \Omega + m' \left\{ x \frac{d \Omega}{d x} + y \frac{d \Omega}{d y} + z \frac{d \Omega}{d z} \right\}$$

Now, if we put $x = \rho \cdot \cos \theta$, $y = \rho \cdot \sin \theta$ and $z = \rho \cdot s$, or suppose ρ = the projected radius vector r , (see fig. 11.) θ the angle it makes with the axis of the x , and s = the tangent of the latitude of m , we have

$$\rho = \frac{r}{\sqrt{1 + s^2}}; r \frac{d \rho}{d r} = r \cdot \frac{1}{\sqrt{1 + s^2}} = \rho$$

But

$$r \frac{d x}{d r} = r \frac{d x}{d \rho} \cdot \frac{d \rho}{d r} = \rho \frac{d x}{d \rho} = \rho \cdot \cos \theta = x$$

and similarly,

$$r \frac{d y}{d r} = y \quad \text{and} \quad r \frac{d z}{d r} = z$$

Consequently, substituting for x, y, z , these values

$$x \frac{d \Omega}{d x} + y \frac{d \Omega}{d y} + z \frac{d \Omega}{d z} = r \left\{ \frac{d \Omega}{d x} \cdot \frac{d x}{d r} + \frac{d \Omega}{d y} \cdot \frac{d y}{d r} + \frac{d \Omega}{d z} \cdot \frac{d z}{d r} \right\} = r \cdot \frac{d \Omega}{d r}$$

and it will be observed, that this property is altogether independent of the nature of the function Ω , and belongs to every possible function of the co-ordinates x, y, z, x', y', z' .

Thus we see, that

$$2 m' \int d \Omega + m' \left\{ x \frac{d \Omega}{d x} + y \frac{d \Omega}{d y} + z \frac{d \Omega}{d z} \right\} = m' \cdot \left\{ 2 \int d \Omega + r \frac{d \Omega}{d r} \right\}$$

Hence, if we put

$$Q = 2 \int d \Omega + r \frac{d \Omega}{d r}, \quad (110)$$

our differential equation becomes

$$o = \frac{1}{2} \frac{d^2 \cdot r^2}{d t^2} - \frac{\mu}{r} + \frac{\mu}{a} + m' Q; \quad (111)$$

Let us now suppose that r represents only the elliptic value of r , and x, y, z, Ω , the elliptic values of the co-ordinates, and the value of the function Ω , which would arise from writing the elliptic values of x, y, z, x', y', z' , in their expressions; and let $r + m' \delta r$, $x + m' \delta x$, $y + m' \delta y$, $z + m' \delta z$, $\Omega + m' \delta \Omega$, &c. represent the disturbed values of these quantities; then, if we neglect m'^2 , we shall get by substitution

$$o = \frac{1}{2} \frac{d^2 \cdot r^2}{d t^2} - \frac{\mu}{r} + \frac{\mu}{a} + m' \cdot \left\{ \frac{d^2 \cdot r \delta r}{d t^2} + \frac{\mu \cdot r \delta r}{r^3} + Q \right\}$$

But since r represents the elliptic value of r , the first part of this equation vanishes of itself, and to determine δr or $r \delta r$, we have the differential equation

$$o = \frac{d^2 \cdot r \delta r}{d t^2} + \frac{\mu}{r^3} \cdot r \delta r + Q; \quad (112)$$

This equation being linear, and of the second order, is immediately integrable by our general formula, equation (100) provided we can find the two particular integrals u' and u'' of the equation

$$\frac{d^2 u}{d t^2} + \frac{\mu}{r^3} u = o$$

but since r , on the supposition of the term Q bearing zero, may be taken for the radius vector on the hypothesis of elliptic motion, it is obvious that the elliptic values of either x , or y , or z , will satisfy this equation, because these values are, in fact, no other than what are derived from the integration of equations precisely similar, viz.

$$\frac{d^2 x}{d t^2} + \frac{\mu}{r^3} x = o$$

$$\frac{d^2 y}{d t^2} + \frac{\mu}{r^3} y = o$$

$$\frac{d^2 z}{d t^2} + \frac{\mu}{r^3} z = o$$

Consequently, we may take $u' = x$ and $u'' = y$, whence we get $u'' \frac{d u'}{u''} = \frac{y d x - x d y}{y} = \frac{h d t}{y}$ because on the hypothesis of elliptic motion $y d x - x d y = h d t$, and it is of the elliptic values of x and y that we are now speaking. Similarly, $u' d \frac{u''}{u'} = \frac{x d y - y d x}{x} = -\frac{h d t}{x}$, and the formula (100) gives

$$C' \cdot x + C'' \cdot y - x \int \frac{Q y d t}{h} + y \int \frac{Q x d t}{h}$$

Astronomy. for the complete value of u or $r \delta r$. The constants then being included under the integral sign, we have

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$$r \delta r = \frac{y \int Q x dt - x \int Q y dt}{h}; \quad (113)$$

Such is the value of $r \delta r$ when we consider only the first power of the disturbing force. It would be absolutely exact, but that x and y are only particular values of u on the hypothesis of r having its elliptic value. If this supposition were not made, we should have $u' = x + m' \delta x$ and $u'' = y + m' \delta y$; and substituting these values, we should obtain terms in the expression of $m' \delta r$ depending on the square of the disturbing force, but with these we have no concern.

The perturbation in longitude ($m' \delta \theta$) is easily obtained when the value of δr is found. In fact, we have

$$d x^2 + d y^2 + d z^2 = d r^2 + r^2 d \theta^2$$

whence we get

$$0 = \frac{r^2 d \theta^2 + d r^2}{d t^2} - \frac{2 \mu}{r} + \frac{\mu}{a} + 2 m' \int d \Omega \quad (114)$$

and if we subtract this from the equation (111), we find

$$0 = \frac{r^2 d \theta^2}{d t^2} - \frac{r d^2 r}{d t^2} - \frac{\mu}{r} - m' \cdot r \frac{d \Omega}{d r} \quad (115)$$

If in this we substitute $r + m' \delta r$ for r and $\theta + m' \delta \theta$ for θ we get, (after obliterating the terms which destroy each other by reason of the properties of elliptic motion, and those which contain m'^2)

$$0 = \frac{2 r^2 d \theta d \delta \theta}{d t^2} + \frac{2 r \delta r d \theta^2}{d t^2} - \frac{r d^2 \delta r}{d t^2} - \frac{d^2 r \cdot \delta r}{d t^2} + \frac{\mu r \delta r}{r^3} - r \frac{d \Omega}{d r}; \quad (116)$$

From this, let the term multiplied by $\frac{d \theta^2}{d t^2}$ be eliminated by means of equation (115) and we get

$$\frac{2 r^2 d \theta d \delta \theta}{d t^2} = \frac{r d^2 \delta r - \delta r \cdot d^2 r}{d t^2} - \frac{3 \mu r \delta r}{r^3} + r \frac{d \Omega}{d r}; \quad (117)$$

And if in this we substitute for $\frac{\mu r \delta r}{r^3}$ its value given by the equation (112) we obtain, (restoring the value of Q)

$$d \delta \theta = \frac{d \{ d r \delta r + 2 r d \delta r \} + \left(3 \int d \Omega + 2 r \frac{d \Omega}{d r} \right) \cdot d t^2}{r^2 d \theta}; \quad (118)$$

but $r^2 d \theta = h d t$, elliptic values only being considered in the second member of this equation, and consequently integrating, we have

$$h \delta \theta = \frac{d r}{d t} \delta r + 2 r \frac{d \delta r}{d t} + \int \left(3 \int d \Omega + 2 r \frac{d \Omega}{d r} \right) d t; \quad (119)$$

Now, we have $h = \sqrt{\mu a (1 - e^2)}$, and if we put $n t$ for the mean motion of the disturbed planet m , we have

$$n = \sqrt{\frac{\mu}{a^3}} \text{ so that } h = n a^2 \cdot \sqrt{1 - e^2}; \text{ and } \frac{1}{h} = \frac{n a}{\mu \sqrt{1 - e^2}}$$

Consequently we get, for the perturbation of the radius vector,

$$m' \delta r = \frac{m' a \left\{ \cos \theta \int r \cdot \sin \theta \cdot Q n dt - \sin \theta \int r \cdot \cos \theta \cdot Q n dt \right\}}{\mu \sqrt{1 - e^2}} \quad (120)$$

and the formula expressing the perturbation in longitude, will become

$$m' \delta \theta = \frac{m'}{n a^2 \sqrt{1 - e^2}} \left(\frac{d r}{d t} \delta r + 2 r \frac{d \delta r}{d t} \right) + \frac{a m'}{\mu \sqrt{1 - e^2}} \left\{ 2 \int r \frac{d \Omega}{d r} \cdot n dt + 3 \iint d \Omega \cdot n dt \right\}; \quad (121)$$

It only remains to determine the amount of the perturbation in latitude, or the value of z or of δs , if we put $z = r \delta s$, in which case δs represents the tangent of the heliocentric latitude of m in its disturbed orbit above the plane of its elliptic motion. Now, if we treat the equation

$$0 = \frac{d^2 z}{d t^2} + \frac{\mu z}{r^3} + m' \frac{d \Omega}{d z}$$

in the same manner as the equation by which the value of $r \delta r$ was found, viz. regarding x and y as two particular integrals of the equation $\frac{d^2 z}{d t^2} + \frac{\mu z}{r^3} = 0$, and then completing the integration by the general

Astronomy formula (100)* we shall find (putting $r \cdot \cos \theta$ and $r \cdot \sin \theta$, for x and y)

$$m' \delta s = \frac{m' a \left\{ \cos \theta \int y \frac{d\Omega}{dz} \cdot n dt - \sin \theta \int x \frac{d\Omega}{dz} \cdot n dt \right\}}{\mu \sqrt{1 - e^2}} \quad (122)$$

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This is the latitude of m above its primitive orbit; and if we denote by s its undisturbed latitude above any fixed plane (as that of the ecliptic) slightly inclined to this orbit, $s + \delta s$ will be its latitude when subjected to the action of the disturbing forces.

The equation (121) gives the perturbation in longitude when that of the radius vector is known, and the latter may be computed from the expression (120) which is general; and considering the complication of the subject, as simple as can be expected. Its form enables us to compute the amount of perturbation even in the most difficult cases, as in that of a comet, by the application of the method of quadratures. Meanwhile, in the theory of the planets, where it is required to develop the value of δr in series of sines and cosines of arcs depending on the configurations of the disturbed and disturbing planet, it will be found much simpler to set out immediately from the differential equations for the disturbed radius, and proceed in the manner now to be explained.

Since the form of this equation is not precisely that of the equation $\frac{d^2 u}{dt^2} + n^2 u + \Pi = 0$, the coefficient of the second term instead of being constant, being $\frac{\mu}{r^3}$ a variable quantity, we must first endeavour to transform it by substitution into one of this form. Assuming then that u is such a quantity that its elliptic value shall satisfy the equation $\frac{d^2 u}{dt^2} + n^2 u = 0$, and its disturbed value (or $u + m' \delta u$) the equation $\frac{d^2 (u + m' \delta u)}{dt^2} + n^2 (u + m' \delta u) + m' \Pi = 0$, which gives $\frac{d^2 \delta u}{dt^2} + n^2 \delta u + \Pi = 0$ we must inquire first, the relation between r and u ; and secondly, the value of Π .

A satisfactory relation between r and u is easily found. In fact, since u is to satisfy $\frac{d^2 u}{dt^2} + n^2 u = 0$, it must be of the form $u = \text{const.} \cos (nt + \text{const.})$. Now the development of r in terms of the mean longitude gives, putting ϵ for the longitude at the epoch of the commencement of the time t , and π for the longitude of the perihelion,

$$r = a \left\{ 1 - e \cdot \cos (nt + \epsilon - \pi) + e^2 \sin^2 (nt + \epsilon - \pi) + \&c. \right\}$$

So that, if we take $u = e \cdot \cos (nt + \epsilon - \pi)$, we shall have

$$r = a \left\{ (1 + e^2) - u (1 - \frac{3}{2} e^2) - u^2 + \&c. \right\} \quad (123)$$

This gives at once

$$\begin{aligned} \delta r &= -a \delta u \left\{ 1 + 2u + e^2 \times \&c. \right\} \\ &= -a \delta u \left\{ 1 + 2e \cdot \cos (nt + \epsilon - \pi) + e^2 \cdot \&c. \right\}; \end{aligned} \quad (124)$$

by which, when δu is found, δr may be had at once.

It only remains to discover Π . Now, since our equation (111), if we put $r^3 = v$ and $m' \left\{ 2 \int d\Omega + r \frac{d\Omega}{dr} \right\} = m' Q$, becomes

$$\frac{d^2 v}{dt^2} = \frac{2\mu}{\sqrt{v}} - \frac{2\mu}{a} - 2m' Q \quad (125)$$

if we multiply by $2 dv$, and integrate, we shall find

$$\left(\frac{dv}{dt} \right)^2 = 8\mu \sqrt{v} - \frac{4\mu v}{a} - 4m' \int Q dv \quad (126)$$

but because u is a function of r , and therefore of r^2 or v , we have

$$\begin{aligned} \frac{du}{dt} &= \frac{du}{dv} \cdot \frac{dv}{dt} \\ \frac{d^2 u}{dt^2} &= \frac{d^2 u}{dv^2} \left(\frac{dv}{dt} \right)^2 + \frac{du}{dv} \cdot \frac{d^2 v}{dt^2} \end{aligned}$$

so that

$$\begin{aligned} \frac{d^2 u}{dt^2} + n^2 u &= \frac{d^2 u}{dv^2} \left\{ 8\mu \sqrt{v} - \frac{4\mu v}{a} - 4m' \int Q dv \right\} \\ &+ \frac{du}{dv} \left\{ \frac{2\mu}{\sqrt{v}} - \frac{2\mu}{a} - 2m' Q \right\} + n^2 u \end{aligned}$$

Now u is a certain function of v (or of r) whose form is determined by the reversion of the series in (123) and is independent of the disturbing forces. But were these forces zero, we should have $\frac{d^2 u}{dt^2} + n^2 u = 0$.

* This formula, which perhaps is new, and which has stood us in some stead in the explanation of that chapter of the *Mécanique Céleste*, (cap. vi. liv. 2.) which I have adopted for the groundwork of this part of the present essay, may be deduced at once from the general theory of linear equations, in my paper *On various points of Analysis*.—*Phil. Trans.* 1814.

Astronomy. Hence the portion of the right hand member of the equation just deduced, which does not depend on the disturbing forces, must be identically zero, in virtue of the relation between u and v ; and that it is so, we may assure ourselves by actual substitution. Consequently we must have, when the disturbing forces are regarded, **Physical Astronomy**

$$\frac{d^2 u}{dt^2} + n^2 u = \left(-2Q \frac{du}{dv} - 4 \int Q dv \times \frac{d^2 u}{dv^2} \right) \times m'$$

that is (putting $u + m' \delta u$ for u , and disregarding terms depending on m'^2)

$$\frac{d^2 \delta u}{dt^2} + n^2 \delta u = -2Q \frac{du}{dv} - 4 \frac{d^2 u}{dv^2} \int Q dv$$

Comparing this with the equation $\frac{d^2 \delta u}{dt^2} + n^2 \delta u + \Pi = 0$ we have

$$\Pi = 2Q \cdot \frac{du}{dv} + 4 \frac{d^2 u}{dv^2} \int Q dv$$

It is desirable to express this in terms of v . Now, as u is a function of v , v is reciprocally a function of u , and

$$\frac{du}{dv} = \frac{1}{\frac{dv}{du}}; \quad \frac{d^2 u}{dv^2} = - \frac{\frac{d^2 v}{du^2}}{\left(\frac{dv}{du} \right)^3}$$

and

$$\Pi = \frac{2}{\frac{dv}{du}} Q - 4 \frac{\frac{d^2 v}{du^2}}{\left(\frac{dv}{du} \right)^3} \int Q dv \quad (127)$$

but since $v = r^2$, if for r we put its value in terms of u (123), we get, neglecting higher powers of u than the first, $v = a^2 (1 - 2u + \&c.)$ and substituting this in Π , and after the differentiations writing for u its value $e \cdot \cos (nt + e - \pi)$ we get

$$\begin{aligned} \Pi = & -\frac{Q}{a^2} \left(1 - e \cdot \cos (nt + e - \pi) \right) - \frac{1}{4} e^2 \cdot \cos (2nt + 2e - 2\pi) \\ & - \frac{2e}{a^2} \int Q n dt \cdot \sin (nt + e - \pi) \{ 1 + e \cdot \cos (nt + e - \pi) \} \end{aligned} \quad (128)$$

and this being the value of Π , we find δu from the equation

$$\frac{d^2 \delta u}{dt^2} + n^2 \delta u + \Pi = 0. \quad (129)$$

SECTION III.

Reduction of the perturbative function Q or $2 \int d\Omega + r \frac{d\Omega}{dr}$ to a series of sines and cosines, and investigation of the perturbations, neglecting the eccentricities and inclinations of both orbits.

WE have now reduced the investigation of the perturbation to the integration of the linear equation (129) and we have before seen that this is accomplished without difficulty, when Π , the last term, is reducible to sines and cosines of the independent variable and its multiples. All then that remains to be done to get their actual expressions, is to execute this reduction. This, however, is by no means a simple process; and in an essay like the present it is not possible to pursue it into all its details: we shall therefore only carry it to a certain extent necessary for our future reference, and point out the principles by which it may be, if required, carried farther.

Let us consider then the value of the function Ω . If we write in it $r \cdot \cos \theta, r' \cdot \sin \theta, r' \cdot \cos \theta'$ and $r' \cdot \sin \theta'$ for x, y, x', y' , and neglect $zz', z'^2, (z - z')^2$, which are either of the order of the squares of the disturbing forces, or of the products of these forces by the mutual inclination of the orbits, and put $\theta - \theta' = w$, we get

$$xx' + yy' + zz' = rr' \cdot \cos w$$

$$\lambda = \sqrt{(x - x')^2 + (y - y')^2 + (z - z')^2} = \sqrt{r^2 - 2rr' \cdot \cos w + r'^2}$$

and

$$\Omega = \frac{r}{r^2} \cdot \cos w - \frac{1}{\sqrt{r^2 - 2rr' \cdot \cos w + r'^2}}$$

Let us conceive this function developed in a series of cosines of w , and its positive and negative multiples to infinity; then, since the cosines of the negative are equal to those of the positive multiples, we may represent Ω as follows:

$$\Omega = R + R' \cdot \cos w + R'' \cdot \cos 2w + R''' \cdot \cos 3w + \&c.; \quad (130)$$

Astronomy. where $R, R', R'',$ &c. are certain given, explicit functions of r and r' and of these alone, depending solely on the peculiar form of Ω , let $A, A', A'',$ &c. represent the same functions of a and a' . Then, if the eccentricities of the orbits were nothing, we should have **Physical Astronomy**

$$\Omega = \frac{a}{a'^2} \cdot \cos w - \frac{1}{\sqrt{a^2 - 2aa' \cos w + a'^2}} \left. \vphantom{\frac{a}{a'^2}} \right\} ; \dots \dots \dots (131)$$

$$= A + A' \cdot \cos w + A'' \cdot \cos 2w + \&c.$$

Were the eccentricities nothing, the orbits would be circles, and the motion in them uniform. We should therefore have $\theta = nt$ and $\theta' = n't$, whence $w = \theta - \theta' = (n - n')t$, so that Ω in this case would be expressed in the very simple series

$$\Omega = A + A' \cdot \cos (n - n')t + A'' \cdot \cos 2(n - n')t + \&c. \quad (132)$$

Moreover, since $d\Omega$ represents the differential of Ω taken on the hypothesis that only the disturbed body moves, we should then have

$$d\Omega = \frac{d\Omega}{dr} dr + \frac{d\Omega}{d\theta} d\theta = \frac{d\Omega}{d\theta} d\theta = n dt \cdot \frac{d\Omega}{d\theta}$$

because, in the case of circular orbits $dr = 0$. Thus we should have

$$d\Omega = -n dt \cdot \{A' \cdot \sin (n - n')t + 2 \cdot A'' \cdot \sin (2n - 2n')t + \&c.\}$$

and integrating relative to t

$$\int d\Omega = \frac{g}{a} + \frac{n}{n - n'} \{A' \cdot \cos (n - n')t + A'' \cdot \cos 2(n - n')t + \&c.\}; \quad (133)$$

$\frac{g}{a}$ being an arbitrary constant.

Again, since $R, R', R'',$ &c. are explicit functions of r, r' , and Ω is only so far a function of r , as this symbol is contained in them, we must have

$$r \frac{d\Omega}{dr} = r \frac{dR}{dr} + r \frac{dR'}{dr} \cdot \cos w + \&c. \quad (134)$$

and in the case of circular orbits, denoting by $\frac{dA}{da}$, &c. the same functions of a, a' that $\frac{dR}{dr}$ denote of r, r' ,

$$r \frac{d\Omega}{dr} = a \frac{dA}{da} + a \frac{dA'}{da} \cdot \cos w + \&c.$$

The values of $\frac{dA}{da}$, $\frac{dA'}{da}$, &c. are easily had in functions of a, a' , when those of $A, A',$ &c. are found. Now, to obtain these, we may proceed as follows:

Take $c = \cos w + \sqrt{-1} \cdot \sin w$. Then (by trigonometry) we shall have $\frac{1}{c} = \cos w - \sqrt{-1} \cdot \sin w$; now, let us consider the function $(a^2 - 2aa' \cos w + a'^2)^{-s}$ which agrees with the second term of Ω if $s = \frac{1}{2}$. This equals $a^{-2s} \left(1 - 2 \frac{a'}{a} \cos w + \left(\frac{a'}{a}\right)^2\right)^{-s}$ or $a^{-2s} (1 - 2a \cos w + a^2)^{-s}$ putting $a = \frac{a'}{a}$. But we have

$$1 - 2a \cos w + a^2 = (-ac) \left(1 - a \cdot \frac{1}{c}\right)$$

because $c + \frac{1}{c} = 2 \cos w$. Hence

$$\begin{aligned} (a^2 + 2aa' \cos w + a'^2)^{-s} &= a^{-2s} \cdot (1 - ac)^{-s} \cdot \left(1 - \frac{a}{c}\right)^{-s} \\ &= a^{-2s} \times \left\{1 + \frac{s}{1} \cdot ac + \frac{s(s+1)}{1 \cdot 2} a^2 c^2 + \&c.\right\} \times \left\{1 + \frac{s}{1} \cdot \frac{a}{c} + \frac{s(s+1)}{1 \cdot 2} \cdot \frac{a^2}{c^2} + \&c.\right\} \\ &= a^{-2s} \left\{1 + \left(\frac{s}{1}\right)^2 \cdot a^2 + \left(\frac{s(s+1)}{1 \cdot 2}\right)^2 a^4 + \&c.\right\} \\ &+ a^{-2s} \left\{\frac{s}{1} a + \frac{s(s+1)}{1 \cdot 2} \cdot \frac{s}{1} a^3 + \frac{s(s+1)(s+2)}{1 \cdot 2 \cdot 3} \times \frac{s(s+1)}{1 \cdot 2} a^5 + \&c.\right\} \left(c + \frac{1}{c}\right) \\ &+ a^{-2s} \left\{\frac{s(s+1)}{1 \cdot 2} a^3 + \frac{s(s+1)(s+2)}{1 \cdot 2 \cdot 3} \times \frac{s}{1} \cdot a^5 + \&c.\right\} \left(c^2 + \frac{1}{c^2}\right) \\ &+ \&c. \end{aligned}$$

but $c + \frac{1}{c} = 2 \cos w$; $c^2 + \frac{1}{c^2} = 2 \cos 2w$, &c. (by trigonometry); consequently we have

$$\begin{aligned}
 (a^2 - 2 a a' \cos w + a'^2)^{-s} &= a^{-2s} \left\{ 1 + \left(\frac{s}{1}\right)^2 a^2 + \&c. \right\} \\
 &+ 2a^{-2s} \left\{ \frac{s}{1} a + \frac{s(s+1)}{1 \cdot 2} \cdot \frac{s}{1} a^3 + \&c. \right\} \cos w \\
 &+ 2a^{-2s} \left\{ \frac{s(s+1)}{1 \cdot 2} a^2 + \&c. \right\} \cos 2w + \&c.
 \end{aligned}
 \tag{135}$$

In the case before us, where $s = \frac{1}{2}$, this gives at once

$$A = \frac{-1}{a} \left\{ 1 + \left(\frac{1}{2}\right)^2 a^2 + \left(\frac{1 \cdot 3}{2 \cdot 4}\right)^2 a^4 + \&c. \right\} \tag{136, 1}$$

$$A' = \frac{-2}{a} \left\{ \frac{1}{2} a + \frac{1^2 \cdot 3}{2^2 \cdot 4} a^3 + \frac{1^2 \cdot 3^2 \cdot 5}{2^2 \cdot 4^2 \cdot 6} a^5 + \&c. \right\} + \frac{a}{a'^2} \tag{136, 2}$$

$$A'' = \frac{-2}{a} \left\{ \frac{1 \cdot 3}{2 \cdot 4} a^2 + \frac{1^2 \cdot 3 \cdot 5}{2^2 \cdot 4 \cdot 6} a^4 + \&c. \right\} \tag{136, 3}$$

If a be less than unity, these series are convergent; but if greater, we have only to throw the expression into the form $a'^{-2s} \left(1 - 2 \frac{a}{a'} \cos w + \left(\frac{a}{a'}\right)^2 \right)^{-s}$ previous to development, and taking $a = \frac{a}{a'}$, instead of $\frac{a'}{a}$, a will now be less than 1, and we shall have $A = \frac{1}{a'} \left(1 + \left(\frac{1}{2}\right)^2 a^2 + \&c. \right) \&c.$

In the former case, when $\frac{a}{a'}$ is less than 1, or the orbit of the disturbed planet is interior to that of the disturbing, we have

$$-A = \frac{1}{a} + \left(\frac{1}{2}\right)^2 \cdot \frac{a'^2}{a^3} + \left(\frac{1 \cdot 3}{2 \cdot 4}\right)^2 \cdot \frac{a'^4}{a^5} + \&c.; \tag{137, 1}$$

$$\frac{dA}{da} = + \frac{1}{a^2} \left\{ 1 + 3 \cdot \left(\frac{1}{2}\right)^2 \cdot \left(\frac{a'}{a}\right)^2 + \&c. \right\}; \tag{137, 2}$$

$$A' = -2 \left\{ \frac{1}{2} \frac{a'}{a^2} + \frac{1^2 \cdot 3}{2^2 \cdot 4} \cdot \frac{a'^3}{a^4} + \&c. \right\} + \frac{a}{a'^2}; \tag{137, 3}$$

$$\frac{dA'}{da} = + \frac{2}{a^2} \left\{ 1 \cdot \frac{a'}{a} + \frac{1^2 \cdot 3}{2^2} \cdot \left(\frac{a'}{a}\right)^3 + \&c. \right\} + \frac{1}{a'^2} \tag{137, 4}$$

and so on; and similar expressions are also readily obtained in the case when the orbit of the disturbed planet is the exterior. Thus, when the mean distances of the two planets are given, the values of A , A' , &c. and their differential co-efficients $\frac{dA}{da}$, &c. are reducible to numerical evaluation, and may therefore be regarded as known quantities. The properties of the series on which they depend, afford many resources for facilitating their evaluation, and rules for deriving one of these quantities from another, but these we shall not stay to explain. The reader will find them with every development in the second book of Laplace's *Mécanique Céleste*, art. 49.

These values once determined, we have Q , or $2 \int d\Omega + r \frac{d\Omega}{d\tau}$, expressed as follows:

$$\begin{aligned}
 Q = \frac{2g}{a} + a \frac{dA}{da} + \left\{ a \frac{dA'}{da} + \frac{2nA'}{n-n'} \right\} \cos(n-n')t \\
 + \left\{ a \frac{dA''}{da} + \frac{2nA''}{n-n'} \right\} \cos 2(n-n')t + \&c. \left. \vphantom{\frac{2g}{a}} \right\} \dots \dots \dots \tag{138, 1}
 \end{aligned}$$

Such is the value of the perturbative function when the eccentricities and inclinations of the orbits are neglected. Let us, for the present, confine ourselves to this case; and, writing M , M' , M'' , &c. for the successive co-efficients, we have

$$Q = M + M' \cos(n-n')t + M'' \cos 2(n-n')t + \&c. \tag{138, 2}$$

Now the equation (128) gives, when $e = 0$,

$$\Pi = -\frac{Q}{a^2} = -n^2 a Q \left(\text{because } n^2 = \frac{1}{a^3} \right) \text{ or,}$$

$$\Pi = -n^2 a \{ M + M' \cos(n-n')t + \&c. \}$$

so that the differential equation in δu becomes

$$0 = \frac{d^2 \delta u}{dt^2} + n^2 \delta u - n^2 a \{ M + M' \cos(n-n')t + \&c. \} \tag{139}$$

Astronomy. and integrating,

$$\delta u = a M - \frac{n^2 a M'}{(n - n')^2 - n^2} \cdot \cos (n - n') t - \frac{n^2 a M''}{4 (n - n')^2 - n^2} \cos 2 (n - n') t - \&c.$$

Consequently, since by equation (124) $\delta r = -a \delta u$ when $e = 0$,

$$\delta r = -a^2 \cdot M + \frac{n^2 a^2 M'}{(n - n')^2 - n^2} \cdot \cos (n - n') t - \frac{n^2 a^2 M''}{4 (n - n')^2 - n^2} \cdot \cos 2 (n - n') t + \&c. \quad (140)$$

The value of δr thus obtained, $\delta \theta$ is easily got from (121); for in this case $\frac{dr}{dt} = 0$,

$$2r \frac{d\delta r}{dt} = -2n^2 a^3 (n - n') \cdot \Sigma \frac{i M^i}{i^3 (n - n')^2 - n^2} \cdot \sin i w$$

where Σ is used to express the sum of all similar terms from $i = 1$ to $i = \infty$ inclusive, and M^i represents the i^{th} in order of the co-efficients M' , M'' , M''' , &c. Moreover

$$\int r \frac{d\Omega}{dr} n dt = a \frac{dA}{da} \cdot n t + \Sigma \frac{a n}{i (n - n')} \frac{dA^i}{da} \cdot \sin i w$$

$$\int n dt \int d\Omega = \frac{g}{a} \cdot n t + \frac{n^2}{(n - n')^2} \Sigma \frac{A^i}{i} \sin i w$$

Uniting therefore these several parts, we get

$$\delta \theta = \frac{1}{n a^2} \times 2r \frac{d\delta r}{dt} + 2a \int r \frac{d\Omega}{dr} \cdot n dt + 3a \int n dt \int d\Omega$$

$$= \left(3g + 2a^2 \frac{dA}{da} \right) \times n t$$

$$+ \Sigma \left\{ \frac{3a n^2 A^i}{i (n - n')^2} + \frac{2a^2 n}{i (n - n')} \frac{dA^i}{da} - \frac{2na (n - n') i}{i^3 (n - n')^2 - n^2} M^i \right\} \sin i w$$

Now, first, since nt represents the mean longitude of m as deduced from observation, the quantity n is already affected with the whole influence of the planetary perturbation, and consequently the part multiplied by nt in this expression of the perturbation in longitude, and which, if allowed to remain, would express an additional perturbation, is superfluous. This furnishes the condition

$$g = -\frac{2}{3} a^2 \frac{dA}{da}; \quad (141)$$

which determines the constant g . Moreover, in the latter part of this expression, if we write for M^i its value

$$M^i = a \cdot \frac{dA^i}{da} + \frac{2nA^i}{n - n'}$$

it will admit reductions, and the value of $\delta \theta$ will at length be found as follows:

$$\delta \theta = \Sigma \left\{ \frac{2n^3 a^2 \cdot \frac{dA^i}{da}}{i (n - n') (i^2 (n - n')^2 - n^2)} + \frac{(3n^2 + i^2 (n - n')^2) \cdot n^2 a A^i}{i (n - n')^2 (i^2 (n - n')^2 - n^2)} \right\} \sin i w \quad (142)$$

If these expressions of δr and $\delta \theta$ be each multiplied by m' , we have the values of $m' \delta r$ and $m' \delta \theta$ the perturbations of the radius vector and the longitude, *i. e.* those parts of them which are independent of the eccentricities of the orbits. These expressions give room for some remarks. The perturbation in longitude as we observe is wholly periodical and dependent on a single angle w and its multiples. In forming then a table of the values of $m' \delta \theta$, the numerical co-efficients being computed, and the value of $m' \delta \theta$ thus reduced to the form $p \cdot \sin w + q \cdot \sin 2w + \&c.$ we may include the whole of this in one column, entered under the general argument w , instead of regarding it as consisting of an infinite number of separate inequalities.

The same remark extends to the periodical part of $m' \delta r$, its arguments are the same as in the formula for $m' \delta \theta$; but besides this, δr includes, as we have seen, a constant part $-a^2 \cdot M$ or $-2ga \cdot a^3 \cdot \frac{dA}{da}$ which becomes by substituting for g its value in (141)

$$\text{const. part of } \delta r = \frac{1}{3} a^3 \cdot \frac{dA}{da},$$

and therefore

$$\delta r = \frac{a^3}{3} \cdot \frac{dA}{da} + n^2 a^2 \cdot \Sigma \frac{1}{i^3 (n - n')^2 - n^2} \left(a \frac{dA^i}{da} + \frac{2nA^i}{n - n'} \right) \cos i w; \quad (143)$$

In the formation of a table of $m' \delta r$ this constant part is of course included with the variable one, but the effect is remarkable. It appears that the action of the disturbing planet alters the mean distance from the sun of the disturbed, and, of course, its mean motion and periodical time from what they would have been had the disturbing planet no existence. At the same time, it will be demonstrated in the following pages, that

Astronomy. these alterations once produced, are permanent and unchangeable in their quantity by the subsequent Physical Astronomy.
actions of the bodies composing the system.

The angle w is the difference of longitudes of the two planets, or their heliocentric elongation from each other. If we call ϵ and ϵ' their epochs, or their actual longitudes at the commencement of the time t , we have

$$\begin{aligned} w &= n t + \epsilon - (n' t + \epsilon') \\ &= n t - n' t + \epsilon - \epsilon' \end{aligned}$$

and this, in fact, is the argument of the perturbations when we neglect the eccentricities and inclinations of the orbits.

Let us next examine the perturbation in latitude, we have

$$\frac{d\Omega}{dz} = \frac{z'}{r^3} - \frac{z' - z}{\lambda^3}$$

Hence, it is evident, that if we neglect the inclinations and eccentricities, we have $\frac{d\Omega}{dz} = 0$, and the plane of the disturbed orbit does not change.

We have thus determined the effect of the action of a third body on the orbit and motion of m , on the simplest supposition, and our results (to recapitulate them) amount to this.

1st. That the radius vector undergoes a permanent change in its mean value, and, of course, that the period and mean motion of m are permanently altered.

2d. That the elliptic value of the radius vector receives an accession of terms, of the form

$$p + q \cdot \cos w + r \cdot \cos 2w + s \cdot \cos 3w + \&c.$$

and that of the true longitude, a series of terms of the form

$$q' \cdot \sin w + r' \cdot \sin 2w + s' \cdot \sin 3w + \&c.$$

w being the difference of longitudes, or mutual elongation of the planets, from each other.

3dly. That to express the several co-efficients of these formulæ in numbers, nothing more is required than a knowledge of the mass of the disturbing planet, and the mean distances and mean motions of both.

In the cases then where the disturbing planet has satellites, as in those of Jupiter, Saturn, and Uranus, the mass is known, and the reduction of the formulæ to numbers is complete. It is fortunate that these are by far the most considerable bodies of our system, but proximity to a certain extent supplies the place of intrinsic energy; and, in the case of the perturbations of the earth, our uncertainty of the masses of Mars and Venus leaves us in some degree at a loss. But physical astronomy furnishes us in this dilemma with considerable aid. Regarding the masses of these planets as unknown quantities, we may calculate in general terms their effect, either on the places of the other planets at assigned instants, or, on the elements of their orbits themselves, which are susceptible of much more accurate determination, by bringing a long series of observations to bear on them, and comparing the variations in their values after long intervals, as computed by theory, and as given by observation, we obtain data for the determination of these delicate quantities, so much the more precise as the variations observed in the elements are greater, or, in other words, as the interval of time in which they are observed is longer. It is thus that the lapse of ages is necessary to give precision to the numerical data of our system, and that continual and patient observation must ultimately lead us to the knowledge of points which elude the direct cognizance of our senses, and defy any investigation but those in which successive generations of mankind bear a part.

In fact, if we regard the masses of all the planets as unknown quantities, but their mean distances and periodic times as known ones;—the latter afford us the means of computing the values of δr and $\delta \theta$ in any assigned case, independent of the value of m' , which does not enter into their expressions. Let us therefore represent by $m' \delta' r$ and $m' \delta' \theta$, the perturbations of the radius vector and longitude produced by the planet m' ; by $m'' \delta'' r$ and $m'' \delta'' \theta$ those produced by m'' , and so on. Then will the true values of these quantities at any assigned instant be

$$\begin{aligned} r + m' \delta' r + m'' \delta'' r + m''' \delta''' r + \&c. \\ \theta + m' \delta' \theta + m'' \delta'' \theta + m''' \delta''' \theta + \&c. \end{aligned}$$

in which $r, \delta' r, \delta'' r, \&c.$ and $\theta, \delta' \theta, \delta'' \theta, \&c.$ are quantities susceptible of calculation from theory. Suppose now that we construct tables of the values of $\theta, \delta' \theta, \delta'' \theta, \&c.$ (or, as we will for a moment write these latter quantities, $\theta, \phi, \psi, \&c.$) then, at any assigned instant, we have only to take out of these tables the values of $\theta, \phi, \psi, \&c.$; and the true longitude of m will be

$$\theta + m' \phi + m'' \psi + \&c. = L$$

Suppose now we compare this formula with a great multitude of observations, and thus obtain a series of equations,

$$\begin{aligned} m' \cdot \phi_1 + m'' \cdot \psi_1 + m''' \cdot \chi_1 + \&c. &= L_1 - \theta_1 \\ m' \cdot \phi_2 + m'' \cdot \psi_2 + m''' \cdot \chi_2 + \&c. &= L_2 - \theta_2 \\ \&c. &= \&c. \end{aligned}$$

The only unknown quantities in these will be the masses of the disturbing planets $m', m'', \&c.$ and by resolving these, we may determine their values, and thus estimate the masses of the planets by the perturbations they produce.

In this, as in almost all such delicate inquiries, where the quantities to be determined are exceedingly small, and the observations from which they are to be discovered liable to inaccuracies, bearing a sensible proportion

Astronomy. to the thing observed, (which in this case is $L - \theta$, or the total perturbation arising from the united action of all the planets,) we are obliged to employ a great many more observations than would be, mathematically speaking, sufficient, if each were perfect, with a view to destroy the errors of observation in the mean result. The number of disturbing planets in our system at present known does not exceed ten; and it would therefore appear that ten observations of the longitude of one disturbed planet would enable us to determine the masses of all the rest; and so they would, were the observations mathematically exact, the elements of the orbits exactly known, and the theory by which the values of ϕ , ψ , χ , &c. are computed, complete. But each of these conditions is far from being fulfilled in the present state of astronomy; and if we would use this method to determine the masses of the planets, we must accumulate many hundreds of observations made, not on one, but on all of them, especially on those subject to the greatest perturbations.

The method of treating a series of equations more numerous than the unknown quantities they contain, so as to give them all their proper influence on the result, and obtain from them a set of values which, though satisfying neither of them separately, yet when substituted in all of them, shall, on the whole, give more satisfactory results than any other, depends on the theory of probabilities and may be found in Laplace's *Theorie Analytique des Probabilités*.

If the mass of any one or more of the planets (m') for instance, be regarded as sufficiently known from other methods, we need only employ this mode for the rest, and regarding the perturbation $m' \delta' \theta$ produced by it as known, place it on the known side of the equation, which will thus become

$$m'' \delta'' \theta + m''' \delta''' \theta + \&c. = L - \theta - m' \delta' \theta$$

Thus we may determine for instance, the masses of Mars and Venus, by means of an extensive series of observations of the sun's longitude, or (which is the same thing) by employing to that end the perturbations they produce on the earth. For the masses of Jupiter, Saturn, and Uranus, being known from the periods of their satellites; and those of Mercury, and the four new planets—Ceres, Pallas, Juno, and Vesta, being so small, as to have little or no influence, we have only two unknown quantities (m' , m'') to determine.

This method is laborious, certainly; but considering the perfection of modern observations, the great multitude of them which may be brought to bear upon this point, and the considerable degree of exactness which the theory of the planetary perturbations has now attained, it is not impossible that it may one day be made to render the best service in determining the masses even of those planets which have satellites. At all events, it is highly desirable that it should be applied for that purpose, as its results would lead us to judge how far the latter method can be depended on in cases like that of Jupiter and Saturn, where the great deviation from

sphericity of the central body renders the application of the formula $t = \frac{2\pi \times a^{\frac{3}{2}}}{\sqrt{M+m}}$ somewhat liable to error.

In fact, this formula is derived on the hypothesis of a force represented by $\frac{M+m}{(\text{dist})^2}$ but in the case of spherical bodies only does the total attraction vary precisely in that ratio*. This alone, however, will not account for the great difference which Mr. Gauss has lately found between the mass of Jupiter, as obtained from observations of its satellites, and that deduced from the perturbations of the small planets intermediate between Jupiter and Mars, so that the subject must be regarded as open to further investigation, should the calculations of the last named eminent geometer be found to coincide with a more extensive series of observations of those interesting bodies than the shortness of the time they have been known has hitherto allowed.

SECTION IV.

Of the method of taking into account the effect of the eccentricities of the orbits on the planetary perturbations, and of the origin of the secular equations of their motions.

WHEN we regard the orbits as elliptical, the whole of the foregoing investigations require modification, the value of the perturbative function, and, of course, of the perturbations themselves, receiving accessions of terms depending on the powers and products of the eccentricities. We will here endeavour to explain the manner in which these terms originate; and to a certain (though limited) extent, the course pursued by geometers in determining their form and value.

The functions Ω , $\int d\Omega$, $r \frac{d\Omega}{dr}$, are explicitly given in terms of r , r' , and w or $\theta - \theta'$, and contain no other symbols. Hence it arose, that when r , r' were supposed constant, the only cause of the variation of these functions consisted in that of w ; and θ and θ' being in this case each expressed by an arc proportional to the time, it was sufficient to develop them in cosines and sines of w , to have at once the expression of the function Π in such a form as we required for integrating our equations. When, however, the eccentricities are introduced, all these facilities are at an end; r , r' , and w , are no longer constant quantities and simple functions of the time, but each of them branches out into a series of powers of the eccentricities, and sines and cosines of variable arcs.

* Laplace (*Theorie des Satellites de Jupiter*, p. 102.) makes the deviation of the attraction of the first, compared with the fourth satellite from the law of the inverse squares of the distances, only $\frac{1}{80000}$ of the whole attraction of the former, supposing Jupiter homogeneous. In Saturn, the attraction of the ring must cause a much more considerable deviation from this law.

Astronomy. Our object being to reduce Ω , &c. to sines and cosines of arcs proportional to the time, or of the form $A t + B$, it is evident that we must substitute for r, r' , and w , their values so expressed, and then develop each term of Ω to the extent we wish. At present we will confine ourselves to the first powers of the eccentricities. **Physical Astronomy.**

Now we have
$$\Omega = R + R' \cdot \cos w + R'' \cdot \cos 2w + \&c.$$

in which R, R', R'' , &c. and w are explicit functions of r, r' , &c. and $w = \theta - \theta'$ putting θ and θ' for the true longitudes of the two planets. Now, if we call ϵ , and ϵ' , the longitudes at the commencement of the time t , $n t + \epsilon$ and $n' t + \epsilon'$ will be their mean longitudes after the lapse of that time, and calling π and π' the longitudes of the perihelion, the mean anomalies will be $n t + \epsilon - \pi$ and $n' t + \epsilon' - \pi'$. Hence, the true anomalies will be (by equation 30),

$$\begin{aligned} n t + \epsilon - \pi + 2 e \cdot \sin (n t + \epsilon - \pi) + e^2 \times \&c. \\ n' t + \epsilon' - \pi' + 2 e' \cdot \sin (n' t + \epsilon' - \pi') + e'^2 \times \&c. \end{aligned}$$

and the true longitudes of course are

$$\begin{aligned} (n t + \epsilon) + 2 e \cdot \sin (n t + \epsilon - \pi) + e^2 \times \&c. \\ (n' t + \epsilon') + 2 e' \cdot \sin (n' t + \epsilon' - \pi') + e'^2 \times \&c. \end{aligned}$$

Hence, if we neglect the eccentricities, we have simply $w = (n t - n' t + \epsilon - \epsilon')$ and as R, R' , &c. in this case assume their circular values A, A' , &c., the terms of Ω not depending on the eccentricities will remain as before,

$$A + A' \cdot \cos (n t - n' t + \epsilon - \epsilon') + A'' \cdot \cos 2 (n t - n' t + \epsilon - \epsilon') + \&c.$$

On the other hand, the terms depending on the eccentricities have their origin,

1st. In the development of the functions R, R' , &c.; when, instead of r, r' , we put their elliptic values,

$$r = a + \Delta r, \text{ and } r' = a' + \Delta r'$$

denoting by Δr and $\Delta r'$ the parts of r, r' arising from the eccentricities.

2d. In the substitution of $W + \Delta w$ for w in $\cos w, \cos 2w$, &c. W being the part of w independent of the eccentricities, or

$$W = (n - n') \cdot t + (\epsilon - \epsilon')$$

and Δw being the part depending on them, or

$$\Delta w = 2 e \cdot \sin (n t + \epsilon - \pi) - 2 e' \cdot \sin (n' t + \epsilon' - \pi') + e^2 \times \&c. \&c.$$

3rd. In the multiplication of these terms together.

Now, if we still continue to denote by Δ that variation in Ω, r, r' , &c. which arises from the eccentricities, we have

$$\Delta \Omega = \frac{d \Omega}{d r} \Delta r + \frac{d \Omega}{d r'} \Delta r' + \frac{d \Omega}{d w} \Delta w$$

in which the differential co-efficients $\frac{d \Omega}{d r}, \frac{d \Omega}{d r'}, \frac{d \Omega}{d w}$, are to have their circular values. If we would

pursue the investigation further, we must add to $\Delta \Omega$ the terms $\frac{d^2 \Omega}{d r^2} \cdot \frac{(\Delta r)^2}{1 \cdot 2}$, &c.

Now, in general, we have

$$\begin{aligned} \frac{d \Omega}{d r} &= \frac{d R}{d r} + \frac{d R'}{d r} \cdot \cos w + \frac{d R''}{d r} \cdot \cos 2w + \&c. \\ \frac{d \Omega}{d r'} &= \frac{d R}{d r'} + \frac{d R'}{d r'} \cdot \cos w + \frac{d R''}{d r'} \cdot \cos 2w + \&c. \end{aligned}$$

the circular values of which are respectively

$$\frac{d A}{d a} + \frac{d A'}{d a} \cdot \cos W + \frac{d A''}{d a} \cdot \cos 2W + \&c. \quad (144, 1)$$

and

$$\frac{d A}{d a'} + \frac{d A'}{d a'} \cdot \cos W + \frac{d A''}{d a'} \cdot \cos 2W + \&c. \quad (144, 2)$$

and, in like manner, the circular value of $\frac{d \Omega}{d w}$ is

$$- \{ 1 \cdot A' \cdot \sin W + 2 \cdot A'' \cdot \sin 2W + 3 \cdot A''' \cdot \sin 3W + \&c. \} \quad (144, 3)$$

The values of Δr and $\Delta r'$ are given by equation (28) if we substitute merely for $n t$ the expressions $n t + \epsilon - \pi$, and $n' t + \epsilon' - \pi'$; which in this case are the mean anomalies, because the mean longitudes $n t, n' t$, and the constants $\epsilon, \epsilon', \pi, \pi'$, are reckoned, not from the perihelion of the orbits, as in that equation, but from the line of the equinoxes. If then we put $V = n t + \epsilon - \pi$, $V' = n' t + \epsilon' - \pi'$, we have

$$\left. \begin{aligned} \Delta r &= -a e \cdot \cos V + e^2 \times \&c. \\ \Delta r' &= -a' e' \cdot \cos V' + e'^2 \times \&c. \end{aligned} \right\}; \dots\dots\dots (145)$$

and

$$\Delta w = 2 e \cdot \sin V - 2 e' \cdot \sin V' + e^2 \times \&c.$$

Substituting therefore in $\Delta \Omega$ for the differential co-efficients, their circular values in (144, 1, 2, 3,) and for $\Delta r, \Delta r'$, and Δw , the values just now found, it will become

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$$\begin{aligned} \Delta \Omega = & -ae \cdot \cos V \left\{ \frac{dA}{da} + \frac{dA'}{da} \cdot \cos W + \frac{dA''}{da} \cdot \cos 2W + \&c. \right\} \\ & - a'e' \cdot \cos V' \left\{ \frac{dA}{da'} + \frac{dA'}{da'} \cdot \cos W + \frac{dA''}{da'} \cdot \cos 2W + \&c. \right\} \\ & - (2e \cdot \sin V - 2e' \cdot \sin V') \{ A' \cdot \sin W + 2A'' \sin 2W + \&c. \} \\ & + e^2 \times \&c. \end{aligned} \quad (146)$$

Now these series are not precisely in the form we want them; the cosines and sines of V and W being multiplied together; and to disengage them, we must employ the well known trigonometrical formula

$$\begin{aligned} \cos A \cdot \cos B &= \frac{1}{2} \{ \cos (A+B) + \cos (A-B) \} \\ - \sin A \cdot \sin B &= \frac{1}{2} \{ \cos (A+B) - \cos (A-B) \} \end{aligned} \quad (A)$$

By the aid of this, we get

$$\begin{aligned} \cos V &= \cos V & \cos V' &= \cos V' \\ \cos V \cdot \cos W &= \frac{1}{2} \{ \cos (V+W) + \cos (V-W) \}; & \cos V' \cdot \cos W' &= \frac{1}{2} \{ \cos (V'+W') + \cos (V'-W') \} \\ \cos V \cdot \cos 2W &= \frac{1}{2} \{ \cos (V+2W) + \cos (V-2W) \}; & \cos V' \cdot \cos 2W' &= \frac{1}{2} \{ \cos (V'+2W') + \cos (V'-2W') \} \\ \&c. &= \&c. & \&c. &= \&c. \\ \sin V \cdot \sin W &= \frac{1}{2} \{ \cos (V-W) - \cos (V+W) \}; & \sin V' \cdot \sin W' &= \frac{1}{2} \{ \cos (V'-W') - \cos (V'+W') \} \\ \sin V \cdot \sin 2W &= \frac{1}{2} \{ \cos (V-2W) - \cos (V+2W) \}; & \sin V' \cdot \sin 2W' &= \frac{1}{2} \{ \cos (V'-2W') - \cos (V'+2W') \} \\ \&c. &= \&c. & \&c. &= \&c. \end{aligned}$$

Thus we see that $\Delta \Omega$ (if we confine ourselves to the first powers of the eccentricities) is reducible to a series of sines and cosines, whose arguments are all comprehended in the forms $V \pm iW$, and $V' \pm iW'$; or, (since $\cos -A = -\cos A$) in the forms $iW \pm V$ and $iW' \pm V'$. That is, (since $V = nt + \epsilon - \pi$, and $W = nt - n't + \epsilon - \epsilon'$) in the forms

$$\begin{aligned} i(nt - n't + \epsilon - \epsilon') \pm (nt + \epsilon - \pi) \\ i(nt - n't + \epsilon - \epsilon') \pm (n't + \epsilon' - \pi') \end{aligned}$$

If we actually execute the substitutions (still, for brevity, preserving the denominations W and V) we shall obtain for the value of $\Delta \Omega$ the following expression—

$$\begin{aligned} \Delta \Omega = & -ae \cdot \frac{dA}{da} \cdot \cos V & - a'e' \cdot \frac{dA}{da'} \cdot \cos V'; & (147) \\ & - e \left\{ \frac{a}{2} \frac{dA'}{da} - A \right\} \cdot \cos (W+V) & - e' \left\{ \frac{a'}{2} \frac{dA'}{da'} + A' \right\} \cos (W+V') \\ & - e \left\{ \frac{a}{2} \frac{dA'}{da} + A \right\} \cdot \cos (W-V) & - e' \left\{ \frac{a'}{2} \frac{dA'}{da'} - A' \right\} \cos (W-V') \\ & - e \left\{ \frac{a}{2} \frac{dA''}{da} - 2A'' \right\} \cdot \cos (2W+V) & - e' \left\{ \frac{a'}{2} \frac{dA''}{da'} + 2A'' \right\} \cos (2W+V') \\ & - e \left\{ \frac{a}{2} \frac{dA''}{da} + 2A'' \right\} \cdot \cos (2W-V) & - e' \left\{ \frac{a'}{2} \frac{dA''}{da'} - 2A'' \right\} \cos (2W-V') \\ & - \&c. & - \&c. \end{aligned}$$

But our object is to obtain the developement of Q the perturbative function, or $2 \int d\Omega + r \frac{d\Omega}{dr}$. The part of this independent of the eccentricities is already found, and we have now only to consider that depending on them; which, in the notation above adopted, will be expressed by ΔQ , or

$$2 \Delta \int d\Omega + \Delta \left(r \frac{d\Omega}{dr} \right) = 2 \int d\Delta \Omega + \Delta \left(r \frac{d\Omega}{dr} \right); \quad (148)$$

Let us first consider the value of the first part of this expression $2 \int d\Delta \Omega$, and let any term of $\Delta \Omega$ be represented by

$$M \cdot \cos (iW + kV + lV')$$

in which i may be any integer from 0 to infinity, and k and l either ± 1 or 0, an expression which obviously comprehends all the terms of which $\Delta \Omega$ consists. This premised, it is obvious that the co-efficient M being constant, the variation of $\Delta \Omega$ can only arise from the variation of the angle $iW + kV + lV'$; and as this angle is supposed only to vary by the motion of m , we are to suppose nt only to vary, and $n't$ to remain constant. So that we have, for that part of the variation of $\Delta \Omega$ which arises from the term in question,

$$-M (i dW + k dV) \sin (iW + kV + lV')$$

because $dV' = 0$, since $V' = n't + \epsilon' - \pi'$. Now $dW = n dt$ and $dV = n dt$ also. Consequently, this

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$$- (i + k) M . n d t . \sin (i W + k V + l V')$$

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and the part of the expression $\int d \Omega$ arising from this term, will therefore be

$$\frac{2 (i + k) n}{i (n - n') + k n + l n'} . M \cos (i W + k V + l V') \quad (149)$$

We see therefore by the foregoing reasoning, that in order to obtain that part of the perturbative function which originates in the term $\int d \Omega$, and depends on the eccentricities, we have only to take the terms of the expression for $\Delta \Omega$ (147) in their order, and with their proper signs, and multiply each of them respectively by that value of the fraction

$$\frac{2 (i + k) . n}{i (n - n') + k n + l n'}$$

which corresponds to the values of i, k, l , in its argument, represented by $i W + k V + l V'$. For instance, the term which has simply V for its argument, must be multiplied by $\frac{2 n}{1 . n} = 2$, that which has V' by o .— Again, the terms, whose respective arguments are $W + V, W - V, W + V', W - V'$, are to be multiplied, according to this rule, by the respective co-efficients $\frac{4 n}{2 n - n'}, 0, 2$, and $\frac{2 n}{n - 2 n'}$. Similarly, the terms which have $2 W + V, 2 W - V, 2 W + V', 2 W - V'$, for their arguments, acquire the co-efficients

$$\frac{6 n}{3 n - 2 n'}, \frac{2 n}{n - 2 n'}, \frac{4 n}{2 n - n'}, \text{ and } \frac{4 n}{2 n - 3 n'}; \text{ and so on.}$$

The co-efficients thus acquired by integration depend solely on the ratios of the mean motions, or periodic times, of the disturbed and disturbing planet, and are of very great importance in the planetary theory. Were it not for these, the theory of the planetary perturbations would be very simple, as the same treatment, or nearly so, could be applied in every case, and the magnitudes of the several inequalities would go on diminishing with more or less rapidity, as the arguments contained higher multiples of the mean motions. But these co-efficients prevent this regular progression of magnitude from taking place, and render it difficult to foresee without computation whether any assigned inequality may be neglected or not, and impossible to argue from its known minuteness in the case of one pair of planets to its probable smallness in that of another. Thus an inequality, whose maximum we know to be small in the case of Venus disturbed by the earth, may have a considerable magnitude in that of Jupiter disturbed by Saturn, merely from a relation subsisting between the periodic times in the latter case which does not in the former. In fact, it is evident that if the periods should happen to be so related as to render the denominator of any of the foregoing or similar fractions very small compared to the numerator, the inequality into which it enters as a co-efficient will, on this account alone, acquire a very great preponderance. Thus, if the period of the disturbed planet

were very nearly half or double that of the disturbing; the terms, multiplied by $\frac{2 n}{2 n - n'}$ or by $\frac{n}{n - 2 n'}$ would become very large, and the length of the period of the inequality represented by it would be proportionally increased, and in the case of exact commensurability would actually become infinite; that is to say, the disturbance would go on to such an extent, as to make a total subversion of the original conditions of the problem. The physical reason of this is equally obvious. In the case just instanced, the two planets, at every revolution of the exterior, would be placed in exactly the same circumstances—the same disturbing forces would act in the same manner for a series of ages, and their effects, instead of compensating each other by mutual opposition, would go on accumulating in the same direction, till their orbits at length became entirely changed, and the commensurability of their periods at length ceased to subsist. In fact, an equation thus limited by no period, and affecting both the longitude and radius vector of each orbit always the same way, is equivalent to an alteration of the *mean* motion and *mean* distance; and as this would take place in opposite directions on the two planets, shortening the period of one, and lengthening that of the other, their periods would continually recede from commensurability; the magnitude of the inequality in question, as well as the length of its period, would both acquire finite values, which even then would continually diminish, till reduced within limits consistent with the stability of the system. It is probably by some such gradations (if we may hazard a conjecture on a part of the planetary theory so far beyond the reach of analysis or exact reasoning,) that our system has attained, in the course of almost indefinite ages, its present admirable state of equilibrium, in which no inequality of enormous magnitude exists; and those which have any notable value, can be proved to be confined within comparatively narrow bounds.

Let us next consider the part of the developement of Q , which arises from the term $\Delta \left(r \frac{d \Omega}{d r} \right)$. Now, if we regard only the first powers of the eccentricities, and consequently neglect the squares of Δr , &c. we have

$$\begin{aligned} \Delta \left(r \frac{d \Omega}{d r} \right) &= \frac{d \Omega}{d r} . \Delta r + r \Delta \frac{d \Omega}{d r} \\ &= \frac{d \Omega}{d r} \Delta r + r . \frac{d^2 \Omega}{d r^2} \Delta r + r \frac{d^2 \Omega}{d r d r'} \Delta r' + r \frac{d^2 \Omega}{d r d w} \Delta w \end{aligned}$$

Astronomy. in which the differential co-efficients are to have their circular values, which we may represent by $\frac{d\Omega}{da}$, &c. Physical Astronomy.

Now these are,

$$\left. \begin{aligned} \frac{d\Omega}{da} &= \frac{dA}{da} + \frac{dA'}{da} \cos W + \frac{dA''}{da} \cos 2W + \&c. \\ a \frac{d^2\Omega}{da^2} &= a \frac{d^2A}{da^2} + a \frac{d^2A'}{da^2} \cos W + a \frac{d^2A''}{da^2} \cos 2W + \&c. \\ a \frac{d^2\Omega}{da da'} &= a \frac{d^2A}{da da'} + a \frac{d^2A'}{da da'} \cos W + a \frac{d^2A''}{da da'} \cos 2W + \&c. \\ a \frac{d^2\Omega}{da dW} &= 1 \cdot a \frac{dA'}{da} \sin W + 2 \cdot a \frac{dA''}{da} \sin 2W + \&c. \end{aligned} \right\} \dots\dots\dots (150)$$

Let these be substituted in the above expression for $\Delta \left(r \frac{d\Omega}{dr} \right)$ and the values of $\Delta r, \Delta r'$ and Δw , given in (145) be put for them, and we shall have, by a process exactly similar to that gone through for $\Delta \Omega$,

$$\begin{aligned} \Delta \left(r \frac{d\Omega}{dr} \right) &= -ae \cos V \cdot \left\{ \left(\frac{dA}{da} + a \frac{d^2A}{da^2} \right) + \left(\frac{dA'}{da} + a \frac{d^2A'}{da^2} \right) \cos W + \&c. \right\} \\ &\quad - a'e' \cos V' \cdot \left\{ a \frac{d^2A}{da da'} + a \frac{d^2A'}{da da'} \cos W + \&c. \right\} \\ &\quad - (2e \sin V - 2e' \sin V') \left\{ 1 a \frac{dA'}{da} \sin W + 2 \cdot \&c. \right\} \end{aligned} \quad (151)$$

which, by the use of the same trigonometrical formulæ, and by properly arranging the terms, becomes

$$\begin{aligned} \Delta \left(r \frac{d\Omega}{dr} \right) &= -e \left\{ a^2 \frac{d^2A}{da^2} + a \frac{dA}{da} \right\} \cos V - e' \cdot a a' \frac{d^2A}{da da'} \cos V' \\ &\quad - \frac{e}{2} \left\{ a^2 \frac{d^2A'}{da^2} - a \frac{dA'}{da} \right\} \cos (W + V) - \frac{e'}{2} \left\{ a a' \frac{d^2A}{da da'} + 2 a \frac{dA'}{da} \right\} \cos (W + V') \\ &\quad - \frac{e}{2} \left\{ a^2 \frac{d^2A'}{da^2} + 3 a \frac{dA'}{da} \right\} \cos (W - V) - \frac{e'}{2} \left\{ a a' \frac{d^2A'}{da da'} - 2 a \frac{dA'}{da} \right\} \cos (W + V') \\ &\quad - \frac{e}{2} \left\{ a^3 \frac{d^2A''}{da^3} - 3 a \frac{dA''}{da} \right\} \cos (2W + V) - \frac{e'}{2} \left\{ a a' \frac{d^2A''}{da da'} + 4 a \frac{dA''}{da} \right\} \cos (2W + V') \\ &\quad - \frac{e}{2} \left\{ a^3 \frac{d^2A''}{da^3} + 5 a \frac{dA''}{da} \right\} \cos (2W - V) - \frac{e'}{2} \left\{ a a' \frac{d^2A''}{da da'} - 4 a \frac{dA''}{da} \right\} \cos (2W - V') \\ &\quad - \&c. \quad \quad \quad - \&c. ; \dots\dots\dots (152) \end{aligned}$$

We are now in a condition to express the whole value of ΔQ , by combining this with the expressions (147) and (149) and it is evident that our result will be of the following form:—

$$\Delta Q = (Ne \cos V + N'e' \cos V') + Oe \cos (W + V) + O'e' \cos (W + V') + Pe \cos (W - V) + P'e' \cos (W - V') + Qe \cos (2W + V) + \&c. \quad (153)$$

and the co-efficients of the several arguments will be

$$\left. \begin{aligned} N &= - \left\{ a^2 \frac{d^2A}{da^2} + 3 a \frac{dA}{da} \right\} \\ O &= - \frac{1}{2} \left\{ a^2 \frac{d^2A'}{da^2} + \frac{2n + n'}{2n - n'} a \frac{dA'}{da} - \frac{8n}{2n - n'} A' \right\} \\ P &= - \frac{1}{2} \left\{ a^2 \frac{d^2A'}{da^2} + 3 a \frac{dA'}{da} \right\}; \\ N' &= - a a' \cdot \frac{d^2A}{da da'} \\ O' &= - \frac{1}{2} \left\{ a a' \cdot \frac{d^2A'}{da da'} + 2 a \frac{dA'}{da} + 2 a' \frac{dA'}{da'} + 4 A' \right\} \\ P' &= - \frac{1}{2} \left\{ a a' \cdot \frac{d^2A'}{da da'} - 2 a \frac{dA'}{da} + \frac{2n}{n - 2n'} a' \frac{dA'}{da'} - \frac{4n}{n - 2n'} A' \right\} \end{aligned} \right\} \dots\dots\dots (154)$$

The value of ΔQ being thus obtained, that of Π is had by mere substitution. The part of Π , independent of the eccentricities, will remain as in the preceding section, changing only in the several arguments w into W ; and if we denote this by Π , and by $\Delta \Pi$ the part of Π which depends on the eccentricities, the equation (128) will give

$$\Delta \Pi = \frac{e}{a^2} \left\{ Q \cdot \cos V - 2 \int Q n dt \cdot \sin V \right\} - \frac{\Delta Q}{a^2}$$

in which Q and ΔQ denote, as we have all along supposed, the parts of Q respectively independent, and dependent, on the eccentricities.

The value of $\Delta \Pi$ therefore consists of two parts; the latter $-\frac{\Delta Q}{a^2}$ is immediately obtained from the expression of ΔQ above found, and is equal to

$$-\frac{e}{a^2} \left\{ N \cdot \cos V + O \cdot \cos (W + V) + P \cdot \cos (W - V) + \&c. \right\} \\ - \frac{e'}{a^2} \left\{ N' \cdot \cos V' + O' \cdot \cos (W + V') + P' \cdot \cos (W - V') + \&c. \right\}$$

The former depends on Q , and must be determined by substituting for Q its value

$$Q = M + M' \cdot \cos W + M'' \cdot \cos 2W + \&c.$$

where M , M' , &c. are co-efficients, whose values are assigned in equation (138, 1). This substitution made, we find

$$Q \cos V = M \cdot \cos V + \frac{M'}{2} \cdot \cos (W + V) + \frac{M'}{2} \cdot \cos (W - V) + \&c. \\ \int Q n dt \cdot \sin V = -M \cdot \cos V - \frac{M' n}{2(2n - n')} \cos (W + V) - \frac{M' n}{2n'} \cos (W - V) - \&c.$$

So that the part of $\Delta \Pi$ now in question becomes

$$\frac{e}{a^2} \left\{ 3M \cdot \cos V + \frac{4n - n'}{4n - 2n'} M' \cdot \cos (W + V) + \frac{2n + n'}{2n'} M' \cdot \cos (W - V) + \&c. \right\}$$

and the whole value of $\Delta \Pi$ will be as follows:—

$$\Delta \Pi = \frac{e}{a^2} \left\{ (3M - N) \cos V + \left(\frac{4n - n'}{4n - 2n'} M' - O \right) \cos (W + V) + \left(\frac{2n + n'}{2n'} M' - P \right) \cos (W - V) + \&c. \right\} \\ - \frac{e'}{a^2} \left\{ N' \cdot \cos V' + O' \cdot \cos (W + V') + P' \cdot \cos (W - V') + \&c. \right\}; \quad (155)$$

This found, the integral of the equation $\frac{d^2 \delta u}{dt^2} + n^2 u + \Pi = 0$ will be obtained by the expressions 106, 107, and 108. The parts of δu , δr , and $\delta \theta$, independent of the eccentricities, have already been found; and calling therefore $\Delta \delta u$, $\Delta \delta r$, and $\Delta \delta \theta$, those parts of these respective quantities which depend on the eccentricities, we shall have (since $V = nt + \epsilon - \pi$, $W = (n - n')t + (\epsilon - \epsilon')$ and $V' = n't + \epsilon' - \pi'$)

$$\Delta \delta u = \frac{e}{a^2} \left\{ (3M - N) \cdot \frac{t}{2} \cos V + \left(\frac{4n - n'}{4n - 2n'} M' - O \right) \frac{\cos (W + V)}{(2n - n')^2 - n^2} + \right. \\ \left. + \left(\frac{2n + n'}{2n'} M' - P \right) \frac{\cos (W - V)}{n'^2 - n^2} + \&c. \right\} \\ - \frac{e'}{a^2} \left\{ \frac{N'}{n'^2 - n^2} \cos V' + \frac{O'}{2} \cdot \frac{t}{2} \cos (W + V') + \frac{P'}{(n - 2n')^2 - n^2} \cos (W - V') + \&c. \right\} \quad (156)$$

The perturbation of the radius vector is now easily found; for, as we have by equation 124,

$$\delta r = -a \delta u \{ 1 + 2e \cdot \cos V + \&c. \}$$

this gives

$$\Delta \delta r = -a \{ \Delta \delta u + 2e \delta u \cdot \cos V \}$$

whence $\Delta \delta r$ is readily obtained. With regard to the perturbation in longitude, or $m' \delta \theta$, no further difficulty than the length of the substitutions remains to be encountered; for the part not depending on the eccentricities being already obtained, that which depends on them will be had by merely substituting for δr ,

$r \frac{d\Omega}{dr}$ and Ω , the values of $\Delta \delta r$, $\Delta \left(r \frac{d\Omega}{dr} \right)$ and $\Delta \Omega$ already obtained in the general expression (121).

But as this process of substitution and reduction presents no difficulties in principle, requiring only patience and exactness in performing the numerous combinations of the terms which occur in it, we shall not pursue it, but content ourselves with observing, that the part of it which depends on the first powers of the eccentricities will, on examination, be found to assume the form

$$e \cdot a t \cdot \cos V + e' \cdot b t \cdot \cos (W + V) \\ + e \left\{ A \cdot \sin (W + V) + B \cdot \sin (W - V) + C \cdot \sin (2W + V) + \&c. \right\} \\ + e' \left\{ A' \cdot \sin V' + B' \cdot \sin (W - V') + C' \cdot \sin (2W + V') + \&c. \right\}; \dots\dots\dots (157)$$

It is here that we first encounter the secular equations of the planetary motions, in the form of two terms containing the time t disengaged from the signs \sin and \cos , and therefore capable of indefinite increase and diminution. They are multiplied by the eccentricities, and therefore originate from the ellipticity of the

Astronomy. planetary orbits; and in the case of strictly circular orbits, would not exist. Similar terms occur of course Physical Astronomy. in the value of δr , being introduced by the integration of the equation for δu . But as the discussion of these terms is one of the most delicate and difficult points of the planetary theory, we shall not enter upon it till we have pointed out the method of taking into account the higher powers of the eccentricities.

SECTION V.

Of the inequalities depending on the squares and higher powers of the eccentricities.

It is not our intention to enter into any detailed account of this very complicated part of the planetary theory. Any such attempt would lead us far beyond our proper limits; and the reader, who is desirous to follow it into its minutiae, must consult the original memoirs of Laplace, Lagrange, &c., the *Mécanique Céleste*, and other works of a similar nature. In the foregoing sections we have however followed, as nearly as possible, the course pursued in the last named immortal work, supplying only such steps in the analysis as cannot be expected to be discovered by the ordinary student, (and they are numerous) and endeavouring throughout to place the principles of the several processes in as strong a light as possible. In the present section, the explanation of the principles on which the process of approximation is to be pursued will be almost our sole object.

Let us resume the consideration of the function Ω .

$$\Omega = R + R' \cdot \cos w + R'' \cdot \cos 2w + R''' \cdot \cos 3w + \&c.$$

When $r + \Delta r$, $r' + \Delta r'$, and $w + \Delta w$, are substituted for r, r', w , in this, Ω becomes $\Omega + \Delta \Omega$, and we have

$$\Delta \Omega = \left. \begin{aligned} & \frac{d\Omega}{da} \cdot \frac{\Delta r}{1} + \frac{d\Omega}{da'} \cdot \frac{\Delta r'}{1} + \frac{d\Omega}{dW} \cdot \frac{\Delta w}{1} \\ & + \frac{d^2\Omega}{da^2} \cdot \frac{(\Delta r)^2}{1 \cdot 2} + \frac{d^2\Omega}{da da'} \cdot \frac{\Delta r \cdot \Delta r'}{1 \cdot 1} + \frac{d^2\Omega}{da'^2} \cdot \frac{(\Delta r')^2}{1 \cdot 2} + \&c. \\ & + \frac{d^3\Omega}{da^3} \cdot \frac{(\Delta r)^3}{1 \cdot 2 \cdot 3} + \&c. \end{aligned} \right\} \dots\dots\dots (158)$$

The differential co-efficients of Ω are here supposed to have their circular values denoted by $\frac{d\Omega}{da}$, $\frac{d\Omega}{da'}$, $\frac{d\Omega}{dW}$, &c.

In like manner, if we consider the value of $\Delta \left(r \frac{d\Omega}{dr} \right)$, or the augmentation of $r \frac{d\Omega}{dr}$ produced by the eccentricities, we have only to substitute for Ω in the expression (158) the circular value of the function in question, or $a \frac{d\Omega}{da}$, and we get

$$\Delta \left(r \frac{d\Omega}{dr} \right) = \frac{\Delta r}{1} \times \frac{d}{da} \left(a \frac{d\Omega}{da} \right) + \frac{\Delta r'}{1} \cdot \frac{d}{da'} \left(a \frac{d\Omega}{da} \right) + \frac{\Delta w}{1} \cdot \frac{d}{dW} \left(a \frac{d\Omega}{da} \right) \\ + \frac{(\Delta r)^2}{1 \cdot 2} \cdot \frac{d^2}{da^2} \left(a \frac{d\Omega}{da} \right) + \frac{\Delta r \Delta r'}{1 \cdot 1} \cdot \frac{d^2}{da da'} \left(a \frac{d\Omega}{da} \right) + \&c. \\ + \&c. \dots\dots\dots (159)$$

in which $\frac{d}{da}$, $\frac{d}{da'}$, &c. denote the differentiation relative to a, a' , &c. respectively of the function to which they are prefixed, and the division of the resulting differential by $da, da', \&c.$ according to the very convenient system of notation explained in (Lacroix, *Differential and Integral Calculus*, 8vo. English translation, Appendix.)

Since $\Omega = A + A' \cdot \cos W + A'' \cdot \cos 2W + \&c.$ we must have

$$\left. \begin{aligned} a \frac{d\Omega}{da} &= a \frac{dA}{da} + a \frac{dA'}{da} \cdot \cos W + a \frac{dA''}{da} \cdot \cos 2W + \&c. \\ \frac{d}{da} \left(a \frac{d\Omega}{da} \right) &= \frac{d}{da} \left(a \frac{dA}{da} \right) + \frac{d}{da} \left(a \frac{dA'}{da} \right) \cos W + \frac{d}{da} \left(a \frac{dA''}{da} \right) \cdot \cos 2W + \&c. \\ \frac{d}{da'} \left(a \frac{d\Omega}{da} \right) &= \frac{d}{da'} \left(a \frac{dA}{da} \right) + \frac{d}{da'} \left(a \frac{dA'}{da} \right) \cos W + \frac{d}{da'} \left(a \frac{dA''}{da} \right) \cdot \cos 2W + \&c. \\ \frac{d}{dW} \left(a \frac{d\Omega}{da} \right) &= -a \frac{dA'}{da} \cdot \sin W - 2a \frac{dA''}{da} \sin 2W - \&c. \\ \frac{d^2}{da^2} \left(a \frac{d\Omega}{da} \right) &= \frac{d^2}{da^2} \left(a \frac{dA}{da} \right) + \frac{d^2}{da^2} \left(a \frac{dA'}{da} \right) \cos W + \frac{d^2}{da^2} \left(a \frac{dA''}{da} \right) \cos 2W + \&c. \\ \frac{d^2}{da dW} \left(a \frac{d\Omega}{da} \right) &= -\frac{d}{da} \left(a \frac{dA'}{da} \right) \sin W - 2 \frac{d}{da} \left(a \frac{dA''}{da} \right) \sin 2W - \&c. \end{aligned} \right\}; (160)$$

Astronomy. and so on. Now these series may be regarded as completely developed; for the co-efficients of their several terms $\cos W$, $\sin W$, &c. are quantities completely given in numbers, when a , a' , A , A' , A'' , &c. and their differential co-efficients are known, by the equations Physical Astronomy.

$$\left. \begin{aligned} \frac{d}{da} \left(a \frac{dA}{da} \right) &= \frac{dA}{da} + a \frac{d^2 A}{da^2}; \quad \frac{d}{da} \left(a \frac{dA'}{da} \right) = \frac{dA'}{da} + a \frac{d^2 A'}{da^2}; \quad \&c. \\ \frac{d}{da'} \left(a \frac{dA}{da} \right) &= \frac{d^2 A}{da da'}; \quad \frac{d}{da'} \left(a \frac{dA'}{da} \right) = \frac{d^2 A'}{da da'}; \quad \&c. \\ \frac{d^2}{da^2} \left(a \frac{dA}{da} \right) &= 2 \frac{d^2 A}{da^2} + a \frac{d^3 A}{da^3}; \\ &\&c. \end{aligned} \right\}; \dots\dots\dots (161)$$

If, instead of $r \frac{d\Omega}{dr}$, we had any other function to develop (such as for instance $r^2 \cdot \frac{d^3 \Omega}{d^2 r dr}$), we might treat it exactly in the same way, and should arrive at corresponding series, in which the sines and cosines of W and its multiples would be combined with co-efficients absolutely constant, and reducible to numbers.

In the values of $\Delta \Omega$ and $\Delta \left(r \frac{d\Omega}{dr} \right)$ in (156) and (159) we see therefore that the differential co-efficients of Ω introduce the sines and cosines of W , and its multiples, combined only with given quantities, and not involving the eccentricities. These latter arise from the factors Δr , $\Delta r'$, Δw , and their powers. Let us now examine these more nearly. Supposing then, as we have done all along,

$$V = nt + e - \pi; \quad V' = n't + e' - \pi'; \quad W = nt - n't + e - e';$$

let us take a , β , &c. as follows:—

$$\left. \begin{aligned} a &= -a \cdot \cos V & a' &= -a' \cdot \cos V' \\ \rho &= \frac{a}{2} (1 - \cos 2V) & \rho' &= \frac{a'}{2} (1 - \cos 2V') \\ \gamma &= -\frac{a}{3} (3 \cos 3V - 3 \cos V) & \gamma' &= -\frac{a'}{3} (3 \cos 3V' - 3 \cos V') \\ &\&c. & &\&c. \\ p &= 2 \sin V & p' &= 2 \sin V' \\ q &= \frac{5}{4} \sin 2V & q' &= \frac{5}{4} \sin 2V' \\ r &= -\frac{1}{4} \sin V + \frac{13}{12} \sin 3V & r' &= -\frac{1}{4} \sin V' + \frac{13}{12} \sin 3V' \\ &\&c. & &\&c. \end{aligned} \right\}; \dots\dots\dots (162)$$

Then we shall have, by the equations (28) and (30)

$$\left. \begin{aligned} \Delta r &= a e + \beta e^2 + \gamma e^3 + \&c. \\ \Delta r' &= a' e' + \beta' e'^2 + \gamma' e'^3 + \&c. \\ \Delta w &= (p e + q e^2 + r e^3 + \&c.) - (p' e' + q' e'^2 + r' e'^3 + \&c.). \end{aligned} \right\}; \dots\dots\dots (163)$$

whence we obtain

$$\begin{aligned} (\Delta r)^2 &= a^2 e^2 + 2 a \beta e^3 + (\beta^2 + 2 a \gamma) e^4 + \&c. \\ \Delta r \Delta r' &= a a' e e' + a \beta' e e'^2 + \beta a' e^2 e' + \&c. \\ \Delta r \Delta w &= p a e^2 + (p \beta + q a) e^3 + \&c. - p' a e e' - \&c. \\ (\Delta w)^2 &= p^2 e^2 + 2 p q e^3 + \&c. + (p'^2 e'^2 + 2 p' q' e'^3 + \&c.) - 2 p p' e e' - \&c. \end{aligned}$$

and so on; and it only remains to substitute these values in the expressions for $\Delta \Omega$ and $\Delta \left(r \frac{d\Omega}{dr} \right)$.

Confining ourselves to $\Delta \Omega$, since it is obvious that the process is exactly similar for the other function, we have, as before, for the part depending on the first powers of the eccentricities,

$$e \left\{ a \frac{d\Omega}{da} + p \frac{d\Omega}{dW} \right\} + e' \left\{ a' \frac{d\Omega}{da'} - p' \frac{d\Omega}{dW} \right\}$$

which, developed into series of sines and cosines, gives the result obtained in the last section.

The part of $\Delta \Omega$ depending on the squares of the eccentricities, consists of three terms multiplied respectively by e^2 , $e e'$, and e'^2 , which originate, 1st. From the terms βe^2 , $\beta' e'^2$, $q e^2$, and $-q' e'^2$, in the simple powers of Δr , $\Delta r'$, Δw , and which are therefore affected with the differential co-efficients of the first order only: 2dly. With the terms $a^2 e^2$, $a'^2 e'^2$, $p^2 e^2$, $p'^2 e'^2$, and $-2 p p' e e'$, in $(\Delta r)^2$, $(\Delta r')^2$, and $(\Delta w)^2$, and which are consequently affected with differential co-efficients of Ω of the second order.

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3dly. With the terms $a p e^2$, $a' p' e'^2$, $a a' e e'$, $a p' e e'$, $a' p e e'$, which arise from combinations of Δr with Δw and with $\Delta r'$.

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The aggregate of all these terms, with their proper co-efficients, is

$$\left. \begin{aligned} & \frac{e^2}{2} \left\{ \left(2\beta \frac{d\Omega}{da} + 2q \frac{d\Omega}{dW} \right) + \left(a^2 \frac{d^2\Omega}{da^2} + 2ap \frac{d^2\Omega}{da dW} + p^2 \frac{d^2\Omega}{dW^2} \right) \right\} \\ & + e e' \left\{ a a' \frac{d^2\Omega}{da da'} - a p' \frac{d^2\Omega}{da dW} + a' p \frac{d^2\Omega}{da' dW} - p p' \frac{d^2\Omega}{dW^2} \right\} \\ & + \frac{e'^2}{2} \left\{ \left(2\beta' \frac{d\Omega}{da'} - 2q' \frac{d\Omega}{dW} \right) + \left(a'^2 \frac{d^2\Omega}{da'^2} - 2a'p' \frac{d^2\Omega}{da' dW} + p'^2 \frac{d^2\Omega}{dW^2} \right) \right\} \end{aligned} \right\} ; \dots\dots\dots (164)$$

The co-efficients of e^2 , $e^2 e'$, $e e' e'$, e'^3 , and of the higher powers and combinations, may in like manner be easily obtained; but the number of terms of which they consist, goes on increasing so rapidly, that they at length become of extreme complexity.

Let us now consider the nature of the terms into which the expressions for Ω and $r \frac{d\Omega}{dr}$ resolve themselves by the process of development, and the manner in which they become modified by the several processes of substitution and integration they have to undergo in obtaining the values of Q , Π , δu , δr , and $\delta \theta$.

It is evident, then, since a , β , γ , a' , β' , &c. are all composed of cosines, and p , q , p' , q' , &c. of sines of V , V' , and their multiples, without W , that any product or combination of these letters, (such as a^2 , $a p$, $a a'$, &c.) is reducible by the trigonometrical formulæ so often employed in the foregoing pages into the simple sines or cosines of arcs, of the form $kV + lV'$. Thus, a^2 or $a^2 \cdot \cos V^2$ becomes $a^2 \cdot \frac{1 + \cos 2V}{2}$, $a p$, or

$-2a \cdot \sin V \cdot \cos V$ becomes $-a \cdot \sin 2V$, $a a'$ or $a a' \cdot \cos V \cdot \cos V'$ is reduced to $\frac{a a'}{2} (\cos (V + V') +$

$\cos (V - V'))$, $a p'$ into $a \cdot \sin (V - V') - a \cdot \sin (V + V')$, and so on. Moreover, it is evident, that whenever the combination in question consists only of the letters a , β , a' , β' , &c. or of these combined with any product of an *even* dimension, in p , q , p' , q' , &c., that the terms into which it is resolved will consist entirely of cosines; but when a product of an *odd* dimension in p , q , p' , q' , &c. occurs, then *of* sines. Now, the differential co-efficient of Ω combined with any such product, will, in the former case, evidently be differentiated an even, and in the latter an odd number of times relatively to W ; so that in the former case it will represent a series of *cosines*, and in the latter, of *sines* of W .

Every term therefore formed by such combination, must be of one or other of the forms

$$\cos iW \cdot \cos (kV + lV') \text{ and } \sin iW \cdot \sin (kV \pm lV')$$

both which, being further resolved, produce terms comprehended in the form

$$\cos (iW \pm kV \pm lV')$$

It is thus demonstrated, that the co-efficients of all the powers of e , e' , in the developements of Ω and $r \frac{d\Omega}{dr}$ are generally reducible to series of cosines of arguments, of the form $iW \pm kV \pm lV'$; but there is a connection between the multiples of V and V' contained in any argument, and the dimension of the power or product of the eccentricities to which it belongs that we must now explain. In fact, it is obvious from the process above pursued, that if we regard a , a' , p , p' , as quantities of one dimension; β , β' , q , q' , as of two; γ , γ' , r , r' of three, and so on, the dimension of every term multiplied by e , or e' , will be one; that of the terms multiplied by e^2 , $e e'$, e'^2 , will be two, and so on. Now, the expressions of these quantities in V and V' involve, each, the sines or cosines of multiples of V , V' , as far as the number expressing its own dimension; and when these come to be combined by multiplication, and then resolved by the usual formula, it is obvious that the resulting terms of the form $\cos (kV \pm lV')$ and $\sin (kV \pm lV')$ can only contain such multiples kV and lV' of V , V' , as together (without regard to their signs) do not exceed the dimension of the combination from which they arose.

Moreover, since the alternate multiples of V , V' are absent in the expressions of a , β , &c. the same law will hold good in any combination of them when developed. Hence we may state it as a general law, that

The co-efficient of any power or product of the eccentricities of the dimension n , in the developement of Ω or $r \frac{d\Omega}{dr}$ will consist of a series of cosines, the form of whose argument is

$$iW \pm kV \pm lV'$$

in which i may have every possible value from 0 to infinity, but k and l are restricted to certain particular values, viz. those which satisfy one or other of the equations

$$k + l = n, \quad k + l = n - 2, \quad k + l = n - 4, \text{ \&c.}$$

down to $k + l = 1$, *or* $k + l = 0$, *according as n is odd or even.*

Thus, as we have already seen, the parts independent of the eccentricities consist of terms of the form \cos

Astronomy. iW only, and those depending on the first powers involve the arguments

$$iW + V, iW - V, iW + V', iW - V'$$

and no other. Similarly, in the part dependent on the squares and product of the eccentricities, the arguments which can occur, are only

$$iW, iW + 2V, iW - 2V, iW + V + V', iW + V - V', iW - V + V', iW - V - V', iW + 2V' \text{ and } iW - 2V', \text{ and so on.}$$

Here, it will be observed, we have again the argument iW , which occurred in the part independent on the eccentricities, and it is easily seen to be a general law, that *any particular argument which first occurs combined with a power or product of the eccentricities of the dimension n will occur again, combined with products of the dimensions $n + 2, n + 4, \&c.$ to infinity, but not with $n + 1, n + 3, \&c.$* For instance, the argument $iW + 3V$ cannot occur combined with any dimension of the eccentricities less than the third, and will occur again in the terms multiplied by the 5th, 7th, &c. dimensions, but not by the 4th, 6th, or any even dimensions.

Since $W = nt - n't + e - e', V = nt + e - \pi$, and $V' = n't + e' - \pi'$, the argument $iW + kV + lV'$ is equal to

$$(i + k) . nt - (i - l) n't + (i + k) e - (i - l) e' - k\pi - l\pi'$$

If then we would inquire in what terms any proposed combination of n and n' , such, for instance, as $(fn - gn')t$ can originate, we have only to put

$$i + k = f, \quad i - l = g \\ i = f - k, \quad i = i - g = (f - g) - k$$

which give

taking then in succession $k = 0, \quad k = \pm 1, \quad k = \pm 2, \&c.$ we get

$$i = f, \quad i = f \mp 1, \quad i = f \mp 2, \&c. \\ l = f - g, \quad l = f - g \mp 1, \quad l = f - g \mp 2, \&c.$$

For instance, if we would know from what terms the combination $(2n - n')t$ can originate, the corresponding values of i, k, l , are

$$1st. i = 2, \quad k = 0, \quad l = -3; \quad 2dly. \begin{cases} i = 1, & k = 1, & l = -4 \\ i = 3, & k = -1, & l = -2 \end{cases} \quad 3dly. \begin{cases} i = 0, & k = 2, & l = -5 \\ i = 4, & k = -2, & l = -1 \end{cases}$$

So that any of the arguments comprised in the following series,

$$2W - 3V', \begin{cases} W + V - 4V'; & 2V - 5V'; \\ 3W - V - 2V'; & 4W - 2V - V'; \end{cases} 5W - 3V, \&c.$$

will produce the combination in question. Now, the lowest sum of the co-efficients of V, V' in these arguments taken without regard to their signs, is 3: consequently, the combination $2nt - 5n't$ will first occur among the inequalities multiplied by the cubes or products of three dimensions of the eccentricities, and among them only in such terms as produce the arguments

$$2W - 3V', \quad 3W - V - 2V', \quad 4W - 2V - V', \quad 5W - 3V.$$

Let us now examine the co-efficients of the several arguments as they occur in the values of $Q, \Pi, \&c.$

The co-efficient of any argument, such as $iW \pm kV \pm lV'$ in the developement of Ω or $r \frac{d\Omega}{dr}$ will obviously consist only of combinations of $a, a', A, A', A'', \&c.$ and their differential co-efficients with a power or product of e, e' , and may therefore be regarded as a given quantity, and its value, with more or less trouble, numerically computed. Taking A for the general representative of such a combination, M will be a function of a, a', e, e' , and of these only, and

$$M \cos (iW \pm kV \pm lV')$$

will be the general form of any term of Ω or $r \frac{d\Omega}{dr}$.

In the value of Q , the terms of $r \frac{d\Omega}{dr}$ enter unchanged; but since

$$dW = ndt, \quad dV = ndt, \quad dV' = 0$$

the term under consideration will produce in $d\Omega$ the term $-M \cdot (i \pm k) n dt \cdot \sin (iW \pm kV \pm lV')$ and in Q , the term

$$\frac{2(i \pm k) n}{(i \pm k) n \pm ln'} M \cos (iW \pm kV \pm lV')$$

so that Q will contain two species of terms, those whose co-efficients are of the form M , and $M \frac{(i \pm k) n}{(i \pm k) n \pm ln'}$

The value of Q substituted in Π (equation 128) will produce terms comprised in one or other of the forms

$$M \cdot \cos (iW \pm kV \pm lV')$$

$$M \cdot \frac{(i \pm k) n}{(i \pm k) n \pm ln'} \cdot \cos (iW \pm kV \pm lV')$$

$$M \cdot \frac{n}{(i \pm k) n \pm ln'} \cdot \cos (iW \pm kV \pm lV')$$

$$M \cdot \frac{(i \pm k) n^2}{\{(i \pm k) n \pm ln'\} \{(i \pm k') n \pm ln'\}} \cdot \cos (iW \pm kV \pm lV')$$

Astronomy. In the process of integration by which δu is derived from Π (or in the integration of the equation (129)) these terms again acquire factors of the form $\frac{1}{\{(i \pm k) n \pm l n'\}^2 - n^2}$ or $\frac{1}{\{(i \pm k') n \pm l n'\}^2 - n^2}$ Physical Astronomy. according as k or k' occurs in the argument; and as these forms are obviously not altered in the transition from δu to δr , the terms of δr will necessarily be included in one of the forms

$$\begin{aligned} M. & \frac{1}{\{(i \pm k) n \pm l n'\}^2 - n^2} \cos (i W \pm k V \pm l V') \\ M. & \frac{(i \pm k') n}{\{(i \pm k) n \pm l n'\} \{(i \pm k) n \pm l n'\}^2 - n^2} \cos (i W \pm k V \pm l V') \\ M. & \frac{n}{\{(i \pm k) n \pm l n'\} \{(i \pm k) n \pm l n'\}^2 - n^2} \cos (i W \pm k V \pm l V') \\ M. & \frac{(i \pm k') n^2}{\{(i \pm k') n \pm l n'\} \{(i \pm k) n \pm l n'\} \{(i \pm k) n \pm l n'\}^2 - n^2} \cos (i W \pm k V \pm l V') \end{aligned}$$

Let the several functions of n, n' , in the co-efficients of these terms be represented indiscriminately by N , then will the general form of the terms of δr be $M. N. \cos (i W \pm k V \pm l V')$.

It remains only to consider the nature of the terms of which $\delta \theta$ consists. Now these will be,

1st. Those arising from the term $\frac{dr}{dt} \delta r$, which are of the form

$$\frac{M. k' n. N.}{n a^2 \sqrt{1 - e^2}} \sin (i W \pm (k \pm k') V \pm l V')$$

2dly. Those arising from $r \frac{d \delta r}{dt}$, whose form is

$$\frac{2 M \{(i \pm k) n \pm l n'\} N}{n a^2 \sqrt{1 - e^2}} \sin (i W \pm (k \pm k') V \pm l V')$$

3dly. Those arising from $\int r \frac{d \Omega}{dr} n dt$ and $\iint d \Omega. n dt$ whose forms are respectively

$$\frac{2 a M}{\mu \sqrt{1 - e^2}} \cdot \frac{n}{(i \pm k) n \pm l n'} \sin (i W \pm k V \pm l V'), \text{ and } \frac{3 a M}{\mu \sqrt{1 - e^2}} \cdot \frac{(i \pm k) n^2}{\{(i \pm k) n \pm l n'\}^2} \sin (i W \pm k V \pm l V')$$

and lastly those peculiar terms containing t out of the signs \sin and \cos which we have already noticed as giving rise to the secular equations.

The complete enumeration of all the possible varieties of terms which $\delta \theta$ may contain, will therefore be had by putting for N each of the four forms above assigned to it; but as those only really differ importantly in which the denominators of the fractions differ, we need only enumerate the latter quantities; which are,

$$\begin{aligned} & (i \pm k) n \pm l n'; \quad \{(i \pm k) n \pm l n'\}^2; \\ & \{(i \pm k) n \pm l n'\}^2 - n^2; \quad \{(i \pm k') n \pm l n'\} \{(i \pm k) n \pm l n'\}^2 - n^2\} \\ & \{(i \pm k') n \pm l n'\} \{(i \pm k) n \pm l n'\} \{(i \pm k) n \pm l n'\}^2 - n^2\} \end{aligned}$$

These then are the various forms of the divisors with which the processes of integration, &c., affect the inequalities in longitude. They are, as we have already remarked, of the highest importance in the theory of the planets, by reason of their effect on the numerical values of the maxima of the perturbations to which they belong. Such is the immense number of terms, or rather of series of terms, branching out in all directions, of which the perturbations consist, that it is manifestly in vain to attempt to take account of them all. It is therefore of the highest consequence to have some guiding principle to direct us in our choice of the terms to be retained or neglected. Were it not for these divisors, we might safely rely on the rapid convergency of the powers and products of the eccentricities; and reject, without further examination, all in which their dimension exceeded a certain limit; but should there be an approach to commensurability in the periodic times of the two planets, (as, for instance, should five times the mean motion of the disturbed planet ($5 n' t$) be very nearly equal to twice that of the disturbing, ($2 n t$)) this circumstance will render some one of their factors ($5 n' - 2 n$) very small. In consequence, all the divisors into which this factor enters will become very small, and the inequalities affected by them will, in consequence, acquire from this cause an unnatural magnitude (if we may use such an expression) and must be retained, even though of such an order as would otherwise authorize their rejection.—The terms so affected too, will originate in a great variety of manners from the developements, and may be affected with various powers of the eccentricities; so that their number will necessarily be infinite, and for the purpose of approximation only the most prominent can be selected.

The equations of the motions of Jupiter and Saturn, known by the name of the great inequalities of these planets, were long a difficulty in the way of the theoretical astronomer, and even a stumbling block in the

Astronomy. way of the Newtonian philosophy. It was observed, on comparing very ancient observations of the oppositions of these planets with more modern ones, that their *mean motions* had undergone an apparent alteration; that of Saturn appearing to have been retarded, and that of Jupiter accelerated. In other words, that Saturn perpetually lagged behind, and Jupiter as constantly surpassed the places, when they ought to have been, on the hypothesis of the mean motion, or periodic time, remaining invariable. **Physical Astronomy.**

We have seen that every inequality of very long period will appear, while on the increase, to affect the mean motion; it is obvious it must, if the latter, as determined by observations comprised within the periods of its increase, be compared with the result of similar observations made while its value is on the diminution. Now, the length of the period of any inequality depends on the multiples of the mean motions found in its argument; and it was not difficult for geometers to shew, that, so far as the first powers or squares of the eccentricities were concerned, no inequalities of such very long periods as the case required, could be found in the motion of either planet. The cubes and higher powers had all along been neglected without fear of error; but Laplace having, from other considerations, ascertained that an acceleration in Jupiter's motion *being supposed*, a retardation in Saturn's *must follow* of course, and that in the very proportion observed; and that therefore the phenomenon was not altogether inconsistent with the laws of gravity, set himself to examine the terms multiplied by the cubes of the eccentricities. Here he immediately encountered the argument ($5n't - 2n't + \text{const.}$); and the mean motion of Jupiter being to that of Saturn nearly in the proportion of 5 to 2, if we suppose n' to correspond to Jupiter's and n to Saturn's motion the co-efficient $5n - 2n'$ is very small, and the corresponding period is found on calculation to amount to 918 years. The resulting inequality has also $5n - 2n'$ for its divisor, and its magnitude is thus increased as well as its period lengthened. On executing the calculation, the inequalities of both planets were found to be such as would completely account for the apparent accelerations and retardations observed.

SECTION VI.

Of the variations of the elements of the planetary orbits, and the secular equations of their motions. Theory of the major axes, inclinations, nodes, eccentricities, and aphelia.

WE have already taken occasion to observe, that the motions of the planets may be regarded as performed in ellipses, whose positions and magnitudes are continually, but very slowly, changing, by the effects of the disturbing forces. These forces are so small, that in a moderate period of time, as for instance, in a single revolution of a planet, the change is insensible; and, if we allow for those inequalities which depend on the configurations of the disturbed and disturbing planets, the theory of which has been exposed in the foregoing sections, the motion, equated by the application of these corrections, will coincide with almost rigorous exactness with the elliptic theory. But after the expiration of many revolutions of the planet, this exact coincidence will cease to take place, even when its place is corrected for such periodical inequalities. The place so corrected is found, it is true, in the circumference of an ellipse with the sun in its focus, but it is not an ellipse of precisely the same form and position as before. Its elements have undergone a change, and this change, though imperceptible in a single revolution, becomes gradually more and more evident, till at length it is too remarkable to be overlooked.

Such slow changes are what we understand by the secular variations of the elements of the planet's orbits. But there is another point of view in which we may consider the subject, which presents peculiar facilities to the application of mathematical investigation. It consists in referring *all* the inequalities resulting from perturbation, to the variation of the elliptic elements, not merely those of long periods, but those which pass rapidly from their maxima to their minima, and depend on the configurations of the bodies. We are indebted to Lagrange for this view of the subject, and shall endeavour to give an idea in this section of the luminous analysis of that great geometer.

If, at the expiration of any instant, the disturbing forces were to cease acting, the planet would go on describing an exact ellipse, of which the infinitesimal arc described in the last instant would be an elementary portion. The plane of the ellipse would be that in which this portion and the sun's centre lie; its eccentricity, position, and magnitude, would all be determined from the *position, magnitude, and curvature* of this element, and the laws of elliptic motion. In a word, it would be a *real ellipse of curvature* to the actual curve described by the planet, at that particular instant, subjected to the conditions of having its focus in the sun, and satisfying the other laws of elliptic motion during that moment. The elements of the planet's orbit then, at any moment, are no other than the elements of this ellipse, and are determined from three consecutive places of the planet, infinitely near each other. Thus every inequality in its motion will produce a corresponding fluctuation in the elements, which will thus be subject to as many equations, periodical or otherwise, as the planet's motion itself is affected with.

The periodical terms thus originating in the elements, will, in the course of many revolutions, compensate each other; but if it should happen that terms not periodical should find their way into their values, these will express *secular* changes, which it becomes of the utmost importance to investigate.

To determine the ellipse of curvature at any instant, is a matter of no difficulty; we have only to call to mind that any one of the constants in its equation may be insulated, and expressed in terms of the co-ordinates and their differential co-efficients, by the mere operations of differentiating and eliminating; so that,

Astronomy supposing the co-ordinates and their differential co-efficients given at any instant, any one of the constants may have its value ascertained by simple substitution. But it is not necessary to go through these processes —we may avoid that trouble by recurring to the origin of the ellipse itself. Now all we know of it is, that it satisfies the dynamical relations of the problem, on the supposition of the disturbing forces ceasing to act at the instant dt . Consequently its equations, however transformed, must be such as to satisfy the differential equations

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$$\frac{d^2 x}{dt^2} + \frac{\mu x}{r^3} = 0$$

$$\frac{d^2 y}{dt^2} + \frac{\mu y}{r^3} = 0$$

$$\frac{d^2 z}{dt^2} + \frac{\mu z}{r^3} = 0$$

and its elements will be the constants introduced by the integration of these, or known functions of them.

For instance, let us consider its major semiaxis. If we pursue with these the same process of integration by which equation (109) was obtained, viz. multiply the first by dx , the second by dy , and the third by dz and add, and integrate, we find

$$\frac{\mu}{a} = \frac{2\mu}{r} - \frac{dx^2 + dy^2 + dz^2}{dt^2}$$

a being the arbitrary constant introduced by integration; and if we compare this with (33), we shall see that a is the semiaxis of the ellipse. Thus we know, that all that is necessary to obtain the semiaxis of the ellipse of curvature at any instant (let the body describe what curve it will) is merely to substitute for

r , $\frac{dx}{dt}$, $\frac{dy}{dt}$, $\frac{dz}{dt}$, in the expression

$$\frac{2\mu}{r} - \left\{ \left(\frac{dx}{dt} \right)^2 + \left(\frac{dy}{dt} \right)^2 + \left(\frac{dz}{dt} \right)^2 \right\}$$

those values which, in the case proposed, they *actually have* in virtue of the real motion of the body, such as the forces in action make it.

Now, in the case of disturbed motion, dx , dy , dz , are given by the equation (94); for if we integrate these after multiplying them respectively by $2dx$, $2dy$, $2dz$, we find

$$\left(\frac{dx}{dt} \right)^2 = -\mu \int \frac{2x dx}{r^3} - 2m \int \frac{d\Omega}{dx} dx$$

$$\left(\frac{dy}{dt} \right)^2 = -\mu \int \frac{2y dy}{r^3} - 2m' \int \frac{d\Omega}{dy} dy$$

$$\left(\frac{dz}{dt} \right)^2 = -\mu \int \frac{2z dz}{r^3} - 2m' \int \frac{d\Omega}{dz} dz$$

consequently adding all together,

$$\begin{aligned} \left(\frac{dx}{dt} \right)^2 + \left(\frac{dy}{dt} \right)^2 + \left(\frac{dz}{dt} \right)^2 &= -\mu \int \frac{2x dx + 2y dy + 2z dz}{r^3} - 2m' \int \left(\frac{d\Omega}{dx} dx + \frac{d\Omega}{dy} dy + \frac{d\Omega}{dz} dz \right) \\ &= \frac{2\mu}{r} - 2m' \int d\Omega \end{aligned}$$

and substituting, we get

$$a = \frac{\mu}{2m' \int d\Omega}, \text{ or } \frac{\mu}{a} = 2m' \int d\Omega; \quad (165, 1)$$

The same result will be obtained as follows:—If we integrate, as in (109), and instead of adding explicitly the arbitrary quantity $\frac{\mu}{a}$ to complete the integral, regard it as included under the sign \int we have

$$\frac{2\mu}{r} - \frac{dx^2 + dy^2 + dz^2}{dt^2} = 2m' \int d\Omega$$

but if a be the major semiaxis of the ellipse of curvature,

$$\frac{\mu}{a} = \frac{2\mu}{r} - \frac{dx^2 + dy^2 + dz^2}{dt^2}$$

Hence we get

$$\frac{\mu}{a} = 2m' \int d\Omega$$

If (a) be the major semiaxis at the commencement of the time t , and $\int a \Omega$ be taken, so as to vanish when

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$$\frac{\mu}{a} = \frac{\mu}{(a)} + 2 m' \int d \Omega$$

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In this instance we have had no difficulty in arriving at once at the finite expression for the varied element. But it is in other cases more commodious to express its *momentary* variation. Let us therefore denote by the characteristic δ , that peculiar variation by which the ellipse of curvature passes from the form and position it had during dt to that which it has in the consecutive instant, then δa , δe , $\delta \pi$, &c. will be the momentary variations of its semiaxis, eccentricity, perihelion, &c. Moreover, δx , δy , δz , and δr , will represent the excesses of the values of x , y , z , r , in the varied ellipse; not over what they *were* in the former instant, but over what they *would have been* had that ellipse remained unaltered. But, both in the one case and the other, the point in the ellipse to which they correspond, coincides with the real place of the planet. Hence the lines x , y , z , r , are the same in the varied ellipse, in the unvaried, and in the actual curve described; so that

$\delta x = 0$, $\delta y = 0$, $\delta z = 0$, and $\delta r = 0$. Again, $\delta \frac{dx}{dt}$, $\delta \frac{dy}{dt}$, $\delta \frac{dz}{dt}$, are the excesses of the values of

$\frac{dx}{dt}$, $\frac{dy}{dt}$, $\frac{dz}{dt}$, in the varied ellipse over what they would have been had the ellipse not varied—that is, had the disturbing force not acted,—in the instant consecutive to dt . Their values therefore would vanish, had the body remained in its former ellipse, and, in general, will be obtained by subtracting from the values actually assumed by $\frac{dx}{dt}$, &c. in the consecutive instant in the curve, what *would have been* assumed by them had the body continued in the same ellipse, or had the disturbing forces ceased to act at the end of dt . Now, in the *curve* the consecutive value of $\frac{dx}{dt}$ is

$$\frac{dx}{dt} + d \frac{dx}{dt} = \frac{dx}{dt} - \left\{ \frac{\mu x}{r^3} + m' \frac{d \Omega}{dx} \right\} dt$$

because $\frac{\mu x}{r^3} + m' \frac{d \Omega}{dx}$ being the force in the curve, we must have $d \frac{dx}{dt} = - \left\{ \frac{\mu x}{r^3} + m' \frac{d \Omega}{dx} \right\} dt$. On

the other hand, had the disturbing force ceased acting, we should have had simply $d \frac{dx}{dt} = - \frac{\mu x}{r^3} dt$,

so that the consecutive value of $\frac{dx}{dt}$ *would have been* merely $\frac{dx}{dt} - \frac{\mu x}{r^3} dt$.

Hence we have

$$\left. \begin{aligned} \delta \frac{dx}{dt} &= - m' \frac{d \Omega}{dx} dt \\ \delta \frac{dy}{dt} &= - m' \frac{d \Omega}{dy} dt \\ \delta \frac{dz}{dt} &= - m' \frac{d \Omega}{dz} dt \end{aligned} \right\} \dots \dots \dots (166)$$

To explain how the variations so obtained may be employed, let us take again the case already treated.

$$\frac{\mu}{a} = \frac{2 \mu}{r} - \frac{d x^2 + d y^2 + d z^2}{d t^2}$$

If we differentiate this relative to the characteristic δ , we get

$$-\frac{\mu \delta a}{a^2} = -\frac{2 \mu \delta r}{r^3} - 2 \left\{ \frac{dx}{dt} \delta \frac{dx}{dt} + \frac{dy}{dt} \delta \frac{dy}{dt} + \frac{dz}{dt} \delta \frac{dz}{dt} \right\}$$

in which, putting $\delta r = 0$, and for $\delta \frac{dx}{dt}$ &c., their values above found, we find

$$-\frac{\mu \delta a}{a^2} = + 2 m' \left\{ \frac{d \Omega}{dx} dx + \frac{d \Omega}{dy} dy + \frac{d \Omega}{dz} dz \right\} = 2 m' d \Omega$$

δa is the *momentary* variation of a in the instant dt , so that this equation may be integrated relative to t , and we get

$$\frac{\mu}{a} = 2 m' \int d \Omega$$

the same result as before.

Before we proceed farther, we will stop to draw from this expression of the reciprocal axis, a most important conclusion. It is this—that all the variations to which the major axes of the planetary orbits are subjected by their mutual attraction are periodical—and that the mean distances, and consequently the mean motions of the planets are subject to no *secular* variations. In fact, when we consider only the first power of the disturbing forces, we have already proved that the development of Ω is entirely composed of terms of the form

$$A \cdot \cos (i W + k V + l V')$$

hence

$$\begin{aligned} d \Omega &= A \cdot \left(i \frac{d W}{n dt} + k \frac{d V}{n dt} \right) \cdot \sin (i W + k V + l V') \\ &= A \cdot (i + k) \cdot \sin \{ (i + k) n t - (i - l) n' t + (i + k) e - (l - l') e' - k \pi - l \pi' \} \end{aligned}$$

Astronomy. which is always periodic unless $i + k = 0$ and $i - l = 0$, when the argument becomes simply $i(\pi - \pi')$ a **Physical Astronomy.** constant quantity; but in this case the whole term vanishes by the disappearance of its co-efficient. Thus

$d\Omega$ and, of course, $\int d\Omega$ contains no term multiplied by t , and none but what is periodic; consequently,

$\frac{\mu}{a}$ is periodic also.

This beautiful result, the demonstration of which is of almost elementary simplicity, assures us of the impossibility of any of the bodies of our system ever leaving it in consequence of the disturbances it may experience, and secures the general permanence of the whole, by keeping the mean distances and periodic times perpetually fluctuating between certain limits (very restricted ones) which they can neither exceed nor fall short of.

Let us next consider the variation in the position of the plane of the disturbed orbit. The equations (4), (5), and (6), give

$$h = y \frac{dx}{dt} - x \frac{dy}{dt}, \quad h' = z \frac{dx}{dt} - x \frac{dz}{dt}, \quad h'' = z \frac{dy}{dt} - y \frac{dz}{dt}$$

These quantities in the case of elliptic motion are constant, but in that of disturbed motion they will equally hold good, if h, h', h'' , be regarded as variable. Now, either on the one or the other supposition, if we multiply the first by z , the second by $-y$, and the third by x , and add, we get

$$hz - h'y + h''x = 0$$

which is the same with equation (7). But if we multiply the first by dz , the second by $-dy$, and the third by dx , we shall also obtain by addition,

$$h dz - h' dy + h'' dx = 0$$

Consequently, even when we regard h, h', h'' as variable, still the equation $hz - h'y + h''x = 0$ and its differential relative to x, y, z , subsist together, just as if h, h', h'' were constant. Hence, it appears, that the body at the end of the instant dt is still found in the plane represented by $hz - h'y + h''x = 0$, or, that this plane is the plane in which the elementary arc described in the instant dt , lies. If therefore we call ϕ its inclination to that of the x and y taken as a fixed plane, and ω the longitude of its ascending node, we have

$$\tan \phi = \frac{\sqrt{h'^2 + h''^2}}{h}; \quad \tan \omega = \frac{h''}{h'}; \quad \mu a (1 - e^2) = h^2 + h'^2 + h''^2; \quad (167)$$

whence these elements (viz. the inclination, the longitude of the node, and the semiparameter,) are expressed in terms of h, h', h'' .

That the equation $h dz - h' dy + h'' dx = 0$ must hold good at the same time with $hz - h'y + h''x = 0$ is also evident from this consideration, that although the ellipse, it is true, varies from one instant to another, yet it must be regarded as invariable, while the body describes each of its elementary portions; because, by hypothesis, it is so adjusted that the body shall remain in its circumference during the whole of the instant dt , and it is not till the consecutive instant that it is necessitated to change its form, &c. to accommodate itself to the new course taken by the body. The same reasoning holds for any other finite equation of elliptic motion. Its first differential may be taken, as if the arbitrary constants it involves were rigorously such. The disturbing forces make no change on x, y, z, r , their consecutive values remain $x + dx, y + dy, z + dz, r + dr$, as before, and are common both to the curve and the ellipse, it is for the consecutive values of dx, dy, dz , that they differ, these becoming $dx + d^2x + \delta dx$ for the curve, and $dx + d^2x$ for the ellipse.

This premised, we have only to inquire the variations of h, h', h'' ; and to this end, by the equations (4, 5, 6,) we have

$$\delta h = y \delta \frac{dx}{dt} - x \delta \frac{dy}{dt}; \quad \delta h' = z \delta \frac{dx}{dt} - x \delta \frac{dz}{dt}; \quad \delta h'' = z \delta \frac{dy}{dt} - y \delta \frac{dz}{dt}.$$

In which, substituting for $\delta \frac{dx}{dt}, \delta \frac{dy}{dt}, \delta \frac{dz}{dt}$, their values $-m' \frac{d\Omega}{dx} dt, -m' \frac{d\Omega}{dy} dt$, and $m' \frac{d\Omega}{dz} dt$, we find

$$\left. \begin{aligned} \delta h &= m' \left\{ x \frac{d\Omega}{dy} - y \frac{d\Omega}{dx} \right\} dt; \\ \delta h' &= m' \left\{ x \frac{d\Omega}{dz} - z \frac{d\Omega}{dx} \right\} dt; \\ \delta h'' &= m' \left\{ y \frac{d\Omega}{dz} - z \frac{d\Omega}{dy} \right\} dt; \end{aligned} \right\} \dots\dots\dots (168)$$

and integrating relative to t ,

$$h = (h) + m' \int \left\{ x \frac{d\Omega}{dy} - y \frac{d\Omega}{dx} \right\} dt; \quad (169, 1)$$

$$h' = (h') + m' \int \left\{ x \frac{d\Omega}{dz} - z \frac{d\Omega}{dx} \right\} dt; \quad (169, 2)$$

$$h'' = (h'') + m' \int \left\{ y \frac{d\Omega}{dz} - z \frac{d\Omega}{dy} \right\} dt; \quad (169, 3)$$

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As these expressions are rigorous, we may derive from them all the laws which regulate the motion of the nodes and the inclinations of the planes of the orbits; but as it is only the secular variations which concern us at present, we shall not regard the periodical parts of the expressions within the brackets under the integral signs. To developpe them, we must consider the orbits as inclined to the plane of the x, y ; but if we take the undisturbed orbit of m for this plane, the value of z at any time t will be of the order of the disturbing forces, and z' will be a very small quantity; so that z^2 , $z z'$, and z'^2 , may be neglected. Hence r ($= \sqrt{x^2 + y^2 + z^2}$) and r' , will represent with this degree of approximation, their projections on the plane of the x, y , and we have (calling s the tangent of m 's latitude,)

$$x = r \cdot \cos \theta, \quad y = r \cdot \sin \theta, \quad z = r s; \quad x' = r' \cdot \cos \theta', \quad y' = r' \cdot \sin \theta', \quad z' = r' s'$$

If then we recur to the expressions for the disturbing forces $m' \frac{d\Omega}{dx}$, &c. in Section I. Part II. we shall find

$$y \frac{d\Omega}{dx} - x \frac{d\Omega}{dy} = (y x' - x y') \left(\frac{1}{r'^3} - \frac{1}{\lambda^3} \right); \quad (170, 1)$$

$$z \frac{d\Omega}{dx} - x \frac{d\Omega}{dz} = (z x' - x z') \left(\frac{1}{r'^3} - \frac{1}{\lambda^3} \right); \quad (170, 2)$$

$$z \frac{d\Omega}{dy} - y \frac{d\Omega}{dz} = (z y' - y z') \left(\frac{1}{r'^3} - \frac{1}{\lambda^3} \right); \quad (170, 3)$$

Now, since $h z - h' y + h'' x = 0$, we have $z = -\frac{h''}{h} x + \frac{h'}{h} y$

Suppose then, $\frac{h''}{h} = p$ and $\frac{h'}{h} = q$, and let the quantities corresponding to p and q in the orbit of m' be p' and q' ; then, if ϕ, ϕ' , be the inclinations of the two orbits to the fixed plane, and w, w' , the longitudes of their ascending nodes, we shall have $\tan \phi = \sqrt{p^2 + q^2}$, $\tan w = \frac{p}{q}$, $\tan \phi' = \sqrt{p'^2 + q'^2}$, $\tan w' = \frac{p'}{q'}$, and thus, when p and q are determined, the inclinations and places of the nodes are easily found. We have, moreover,

$$\begin{aligned} z &= q y - p x, & z' &= q' y' - p' x' \\ z x' - x z' &= (p' - p) x x' + q y x' - q' x y' \\ z y' - y z' &= - (q' - q) y y' - p x y' + p' y x' \end{aligned}$$

and if we therefore suppose for a moment $M = \frac{1}{r'^3} - \frac{1}{\lambda^3}$,

$$\begin{aligned} \frac{d h}{d t} &= m' M \cdot (y x' - x y') \\ \frac{d h'}{d t} &= m' M \cdot \{ (p' - p) x x' + q y x' - q' x y' \} \\ \frac{d h''}{d t} &= m' M \cdot \{ - (q' - q) y y' - p x y' + p' y x' \} \end{aligned}$$

Now, our design being to eliminate h, h', h'' , from the formulæ, and obtain expressions involving only p and q , from which h, h', h'' , may be deduced if wanted, we differentiate the equations $p = \frac{h''}{h}$ and $q = \frac{h'}{h}$, when we find

$$\frac{d p}{d t} = -\frac{1}{h} \left(\frac{d h''}{d t} - p \frac{d h}{d t} \right) \quad \frac{d q}{d t} = \frac{1}{h} \left(\frac{d h'}{d t} - q \frac{d h}{d t} \right)$$

and if we substitute in these, the values of $\frac{d h}{d t}$, $\frac{d h'}{d t}$, and $\frac{d h''}{d t}$, as above found, we shall get

$$\begin{aligned} \frac{d p}{d t} &= m' M \cdot \frac{(p' - p) y x' - (q' - q) y y'}{h} = \frac{m' M}{h} y \{ (p' - p) x' - (q' - q) y' \} \\ \frac{d q}{d t} &= m' M \cdot \frac{(p' - p) x x' - (q' - q) x y'}{h} = \frac{m' M}{h} x \{ (p' - p) x' - (q' - q) y' \} \end{aligned}$$

But if we neglect the squares of the disturbing forces, and the eccentricities and inclinations of the orbits, we

have $\frac{m'}{h} = \frac{m'}{\sqrt{(1+m)(1-e^2)} \cdot a} = \frac{m'}{\sqrt{a}}$, so that these expressions become

$$\frac{d p}{d t} = m' M \cdot \frac{(p' - p) x' - (q' - q) y'}{\sqrt{a}} y \quad (171)$$

$$\frac{d q}{d t} = m' M \cdot \frac{(p' - p) x' - (q' - q) y'}{\sqrt{a}} x$$

Let us suppose the function $(a^2 - 2 a a' \cdot \cos w + a'^2)^{-\frac{3}{2}}$ developed in a series of cosines of w , and its multiples

$$S + S' \cdot \cos w + S'' \cdot \cos 2 w + \&c.$$

then we shall have, by (135)

$$\left. \begin{aligned} S &= \frac{1}{a^3} \left\{ 1 + \left(\frac{3}{2} \right)^2 a^2 + \left(\frac{3 \cdot 5}{2 \cdot 4} \right)^2 a^4 + \left(\frac{3 \cdot 5 \cdot 7}{2 \cdot 4 \cdot 6} \right)^2 a^6 + \&c. \right\} \\ S' &= \frac{2}{a^3} \left\{ \frac{3}{2} a + \frac{3^2 \cdot 5}{2^2 \cdot 4} a^3 + \frac{3^2 \cdot 5^2 \cdot 7}{2^3 \cdot 4^2 \cdot 6} a^5 + \&c. \right\} \\ S'' &= \frac{2}{a^3} \left\{ \frac{3 \cdot 5}{2 \cdot 4} a^2 + \frac{3^2 \cdot 5 \cdot 7}{2^2 \cdot 4 \cdot 6} a^4 + \&c. \right\} \end{aligned} \right\} \dots\dots\dots (173)$$

and if we neglect in the present research, as is allowable, the eccentricities and inclinations, or suppose the orbits circular and in one plane,

$$M = \frac{1}{a^3} - \frac{1}{\lambda^3} = \left(\frac{1}{a^3} - S \right) - S' \cdot \cos (\theta' - \theta) - S'' \cdot \cos 2 (\theta' - \theta) - \&c. \quad (174)$$

$$r = a, \quad r' = a', \quad \theta = n t + \epsilon, \quad \theta' = n' t + \epsilon', \quad \theta - \theta' = n t - n' t + \epsilon - \epsilon' = W$$

$$\begin{aligned} y \{ (p' - p) x' - (q' - q) y' \} &= a a' \cdot \frac{q' - q}{2} \{ \cos (\theta' + \theta) - \cos (\theta' - \theta) \}; \\ &+ a a' \cdot \frac{p' - p}{2} \{ \sin (\theta' + \theta) - \sin (\theta' - \theta) \} \end{aligned} \quad (175)$$

$$\begin{aligned} x \{ (p' - p) x' - (q' - q) y' \} &= a a' \cdot \frac{p' - p}{2} \{ \cos (\theta' + \theta) + \cos (\theta' - \theta) \}; \\ &- a a' \cdot \frac{q' - q}{2} \{ \sin (\theta' + \theta) + \sin (\theta' - \theta) \} \end{aligned} \quad (176)$$

Each of these latter quantities is to be combined with M by multiplication, and in resolving each of the products of sines and cosines so originating into simple sines and cosines of sums and differences, it is obvious that constant terms will arise whenever similar terms are combined, by reason of the property $\cos A \times \cos A = \frac{1}{2} \cos 2 A + \frac{1}{2}$. Now the only argument common to both factors is $\theta' - \theta$; and, of course, the only terms in (174) and (175) from whose combination a constant term can originate, are $- S' \cdot \cos (\theta' - \theta)$ and $- a a' \cdot \frac{q' - q}{2} \cdot \cos (\theta' - \theta)$. Consequently, if we reject all the periodical terms, and put $I = \frac{a a'}{4} S'$, we have

$$\frac{d p}{d t} = \frac{m'}{\sqrt{a}} \cdot I \cdot (q' - q); \quad (177, 1)$$

and similarly

$$- \frac{d q}{d t} = \frac{m'}{\sqrt{a}} \cdot I (p' - p); \quad (177, 2)$$

If we go through a process exactly analogous, so as to obtain differential equations for determining p' and q' relative to the orbit of m' , we shall find them to be

$$\frac{d p'}{d t} = \frac{m}{\sqrt{a'}} I (q - q'); \quad (177, 3)$$

$$- \frac{d q'}{d t} = \frac{m}{\sqrt{a'}} I (p - p'); \quad (177, 4)$$

These four equations, being of the first order, and with constant co-efficients (for the secular variations of a, a' , and therefore of S' and of $\frac{a a'}{4} \times S'$ or I , which are symmetrical functions of a, a' , vanish,) are easily integrated. As they subsist simultaneously among the four variables p, p', q, q' , we may integrate them all together, if we assume

$$\begin{aligned} p &= A \cdot \sin (g t + k) & p' &= A' \cdot \sin (g t + k) \\ q &= A \cdot \cos (g t + k) & q' &= A' \cdot \cos (g t + k) \end{aligned}$$

for if we substitute these values, we find that the variable part divides off, and there remain the following equations of condition between the constants A, A' , and g ,

$$g A = \frac{m'}{\sqrt{a}} I (A' - A); \quad (178, 1)$$

$$g A' = \frac{m}{\sqrt{a'}} I (A - A'); \quad (178, 2)$$

In these one of the constants A, A', g , remains indeterminate. Let this be A , then eliminating A' , we get for determining g ,

$$\left. \begin{aligned} g^2 + \left(\frac{m'}{\sqrt{a}} + \frac{m}{\sqrt{a'}} \right) I g &= 0 \\ g = 0, \text{ or } g &= - \frac{m \sqrt{a} + m' \sqrt{a'}}{\sqrt{a} a'} \cdot I \end{aligned} \right\} \dots\dots\dots (179, 1, 2)$$

Now it is evident, that if g and g' be two values of g which satisfy the equations of condition, then $A \cdot \sin (g t + k)$ and $B \cdot \sin (g' t + k')$ will each be satisfactory values of p , and so for the rest; consequently (the equations being linear) their sum will be so, and we have

$$\begin{cases} p = A \cdot \sin (g t + k) + B \cdot \sin k' & (180, 1) \\ q = A \cdot \cos (g t + k) + B \cdot \cos k' & (180, 2) \end{cases}$$

$$\begin{cases} p' = A' \cdot \sin (g t + k) + B \cdot \sin k' & (180, 3) \\ q' = A' \cdot \cos (g t + k) + B \cdot \cos k' & (180, 4) \end{cases}$$

where A, B, k, k' , are four arbitrary constants, and

$$g = - I \cdot \frac{m \sqrt{a} + m' \sqrt{a'}}{\sqrt{a} a'}; \quad A' = - \frac{m \sqrt{a}}{m' \sqrt{a'}} \cdot A; \quad (181, 1, 2)$$

From these values of p and q, p' and q' it is easy to eliminate $\sin (g t + k)$ and $\cos (g t + k)$; for if we multiply (180, 1) by $m \sqrt{a}$, and (180, 3) by $m' \sqrt{a'}$, and add, noticing that $A \cdot m \sqrt{a} + A' \cdot m' \sqrt{a'} = 0$ by reason of the relation between A and A' (181, 2) we get

$$m \sqrt{a} \cdot p + m' \sqrt{a'} \cdot p' = (m \sqrt{a} + m' \sqrt{a'}) B \cdot \sin k' = \text{constant}; \quad (182)$$

and similarly,

$$m \sqrt{a} \cdot q + m' \sqrt{a'} \cdot q' = (m \sqrt{a} + m' \sqrt{a'}) B \cdot \cos k' = \text{const.}; \quad (183)$$

Moreover, we have

$$\tan \phi^2 = p^2 + q^2 = (A^2 + B^2) + 2 AB \cdot \cos (g t + k - k'); \quad (184, 1)$$

$$\tan \phi'^2 = p'^2 + q'^2 = (A'^2 + B^2) + 2 A'B \cdot \cos (g t + k - k'); \quad (184, 2)$$

Consequently,

$$m \sqrt{a} \cdot \tan \phi^2 + m' \sqrt{a'} \cdot \tan \phi'^2 = m \sqrt{a} (A^2 + B^2) + m' \sqrt{a'} (A'^2 + B^2) = \text{constant}; \quad (185)$$

The arbitrary constants A, B, k, k' , may be determined in any particular case, either by comparing the general expressions for p, p', q, q' , with their actual values at any assigned instant as derived from observation, or from these last derived equations; for since $p^2 + q^2 = \tan \phi^2$ and $\frac{p}{q} = \tan \omega$, we have

$$p = \tan \phi \cdot \sin \omega; \quad q = \tan \phi \cdot \cos \omega; \quad p' = \tan \phi' \cdot \sin \omega'; \quad q' = \tan \phi' \cdot \cos \omega'; \quad (186)$$

Now the equation (180) gives

$$(p' - p) = (A' - A) \cdot \sin (g t + k); \quad q' - q = (A' - A) \cdot \cos (g t + k)$$

consequently, if we take $t = 0$, or if we assume for our data, the elements $\phi, \phi', \omega, \omega'$, as they were observed at the epoch or origin of the time t , we find

$$\tan k = \frac{p' - p}{q' - q} = \frac{\tan \phi' \cdot \sin \omega' - \tan \phi \cdot \sin \omega}{\tan \phi' \cdot \cos \omega' - \tan \phi \cdot \cos \omega}; \quad (187)$$

Hence A and A' are found; for, the values of S, S' , &c. being known from equations (173), $I = \frac{a a'}{4} S'$ is also known; and since by (181, 2)

$$A' - A = - \frac{m \sqrt{a} + m' \sqrt{a'}}{m' \sqrt{a'}} \cdot A \text{ we get } A = - \frac{m' \sqrt{a'} (p' - p)}{(m \sqrt{a} + m' \sqrt{a'}) \sin k} \quad (188)$$

Again, if we divide (182) by (183), we find

$$\tan k' = \frac{m \sqrt{a} \cdot \tan \phi \cdot \sin \omega + m' \sqrt{a'} \cdot \tan \phi' \cdot \sin \omega'}{m \sqrt{a} \cdot \tan \phi \cdot \cos \omega + m' \sqrt{a'} \cdot \tan \phi' \cdot \cos \omega'}; \quad (189)$$

$$B = \frac{m \sqrt{a} \cdot \tan \phi \cdot \sin \omega + m' \sqrt{a'} \cdot \tan \phi' \cdot \sin \omega'}{(m \sqrt{a} + m' \sqrt{a'}) \cdot \sin k'}; \quad (190)$$

These constants once computed, the laws, periods, and limits of the motions of the planes of both orbits are known. The period in which the inequalities recur is deducible at once from the value of g . If we express the time t in Julian years, n, n' represent the mean motions of the planets m, m' , in one such year, in parts

Astronomy of a whole circumference, and $n^2 a^3 = 1$, that is, $\sqrt{a} = \frac{1}{na}$, hence $g \left(= -I \cdot \left(\frac{m}{\sqrt{a}} + \frac{m'}{\sqrt{a'}} \right) \right)$ will be Physical Astronomy

found as follows :—

$$g = -(m n a' + m' n a) \cdot I$$

and if we call T the whole period, $-g T = 1$ circumference = 1 ; and

$$T = -\frac{1}{g} = \frac{1}{(m n' a' + m' n a) \cdot I} ; \quad (191)$$

where

$$I = \frac{a a'}{4} S'$$

If n, n' be expressed in seconds, the numerator of T instead of being unity, must be $360 \times 60 \times 60'' = 2196000''$.

The limits of the variations of the inclinations are readily found ; for it appears from the equations (184, 1, 2) that their maxima and minima occur when $g t + k - k' = 0$ and 180° , and have for their corresponding values

$$\left. \begin{array}{l} A + B \\ A' + B' \end{array} \right\} \text{ and } \left. \begin{array}{l} A - B \\ A' - B' \end{array} \right\} ; \dots\dots\dots (192)$$

Now, it is obvious from the values of p and q , (180, 1, 2) that A and B are small quantities of the same order as p, q ; so that the inclinations can never increase or diminish beyond certain very narrow limits. This follows too from the equation (185) ; for in the present state of our system, $\tan \phi$ and $\tan \phi'$ being extremely small, the sum $m \sqrt{a} \cdot \tan \phi^2 + m' \sqrt{a'} \cdot \tan \phi'^2$ is always a very minute quantity ; and since \sqrt{a} and $\sqrt{a'}$

must both be taken positively, (for $\sqrt{a} = \frac{1}{na}$ and n is positive for both planets, the motions being both in the same direction) neither term separately can exceed the value of the constant ; so that ϕ or ϕ' must remain for ever confined to a value not greatly different from what it now has, and the planes of the planetary orbits must keep for ever oscillating within very confined limits about their mean positions.

With regard to the nodes it is different. These are liable to great changes of place, and may even circulate for ever in one direction without returning. In fact, if we would determine the maxima and minima of their longitudes, we have only to put $d w = 0, d w' = 0$; the roots of which equations, if real, will indicate the stationary points ; and, if imaginary, will shew that such points do not exist, or that the nodes circulate. Now we have

$$d w = \frac{d \cdot \tan w}{1 + \tan^2 w} = 0, \text{ or } d \cdot \tan w = 0, \text{ or } \frac{d p}{p} = \frac{d q}{q}, \text{ or } q \frac{d p}{d t} - p \frac{d q}{d t} = 0$$

in which equation, substituting for $\frac{d p}{d t}$ and $\frac{d q}{d t}$ their values in equations (177, 1, 2) we find

$$q q' - q^2 + p p' - p^2 = 0, \text{ or } p p' + q q' = p^2 + q^2 ; \quad (193)$$

in which, substituting for p, p', q, q' , their values (180) we get ultimately

$$A + B \cdot \cos (g t + k - k') = 0 \text{ which gives } \cos (g t + k - k') = -\frac{A}{B} ; \quad (194)$$

Hence, if $B > A$ (no regard being had to the signs) this will correspond to a real value of t , and the node will then merely have a libratory motion, advancing and receding alternately ; if $B < A$, they will circulate always in one direction. In the former case, if we substitute in the value of $\tan \phi^2$ (equation 184, 1) this value of $\cos (g t + k - k')$ we shall find $\tan \phi^2 = B^2 - A^2$; $\tan \phi = \sqrt{B^2 - A^2}$, which gives the inclination corresponding to the stationary points of the node. These points are attained when $\cos (g t + k - k') = -\frac{A}{B}$

while the maxima and minima of the inclinations happen when $\cos (g t + k - k') = \pm 1$. The stationary positions of the node therefore do not correspond either to the maxima and minima of the inclinations, or to the semi-intervals between them.

If we had considered more than two bodies the results would have been analogous, and we should have arrived at similar expressions for p and q only containing more terms, and analogous equations to those in 182, 183, and 185, viz.

$$\left. \begin{array}{l} m \sqrt{a} \cdot p + m' \sqrt{a'} \cdot p' + m'' \sqrt{a''} \cdot p'' + \&c. = \text{const.} \\ m \sqrt{a} \cdot q + m' \sqrt{a'} \cdot q' + m'' \sqrt{a''} \cdot q'' + \&c. = \text{const.} \\ m \sqrt{a} \cdot \tan \phi^2 + m' \sqrt{a'} \cdot \tan \phi'^2 + m'' \sqrt{a''} \cdot \tan \phi''^2 + \&c. = \text{const.} \end{array} \right\} , \dots\dots\dots (195)$$

Let us apply this theory to an example, and we will take that of the orbits of Jupiter and Saturn, the two principal planets of our system. If we take for our epoch the year 1700, we have, by Halley's tables,

$$w = 101^\circ 5' 6'' ; \phi = 2^\circ 30' 10'' ; a = 9.54007$$

$$w' = 97^\circ 34' 9'' ; \phi' = 1^\circ 19' 10'' ; a' = 5.20098$$

Hence we find the values of p, p', q, q' , for that epoch, by the equations (186) as follows :—

$$p = 0.04078, q = -0.01573 ; p' = 0.02283, q' = -0.00303$$

whence, having computed I , and assuming $m' = \frac{1}{1067}$ and $m = \frac{1}{3358}$, we shall find A, B, k, k' and g as follows :

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$B = 0.02905$; $A = 0.01537$; $k = 125^\circ 15' 40''$, $k' = 103^\circ 38' 40''$
 and $A' = -0.00661$, and finally $g = -25''.5756$

Hence we obtain, in the case of Saturn,

$$\tan \phi = 0.03287 \cdot \sqrt{1 + 0.82665 \cdot \cos \{21^\circ 37' - t \times 25''.5756\}}$$

and for Jupiter

$$\tan \phi' = 0.02980 \cdot \sqrt{1 - 0.43290 \cdot \cos \{21^\circ 37' - t \times 25''.5756\}}$$

Also we have $B + A = 0.04442$; $B - A = 0.01368$, so that the maxima and minima of the inclinations of Saturn's orbit are $2^\circ 32' 40''$ and $0^\circ 47'$, and its greatest deviation from the mean state will not exceed $52' 50''$. In Jupiter's orbit the maximum is $2^\circ 2' 30''$, and the minimum $1^\circ 17' 10''$, and the greatest deviation from a mean state $0^\circ 22' 40''$.

The longitude of the node, ω , has a maximum and a minimum in both orbits, because $B > A$; and the extent of its librations will be, in the case of Jupiter's orbit, $13^\circ 9' 40''$, and in that of Saturn's, $31^\circ 56' 20''$ on either side of its mean station, on the plane of the ecliptic supposed immovable.

The period in which these changes take place, or the whole time in which the inclinations vary from their greatest to their least values, and the nodes from their greatest to their least longitudes, and back again, is equal to $\frac{360^\circ}{g} = \frac{360^\circ}{25''.5756} = 50673$ Julian years; an immense period, and which may serve to give some

idea of the extent to which the Newtonian theory, assisted by the refined methods of the modern analysis, enables us to carry our views of the past and future condition of our system; as this, though subject of course to some subordinate corrections, is perhaps one of the least uncertain of the results of perturbation.

Let us now consider the secular variations of the eccentricities and aphelia. Our first object, agreeable to the theory of the variation of the arbitrary constants already exposed, must be to obtain such an equation of the ellipse of curvature as shall be adapted to our purpose, by containing these elements (or convenient functions of them) in a state sufficiently disengaged from the variables x, y, z , and their differentials. Now if we

differentiate the quantities $\frac{x}{r}, \frac{y}{r}, \frac{z}{r}$, noticing that $r^2 = x^2 + y^2 + z^2$, we get

$$\begin{aligned} d \frac{x}{r} &= \frac{r dx - x dr}{r^2} = \frac{r^2 dx - x \cdot r dr}{r^3} = \frac{r^2 dx - \frac{1}{2} x d(r^2)}{r^3} \\ &= \frac{(x^2 + y^2 + z^2) dx - (x dx + y dy + z dz) \cdot x}{r^3} \\ &= \frac{y(y dx - x dy) + z(z dx - x dz)}{r^3} \end{aligned}$$

That is, substituting for $y dx - x dy$ and $z dx - x dz$ their values $h dt, h' dt$,

$$\begin{aligned} d \left(\frac{x}{r} \right) &= \frac{h y + h' z}{r^3} dt \\ \text{similarly, } d \left(\frac{y}{r} \right) &= \frac{-h x + h'' z}{r^3} dt \\ d \left(\frac{z}{r} \right) &= \frac{-h' x - h'' y}{r^3} dt \end{aligned} \quad \left\{ \begin{array}{l} \mu d \left(\frac{x}{r} \right) = \left(h \frac{\mu y}{r^3} + h' \frac{\mu z}{r^3} \right) dt \\ \mu d \left(\frac{y}{r} \right) = \left(-h \frac{\mu x}{r^3} + h'' \frac{\mu z}{r^3} \right) dt \\ \mu d \left(\frac{z}{r} \right) = \left(-h' \frac{\mu x}{r^3} - h'' \frac{\mu y}{r^3} \right) dt \end{array} \right.$$

Now, in the case of elliptic motion, $\frac{\mu x}{r^3} = -\frac{d^2 x}{dt^2}$; $\frac{\mu y}{r^3} = -\frac{d^2 y}{dt^2}$ and $\frac{\mu z}{r^3} = -\frac{d^2 z}{dt^2}$, consequently substituting these, we obtain

$$\begin{aligned} \mu d \left(\frac{x}{r} \right) &= d \left\{ -h \frac{dy}{dt} - h' \frac{dz}{dt} \right\}; \mu d \left(\frac{y}{r} \right) = d \left\{ h \frac{dx}{dt} - h'' \frac{dz}{dt} \right\} \\ \mu d \left(\frac{z}{r} \right) &= d \left\{ h' \frac{dx}{dt} + h'' \frac{dy}{dt} \right\} \end{aligned}$$

and integrating,

$$f = \mu \frac{x}{r} + \left(h \frac{dy}{dt} + h' \frac{dz}{dt} \right); \quad (196, 1)$$

$$f' = \mu \frac{y}{r} + \left(-h \frac{dx}{dt} + h'' \frac{dz}{dt} \right); \quad (196, 2)$$

$$f'' = \mu \frac{z}{r} + \left(-h' \frac{dx}{dt} - h'' \frac{dy}{dt} \right); \quad (196, 3)$$

These equations will serve our purpose, as the arbitrary constants f, f', f'' , are completely disengaged; but before we proceed to employ them, we must determine the values of f, f', f'' , in terms of the elements, and *vice versa*; and for this purpose must first eliminate the differential co-efficients, which (from the peculiar form of the equations) is practicable, by merely multiplying the first by x , the second by y , and the third by

Astronomy. z ; and adding the results, when we get

$$\mu r - h \frac{y dx - x dy}{dt} - h' \cdot \frac{z dx - x dz}{dt} - h'' \cdot \frac{z dy - y dz}{dt} = f x + f' y + f'' z$$

or, by reason of the equations (4) (5) (6),

$$\mu r = f x + f' y + f'' z + (h^2 + h'^2 + h''^2); \quad (197)$$

This equation expresses the general property of the conic sections, in virtue of which a line drawn from the focus to the circumference is always in a given ratio to a perpendicular let fall from that point on the directrix.

If we multiply (196, 1) by h'' , (196, 2) by $-h'$, and (196, 3) by h , and take the sum of the results, (observing that $h'' x - h' y + h z = 0$) it will be found that all the variable terms will destroy each other, leaving simply the equation of condition,

$$h'' f - h' f' + h f'' = 0 \quad (198)$$

Suppose X, Y, Z , the co-ordinates of the perihelion. At this point $dr = 0$, or $x dx + y dy + z dz = 0$, but if for h, h', h'' , we write their values in (4) (5) (6) we have

$$h \frac{dy}{dt} + h' \frac{dz}{dt} = -x \cdot \frac{dy^2 + dz^2}{dt^2} + dx \cdot \frac{y dy + z dz}{dt^2} \quad (199, 1)$$

$$-h \frac{dx}{dt} + h' \frac{dz}{dt} = -y \frac{dx^2 + dz^2}{dt^2} + dy \cdot \frac{x dx + z dz}{dt^2} \quad (199, 2)$$

$$-h' \frac{dx}{dt} - h'' \frac{dy}{dt} = -z \frac{dx^2 + dy^2}{dt^2} + z \frac{x dx + y dy}{dt^2} \quad (199, 3)$$

At the perihelion therefore, substituting for $y dy + z dz$, $x dx + z dz$, $x dx + y dy$, their equals $-X dX$, $-Y dY$, and $-Z dZ$, respectively, these quantities become

$$-X \cdot \frac{dX^2 + dY^2 + dZ^2}{dt^2}$$

$$-Y \cdot \frac{dX^2 + dY^2 + dZ^2}{dt^2}$$

$$-Z \cdot \frac{dX^2 + dY^2 + dZ^2}{dt^2}$$

Consequently we have, putting $\frac{dX^2 + dY^2 + dZ^2}{dt^2} = V^2$, $X^2 + Y^2 + Z^2 = R^2$

$$f = X \left(\frac{\mu}{R} - V^2 \right); \quad f' = Y \cdot \left(\frac{\mu}{R} - V^2 \right); \quad f'' = Z \left(\frac{\mu}{R} - V^2 \right); \quad (200)$$

Hence it is easy to obtain the following equation,

$$f^2 + f'^2 + f''^2 = R^2 \left(\frac{\mu}{R} - V^2 \right)^2; \quad (201)$$

Now R is the perihelion distance, $R = a(1 - e)$ and V being the velocity at the perihelion, we have, by

$$(33) V^2 = \mu \left(\frac{2}{R} - \frac{1}{a} \right) \quad \text{Hence we get}$$

$$f^2 + f'^2 + f''^2 = R^2 \left(\frac{\mu}{a} - \frac{\mu}{R} \right)^2 = \frac{\mu^2 (R - a)^2}{a^2} = \mu^2 e^2$$

$$\mu e = \sqrt{f^2 + f'^2 + f''^2}; \quad (202)$$

Moreover we have

$$\frac{f'}{f} = \frac{Y}{X}, \quad \frac{f''}{f} = \frac{Z}{X}; \quad \frac{f''}{\sqrt{f^2 + f'^2}} = \frac{Z}{\sqrt{X^2 + Y^2}}; \quad (203)$$

But $\frac{Y}{X}$ is the tangent of the longitude of the perihelion, or of the angle which the projection of the perihelion distance makes with the axis of the x : also, π is the longitude of the perihelion reckoned on the orbit; and if its plane is but very little inclined to that of the x, y , this angle differs from its projection on that plane only by a very small quantity of the second order; so that if we disregard, as we have hitherto done in this research, the squares of the eccentricities, inclinations, and disturbing forces, we have

$$\tan \pi = \frac{f'}{f}; \quad (204)$$

These three equations, viz. $\tan \pi = \frac{f'}{f}$, $\mu e = \sqrt{f^2 + f'^2 + f''^2}$, and $h f'' - h' f' + h'' f = 0$, determine f, f', f'' , in terms of the elements of the orbit. The variations of the eccentricities and longitudes of the

Astronomy. perihelia, e and π , may be immediately determined from those of f and f' , so that it is to these we shall now turn our attention. Resuming then the equations (196, 1, 2, 3) let us differentiate them on the supposition of f, f', f'' , being variable quantities, their variations being such as to express the effect of the disturbing forces only. To express this we have used the characteristic δ , and we shall continue to do so, to keep the principle on which the process is founded distinctly in view. No inconvenience can arise from the confusion of two symbols, δ and d , in the same investigation, when we bear in mind that any expression such as δf denotes, strictly, the whole amount of the momentary change which the quantity f undergoes while the planet passes from one elementary portion of its unknown curve to that immediately consecutive to it, while such expressions as dx, dy , &c., denote the changes which x, y , &c. undergo while it passes from one end to the other of the same elementary portion. In both points of view, the accumulated effects during a finite time are obtained

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by the same rules of integration; the whole variation of f is legitimately expressed by $\int \delta f$ just as that of x is by $\int dx$, and when δf is expressed in terms of dt , the integration must be performed in the usual manner, as on a function of t .

We have therefore, since $\delta \frac{x}{r} = 0, \delta \frac{y}{r} = 0, \delta \frac{z}{r} = 0$

$$\delta f = \frac{dy}{dt} \delta h + \frac{dz}{dt} \delta h' + h \delta \frac{dy}{dt} + h' \delta \frac{dz}{dt}$$

Substituting therefore for $\delta h, \delta h'$ their values in the equations (168) for $\delta \frac{dy}{dt}, \delta \frac{dz}{dt}, -m' \frac{d\Omega}{dy}$ and $-m' \frac{d\Omega}{dz}$ respectively; and for h, h' , their values $\frac{y dx - x dy}{dt}$ and $\frac{z dx - x dz}{dt}$, we get

$$\delta f = - \left\{ dy \left\{ y \frac{d\Omega}{dx} - x \frac{d\Omega}{dy} \right\} + dz \left\{ z \frac{d\Omega}{dx} - x \frac{d\Omega}{dz} \right\} \right\} \times m' \dots \dots \dots (205, 1)$$

and similarly,

$$\delta f' = - \left\{ dx \left\{ x \frac{d\Omega}{dy} - y \frac{d\Omega}{dx} \right\} + dz \left\{ z \frac{d\Omega}{dy} - y \frac{d\Omega}{dz} \right\} \right\} \times m' \dots \dots \dots (205, 2)$$

These are the rigorous values of $\delta f, \delta f'$, and their integrals regarded as functions of t , express the total variations of these elements produced in that time. But, as we have done in the less complicated theory of the nodes and inclinations we shall neglect, in developing them, all their periodical terms, at least such as depend on the configurations of the planets, as well as all terms containing the squares of the disturbing forces, eccentricities, and inclinations, and take the primitive orbit of m for our fixed plane. This will simplify them greatly; for in this case z , and $\frac{d\Omega}{dz}$ are quantities of the order of the disturbing forces, and therefore, when multiplied by m' , may be rejected. Moreover, on this supposition, it is indifferent whether we reckon the longitudes on the orbit or on the fixed plane, since the difference is of the order z^2 ; consequently, representing that of m by θ , we have

$$x = r \cdot \cos \theta, y = r \cdot \sin \theta;$$

whence we find

$$\begin{aligned} x \frac{d\Omega}{dy} - y \frac{d\Omega}{dx} &= (y'x - x'y) \cdot \left(\frac{1}{r^3} - \frac{1}{\lambda^3} \right) \\ &= r' \cdot \sin (\theta' - \theta) \cdot \left(\frac{1}{r^3} - \frac{1}{\lambda^3} \right) = \frac{d\Omega}{d\theta}; \end{aligned} \quad (206)$$

$$\delta f = - dy \frac{d\Omega}{d\theta} - h \frac{d\Omega}{dy} dt$$

$$\text{Now we have } \frac{d\Omega}{dy} = \frac{y'}{r^3} - \frac{y' - y}{\lambda^3} = \frac{\sin \theta'}{r^2} - \frac{r' \cdot \sin \theta' - r \cdot \sin \theta}{\lambda^3}$$

But since

$$\lambda = \sqrt{r^2 - 2rr' \cdot \cos (\theta - \theta') + r'^2}$$

$$\Omega = \frac{r \cdot \cos (\theta - \theta')}{r'^2} - \frac{1}{\lambda}$$

$$\frac{d\Omega}{dr} = \frac{\cos (\theta - \theta')}{r^2} + \frac{r - r' \cdot \cos (\theta - \theta')}{\lambda^3} \quad (207)$$

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$$\frac{d\Omega}{d\theta} = -\frac{r \cdot \sin(\theta - \theta')}{r'^2} + \frac{r r' \cdot \sin(\theta - \theta')}{\lambda^3} \quad (208) \quad \text{Physical Astronomy}$$

Consequently

$$\frac{d\Omega}{dy} = \frac{1}{r} \left\{ r \frac{d\Omega}{dr} \cdot \sin \theta + \frac{d\Omega}{dt} \cdot \cos \theta \right\} \quad (209)$$

and as we have $dy = dr \cdot \sin \theta + r d\theta \cdot \cos \theta$, and $h dt = r^2 d\theta$,

$$\delta f = - \left\{ dr \cdot \sin \theta + 2r d\theta \cdot \cos \theta \right\} \frac{d\Omega}{d\theta} - r^2 d\theta \frac{d\Omega}{dr} \cdot \sin \theta; \quad (210)$$

In like manner we should have found

$$\delta f' = + \left\{ dr \cdot \cos \theta - 2r d\theta \cdot \sin \theta \right\} \frac{d\Omega}{d\theta} + r^2 d\theta \frac{d\Omega}{dr} \cdot \cos \theta; \quad (211)$$

Now, if we neglect the squares of the eccentricities, we have

$$r = a \{ 1 - e \cdot \cos(\theta - \pi) \}, \quad dr = ae \cdot \sin(\theta - \pi) d\theta, \quad r^2 d\theta = na^2 dt,$$

whence we have

$$d\theta = n dt (1 + 2e \cdot \cos(\theta - \pi)) \text{ and } dr = nae dt \cdot \sin(\theta - \pi)$$

and consequently,

$$\delta f = -na dt \cdot \frac{d\Omega}{d\theta} \{ (2 - 2e \cdot \cos \overline{\theta - \pi})(1 + 2e \cdot \cos \overline{\theta - \pi}) \cdot \cos \theta + e \cdot \sin \theta \cdot \sin \overline{\theta - \pi} \} - na^2 dt \cdot \frac{d\Omega}{dr};$$

which reduced, rejecting e^2 , becomes

$$\delta f = -m' n dt \left\{ a \frac{d\Omega}{d\theta} \left(2 \cos \theta + \frac{3}{2} e \cdot \cos \pi + \frac{e}{2} \cdot \cos(2\theta - \pi) \right) + a^2 \frac{d\Omega}{dr} \cdot \sin \theta \right\} \quad (212)$$

In order to find the secular part of the variation of f , we must developpe this expression, retaining only such terms as are not periodical. Now, recurring to the notation of Sections 3, 4, 5, Part II. we have

$$\theta = nt + e + 2e \cdot \sin V = V + \pi + 2e \cdot \sin V$$

therefore

$$\sin \theta = \sin(V + \pi) + 2e \cdot \sin V \cdot \cos(V + \pi)$$

$$= \sin(V + \pi) + e \cdot \sin(2V + \pi) - e \cdot \sin \pi$$

$$2 \cos \theta = 2 \cos(V + \pi) - 4e \cdot \sin V \cdot \sin(V + \pi)$$

$$= 2 \cos(V + \pi) + 2e \cos(2V + \pi) - 2e \cdot \cos \pi$$

$$\frac{e}{2} \cdot \cos(2\theta - \pi) = \frac{e}{2} \cdot \cos(2V + \pi); \quad \frac{3e}{2} \cdot \cos \pi = \frac{3e}{2} \cdot \cos \pi$$

neglecting e^2 &c. Again we have

$$a^2 \frac{d\Omega}{dr} = a^2 \frac{d\Omega}{da} + a^2 \Delta \frac{d\Omega}{dr}$$

When this therefore is multiplied by $\sin \theta$, the term $-e \cdot \sin \pi$ will combine with the constant term $a^2 \frac{d\Omega}{da}$ of $a^2 \frac{d\Omega}{da}$, and produce $-a^2 e \cdot \frac{d\Omega}{da} \cdot \sin \pi$, and this is the only constant term which can arise from

$a^2 \frac{d\Omega}{da}$, because every other term contains W , and therefore both V and V' ; while the value of $\sin \theta$ contains only V , and, of course, V' cannot be destroyed by their combination. Again, with respect to $\Delta \frac{d\Omega}{dr}$, it has been proved, that (regarding only the first powers of the eccentricities,) this is resolvable

into a series of terms of the form $A \cdot \cos(iW \pm V)$, $A \cdot \cos(iW \pm V')$. Now, since $\Delta \frac{d\Omega}{dr}$ is of the order of the eccentricities, we may disregard the term of $\sin \theta$ multiplied by e , and take it simply equal to $\sin(V + \pi)$. If we multiply this into the series $\Sigma A \cdot \cos(iW \pm V)$ since $W = V - V' + \pi - \pi'$, only one term can produce a constant argument, viz. that into which V' does not enter, or in which $i = 0$; that is, the term multiplied by $\cos V$. The co-efficient in this term in $\Delta \frac{d\Omega}{dr}$ is $-ae \frac{d^2 A}{da^2}$, (as appears from (147),

writing $\frac{d\Omega}{dr}$ for Ω , and of course $\frac{dA}{da}$, &c. for A , &c.) Hence the term so originating in $a^2 \frac{d\Omega}{dr} \cdot \sin \theta$ will

be equal to the constant part of $-a^3 e \frac{d^2 A}{da^2} \cdot \cos V \cdot \sin(V + \pi)$, that is, to $-\frac{a^3 e}{2} \cdot \frac{d^2 A}{da^2} \cdot \sin \pi$. Again the multiplication of $\sin(V + \pi)$ with the series $\Sigma A \cdot \cos(iW \pm V')$ will produce a constant term by the combination of $\sin(V + \pi)$ with $\cos(W + V') = \cos(V + \pi - \pi')$. The co-efficient of $\cos(W + V')$ in the developement of $\Delta \frac{d\Omega}{dr}$ is $-\frac{e'}{2} \left(a \frac{d^2 A'}{da da'} + 2 \frac{dA'}{da} \right)$. Hence the constant part of $a^2 \frac{d\Omega}{dr} \cdot \sin \theta$ so

$$\underbrace{\text{Astronomy}} \text{originating, will be } -\frac{a^2 e'}{2} \left(a' \frac{d^2 A'}{d a d a'} + 2 \frac{d A'}{d a} \right) \times (\text{the constant part of } \sin(V + \pi) \cdot \cos(V + \pi - \pi')) \underbrace{\text{Physical Astronomy.}} \\ = -a^2 \cdot \frac{e'}{4} \sin \pi' \left\{ 2 \frac{d A'}{d a} + a' \frac{d^2 A'}{d a d a'} \right\}$$

Let us next consider the constant terms which can originate from the part multiplied by $a \frac{d \Omega}{d \theta}$. This quantity is equal to $a \frac{d \Omega}{d w} \cdot \frac{d w}{d \theta} = a \frac{d \Omega}{d w} = a \frac{d \Omega}{d W} + a \Delta \frac{d \Omega}{d w}$. But we have

$$\frac{d \Omega}{d W} = -A \cdot \sin W - 2 A' \cdot \sin 2 W - \&c.$$

and as all these contain V' , and this arc does not enter into any of the terms multiplying $\frac{d \Omega}{d \theta}$ in the expression of δf , it is obvious that no combination not containing V' , can arise in this way, *i. e.* no constant term can occur in the part of the developement depending on $\frac{d \Omega}{d W}$. With regard to $\Delta \frac{d \Omega}{d w}$ being already of the order of the eccentricities, it need only be combined with the first term, $2 \cos \theta$, and this term may be regarded as equal simply to $2 \cdot \cos(V + \pi)$. We have only then to investigate the constant part of

$$2 a \cdot \cos(V + \pi) \left\{ \Delta \frac{d \Omega}{d w} = \frac{d^2 \Omega}{d a d W} \Delta r + \frac{d^2 \Omega}{d a' d W} \Delta r' + \frac{d^2 \Omega}{d W^2} \Delta w \right\}$$

Now $\Delta r = -a e \cdot \sin V$, $\Delta r' = -a' e' \cdot \sin V'$, $\Delta w = 2 e \cdot \sin V - 2 e' \cdot \sin V'$

$$1st. \text{ The combination } -2 a^2 e \cdot \cos V \cdot \cos(V + \pi) \cdot \Sigma -i \frac{d A^i}{d a'} \sin i W$$

can contain no constant term, because the terms $\sin V$ and $\cos(V + \pi)$ cannot eliminate V' from $\sin i W$.

$$2dly. -2 a a' e' \cdot \cos(V + \pi) \cdot \cos V' \cdot \Sigma -i \frac{d A^i}{d a'} \sin i W$$

will contain a constant term, for taking $i = 1$, $\cos V' \times \sin W = \frac{1}{2} \sin(W + V') + \frac{1}{2} \sin(W - V')$, the last term of which does not contain V' , being equal to $+\frac{1}{2} \sin(V + \pi - \pi')$. The combination now under consideration, will therefore contain the term

$$+ a a' e' \cdot 1 \cdot \frac{d A'}{d a'} \cdot \cos(V + \pi) \cdot \sin(V + \pi - \pi')$$

which, developed, will produce a constant term,

$$- a a' \cdot \frac{e}{2} \cdot \frac{d A'}{d a'} \cdot \sin \pi'$$

Lastly, the combination

$$4 a \cdot \cos(V + \pi) \{ e \cdot \sin V - e' \cdot \sin V' \} \cdot \Sigma i^2 \cdot A^i \cdot \cos i W$$

will produce a constant term; viz. that arising from

$$-4 a e' \cdot \cos(V + \pi) \cdot \sin V' \times -1^2 A' \cdot \cos W$$

for $\sin V' \cdot \cos W$ produces a term free from V' , viz. $\frac{1}{2} \sin(V + \pi - \pi')$ which again combined with $\cos(V + \pi)$ produces $\frac{-1}{4} \sin \pi'$; so that the complete result of this combination will be

$$- a e' \cdot A' \cdot \sin \pi'$$

Assembling therefore all the constant terms so found, we get

$$\delta f = m' \cdot a n d t \cdot e \sin \pi \left\{ a \frac{d A}{d a} + \frac{a^2}{2} \frac{d^2 A}{d a^2} \right\} \quad (213) \\ + m' \cdot a n d t \cdot e' \sin \pi' \left\{ A' + \frac{a}{2} \frac{d A'}{d a} + \frac{a'}{2} \frac{d A'}{d a'} + \frac{a a'}{4} \cdot \frac{d^2 A'}{d a d a'} \right\}$$

The value of $\delta f'$ may be derived by a similar process, but without executing the whole process, we may obtain it more readily by considering the expressions (210) and (211) we then see that δf changes to $\delta f'$ simply by changing θ to $\theta - 90^\circ$, but leaving r , $\frac{d \Omega}{d r}$ and $\frac{d \Omega}{d \theta}$ unchanged. Now Ω , $\frac{d \Omega}{d r}$ and $\frac{d \Omega}{d \theta}$ remain manifestly

unaltered, whatever direction we assume as that from which we reckon the longitudes. Suppose then that we shift this direction 90° , then will each of the angles θ , θ' , e , e' , π , π' , be diminished 90° , and this will produce no alteration in r , because $\theta - \pi$ will remain unchanged. Hence, the effect produced by this change on δf will transform it into $\delta f'$. The value of $\delta f'$ therefore will be had at once by changing $\sin \pi$ and $\sin \pi'$ into $-\cos \pi$, and $-\cos \pi'$ in that of δf ; and this will be found to be confirmed by the direct process.

Suppose now we denote the constant co-efficients within the brackets for brevity by $-I$ and K , so as to have

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$$-I = a \frac{dA}{da} + \frac{a^2}{2} \cdot \frac{d^2 A}{da^2} \quad (214) \quad \text{Physical Astronomy.}$$

$$K = A' + \frac{a}{2} \frac{dA'}{da} + \frac{a'}{2} \cdot \frac{dA'}{da'} + \frac{aa'}{4} \cdot \frac{d^2 A'}{da da'} \quad (215)$$

Then it will be observed, that K is obviously symmetrical with respect to a, a' , because the co-efficient A is so; and the same will be found to hold good with I ; for if we write it as follows,—

$$I = -\frac{1}{2} \frac{d}{da} \left\{ a^2 \frac{dA}{da} \right\} \quad (216)$$

and for A substitute its value given in (137, 1) it will be found that the execution of the operations here indicated will lead to precisely the same result as those denoted by $\frac{d}{da'} \left\{ a'^2 \frac{dA}{da'} \right\}$: consequently I and K are both symmetrical functions of a , and a' , and do not alter when these elements are mutually transposed. With regard to I , it may be easily shewn to be identical with the function $\frac{aa'}{4} S'$ which we have already denoted by the same symbol. To see this, we need only write for A and S' their values in (136, 1) and (173, 2), and executing on the former series the operations indicated by $-\frac{1}{2} \cdot \frac{d}{da} \left(a^2 \frac{d}{da} \right)$, (remembering that $a = \frac{a'}{a}, \frac{da}{da} = -\frac{a}{a}; \frac{d^2 a}{da^2} = +\frac{2a}{a^3}$) and multiplying the latter by $\frac{aa'}{4}$, the result will be identical. We have then

$$\begin{aligned} \delta f &= -m' \cdot na \cdot Idt \times e \cdot \sin \pi + m' \cdot na \cdot Kdt \times e' \sin \pi \\ \delta f' &= +m' \cdot na \cdot Idt \times e \cos \pi - m' \cdot na \cdot Kdt \times e' \cos \pi' \end{aligned} \quad (217)$$

Now, if we recur to the values of f, f', f'' , we find

$$\frac{f'}{f} = \tan \pi; \quad \frac{f''}{\sqrt{f^2 + f'^2}} = \sin \pi, \quad f'' = \frac{h'f' - h''f}{h}$$

but h' and h'' are of the order of the disturbing forces when we take the plane of m 's undisturbed orbit for our fixed plane; consequently, f'' is so, and therefore $\sqrt{f^2 + f'^2} = \sqrt{f^2 + f'^2} = \mu e$, neglecting the squares of the disturbing forces. Hence we have $f = \mu e \cdot \cos \pi$ and $f' = \mu e \cdot \sin \pi$.

In imitation of what we did in the theory of the nodes and inclinations, let q and p represent these quantities; then dq and dp being the momentary variations of q and p , will replace δf and $\delta f'$; and if in like manner we put q' and p' for $\mu' e' \cdot \cos \pi'$ and $\mu' e' \cdot \sin \pi'$, the corresponding quantities in the orbit of m' , we shall have (putting $\frac{1}{\sqrt{a}}, \frac{1}{\sqrt{a'}}$, for na, na' ; and supposing $\mu = 1$, and $\mu' = 1$, which is allowable since the squares of the disturbing forces are neglected, and the quantities under consideration are all of the first order of these forces,)

$$\left. \begin{aligned} + \frac{dp}{dt} &= \frac{m'}{\sqrt{a}} \left\{ Iq - Kq' \right\} \\ - \frac{dq}{dt} &= \frac{m'}{\sqrt{a}} \left\{ Ip - Kp' \right\} \\ + \frac{dp'}{dt} &= \frac{m}{\sqrt{a'}} \left\{ Iq' - Kq \right\} \\ - \frac{dq'}{dt} &= \frac{m}{\sqrt{a'}} \left\{ Ip' - Kp \right\} \end{aligned} \right\} \dots\dots\dots (218, 1, 2, 3, 4)$$

Thus we see that the problem of the secular variations of the eccentricities and apelia, depends, exactly as in the case of the nodes and inclinations, on the simultaneous integration of four differential equations of the first order with constant co-efficients. If we compare these, however, with those of (177), we shall observe that the former research is rather more complicated than the latter, by reason of the two co-efficients, I and K , which it involves, while the system of equations on which the nodes and inclinations depend, involves only the first. However, this circumstance does not render the integration more difficult; the same substitution succeeds, and the integrals have a form almost exactly similar; we have only to take

$$\left. \begin{aligned} p &= A \cdot \sin(gt + k) + B \cdot \sin(ht + l) \\ q &= A \cdot \cos(gt + k) + B \cdot \cos(ht + l) \\ p' &= A' \cdot \sin(gt + k) + B' \cdot \sin(ht + l) \\ q' &= A' \cdot \cos(gt + k) + B' \cdot \cos(ht + l) \end{aligned} \right\} \dots\dots\dots (219)$$

If these be substituted, the equations will be satisfied, provided the following equations hold good between the constants,

$$\left. \begin{aligned} g A &= \frac{m'}{\sqrt{a}} (IA - KA') \\ g A' &= \frac{m}{\sqrt{a'}} (IA' - KA) \end{aligned} \right\} \text{ and } \left\{ \begin{aligned} h B &= \frac{m'}{\sqrt{a'}} (IB - KB') \\ h B' &= \frac{m}{\sqrt{a'}} (IB' - KB) \end{aligned} \right.$$

It is obvious that the elimination of A' from the first pair, and of B' from the second pair, will lead to equations of exactly the same form for g and h , which are therefore the two roots of the quadratic,

$$g^2 - I g \cdot \frac{m \sqrt{a} + m' \sqrt{a'}}{\sqrt{a a'}} + \frac{m m'}{\sqrt{a a'}} (I^2 - K^2) = 0 \quad (220)$$

and g and h being found, and A, B, k, l , remaining arbitrary, A' and B' are easily found.

To adapt these values to any particular case, the general values of p, q, p', q' , at any assigned epoch, must be made to coincide with the observed values of $e \cdot \sin \pi, e \cdot \cos \pi, e' \cdot \sin \pi',$ and $e' \cdot \cos \pi'$; which condition will furnish equations to determine all the arbitrary constants.

For example, in the case of Jupiter and Saturn, we shall find on computation $g = 21''.9905, h = 3''.5851, A = 0.04877; B = 0.03532; A' = -0.01715; B' = 0.04321; k = 306^\circ 34' 40''; l = 210^\circ 16' 40''$. Now, since $p = e \cdot \sin \pi$, and $q = e \cdot \cos \pi$, we have

$$\tan \pi = \frac{p}{q}, \text{ and } e = \sqrt{p^2 + q^2} = \sqrt{A^2 + B^2 + 2AB \cos \{(g-h)t + k-l\}}$$

which gives, by substituting the numbers already found,

$$e = 0.06021 \cdot \sqrt{1 - 0.95009 \times \cos (83^\circ 42' - 18''.4054 \times t)}$$

which is the eccentricity of Saturn's orbit; and similarly,

$$e' = 0.04649 \cdot \sqrt{1 + 0.68592 \cdot \cos (83^\circ 42' - 18''.4054 \times t)}$$

for that of Jupiter, after any number t of years since 1700.

The longitudes of their respective apelia, will also be known by the formulæ

$$\tan \pi = \frac{A \cdot \sin (gt + k) + B \cdot \sin (ht + l)}{A \cdot \cos (gt + k) + B \cdot \cos (ht + l)}; \quad (221)$$

$$\tan \pi' = \frac{A' \cdot \sin (gt + k) + B' \cdot \sin (ht + l)}{A' \cdot \cos (gt + k) + B' \cdot \cos (ht + l)}; \quad (222)$$

and the maxima and minima of these, or their greatest deviations from their mean places will take place, when

$$gA^2 + hB^2 + AB(g+h) \cdot \cos \{(g-h)t + (k-l)\} = 0$$

that is, when

$$\cos \{(g-h)t + (k-l)\} = -\frac{gA^2 + hB^2}{(g+h)AB}; \quad (223)$$

If this fraction be less than unity, the apelia will librate, as in the case of the nodes about a mean position, if not, they move in one direction continually. In the case of Jupiter and Saturn now before us, $gA^2 + hB^2 > 7(g+h)AB$, so that the apelia go on for ever in one direction.

The period in which the eccentricities go through all their evolutions, and return to the same state, is represented by $\frac{360^\circ}{g-h} = \frac{360^\circ}{18''.4054} = 70414$ Julian years.

The greatest and least values of the eccentricities are respectively $A \pm B$ and $A' \pm B'$, for the two planets; and in the case before us

for Saturn, 0.08409 and 0.01345 } the maximum of the one corresponding to the minimum of
and for Jupiter, 0.06036 and 0.02606 } the other planet.

Finally, we may derive relations between the eccentricities, masses, and semiaxes, similar to those obtained between the inclinations, &c. in equations (182, 183, &c.) for since $p^2 + q^2 = e^2$ it is easy to see that we must have

$$m \cdot e^2 \sqrt{a} + m' \cdot e'^2 \sqrt{a'} = m \sqrt{a} (A^2 + B^2) + m' \sqrt{a'} (A'^2 + B'^2) + 2(m \sqrt{a} \cdot AB + m' \sqrt{a'} \cdot A'B') \cdot \cos \{(g-h)t + (k-l)\} \quad (224)$$

Now we have, from the equations of condition,

$$\left. \begin{aligned} A \left(g - \frac{m'}{\sqrt{a}} \cdot I \right) &= -A' \cdot \frac{m'}{\sqrt{a}} K \\ B \left(h - \frac{m'}{\sqrt{a}} \cdot I \right) &= -B' \cdot \frac{m'}{\sqrt{a}} K \end{aligned} \right\}; \dots\dots\dots (225)$$

so that

$$AB \left(g - \frac{m'}{\sqrt{a}} I \right) \left(h - \frac{m'}{\sqrt{a}} I \right) = A' B' \cdot \frac{m'^2}{a} K^2$$

Astronomy. but g and h being the roots of (220), we have

$$gh = \frac{m m'}{\sqrt{a a'}} (I^2 - K^2) \quad g + h = I \cdot \frac{m \sqrt{a} + m' \sqrt{a'}}{\sqrt{a a'}}$$

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Consequently, our equation, on executing the multiplications indicated, and substituting, reduces itself to

$$m \sqrt{a} \cdot AB + m' \sqrt{a'} \cdot A'B' = 0$$

so that the equation (224) becomes simply

$$m e^2 \cdot \sqrt{a} + m' e'^2 \sqrt{a'} = m \sqrt{a} (A^2 + B^2) + m' \sqrt{a'} (A'^2 + B'^2) = \text{constant}, \quad (226)$$

because the major axes are constant; and had we considered a greater number of planets than two, we should, in like manner, have arrived at the equation

$$m e^2 \sqrt{a} + m' e'^2 \sqrt{a'} + m'' e''^2 \sqrt{a''} + \&c. = \text{constant}; \quad (227)$$

PART III.

OF THE THEORY OF THE MOON.

SECTION I.

Rigorous investigation of the differential equations of the moon's motion, and general expression of the disturbing forces.

THE extreme minuteness of the masses of the planets, as well as their great distance, renders it allowable in the theory of their perturbations to neglect altogether the squares of the disturbing forces, and affords such facilities to the whole investigation, as to permit us to express at once the true longitude and radius vector in explicit functions of the mean longitude or the time. This is not the case in the lunar theory, in which one of the most remarkable of the inequalities depends for nearly half its value on the square of the disturbing force, and in which the whole investigation is so delicate, as to render it necessary to abandon the direct process followed in the planetary theory and adopt a route, apparently more circuitous, but possessing advantages of a peculiar kind. It consists in expressing the radius vector and the time or *mean* longitude in functions of the *true*, and thence by the reversion of series deducing the true longitude in terms of the mean. The advantage of this is, that we are thus enabled to set out from differential equations in which nothing has been neglected, and consequently have it in our power fully to appreciate the influence of all the terms we may afterwards neglect in their integration on the final result.

Our directing principle in this investigation will be to follow out, step by step, with the proper modifications, the same system of transformations, by which the differential equations 1, 2, 3, of undisturbed motion were converted into (9) and (11) expressing the radius vector and time in terms of θ .

Our first step will consist in the investigation of an equation corresponding to (9): to this end (as we have supposed $d t$ constant) multiply the first of the equations (94) by y , and the second by $-x$, and add, and we get

$$0 = \frac{y d^2 x - x d^2 y}{d t^2} + m' \cdot \left\{ y \frac{d \Omega}{d x} - x \frac{d \Omega}{d y} \right\}$$

and integrating,

$$\frac{y d x - x d y}{d t} = h + m' \int \left\{ x \frac{d \Omega}{d y} - y \frac{d \Omega}{d x} \right\} d t; \quad (228)$$

but $y d x - x d y = \rho^2 d \theta$, so that this equation becomes

$$\rho^2 \frac{d \theta}{d t} = h + m' \int T \rho d t \quad (229)$$

if we assume

$$T = \frac{x \frac{d \Omega}{d y} - y \frac{d \Omega}{d x}}{\rho}. \quad (230)$$

The function T is the measure, and $m' \cdot T$ the quantity, of what, in the lunar theory, is called the tangential force, or that part of the disturbing force acting on m , which is perpendicular to the direction of the radius vector.

This is evident, if we consider that $m' \frac{d \Omega}{d y}$ and $m' \frac{d \Omega}{d x}$ representing the disturbing forces in the directions QP

Fig. A. and PM (fig. A) if we resolve these each into the direction KQ perpendicular to QM they will become respectively

$$-m' \frac{x}{\rho} \cdot \frac{d \Omega}{d y} \text{ and } +m' \frac{y}{\rho} \cdot \frac{d \Omega}{d x}$$

so that their aggregate, or the whole *tangential force*, will have for its expression, the quantity above represented by T in equation (230).

The equation (229), if m' were zero, would coincide with (9), and would express the proportionality of the

Astronomy areas to the times. The term $m' \int T \rho \, dt$ then expresses the momentary effect of the disturbing forces in Physical Astronomy.
deranging this proportionality; and since

$$\int \rho^3 \, d\theta = h(t + \text{const.}) + m' \int dt \int T \rho \, dt \quad (231)$$

the term $m' \int dt \int T \rho \, dt$ expresses their total effect after the lapse of the time t ; and we see that this effect takes place solely in virtue of the tangential force, agreeably to what Newton has advanced in the 11th section of the *Principia*.

Our next step is to investigate an equation for the disturbed motion, corresponding to (11) in the undisturbed. The process we followed in the deduction of (11), it will be remembered, consisted

1st. In making θ the independent variable.

2dly. In eliminating t and its differentials.

Now, if in the equations (94) instead of supposing $d\theta$ constant, we regard it as variable, we must change $\frac{d^2 x}{dt^2}$ into $\frac{1}{dt} \frac{d}{dt} \frac{dx}{dt}$ or $\frac{d^2 x}{dt^2} - \frac{dx}{dt} \frac{d^2 t}{dt^2}$, and so for $\frac{d^2 y}{dt^2}$, &c. This done, let the first of the equations (94) be multiplied by x , and the second by y ; their sum will be

$$0 = \left\{ (x d^2 x + y d^2 y) - (x dx + y dy) \cdot \frac{d^2 t}{dt} \right\} \frac{1}{dt^2} \\ + \mu \cdot \frac{x^2 + y^2}{r^3} + m' \cdot \left\{ x \frac{d\Omega}{dx} + y \frac{d\Omega}{dy} \right\}$$

but since $x = \rho \cdot \cos \theta$, and $y = \rho \cdot \sin \theta$, $r = \sqrt{x^2 + y^2 + z^2}$
 $= \sqrt{\rho^2 + z^2}$

this equation will become (supposing $d\theta$ constant),

$$0 = \left\{ \rho d^2 \rho - \rho^2 d\theta^2 - \rho d\rho \cdot \frac{d^2 t}{dt} \right\} \frac{1}{dt^2} \\ + \mu \cdot \frac{\rho^2}{(\rho^2 + z^2)^{\frac{3}{2}}} + m' \left\{ x \frac{d\Omega}{dx} + y \frac{d\Omega}{dy} \right\} \quad (232)$$

We have only now to find the value of $\frac{d^2 t}{dt}$ on the supposition of $d\theta$ being invariable, and substitute it in this equation; and for this purpose we must employ our equation (229) which gives (multiplying it by $T \rho \, dt$)

$$T \rho^3 \, d\theta = h \cdot T \rho \, dt + m' \cdot T \rho \, dt \cdot \int T \rho \, dt$$

and integrating,

$$\int T \rho^3 \, d\theta = h \int T \rho \, dt + \frac{m'}{2} \left(\int T \rho \, dt \right)^2$$

from which we get, by the solution of a quadratic equation,

$$m' \cdot \int T \rho \, dt = -h + \sqrt{h^2 + 2m' \int T \rho^3 \, d\theta}$$

and differentiating, and dividing by $m' T \rho$

$$dt = \frac{\rho^2 \, d\theta}{\sqrt{h^2 + 2m' \int T \rho^3 \, d\theta}}; \quad (233)$$

If we take the logarithmic differential on both sides of this, supposing $d\theta$ constant, we get

$$\frac{d^2 t}{dt} = 2 \frac{d\rho}{\rho} \cdot m' \cdot \frac{T \rho^3 \, d\theta}{h^2 + 2m' \int T \rho^3 \, d\theta}; \quad (234)$$

Substituting this then in equation (232), and for $\frac{1}{dt^2}$, writing its value $\frac{h^2}{\rho^4 \, d\theta^2} + 2m' \cdot \frac{\int T \rho^3 \, d\theta}{\rho^4 \, d\theta^2}$ given by (233), it becomes

$$0 = \frac{1}{\rho \, d\theta^2} \left\{ \frac{d^2 \rho}{\rho^2} - 2 \frac{d\rho^2}{\rho^3} - \frac{d\theta^2}{\rho} \right\} (h^2 + 2m' \int T \rho^3 \, d\theta) \\ + \frac{\mu}{\rho} \cdot \frac{1}{\left(1 + \frac{z^2}{\rho^2}\right)^{\frac{3}{2}}} + m' \cdot \left\{ \left(x \frac{d\Omega}{dx} + y \frac{d\Omega}{dy} \right) + 1 \cdot \frac{d\rho}{d\theta} \right\}, \quad (235)$$

Astronomy In this equation let $\frac{z}{\rho} = s$, or $z = \rho s$, so that $s = \tan$ heliocentric elevation of m above the plane of the Physical Astronomy.
 x and y , $= \tan m$ MQ fig. (A). Moreover, let

$$V = \frac{x \frac{d\Omega}{dx} + y \frac{d\Omega}{dy}}{\rho}; \quad (236)$$

then will V be the measure, and $m' \cdot V$ the quantity, of that part of the disturbing force which operates to increase the gravity of m towards the central body. For if we resolve the forces $m' \cdot \frac{d\Omega}{dx}$ and $m' \cdot \frac{d\Omega}{dy}$ which act in the directions of the co-ordinates x and y into the direction of the distance ρ , they will become respectively

$$+ m' \cdot \frac{x}{\rho} \cdot \frac{d\Omega}{dx} \text{ and } + m' \cdot \frac{y}{\rho} \cdot \frac{d\Omega}{dy}$$

whose sum is $m' \cdot V$.

These substitutions made, and ρ being put (as in the elliptic theory) equal to $\frac{1}{\mu}$, the equation will become when multiplied by $-\frac{\rho}{h^2}$,

$$0 = \frac{d^2 u}{d\theta^2} + u - \frac{\mu}{h^2} \times \frac{1}{(1+s^2)^{\frac{3}{2}}} - \frac{m'}{h^2} \left\{ \frac{V}{u^2} - \frac{T}{u^3} \cdot \frac{du}{d\theta} - 2 \left(\frac{d^2 u}{d\theta^2} + u \right) \int \frac{T d\theta}{u^3} \right\}; \quad (237)$$

which equation is rigorous, nothing having been neglected in the previous process. But if the squares of the disturbing forces be neglected, we have $\frac{1}{(1+s^2)^{\frac{3}{2}}} = 1$ because s or $\frac{z}{\rho}$ is of the order of the disturbing forces; and since, in virtue of this very equation,

$$\begin{aligned} m' \left(\frac{d^2 u}{d\theta^2} + u \right) &= m' \left\{ \frac{\mu}{h^2} \cdot \frac{1}{(1+s^2)^{\frac{3}{2}}} + m' \times \&c. \right\} \\ &= m' \cdot \frac{\mu}{h^2} + m'^2 \times \&c. \end{aligned}$$

the equation becomes

$$0 = \frac{d^2 u}{d\theta^2} + u - \frac{\mu}{h^2} - \frac{m'}{h^2} \left\{ \frac{V}{u^2} - \frac{T}{u^3} \cdot \frac{du}{d\theta} - \frac{2\mu}{h^2} \int \frac{T d\theta}{u^3} \right\}; \quad (238)$$

Now $\mu = 1 + m$, and if we neglect the powers and products of the disturbing forces and the mass m , the equation becomes

$$0 = \frac{d^2 u}{d\theta^2} + u - \frac{\mu}{h^2} - \frac{m'}{h^2} \left\{ \frac{V}{u^2} - \frac{T}{u^3} \cdot \frac{du}{d\theta} - \frac{2}{h^2} \int \frac{T d\theta}{u^3} \right\}; \quad (239)$$

It is on this equation, (or the rigorous equation (237)) that the perturbation of the radius vector is made to depend in the lunar theory. It remains to derive an equation from which the perturbation in latitude can be obtained. For this purpose, all that is necessary will be to regard x and y as known, and to employ the third of our general equations (94) to obtain a value of z , or, which comes to the same, of s , since $z = \rho s$, and ρ is given by the previous theory.

Now, if in our equation (94) we change the independent variable from t to θ , it becomes

$$0 = \frac{d^2 z}{dt^2} - dz \cdot \frac{d^2 t}{dt^3} + \mu \cdot \frac{z}{r^3} + m' \cdot \frac{d\Omega}{dz}$$

but since

$$\left. \begin{aligned} \frac{1}{d t^2} &= \frac{h^2 + 2 m' \int T \rho^3 d\theta}{\rho^4 d\theta^2} \\ \frac{d^2 t}{dt^3} &= \frac{1}{d\theta} \left\{ \frac{2 d\rho}{\rho^5 d\theta} (h^2 + 2 m' \int T \rho^3 d\theta) - m' \cdot \frac{T}{\rho} \right\} \end{aligned} \right\}; \dots\dots\dots (240)$$

as appears from equations (233) and (234); if we substitute these, we shall get

$$\begin{aligned} 0 &= \left\{ h^2 + 2 m' \int T \rho^3 d\theta \right\} \left(\frac{1}{\rho} \cdot \frac{d^2 z}{d\theta^2} - \frac{2 d\rho}{\rho^2 d\theta} \cdot \frac{dz}{d\theta} \right) \\ &\quad + \rho^3 \left\{ \frac{\mu z}{r^3} + m' \cdot \frac{d\Omega}{dz} + m' \cdot \frac{T}{\rho} \cdot \frac{dz}{d\theta} \right\} \end{aligned}$$

Now we have $z = \rho s$, $dz = \rho ds + s d\rho$, $d^2 z = \rho d^2 s + 2 d\rho ds + s d^2 \rho$, whence we find

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$$\frac{1}{\rho} \cdot \frac{d^2 z}{d \theta^2} - \frac{2}{\rho^2} \frac{d \rho}{d \theta} \cdot \frac{d z}{d \theta} = \frac{d^2 s}{d \theta^2} + \frac{\rho^2 s}{\rho d \theta^2} \left\{ \frac{d^2 \rho}{\rho^2} - \frac{2 d \rho^2}{\rho^3} \right\}$$

But, by (235) we have

$$\begin{aligned} & (h^2 + 2 m' \int T \rho^3 d \theta) \times \frac{1}{\rho d \theta^2} \left\{ \frac{d^2 \rho}{\rho^2} - \frac{2 d \rho^2}{\rho^3} \right\} = \\ & = \frac{1}{\rho^2} (h^2 + 2 m' \int T \rho^3 d \theta) - \frac{\mu}{\rho} \cdot \frac{\rho^3}{r^3} - m' \left\{ \rho V + T \cdot \frac{d \rho}{d \theta} \right\} \end{aligned}$$

so that, by substitution we obtain,

$$\begin{aligned} 0 &= (h^2 + 2 m' \int T \rho^3 d \theta) \left(\frac{d^2 s}{d \theta^2} + s \right) \\ &+ m' \rho^3 \left\{ \frac{d \Omega}{d z} - s V \right\} + m' T \rho^2 \left\{ \frac{d}{d \theta} - \frac{s d \rho}{d \theta} \right\} \end{aligned}$$

but $d z = \rho d s + s d \rho$; so that $\frac{d z}{d \theta} - \frac{s d \rho}{d \theta} = \frac{\rho d s}{d \theta}$, and we get

$$0 = \left\{ h^2 + 2 m' \int T \rho^3 d \theta \right\} \left(\frac{d^2 s}{d \theta^2} + s \right) + m' \rho^2 \left\{ T \frac{d s}{d \theta} - s V + \frac{d \Omega}{d z} \right\}$$

this equation (writing u for $\frac{1}{\rho}$) becomes

$$0 = \frac{d^2 s}{d \theta^2} + s + m' \cdot \frac{T \frac{d s}{d \theta} - s V + \frac{d \Omega}{d z}}{u^3 \left\{ h^2 + 2 m' \int T \frac{d \theta}{u^3} \right\}}; \dots\dots\dots (250)$$

and when the square of the disturbing force is neglected,

$$0 = \frac{d^2 s}{d \theta^2} + s + \frac{m'}{h^2 u^3} \left\{ T \frac{d s}{d \theta} - s V + \frac{d \Omega}{d z} \right\} \quad (241)$$

Such are the fundamental equations on which the problem of the moon's motion depends. They were first investigated by Clairaut, in a piece which gained the prize of the Petersburg Academy for 1750, and have been deduced by almost every writer since his time, as the groundwork of the lunar theory. The method we have followed shews how the expressions of the tangential and centripetal disturbing forces peculiar to this theory, originate in the general equations of the problem of these bodies; and thus connects the modern with the more ancient methods followed by Newton, Clairaut, &c.

Hitherto we have only considered the rigorous equations of the moon's motion. It remains to apply to them the methods of approximation appropriate to the case, and deduce in succession the equations of the disturbed motion. The great length to which the details of this complicated process would lead, will preclude our entering minutely into it. We must content ourselves with leading results and general principles. One great peculiarity of the lunar theory consists in this—that the mass of the disturbing body, instead of being, as in that of the planets, small in comparison with the disturbed and central bodies, exceeds them both several hundred thousand times. Its vast distance, however, makes up for this, and renders its disturbing effect small in comparison with the attraction of the central body. In the foregoing equations then the quantities m , and m' , are not to be our guides in regulating the orders of the corrections. The very small quantities, whose powers determine the convergency of our approximations are T and V , or rather $m' T$, and $m' V$, for m' never occurs unmultiplied by one or other of these quantities. Let us therefore, first of all, consider the nature and magnitude of these forces.

Expression of the centripetal disturbing force, $m' V$.

$$\begin{aligned} m' V &= m' \left(\frac{x}{\rho} \frac{d \Omega}{d x} + \frac{y}{\rho} \frac{d \Omega}{d y} \right) = \frac{m'}{\rho} (x r' + y y') \left(\frac{1}{r^3} - \frac{1}{\lambda^3} \right) + \frac{m' \rho}{\lambda^3} \\ &= m' \rho' \cdot \cos (\theta - \theta') \left\{ \frac{1}{r^3} - \frac{1}{\lambda^3} \right\} + \frac{m'}{\lambda^3} \end{aligned}$$

Now, if we neglect the inclination of the moon's orbit, and take the plane of the ecliptic for our fixed plane, we have

$$z = 0, \quad z' = 0, \quad r = \rho, \quad r' = \rho', \quad s = 0,$$

and

$$\frac{1}{r^3} = \left\{ r'^2 - 2 r r' \cdot \cos (\theta - \theta') + r^2 \right\}^{-\frac{3}{2}}$$

$$\begin{aligned}
&= \frac{1}{r'^3} \left\{ 1 - \left(2 \frac{r}{r'} \cos (\theta - \theta') - \left(\frac{r}{r'} \right)^2 \right) \right\}^{-\frac{1}{2}} \\
&= \frac{1}{r'^3} \left\{ 1 - \left(2 \frac{r}{r'} \cos w - \left(\frac{r}{r'} \right)^2 \right) \right\}^{-\frac{1}{2}} \\
&= \frac{1}{r'^3} \left\{ 1 + \left(3 \frac{r}{r'} \cos w - \frac{3}{2} \left(\frac{r}{r'} \right)^2 \right) + \frac{3 \cdot 5}{2 \cdot 4} \left(2 \frac{r}{r'} \cos w - \left(\frac{r}{r'} \right)^2 \right)^2 + \&c. \right\} \\
&= \frac{1}{r'^3} \left\{ 1 + 3 \frac{r}{r'} \cos w + \frac{9 + 15 \cos 2w}{4} \cdot \left(\frac{r}{r'} \right)^2 + \&c. \right\}
\end{aligned}$$

putting $w = \theta - \theta'$, and developing in powers of $\frac{r}{r'}$. Hence, we obtain (writing r for ρ in the expression for V) after all reductions,

$$m' V = - \frac{m'}{r'^2} \left\{ \frac{1 + 3 \cos 2w}{2} \cdot \frac{r}{r'} + \frac{9 \cos w + 15 \cos 3w}{8} \left(\frac{r}{r'} \right)^2 + \&c. \right\} \quad (242)$$

Now $\frac{m'}{r'^2}$ is the accelerating force exerted by the sun on the earth, whose intensity, as we have shewn in Section 1. Part I., is represented by 2.17399 (in its mean state) on the supposition of the earth's attraction on the moon being unity. Also, $\frac{r}{r'} = \frac{60 \cdot 23799}{23405} = \frac{1}{388}$ nearly, and $\left(\frac{r}{r'} \right)^2 = \frac{1}{119920}$; so that we may safely neglect $\left(\frac{r}{r'} \right)^2$ &c. The constant, or non-periodical parts of $m' V$ is therefore equal to $-\frac{1}{2} \cdot \frac{m' r}{r'^2} = -\frac{1}{2} m' \left(\frac{r}{r'} \right)^3 \times \frac{1}{r^2} = \text{nearly} - \frac{1}{357 \cdot 5} \times \frac{1}{r^2}$.

Let a and a' be the actual mean distances of the moon and sun from the earth, and supposing

$$a = m' \left(\frac{a}{a'} \right)^3$$

a will be nearly $\frac{1}{178 \cdot 7}$, and the mean effect of the centripetal disturbing force will be $-\frac{\frac{1}{2} a}{r^2}$ indicating a diminution of the moon's mean gravity, amounting to $\frac{1}{357 \cdot 5}$ th of its whole quantity.

The moon's mean gravity being diminished by the action of the sun, and the velocity remaining nearly the same, the centrifugal force which would be balanced by the centripetal in a circular orbit without the sun's action, will obtain the preponderance when the sun acts and carry the moon farther out; thus the disturbing force has the effect of increasing the distance and periodic time above what they would be in the undisturbed orbit.

Let us next consider the tangential disturbing force $m' T$. Now we have

$$\begin{aligned}
m' T &= m' \left(\frac{x}{\rho} \frac{d \Omega}{d y} - \frac{y}{\rho} \frac{d \Omega}{d x} \right) \\
&= m' \cdot \frac{x y' - y x'}{\rho} \left(\frac{1}{r'^3} - \frac{1}{\lambda^3} \right) = - m' r' \cdot \sin w \left(\frac{1}{r'^3} - \frac{1}{\lambda^3} \right) \\
&= + \frac{m'}{r'^2} \left(3 \frac{r}{r'} \cos w + \left(\frac{r}{r'} \right)^2 \cdot \&c. \right) \cdot \sin w
\end{aligned}$$

that is

$$m' T = \frac{3}{2} \frac{m'}{r'^2} \cdot \frac{r}{r'} \cdot \sin 2w; \quad (243)$$

The tangential force therefore vanishes both in syzgies and quadratures, and is a maximum in octants, or at 45° of elongation. Now we have seen, that the description of areas would be equable, were it not for this tangential force. Hence, at the former points, the equable description of areas still holds good; while, at the latter, an acceleration or retardation takes place. This, of course, produces an equation in the moon's longitude, whose period is semimenstrual, as during the first quarter of its synodic revolution, the moon is continually retarded in its orbit (supposed circular): at the first quadrature the momentary retardation ceases and changes to an acceleration—and here the effect accumulated through the whole of the quadrant has reached its maximum. In the next quadrant, the moon is gradually accelerated; and at the full moon, its motion has regained all it had lost, and is restored to its mean state, so that here the equation is nothing, and so on. This is the origin of that equation of the moon's motion to which astronomers have attached the name of the variation.

Approximate integration of the equation of the moon's orbit.

IN this inquiry we will commence with the equation (238), in which the square of the disturbing force is neglected. Now, the obvious mode of beginning the approximation, would be (in analogy with what we did in the theory of the planets) to substitute for u its undisturbed, or elliptic value, $\frac{1 + e \cdot \cos \theta}{a(1 - e^2)}$. But, as it is obvious that the more nearly our first approximation approaches the truth, the more rapid will be the convergency of all the succeeding ones, we may take advantage of what observation has taught us respecting the form of the lunar orbit, and assume, for our first value of u , an expression representing that form, more nearly than the ellipse. Now, we learn from observation, that during a single revolution, it is true, the orbit does not deviate materially from an ellipse, but that in a very few revolutions the elliptic radius vector differs very sensibly from the real one, by reason of the rapid motion of the lunar apsides, which perform a whole circuit in little more than nine years. Instead then of representing by u the inverse radius vector of a *fixed* ellipse, we will give it such a form, in terms of θ , as shall express that of an ellipse in motion, and revolving in its own plane in the direction of the moon's motion, as observation informs us it does.

Let us then take c , such that $1 : 1 - c ::$ the moon's motion in longitude to the motion of the apsides, then will $\theta(1 - c)$ be the angle described by the apsis, while the moon describes θ ; and, consequently, supposing (as we may) that the origin of the time t is fixed at the epoch when the apsis coincided with the axis of the x , we shall have $\theta - (1 - c)\theta = c\theta$ for the moon's true anomaly, and in the place of taking

$$u = \frac{1 + e \cdot \cos \theta}{a(1 - e^2)}, \text{ we may take } u = \frac{1 + e \cdot \cos c\theta}{a(1 - e^2)} \text{ for our initial value of } u.$$

Although the idea of commencing the approximation with this value of u is, it is true, taken from observation, yet, when once suggested, no matter from what source, we may look on it as a mere analytical artifice; and the truth or falsehood, convenience or inconvenience of the assumption will be tried by actual substitution; when, if we find that it reproduces the original expression with a train of small corrections multiplied by the disturbing force, or can be made to do so by a proper assumption of the constant c , we may rest assured that the assumption is (mathematically speaking) a legitimate one, and concern ourselves no farther with the way in which we arrived at it.

Now, neglecting e^2 , we have

$$u = \frac{1}{a} (1 + e \cdot \cos c\theta)$$

whence we get

$$\frac{1}{u^3} = a^3 (1 - 3e \cdot \cos c\theta); \quad \frac{1}{u^4} = a^4 (1 - 4e \cdot \cos c\theta)$$

$$\frac{du}{d\theta} = -\frac{ce}{a} \cdot \sin c\theta$$

and, by substituting this in (238) which is equivalent to

$$0 = \frac{d^2 u}{d\theta^2} + u - \frac{1}{h^2} + m' \frac{u^3}{h^2} \cdot \frac{1 + 3 \cos 2w}{2} - \frac{3m' u^3}{2h^2 a} \cdot \frac{ce}{a} \sin c\theta \cdot \sin 2w \\ + \frac{3}{h^4} \int m' \frac{u^3}{u^4} \sin 2w \cdot d\theta$$

Now, if we neglect the eccentricity of the earth's orbit, $u' = \frac{1}{a}$, so that (recollecting that $m' \cdot \left(\frac{a}{a'}\right)^3 =$

$$a = \frac{1}{179})$$

$$0 = \frac{d^2 u}{d\theta^2} + u - \frac{1}{h^2} + \frac{a}{2h^2} (1 + 3 \cos 2w) (1 - 3e \cdot \cos c\theta) \\ - \frac{3ce}{2h^2} a \cdot \sin c\theta \cdot \sin 2w \\ + \frac{3a}{h^4} \cdot \int (1 - 4e \cos c\theta) \sin 2w \cdot d\theta \quad (244)$$

The only difficulty now in the way of integrating this, is to express $\cos 2w$ and $\sin 2w$ in terms of θ . Now we have $w = \theta - \theta'$; and since we neglect the eccentricity of the earth's orbit, θ' is proportional to the time or $\theta' = n't$; also we have θ by the equation, $h dt = r^2 d\theta$, or

$$t = \frac{1}{h} \int \frac{d\theta}{u^2} = \frac{1}{h} \int d\theta (1 - 2e \cdot \cos c\theta) a^2 = \frac{a^2}{h} \left(\theta - \frac{2e}{c} \cdot \sin c\theta \right)$$

Hence, since $h = \sqrt{a}$ very nearly, and the motion of the apsides being slow in comparison with that of the moon itself, c is nearly unity, at least sufficiently so for a first approximation; consequently,

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$$t = a^{\frac{1}{2}} (\theta - 2e \cdot \sin c \theta)$$

$$\text{or } n t = \theta - 2e \cdot \sin c \theta$$

and eliminating t between this and the equation $\theta' = n' t$, we have

$$\frac{n}{n'} \theta' = \theta - 2e \cdot \sin c \theta$$

and therefore,

$$\theta - \theta' = \theta \left(1 - \frac{n'}{n} \right) + 2e \cdot \frac{n'}{n} \cdot \sin c \theta$$

putting n and n' for the mean motions of the moon and earth. Hence we find, putting $\frac{n'}{n} = k$

$$\begin{aligned} \cos 2w &= \cos 2(1-k)\theta - \sin 2(1-k)\theta \cdot \sin \left\{ 4e \cdot \frac{n'}{n} \cdot \sin c \theta \right\} \\ &= \cos 2(1-k)\theta + 2ke \{ \cos(2-2k+c)\theta - \cos(2-2k-c)\theta \}; \end{aligned} \quad (245)$$

and similarly,

$$\sin 2w = \sin 2(1-k)\theta + 2ke \{ \sin(2-2k+c)\theta - \sin(2-2k-c)\theta \} \quad (246)$$

It only remains to substitute these values in (244) where it becomes, after reducing all the products of sines and cosines to simple multiples, and rejecting e^2 ,

$$o = \frac{d^2 u}{d\theta^2} + u - \frac{1}{h^2} \left(1 - \frac{a}{2} \right) - \frac{3e}{2h^2} a \cdot \cos c\theta \quad (247)$$

$$+ a \{ A \cdot \cos(2-2k)\theta + B e \cdot \cos(2-2k-c)\theta + C e \cos(2-2k+c)\theta \}$$

where A, B, C , are co-efficients easily expressed in functions of k and c . If we integrate this, we get

$$u = \frac{1}{h^2} \left(1 - \frac{a}{2} \right) - \frac{3e}{2h^2} \cdot \frac{a}{c^2 - 1} \cdot \cos c\theta \quad (248)$$

$$+ a \{ A' \cdot \cos(2-2k)\theta + B' e \cdot \cos(2-2k-c)\theta + C' \cdot e \cdot \cos(2-2k+c)\theta \}$$

Let us now consider the nature of this result. We set out with $u = \frac{1}{a} + \frac{e}{a} \cdot \cos c\theta$ for a first assumed value of u ; and having substituted this, have deduced another value, which ought to be a nearer approximation, and which ought to coincide with the former, if $a = 0$.

Now, the terms within the brackets being all multiplied by a , vanish as they should do when $a = 0$, and the two first terms of this expression will coincide with the uncorrected value of u , provided we take

$$\frac{1}{h^2} \left(1 - \frac{a}{2} \right) = \frac{1}{a} \quad \text{or } h = \sqrt{a} \cdot \sqrt{1 - \frac{a}{2}} = \sqrt{a} \left(1 - \frac{a}{4} \right)$$

and

$$\frac{3e}{2h^2} \cdot \frac{a}{1-c^2} = \frac{e}{a}, \quad \text{or } 1-c^2 = \frac{3}{2} a \cdot \frac{a}{h^2}$$

or

$$1-c^2 = \frac{3}{2} a \cdot \frac{a}{h^2} = \frac{3}{2} \frac{a}{1-\frac{a}{2}} = \frac{3}{2} a + \frac{3}{4} a^2$$

$$1-c = \frac{3}{4} a + \frac{21}{32} a^2, \quad \text{or } 1-c = \frac{3}{4} a$$

neglecting a^2 ; that is, since $a = \frac{1}{178.7}$, $c = 0.99580$, and $1-c = 0.0042$. Hence, according to this view of the subject, the progression of the apsides in one revolution of the moon is $0.0042 \times 360^\circ$, or about $1^\circ 31'$.

Now, it is remarkable that this progression of the lunar apsides, as determined by a first approximation, is only half the quantity actually observed; and this is the conclusion Newton had arrived at, where, after going through a process capable of being identified with that now under consideration, he remarks (in the 9th section of the *Principia*,) "*Apsis lunæ est duplo velocior circiter.*" Every attempt to obtain a nearer coincidence, by taking into account the higher powers of the eccentricities, the inclination of the moon's orbit (which is pretty considerable, and materially modifies some of its inequalities) &c, failed; and this difficulty, which Newton evidently felt, though he had passed it in this apparently negligent manner, as being at that time beyond his reach, or deferred for farther consideration, became so great a stumbling-block in the way of succeeding geometers, as to shake their faith in the theory of gravity, till Clairaut shewed that it would vanish on pushing the process to a second approximation, and that, in fact, the value of c depends, for nearly half its quantity, on a term containing the square of the disturbing force.

To shew this, we have only to substitute the approximate value of u (248) back into the original equation

$$o = \frac{d^2 u}{d\theta^2} + u - \frac{1}{h^2} + \left\{ \frac{m' T}{h^2 u^3} \frac{du}{d\theta} - \frac{m' V}{h^2 u^2} + 2 \left(\frac{d^2 u}{d\theta^2} + u \right) \int \frac{m' T}{h^2 u^3} d\theta \right\}$$

Astronomy. Supposing now we represent the first assumed value, or $\frac{1}{a} (1 + e \cdot \cos c \theta)$ by u and by δu all the terms **Physical Astronomy.** added by the first approximation, then will the value resulting from a second approximation be had by substituting $u + \delta u$ for u in all the part of the above equation within the brackets, and integrating. As we only require at present to know the terms introduced by this second approximation, let us call them $\delta^2 u$, and the value of $\delta^2 u$ will be had at once by integrating the equation

$$0 = \frac{d^2 \delta^2 u}{d \theta^2} + \delta^2 u + \frac{m'}{h^2} \left\{ \delta \cdot \frac{T}{u^3} \frac{d u}{d \theta} - \delta \frac{V}{u^2} + 2 \left(\frac{d^2 \delta u}{d \theta^2} + \delta u \right) \int \frac{T}{u^3} d \theta \right. \\ \left. + 2 \left(\frac{d^2 u}{d \theta^2} + u \right) \int \delta \cdot \frac{T}{u^3} d \theta \right\} ; \dots \dots (249)$$

Now, since $h^2 = a$ very nearly, and $a = m' \cdot \frac{a^3}{a'^3}$, we have

$$\frac{m'}{h^2} \delta \left(\frac{T}{u^3} \frac{d u}{d \theta} \right) = \frac{3}{2} a \delta \left\{ \frac{\sin 2 w}{(a u)^4} \frac{d u}{d \theta} \right\} \\ = \frac{3}{2} a \left\{ \frac{2 \cos 2 w}{(a u)^4} \frac{d u}{d \theta} \delta w + \frac{\sin 2 w}{(a u)^4} \cdot \frac{d \delta u}{d \theta} - 4 a \frac{\sin 2 w}{(a u)^5} \frac{d u}{d \theta} \delta u \right\} \\ - \frac{m'}{h^2} \delta \frac{V}{u^2} = - \frac{a}{2 a} \delta \frac{1 + 3 \cos 2 w}{(a u)^3} = \frac{3 a}{2 a} \left\{ \frac{2 \sin 2 w}{(a u)^3} \delta w + a \cdot \frac{1 + 3 \cos 2 w}{(a u)^4} \delta u \right\} \\ \frac{m' T}{h^2 u^3} = \frac{3 a}{2} \cdot \frac{\sin 2 w}{(a u)^4} \\ \int \delta \frac{m' T}{h^2 u^3} d \theta = 3 a \int \left\{ \frac{\cos 2 w}{(a u)^4} \delta w - 2 a \frac{\sin 2 w}{(a u)^5} \delta u \right\} d \theta$$

Let these values be substituted in the general equation (249), and it becomes

$$0 = \frac{d^2 \delta^2 u}{d \theta^2} + \delta^2 u + \frac{3}{2} a \left\{ \frac{2 \cos 2 w}{(a u)^4} \frac{d u}{d \theta} \delta w + \frac{\sin 2 w}{(a u)^4} \frac{d \delta u}{d \theta} - 4 a \frac{\sin 2 w}{(a u)^5} \frac{d u}{d \theta} \delta u \right\} \\ + \frac{3}{2} a \left\{ \frac{2 \sin 2 w}{a (a u)^3} \delta w + \frac{1 + 3 \cos 2 w}{(a u)^4} \delta u \right\} \\ + 3 a \left(\frac{d^2 \delta u}{d \theta^2} + \delta u \right) \int \frac{\sin 2 w}{(a u)^4} d \theta \\ + 6 a \left(\frac{d^2 u}{d \theta^2} + u \right) \int \left\{ \frac{\cos 2 w}{(a u)^4} \delta w - 2 a \frac{\sin 2 w}{(a u)^5} \delta u \right\} d \theta \quad (250)$$

To reduce this equation to the degree of approximation required, let us agree to neglect all terms which contain a^2 multiplied by any power of the eccentricity; or in that part of the approximation which depends on the square of the disturbing force, to regard the orbits as circular. We shall have then $u = \frac{1}{a} a u = 1$, and

supposing $\delta u = \frac{a}{a} \cdot \phi$, where ϕ is the sum of any number of terms of the form $A \cdot \cos i \theta$, we have

$$\delta t = \frac{1}{h} \int \delta \frac{d \theta}{u^2} = - \frac{1}{h} \int \frac{2 \delta u}{u^3} d \theta = - \frac{2 a^2}{h} a \int \phi d \theta$$

or since $n = \frac{1}{a^{\frac{1}{2}}} = \frac{h}{a^2}$

$$\delta \cdot n t = - 2 a \int \phi d \theta$$

consequently, since $w = \theta - \theta = \theta - \frac{n'}{n} t = \theta - k \cdot n t$, we get

$$w = \theta (1 - k), \text{ and } \delta w = - k \cdot \delta n t = 2 k a \int \phi d \theta$$

We have also $\frac{d u}{d \theta} = 0$ and $\frac{1}{(a u)^n} = 1$; so that the equation (250) reduces itself to

$$0 = \frac{d^2 \delta^2 u}{d \theta^2} + \delta^2 u + \frac{3 a^2}{2 a} \sin 2 w \cdot \frac{d \phi}{d \theta} + \frac{6 k a^2}{a} \sin 2 w \cdot \int \phi d \theta + \frac{3 a^2}{2 a} (1 + 3 \cos 2 w) \cdot \phi \\ + \frac{3 a^2}{a} \left(\frac{d^2 \phi}{d \theta^2} + \phi \right) \int d \theta \cdot \sin 2 w \\ + \frac{12 a^2}{a} \int \left\{ k \cos 2 w \int \phi d \theta - \sin 2 w \cdot \phi \right\} d \theta \quad (251)$$

Astronomy Let us now, more particularly, consider any one of the terms to which ϕ consists; as, for instance, $A \cdot \cos B \theta$. This gives Physical Astronomy.

$$\phi = A \cdot \cos B \theta, \quad \frac{d\phi}{d\theta} = -AB \cdot \sin B \theta, \quad \frac{d^2\phi}{d\theta^2} = -AB^2 \cdot \cos B \theta; \quad \int \phi d\theta = \frac{A}{B} \cdot \sin B \theta, \text{ \&c.}$$

and the equation will give, on substitution and reduction,

$$\begin{aligned} 0 = \frac{d^2 \delta^2 u}{d\theta^2} + \delta^2 u + \left\{ \frac{6ka^2}{a} \cdot \frac{A}{B} - \frac{3a^2}{2a} \cdot AB \right\} \sin 2w \cdot \sin B \theta \\ + \left\{ \frac{3a^2}{2a} (1 + 3 \cdot \cos 2w) + (B^2 - 1) \cdot \frac{3a^2}{a} \cdot \frac{\cos 2w}{2(1-k)} \right\} A \cdot \cos B \theta \\ - \frac{6Aa^2}{a} \left\{ \frac{k-B}{B(B+2-2k)} \cos(B+2-2k)\theta + \frac{k+B}{B(B-2+2k)} \cos(B-2+2k)\theta \right\} \end{aligned}$$

or resolving the products of sines and cosines by the formula (A), page 690,

$$\begin{aligned} 0 = \frac{d^2 \delta^2 u}{d\theta^2} + \delta^2 u + \frac{3Aa^2}{2a} \left(\frac{2k}{B} - \frac{B}{2} \right) \{ \cos(B-2+2k)\theta - \cos(B+2-2k)\theta \} \\ + \frac{3Aa^2}{2a} \cdot \cos B \theta \\ + \frac{3Aa^2}{4a} \left\{ 3 + \frac{B^2-1}{1-k} \right\} \{ \cos(B-2+2k)\theta + \cos(B+2-2k)\theta \} \\ - \frac{6Aa^2}{a} \left\{ \frac{k+B}{B(B+2-2k)} \cdot \cos(B+2-2k)\theta + \frac{k+B}{B(B-2+2k)} \cos(B-2+2k)\theta \right\} \end{aligned}$$

That is, assembling together the co-efficients of like terms,

$$\begin{aligned} 0 = \frac{d^2 \delta^2 u}{d\theta^2} + \delta^2 u + \frac{3Aa^2}{a} \left\{ \frac{k}{B} - \frac{B}{4} + \frac{3}{4} + \frac{B^2-1}{4(1-k)} - \frac{2(k+B)}{B(B-2+2k)} \right\} \cos(B-2+2k)\theta \\ + \frac{3Aa^2}{a} \left\{ -\frac{k}{B} + \frac{B}{4} + \frac{3}{4} + \frac{B^2-1}{4(1-k)} - \frac{2(k-B)}{B(B+2-2k)} \right\} \cos(B+2-2k)\theta \\ + \frac{3Aa^2}{2a} \cdot \cos B \theta \end{aligned}$$

Which equation, integrated, gives

$$\begin{aligned} \delta^2 u = \frac{3Aa^2}{2a} \cdot \left\{ \frac{\cos B \theta}{B^2-1} + \left(\frac{2k}{B} + \frac{3-B}{2} + \frac{B^2-1}{2(1-k)} - \frac{4(k+B)}{B(B-2+2k)} \right) \frac{\cos(B-2+2k)\theta}{(B-2+2k)^2-1} \right. \\ \left. + \left(-\frac{2k}{B} + \frac{3+B}{2} + \frac{B^2-1}{2(1-k)} - \frac{4(k-B)}{B(B+2-2k)} \right) \frac{\cos(B+2-2k)\theta}{(B+2-2k)^2-1} \right\} \quad (252) \end{aligned}$$

Such is the value of that part of $\delta^2 u$, which originates in the term $A \cdot \cos B \theta$ in the value of ϕ or in the term $\frac{a}{a} \cdot A \cos B \theta$ in u itself. A similar set of terms, having for their arguments respectively,

$$B' \theta, \quad (B' - 2 + 2k) \theta, \quad (B' + 2 - 2k) \theta$$

and

$$B'' \theta, \quad (B'' - 2 + 2k) \theta, \quad (B'' + 2 - 2k) \theta; \text{ \&c.}$$

will arise from the other terms $\frac{a}{a} A' \cdot \cos B' \theta$, $\frac{a}{a} A'' \cdot \cos B'' \theta$, &c. Consequently, by substituting

for $B, B', B'', \text{ \&c.}$ the several values which they actually have in the first approximate value of u (equation 248), or $2-2k, 2-2k-c, 2-2k+c$, we shall obtain the arguments of the several terms in the second approximation, and so on. These are then as follows:—

From $c \theta$ arise the arguments $c \theta, 2c \theta$

From $(2-2k) \theta$ arise the arguments $(2-2k) \theta; c \theta; (4-4k) \theta$

From $(2-2k-c) \theta$ arise $(2-2k-c) \theta; c \theta; (4-4k-c) \theta$

From $(2-2k+c) \theta$ arise $(2-2k+c) \theta; c \theta; (4-4k+c) \theta$

Thus we see, 1st. that each term which in the expression of u is multiplied by the first power of the disturbing force a , reproduces itself on a second approximation, multiplied by the square of a , and thus the co-efficient of every term in the general expression of u , consists in fact of an infinite series of powers of a ; for what takes place at the second approximation will, of course, do so at every succeeding step; and, in fact, the very same analysis, and the same resulting formula, may be applied to deduce the terms depending on $a^3, a^4, \text{ \&c.}$, in all of which the same arguments will occur.

Astronomy. Hence, it follows, that if we set out, as a first hypothesis, with $u = \frac{1}{a} (1 + e \cdot \cos c \theta)$ the process of Physical Astronomy. approximation leads us to a value of u , of the form

$$\begin{aligned} u = & \frac{1}{a} \{ A + B a + C a^2 + D a^3 + \&c. \} - \frac{1}{h^2} \\ & + \cos c \theta \{ A' a + B' a^2 + C' a^3 + \&c. \} \\ & + \cos (2 - 2k) \theta \cdot \{ A'' a + B'' a^2 + \&c. \} \\ & + \cos (4 - 4k) \theta \cdot \{ B''' a^2 + \&c. \} \\ & + \cos (2 - 2k + c) \{ A^{iv} a + B^{iv} a^2 + \&c. \} \\ & + \cos (2 - 2k - c) \{ A^v a + B^v a^2 + \&c. \} \\ & + \&c. \end{aligned}$$

We have therefore only to take care that this shall not contradict the original hypothesis, and therefore we must have

$$1 = A + B a + C a^2 + \&c. \cdot \frac{a}{h^2} \quad (253)$$

$$\frac{e}{a} = A' a + B' a^2 + C' a^3 + \&c. \quad (254)$$

These equations established, the two first terms of u agree with the original value, and the remaining ones are the equations due to the effect of perturbation, not capable of being expressed by the hypothesis of a revolving ellipse.

The equation (253) expresses a relation between a and h , which we shall consider presently; the other (254) gives the value of $1 - c$, or the velocity of the motion of the lunar apogee. We have already seen

(248) that the value of A' is $\frac{3e}{2h^2(1-c^2)}$, and if we break off the equation (254) at the first power of a we

have also seen that the resulting value of $1 - c$ is $\frac{3}{2}a$, amounting to only half the value given by observation.

If we make B successively equal to $2 - 2k - c$ and $2 - 2k + c$ in (252), we find for the co-efficients of the argument $c\theta$, respectively,

$$\frac{1}{c^2 - 1} \cdot \frac{3a^2e}{2a} B' \left\{ \frac{2 + 6k + c - 4k^2 - 4kc - c^2}{2(2 - 2k - c)} + \frac{(2 - 2k - c)^2 - 1}{2(1 - k)} + \frac{4(2 - k - c)}{c(2 - 2k - c)} \right\}$$

and

$$\frac{1}{c^2 - 1} \cdot \frac{3a^2e}{2a} C' \left\{ \frac{2 + 6k - c - 4k^2 + 4kc - c^2}{2(2 - 2k + c)} + \frac{(2 - 2k + c)^2 - 1}{2(1 - k)} - \frac{4(2 - k + c)}{c(2 - 2k + c)} \right\}$$

when B' and C' are the co-efficients of $e \cdot \cos (2 - 2k - c) \theta$ and $e \cdot \cos (2 - 2k + c) \theta$ within the parentheses of equation (248).

Thus our equation (254) carried as far as a^2 , will become

$$1 - c^2 = \frac{3}{2}a + \frac{3}{2}a^2 \cdot \left\{ \begin{aligned} & B' \left(\frac{2 - 6k + c - 4k^2 - 4kc - c^2}{2(2 - 2k - c)} + \frac{(2 - 2k - c)^2 - 1}{2(1 - k)} + \frac{4(2 - k - c)}{c(2 - 2k - c)} \right) \\ & + C' \left(\frac{2 + 6k - c - 4k^2 + 4kc - c^2}{2(2 - 2k + c)} + \frac{(2 - 2k + c)^2 - 1}{2(1 - k)} - \frac{4(2 - k + c)}{c(2 - 2k + c)} \right) \end{aligned} \right\} \quad (25b)$$

This equation is to be resolved by approximation, so as to find c ; and as a is a very small quantity, the approximation may be performed easily; for we have only to rescind a^2 , and take $c = 1 - \frac{3}{4}a$ for a first

approximation = 0.9957, and then substitute this for c in the co-efficient of a^2 , which will then become a given number (B' , C' , being previously computed, and k being known,) after which, if this co-efficient be called β , we have

$$1 - c^2 = \frac{3}{2}(a + \beta a^2) \quad c = \sqrt{1 - \frac{3}{2}(a + \beta a^2)} = 1 - \frac{3}{4}a - \frac{6\beta + 1}{8}a^2$$

and, again substituting this for c , we find a nearer approximation, and so on.

Now, it is a remarkable circumstance, that when this process is executed in numbers, the term $-\frac{6\beta + 1}{8}a^2$ thus added to the value of c is found to be very nearly equal to the term $-\frac{3}{4}a$ introduced by

the first step of the approximation; and on pushing it still farther, the subsequent corrections are found to be inconsiderable. Now the velocity of the apogee of the lunar orbit is expressed by $1 - c$, or by

$\frac{3}{4}a + \frac{6\beta + 1}{8}a^2 + \&c.$ We have already seen that the value of this, as deduced from a first approximation

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(or the part $\frac{3}{4}a$) came out only half what observation has assigned as its real amount, and that this was regarded as a great difficulty in the way of the full admission of the Newtonian law of gravity. We now see the reason of this; the remaining half is accounted for by the term $\frac{6\beta+1}{8}a^2$ arising from a second approximation; and the fact, so far from being an objection against gravity, is thus converted into a most cogent argument in its favour—its effects being thus shewn to correspond not merely to general results and first approximations, but to the refinements and niceties of theory.

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SECTION III.

Expression of the moon's mean longitude in terms of the true, and vice versa; and of the variation and evection of the moon.

THE purposes of astronomy require us to know the moon's true longitude and latitude at any assigned instant, or to know the angle θ in terms of the time t . For this we must have recourse to the equation.

$$t = \int \frac{\rho^2 d\theta}{\sqrt{h^2 + 2m' \int T \rho^3 d\theta}} = \int \frac{d\theta}{u^2 \sqrt{h^2 + 2m' \int \frac{T}{u^3} d\theta}}$$

If we developpe this, and retain only the first power of the disturbing force, we get

$$t = \frac{1}{h} \int \frac{d\theta}{u^2} \left\{ 1 - \frac{m'}{h^2} \int \frac{T}{u^3} d\theta \right\} = \frac{1}{h} \int \frac{d\theta}{u^2} \left\{ 1 - \frac{m'}{a} \int \frac{T}{u^3} d\theta \right\}$$

because $h^2 = a$, if the disturbing force and square of the eccentricity be neglected. Now, in this for u let us substitute

$$u = \frac{1}{a} \cdot (1 + e \cdot \cos c\theta) + a \{ A' \cdot \cos (2 - 2k)\theta + B'e \cdot \cos (2 - 2k - c)\theta + C'e \cdot \cos (2 - 2k + c)\theta \}$$

$$\text{and for } \frac{m'}{h^2} \int \frac{T}{u^3} d\theta \text{ its value } \frac{3a}{2} \int \frac{\sin 2w}{(au)^4} d\theta$$

$$= \frac{3a}{2} \int \sin 2w (1 - 4e \cdot \cos c\theta) d\theta$$

$$= \frac{3a}{2} \int \{ \sin (2 - 2k)\theta - (2 - 2k)e \sin (2 - 2k + c)\theta - (2 + 2k)e \sin (2 - 2k - c)\theta \} d\theta$$

$$= \frac{3a}{2} \left\{ \frac{\cos (2 - 2k)\theta}{2(k-1)} - \frac{2(k-1)e}{2-2k+c} \cos (2 - 2k + c)\theta + \frac{2(k+1)e}{2-2k-c} \cos (2 - 2k - c)\theta \right\}$$

whence we find

$$\frac{ht}{a^2} = \int d\theta \left\{ \begin{aligned} &1 - 2e \cdot \cos c\theta - 2aa \{ A' \cdot \cos (2 - 2k)\theta + B'e \cdot \cos (2 - 2k - c)\theta + C'e \cdot \cos (2 - 2k + c)\theta \} \\ &+ 3A'a \cdot a e \{ \cos (2 - 2k + c)\theta + \cos (2 - 2k - c)\theta \} \\ &- \frac{3a}{2} (1 - 2e \cos c\theta) \left(\frac{\cos 2(1-k)\theta}{2(k-1)} - \frac{2(k-1)e}{2-2k+c} \cos (2 - 2k + c)\theta + \frac{2(k+1)e}{2-2k-c} \cos (2 - 2k - c)\theta \right) \end{aligned} \right\} \quad (256)$$

$$\text{or putting } \frac{h^2}{a^2} = n$$

$$nt = \theta - \frac{2e}{c} \cdot \sin c\theta + Pa \cdot \sin (2 - 2k)\theta + Qae \cdot \sin (2 - 2k + c)\theta + Rae \cdot \sin (2 - 2k - c)\theta \} \dots (257)$$

A', B', C' , being as in (248)

where

$$\left. \begin{aligned} P &= -\frac{aA'}{1-k} + \frac{3}{8(1-k)^2} \\ Q &= \frac{a(A' - 2C')}{2-2k+c} + \frac{3}{4(k-1)(2-2k+c)} + \frac{3(k-1)}{(2-2k+c)^2} \\ R &= \frac{a(A' - 2B')}{2-2k-c} + \frac{3}{4(k-1)(2-2k-c)} - \frac{3(k+1)}{(2-2k-c)^2} \end{aligned} \right\} \dots (258)$$

In order to have the value of θ in terms of nt , we must revert this series, and the simplest mode of accomplishing this will be to follow the steps by which we resolved the equation $nt = v - e \cdot \sin v$ in the elliptic theory (vide page 656, col. 1, equation 26). Putting it therefore into the form

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$$\begin{aligned}\theta = nt + \frac{2e}{c} \cdot \sin c\theta - Pa \sin (2 - 2k)\theta \\ - Qae \cdot \sin (2 - 2k + c)\theta \\ - Rae \cdot \sin (2 - 2k - c)\theta\end{aligned}$$

nt appears to be the first approximate value of θ , neglecting e and a . Let this be substituted in the whole of the second member, and we get

$$\begin{aligned}\theta = nt + \frac{2e}{c} \cdot \sin cnt - Pa \cdot \sin (2 - 2k)nt \\ - Qae \cdot \sin (2 - 2k + c)nt \\ - Rae \cdot \sin (2 - 2k - c)nt\end{aligned}\quad (260)$$

for a second approximation carried as far as the first powers of e and a . If we now substitute this value in (259) the terms multiplied by e and a only will of course remain; but in addition to those multiplied by ae , others will be introduced; and if we reject e^3 and a^2 , and retain only ae , these terms will be, (putting $c = 1$)

$$\begin{aligned}- 2Pae \cdot \sin (2 - 2k)nt \cdot \cos cnt \\ - 2Pae (1 - k) \cdot \cos (2 - 2k)nt \cdot \sin cnt\end{aligned}$$

whose sum, after the usual reductions, is

$$- P(2 - k)ae \cdot \sin (2 - 2k + c)nt - Pkae \cdot \sin (2 - 2k - c)nt$$

Consequently, if we push the approximation only to such terms, we get

$$\begin{aligned}\theta = nt + \frac{2e}{c} \cdot \sin cnt - Pa \cdot \sin (2 - 2k)nt \\ - \{Q + (2 - k)P\}ae \cdot \sin (2 - 2k + c)nt \\ - \{R + kP\}ae \cdot \sin (2 - 2k - c)nt\end{aligned}\quad (261)$$

The variation of the moon is that inequality which is represented by the term $- Pa \cdot \sin (2 - 2k)nt$. The numerical magnitude of the co-efficient being considerable, (2146'') it, (as well as the inequality represented by the term, whose argument is $(2 - 2k - c)nt$, which is called the *evection*, and whose co-efficient is still greater (4830'') was long observed before its cause was known. If we observe that nt being the mean longitude of the moon knt will be that of the sun, and if we put \textcircled{D} for the former, and \textcircled{C} for the latter, the argument of the variation will be $2\textcircled{D} - 2\textcircled{C}$, or simply $\textcircled{D} - \textcircled{C}$, and its *period*, half a synodic revolution of the moon = $14^d.765$. The argument of the evection is in like manner

$$2\textcircled{D} - 2\textcircled{C} - c$$

where c is the longitude of the moon's perigee, and its period is nearly that of the variation, a little longer, however, on account of the progression of the perigee during a revolution of the moon. The term depending on the argument $(2 - 2k + c)nt$, or $2\textcircled{D} - 2\textcircled{C} + c$, is found on calculation to have its co-efficient very small, and has no particular name.

In the theory we have now investigated we have been only anxious to simplify and curtail as much as possible the developements, it being far beyond our design to enter into minutiae on so complicated a problem. It must suffice to have shewn in rather more than a general way the origin of the principal inequalities, and traced them from the differential equations to their final expression in the value of the longitude. In so doing, we have purposely neglected the inclination of the moon's orbit, and its effect both in modifying the disturbing forces which act in the plane of the orbit, and in altering the longitudes by reducing them from their actual values to their projections on the plane of the ecliptic, or by what is called in astronomy, the "Reduction to the ecliptic." The effect of the inclination, however, will be briefly touched on in the next section.

SECTION IV

Of the effect of the inclination of the plane of the moon's orbit.—Of the motion of the nodes, and the precession of the equinoxes.

In determining the effect of the inclination of the lunar orbit to the plane of the ecliptic and the motion of its nodes, we shall suppose the orbit circular, and neglect the square of the disturbing force, the cube and higher power of the inclination, and the product of the disturbing force and inclination. Thus our equation (237) will become simply,

$$0 = \frac{d^2 u}{d\theta^2} + u - \frac{1}{h^2(1 + s^2)^{\frac{3}{2}}} - \frac{m'}{h^2} \left\{ \frac{V}{u^2} - \frac{2}{h^2} \int \frac{T d\theta}{u^3} \right\}; \quad (262)$$

The value of the part depending on the disturbing forces $m'V$ and $m'T$ in this equation remains unaltered, if we neglect as^2 ; and therefore the part of u arising from perturbation is not altered when we only push the approximation to the extent now proposed. The only new terms arising in the value of u from the introduction

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$$\frac{d^2 u}{d \theta^2} + u - \frac{1}{h^2} \left(1 - \frac{3}{2} s^2 \right) = 0$$

neglecting s^4 . Now, in this we must put for s its undisturbed value, which is given by the equation

$$\frac{d^2 s}{d \theta^2} + s = 0 \text{ (making } m' = 0, \text{ in equation 240) and which is}$$

$$s = \gamma \cdot \sin (\theta - N) \quad (263)$$

where γ is the tangent of the inclination, and N the longitude of the ascending node. This gives

$$s^2 = \gamma^2 \cdot \frac{1 - \cos 2 (\theta - N)}{2}$$

$$\frac{d^2 u}{d \theta^2} + u - \frac{1}{h^2} \left(1 - \frac{3 \gamma^2}{4} \right) - \frac{3 \gamma^2}{4} \cdot \cos 2 (\theta - N) = 0$$

which integrated, gives

$$u = \frac{1}{h^2} - \frac{3 \gamma^2}{4 h^2} - \frac{\gamma^2}{4 h^2} \cos 2 (\theta - N) \quad (264)$$

Thus we see that the effect of the inclination is to add to

$$u = \frac{1}{h^2} \text{ or } u = \frac{1}{a} \text{ the terms } - \frac{\gamma^2}{4 h^2} (3 + \cos 2 (\theta - N))$$

Now, in approximating to u , we assumed in the foregoing pages an expression for a first value derived not *a priori* from theory, but from the *observed* motion of the lunar apsis. It is equally a matter of observation, that the node of the lunar orbit does not remain fixed, but is continually retreating on the ecliptic, in a direction contrary to the motion of the moon in its orbit. Let us then introduce this modification into our expression for u above found, and taking for the origin of the time, the moment when the node coincided with the axis of the x , we have only to write $g \theta$ in place of $\theta - N$, when we get

$$u = \frac{1}{h^2} \left\{ 1 - \frac{3}{4} \gamma^2 - \frac{\gamma^2}{4} \cdot \cos 2 g \theta \right\} \quad (265)$$

Let us next investigate the effect of the inclination on that part of u which arises from perturbation. Now this may be done at once, by merely taking in 250, 251,

$$\delta u = \frac{a}{r} \phi = - \frac{1}{a} \left(\frac{3}{4} \gamma^2 + \frac{1}{4} \gamma^2 \cos 2 g \theta \right)$$

or

$$\phi = - \frac{3}{4} \frac{\gamma^2}{a} \cdot \cos 0 \cdot \theta - \frac{1}{4} \frac{\gamma^2}{a} \cos 2 g \theta$$

that is, supposing in 252, 1st. $A = - \frac{3 \gamma^2}{4 a}$, $B = 0$, and 2dly, $A = - \frac{\gamma^2}{4 a}$ and $B = 2 g$

The first supposition gives

$$\delta^2 u = \frac{9}{8 a} a \gamma^2 \left\{ 1 - \left(3 - \frac{1}{1 - k} \right) \frac{\cos (2 - 2 k) \theta}{(2 - 2 k)^2 - 1} \right\} \quad (266)$$

and the second in like manner introduces into the value of u the new arguments $\cos (2 g - 2 + 2 k) \theta$ and $\cos (2 g + 2 - 2 k) \theta$ with co-efficients affected with $a \gamma^2$.

These arguments existing in the expression of u will in like manner be introduced into the value of t or $\int \frac{d \theta}{h u^2}$ and thence, by the process of reversion before explained, will arise terms of the form $a \gamma^2 \cdot \sin (2 g - 2 + 2 k) n t$ and $a \gamma^2 \cdot \sin (2 g + 2 - 2 k) n t$ in the longitude, of which this mention must suffice, as their co-efficients may easily be calculated from the principles above laid down, and nothing very remarkable in the lunar theory depends on them. Other terms also will arise from the part of (250) multiplied by $\frac{d u}{d \theta}$ which in this case is not $= 0$ as was supposed in (251) but is equal to $\frac{g \gamma^2}{2 h^2} \sin 2 g \theta$.

Let us now examine the manner in which the disturbing forces affect the moon's latitude. For this purpose we must take the equation (241) in which writing for s the expression $\gamma \cdot \sin (g v - N)$, for $\frac{d s}{d \theta}$ its value

$$g \gamma \cdot \cos (g v - N) \text{ for } m' T \text{ and } m' V \text{ their values } \frac{3}{2} \frac{a}{a^2} \sin 2 (1 - k) \theta \text{ and } - \frac{1}{2} \frac{a}{a^2} (1 + 3 \cos 2 (1 - k) \theta)$$

Astronomy. and for $m' \frac{d\Omega}{dz}$ its value $\frac{m'z}{\lambda^3} = \frac{m'\rho s}{\lambda^3} = \frac{m'as}{\lambda^3} = \frac{m'as}{a^3} \left(1 + 3 \frac{a}{a'} \cos w\right)$ (see page 718, line 4,) Physical Astronomy

or, neglecting the product $s \times \frac{a}{a'}$, simply

$$m' \frac{d\Omega}{dz} = \frac{m'as}{a^3} = \frac{a}{a'^2} s$$

we shall find

$$o = \frac{d^2 s}{d\theta^2} + s + \frac{3}{2} a \gamma \cdot g \sin 2(1-k) \cos \theta (g\theta - N) \\ + \frac{a}{2} s (1 + 3 \cos 2(1-k)\theta) + as$$

and resolving the products of sines and cosines into simple sines,

$$o = \frac{d^2 s}{d\theta^2} + s + \frac{3a}{4} a \gamma \cdot \{ \sin(\overline{2-2k+g} \cdot \theta - N) + \sin(\overline{2-2k-g} \cdot \theta + N) \} \\ + \frac{3}{2} a \cdot \gamma \sin(g\theta - N) + \frac{3}{4} a \gamma \{ \sin(\overline{2-2k+g} \cdot \theta - N) - \sin(\overline{2-2k-g} \cdot \theta + N) \}$$

or

$$o = \frac{d^2 s}{d\theta^2} + s + \frac{3}{2} a \gamma \cdot \sin(g\theta - N) \quad (267) \\ + \frac{3}{4} a \gamma (1+g) \cdot \sin(\overline{2-2k+g} \cdot \theta - N) \\ - \frac{3}{4} a \gamma (1-g) \cdot \sin(\overline{2-2k-g} \cdot \theta + N)$$

The integration of this gives

$$s = \frac{3a\gamma}{2(g^2-1)} \sin(g\theta - N) + a\gamma \sin(\overline{2-2k+g} \cdot \theta - N) \times \&c. \\ + a\gamma \sin(\overline{2-2k-g} \cdot \theta + N) \times \&c.$$

Now the process of approximation ought necessarily to reproduce the original value $s = \gamma \cdot \sin(g\theta - N)$. Hence, exactly as in the case of the apogee, we ought to have

$$\frac{3a\gamma}{2(g^2-1)} \cdot \sin(g\theta - N) = \gamma \cdot \sin(g\theta - N)$$

and

$$g^2 = 1 + \frac{3a}{2}, \quad g = 1 + \frac{3}{4} a$$

Thus we obtain the motion of the node, for $g-1$ or $\frac{3}{4}a$ expresses the ratio of the retrograde velocity of the

node to that of the moon in its orbit. If we execute the calculation, we find this ratio to be that of 0.0042 : 1 (see page 720, line 15 from bottom,) and it is remarkable, that it is exactly the same as a first approximation gives for the direct motion of the apogee; but there is one remarkable difference, that in the present case, this first approximated value is very near the truth as given by observation, and is little altered by a second approximation; whereas, in the other, the repetition of the process doubles the value.

The regression of the moon's nodes is a circumstance in their theory which admits of a very easy and familiar illustration.

The part of the disturbing force which acts in the direction of the z , or which tends to draw the moon out of the plane of its orbit is $m' \frac{d\Omega}{dz} = m' \cdot \left(\frac{z'}{r^3} - \frac{z' - z}{\lambda^3} \right)$. If we wish to know the whole force which tends to draw the moon in a direction at right angles to the plane of its own orbit, we must assume that plane for the plane of the x, y ; then will $z = o$, and

$$m' \frac{d\Omega}{dz} = m' z' \left(\frac{1}{r^3} - \frac{1}{\lambda^3} \right)$$

Now $m' \left(\frac{1}{r^3} - \frac{1}{\lambda^3} \right) = -\frac{3m'r}{r^4} \cos w = \text{nearly} -\frac{3a}{a^3 a'} \cos w$, and z' being a perpendicular let fall from the sun on the plane of the moon's orbit is equal to r' or a' multiplied by the sine of the sun's angular distance from the node ($\sin(\theta' - N)$) and by that of the inclination or by γ . Hence, the expression of this force is

$$-\frac{3a}{a^2} \times \cos w \cdot \sin(\theta' - N) \cdot \sin I$$

w is the moon's angular distance from the sun as seen from the earth, and therefore $\cos w$ is equal to sine of

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$$- \frac{3a}{a^2} \times \sin \text{☾'s dist. from quadrat.} \times \sin \text{☉'s dist. from ☿} \times \sin \text{inclin.};$$

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for the value of the force in question. Let us now examine how this force tends to produce a motion in the node. To this end, suppose the moon to set off from its ascending node, and at any point in its orbit (supposed circular) to describe an infinitely small arc which will be a portion of a great circle seen from the earth's centre. If the disturbing force at this moment ceased to act, this arc would be a portion of the undisturbed orbit, and being prolonged backwards would cut the ecliptic in the ascending node, ☿, whose position therefore would be the same at the beginning and end of this infinitesimal instant. To limit our ideas, let us conceive the moon to be within 90° from the ☿, and therefore its motion is from the plane of the ecliptic. Now, let the disturbing force act, and suppose its direction to be such as to draw the moon out of its orbit towards the plane of the ecliptic, then will the elementary arc actually described by the moon in virtue of its own inertia combined with the new impulse given by this force, be *less inclined* to the ecliptic than the last described portion; and therefore being produced backwards, will cut the ecliptic in a point *behind* the former place of the node. This point is the new or consecutive place of the node, which therefore has retreated on the plane of the ecliptic by the action of the disturbing planet.

On the other hand, had the disturbing force been directed from the plane of the ecliptic, the new path of the moon would be more inclined than in the preceding instant; and therefore, produced backwards, would cut the ecliptic in a point more *forward* than the previous place of the node, so that under these circumstances the node advances.

Thus we see that when the moon moves from the ecliptic and the force acts *to* it, the node retreats; but when the force acts *from* it, the node advances.

Again, let the moon be approaching the ecliptic; then, if the force act *to* that plane, it will approach it more rapidly, and will cut it in a point nearer than the node to which it is approaching. This node (and of course both) will then move to meet the moon, or in a direction contrary to its motion, *i.e.* will retreat; and *vice versa*, if the force act from the plane, the node advances.

From this analysis of all the cases, we see that whenever the disturbing force tends to elevate the moon from the plane of the ecliptic, the node advances, and in every other case retreats. Now, it is easily seen, that the former condition never holds good unless the moon is between the node and the quadratures; and as the extent of the angle in which this can happen during a whole revolution of the moon in its orbit, is necessarily less than two right angles, the preponderant tendency of the node on the average of a whole revolution is always in favour of its retreat. In fact, when the node is in quadratures, it retreats at every instant of the lunation; and in the most unfavourable case, when the node is in syzgies, its retreat is barely counterbalanced by its advance, and the node only rests for an instant; the sun being then for a moment in the plane of the lunar orbit.

The general tendency of the node to recede on the ecliptic is thus clearly made out; but we may go further on these principles, and make the quantity of its recess a matter of calculation. For, let us denote by $\frac{K}{a^2}$ the force expressed in (268). Then, since the lunar gravity $\left(\frac{1}{a^2}\right)$ draws the moon in the instant of time dt through the versed sine of an arc $= a d\theta$ or through a space equal to $\frac{(a d\theta)^2}{2a}$, the force $\frac{K}{a^2}$ will draw it in the same time through the space $K \cdot \frac{(a d\theta)^2}{2a}$. The inclination therefore of its new path to its old will be represented by the infinitely small angle

$$K \cdot \frac{(a d\theta)^2}{2a} = K d\theta$$

Let the new elementary portion be prolonged till it meets the ecliptic in ☿', ☿ being the former place of the node, then we shall have ☿ ☿' for the momentary change of the node's place. Now this is the side of a spherical triangle opposite to the infinitely small angle $K d\theta$. The included side is the arc of the moon's orbit between the moon and node, or the moon's distance from its node, while the included angle is I the inclination. Hence, if we call L the longitude of the node, we shall have, by spherical trigonometry,

$$- dL = K d\theta \times \frac{\sin(\theta - L)}{\sin I}$$

but $K = -3a \cdot \sin(\text{☾} - \text{☉}) \cdot \sin(\text{☉} - \text{☿}) \cdot \sin I$. Hence, we have

$$dL = 3a \cdot \sin(\theta - k\theta) \cdot \sin(k\theta - L) \cdot \sin(\theta - L) d\theta$$

from which differential equation the relation between L and θ may be deduced.

If we assume L as constant during one lunation in the second member, we get by integration the whole change of L in that interval approximately, or the mean motion of the node in a lunation, which we will call ΔL

$$\Delta L = 3a \cdot \int d\theta \cdot \sin(1-k)\theta \cdot \sin(k\theta - L) \cdot \sin(\theta - L)$$

This is in principle the method followed by Newton in that part of the third book of the *Principia*, where he treats of the motion of the moon's node; the most elegant and satisfactory instance of the application of his geometry to the lunar theory.

The precession of the equinoxes is explicable on the same principles as the motion of the moon's nodes. The centrifugal force at the earth's equator throws out a portion of the matter of which it consists into the form of an oblate or flattened spheroid; and we may conceive this redundant matter, as a spheroidal shell investing an inscribed sphere. Suppose now every particle of this shell at liberty to obey any impulse, unfettered by the others, or conceive it to consist of an infinite number of infinitely small moons: each of these will describe an orbit, whose nodes have a tendency to recede on the plane of the ecliptic, and though some (those which happen to lie between their quadratures with the sun and their nodes,) will have their nodes in a state of advance; all the rest, which constitute the greater number, will have theirs in a state of recess. Conceive now the particles to cohere and form a solid ring, *unconnected* with the central globe, the motion of this ring will be a mean among all the motions of its parts, and the nodes of the ring will in consequence continually recede, with a certain velocity. Now, let the ring adhere to the sphere, then must all its motion be divided between itself and the whole mass of the earth, which *alone* has no such tendency. The redundant matter at the equator, however, bears a very small ratio to the whole mass of the earth; so that owing to this cause, the retrogradation is exceedingly diminished in rapidity; and owing to the enormous distance of the sun and the smallness of the earth (whose radius being only $\frac{1}{23405}$ -th part of the

distance; so that here, $a = m' \cdot \left(\frac{a}{a'}\right)^3 = \frac{m'}{(23405)^3}$ a quantity quite insensible,) is rendered too small to be distinctly perceived.

But the moon also exerts a disturbing force. That luminary is to our imaginary moons, or terrestrial molecules, what the sun is to the moon itself in the theory of the lunar perturbation. If we reduce these principles to calculation, assuming such a mass and distance of the moon as we know to be near the truth, we shall find that a retrograde motion of the earth's equator on the ecliptic of about 50" per annum will actually result. Now this is the very phenomenon known by the name of the precession of the equinoxes; and though its strict theory is much more complicated and difficult than the general view here taken, its accordance with observation is perfect, and affords one of the most refined verifications of that admirable law which holds the frame of nature in the harmony we are now so well able to appreciate.

T A B L E

OF THE

PRINCIPAL MATTERS IN THE

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ASTRONOMY.

Physical.

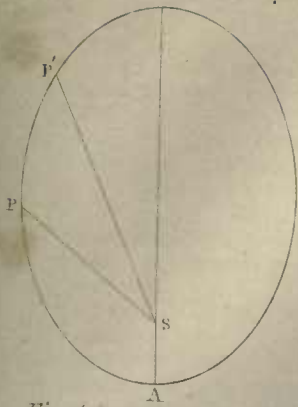
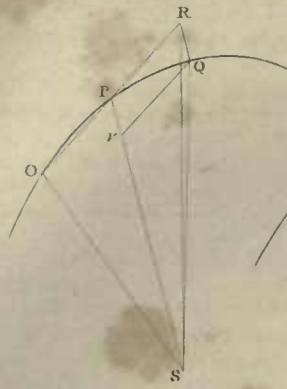
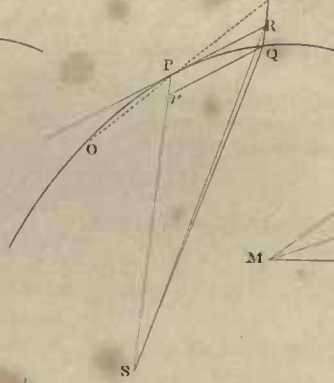


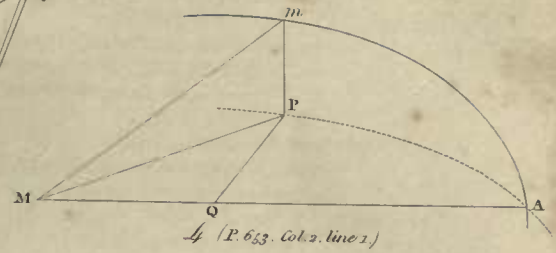
Fig. 1. (P. 649. Col. 2. line 34.)



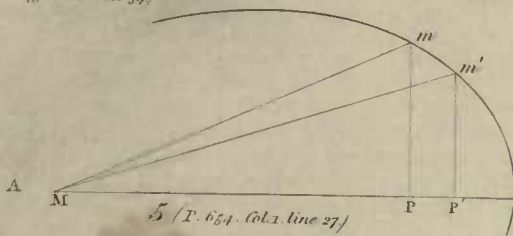
2 No. 1. (P. 649. Col. 2. line 32.)



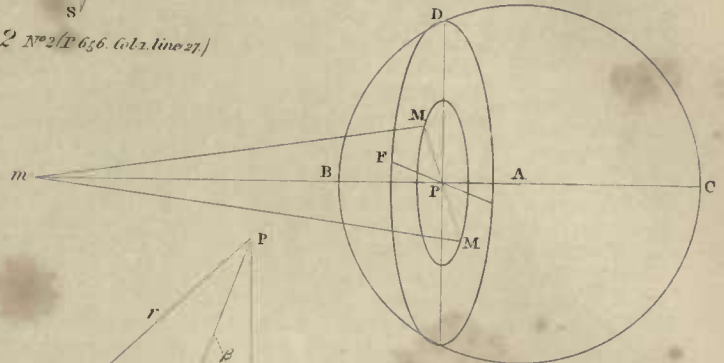
2 No. 2. (P. 656. Col. 2. line 27.)



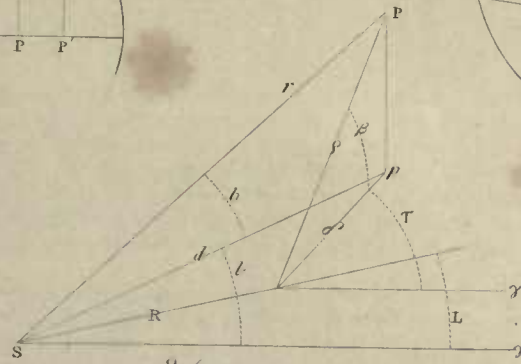
4 (P. 653. Col. 2. line 1.)



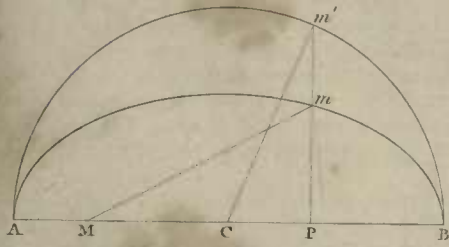
5 (P. 654. Col. 1. line 27.)



3 (P. 651. Col. 2. line 33.)

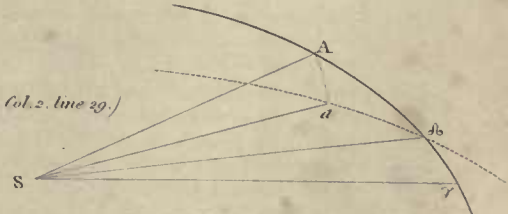


8 (P. 664. Col. 2. line 11.)



6 (P. 655. Col. 2. line 21.)

9 (P. 664. Col. 2. line 29.)



7 (P. 659. Col. 2. line 15.)

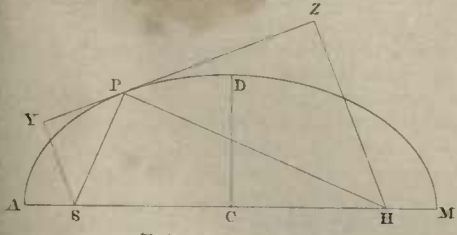
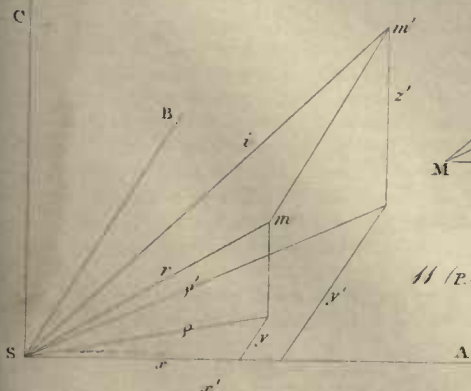
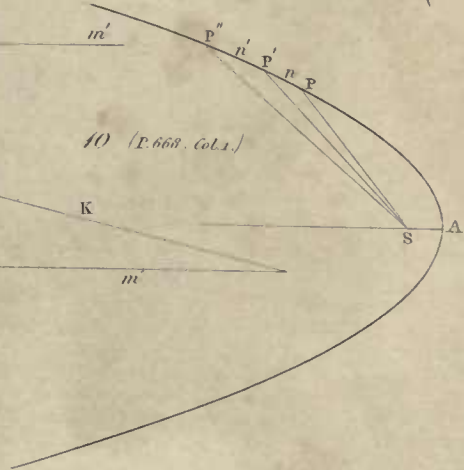


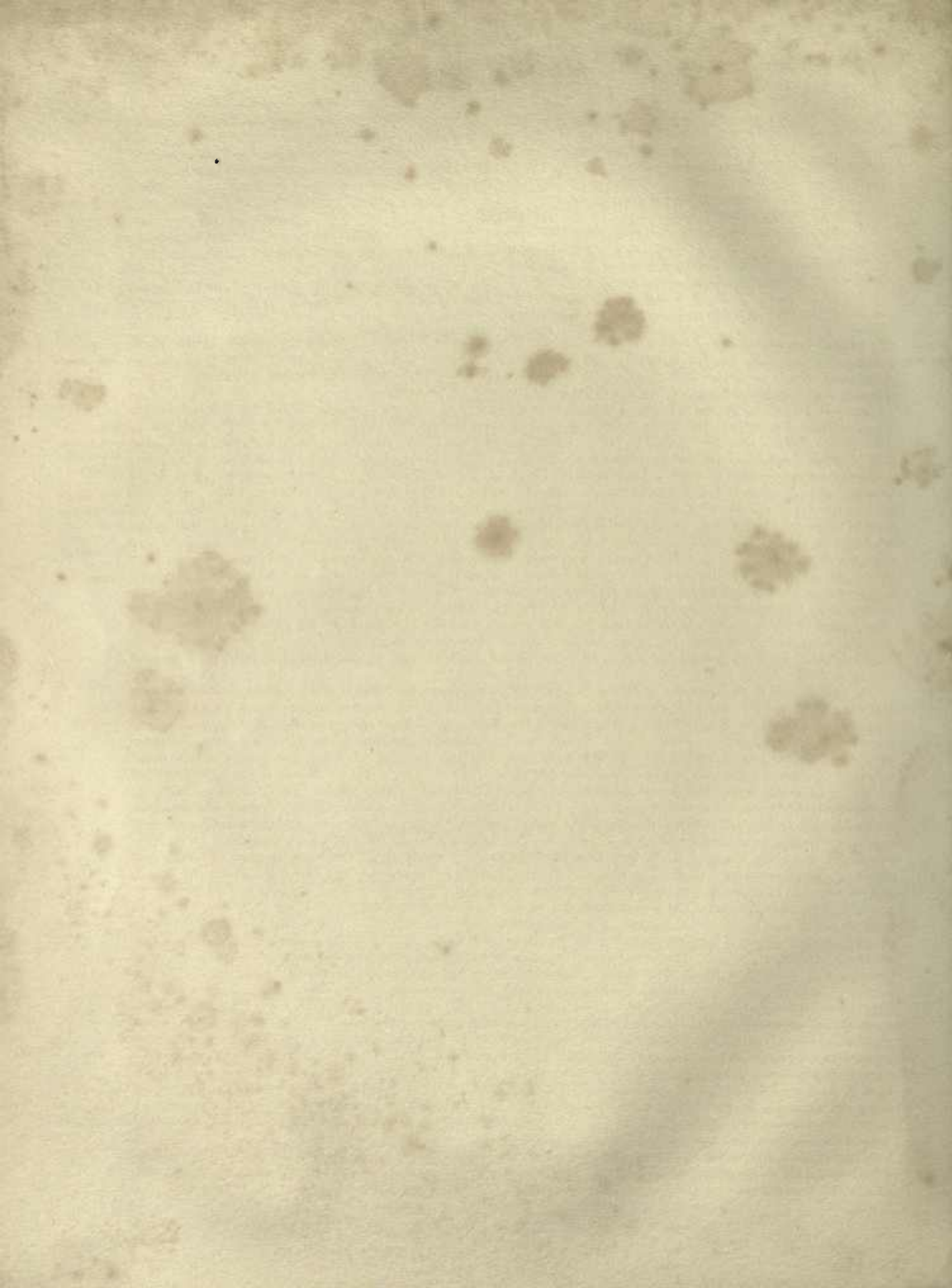
Fig. 4.

11 (P. 674. line 7.)



10 (P. 668. Col. 1.)





SOUND.

PART I.

OF THE PROPAGATION OF SOUND IN GENERAL.

§ I. Of the Propagation and Velocity of Sound in Air.

Sound.

To explain the nature and production of Sound, the laws of its propagation through the various media which convey it to our ears, and the manner of its action on those organs; the modifications of which it is susceptible in speech, in music, or in inarticulate and unmeaning noises; and the means, natural or artificial, of producing, regulating, or estimating them, are the proper objects of Acoustics.

Every body knows that Sounds are conveyed to our ears from a distance through the air, but it is not equally apparent that they would not reach us as well through a space perfectly void; or, in other words, that the air itself is the vehicle, or active agent, by whose operation they are conveyed to us. Such, however, is the case. Shortly after the invention of the air-pump, it was found that the collision of hard bodies in an exhausted receiver produced no appreciable Sound. Hanksbee (*Philosophical Transactions*, 1705) having suspended a bell in the receiver of an air-pump, found the Sound die away by degrees, as the air was exhausted, and again increase on its readmission; and when made to sound in a vessel full of air, the Sound was not transmitted through the interval between that and an exterior vessel from which the air had been extracted, though it passed freely when readmitted. On the other hand, when the air was condensed in a receiver, the Sound of a suspended bell was stronger than in natural air, and its intensity increased with the degree of condensation. Roebuck, (*Transactions of the Royal Society, Edinburgh*, vol. v. p. 34.) when shut up in a cavity excavated in a rock, which served as a reservoir of air for an iron foundry in Devonshire to equalize the blast of the bellows, observed the intensity of Sound to be considerably augmented in the air thus compressed by their action. The same effect has been experienced in diving-bells. More recently M. Biot has repeated the experiment of the exhausted receiver, with a much more perfect vacuum than could be procured in Hanksbee's time; and found the Sound to be quite imperceptible, even when the ear was held close to the receiver, and all attention paid. (*Mém. d'Arcueil*, vol. ii. p. 97.)

The diminution of the intensity of Sound in a rarefied atmosphere is a familiar phenomenon to those who are accustomed to ascend very high mountains. The deep silence of those elevated regions has a physical cause, independent of their habitual solitude. Saussure relates, that a pistol fired on the summit of Mont Blanc, produced no greater report than a little Indian cracker (*petit petard de Chine*) would have done in a room. (*Voyage dans les Alpes*, vol. vii. p. 337, § 2020.) We have ourselves had occasion to notice the comparatively small extent to which the voice can be heard, at an altitude of upwards of 13,000 feet on Monte Rosa. Observations on this point in the elevated passes of the Himalaya Mountains would be interesting. They should be made by the explosions of a small detonating pistol, loaded with a constant charge, and the distances should be measured; for the voice loses much of its force from the diminution of muscular energy in rarefied air, and distances are extravagantly underrated by estimation in such situations. The height, however, to which an atmosphere, or medium capable of conveying Sound extends, far exceeds any attainable on mountains, by balloons, or even by the lightest clouds. The great meteor of 1753 produced a distinct rumbling Sound, although its height above the earth's surface was full 50 miles at the time of its explosion. (See Sir Charles Blagden's interesting Paper, *Philosophical Transactions*, 1784.) The Sound produced by the explosion of the meteor of 1719, at an elevation of at least 69 miles, was heard as "the report of a very great cannon, or broadside;" shook the doors and windows of houses, and threw a looking-glass out of its frame and broke it. (Halley, *Philosophical Transactions*, vol. xxx. p. 978.) These heights are deduced by calculation from observations too unequivocal, and agreeing too well with each other, to allow of doubt. Scarcely less violent was the Sound caused by the bursting of the meteor of July 17, 1771, near Paris; the height of which, at the moment of the explosion, is assigned by Le Roy at 20,598 toises, or about 25 miles. (*Mém. Acad. Par.* 1771, p. 668.) The report of a meteor, in 1756, threw down several chimneys at Aix in Provence, and was taken for an earthquake. These instances, and others which might be adduced, are sufficient to show that Sound can be excited in, and conveyed by, air of an almost inconceivable tenuity (for such it must be at the heights here spoken of) provided the exciting cause be sufficiently powerful and extensive, neither of which qualities can be regarded as deficient in the case of fire-balls, such as those of 1719 and 1783, the latter of which was half a mile in diameter, and moved at the rate of 20 miles in a second. It may, however, be contended, and not without some probability, that at these enormous heights Sound may owe its propagation to some other medium more rare and elastic than air, and extending beyond the limits of the atmosphere of air and vapour.

Sound is not *instantaneously* conveyed from the sounding body to the ear. It requires time for its propagation. This is a matter of the most ordinary observation. We hear the blows of a hammer at a distance, a very sensible interval of time after we see them struck. The report of a gun is always heard later than the flash is seen, and

Part I.

1.

2.
Sound is conveyed to us by means of the air.

Diminution of Sound in rarefied air.

Its increase in condensed air.

3.

On high mountains.

Extent of the Sound-ing atmosphere.

Sound.
Sounds not
instanta-
neously con-
veyed.

5.
Velocity of
Sound.

Influence of
the wind
on it.

6.
Various de-
terminations
of the velo-
city of
Sound.

7.
Derham's
experi-
ments.
Circum-
stances in-
fluencing
the propa-
gation of
Sound.

9.
Experi-
ments of the
French Aca-
demicians,
1738.
Mode of
observing
by reciproc-
al signals.

the interval is longer the more distant the gun. We estimate the distance of a thunder-storm by the length of the interval between the lightning and the thunder-clap, which often arrives when we have ceased to expect it. The report of the meteor of 1783 was heard at Windsor castle, ten minutes after its disappearance. This is, probably, the longest interval yet observed.

A great multitude of experiments have been made to determine the precise velocity of Sound. The earlier results differ more than might have been expected, from the influence of several causes not immediately obvious, but chiefly from want of due attention to the influence of the wind. It is evident from the mechanical concussion attending loud noises, that Sound consists in a motion of the air itself communicated along it by virtue of its elasticity, as a tremor runs along a stretched rope. If, then, the whole body of the air were moving in a contrary direction, with the velocity of Sound, it would never make its way against the stream at all; and, on the other hand, when the wind blows from the sounding body direct towards the ear, it is equally clear that the velocity of the wind itself will be added to that of Sound in still air. If a stone be thrown into a still lake, the waves spread with equal rapidity in all directions, in circles whose centre is the stone. If into a running river, they still form circles, but their centre is carried down the stream; and, in point of fact, the wave arrives opposite to a point of the bank *above* the place where the stone fell, later than a point at the same distance below it in proportion to the rapidity of the stream. Hence all experiments on the velocity of Sound ought to be made, if possible, either in calm weather, or in a direction at right angles to that of the wind.

The assumption of 1300 feet per second for the velocity of Sound by Roberts, (*Phil. Trans.* 1694,) and the inaccurate determinations of Mersenne, Bayle, and Walker, (*Phil. Trans.* 1698,) which give respectively 1474, 1200, and 1305 feet, (the latter by a mean of 12 experiments disagreeing no less than 370 feet *inter se*.) scarcely deserve more mention than the rude guesses of Roberval and Gassendi, which differ by an amount nearly equal to the whole quantity to be measured; the former fixing it at 560 feet, the latter at 1473. The first experiments which appear to have been made with any degree of care, were those instituted by the Florentine Academy *Del Cimento*. It was observed in these that at a distance equal to 5739 English feet, the Sound of a harquebuss arrived five seconds after the flash; and repeating the experiment at half the distance, they found exactly half the time to be required. This gives, for the velocity of Sound, 1148 feet per second.

Cassini the Elder, Picard, and Roëmer, from experiments made at a distance of 1280 toises, as related by Duhamel in the *Hist. de l'Acad. Par.* assign 1172; while Flamsteed and Halley, from a series of observations at the Royal Observatory, the origin of the Sound being three miles distant, concluded the velocity to be 1142 feet per second.

In a Paper communicated to the Royal Society in 1708, by Dr. Derham, the subject of the velocity of Sound is investigated more fully and distinctly than had before been done, and with some degree of attention to a variety of circumstances which appear likely to influence its propagation. These are chiefly

1. The direction and velocity of the wind.
2. The amount of barometric pressure.
3. The temperature of the air through which the sound is conveyed.
4. Its hygrometrical state of moisture and dryness.
5. The actual weather, whether fog, rain, snow, sunshine, &c.
6. The nature of the Sound itself, whether produced by a blow, a gunshot, the voice, a musical instrument; its pitch, quality, and intensity.
7. The original direction impressed on the Sound—by turning, for instance, the muzzle of a gun one way or the other.
8. The nature and position of the surface over which the Sound is conveyed; whether smooth or rough, horizontal or sloping; moist or dry, &c.

8. To all these circumstances, except the wind, Derham attributes no effect; and, in fact, none of them, except the temperature of the air, have been ascertained to exercise any material influence on the velocity; though many, indeed all, have a very powerful one on its intensity, or the loudness of the Sound as it reaches the ear from a given distance. The quantity of aqueous vapour indeed ought (as we shall see) to affect the velocity, but in a degree only appreciable in the most delicate experiments. Derham concludes, from the whole of his observations, that Sound is propagated at the rate of 1142 feet per second, agreeing with the result of Flamsteed and Halley, and with that of the Florentine Academicians; and as the distances of the stations employed were considerable, in one case amounting to upwards of 12 miles, this determination appears deserving of some reliance. The temperature, unfortunately, was not registered; so that the experiment loses much of its value from the impossibility of applying with certainty the requisite correction.

9. In 1737-1738, the Academy of Paris directed a reinvestigation of the subject, and Messrs. Cassini de Thury, Maraldi, and La Caille, who were at that time engaged in the triangulation of France, were charged with the conduct of the experiments; an account of which, by Cassini, is to be found in the volumes of the *Histoire de l'Acad.* for the latter year and for 1739. Their observations were carefully made, and the distance of the stations was considerable, (from 2931 to 16,079 toises.) In these experiments we find the first example of observations so disposed as to eliminate in some measure the disturbing effect of the wind. To apprehend how this may be done, let us suppose a current of wind to blow *uniformly* with any velocity from one station A to another B at any distance, and at these two stations let shots be fired. The Sound of the shot fired at A will then be accelerated, and that of the signal at B will be retarded, in traversing the interval, by equal quantities; and consequently (since the velocity of Sound is very much greater than that of the most violent wind) the time in which the Sound runs over the line AB will be diminished, and that in which it traverses BA increased, by equal quantities; so that the mean will be unaffected by the wind's velocity. In fact, supposing V to be the velocity of Sound, v that of the wind, and S the space described, the velocities of the Sound in the two opposite directions

Sound.

will be $V + v$ and $V - v$; and the times of description of the space AB will be $\frac{S}{V + v}$ and $\frac{S}{V - v}$ whose Part 1.
 mean is equal to $\frac{SV}{V^2 - v^2}$, or to $\frac{S}{V} \left\{ 1 + \left(\frac{v}{V} \right)^2 + \left(\frac{v}{V} \right)^4 + \&c. \right\}$, which when v is small with respect to V , reduces itself simply to $\frac{S}{V}$. The most violent hurricane moves at a rate less than one-tenth that of Sound;

so that in the worst case the neglect of the terms depending on the velocity of the wind will entail an error less than $\frac{1}{10}$ of the whole result, or about 11 feet; and under ordinary circumstances such as are likely to be selected for experiment, their influence is quite inappreciable.

It is evident, however, that any want of uniformity in the rate of the wind will destroy, so far as it goes, the precision of the result so obtained; and that, in consequence, if the corresponding signals are not *strictly simultaneous* so as to make the Sound traverse the same identical portion of the aerial current, a great part of the advantage of this mode of experimenting is lost. M. Arago has indeed remarked, that even in that case, if the wind be very irregular, and in sudden gusts, it will still affect the result; to conceive which, we will suppose a gust of wind to arise suddenly at the station A at the moment of firing the signals both at A and B . The Sound which proceeds in the direction AB , as it runs quicker than the wind, will leave it behind, and be propagated at every point of AB in still air, before the agitation of the wind has had time to reach it. On the other hand, the Sound from B will meet the wind; and, during the latter part of its course, at least, will be propagated in a moving atmosphere. Still, it will be observed, that it can be *only* the latter part of its course which can be thus affected, less, at all events, than one-tenth of the whole space; and the effect during that tenth being to retard the Sound by one-tenth at most of *that interval*, will produce a total effect, not exceeding a hundredth of the whole time of traversing AB ; and, consequently, will affect the mean of the two deduced velocities by a quantity not exceeding a two-hundredth part of its value, or about five feet per second. We have already seen that the neglect of the square and higher powers of the velocity of the wind may in the same extreme case produce double this amount of error. This, however, is the error produced by a sudden gust equal to the most violent tornado. In ordinary winds, then, it must be quite inappreciable; and the method of simultaneous observations at opposite stations, provided they be *strictly* such, may be regarded as completely eliminating the wind's influence.

In the experiments of Cassini and his colleagues, however, none of these niceties were attended to; a long interval elapsed between the corresponding observations when obtained; and, indeed, the greater part of their series was made without any regard to correspondence at all. They conclude the velocity of Sound to be 173 toises, or 1106 British feet per second, at the temperature between 4° and 6° Reaum. at which the experiments were made. The extreme difference of velocities due to a favourable and a contrary wind they found to be about one-eleventh of the whole, giving $\frac{1}{11}$ for the ratio of the velocity of the wind to that of Sound as their maximum, or 50 feet per second. When the correction for the temperature of the air is applied, it will be seen presently that their result justifies the reliance placed on it by its authors; being, in fact, within about a yard of the truth.

Nearly about the same time Bianconi in Italy, and La Condamine at Quito and at Cayenne, instituted a series of experiments for the same purpose, of which accounts will be found in the *Comment. Bonon.* ii. p. 365; in La Condamine's *Introduction Historique*, &c. 1751, p. 98; and in the *Mém. Acad. Par.* 1745, p. 448. But the theory of Sound being at that time but imperfectly understood, and the necessary corrections in consequence being not sufficiently, or not at all attended to, the subject has been regarded as still open to further discussion; and accordingly a great number of researches by later experimenters have been instituted, of which the principal are those by Müller in 1791, (*Götting. Gelehrte. Anzeigen*, 1791, No. 159;) by Espinosa and Bauza, in Chili, in 1794, (*Ann. de Chim.* vii. 93;) by Benzenberg in 1809, (*Gilbert's Annalen*, new series, v. 383;) by Arago, Bouvard, Matthieu, Prony and Humboldt, and Gay Lussac, in 1822, (*Connaiss. des Temps*, 1825, p. 361;) by Moll, Vanbeek, and Kuytenbrouwer in Holland, in 1822, (*Phil. Trans.* 1824, p. 424;) by Mr. Goldingham, in 1820, at Madras, (*Phil. Trans.* 1823, p. 96;) by Dr. Gregory, at Woolwich, in 1823, (*Trans. of Cambridge Phil. Soc.* 1824;) and by General Myrbach and Professor Stampfer, at Salzburg, (*Jahrbuch des Polytechn. Instituts zu Wien*, vol. vii.)

Of these by far the most considerable and circumstantial, as well as in all probability, from the instrumental means employed and precautions used, the most exact, are those of the Dutch and the Parisian Philosophers. Every attention was paid in them (at least in the case of the Dutch experimenters) to obtain signals *strictly* reciprocal, by guns fired at the same instant of time at the two ends of the line of observation; all those corrections depending on Meteorological circumstances which theory points out, and which it will be the object of subsequent parts of this Essay to explain, being carefully applied; and the distances of the stations being at once considerable, and determined with sufficient exactness by Trigonometrical operations.

One very material difficulty in the way of former observers (Benzenberg excepted) lay in the want of adequate means of measuring with precision intervals of time to a minute fraction of a second. This difficulty was obviated in the experiments of the French Commissioners, by the use of the stop-watch of Breguet, and the *chronograph* of Rieussec, a species of watch, one of whose hands performs a revolution per second, and can be made to touch with its extremity the dial-plate, at any instant, and leave there a dot, without interrupting its motion of rotation, by the sudden pressure of a small lever; to effect which it carries with it a drop of printer's ink in a peculiar and ingenious species of dotting pen. In the Dutch experiments, a clock with a conical pendulum was used, capable of determining intervals to the hundredth of a second, by suddenly suspending the motion of the index, without stopping the clock. By the use of these instruments it was found practicable to ascertain the

10.
Influence of
sudden gusts
of wind.

11.
Cassini's
result.

12.
Other deter-
minations.

13.
Experi-
ments of
Moll, Van-
beek, &c.
and of
Arago, Mat-
thieu, &c.

14.
Methods of
measuring
very small
portions of
time.

Sound. interval between the sight of the flash, and the arrival of the report, of a gun, with such precision as to destroy all material error in the result which might arise from this cause; an improvement of great importance, when we consider that an error of a single tenth of a second in the measure of time is equivalent to 110 feet in that of distance.

15. The close agreement of the results of these experiments is a convincing proof of their accuracy. The French Philosophers state 331·05 met. = 1086·1 feet, as the velocity of propagation of Sound in air of the temperature of freezing water, while the Dutch experimenters make it 332·05 met. = 1089·42 feet in perfectly dry air of the same temperature. The latter seems to deserve the preference. if only from the circumstance of the signals from which it is deduced having been strictly simultaneous, the guns at the two extremities of the line (nine miles in length) having been fired at the same second of time, while in the former series this exact coincidence was not obtained.

16. We subjoin a list of the results arrived at in the various determinations above enumerated, with their dates, the distances of the stations employed, &c. to bring the whole subject under one view.

TABLE I.—*Velocity of Sound as determined by various Experiments.*

Observers' Names.	Date of Determination. A. D.	Distance of Stations in Feet.	Velocity in English Feet per second.	Remarks.
Mersenne	1474	
Florentine Academicians	1660	5906	1148	Moll and Vanbeek state this result at 361 metres = 1184 feet. Our authorities are Derham and Walker.
Roberval	560	
Gassendi	1473	
Boyle	1200	<i>Essay of Languid Motion.</i>
Roberts	1694	1300	No experiments stated.
Walker	1698	{ variable 600 to 2370 }	1305	By return of echos in given times and measuring distance.
Cassini, Picard, Roëmer	8186	1172	Duhamel.
Cassini, Huygens	9239	1151	Moll on authority of Duhamel.
Flamsteed, Halley	15840	1142	As stated by Derham.
Derham	1704	{ 5280 to 63360 }	1141	
Cassini de Thury, Maraldi, Lacaille	1738	{ 18744 to 102824 }	1106	Near Paris at Montlhery, Dammartin, &c. Therm. +5° Reaum. consider their result as within a fathom of the truth.
Cassini, Lacaille	1739	144124	1093 1110	Do. reduced to freezing temperature. Between Sette and Aiguesmortes, <i>Mém. Acad. Par.</i> 1739, p. 127, temperature not stated, probably about + 6° R.
Bianconi	1740	78740	1043	
La Condamine	1740	67400	1112	At Quito.
La Condamine	1744	129360	1175	At Cayenne.
Millington	1130	Cited by Goldingham, (<i>Phil. Trans.</i> 1823.)
T. F. Mayer	1778	3412	1105	
G. E. Müller	1791	8530	1109	
Pictet	1130	Cited by Dr. Gregory, (<i>Trans. Phil. Soc. Cambridge</i> , ii. 120.)
Espinosa and Bauza	1794	{ 53626 to 14071 }	1222·23	At Chili, at a temperature = 74° 7' Fahr. mean of four determinations, and mean temperature, the mean taken giving a weight to each proportional to the distance.
Benzenberg, (Dusseldorf)	1809	29764	1093	At freezing temperature.
Arago, Matthieu, Prony, Bouvard, Humboldt	1822	61064	1086·1	At freezing temperature, (between Villejuif and Montlhery.)
Moll, Vanbeek, Kuytenbrouwer	1823	57839	1089·42	In dry air, at freezing temperature.
Gregory	1823	{ Various 2700 to 13460 }	1088·05	Mean of eight results given by Dr. Gregory, each separately reduced to the freezing temperature.
Myrbach	1822	32615	1092·1	Mean of 88 observations reduced to the freezing temperature difference of level of stations = 4474
Goldingham, (Madras)	1821	{ 29547 13932 mean }	1089·9 1079·9 1086·7	Hygrom. 20·31 } Reduced to the freezing temperature. The mean taken by attributing to each determination a weight proportional to the distance of the stations. The nature of the hygrometer not stated.

17. The agreement between such of the above results as are reduced to the standard or freezing temperature, *i. e.* of the last six, and the first determination of Cassini at Paris, is very close; their extreme discrepancy being less than seven feet, or a 160th of the whole amount, and their mean (1089·7) agreeing almost precisely with the result of Moll, Vanbeek, &c.; we may, therefore, adopt 1090 feet without hesitation (as a whole number) as no doubt

Sound

within a yard of the truth, and probably within a foot. The reduction to the zero of temperature has been made (when not performed by the authors themselves) on the supposition that every additional degree of atmospheric temperature, on Fahrenheit's scale, adds 1.14 foot to the velocity, a correction of which the grounds will be hereafter explained. (See Art. 68.)

It may, therefore, be stated in round numbers, that Sound, in dry air and at the freezing temperature, travels at the rate of 1090 feet, or 363 yards per second, and that at 62° Fahrenheit (which is the standard temperature of the British metrical system) it runs over 9000 feet in eight seconds, $12\frac{3}{4}$ British standard miles in a minute, or 765 miles in an hour, which is about three-fourths of the diurnal velocity of the Earth's equator.

Hence, in latitude $42\frac{1}{2}^{\circ}$, ($42^{\circ} 29' 40''$.) if a gun be fired at the moment a star passes the meridian of any station, the Sound will reach any other station exactly west of it at the precise instant of the same star's arriving on its meridian.

In the experiments of Dr. Gregory, the velocity of the wind was measured by an anemometer, and allowed for. The close agreement of their results with those of the Dutch and French observers, when the smallness of the distances is taken into consideration, is a strong proof of the care and accuracy with which they were made. The observations of Mr. Goldingham, or at least his mode of stating and reducing them, has been strongly, but we think undeservedly, censured in Poggendorff's *Annalen der Physik*, 81. Band. s. 490. He takes a mean of all the velocities observed daily, in calm weather, during a very long time, by the firing of a morning and evening gun at two stations visible from Madras, and a mean of all the temperatures, pressures, and hygrometer-readings. All that we have done is to apply the correction for this mean temperature to his mean velocities, as if they had been given by a single observation, a course, no doubt, perfectly legitimate, and saving a world of calculation. It is to be lamented that the nature of his hygrometer is not stated, as its indications at present are perfectly useless. The experiments of Espinosa and Bauza differ so enormously in their result from the rest, even when reduced to the freezing temperature, that most probably some fundamental mistake, either in their measurement of the distances, or in the calculations founded on them, must have been committed. Our authority is the *Annales de Chimie*, vol. vii. (N. S.) p. 93.

Derham found that fogs and falling rain, but especially snow, tend powerfully to obstruct the free propagation of Sound, and that the same effect was likewise produced by a coating of fresh fallen snow on the ground, though when glazed and hardened at the surface by freezing it had no such influence. Over water, or a surface of ice, Sound is propagated with remarkable clearness and strength. Dr. Hutton relates, that on a quiet part of the Thames, near Chelsea, he could hear a person read distinctly at 140 feet distance, while on the land the same could only be heard at 76. Lieutenant Foster, in the third Polar Expedition of Captain Parry, found that he could hold a conversation with a man across the harbour of Port Bowen, a distance of 6696 feet, or about a mile and a quarter. This, however remarkable, falls far short of what is related by Dr. Young on the authority of Derham, viz. that at Gibraltar the human voice has been heard ten miles, (perhaps across the Strait.) We have not been able to find the original passage either in his *Physico-Theology*, or in his dissertation *De Soni Motu*, in both which very remarkable instances are adduced, of which the following will suffice as specimens.

Guns fired at Carlscroon were heard across the southern extremity of Sweden as far as Denmark; 80 miles, as Derham states from memory, but according to the map at least 120.

Dr. Hearn, a Swedish physician, relates that he heard guns fired at Stockholm, on the occasion of the death of one of the Royal family in 1685, at the distance of 30 Swedish, or 180 British miles.

The cannonade of a sea-fight between the English and Dutch, in 1672, was heard across England as far as Shrewsbury, and even in Wales, a distance of upwards of 200 miles from the scene of action.

That Sounds of all pitches, and of every quality, travel with equal speed, we have a convincing proof in the performance of a rapid piece of music by a band at a distance. Were there the slightest difference of velocity in the Sounds of different notes, they could not reach our ears in the same precise order, and at the exact intervals of time in which they are played, nor would the component notes of a harmony, in which several Sounds of different pitch concur, arrive at once. M. Biot caused several airs to be played on a flute at the end of a pipe 951 metres, or 3120 feet, long, which were distinctly heard by him at the other end, without the slightest derangement in the order or intervals of sequence of the high and low notes. (*Mém. d'Arcueil*, ii. 422.) A better form of the experiment would have been to strike two bells of very different pitch one against the other, having removed their clappers. Both their sounds would (no doubt) arrive together.

A very material difference, however, is observed in the *intensity* with which Sounds are propagated, or the distances to which they may be heard with equal distinctness according to a great variety of circumstances. Thus, if a Sound be prevented from spreading and losing itself in the air, whether by a pipe, by the vicinity of an extensive flat surface, as a wall, or otherwise, it may be conveyed to very great distances with little diminution of force. This we observe familiarly in speaking pipes conducted from one apartment to another of a building. In the experiments already cited of M. Biot, a person being stationed at one end of the enormous tube above mentioned, (which was a combination of cast iron conduit pipes laid down for the supply of Paris with water, forming a continuous canal of equal internal diameter throughout, and having two flexures about the middle of its length) the lowest whisper at one end was distinctly heard at the other, so that, in fact, the only way not to be heard was not to speak at all. Nay, so faithful was the transmission of every agitation of the air, whether sonorous, or otherwise, along the pipe, that a pistol fired at one end actually blew out a candle at the other, and drove out light substances placed there with considerable violence.

At Carisbrook Castle, near Newport, in the Isle of Wight, is a well, 210 feet in depth, and 12 in diameter, into which if a pin be dropped, it will be distinctly heard to strike the water. The interior is lined with very smooth masonry.

It is evident, without entering into any nice theoretical considerations, that a mechanical impulse of whatever nature impressed on any portion of the air or other medium, whether fluid or solid, and thence communicated

Part I.

Velocity finally adopted = 1090 feet per second.

18.

Approximations in round numbers.

19.

Comparison with the Earth's diurnal motion.

20.

Remarks on some of the above results.

21.

Effect of fogs, &c. to obstruct Sound. Sound well conveyed over water and smooth ice.

22.

Distances at which Sounds have been heard.

23.

All sounds travel with equal velocity.

24.

Effect of pipes in conveying Sound.

25.

Sound in Carisbrook Well.

26.

Sound. to the surrounding parts, if allowed to spread in all directions as from a centre, must reach every more distant point with an energy continually less and less, because the same quantity of motion is communicated in succession to a larger and larger sphere of inert matter, but if only allowed to spread in certain directions, its diminution will be less rapid in proportion as the quantity of matter successively put in motion increases less rapidly. Hence a Sound might be expected to be conveyed with less diminution along a wall than in the open air, the trough or angle between the wall and the ground, in fact, forming two sides of a square pipe, and the divergence of the Sound in two directions being thereby in great measure prevented. Dr. Hutton relates that part of the wall of a garden, formerly in the possession of W. Pitt, Esq. of Kingston, in Dorsetshire, conveys a whisper in this way nearly 200 feet. (*Mathematical Dictionary*, Article *Sound*.) It is probably to some such principle that we must refer a fact mentioned by the last-named author, which at first sight appears surprising enough. He relates that when a canal of water was laid under the pit floor of the Theatre *Del Argentino*, at Rome, a surprising difference was observed. The voice has since been heard very distinctly when it was before scarcely distinguishable. It is a general remark that Sounds are well heard in buildings which stand on arches over water. The cause of this, however, seems to be the echo produced between the water and the arch which unites with, and reinforces, the original Sound. The Work just referred to contains many curious instances of the kind.

27. When Sound in the course of its propagation meets with an obstacle of sufficient extent and regularity it is reflected, producing the phenomenon we call an *Echo*. A wall, the side of a house, or the surface of a rock, the ceiling, floor, and walls of an apartment, the vaulted roof of a church, all, under proper circumstances, give rise to Echos more or less audible. The reflected Sound meeting another such obstacle is again reflected, and thus the Echo may be repeated many times in succession, becoming, however, fainter at each repetition till it dies away altogether. We shall here set down a few localities in which Echos, remarkable either for distinctness, or frequency of repetition, may be heard.

28. An Echo in Woodstock Park, (Oxfordshire,) repeats 17 syllables by day, and 20 by night, (Plot, *Nat. Hist. Oxford*, ch. i. p. 7.) One on the banks of the Lago del Lupo, above the fall of Terni, repeats 15.

Echos. Their nature. In the Abbey Church of St. Alban's is a curious Echo. The tick of a watch may be heard from one end of the church to the other. In Gloucester Cathedral, a gallery of an octagonal form conveys a whisper 75 feet across the nave.

29. An Echo on the north side of Shipley Church, in Sussex, repeats 21 syllables. (Cavallo, citing Plot and Harris.)


30. In the Cathedral of Girgenti, in Sicily, the slightest whisper is borne with perfect distinctness from the great western door to the cornice behind the high altar, a distance of 250 feet. By a most unlucky coincidence the precise focus of divergence at the former station was chosen for the place of the confessional. Secrets never intended for the public ear thus became known, to the dismay of the confessors, and the scandal of the people, by the resort of the curious to the opposite point, (which seems to have been discovered accidentally,) till at length, one listener having had his curiosity somewhat over-gratified by hearing his wife's avowal of her own infidelity, this tell-tale peculiarity became generally known, and the confessional was removed. (*Travels through Sicily and the Lipari Islands, in the Month of December, 1824*. By a Naval Officer. 1 vol. 8vo. London, 1827.)

31. In the Whispering Gallery of St. Paul's, London, the faintest Sound is faithfully conveyed from one side to the other of the dome, but is not heard at any intermediate point.

32. In the Manfroni Palace at Venice is a square room about 25 feet high, with a concave roof, in which a person standing in the centre, and stamping gently with his foot on the floor, hears the Sound repeated a great many times, but as his position deviates from the centre the reflected Sounds grow fainter, and at a short distance wholly cease. The same phenomenon occurs in the large room of the Library of the Museum at Naples.

33. Southwell (*Phil. Trans.* 1746, 223.) describes an Echo in an old Palace near Milan, which repeated the report of a pistol 56, or even 60 times. His description is singularly confused, but the palace is no doubt that of Simonetta, mentioned by Addison in his *Travels*. This was a building with two wings, forming three sides of a square. The pistol was discharged from a window in one wing, the Sound was returned from a dead wall in the other wing, and heard from a window in the back front. (Hutton, Art. *Echo*. Misson, *Voy. d'Ital.* ii. 196.) The Palace still exists, but ear-witnesses have described the phenomenon to us somewhat differently. The Echos are heard at the window whence the pistol is fired.

34. Beneath the Suspension Bridge across the Menai Strait, in Wales, close to one of the main piers, is a remarkably fine Echo. The Sound of a blow on the pier with a hammer, is returned in succession from each of the cross-beams which support the roadway, and from the opposite pier at a distance of 576 feet, and, in addition to this, the Sound is many times repeated between the water and the roadway. The effect is a series of

Sounds which may be thus written  &c. ; the first return

is sharp and strong, from the roadway over head, the rattling which succeeds dies away rapidly, but the single repercussion from the opposite pier is very strong, and is succeeded by a faint palpitation, repeating the Sound at the rate of 28 times in five seconds, and which therefore corresponds to a distance of 184 feet, or very nearly the double interval from the roadway to the water. Thus it appears, that in the repercussion between the water and roadway, that from the latter only affects the ear, the line drawn from the auditor to the water being too oblique for the Sound to diverge sufficiently in that direction. Another peculiarity deserves especial notice; viz. that the Echo from the opposite pier is best heard when the auditor stands precisely opposite to the middle of the breadth of the pier and strikes just on that point. As he deviates to one or the other side the return is proportionably fainter, and is scarcely heard by him when his station is a little beyond the extreme edge of

Sound,

the pier, though another person stationed (on the same side of the water) at an equal distance from the central point, so as to have the pier between them, hears it well. Thus, in the reflexion of Sound, there is an evident approach to the law of equality between the angles of incidence and reflexion which obtains in that of Light; and a tendency in the reflected Sound to confine itself to the direction which a ray of Light regularly reflected at the echoing surface would follow, and not to spread into the surrounding air equally in all directions. This experiment (which we had an opportunity of making, with the assistance of Mr. Babbage, in 1827) might be carried much farther under more favourable circumstances; and, we doubt not, would lead to remarkable confirmations of the law of interference, and the general analogy between Sound and Light, to which all Optical and Acoustical phenomena point, and of which we shall have occasion to say more hereafter. (See also our Essay on LIGHT.) The span of the bridge between the piers is 576 feet, and the breadth of each pier about 30 feet.

The most favourable position for the production of a distinct Echo from plane surfaces is, when the auditor is placed between two such, exactly half way. In this situation the Sounds reverberated from both will reach him at the same instant, and reinforce each other. If nearer to one surface than the other, the one will reach him sooner than the other, and the Echo will be double and confused. If the Echoing surface be concave towards him, the sounds reflected from its several points will, after reflexion, converge towards him, exactly as reflected rays of Light do; and he will receive a Sound more intense than if the surface were plane, and the more so the nearer it approaches to a sphere concentric with himself: the contrary if convex. If the Echo of a Sound excited at one station be required to be heard most intensely at another, the two stations ought to be *conjugate foci* of the reflecting surface, *i. e.* such that if the reflecting surface were polished, rays of Light diverging from one would be made after reflection to converge to the other. Hence if a vault be in the form of a hollow ellipsoid of revolution, and a speaker be placed in one focus, his words will be heard by an auditor in the other as if his ear were close to the other's lips. The same will hold good if the vault be composed of two segments of paraboloids, having a common axis, and their concavities turned towards each other; only in this case, Sounds excited in the focus of one segment will be collected in the focus of the other, after two reflexions.

An attention to the doctrine of Echos is of some, though we think a rather overrated, importance to the architect in the construction of buildings intended for public speaking, or music, especially if they be large. In small buildings, the velocity of Sound is such that the dimensions of the building are traversed by the reflected Sound in a time too small to admit of the Echo being distinguished from the principal Sound. In great ones, on the other hand, as in Churches, Theatres, and Concert rooms, the Echo is heard after the principal Sound has ceased; and if the building be so constructed as to return several Echos in very different times, the effect will be unpleasant. It is owing to this that in Cathedrals the service is usually read in a sustained uniform tone, rather that of singing than speaking, the voice being thus blended in unison with its Echo. A good reader will time his syllables, if possible, so as to make one fall in with the Echo of the last, which will thus be merged in the louder Sound, and produce less confusion in his delivery. For music, in apartments of moderate size, all objects which can obstruct the free reflexion of Sound from the walls, floor, and ceiling are injurious. The Echo is not sensibly prolonged after the original Sound, and therefore only tends to reinforce it, and is of course highly advantageous. In large ones, an Echo can only be advantageous in the performance of slow pieces, (as Church music.) The prolongation of a chord, after the harmony is changed, can be productive of nothing but dissonance. When ten notes succeed one another in a second, as is often the case in modern music, the longitudinal Echo of a room 55 feet long, will precisely throw the second reverberation of each note on the principal Sound of the following one wherever the auditor be placed; which, in most cases, will produce (in so far as it is heard) only discord. Much mistake seems to be prevalent on this subject. Thus it is said that the form of an orchestra should be parabolic, &c. that the rays of Sound should be reflected out in parallel lines to the audience. But even if they were so, the reflected Sound cannot possibly reach them in the same *time* with the direct; and in Acoustics it is of little moment in what direction sounds reach the ear, which is not, like the eye, capable of appreciating direction with any precision, or collecting the rays or waves of Sound to a focus within the ear. It is not possible to place a whole band in the focus of a parabolic or elliptic orchestra, or a whole audience in that of a corresponding opposite segment. We may add, too, that an apartment would be worse lighted, were its internal surface a polished semi-ellipsoid, with a candle in the focus, than if it were of the usual shape, and its walls and ceiling a dead white. The object to be aimed at in a Concert-room is, not to deafen a favoured few, but to fill the whole chamber equally with Sound, and yet allow the Echo as little power to disturb the principal Sound, by a lingering after-twang, as possible. But, whether for music or for oratory, open windows, deep recesses, hangings, or carpeting, and a numerous audience in woollen clothing, are all unfavourable to good hearing. They are to Sound, what black spaces in an apartment would be to light; they return back none, or next to none, of what falls on them. Their fault is not so much that they reflect it irregularly, as that they do not reflect it at all.

The rolling of thunder has been attributed to Echos among the clouds; and if it is considered that a cloud is a collection of particles of water, however minute, yet in a liquid state, and therefore each individually capable of reflecting Sound, there is no reason why very loud Sounds should not be reverberated confusedly (like bright lights) from a cloud. And that such is the case, has been ascertained by direct observation on the Sound of cannon. Messrs. Arago, Matthieu, and Prony, in their experiments on the velocity of Sound, observed, that under a perfectly clear sky, the explosions of their guns were always heard single and sharp, whereas when the sky was overcast, or even when a cloud came in sight over any considerable part of the horizon, they were frequently accompanied with a long continued roll like thunder, and occasionally a double Sound would arrive from a single shot.

But there is, doubtless, also another cause for the rolling of thunder, as well as for all its sudden and capricious bursts and variations of intensity, of which our knowledge of the velocity of Sound furnishes a perfect explanation. To understand this, we must premise that, *cæteris paribus*, the estimated intensity of a Sound will

Part I.

Equality of the angles of incidence and reflexion.

36.

Situations favourable to Echos.

37.

Effect of Echos in Churches and public buildings.

38.

Reverberation of Sound from the clouds.

39.

Explanation of thunder

Sound.

be proportional to the quantity of it (if we may so express ourselves) which reaches the ear in a given time. Two blows equally loud, at *precisely the same distance from the ear*, will Sound as one of double the intensity; a hundred, struck in an instant of time, will sound as one blow a hundred times more intense than if they followed in such slow succession that the ear could appreciate them singly. Now let us conceive two equal flashes of lightning, each four miles long, both beginning at points equidistant from the auditor, but the one running out in a straight line directly away from him; the other describing an arc of a circle having him in its centre. Since the velocity of Electricity is incomparably greater than that of Sound, the thunder may be regarded as *originating* at one and the same instant in every point of the course of either flash. But it will reach the ear under very different circumstances in the two cases. In that of the circular flash, the Sound from every point will arrive at the same instant, and affect the ear as a single explosion of stunning loudness. In that of the rectilinear flash, on the other hand, the Sound from the nearest point will arrive sooner than from those at a greater distance; and those from different points will arrive in succession, occupying altogether a time equal to that required by Sound to run over four miles, or about 20 seconds. Thus the same *amount* of Sound is in the latter case distributed uniformly over 20 seconds of time, which in the former arrives at a single burst; of course, it will have the effect of a long roar, diminishing in intensity as it comes from a greater and greater distance. If the flash be inclined in direction, the Sound will reach the ear *more compactly*, (*i. e.* in shorter time from its commencement,) and be proportionally more intense. If (as is almost always the case) the flash be zigzag, and composed of broken rectilinear and curvilinear portions, some concave, some convex to the ear; and if, especially, the principal trunk separates into many branches, each breaking its own way through the air, and each becoming a separate source of thunder, all the varieties of that awful Sound are easily accounted for.

40.
Phenomenon observed in the eruptions of volcanoes.

We will only mention one other phenomenon which is accountable for on the same principle. In the eruption of a volcano it is often remarked, that every ejection of stones, &c. is accompanied with an explosion like artillery when heard at a distance; but when near, the Sound resembles rather that of a loud and deep sigh, unaccompanied with any sudden burst. In both cases the cause of Sound is the same, the upward rush and displacement of the air by the stone; but where the auditor is near the bottom of the column of Sound, it reaches his ear more in detail than when at a distance, and therefore nearly equidistant from all its parts. In fact, let t = the time taken by the stone to rise to a height x , and let a be the distance of the observer from the bottom

of the column, and v the velocity of Sound, then will $t + \frac{\sqrt{a^2 + x^2}}{v}$ = time elapsed from the moment of ejection to that of the Sound of the column at the height x reaching the ear. Hence the whole Sound of the

portion x of the column will arrive in an interval of time represented by $t + \frac{\sqrt{a^2 + x^2} - a}{v}$. Now, as a

increases, x , and therefore t remaining constant, this function diminishes rapidly, and ultimately approaches t as its limit. Thus the Sound arrives continually more and more condensed. Should any discharge be made obliquely towards the observer's station, a still greater concentration of the noise will happen, as may be easily seen by considering that if shot directly towards him, with the velocity of Sound, the report would reach him from every part of the line strictly at the same moment. Now, as these ejections have been known to rise to a height of 10,000 feet, in spite of the resistance of the air, their initial velocity must be, at least, equal to that of Sound. At great distances it is probable that only the Sounds produced by such oblique ejections have intensity or (as we may express it) body enough to affect the sense.

§ II. Mathematical Theory of the Propagation of Sound in Air, and other Elastic Fluid Media.

41.
General notion of the communication of motions in elastic media. Propagation of tremors along a stretched cord. Fig. 1. Fig. 2.

A general notion of the mode in which an impulse communicated to one portion of the air, or other elastic fluid, is diffused through the surrounding portions, and successively propagated to portions at a greater and greater distance from the original source of the motion, may be obtained by considering the way in which a tremor runs along a stretched cord, or in which waves excited in the surface of still water dilate themselves circularly, and propagate a motion impressed on one point of the surface, in all directions to a distance. In the former case, conceive a blow given to a point in the middle of the cord, transversely to its length. The point to which the blow is given will be thrown out of the straight line, and a flexure, or angle, will be formed in that part. Owing, however, to the inertia of the cord, the displacement of the particles in the first instant will be confined to the immediate neighbourhood of the point of impulse; so that the cord will not at once assume the state represented in fig. 1, consisting of two straight portions AB, BC, forming a very obtuse angle ABC; but rather that in fig. 2, in which the greater part on either side AD, DC, retain their original position; and a small part DBE, proportioned to the violence and suddenness of the blow, is, as it were, bulged out into an angular form DBE. The particle at B then is solicited on both sides by the tension of the cord in directions BD, BE; but these tensions, which in the quiescent state of the string exactly counteracted each other, now only do so in respect of those parts of each which, when resolved, act in directions parallel to DA, EC respectively. The other resolved portions, perpendicular to these, conspire and urge the point B towards its point of departure b . As there is no force to counteract this (the impulse being supposed momentary) B will obey their solicitation, and approach b with an accelerated velocity. But, action and reaction being equal and contrary, the same force by which the molecule E drags B down, will be exerted on E to drag it up, or out of the line; so that

by the time B has performed half its course towards *b*, E will have been raised above the line, and will have acquired a velocity capable of carrying it still further in that direction. At this instant the cord will have assumed the figure A D' D B E E' C. At the next moment the forces are reversed, B then tends to drag both D and E down to the line; but its own acquired momentum is expended in the effort, and by the time it has reached its original place in the line, its inertia is destroyed, and it rests there without a tendency to go beyond it on the other side. Meanwhile, however, D and E have attained their greatest elevation; and thus the protuberance D B E is resolved into two D' D B and B E E' (of less height, however) on either side. In like manner the particles D and E, in returning to their places, drag up the next adjoining D' and E', and then the next, and so on; and thus the summits of the protuberances advance along the line, and correspond in succession to all its points; and the visible effect is an undulation, or wave, which runs along the cord with a velocity greater the greater is the force with which the cord is strained, as it manifestly ought to be, since the rapidity with which each molecule returns from its displaced situation is greater as the force urging it is so; and this force is nothing more than the resolved part of the tension.

In like manner, when a wave is excited in the surface of water, as when by throwing in a stone one portion is violently driven down, and the surrounding part heaped up above its natural level; this subsides and fills up the vacuity; but as its pressure takes place alike on both sides of the ridge, the fluid on the outside of the ridge is also pressed on, from below upwards, by the reaction of the fluid stratum which sustains the ridge, and whose pressure is propagated equally in all directions. Thus the ridge, in subsiding, not only fills up the central vacancy, but forces up another ridge exterior to it; and this, in subsiding, another, and so on; and thus an advancing wave is formed; and the same action taking place on all sides of the centre, the wave can advance no otherwise than in the direction of radii on all sides diverging therefrom.

It is by no means intended, in what is here said, to give an accurate account of what passes in either of these cases, (in fact, it is very far from being so, as the reader by a little attention will soon perceive,) but only to give a first conception of the propagation of motion by undulations or waves.

In this general account of the above cases, one thing, however, cannot fail to strike the reader, that the *wave* which advances on the surface of water—the sinuosity which runs along the stretched cord—are neither of them things, but *forms*. They are not moving masses advancing in the direction in which they appear to run, but outlines, or figures, which at each instant of time include all the particles of the water or the cord which, it is true, *are* moving, but whose motion is in fact *transverse* to the direction in which the waves advance. But this is by no means an essential condition. We may generalize this idea of a *wave*, and conceive it as the form, space, or outline, whether linear or superficial, comprehending all the particles of an undulating body which are at once in motion, (supposing, for the present, that the motion of each consists of a simple displacement and return to quiescence, and not in a repetition of several such displacements and returns in succession.)

The waves in a field of standing corn, as a gust of wind passes over it, afford a familiar example of the relation between the motion of the wave, and that of the particles of the waving body comprised within its limits, and of the mutual independence which may in certain cases subsist between these two motions. The gust in its progress depresses each ear, in its own direction, which, so soon as the pressure is removed, not only returns, by its elasticity, to its original upright situation, but by the impetus it has thus acquired, surpasses it, and bends over as much, or nearly as much, on the other side; and so on alternately, oscillating backwards and forwards in equal times, but continually through less and less spaces, till it is reduced to rest by the resistance of the air. Such is the motion of each individual ear; and as the wind passes over all of them in succession, and bends each equally, all their motions are so far similar. But they differ in this, that they commence not at once but successively. Suppose (to fix our ideas) the wind runs over 100 feet in a second, and that the ears stand one foot asunder, and each makes one complete vibration to and fro in a second. Suppose A (fig. 3) to be the furthest point which the wind at any given instant of time has reached, or the last ear which it has just bent, and let the action of the wind be regarded as lasting only for a single instant. Then will the next preceding ear B have already begun to rise from its bent position, the next C will have risen rather more, and the 25th ear G (since the distance A F is 25 feet, and consequently since $\frac{25}{100} = \frac{1}{4}$ of a second have elapsed since the wind was at G) will have gone through one-fourth of its complete vibration to and fro, and will have therefore just attained its upright position; so that the ears F, E immediately adjacent towards A will not yet have quite recovered their perpendicularity, but still lean somewhat forwards; while those on the other side H, I will have surpassed the perpendicular, and have begun to sway backwards; consequently at G the stalks will on both sides be convex towards G, and the ears in that place will be further asunder than in their state of rest, and will appear as it were *rarefied* when viewed by a spectator so distant as to take in a great extent at once. Still further in rear of the wind, as 50 feet, at L, the 50th ear will have swung backwards as far as possible, and will just have its motion destroyed. The preceding stalk, K, will still want somewhat of its extreme backward flexure; the subsequent one, M, will already have risen a little, and therefore the interval of the ears K, N will be just what it was in the state of rest. At L, then, the spectator will see the ears at their natural distances from each other. Again the 75th stalk, Q, in rear of the wind will have had time to rise again erect from its *backward* inclination, three-fourths of a second having elapsed since its first bending forward. The 74th, P, will not be quite erected; the 76th will have surpassed the erect state, and have again begun to lean forward. The stalks then on both sides of Q will curve *towards* Q, and their ears will therefore be closer together than in their natural state, and will appear *condensed* to the spectator above mentioned. Finally, the 99th, 100th, and 101st ears will be again in the same relative state as the 49th, 50th, and 51st; only leaning forwards instead of backwards, and therefore neither condensed nor rarefied. The field, then, will present to the spectator a series of alternate condensations and rarefactions of the corn ears, separated by intervals in their natural state of density; and this series will extend so far in rear of the wind, till the resistance of the air and want of perfect elasticity in the stalks shall

Part I.

42.

Propagation of waves in still water.

43.

44.

A wave not a progressive moving body but an advancing form.

45.

Example. Waves in a field of standing corn.

Fig. 3.

Sound. have reduced them to rest, and these alternations, by the difference of shading they offer, will become apparent to his sight as dark and bright zones.

46.
Velocity of the wave distinguished from that of its component parts.

It matters not, for our present purpose, that the impulse is, in the case here taken, not propagated mechanically from ear to ear by mutual impulse, but that each moves independently of all the rest. All we want to illustrate is the distinction between the *wave* and the *moving matter*, and the independence of their motions. The waves here run along with the speed of the wind, whatever that may be; for it is always the point 25 feet in rear of the wind that is most rarefied, and that at 75 that is most condensed; and the interval between the first and 100th ear, comprehending ears in every state or *phase* of their vibrations, is what we term a wave. The *velocity* of the wave, then, is, in this case, that of the wind; and is totally distinct from, and independent of, that of each or any particular ear. The one is a constant, the other a variable quantity; the one a general resulting phenomenon, the other a particular, individual, mechanical process, going on according to its own laws.

47.
And independent of their excursions from the state of rest.

Neither is it of the least consequence whether the excursions of the several stalks from their position of rest be great or little; whether the degree of bending, or force of the wind, be great or small, provided its velocity be constant. In the case of wind, indeed, the force depends on the velocity; but if we conceive the impulse given by a rigid rod made to sweep across the field, any greater or less degree of flexure might be given, with the same velocity, by a mere change of its level; but the velocity of the wave would still be that of the rod in every case.

48.
Breadth of the waves.

But with respect to the *breadth* of the wave, or the magnitude of that interval which comprises particles in every *phase* or state of their motion, going and returning, it is otherwise. This is a result depending essentially on the motions of the particles themselves; for we see evidently in the above instance, that this breadth, which is 100 feet, is equal to the space run over by the wind in a time equal to that of one complete vibration, going and returning, of each individual ear. Now this time depends only on the elasticity of the stalk, and the weight of the ear it carries. In general then we may state, that "The breadth of a wave is equal to the space run over by it in a time equal to that in which any molecule of the waving body performs one complete vibration, going and returning, through all the phases of its motion." In the case here taken, the motion of the individual molecules is not, as in the former instances, transverse to that of the wave, but parallel to it. It is then hardly to be termed a *form*, or an *outline*. To such a wave, the term *pulse* is often applied. Whatever be the nature of the internal motions, however, the general name wave or undulation will equally apply, and will be used in future indiscriminately for all sorts of propagated impulses. It is not even necessary that the motions of the constituent particles should be rectilinear, or even lie in one plane. We may suppose the impelling cause to be a whirlwind. In this case each ear will have a rotatory or twirling motion, or the stalk a conical one, simply, or in addition to its flexure in a vertical plane; but the wave is independent of these particularities.

Various species of waves.

49.
Sonorous waves propagated in air.

In the case just described, each particle is supposed to be set in motion by an external cause, and to be uninfluenced in its motions by the rest. It is, therefore, not a case of the *propagation* of motion at all. It is quite otherwise with Sound, or other similar cases, where every particle of a medium receives its whole motion from those which were moving before, and transmits it to others previously at rest. The problem to investigate the general laws of the communication of motion under such circumstances is one of the utmost complexity, and at present has been only resolved under very restricted conditions; enough, however, to verify principal facts, and establish leading points, in the doctrine of Acoustics. We shall be far from attempting to present here any thing approaching to a sketch of the profound geometrical researches which have been bestowed on this department of Physics, contenting ourselves with referring the reader for a knowledge of them to the various Memoirs of Euler, D. Bernouilli, Lagrange, Poisson, &c. See (1.) *Recherches sur la Nature et la Propagation du Son*, par L. de Lagrange, *Mém. Acad. Turin*, i. 247. (2.) Euler, *Recherches sur la Propagation des Ebranlemens dans un Milieu Elastique*, *Miscel. Turin*, ii. (3.) *Nouvelles Recherches sur la Propagation du Son*, par M. Lagrange, *Miscel. Turin*, ii. (4.) Euler, *De la Propagation du Son*, *Mém. Acad. Berlin*, 1759, p. 185, and *Supplement*, p. 210, and *Continuation*, p. 241. (5.) Euler, *Eclaircissement plus détaillé sur la Propagation du Son et sur l'Echo*, *Mém. Acad. Berlin*, 1765, p. 335. (6.) Poisson, *Sur l'Intégration de quelques Equations Linéaires des Différens Partielles*, *Mém. de l'Acad. Paris*, 1818, p. 121. (7.) Poisson, *Sur la Théorie du Son*, *Journal de l'Ecole Polytechn.* xiii. 319.: while we confine ourselves to just so much developement of the mathematical analysis of the subject as will suffice for the demonstration of the chief theoretical propositions we shall have occasion for in the sequel.

50.
Propagation of Sound in air in one dimension. Fig. 4.

Let us then consider, as the simplest case, the propagation of Sound in a straight canal of equal bore throughout, filled with air or any other elastic fluid of equable density and elasticity, unacted on by gravity, and of which the transverse section is so small, and the sides so perfectly polished, that we may regard the motions of all particles in the same section as exactly similar; so that each section shall merely advance and recede in the pipe, without any lateral change of place of its constituent molecules *inter se*. Let AB (fig. 4) be such a pipe, and let any section of it, as A, be agitated by an external cause, with any arbitrary motion, *i. e.* one whose duration and extent, and whose velocity at every instant, shall be entirely dependent on the will, or, if we please, the caprice of an external operator sufficiently powerful to command it; and let us inquire how any other section whatever, situated at any assigned distance, x , from A. will move in consequence of this arbitrary motion of A.

51.
Analysis of this case.

Let us then conceive, that, in general, the section or stratum of molecules $aabb$, whose distance from the initial place A of the section A is represented by x , shall, after the lapse of any time t , have been transported into the situation $a\alpha\beta\beta$, at a distance $Aa = y$ from the same fixed point A. Let x' , x'' , &c. be the distances of the next consecutive sections from the fixed point A, in their state of rest, and y' , y'' , &c. their distances after the lapse of the same time t . Then wil. $x' - x = dx$, $x'' - x' = dx'$, $x''' - x'' = dx''$, &c. be the thicknesses (supposed infinitely small) of these strata, or the spaces occupied by them (taking the area of the section for unity) in their quiescent state, and $y' - y = dy$, $y'' - y' = dy'$, $y''' - y'' = dy''$, &c. the same in their state of

Sound.

motion. Now as these strata were in contact at the origin of the motion, and are held together by the pressure of the surrounding fluid, they will remain in contact, and advance and recede along the pipe as one mass, only the space they will occupy at different points of their motion will be variable, according to the degree of condensation or dilatation they may have undergone in virtue of their motion itself. If, for instance, at any moment the hinder of them dy be in the act of urging forward the next dy' , it will be condensed; if retreating, rarefied in comparison with the state of the preceding one dy'' .

Part I.

Now any stratum of molecules dy' interjacent between two others dy and dy'' can only undergo a change in its velocity when urged by some force, and the only force which can urge it is the difference of pressures it may experience on its two faces by the difference (if any) of the elasticities of the adjacent strata dy'' and dy . If we can estimate this, the laws of Dynamics will enable us to express the consequent change of motion. To this end, then, let the elasticity of the air in its quiescent state be represented by E , which is a given quantity, and is measured by the weight of a column of mercury sustained by it, or by the length of a homogeneous column of air of the same density, whose weight shall suffice to keep it so compressed, or be equal to that of the column of mercury in the barometer. Then, since the elasticity of air is inversely as the space it occupies, (*cæteris paribus*), the elasticity of the air when occupying the stratum dx : its elasticity when occupying

52.

Expression of the acting forces.

dy : dy : dx , and therefore the elasticity when occupying the space $dy = E \cdot \frac{dx}{dy}$. Similarly the elasticities

of the air occupying dy' and dy'' will be represented by $E \cdot \frac{dx'}{dy'}$ and $E \cdot \frac{dx''}{dy''}$. Hence the plane separating

the strata dy and dy' will be pressed *forward* by the elasticity $E \cdot \frac{dx}{dy}$, and *backward* by $E \cdot \frac{dx'}{dy'}$. So that

it will, on the whole, be urged forward by $-E \left(\frac{dx'}{dy'} - \frac{dx}{dy} \right)$ that is, by $-E d \frac{dx}{dy}$, the differentials being all on

the supposition of t , the time being constant, and x and y only variable. Now, if we denote by H the length of a homogeneous column of air necessary to counterbalance the elasticity of the quiescent air, and by D its density, we have $HD = \text{its weight} = \text{the elasticity } E$, and $dx' \cdot D = \text{the weight of the stratum } dx'$, which, substituting

for D its value $\frac{E}{H}$, becomes $dx' \cdot \frac{E}{H}$. Thus, then, the moving force $-E d \frac{dx}{dy}$ is exerted in urging for-

ward a weight $= dx' \cdot \frac{E}{H}$, and is therefore equivalent to an accelerating force

$$-H \cdot \frac{d \left(\frac{dx}{dy} \right)}{dx'} = +H \cdot \left(\frac{dx}{dy} \right)^2 \cdot \frac{d^2 y}{dx^2},$$

regarding dx as constant, or all the strata dx , dx' , dx'' , &c. as originally equal.

Now the distance of the mass thus urged from the fixed point A , at the expiration of the time t , is y' . Hence, if we regard only the motion of the particle dy' (or which comes to the same) of dy , which is in contact with it, we have by Dynamics

53.

Equation deduced.

$$\left(\frac{dy}{dx} \right)^2 \cdot \frac{d^2 y}{dt^2} = 2gH \cdot \frac{d^2 y}{dx^2}, \quad (a)$$

where $2g = 9^{\text{met}}.8088 = 32.180$ British standard feet and gravity, for the unit of accelerating force, and in which equation t is expressed in mean solar seconds; and all linear quantities, such as H , x , y , in metres or feet, according as we take the metre or foot as the unit of linear measure.

This is, in fact, an equation of partial differentials, y being at once a function both of x the original distance of the stratum dx from the origin of the motion, and of t the time elapsed. In its present form, simple as it appears, it is altogether intractable and incapable of integration. In fact, it embraces a class of dynamical problems of very great complexity; for it is evident that since no hypothesis has been made in any way limiting the extent of the excursions of the original or subsequent strata from their points of quiescence, this equation must contain the general expression of all possible motions of elastic fluids in narrow pipes, whether great, as when urged by pistons or driven by bellows, or small, as are the tremors which cause Sound. In the theory of Sound we suppose the agitations of each molecule so minute as not to move it sensibly from its point of rest. Experience confirms this. Sounds transmitted through a smoky or dusty atmosphere cause no visible motion in the smoke or floating dust, unless the source of Sound be so near as to produce a *wind*, which, however, is always insensible at a very moderate distance.

54.

Limitation in the case of Sound.

If we introduce this condition, the equation (a) admits of integration; for the whole amount of motion of each molecule being extremely minute, their differences for consecutive molecules, or the amount of the rarefactions

55.

Simplification of the final equation.

and condensations undergone, must be much more so. Hence the value of $\frac{dy}{dx}$, which expresses the ratio of

the condensation of the stratum dy in motion and in rest, may be regarded as equal to unity, and the equation becomes simply,

$$\frac{d^2 y}{dt^2} = a^2 \frac{d^2 y}{dx^2}, \quad \text{where } a = \sqrt{2gH}; \quad (b)$$

Sound. which is the equation of Sound regarded as propagated in one dimension, that of length, only ; or, as prevented from spreading laterally by a pipe.

56. The complete integral of this equation is well known to be

$$y = F(x + at) + f(x - at), \quad (c)$$

where F and f denote arbitrary functions of the quantities within the parenthesis, and which must be determined by a consideration of the initial state of the fluid, or by the nature of the motion originally communicated to its molecules.

57. Let us then suppose, that, at the commencement of the motion, we have impressed on each section of the fluid, along its whole extent, any arbitrary velocities and condensations, by any means whatever, so as to comprehend in our investigation all possible varieties of initial motion, whether expressible by regular analytical functions, or depending on no regular law whatever. It is manifest that these conditions will be expressed by assuming arbitrary functions of x , such as $\phi(x)$ and $\psi(x)$ for the *initial* values of the two partial differentials

$\frac{dy}{dt}$ and $\frac{dy}{dx}$, whereof the former represents in all cases the velocity (v) of a particle which would be at the

distance x from the origin of the coordinates in the state of equilibrium, and the latter the linear extent (e) of that particle compared with its original extent, to which its density and elasticity are reciprocally proportional. Now, differentiating (c) we get for the general values of v and e

$$v = \frac{dy}{dt} = a \{ F'(x + at) - f'(x - at) \}; \quad (d)$$

$$e = \frac{dy}{dx} = F'(x + at) + f'(x - at); \quad (e)$$

consequently their initial values, making $t = 0$, will be

$$\phi(x) = a \cdot \{ F'(x) - f'(x) \}$$

$$\psi(x) = F'(x) + f'(x),$$

whence we get immediately

$$F'(x) = \frac{1}{2a} \{ a\psi(x) + \phi(x) \}; \quad f'(x) = \frac{1}{2a} \{ a\psi(x) - \phi(x) \}; \quad (f)$$

and multiplying by dx and integrating

$$F(x) = \frac{1}{2a} \int \{ a\psi(x) + \phi(x) \} dx; \quad f(x) = \frac{1}{2a} \int \{ a\psi(x) - \phi(x) \} dx;$$

and thus the forms of the functions F and f become known when those of ϕ and ψ are given.

58. The question of the propagation of Sound, however, does not require us to concern ourselves with these functions, as a knowledge of the actual velocity and density of any molecule at any instant is sufficient for our purpose. Substituting then in (d) and (e) for F' and f' , the forms corresponding in ϕ and ψ , we get

$$v = \frac{dy}{dt} = \frac{a}{2} \{ \psi(x + at) - \psi(x - at) \} + \frac{1}{2} \{ \phi(x + at) + \phi(x - at) \}; \quad (g)$$

$$e = \frac{dy}{dx} = \frac{1}{2} \{ \psi(x + at) + \psi(x - at) \} + \frac{1}{2a} \{ \phi(x + at) - \phi(x - at) \}; \quad (h)$$

or, as it may also be written,

$$v = \frac{1}{2} \{ \phi(x + at) + a\psi(x + at) \} + \frac{1}{2} \{ \phi(x - at) - a\psi(x - at) \}; \quad (i)$$

$$e = \frac{1}{2a} \{ \phi(x + at) + a\psi(x + at) \} - \frac{1}{2a} \{ \phi(x - at) - a\psi(x - at) \}. \quad (j)$$

59. These are essentially the same expressions with those given by Euler in his Paper on the Propagation of Sound, in the *Berlin Memoirs* for 1759, and by Poisson in his elaborate Memoir on the Motion of Elastic Fluids in Pipes, and on the theory of Wind Instruments, and they comprise the whole theory of the linear propagation of Sound. But before we proceed to the interpretation of their meaning in particular cases, we have a few remarks to make on their general form.

60. And, first, it is evident, that since the variable quantity x enters into all the terms both of v and e under the functional characteristics, these quantities, regarded as functions of t , are modified essentially by the value of x , which may be regarded as a parameter, or constant element in the composition of the functions expressing the nature of the motion of any assigned molecule. If only $x + at$, or only $x - at$, separately entered under the

characteristics, since $x + at = a \left(t + \frac{x}{a} \right)$ and $x - at = -a \left(t - \frac{x}{a} \right)$ the variation of x would only vary the *origin* of t ; and the motions of all the successive molecules would be performed according to the same laws, only commencing at a different epoch for each molecule; but, as both these quantities are involved, that will not be universally the case. Consequently, in general, it appears that the undulation, or pulse, as it is propagated onward, becomes modified essentially in its quality by the distance it has passed over, it is no longer

Remarks on these expressions.

Sound. *the same sound, i. e.* not identical with what would be produced by shifting the initial stratum forward. Its velocity, intensity, and pitch, it is true, will remain (as we shall see) unaltered; but its *quality*, its mode of action on the ear, (which must be differently affected by changes in the nature of the impulse made on it,) will undergo a change. This establishes an essential difference between a Sound wave and such a wave as we took for an illustration in Art. 45, where every point was in succession agitated by the same *identical* motion.

Part I.

Consequently every theory of Sound in which it is assumed that the several particles in a sounding column are all in succession agitated alike, is defective. This is the case with Newton's doctrine of the propagation of Sound as delivered in the 47th proposition of the 2nd book of the *Principia*, and, were there no other objection against it, would suffice to vitiate the whole. This, and other unsatisfactory points in the celebrated theory alluded to, were first distinctly perceived and pointed out by Lagrange, in the first volume of the *Turin Miscellanies*, and an exact and rigorous investigation substituted in its place, in which the sounding column is regarded as consisting of a series of finite, insulated particles, mutually repelling each other; a mode of conception which leads, by a very complicated analysis, to the same results with that above stated, but which has the advantage of setting in a distinct light the internal mechanism, if we may so term it, by which Sound is propagated, and to which we therefore willingly refer the reader.

61
Inaccuracies in Newton's theory of Sound.

Moreover, since by differentiating the equation (d) we get

62.

$$\frac{d^2 y}{dt^2} = a^2 \{ F''(x + at) + f''(x - at) \},$$

this will be proportional to the accelerating force acting on the molecule. It is therefore by no means universally proportional to $y - x$, the distance of the molecule from its point of rest; and therefore another assumption on which the Newtonian doctrine of Sound rests, *viz.* that the motion of each molecule necessarily follows the law of a vibrating pendulum, is equally destitute of foundation. In fact, Cramer had shown, before the examination of Lagrange, that any other law of molecular motion might be substituted in Newton's enunciation of his general proposition, and the demonstration would be equally conclusive, and the resulting velocity of Sound the same.

Let us now descend more into particulars; and, first, let us suppose the initial state of the fluid to consist in a general repose of the whole of an infinitely extended column, except a very small portion at A the origin of the coordinates, which we will suppose agitated with any arbitrary motion. This is, in fact, the simplest case of the production of Sound; the initial disturbance of the air being always confined within extremely small limits compared to the distances to which the Sound is propagated. Let us then conceive the initial disturbance to take place over a minute length $2a$ of the column, whose middle we will suppose to be in the origin of the x . This amounts to supposing $\phi(x) = 0$, and $\psi(x) = 1$, for every value of x not comprised within the limits $x = -a$ and $x = +a$, admitting them to have any arbitrary values between these limits.

63.
Case of a very small initial disturbance.

If we suppose now t to be less than $\frac{x-a}{a}$, and regard at first what happens only on the positive side of the origin of the x , since $t < \frac{x-a}{a}$ we have $at < x - a$, and therefore $x - at > +a$, and, *à fortiori*,

64.

$x + at > +a$, consequently for all values of t less than $\frac{x-a}{a}$ we have $\phi(x - at) = 0$, $\phi(x + at) = 0$;

Propagation of a single initial disturbance, beginning and terminating suddenly.

$\psi(x - at) = \psi(x + at) = 1$; and therefore for all values of t less than $\frac{x-a}{a}$ we have $v = 0$, and $e = 1$.

Consequently the molecule at the distance x from the origin of the coordinates, will remain at rest and undensified, or expanded, so long as t remains less than $\frac{x-a}{a}$; that is, for a time proportional to the distance

from the nearest point of the initial disturbance. But the moment t has attained this limit, $\phi(x \pm at)$ will have finite values, and $\psi(x \pm at)$ values differing from unity, and v and e will consequently have such. The particle

then will begin to move, and to undergo a change of density, and will continue to do so till $t = \frac{x+a}{a}$. At

this limit we have $x - at = -a$, $x + at = 2at - a = 2(x + a) - a = 2x + a$, and consequently $x + at > +a$. Hence at this limit we have again $\phi(x - at) = \phi(x + at) = 0$, and $\psi(x - at) = \psi(x + at) = 1$, and the motion and condensation of the particle will cease; and will not be resumed afterwards, because the

supposition $t > \frac{x+a}{a}$ gives $x - at < -a$, and $x + at > 2x + a$. and, *à fortiori*, $> +a$, so that the functions retain their values 0 and 1 from this moment for ever.

Thus we see that the molecule distant by x from the origin of the coordinates will remain at rest for a certain time $t = \frac{x-a}{a}$, will then begin to move, and continue moving, during a time equal to $\frac{x+a}{a} - \frac{x-a}{a}$

65.
Velocity of Sound uniform.

$= \frac{2a}{a}$, or till $t = \frac{x+a}{a}$, and will then return to a state of permanent rest. A similar reasoning will apply for negative values of x . Hence if we consider any two molecules at distances x, x' from A, we see that

Sound. the more distant will commence and terminate its motion later than the nearer, by an interval of time $\frac{x' - x}{a}$. This then is the time required for the propagation of the impulse, or Sound, over the intermediate space $x' - x$, and being proportional to that space, the velocity of propagation must be uniform, and must be represented by the quantity $a \left(= \frac{x' - x}{t' - t} = \frac{\text{space}}{\text{time}} \right)$. Hence it follows that *the velocity of Sound is uniform,—is independent of the nature, extent, and intensity of the primitive disturbance, (for the arbitrary functions do not enter it,) and is expressed by the quantity we have called a, that is $\sqrt{2gH}$.*

Its expression.

66. Let us reduce this to numbers, in order to compare theory with observation. To this end, if we call Δ the density of mercury, h the height of the mercury in a barometer exposed to the same pressure as the sounding column, and D the density of the air in it, we have for the height of a homogeneous column of such air capable of counterbalancing the elasticity of the sounding fluid, the following value

$$H = h \cdot \frac{\Delta}{D};$$

and, calling V the velocity of Sound, we should have

$$V = \sqrt{2g h \cdot \frac{\Delta}{D}}.$$

Now, at the freezing temperature, and in a mean state of barometric pressure, we have, according to Biot,

$h = 0^{\text{met.}}.76$; $2g = 9^{\text{met.}}.8088$; and $\frac{\Delta}{D} = 10463$; so that we obtain, by executing the numerical operations,

$$V = 279^{\text{met.}}.29 = 916^{\text{feet.}}.322.$$

67. The actual value of V obtained by experiment is, as we have seen, 1089.42. The difference, 173 feet, is nearly one-sixth of the whole amount; a discrepancy far too great to be attributed to any inaccuracy in the determination of the data, which are all of the utmost precision. It is evident, then, that there is something radically insufficient in the theory, as above delivered; and, accordingly, Geometers for a long while endeavoured to account for it on various suppositions. Newton, who, by a singularly happy coincidence, which certainly deserves to be called a divination, had, from a theory totally inapplicable in all its points, elicited the correct expression $\sqrt{2gH}$ above demonstrated, for the velocity of Sound, and who immediately encountered this difficulty on deducing its numerical value, endeavours to account for the deficient 173 feet by supposing the molecules of the air to be actual spherical solids of a certain diameter, ($\frac{1}{916}$ of the interval between them,) and that the Sound is propagated through them *instantly*. It is needless to comment on this explanation. Lagrange treats the whole matter lightly, and seems inclined to attribute the deviation of fact from theory to erroneous data; in other words, dissembling the difficulty, which Euler, on the contrary, broadly acknowledged; and

Various attempts to account for it.

considered that it might possibly arise from an incorrectness of analysis, in assuming the factor $\left(\frac{dy}{dx}\right)^2 = 1$ in

the equation (a) Art. 53, previous to integration. The true explanation was reserved for the sagacity of Laplace. But before we state it, it will be necessary to consider what will be the effect of variations of temperature and pressure on the velocity, according to the principles already laid down, and the formula arrived at.

68. With regard to an increase of pressure, its effect is to increase the density of the air; but since at the same time it increases its elasticity, and in exactly the same ratio; the mass to be moved, and the moving force, are increased alike, and therefore the accelerating force remains unaltered. The velocity, therefore, ought to undergo no change by this alteration. On the other hand, an increase of temperature, under a constant pressure, tends to dilate the air, and either renders it more elastic in the same space, or more rare with the same elasticity. Hence, on a variation of temperature, the moving force remains unaltered, while the mass moved decreases, and therefore an acceleration in all the resulting motions must arise. The velocity of Sound then ought to be greater in warm than in cold air, *ceteris paribus*. These two conclusions are both amply confirmed by experiment. They agree too with the formula above stated; for, if we denote by (h) the mean height of the mercury in the barometer ($0^{\text{met.}}.76$), and by (D) the density of air under this pressure at the freezing temperature, since, by the experiments of Gay Lussac, air expands 0.00375 of its volume by every degree *centigrade* of increase of temper-

Effect of variations of temperature and pressure on the velocity of Sound.

ature, its density under the pressure (h) at any other temperature $+ \tau^{\circ}$ (centig.) will be $\frac{(D)}{1 + \tau \cdot 0.00375}$, and

under the pressure h it will be $\frac{h}{(h)} \times \frac{(D)}{1 + \tau \cdot 0.00375} = D$; consequently the expression (Art. 66) for the velocity becomes

$$V = \sqrt{2g(h) \cdot \frac{\Delta}{(D)} \times (1 + \tau \cdot 0.00375)}.$$

Now, if we call (V) the velocity under the mean pressure (h), and at the freezing point, this gives

$$(V) = \sqrt{2g(h) \cdot \frac{\Delta}{(D)}}.$$

Sound. and therefore $V = (V) \cdot \sqrt{1 + \tau \cdot 0.00375} = (V) \{1 + \tau \cdot 0.001875\}$,
 or if τ be expressed in degrees of Fahrenheit's scale,

$$V = (V) \{1 + \tau \cdot 0.001042\};$$

Part I.

which shows, first, that the velocity is independent of the pressure, since h is not contained in its expression; and that, secondly, τ increases by very nearly the 0.001875 part of its whole quantity for every degree centigrade, or $\frac{1}{5} \times 0.001875 = 0.001042$ for every degree Fahrenheit above the freezing point, that is in feet 1.136, (see Art. 17.) and decreases by the same quantity for each degree below freezing.

The law of Mariotte, which makes the elastic force of the air proportional to its density, and which has been employed in estimating the elasticity with which each molecule of the aerial column resists condensation, and transmits it to its neighbour, assumes that the temperature of the whole mass of air is alike, and undergoes no change in the act of condensation, and is therefore only true of masses of air which, after compression, are of the same temperature as before. But it is an ascertained fact, that air and all elastic gaseous fluids give out heat in the act of compression, *i. e.* actually become *hotter*, a part of their latent heat being developed, and acting to raise their temperature. This is rendered evident in the violent and sudden condensation of air by a tight-fitting piston in a cylinder closed at the end. The cylinder, if of metal, becomes strongly heated; and if a piece of tinder be enclosed, on withdrawing the piston it is found to have taken fire; thus proving that a heat, not merely trifling, but actually that of ignition, has been excited, of at least 1000° of Fahrenheit's scale. Now when we consider how small the mass of air in such an experiment is, compared with that of the including vessel, which rapidly carries off the heat generated, it is evident that if air by any cause could be compressed to the same degree without contact of any other body, a very enormous heat would be generated in it. It would, therefore, resist the pressure much more than if cold; and, consequently, would require a much more powerful force to bring it into that state of condensation than, according to Mariotte's law, would be necessary.

69.
Laplace's explanation of the anomaly above mentioned.

Heat developed in air in act of compression

Air, then, when suddenly condensed, and out of contact with conducting bodies, resists pressure more (*i. e.* requires a greater force to condense it equally) than when slowly condensed, and the heat developed carried off by the contact of massive bodies of its original temperature. In other words, it is under such circumstances more elastic, and our analytical expression for its elasticity must be modified accordingly. In fact, the condensation of the aerial molecules in the production of Sound is precisely performed under the circumstances most favourable to give this cause its full influence; the condensations being so momentary that there is no time for any heat to escape by radiation; and the condensed air being in contact with nothing but air, *differing infinitesimally from its own temperature*; so that conduction is out of the question. Let us see now how this will affect the matter in hand.

70.
Influence of this cause in the propagation of Sound.

It was assumed in Art. 35, that the elasticity of the air occupying the space dx , or (E) : its elasticity when occupying $dy :: dy : dx$. But, in fact, the varied temperature being taken into account, the latter ratio should have stood $:: dy (1 + a\tau) : dx (1 + a\tau')$, where a denotes the coefficient 0.00375, and τ and τ' the original and altered temperatures in centigrade degrees. Hence in place of $E \cdot \frac{dx}{dy}$ we must have $E \cdot \frac{dx}{dy} \cdot \frac{1 + a\tau'}{1 + a\tau}$,

71.
Modification of the analysis and formulæ required by it.

that is, $E \cdot \frac{dx}{dy} \{1 + a(\tau' - \tau)\}$, for the elasticity of the molecule of air when occupying the space dy ,

because, the condensations being all along supposed exceedingly small, τ' differs from τ only by a quantity of the same order as the condensations; so that $(\tau' - \tau)^2$ and its higher powers may be neglected.

Now, whatever may be the law according to which the temperature of a mass of air is increased by a sudden diminution of its volume, it is obvious that for very small condensations, such as those considered in the theory of Sound, the rise of temperature will be proportional to the increase of density; because, the quantity of latent heat having sustained only a very minute diminution, by a given extremely small condensation, a repetition of the same condensation will develop a quantity of heat falling short of the first only by a quantity of the second order; so that, neglecting such quantities, double the condensation will develop double the heat, and so

72.
Analysis.

in proportion. Hence we must have $\tau' - \tau = k \left\{1 - \frac{dy}{dx}\right\}$ where k is a constant coefficient, whose magnitude may become known either by direct experiment, or by the very phenomena under consideration. Substituting this for $\tau' - \tau$, we get, for the elasticity of the condensed molecule,

$$E \cdot \frac{dx}{dy} \left\{1 + ka \left(1 - \frac{dy}{dx}\right)\right\} = E(1 + ka) \cdot \frac{dx}{dy} - kaE.$$

And the difference of elasticities on either side of the plane separating the molecules dy and dy' , instead of being, as in (Art. 35.) $-E \cdot d \frac{dx}{dy}$, will be now represented by $-d \left\{E(1 + ka) \frac{dx}{dy} - kaE\right\}$, that is,

$$\text{by} \quad -E(1 + ka) \cdot d \frac{dx}{dy}.$$

This differs from the expression originally obtained only by the constant factor $(1 + ka)$. Without, therefore, going again through all the foregoing analysis, we see at once that the general equations of Sound will be precisely as before, writing only $(1 + ka) \cdot H$ for H throughout; and, therefore, if instead of putting, as before,

73

Sound. $a = \sqrt{2gH}$, we put $a = \sqrt{2gH(1+ka)} = \sqrt{2gH \cdot K}$; when $K = 1 + ka$ the equation (a) will become

$$\left(\frac{dy}{dx}\right)^2 \cdot \frac{d^2y}{dt^2} = a^2 \cdot \frac{d^2y}{dx^2};$$

Final and exact expression of the equation and velocity of Sound.

and all the other equations will remain unaltered, and the velocity of Sound on this new hypothesis will be expressed by the new value ascribed to a , that is, by

$$V = \sqrt{2gH(1+ak)}$$

$$= \sqrt{2g(h) \cdot \frac{\Delta}{(D)} K(1+ak)}$$

where $a = 0.00375$.

74. The actual numerical value of the constant coefficient K may be determined, as we have before said, in two ways; either by direct experiment on the increase of temperature developed in a given volume of air by a given condensation, or by a comparison of the formula to which we have arrived with the known velocity of Sound. As we have already observed, however, the circumstances under which Sound is propagated are far more favourable to the free and full production of the whole effect of the cause in question than those of any experiments in close vessels. We must not, therefore, be surprised, if the value of K as derived from such experiments should differ materially from its value deduced from the velocity of Sound; nor *vice versa*, if the observed velocity of Sound should differ materially from that obtained by calculation, from an experimental value of K . It is sufficient, in a philosophic point of view, to have pointed out a really existing cause, a *vera causa*, which *must* act to increase the velocity, and is fully adequate to do so to the extent observed.

75. We have seen that the numerical value of V neglecting K is equal to 916.322 feet. The observed value on the other hand, is 1089.42. Hence we have the following equation for determining K and k ,

$$1089.42 = 916.322 \times \sqrt{1 + k \cdot 0.00375} = 916.322 \times \sqrt{K},$$

whence we obtain

$$K = \left(\frac{1089.42}{916.32}\right)^2 = 1.4132,$$

and

$$k = \frac{1}{0.00375} \left\{ \left(\frac{1089.42}{916.32}\right)^2 - 1 \right\} = 110.26.$$

Difficulty of its direct determination. The actual amount of heat given out by a given amount of condensation is not an element very easily or exactly determinable by direct experiment with thermometers. If a common mercurial thermometer be enclosed in a receiver, and the air suddenly compressed, the thermometer, it is true, rises; but the amount of its rise is evidently far inferior to the actual increase of temperature; for, first, its mass is enormously greater than that of the air *immediately* in contact with it; secondly, it is brought into contact successively with an unknown, and, no doubt, a variable quantity in different experiments, by the effect of circulation; thirdly, the vessel used carries off by far the greater part of the heat, and one which we have no means of estimating. It is accordingly found that by increasing the sensibility of the thermometer, by extending its surface compared to its mass, higher and higher degrees of temperature are indicated for the same condensation; and highest of all when the delicate pyrometer of Breguet is used, which consists of two extremely thin strips of platina and palladium soldered together over their whole surface, and coiled up in a spiral, which twists and untwists by the different expansions of the metals constituting its inner and outer face. Still, however, though almost all surface, the materials of which this instrument consists are so infinitely denser than air, that its indications must fall far short of the truth.

76. Another very ingenious method has been practised by Messrs. Clement and Desormes. (*Journal de Physique*, November, 1819, p. 334.) Suppose we have any quantity of air enclosed in a receiver communicating, first, with an air-pump, by a valvular orifice, (A); second, with the upper part of a barometer tube containing mercury, whose height therefore measures the elasticity of the air in the receiver by its depression below the barometric level of the external atmosphere; thirdly, with the external air, by a stopcock, or valve, (B,) so large that the pressure within may be instantaneously restored to an equilibrium with that without, on opening it. Let the whole apparatus be at the temperature of the atmosphere, (τ) and suppose the valve (B) open, then will the internal elasticity, or pressure, (P .) be equal to that without, and also the density (D .) Close the valve B, and open A, and, by means of the air-pump, exhaust a small portion of the air; and, again closing the valve A, let the apparatus remain at rest till the whole has attained the temperature τ of the atmosphere. In this state let the internal pressure be observed by the barometer, which call P' ; and D' , the density, will, of course, be equal to

$D \cdot \frac{P'}{P}$, and is therefore known. Now suddenly open the valve B. The external air will rush in and restore

the equilibrium. The moment this is done (which will be known by the cessation of the inward current) let the valve B be closed. It will then be found that the internal temperature is raised by the condensation thus effected, and has become τ' ; and the increase of temperature $\tau' - \tau$ may be measured by a delicate thermometer, and that with the more precision the greater the capacity of the receiver. But it will be much more exactly measured by the following process, which, in fact, amounts to making the receiver itself an *air thermometer*. At the moment of closing the valve the internal pressure is, of course, P . But as the air cools, its elasticity diminishes, and, being cut off from a fresh supply from without, the mercury will rise in the barometer tube till the whole

Sound. of the heat evolved is dissipated. Let the internal pressure, then, be again observed when this state is attained, and call it P'' , then will the corresponding density, or D'' , be equal to $D \cdot \frac{P''}{P}$. It is required from these data

Part I.

(P, P', P'' , being given by observation) to deduce the value of $\tau' - \tau$ and the coefficient k .

Now, this is easy; for, first, since in the final state of the receiver the density is D'' sustaining a pressure P'' at a temperature τ ; therefore, the same quantity of air in the same space, raised to the temperature τ' , would sustain a pressure $P' \times \{1 + a(\tau' - \tau)\}$ where $a = 0.00375$, τ being in centigrade degrees, its density remaining D'' . But at the moment of closing the valve A, the temperature was τ' , and the pressure simply P , we have, therefore,

77.
Analysis of
this experi-
ment.

$$P = P'' \{1 + a(\tau' - \tau)\}, \text{ whence } \tau' - \tau = \frac{P - P''}{a P''}.$$

Now, secondly, this is the elevation of temperature due to the sudden transition of the air from the density D' to the density D'' , by the introduction of that portion of external air which rushed in on opening the valve. Calling l the capacity of the receiver, $l \times D' = D'$ expresses the quantity of air in it before the valve was opened, and $l \times D''$ or D'' the quantity after, so that $D'' - D'$ expresses the quantity of air admitted. Its density before admission being D , and afterwards D'' , it had undergone a dilatation equal to $1 - \frac{D''}{D}$, and therefore its tem-

perature had diminished by an amount represented by $k \left(1 - \frac{D''}{D}\right)$. On the other hand, the quantity of air in the receiver before opening the valve was $l \times D' = D'$, and this quantity having changed its density suddenly from D' to D'' , must have undergone an elevation of temperature represented by $k \left(1 - \frac{D'}{D''}\right)$. These two masses of air, the one cooled by dilatation, the other heated by condensation, became suddenly mixed, and therefore must have undergone a mean rise of temperature = $\frac{\text{sum of products of masses and changes of temp.}}{\text{sum of masses}}$ and, consequently,

$$\begin{aligned} \text{mean elevation of temperature} &= \frac{\left\{ D' \left(1 - \frac{D'}{D''}\right) - (D'' - D') \left(1 - \frac{D''}{D}\right) \right\} k}{D' + D'' - D' = D''} \\ &= \frac{D'' - D'}{D''} \left\{ \frac{D'}{D''} - \frac{D - D''}{D} \right\} k = \frac{D'' - D'}{D''} \left\{ \frac{D'}{D''} + \frac{D''}{D} - 1 \right\} k, \\ &= \left(1 - \frac{D'}{D''}\right) \left\{ \frac{D''}{D} - \left(1 - \frac{D'}{D''}\right) \right\} k = k \left\{ \frac{D''}{D} \left(1 - \frac{D'}{D''}\right) - \left(1 - \frac{D'}{D''}\right)^2 \right\}. \end{aligned}$$

But we have $\frac{D''}{D} = \frac{P''}{P}$, and $\frac{D'}{D} = \frac{P'}{P}$, so that $\frac{D'}{D''} = \frac{P'}{P''}$,

and therefore substituting, we find for the value of the above expression, or $\tau' - \tau$,

$$k \left\{ \frac{P''}{P} \left(1 - \frac{P'}{P''}\right) - \left(1 - \frac{P'}{P''}\right)^2 \right\} = k \cdot \frac{P'' - P'}{P} \left\{ 1 - \frac{P(P'' - P')}{P''^2} \right\}.$$

If we suppose the changes of pressure sufficiently small to allow of their squares being neglected, the value of $\tau' - \tau$ is reduced to $k \cdot \frac{P''}{P} \left(1 - \frac{P'}{P''}\right) = k \cdot \frac{P'' - P'}{P}$. Equating this to $\frac{P - P''}{a P''}$, the previously determined value of $\tau' - \tau$, we get

Value of k
expressed.

$$k = \frac{1}{a} \cdot \frac{P - P''}{P'' - P'} \cdot \frac{P}{P''}; \quad k a = \frac{P(P - P'')}{P''(P'' - P')}.$$

In an experiment of Messrs. Clement and Desormes, on which M. Poisson has grounded his computation of the theoretical velocity of Sound, the values of P, P', P'' were

78.
Numerical
computa-
tion.

$$P = 0^m.7665; \quad P - P' = 0^m.01381; \quad P - P'' = 0^m.00361;$$

and, consequently,

$$P'' - P' = 0.01020,$$

which gives, by the approximate formula,

$$k a = 0.3492, \quad \text{and } 1 + k a = 1.3492;$$

whence the velocity, at a mean pressure and freezing temperature, comes out

$$916^{\text{feet}}.322 \cdot \sqrt{1.3492} = 1064.35,$$

which falls short of the actually observed velocity only by about 25 feet. If the rigorous value of $k a$ be employed, the deficiency is rather less, the velocity coming out 1066.2. In this experiment, the time occupied

Sound. by the intromission of the air was about half a second; the whole elevation of temperature, computed from the Part I.

formula $\tau' - \tau = \frac{P - P''}{a P''}$, must have been $1^{\circ} \cdot 321$ centig. ($= 2^{\circ} \cdot 378$ Fahr.) M. Poisson has shown (*Annales de Chim.* xxiii. 1823, p. 11) that an absorption of $\frac{1}{2}$ of a degree (cent.) by the receiver, which might very well happen, would completely reconcile the observed and theoretical velocities. Laplace, calculating on the experiments of Messrs. Welter and Gay Lussac, has, since, obtained a still nearer approximation to the theoretical velocity, the difference amounting only to about 3 metres. In inquiries of such delicacy, and where the effects of minute errors of experiment become so much magnified, it seems hardly candid to desire a more perfect coincidence.

79. Laplace, guided by peculiar theoretical considerations respecting the constitution of gaseous fluids, has been induced to put the foregoing expression for the velocity of Sound under a somewhat different form. Let K denote the ratio of the specific heat of air under a constant pressure to its specific heat if retained at a constant density; that is, a fraction whose numerator is the quantity of heat requisite to raise a given mass of air 1° in temperature under a constant pressure, (its volume being permitted to increase,) and whose denominator is the quantity necessary to raise it 1° in a constant volume, or when so confined as not to dilate. Then will the velocity of Sound be

$$V = \sqrt{2g(h) \cdot \frac{\Delta}{(D)} (1 + a\tau) \cdot K}.$$

To show this, let Q and q be the quantities of heat above mentioned. It is evident, first, that when forcibly prevented from expanding, and thereby absorbing heat and rendering it latent, a less quantity of heat will suffice to raise the temperature of a given mass of air any given quantity, as 1° , than if unconfined. In fact, suppose it heated 1° , and allowed meanwhile to dilate, so that the temperature of the dilated air shall be 1° above its primitive state, then, if compressed back into its original volume, the whole quantity of heat developed by the condensation will be employed in raising the temperature still higher. If then the quantity Q of caloric raise the temperature 1° under a given pressure, it will raise it more than 1° when confined to a given volume, by the whole amount of temperature due to a compression equal to its dilatation in the former case. Suppose the initial temperature freezing, then if $a = 0.00375$, an increase of temperature of 1° cent. will produce, under a constant pressure, a dilatation = a, and the volume from 1 will become $1 + a$. Let the air so dilated and raised in temperature be compressed back to its former volume, then will its temperature be further increased by ka , k denoting as before; so that the quantity of caloric Q will have ultimately produced a rise of temperature

$= 1 + ka$, under a constant volume; and therefore a quantity $= \frac{Q}{1 + ka}$ only would be required to raise it 1° . Hence $q = \frac{Q}{1 + ka}$, and $1 + ka = \frac{Q}{q} = K$. This demonstration assumes, as an axiom, that the

temperature produced by the introduction of the same quantity of caloric is the same, whether it be introduced into air confined in a given space, or into air allowed to expand freely, and then forcibly compressed back; which it evidently is, since the heat given out by the compression must of necessity exactly equal that absorbed and rendered latent in the act of expansion.

§ II. Of the Linear Propagation of Sound in Gases and Vapours.

80. The analysis by which we have in the foregoing articles determined the laws and velocity of the propagation of Sound in air, applies equally, *mutatis mutandis*, to its propagation in all permanently elastic fluids, and in vapours, in so far as their properties are the same as those of gases. The formula so often referred to then

$$V = \sqrt{2g(h) \cdot \frac{\Delta}{(D)} \cdot K \cdot (1 + a\tau)}$$

expresses the velocity of Sound in all such media, provided for (D) we write instead of the density of atmospheric air that of the gas at the freezing temperature, and under the mean pressure (h). In the case of vapours, we must suppose in calculating the value of (D) that they follow the law of gases in their condensation, and that no portion of them undergoes a change of state to a liquid, by reduction to the standard temperature and pressure. Suppose, then, the specific gravity of atmospheric air to be denoted by s, and that of any gas or vapour under the same temperature and pressure by s'; then if V and V' be the velocities of Sound in air, and in the gas or vapour, we have

$$V = \sqrt{2g(h) \cdot \frac{\Delta}{s} K (1 + a\tau)}, \quad V' = \sqrt{2g(h) \cdot \frac{\Delta}{s'} K (1 + a\tau)},$$

because (see PNEUMATICS, HEAT) the law of dilatation, or the value of a, is alike in all. Consequently, we have

$$\frac{V'}{V} = \sqrt{\frac{s}{s'} \cdot \frac{K'}{K}},$$

Sound.

or

$$V' : V :: \sqrt{\frac{K'}{s'}} : \sqrt{\frac{K}{s}};$$

Part I.

whence the ratios of $s' : s$ and of $K' : K$ being known, the ratio of the velocities is also known, being, *cæteris paribus*, in the inverse subduplicate ratio of the specific gravities.

To compare this with experiment directly is impracticable, as no column of any gas but atmospheric air can be obtained of sufficient length and purity to determine the velocity of Sound in it by direct measure. Indirectly, however, the comparison may be performed by comparing the Sounds of one and the same organ-pipe, filled with the gases to be compared, successively, or by other means of a similar kind, of which more hereafter. (See INDEX, under the heads *Gases, Vapours, Sounds of Pipes.*)

The following Table exhibits the Velocities of Sound, as deduced from theory, and compared with experiments instituted by M. Van Rees, in conjunction with Messrs. Frameyer and Moll.

81.
How com-
parable with
experiment.

82.
Velocity of
Sound in
various
media.

Gas, or Vapour.	Velocity of Sound, reduced to 0° R. (freezing.)	Velocity of Sound, reduced to 0° R.	Velocity assigned by Chladni, <i>Acoustics</i> , p. 274.
	Theory.	Experiment.	
	Metres.	Metres.	Metres.
Oxygen (from Manganese, therefore impure)...	317·7	316·6	310
Azote.....	339·0	338·1	310
Hydrogen.....	1233·3	914·2	{ 680 } according to { 820 } its purity.
Carbonic acid.....	270·7	275·3	269
Oxide of carbon (from zinc and chalk).....	341·1	316·9	
Protoxide of azote (from nitrate of ammon.)...	270·6	281·4	
Deutoxide of azote (nitrous gas).....	317·4	309·8	320
Carburetted hydrogen.....	337·4	317·8	
Sulphuretted hydrogen.....	305·7	318·7	
Sulphurous acid.....	229·2	229·2	
Muriatic acid gas.....	298·8	309·3	
Ammonia.....	432·0	399·4	
Vapour of water at 54° R.....	422·6	369·6	
Vapour of alcohol at 48° R.....	262·7	289·1	

We give this Table, to the best of our comprehension, from a very imperfect and obscure abstract of an inaugural dissertation of M. Van Rees, (printed in 1819,) given in the *Journal de Physique*, 1821, p. 40. We have not been able to procure the original. The differences of the columns probably arise from impurities in the gases, or difficulty in estimating the exact pitch of Sounds propagated by them.

These determinations are, of course, liable to considerable errors; but the difference between the results of theory and experiment in the case of hydrogen is so great as to warrant a conclusion, otherwise not improbable, that the value of the coefficient K in that gas (at least) is materially different from what it is in others. Experiments are hardly yet sufficiently multiplied to enable us to speak with certainty on this point; but if by any means we are enabled to determine precisely the velocity of Sound, in a gas, or indeed in any medium, the ratio of the values of this coefficient in it, and in air, may be obtained by the analogy

$$K : K' :: V^2 : s' ;$$

which expresses that the value of K is as the square of the velocity of Sound, and the specific gravity of the medium jointly. Thus the specific gravity of pure hydrogen being to that of air as 0·0694 : 1, (Thomson, *Attempt to establish the first Principles of Chemistry*, i. 72.) and the velocity of Sound in it being to that in air as 2999·4 to 1089·4, we have

$$K \text{ in hydrogen} : K \text{ in air} :: (2999\cdot4)^2 \times 0\cdot0694 : (1089\cdot4)^2 \times 1, \\ :: 0\cdot526 : 1 :: 1 : 1\cdot901.$$

But not only the velocity of Sound differs in media of different chemical and mechanical natures. Its intensity, *i. e.* the impression it is capable of producing on our organs of hearing, *cæteris paribus*, also varies extremely with a variation in the density of the transmitting medium. This we have already remarked in the case of air, whether rarefied or condensed. Priestley (*Observations and Experiments*, iii. 355.) enclosed a piece of clockwork, by which a hammer could be made to strike at intervals, in a receiver filled successively with different species of gas. The distances at which the Sound ceased to be heard were measured. He thus found that in hydrogen the Sound was scarcely louder than in a vacuum, (such a one as he could produce.) In carbonic acid it was louder than in air, and somewhat louder also in oxygen. Perolle (*Mém. Acad. Toulouse*, 1781; *Mém. Acad. Turin*, 1786-1787) has described some experiments not altogether in agreement with these. The distance at which a given Sound ceased to be heard in atmospheric air being 56 feet, he found that in carbonic acid it was 48 only; while in oxygen and nitrous gas the distance was 63, and in hydrogen only 11. Chladni found the Sound of hydrogen gas in an organ-pipe remarkably feeble and difficult to distinguish, and that of oxygen stronger than that of atmospheric air, but remarked nothing particular in the case of carbonic acid. (*Acoust.* 281.)

83.
Peculiarity
in hydrogen.

Value of K
in hydrogen
nearly dou-
ble of its
value in air.

84.
Intensity of
Sound differs
in differ-
ent media.

Leslie (*Camb. Phil. Trans.* i. 267.) relates some very curious experiments, by which it should appear that

85.

Sound.
Singular
effect of
hydrogen in
enfeebling
Sound.

hydrogen gas is peculiarly indisposed for the conveyance of Sound. He rarefied the air of a receiver in which a piece of clockwork was enclosed, striking a bell every half minute, 100 times; and then introduced hydrogen gas, when *no augmentation whatever* of the Sound took place. Yet more; when the air in the receiver was only half exhausted, and the deficiency filled up with hydrogen gas, not only the Sound was not increased, but was actually diminished *so as to become scarcely audible*. If this last fact be correctly stated, (which from the high character of Mr. Leslie, as an experimenter, we must not doubt,) some peculiar modification of the usual process by which Sound is propagated must have taken place. It is much to be regretted that the circumstances are not more fully stated; the *pitch* of the bell in air, in the mixed gases, and in hydrogen alone; the dimensions of the receiver; the distances at which the Sounds ceased to be heard; and whether the same effect took place when bells of different pitch were struck, and when the bell was muffled so as to produce no *musical* Sound, are all particulars of essential consequence to enable us to form a judgment of what really took place in this interesting experiment, which we venture to express a hope will be repeated and varied by its author on a scale proportioned to its importance. We shall have occasion again to refer to this subject. (See Index, *Interference of Sonorous Vibrations and Propagation of Sound in Mixed Media*.)

86.
Effect of
hydrogen on
the voice
when
breathed.
Musical
Sounds ex-
cited by the
combustion
of hydrogen.

When hydrogen is breathed (which may be done for a short time, but not altogether without inconvenience and even danger) the voice is singularly affected, being rendered extremely feeble, and at the same time raised in pitch. (Otier, *Journal de Physique*, vol. xlviii.) This is just what ought to arise from the lungs, larynx, and fauces being filled with an exceedingly rare medium; but if, as some experimenters relate, the effect subsists long after the hydrogen is expired, and the lungs completely cleared of it, this can only be ascribed to some physiological cause depending on its peculiar action on the organs of the voice. The singular Sounds produced by burning this gas in pipes of proper construction have nothing to do with the propagation of Sound in the gas itself.

87.
Propagation
of Sound in
vapours.

The propagation of Sound in vapours offers two distinct cases in which it would at first appear that very different effects should take place. In the first, in which the vapour is subjected to a less compression than what is sufficient to reduce a portion of it to the liquid state, experiments have sufficiently proved the identity of the laws which regulate the compression and dilatation of this species of elastic fluids with those which prevail in the case of ordinary gases; and, indeed, recent researches have proved that a great number, and rendered it probable that all the latter, are in fact only vapours of certain liquids capable of sustaining a very much greater than the ordinary atmospheric pressure; or, which is the same thing, habitually maintained at a temperature far above their boiling point. In this state, then, the propagation of Sound in vapours differs in no respect from that in gases. But when the pressure sustained by the vapour is sufficient to condense a portion of it, as, for instance, in the upper part of a vessel in which water is kept boiling, and which is therefore full of steam at 212° Fahr., it would seem, at first sight, that no Sound could be propagated through such a medium; for, since the slightest additional pressure is sufficient to reduce a portion of the vapour to the liquid state, it would appear that the whole effect of an impulse suddenly communicated to any portion of the vapour, urging it towards the adjacent stratum, would be, not the compression of the whole of such portion into less dimensions, accompanied with increased elastic force, but the absolute condensation of a small portion into inelastic water, the remainder retaining precisely the same elasticity as before. Thus the necessary conditions for the propagation of the impulse are nullified, and it should seem, therefore, that no Sound could be excited in such a case.

88.
Experiment
proposed
by Biot.

But if in vapours, as in gases, the act of compression develops a certain portion of heat, it is evident that this may be such as to prevent altogether the mechanical condensation of the compressed vapour, and maintain it in its elastic state even under the increased pressure; and therefore Sound ought on this supposition to be propagated freely. Thus it appears that we are furnished with an *experimentum crucis* for deciding on the validity of the explanation above stated of the excess of the observed above the theoretical velocity of Sound. If the momentary condensations and dilatations of an elastic fluid *do*, as supposed in that explanation, give out and absorb heat, Sound should be freely propagated in a *saturated vapour*, (*i. e.* a vapour in contact with liquid, or under a pressure which it can just sustain.) If not, no Sound can be transmitted by it. The experiment has been made with care by M. Biot, assisted by Messrs. Berthollet and Laplace, (*Mém. d'Arcueil*, ii. 99.) by means of a bell suspended in a large glass balloon. When completely exhausted, no Sound was heard on striking the bell; but on the admission of a little water it was feebly heard, and as the water and balloon were warmed, became stronger and stronger. When allowed to cool, the vapour condensed, and the Sound became enfeebled by the same degrees. When alcohol was used instead of water the Sound was more powerful, and still more so when ether was introduced, the vapours of these liquids at a given temperature being more dense than that of water. As in these experiments care was taken to keep the inside of the balloon constantly wet with the liquid, it is evident that the only condition requisite to be observed, that of maintaining the vapour in the interior, at its maximum of pressure, was completely fulfilled. The reader is referred to the original Memoir for an account of the details of this elegant experiment. The reasoning above stated is M. Biot's. We would remark, however, on it, that the developement of the latent heat of a vapour on its condensation into a liquid, though, no doubt, analogous to, is still in a material point different from, the developement of heat in a gas by mere compression, unaccompanied with a change of state. If the latent heat of steam at 212° (amounting to about 945°) be not conducted away, the steam cannot be condensed into water of 212° . A portion will be condensed, but its latent heat will be employed in raising the temperature of the water produced and of the remaining steam, and thus increasing its elasticity and resistance to the pressure. Thus, the propagation of Sound in saturated vapour is not incompatible with the reduction of a portion of the vapour to a liquid state at every condensation caused by the sonorous pulse, and its reconversion into vapour when the condensation goes off: nor is it to be assumed as proving any thing with respect to gases or vapours under less than their maximum pressure. The heat developed may (for any thing this experiment proves) come entirely from the liquefied

Remarks
thereon.

Sound. portion, and have no existence when no portion is liquefied. We do not make this remark as detracting from the merit of M. Biot's ingenious views, in which, on the contrary, we fully coincide as to their result, but as an instance of the circumspection requisite in drawing conclusions in a theory so delicate as that of the propagation of Sound.

Part I.

§ IV. Of the Propagation of Sound through Liquids.

The experiments of Canton, and the more recent ones of Perkins, Oersted, Colladon, and Sturm, have shown that water, alcohol, ether, and, no doubt, all other liquids, are compressible and elastic, though requiring a very much greater force to produce a given diminution of bulk than air. Water, according to the experiments of Perkins, (*Phil. Trans.* 1820, p. 234.) as computed by Dr. Roget, suffers a condensation of $\frac{1}{21236}$ by a pressure of 100 atmospheres. This result agrees sufficiently well with that of Canton, which gave a condensation of 0.000046 for every atmosphere of pressure, (*Phil. Trans.* 1764,) and has been since confirmed by Oersted's researches.

89.

Liquids compressible and elastic.

Since water, then, and other liquids have the essential property of elastic media, on which the propagation of Sound depends, it may be presumed, *à priori*, that Sounds are capable of being conveyed by them as well as by the air; and, indeed, better, by reason of their greater density, pursuant to the same law which obtains in gases. This conclusion is abundantly confirmed by experiment. Hauksbee (*Phil. Trans.* 1726, 371.) ascertained that water would transmit a Sound excited in air. Anderson (*Phil. Trans.* 1748, p. 151.) describes a number of experiments on the hearing of fishes, from which, indeed, he concludes, that they are altogether devoid of this sense. But a very different conclusion really follows from them. Fishes enclosed in a glass jar appeared (says Anderson) utterly insensible to any Sound excited in the air without them, (if unaccompanied with motion,) but the slightest tap with the nail on the edge of the jar, although made in such a situation that the motion could not be seen by them, immediately disturbed them. This is easily explicable; and is, in fact, just what ought to happen. The intensity of Sound excited in any medium must evidently be proportioned to the energy of the original impulse, and must therefore be much greater when arising from the direct impact of a solid body on the water, or its containing vessel, than from that of the particles of the air in a sonorous wave, whose momentum is necessarily very small. As fishes have no external organs of hearing, Sounds must be conveyed to their sensorium by direct propagation, through the bones of their heads; and they are probably insensible to, or habitually careless of, those feeble impulses which are communicated from the air. But that the latter impulses do exist, and are audible by our ears, Anderson's Paper furnishes proof enough. He made three people, stripped quite naked, dive at once, and remain about two feet below the surface of the water. In this situation he spoke to them as loud as he was able. At their coming up they repeated his words, but said he spoke very low. He caused the same persons to dive about 12 feet below the surface, and discharged a gun over them, which they said they heard, but that the noise was scarce perceivable. He further caused a diver to halloo under water, which he did; and the Sound was heard, though faintly. A grenade, exploded about nine feet below the surface, gave a prodigious hollow Sound, with a most violent concussion of the earth around. Lastly, he caused a diver to descend with a bell in his hand, whose ringing he (the diver) assured him he could hear distinctly at all depths; adding, also, that he could hear the rushing of the water through a flood-gate at 20 feet distance from the place he was in.

90.

And therefore capable of conveying Sound.

Hearing of fishes.

Sounds excited in air heard under water.

The Abbé Nollet having descended to various depths, from 4 to 24 inches, could hear all Sounds made in the air (as a clock striking, a hunter's horn, the human voice, &c.) distinctly, but faint and attenuated. (*Brocklesby, Phil. Trans.* 1748, p. 237.)

91.

Nollet's experiments.

Franklin, having plunged his head below water, caused a person to strike two stones together beneath the surface; and at more than half a mile distance heard the blows distinctly. These instances are sufficient to show that Sound is *audibly* conveyed through water as well as through air; and, indeed, if properly excited, much better.

92.

Franklin's.

A series of experiments on the velocity of Sound in sea-water was instituted by M. Beudant, at Marseilles. Two observers, with regulated watches, were stationed in boats at a known distance. Each was accompanied by a diver. A bell was struck at stated intervals at one station; and at the instant of its being heard by the diver at the other he made a signal, and the time was noted by the observer in the boat. Of course, time was lost. The mean result of these observations gives 1500 metres = 4921 feet per second for the velocity.

93.

Velocity of Sound in water. Beudant's experiments

A more careful and no doubt more exact determination was undertaken and executed in 1826, by M. Colladon, in the Lake of Geneva. After trying various means for the production of the Sound, as the explosion of gunpowder, blows on anvils, and bells; the latter were preferred, as giving the most instantaneous, and, at the same time, most intense Sound, the blow being struck about a yard below the surface by means of a metallic lever. The experiments were all made at night, to avoid the interference of extraneous sounds, and for the better observing of the signals made at each blow by the flash of gunpowder.

94.

Colladon and Sturm's experiments

To render audible to an observer out of water (in which situation only can any observations worthy of confidence be made) sounds excited at a great distance, a very ingenious method was practised by M. Colladon. He found, that although the Sound of the blow was well heard directly above the bell, yet the intensity of the Sound so propagated into the air diminished with great rapidity as the observer removed from its immediate neighbourhood, and at two or three hundred yards it could no longer be heard at all. This fact renders it probable, that the waves of Sound, like those of light, in passing from a denser into a rarer medium, undergo, at a certain acuteness of incidence, a total reflexion; (see LIGHT, Art. (184); see also Index to this Article—

95.

Method practised by them to hear Sounds in water at great distances.

Sound.

Reflexion of Sound—Echo;) and cease to penetrate the surface, so that the Sound heard beyond that limit is merely that which diverges, *in the air*, from the point immediately above the bell. Acting on this idea, M. Colladon plunged vertically into the water a thin tin cylinder, about three yards long and eight inches in diameter, closed at the lower end, and open to the air above; thus forming an artificial surface on which the sonorous waves, impinging perpendicularly, might enter the air, and be thence propagated freely as from a new origin; just as we may look into water at any obliquity by using a hollow tube with a glass plate at the end perpendicular to the axis. This contrivance succeeded completely, and he was enabled by its aid to hear the strokes of a bell under water at a distance of 2000, 6000, and even 14,000 metres, (about 9 miles,) viz. across the whole breadth of the lake of Geneva, from Rolle to Thonon. A better spot could not have been found, the water being exceedingly deep, without a trace of any current, and of the most transparent purity. The signals were made by the inflammation of gunpowder, which being performed by the same blow of the hammer by which the bell was struck, all loss of time was effectually avoided. The time was reckoned by a quarter second stop-watch, from the appearance of the flash to the arrival of the Sound.

96.
Result of
their expe-
riments

By the mean of 44 observations on three different days, it appeared that a distance of 13·487 metres was traversed in 9·295 seconds, the greatest deviation being less than three-tenths of a second. M. Colladon assumes 9·4 as the true interval, regarding it as probable that a minute portion of time is necessarily lost in the estimation of the interval. The mean temperature of the water, from trials made at both stations, and half way between, was found to be 8°·1 cent. (= 46°·6 Fahr.) At this temperature, then, the velocity of Sound in the water of the Lake of Geneva was 1435 metres = 4708 feet per second.

97.
Comparison
with theory.

To compare this result with theory, we will take the data afforded by the experiments of Messrs. Colladon and Sturm on this very water; whose foreign contents, as appears by the analysis of M. Tingry, amount only to $\frac{1}{5000}$ of its weight, and which may, therefore, be regarded as pure water, (though, of course, saturated with air.) They state the compressibility, both at this and at the freezing temperature, at 0·0000495 for every atmo-

sphere; *i. e.* that an increase of pressure of one atmosphere produces a diminution of bulk equal to $\frac{49\cdot5}{1\cdot000\cdot000}$ of the whole, or very nearly one two-hundred-thousandth. But as the atmospheric pressures used in their experiments were not *standard* ones, but each equal to a column of mercury 0·76 metres long, at a temperature of 10° cent., instead of 0° the compressibility by one *standard atmosphere* must be equal to

$$0\cdot0000495 \times \frac{\text{specific gravity of mercury at } 0^{\circ} \text{ cent.}}{\text{specific gravity of mercury at } 10^{\circ} \text{ cent.}} = 0\cdot0000495 \times \frac{1\cdot0018}{1\cdot0000} = 0\cdot000049589.$$

98.
Different
mode of ex-
pressing the
coefficient
of elasticity.

To apply the general analysis by which the velocity of Sound in an elastic medium was deduced (Art. 52.) to this case, we must express the elasticity in a form somewhat different from that before employed in the case of aerial fluids. Let us then put e for the *compressibility* of any elastic medium, or, the diminution of bulk it will sustain by an additional pressure of a single atmosphere; or by immersion to the depth of 0·76 metres (= 29·927 inches) in mercury of the freezing temperature, (so that in water $e = 0\cdot000049589$.) Then, if we neglect the heat disengaged by compression, an infinitesimal column dx of the medium, when compressed into a space

$1 - \frac{dy}{dx}$, will exert a resistance on the compressing column equal to *one atmosphere* $\times \frac{1 - \frac{dy}{dx}}{e}$. Let A be the area of the section of the sounding column, then will the weight of the particle dx be represented by $A dx \times D$, where D is the density of the medium; and its elastic pressure on the section A, which separates it from the

preceding particle, will be $A \times \frac{1 - \frac{dy}{dx}}{e} \times (h) \Delta$, where (h) = the *standard* height of mercury in the barometer and Δ the density of mercury *at the freezing temperature*. This, then, is the force mutually exerted between dx and the particle immediately preceding it. Similarly the force exerted between dx' and the particle (dx)

immediately preceding it is represented by $A \times \frac{1 - \frac{dy'}{dx'}}{e} \times (h) \Delta$; and the difference of these, or the whole force by which dx is urged forwards, is therefore

$$- d \left\{ A \Delta \cdot \frac{(h)}{e} \left(1 - \frac{dy}{dx} \right) \right\} = A dx \cdot \frac{(h) \Delta}{e} \cdot \frac{d^2 y}{dx^2};$$

so that the accelerating force acting on dx is

$$\frac{(h)}{e} \cdot \frac{\Delta}{D} \cdot \frac{d^2 y}{dx^2};$$

99.
Resulting
formula for
velocity of
Sound.

If we take into consideration the heat developed by compression, we have only to multiply this by the coefficient K. Finally, therefore, if, as before, we represent by a the velocity of Sound, we shall have

$$a = \sqrt{\frac{2g(h)\Delta}{eD}} K = \sqrt{\frac{9\cdot8088 \times 0\cdot76 \times 13\cdot568}{0\cdot000049589 \times D}} K.$$

Sound.

The specific gravity of the water of the lake at the temperature of the experiment was found to be exactly that of distilled water at its maximum density, the trifling expansion due to the excess of temperature being exactly counterbalanced by the superior density due to the saline contents, so that $D = 1$. Reducing, then, the value of a to numbers, we find

$$a = 1428.2 \text{ met. } (= 4685.6 \text{ feet}) \times \sqrt{K}.$$

As we have seen, the velocity actually observed was 1435 metres. The agreement of this with the coefficient of \sqrt{K} within 7 metres (a space run over by the aqueous pulse in one 200th of a second) is so near, as to authorize the conclusion that in water, at least, the heat developed by compression, and consequent increased resistance to sudden condensation, is insensible.

In the course of these experiments, M. Colladon was led to remark some very curious particulars respecting the nature, intensity, and duration of Sounds propagated by water. He observed, first, that the Sound of a bell struck under water, when heard at a distance, has no resemblance to its Sound in air. Instead of a continued tone, a short sharp sound is heard, like two knife-blades (messerklängen) struck together. The effect produced by hearing such a short dry sound, at a distance of many miles from its origin, he compares to that of seeing, for the first time, very distant objects sharply defined in a telescope. When tried at different distances, it preserved this character, varying only in intensity, so as to render it impossible to distinguish whether the sound heard arose from a violent blow at a great distance, or a gentle one near at hand. It was only when within 200 metres (about a furlong) that the musical tone of the bell was distinguishable after the blow. In air the contrary takes place, as every one knows; the shock of the first impulse of the hammer being heard only in the immediate neighbourhood of the bell, while the continued musical Sound is the only one that affects the hearing at a distance. The reason of this curious difference will be apparent when we come to speak of Musical Sounds. (See Index. *Musical Sounds. Vibrations of Bodies in different media.*)

Another very curious and important observation of M. Colladon, is that of the effect of interposed obstacles. Sounds in air spread round obstacles with great facility, so that by a hearer situated behind a projecting wall, or the corner of a building, sounds excited beyond it are heard with little diminution of intensity. But in water this was far from being the case. When the tin cylinder, or hearing-pipe, already mentioned, was plunged into the water, at a place screened from rectilinear communication with the bell by a projecting wall running out from the shore, whose top rose above the water, M. Colladon assures us, that a very remarkable diminution of intensity in the Sound was perceived, when compared with that heard at a point very near the former, but within reach of direct communication with the bell; or, so to speak, out of the acoustic shadow of the wall. Thus the phenomena of Sound in water approximate in this respect to the rectilinear propagation of light, and may lead us to presume, that in a medium incomparably more elastic than water, the *shadow* would be still more perfect and more sharply defined. A material support is thus afforded to the undulatory doctrine of Light, against one of its earliest and strongest objections—the existence of shadows.

It appears, from these experiments, that the velocity of Sound in water may be correctly computed when its compressibility is known, without the necessity of having regard to the heat developed during compression. From all direct experiments hitherto made, it appears that in water, and all other liquids, the quantity of heat thus developed is either altogether insensible, or at least very minute; so that, most probably, the same thing will hold in other liquids. The Memoir of Messrs. Colladon and Sturm, then, which contains a very elaborate determination of the compressibility of a variety of liquids, will afford the means of computing the velocity of Sound in them. We, therefore, subjoin a Table of their results, and of such others as we have been able to collect.

Part I.

Reduced to numbers.

100.

Heat developed by the compression of water.

101.

Curious phenomena observed in these experiments.

102.

Non-divergence of the rays of Sound round obstacles in water

103.

Velocity of Sound in other liquids.

Sound.
Table of absolute compressibilities of various bodies.

Substance compressed.	Authority.	Absolute compression in millionth parts of the original volume.	Pressure by which the compression in the last column was produced.
Mercury at 0° cent.	Colladon and Sturm.	5.03	A column of mercury 0m.76 high at 10° C.
Water free from air 0°	Ditto.	51.30	Ditto.
Water saturated with air at 0°	Ditto.	49.50	Ditto. (If correctly computed.)
Oil of turpentine at 0°	Ditto.	73.0	Ditto.
Nitric acid S. G. 1.403 at 0°	Ditto.	35.5	Ditto.
Concentrated sulphuric acid at 0° ..	Ditto.	32.0	Ditto. (Query strength.)
Acetic acid at 0°	Ditto.	42.2	Ditto.
Alcohol at 11° 6 cent.	Ditto.	{ 96.2	Ditto. (Under an initial pressure of one atmosphere.)
		{ 93.5	Ditto. (Under 8 atmosph. init. press.)
		{ 89.0	Ditto. (Under 20 ditto.)
Sulphuric ether at 0°	Ditto.	{ 133.3	Ditto. (Under 3 ditto.)
		{ 118.5	Ditto. (Under 24 atmospheres.)
———— at 11° 4	Ditto.	{ 149.3	Ditto. (Under 3 ditto.)
		{ 141.3	Ditto. (Under 24 ditto.)
Water saturated at 20° cent. with } ammonia (temp. 10°)	Ditto.	38.0	Ditto. (Under a mean pressure of about 10 atmospheres. Diminishes rapidly as the pressure increases.)
Nitric ether 0°	Ditto.	71.5	Ditto.
Acetic ether at 0°	Ditto.	{ 79.3	Ditto. (Under 4 atmospheres.)
		{ 71.3	Ditto. (Under 16 ditto.)
Muriatic ether at 11° 2	Ditto.	{ 85.9	Ditto. (Under 2 ditto.)
		{ 82.25	Ditto. (Under 9 ditto.)
Linseed oil	Galy-Calazat.	46.8	One atmosphere, (doubtful.)
Olive oil	Ditto.	47.0	Ditto ditto.
Glass	Colladon and Sturm.	3.30	0.76 met. of mercury of 10° cent.
————	Galy-Calazat.	2.84	One atmosphere (!)
Copper	Ditto.	7.09	Ditto. ditto.
Lead	Ditto.	0.18	Ditto. ditto.
Water	Canton.	46.00	One atmosphere at 50° Fahr.
————	Perkins.	47.09	As computed by Dr. Roget.
————	Oersted.		

Of the Propagation of Sound in Solids and in Mixed Media.

104. Solids, if elastic, are equally well, or better, adapted for the conveyance of Sound with fluids. By elasticity in a solid is not meant a power of undergoing *great* extensions and compressions, after the manner of air, or Indian rubber, and returning readily to its former dimensions; but rather what is commonly called hardness, in contradistinction to toughness, a violent resistance to the displacement of its molecules *inter se* in all directions. Thus the hardest solids are, generally speaking, the most elastic, as glass, steel, and the hard brittle alloys of copper and tin, of which mirrors are made; and in proportion as they are so, they are adapted to the free propagation of Sound through their substance.

105. But an important condition in their constitution is homogeneity of substance; and in a substance perfectly homogeneous, we may add, too, uniformity of structure. The effect of want of homogeneity in a medium, on its power of propagating Sound, is precisely analogous to that of the same cause in obstructing the free passage of light, and (as the undulatory doctrine of light teaches) for the very same reason. The sonorous pulses, in their passage through it, are at every instant changing their medium. Now, at every change of medium, two things happen; first, a portion of the wave is reflected, (see *Reflexion* of Sound,—*Echo*, in the Index,) and the intensity of the transmitted part is thereby diminished; secondly, the direction of propagation of the transmitted part is changed, and the sonorous rays, like those of light, are turned aside from their direct course. (See *Refraction* of Sound, in the Index.) Thus the general wave is broken up into a multitude of non-coincident waves, emanating from different origins, and crossing and interfering with each other in all directions. Now, whenever this takes place, a mutual destruction of the waves, to a greater or less extent, arises, and the Sound is stifled and obstructed. Further yet:—as the parts of a non-homogeneous medium differ in elasticity, the velocities with which they are traversed by the sonorous pulses also differ; and thus, among the waves which do ultimately arrive at the same destination in the same direction, some will arrive sooner, some later. These, by the law of interference, tend mutually to destroy or neutralize each other.

106. But of all causes which obstruct the propagation of Sound, one of the most effective is a want of perfect adhesion at the junctures of the parts of which such a medium consists. The effect of this may be conceived, by regarding the superficial strata of molecules of each medium when in contact, as forming together a thin film of less elasticity than either; at which, therefore, a proportionally greater reflexion of the wave will take place than if the cohesion were perfect,—just as light is much more obstructed by a tissue of cracks pervading a piece of glass, than it would be by any inequality in the composition of the glass itself.

Sound.

A pleasing example of the stifling and obstruction of the pulses propagated through a medium, from the effect of its non-homogeneity, may be seen by filling a tall glass (a Champagne glass, for instance) half full of that sparkling liquid. As long as its effervescence lasts, and the wine is full of air-bubbles, the glass cannot be made to ring by a stroke on its edge, but gives a dead, puffy, disagreeable Sound. As the effervescence subsides the tone becomes clearer, and when the liquid is perfectly tranquil the glass rings as usual; but on reexciting the bubbles by agitation, the musical tone again disappears. To understand the reason of this, we must consider what passes in the communication of vibrations through the liquid from one side of the glass to the other. The glass and contained liquid, to give a musical tone, must vibrate regularly in unison as a system; (see *Vibrations* of a System of bodies;) and it is clear, that if any considerable part of a system be unsuceptible of regular vibration, the whole must be so. This neat experiment seems to have been originally made by Chladni, (*Acoustique*, § 214,) and has been employed by Humboldt, to illustrate by it a natural phenomenon equally familiar and striking; we mean, the greater audibility of distant Sounds by night than by day. This he attributes to the uniformity of temperature in the atmosphere by night, when upward currents of air, heated by their contact with the earth under the influence of the sun's rays, are no longer continually mixing the lower with the upper strata, and disturbing the equilibrium of temperature. It is obvious that Sound, as well as light, must be obstructed, stifled, and dissipated from its one original direction, by the mixture of air of different temperatures, (and consequently elasticities;) and thus the same cause which produces that extreme transparency of the air at night, which astronomers only fully appreciate, renders it also more permeable to Sound. There is no doubt, however, that the universal and dead silence generally prevalent at night renders our auditory nerves sensible to impressions, which would otherwise escape them. The analogy between Sound and light is perfect in this as in so many other respects. In the general light of day the stars disappear. In the continual hum of noises which is always going on by day, and which reach us from all quarters, and never leave the ear time to attain complete tranquillity, those feeble Sounds which catch our attention at night make no impression. The ear, like the eye, requires long and perfect repose to attain its utmost sensibility.

To a cause of the same kind, particularly modified, may possibly be attributable the singular effect of hydrogen gas when mixed with air, already described, Art. 85, in unfitting it for the free propagation of Sound. Chemists maintain that when gases are mixed, the molecules of each form separate and independent systems, being mutually inelastic, and each sustaining a part of the pressure proportional to its own density. They admit, however, that the molecules of one gas (A) act as obstacles, to obstruct the free motion of those of another (H;) and on this principle they explain the *slow* mixture of two gases in separate vessels communicating by a narrow aperture. Granting these postulates, let us conceive a pulse excited in a mixture of equal volumes of two gases. If the velocity of Sound in both be alike, the pulse will run on in each, although independently, yet with the same speed, and at any instant, and at any point of the medium, the contiguous molecules of both gases will be moving in the same direction and with the same velocity. They will, therefore, offer no mechanical obstruction to each other's motion, and Sound will be freely propagated. But if they differ in their specific elasticity, the case will be altered. Each being non-elastic to the other, two distinct pulses will be propagated, and will run on with different velocities; the molecules of either gas, at different points beginning, and ceasing to be agitated with the pulsation at different instants. Thus an internal motion, a change of relative position among the molecules of the gas (H) and those of the gas (A) will take place, the one set being obliged to force themselves a passage between the other; in which, of course, a portion of their motion will be diverted in all sorts of lateral directions, and will be mutually destroyed. It is evident that the greater the difference of specific elasticities, the greater will be the effect of this cause. In hydrogen the velocity of the pulse is nearly three times its velocity in atmospheric air; and, of course, it may be expected in this case to act with great efficacy. In azote and oxygen the velocities are so nearly alike, that very little obstruction can arise from its influence; so that, in so far as the phenomena of Sound are concerned, atmospheric air may be looked upon as a homogeneous medium.

If saturated with aqueous vapour, at high temperatures, however, it is possible that the effect may become sensible, and, perhaps, to this cause may be attributed a phenomenon, mentioned by more than one experimenter on this branch of Physics, of the occasional duplication of the Sound of a gunshot heard from a great distance, a part of the Sound being transmitted quicker than the rest by aqueous vapour, or even by water in the liquid state suspended in the air. If this be the case, Sounds might be expected to be heard double in thick fogs, or in a snow-storm. But the remarkable obstruction to Sound caused by fog, and especially by snow, (see Art. 21,) would, probably, prevent any Sound from being heard far enough to permit the interval of the two pulses to be distinguishable. This latter phenomenon, we may here observe, affords another and very satisfactory illustration of the general principle explained in Art. 107. To it we may add the well-known effect of carpeting, or woollen cloth of any kind, in deadening the Sound of music in an apartment. The intermixture of air and solid fibres in the carpets through which the Sound has to pass, deadens the Echo between the ceiling and floor by which the original Sound is swelled.

A phenomenon noticed by every traveller who visits the Solfaterra near Naples, but whose true nature has been much misconceived, is easily explicable on this principle. The Solfaterra is an amphitheatre, or extinct crater, surrounded by hills of lava, in a rapid state of decomposition by the action of acid vapours issuing from one principal and many subordinate vents and cracks. The whole soil of the level at its bottom consists of this decomposed lava, whose disintegration, however, is not so complete as to reduce it to powder; but leaves it in coherent white masses of a very loose friable structure. At a particular spot, a large stone violently thrown against the soil, is observed to produce a peculiar hollow Sound, as if some great vault were below. Accordingly it is usually cited as a proof of the existence of some vast cavity below, communicating with the ancient vent of the volcano, and perhaps with subterraneous fires; while others ascribe it to a reverberation from the surrounding hills, with which it is nearly concentric; and others to a variety of causes more or less fanciful. It seems most

Part I.

107.
Experiment
in illustration
of.

Greater
audibility of
sounds by
night than
by day.
Humboldt's
explanation

Another
reason.

108.
Sounds in
mixed
gases.

Obstruction
of Sound by
hydrogen
gas mixed
with air ex-
plained.

109.
Duplication
of Sounds
occasionally
observed.

Effect of
carpeting,
&c. in dead-
ening
Sound.

110.
Pheno-
menon
observed at
Solfaterra.

Explained

Sound.

probable, however, that the hollow reverberation is nothing more than an assemblage of partial echoes arising from the reflexion of successive portions of the original impulse in its progress through the soil at the innumerable half-coherent surfaces composing it; were the whole soil a mass of sand, these reflexions would be so strong and frequent as to destroy the whole impulse in too short an interval to allow of a distinguishable after-sound. It is a case analogous to that of a strong light thrown into a milky medium, or smoky atmosphere; the whole medium appears to shine with a nebulous undefined light. This is to the eye, what such a hollow Sound is to the ear.

111.
Essential
difference
in the con-
stitution of
fluids and
solids.

The general principle on which the conveyance of Sound through solids depends, is precisely the same as in fluids; and the same formula may be used to express its velocity when the specific elasticity is known. There are, however, two very important particulars in which they differ; first, the molecules of fluids are capable of displacement *inter se*. Those of solids, on the other hand, are subjected to the condition of never changing their order of arrangement. Secondly, each molecule of a fluid is similarly related to those around it in all directions; in solids each molecule has distinct sides, and different relations to space and to the surrounding particles. Hence arise a multitude of modifying causes, which must necessarily affect the propagation of sonorous pulses through solids, which have no place in fluids, and modes of vibration become possible in the former, which it is difficult to conceive in the latter, whose parts have no lateral adhesion. Thus we may conceive pulses propagated in solids, like those of a cord vibrating transversely, in which the motion of each molecule is transverse, or oblique, to the direction in which the general pulse is advancing. Again, the cohesion of the molecules of crystallized bodies is different on their different sides, as their greater facility of cleavage in some directions than in others indisputably proves. They must, in consequence, have unequal elasticities in different directions; and thus the velocity of the pulse propagated through a crystallized solid will depend on its direction with respect to the axes of crystallization. Among uncrystallized solids, too, there are many, such as wood, whalebone, &c. which have a fibrous structure, in virtue of which, it is evident, they are very differently adapted to convey an impulse longitudinally and transversely.

Propagation
of oblique
or transverse
undulations.

Propagation
of pulses
in crystal-
lized media.

112.
Wood an
excellent
conductor
of Sound.

Interruptions of crystalline structure, then, ought to produce an effect on the conveyance of Sound analogous to that of the mixture of extraneous matter in a medium. The conducting power of wood along the grain is certainly very surprising. A simple experiment will show it. Let any one apply his ear close to one end of the longest stick of sound timber, and let an assistant at the other end scratch with the point of a pin, or tap so lightly with its head as to be inaudible to himself. Every scratch or tap will be distinctly, nay loudly, heard at the other end, as if close to the head. In general, however, all solids tolerably compact conduct Sound well, and transmit it rapidly.

113.
Conduction
of Sound
along a wire.

Chladni relates an experiment made by Messrs. Herhold and Rafn, in Denmark, where a metallic wire 600 feet long was stretched horizontally. At one end a plate of sonorous metal was suspended, and slightly struck; an auditor placed at the other, and holding the wire in his teeth, heard at every blow two distinct sounds; the first transmitted almost instantaneously by the metal, the other arriving later through the air. Messrs. Hassenfratz and Gay Lussac made a similar experiment in the quarries at Paris; a blow of a hammer against the rock produced two Sounds, which separated in their progress; that propagated through the stone arriving almost instantly, while the Sound conveyed by the air lagged behind. The same thing has been observed in the blasting of rocks in the deep mines of Cornwall. These experiments were, however, made at intervals too short to give any numerical estimate of the velocity of transmission of Sound in the iron or stone. The only direct experiments we have on this subject are those of M. Biot himself, who, assisted by Messrs. Bouvard, Malus, and Martin, ascertained the interval required for the Sound of a blow on the cast-iron conduit pipe already spoken of, Art. 24, to traverse measured lengths of it. The pipe consisted of joints of cast iron, each $2^{\text{met}}.515 = 8.2514$ feet long, and connected by flanches with collars of lead covered with tarred cloth interposed, and strongly screwed home; each collar measured $0^{\text{met}}.14256 = 0^{\text{f}}.46773$. A blow being struck at one end, and heard at the other, the interval between the arrival of the Sound through the air and through the iron was noted. The length being known, the time required for the transmission of the aerial Sound became known with great precision, and thence the time of transmission through the iron became known also. The following is a statement of the results:

Through
rocks.

Through
cast iron.
Biot's ex-
periments.

Observers' names.	Number of iron joints.	Number of leaden collars.	Total length when con- nected in metres.	Observed in- terval of the sounds in seconds.	Number of observa- tions.	Computed time of trans- mission in air. Seconds.	Deducted time of transmis- sion through the compound solid.
Biot, Bouvard ...	78	77	197.27	0.542	53	0.579	0.037
Bouvard, Malus...	156	155	394.55	0.810	64	1.158	0.348
Biot, Martin	376	375	951.25	2.500	200	2.790	0.290
Ditto do.	ditto.	ditto.	ditto.	Time directly observed by a different method.			0.260

114.
Actual ob-
served velo-
city of
Sound in
cast iron.

The last result was obtained as follows. Each observer holding in one hand a chronometer and in the other a hammer, (the chronometers being carefully compared,) the one (M) at the precise beats of 0° and 30° struck on the pipe, and the other noted the moment of arrival of that Sound *only* which was propagated through the solid, (*i. e.* the first.) At every 15° and 45° , and also precisely on the beat of his chronometer, the observer (B) struck the pipe, and (M) noted in the same manner the moment of arrival of the metallic Sound by *his* watch. From

Part I.

Sound.

such reciprocal observations, a very little consideration will show that the exact time required for the Sound's propagation through the solid may be obtained, independent of any observation of the aerial Sound, as well as of the rates of the watches. The agreement of the results obtained by the two methods sufficiently proves that the result of Messrs. Bouvard and Malus, in the above Table, is too large; rejecting this for that reason, and the first on account of the shortness of the pipe, we have, as a mean result, 0.3275 for the time required to traverse 951.25 metres, which gives a velocity of $3459.1 = 11090$ feet per second for the velocity of Sound in cast iron at the temperature of the experiment, (11° cent. $= 51^{\circ} 8$ Fahr.) and neglecting the very small retardation due to the collars, whose united thickness was $5.161 = 18.41$ only. This is about $10\frac{1}{2}$ times its velocity in air. Chladni assigns 3597 metres for the velocity of Sound in brass. Laplace, calculating on an experiment of Borda, on the compressibility of brass, makes it 3560.4 . According to Chladni, the following are the velocities of Sound in different solids, that in air being taken for unity: tin $= 7\frac{1}{2}$, silver $= 9$, copper $= 12$, iron $= 17$, glass $= 17$, baked clay (porcelain?) $10 \dots 12$, woods of various species $= 11 \dots 17$. The error in the case of iron throws a doubt on all the rest; unless, perhaps, steel be meant. (*Acoust.* § 219.)

Its velocity in brass.

From this determination we may estimate the time it requires to transmit force, whether by pulling, pushing, or by a blow, to any distance, by means of iron bars or chains. For every 11090 feet of distance the pull, push, or blow, will reach its point of action one second after the moment of its first emanation from the first mover. In all moderate distances, then, the interval is utterly insensible. But were the sun and the earth connected by an iron bar, no less than 1074 days, or nearly three years, must elapse before a force applied at the sun could reach the earth. The force actually exerted by their mutual gravity may be proved to require no appreciable time for its transmission. How wonderful is this connection!

115. Time required to transmit force by iron rods, levers, &c.

§ VI. Of the Divergence and Decay of Sound.

Hitherto we have taken no account of the lateral divergence of Sound, which we have supposed confined by a pipe; but it is evident that condensation taking place in any section of such a channel will urge the contained air laterally against the side of the pipe, as well as forward along its axis; and, consequently, if the pipe were cut off at any point, the Sound would diverge from that point into the surrounding air. Accordingly, when any one speaks through a long straight tube the voice is heard laterally, as if proceeding from the mouth of a speaker at the orifice.

116. Divergence of Sound from the end of a pipe.

In general, a Sound excited in, or impulse communicated to, any portion of the air or other elastic medium, spreads, more or less perfectly, in all directions in space. We say more or less perfectly; for though there are Sounds, as the blow of a hammer, the explosion of gunpowder, &c. which spread equally in all directions, yet there are others which are far from being in that predicament. For instance, a common tuning-fork (a piece of steel in the shape represented in fig. 6) being struck sharply, when held by the handle (A) against a substance, is set in vibration, the two branches of the fork alternately approaching to and receding from each other. Each of them, consequently, sets the air in vibration, and a musical tone is produced. But this Sound is very unequally audible in different directions. If the axis of the fork, or the line to which it is symmetrical, be held upright about a foot from the ear, and it be turned round this axis while vibrating, at every quarter revolution the Sound will become so faint as scarcely to be heard, while in the intermediate axes of rotation it is heard clear and strong. The audible situations lie in lines perpendicular and parallel to the flat faces of the fork, the inaudible at 45° inclined to them. This elegant experiment, due originally to Dr. Young, has recently been called into notice by Weber. (*Wellenlehre*, § 271.)

117. Unequal divergence of certain Sounds. Fig. 6.

Exemplified in Sound of a tuning-fork.

The non-uniformity of the divergent pulses which constitute certain Sounds is easily demonstrated by considering what happens when a small disc is moved to and fro in a line perpendicular to its surface. The aerial molecules in front of the disc are necessarily in an opposite state of motion from those similarly situated behind it. Hence, if we conceive a wave propagated spherically all around it, the vertices of the two hemispheres in front and behind are in opposite motions with respect to the centre. But with regard to that wave of the sphere where the vibrating plate prolonged cuts it, there is evidently no reason why its molecules should approach to or recede from the centre, or, rather, there is as much reason for one as for the other. They will therefore either remain at rest, or move tangentially; so that the motion of the whole sounding surface, or wave, will, in this case, be rather as in fig. 7 than in fig. 8; and a corresponding difference, both in the intensity and character of the Sound heard in different directions, may be fairly expected.

118. A priori considerations.

Fig. 7.
Fig. 8.

The mathematical theory of such pulses as these is of the utmost complication and difficulty, depending on the integration of partial differential equations with four independent variables, viz. the time and the three coordinates of the moving molecules. It is therefore of much too high a nature to have any place in an Essay like the present. We shall merely content ourselves with stating the following as general results in which mathematicians are agreed.

119.
120.

Velocity of Sound in free air.

121.

1st. The velocity of propagation of a sonorous pulse is the same, whether we regard it as propagated in one, two, or three dimensions, i. e. in a pipe, a lamina, or a mass of air.

Law of the decay of Sound.

2nd. Sounds propagated in a free mass of air diminish in intensity as they advance further from the sonorous centre, and their energy is in the inverse duplicate ratio of this distance, *ceteris paribus*.

122.

We shall not attempt a proof of these propositions in the general cases, but content ourselves with illustrating them in one particular but important case, viz. when the initial impulse is confined to a very small space, and consists in any small radiant motion of all the particles of a spherical surface in all directions equally from the centre.

Case of a spherical undulation alike on all sides.

Since the initial wave is spherical, and similar in all its parts, it will evidently retain this property as it dilates

123.

Sound. by the progress of the impulse. If, then, it be conceived to be divided into its infinitesimal elements by a system of pyramidally disposed plane surfaces, having the common vertex in the centre of the sphere, each of these elements will form the base of one of the pyramids, and its molecules will advance and recede along its axis, as the pulse traverses them, without any change of their relative positions, *inter se*; so that the whole wave may be regarded as broken up into partial waves, each advancing as if confined within a pyramidal pipe, independently of all the rest.

124.
Velocity of
a pulse in
any pyra-
midal pipe.

Now in any one of these imaginary pipes the pulse will be propagated from layer to layer of the included particles with the same velocity as if the pipe were cylindrical, for the divergence of the sides of the pipe can only cause a lateral extension, and thence a diminished thickness, of the stratum, and will, therefore, alter the *velocity* of each of its molecules and the extent and law of its motion from what it would be in a cylindrical pipe. But if we consider a row of particles situated in the axis of the pyramid, the propagation of a pulse along them depends, as we have seen, neither on the velocity nor extent or law of excursion of the individual molecules, but only on their *intrinsic* elasticity. The latter, however, is not altered by the shifting of the whole vibrating fibre into a wider or narrower part of the pipe, since, from *this* cause, (its excursions from its original place being supposed infinitely small,) the whole dilates or contracts together, as if by an external compressing or rarefying force. Now we have seen that a variation in the *general* density of the medium in which a pulse is propagated from external pressure makes no change in its velocity. It follows, then, that the pulse will be propagated with equal velocity along the line of molecules in question, whether the pipe be cylindrical or pyramidal, or, indeed, of any shape; and as it runs equally fast in each of the imaginary pyramids into which the sphere is divided, the wave, of which it is an element, will dilate itself spherically with the uniform velocity of Sound in a straight tube. See also Euler, *Comm. Petrop.* 1771, cap. iv. &c., where the general equations for the motions of air in tubes of any figure are deduced, and the above proposition proved therefrom, in the case of hyperbolic tubes (p. 391) and conical or pyramidal ones (p. 418.)

125.
Application
of the law of
vis viva.

Let us now conceive a spherical wave by any means excited, such that the whole interval, reckoned along its radius, within which the motion to and fro of the molecules is comprised, shall be equal to $2a$. This, then, will be the breadth of the wave, and as all its parts dilate equally fast, this will continue to be its breadth throughout its whole progress. Its surface increases in the ratio of the square of the radius, and, therefore, calling r this radius, $2a r^2$ will represent the quantity of matter in motion at the moment the Sound has reached the distance r from its origin. Now, as all the air within and beyond the wave is quiescent, the whole impulse, or *vis viva*, originally communicated to the sphere first set in motion, is successively transferred to all the rest without loss or increase, (by the general law of the conservation of the *vis viva*. See MECHANICS.) And since it is distributed equally over the whole spherical surface, any portion of it, of given magnitude, (that of the aperture of the ear,

for instance,) will receive a part of the whole, proportional to $\frac{1}{2a r^2}$, or to $\frac{1}{r^2}$. Thus the whole shock or impulse

given to the ear, while the wave passes over it, is as the inverse square of the distance from its origin, and the absolute velocity of each molecule in any determinate phase of its motion inversely as the distance itself.

126.
Effect on
the ear esti-
mated by the
vis viva.

In the theory of Sound, as in that of Light, the intensity of the impression made on our organs is estimated by the shock, impetus, or *vis viva*, of the impinging molecules, which is as the square of their velocity; and not by their inertia, which is as the velocity simply. Were the latter the case, there could be no such thing as Sound or Light, since the negative inertia of the receding molecules would exactly equal and destroy their positive effect in their advance. (*Vide* LIGHT, Art. 578.) We conclude, then, that the intensity of Sound decays, in receding from its origin, as the square of the distance increases. It is exceedingly difficult to subject this law to satisfactory experimental tests, and we know of no attempt that has yet been made for the purpose.

Law of de-
cay of
Sound.

§ VII. Of the Reflexion and Refraction of Sound, and of Echos.

127.
Reflexion of
Sound at the
confines of
two media.

As there is no body in nature absolutely hard and inelastic, whenever the particles of a vibrating medium impinge on the solid or fluid matter which contains or limits it, they will agitate those of the latter with motions similar to their own, but modified by their greater or less density and mobility. A pulse, then, will be propagated into the solid or fluid according to its own laws, but this will not take place without the propagation back again of a pulse in the original medium, which may be regarded as the reflexion or echo of the first. To understand how this happens, let us consider what takes place when a motion is first impressed on any small stratum, whose thickness is $2a$ (as in Art. 63) of a sounding column, and let its law be as there expressed, *i. e.* that the velocity of any one of its particles at the distance x from its middle shall be, at the first instant, represented by $\phi(x)$, and the linear extent of the same molecule, compared with its original length, or e , shall equal $\psi(x)$ where $\phi(x) = 0$, and $\psi(x) = 1$ from $x = -\infty$ to $x = -a$, and from $x = +a$ to $x = +\infty$; while from $x = -a$ to $x = +a$ they may have any arbitrary values.

128.
Condition
of the single
propagation
of a pulse.

Since t is always positive, if we take $x > +a$ we have, of necessity, $x + at > +a$, and, therefore, $\phi(x + at) = 0$, and $\psi(x + at) = 1$, so that the values of v and e in equations (i) (j) become

$$v = \frac{1}{2} \left\{ a + \phi(x - at) - \psi(x - at) \right\};$$

$$ae = \frac{1}{2} \left\{ a - \phi(x - at) + \psi(x - at) \right\};$$

which gives

$$ae = a - v, \text{ or } 1 - e = \frac{v}{a}.$$

Again, on the negative side of the x we take $x < -a$, we have, of necessity, $x - at < -a$, and, therefore, $\phi(x - at) = 0$; $\psi(x - at) = 1$, and, consequently,

$$v = \frac{1}{2} \left\{ -a + \phi(x + at) + \psi(x + at) \right\};$$

$$ae = \frac{1}{2} \left\{ a + \phi(x + at) + \psi(x + at) \right\};$$

and, therefore, in this case,

$$1 - e = -\frac{v}{a}.$$

Now $1 - e = 1 - \frac{dy}{dx} = \frac{dx - dy}{dx}$ expresses the condensation the molecule dx has undergone in its disturbed

state. Hence we see, that in each of the two waves into which the primary impulse separates itself, one running towards the positive, the other towards the negative side of the x , there obtains this condition, *viz.* that the condensations of the aerial molecules are proportional to their actual velocities, the fluid being condensed wherever the molecules are moving from the origin of the first impulse, and dilated when returning to it.

This remarkable relation, which does not of necessity hold good within the limits of the first disturbance, establishes a distinction equally marked between the initial impulse and the waves freely propagated from it. The former is subject to no law, the latter must obey this condition. Any impulse, then, in which this condition is not satisfied, will immediately divide itself into two pulses running opposite ways, in each of which the condition in question holds, but so long as this condition obtains, no subdivision of the pulse will take place. This is easily shown, for if we suppose an initial impulse communicated to any portion ($2a$) of the fluid in which this relation is purposely maintained, such supposition is equivalent to making

$$a(1 - \psi(x)) = \phi(x),$$

which, substituted in (i) and (j), give, for all values of x and t ,

$$v = a - a\psi(x - at)$$

$$ae = \psi(x - at),$$

in which, whenever x is negative and $< -a$, we have $v = 0$ and $1 - e = 0$; thus indicating that the molecules on the negative side of such a primitive disturbance as supposed will remain constantly at rest, in other words, that the pulse will only be propagated on the positive side.

Whenever, then, in the progress of a pulse through a medium, it receives, by extraneous causes, any modification which disturbs the condition $1 - e = \frac{v}{a}$, it will undergo subdivision, and a portion will run backward, or be

reflected. Similarly this portion may be again subdivided and undergo partial reflexion, and so on *ad infinitum*, giving rise to a continual series of repetitions or Echos of the original Sound.

Let us now examine more closely what passes at the junction of two media when the pulse arrives there; and, first, in the equations (i) and (j) let us write, instead of $\phi(x)$ and $\psi(x)$, which are arbitrary, the combinations, equally arbitrary,

$$\phi(x) = F(x) + f(x)$$

$$\psi(x) = 1 + \frac{F(x) - f(x)}{a},$$

when it is to be observed that F and f are not the same with the F and f of Art. 57, which we shall have no more occasion to refer to. If, then, we put $s = 1 - e$, so that s shall represent the infinitely small condensation undergone by the molecule dx in its troubled state, those equations will become

$$v = f(x - at) + F(x + at) \}$$

$$as = f(x - at) - F(x + at) \} \quad (A)$$

These represent the state of the molecules of the first medium. Similarly, the state of those in the second will, of necessity, be represented by another system similar in form,

$$v' = f'(x - a't) + F'(x + a't) \}$$

$$a's' = f'(x - a't) - F'(x + a't) \} \quad (B)$$

where a' represents the velocity of Sound in the second medium, but the functions f' and F' (which are not here intended to represent the *derived* functions or differential coefficients of f and F , but others quite distinct) are here no longer arbitrary, because the motion of the particles of the second medium must evidently depend on that of the first, and on their relative elasticities, densities, &c. Let us see, then, what conditions the nature of the case, and their mutual action at their point of junction, will enable us to assign for deducing the forms of these functions from those of f and F , supposed to remain arbitrary.

129.

If this condition obtains, a pulse will not divide itself.

130.

If disturbed the pulse will divide and a part run back.

131.

General equations of propagation of Sound along pipes filled with different media.

Sound.

Now, first, the condition of continuity of the two media requires that the strata in contact should always have a common motion, or that for the value $x = l$, corresponding to the place of junction, we should have $v = v'$,

Part I.

132. Condition of continuity.

which gives

$$f(l - at) + F(l + at) = f'(l - a't) + F'(l + a't); \quad (C).$$

133. Must have a common elasticity at their junction.

Again, they must not only have a common motion, but a common elasticity, at this point. Now, if we call E the natural elasticity of the first medium and E' that of the second, the elasticities in the disturbed state will be expressed by $E(1 + \beta s)$ and $E'(1 + \beta' s')$, where β and β' are constant coefficients depending on the nature of the media and the heat developed in them by compression, and which would each be unity were no heat so developed. Hence we must have $E(1 + \beta s) = E'(1 + \beta' s')$, and since in the state of equilibrium $E = E'$, we must also have $\beta s = \beta' s'$, that is

$$\frac{\beta}{a} \{ f(l - at) - F(l + at) \} = \frac{\beta'}{a'} \{ f'(l - a't) - F'(l + a't) \};$$

or, putting

$$c = \frac{\beta'}{\beta} \times \frac{a}{a'}$$

$$f(l - at) - F(l + at) = c \{ f'(l - a't) - F'(l + a't) \}; \quad (D).$$

134.

Suppose the whole extent of both media to be initially at rest (and, therefore, $v = v' = s = s' = 0$), for every value of x , but those comprised within the region of the primitive disturbance ($x = \pm a$), supposed very minute and situated at the origin of the x , we shall have then

$$f'(x) = 0 \text{ and } F'(x) = 0 \text{ from } x = l \text{ to } x = \infty,$$

and since t is necessarily positive, and also a' , therefore

$$f'(l + a't) = 0 \text{ and } F'(l + a't) = 0.$$

The equations C and D then become

$$\begin{aligned} f(l - at) + F(l + at) &= f'(l - a't) \\ f(l - at) - F(l + at) &= c f'(l - a't); \end{aligned}$$

and, consequently,

$$f(l - at) - F(l + at) = c \{ f(l - at) + F(l + at) \}. \quad (E).$$

135.

Division of the pulse on encountering an obstacle.

Now this equation is equivalent to $as = cv$; x being supposed $= l$, (equation A.) Consequently, whatever be the motion of the first medium, the existence of a second, in contact with it, establishes at their point of junction a relation between the velocity v and the condensation s of its terminal stratum, which is incompatible with the condition $as = v$, (unless in the very peculiar case where $c = 1$.) which we have shown to be essential to the total propagation of the pulse forward. It will, therefore, divide itself conformably to what was said in Art. 129, and a portion will run back in the first medium and cause an Echo.

136.

But is propagated singly beyond it.

In the second medium, on the other hand, we have constantly $x > l$, and, therefore, $x + a't > l$, so that $F'(x + a't) = 0$, and, therefore, the equations (B) give

$$v' = a's' = f'(x - a't); \quad (F).$$

The condition of the single propagation of the pulse onward in this medium $v' = a's'$ being therefore satisfied, no further subdivision of the pulse will take place, and each particle of the second medium will be agitated once and no more. The reader who would pursue this discussion (a very delicate one) further, is referred to M. Poisson's Memoir, *Sur le Mouvement des Fluides Elastiques dans des Tuyaux Cyliindriques*, *Mém. Acad. Par.* 1818, 1819. See also a very curious Paper by Euler, *Sur la Propagation du Son et sur la Formation de l'Echo*, *Mém. Acad. Berlin*, 1765, p. 355; where he shows how an echo may be formed at the open mouth of a tube, by the mere conditions to be satisfied by the arbitrary functions, and without any reflexion properly so called. It is enough that the condition $as = v$ should be disturbed (as it will by the sudden breaking off of the pipe) to cause an echo. See also Weber, *Wellenlehre*, § 276, who shows how this disturbance takes place, owing to the greater freedom of motion suddenly attained by the particles when the pulse reaches the free air.

137.

Oblique refraction of Sound.

If we suppose a plane wave of indefinite extent to fall obliquely on the surface of a second elastic medium, each particle of this surface may be regarded as being put in agitation by it and becoming a separate and independent centre, from which spherical waves originate and are thence propagated in either medium with the velocity peculiar to it. Now, if we investigate the surfaces which in either medium are common tangents to all these spheres, and which, therefore, will be the form of the general or resulting waves in each, we shall find them to be planes; that in the medium of incidence being inclined to the surface at an angle equal to that made with it by the incident wave, and that in the other medium at an angle whose cosine is to the cosine of that made with it by the incident wave as the velocity of propagation of the wave in the first medium to that in the second. For the demonstration of these propositions we shall refer to our article on LIGHT, Art. 586. Thus the reflexion and refraction of Sound at oblique surfaces obeys the same geometrical laws with those of Light. The observation of Messrs. Colladon and Sturm, above cited, Art. 95, shows that this analogy extends to the case of oblique internal reflexion at the surface of a less elastic medium, which, at a certain incidence, becomes total.

Internal total refraction.

PART II.

OF MUSICAL SOUNDS.

§ I. *Of the Nature and Production of Musical Sounds.*

- Sound.** EVERY impulse mechanically communicated to the air, or other sonorous medium, is propagated onward by its elasticity as a wave or pulse; but, in order that it shall affect the ear as an audible sound, a certain force and suddenness is necessary. Thus the slow waving of the hand through the air is noiseless, but the sudden displacement and collapse of a portion of that medium by the lash of a whip produces the effect of an explosion. It is evident that the impression conveyed to the ear will depend entirely on the nature and law of the original impulse, which being completely arbitrary, both in duration, violence, and character, will account for all the variety we observe in the continuance, loudness, and quality of Sounds. The auditory nerves, by a delicacy of mechanism, of which we can form no conception, appear capable of analyzing every pulsation of the air, and appreciating immediately the law of motion of the particles in contact with the ear. Hence all the qualities we distinguish in Sounds—grave or acute, smooth, harsh, mellow, and all the nameless and fleeting peculiarities which constitute the differences between the tones of different musical instruments—bells, flutes, cords, &c., and between the voices of different individuals or different animals.
- Every irregular impulse communicated to the air produces what we call a *noise*, in contradistinction to a musical Sound. If the impulse be short and single we hear a crack, bounce, or explosion; yet it is worthy of remark, as a proof the extreme sensibility of the ear, that the most short and sudden noise has its peculiar character. The crack of a whip, the blow of a hammer on a stone, and the report of a pistol, are perfectly distinguishable from each other. If the impulse be of sensible duration and very irregular we hear a crash, if long and interrupted, a rattle or a rumble, according as its parts are less or more continuous, and so for other varieties of noise.
- The ear, like the eye, retains for a moment of time, after the impulse on it has ceased, a perception of excitement. In consequence, if a sudden and short impulse be repeated beyond a certain degree of quickness, the ear loses the intervals of silence and the Sound appears continuous. The frequency of repetition necessary for the production of a continued Sound from single impulses is, probably, not less than sixteen times in a second, though the limit would appear to differ in different ears.
- If a succession of impulses occur, at exactly equal intervals of time, and if all the impulses be exactly similar in duration, intensity, and law, the Sound produced is perfectly uniform and sustained, and has that peculiar and pleasing character to which we apply the term musical. In musical Sounds there are three principal points of distinction, the pitch, the intensity, and the quality. Of these, the intensity depends on the violence of the impulses, the quality on their greater or less abruptness, or, generally, on the law which regulates the excursions of the molecules of air originally set in motion. The pitch is determined solely by the frequency of repetition of the impulse, so that all Sounds, whatever be their loudness or quality, in which the elementary impulses occur with the same frequency, are at once pronounced by the ear to have *the same pitch*, or to be *in unison*. It is the pitch only of musical Sounds whose theory is susceptible of exact reasoning, and on this the whole doctrine of harmonics is founded. Of their qualities and the molecular agitations on which they depend, we know too little to subject them to any distinct theoretical discussion.
- The means by which a series of equidistant impulses, or, to speak more generally, by which an initial impulse of a periodical nature (*i. e.* capable of being represented by a periodical function) can be produced mechanically, are extremely various. Thus, if a toothed wheel be turned round with uniform velocity, and a steel spring be made to bear against its circumference with a constant pressure, each tooth, as it passes, will receive an equal blow from the spring, and the number of such blows per second will be known, if the velocity of rotation and number of teeth in the wheel be known.
- The late Professor Robison devised an instrument in which a current of air passing through a pipe was alternately intercepted and permitted to pass by the opening and shutting of a valve or stopcock. When this was performed with sufficient frequency (which could only be done, we presume, by giving a rapid rotatory motion to the stopcock by wheelwork) a musical tone was produced, whose pitch became more acute as the alternations became more frequent. This is precisely the principle of the Sirene of Baron Cagniard de la Tour. In this elegant instrument the wind of a bellows is emitted through a small aperture, before which revolves a circular disc, pierced with a certain number of holes arranged in a circle concentric with the axis of rotation, exactly equidistant from each other, and of the same size, &c. The orifice, through which the air passes, is so situated, that each of these holes, during the rotation of the disc, shall pass over it and let through the air, and the disc is made to revolve so near the orifice, that in the intervals between the holes it shall act as a cover and intercept the air. If the holes be pierced obliquely, the action of the current of air alone will set the disc in motion: if perpendicular to the surface, the disc must be moved by wheelwork, by means of which its velocity of rotation is easily regulated and the number of impulses may be exactly counted. The Sound produced is clear and sweet, like the human voice. If, instead of a single aperture for transmitting the air, there be several, so disposed in a circle of equal dimension with that in which the holes of the disc are situated, that each shall be
- Part II.**
138. Perception of Sounds in general.
139. Noise, as distinguished from musical Sound.
140. Continuous Sound.
141. Periodical impulses produce musical Sounds. Pitch.
142. Means of producing periodical impulses.
143. The Sirene.

Sound. opposite one corresponding hole when at rest, these will all form Sounds of one *pitch*, and being heard together will reinforce each other. The Sirene sounds equally when plunged in water, and fed by a current of that fluid, as in air; thus proving that it is the number of impulses alone, and nothing depending on the nature of the medium in which the Sound is excited, that influences our appreciation of its pitch.

144. In general, whatever cause produces a succession of equidistant impulses on the ear, causes the sensation of a musical Sound, whether such periodicity be a consequence of periodical motions in the origin of the Sound, or of the mode in which a single impulse is multiplied in its conveyance to the ear. For example, a series of broad palisades set edgewise in a line directed from the ear, and equidistant from each other, will reflect the Sound of a blow struck at the end of the line nearest the auditor, producing a succession of echos, which (by reason of the equidistance of the palisades) will reach his ear at equal intervals of time, $\left(= 2 \times \frac{\text{distance of palisades}}{\text{velocity of Sound}}\right)$, and will

therefore produce the effect of a number of single impulses originating in one point. Thus a musical note will be heard whose pitch corresponds to a number of vibrations per second, equal to the quotient of the velocity of Sound by twice the distance of the palisades.

145. A similar account may be given of the singing Sound of a bullet, or other missile, traversing the air with great rapidity. The bullet being in a state of rapid rotation, and not exactly alike in all its parts, presents, periodically, at equal intervals of time and space, some protuberance or roughness first to one side, then to the other. Thus an interruption to the uniformity of its mode of cutting through the air is periodically produced, and reaches the ear in longer or shorter equal intervals of time, according as the rectilinear velocity of the bullet bears a greater or less ratio to the velocity of its rotation about its axis.

146. The echos in a narrow passage, or apartment of regular figure, being regularly repeated at equal very small intervals, always impress the ear with a musical note; and this is, no doubt, one of the means which blind persons have of judging of the size and shape of any room they happen to be in. But the most ordinary ways in which musical Sounds are excited and maintained consist in setting in *vibration* elastic bodies, whether flexible, as stretched strings, or membranes; or rigid, as steel springs, bells, glasses, &c. or columns of air of determinate length enclosed in pipes. All such vibrations consist in a regular alternate motion to and fro of the particles of the vibrating body, and are performed in strictly equal portions of time. They are, therefore, adapted to produce musical sounds by communicating that regularly periodic initial impulse to the aerial molecules in contact with them which such sounds require. We shall, therefore, proceed to consider more particularly the principal of these modes of production; but especially, at present, the first and last, being the most simple cases.

§ II. Of the Vibrations of Musical Strings or Cords.

147. If a string, or wire, be stretched between two fixed pins, or supports, and then struck, or drawn a little out of its straight line, and suddenly let go, it will vibrate to and fro, till its own rigidity, and the resistance of the air, reduce it to rest; but if a *bow* (which is an instrument composed of a bundle of fibres of horse hair, loosely stretched, and rendered adhesive by rubbing with rosin) be drawn across it, the vibrations are continually renewed, and may be maintained for any length of time, and a musical Sound is heard corresponding to the rapidity of the vibration.

148. The mathematical theory of the vibrations of a stretched cord is remarkable, in an historical point of view, as having given rise to the first general solution of an equation of partial differences; and led geometers to the consideration of the nature and management of the arbitrary functions which enter into the integrals of these equations. Such functions, as we have seen, enter into the general expressions for the motion of the air in Sound; and such, as we shall presently show, into that of the molecules of a vibrating cord; and a long and lively discussion, on the degree of generality which ought to be attributed to them, soon arose between Euler, D'Alembert, D. Bernouilli, and Lagrange. It is not, however, our intention in this Article to enter into any points of historical detail, and we shall content ourselves with a reference to the principal Memoirs, &c. on the subject, which the reader may consult for himself; while we proceed to give such a view of the subject as is consistent with the present state of knowledge on this delicate point, and sufficient for the purpose we have in hand. See Taylor, *De Motu Nervi Tensi*, *Phil. Trans.* 1713-26; D'Alembert, *Mém. Acad. Berl.* 1747; Ditto, 1753; Ditto *Opuscles*, tom. i.; Euler, *Mém. Acad. Berl.* 1753; Daniel Bernouilli, Ditto; Lagrange, *Miscellanea Taurin.* vol. i. See also Sauvour, *Mém. Acad.* for 1713, p. 324; J. Bernouilli, on Vibrating Cords, *Petrop. Comm.* iii. 13; Daniel Bernouilli, Ditto, p. 62; Ditto, on Vibrations of Unequal Cords, *Acad. Berl.* 1765, p. 81; Ditto, on Vibrations of Compound Cords, *N. Comm. Petrop.* xvi. 257; Euler, *Acad. Berl.* 1748, p. 69; Ditto, Ditto, 1765, p. 307, 335; Ditto, on Unequal Vibrating Cords, *N. Comm. Petrop.* xvii. 381; Ditto, 1780, iv. ii. 99.

149. Let M N (fig. 9) be a cord maintained by any means in a constant state of equal tension throughout, and disturbed by any external cause from its rectilinear position, and then left to take its own form and motion in consequence of its tension; its gravity, however, being neglected. Let M A B C D N be the figure of the cord after the lapse of any time *t* from the initial disturbance; respecting which we will only suppose that the distance of all its points from the axis V T (the undisturbed rectilinear position of the cord) is extremely small; so that in this theory, as in that of the sonorous vibrations of the air, we concern ourselves only with such excursions of the vibrating molecule as may be considered infinitely minute. Let A B C D be points of the cord infinitely near each other; and erecting the ordinates A P, B Q, C R, D S, and drawing A a, B b, C c, D d, parallel to V T

Sound.

Part II.

put $VP = x$, $VQ = x'$, $VR = x''$, &c. and $AP = y$, $BQ = y'$, $CR = y''$, &c. Let the tension of the cord at rest be represented by c , which (since the cord is infinitely little disturbed from its position of repose) will also be its tension in its disturbed state; and will in this, as in the former state, be uniform over its whole length, the curvature being evanescent. The point B of the cord then will be solicited towards the axis by the tension c applied at B, and acting in the direction BA, and whose resolved value is, therefore,

$$c \cdot \frac{BA}{BA} = c \cdot \frac{dy}{\sqrt{dx^2 + dy^2}} = c \cdot \left\{ \frac{dy}{dx} - \frac{1}{2} \left(\frac{dy}{dx} \right)^3 - \&c. \right\} = c \cdot \frac{dy}{dx},$$

neglecting the higher powers of the quantity $\frac{dy}{dx}$, which (being the tangent of the inclination of the element AB to the axis) is infinitely small. Similarly the point B will be solicited from the axis by the tension c applied at B in the direction BC, whose resolved part in the direction of the ordinate is equal to $c \cdot \frac{dy'}{dx'}$. The resolved parts in directions parallel to the axis, being equal and parallel, destroy each other; consequently, the whole force applied at B will be $c \cdot \left(\frac{dy'}{dx'} - \frac{dy}{dx} \right)$; or, supposing dx constant, $c \cdot \frac{d^2y}{dx^2} dx$, tending to increase the value of y .

Now the motion of the cord will be the same, whether we regard it as a continuous mass, or compound of detached particles situated at A, B, C, D, &c. and connected by filaments AB, BC, &c. without weight. Thus

150.

Its equation derived and integrated.

at B we may conceive to be placed a weight equal to $\frac{AB + BC}{2}$, at C the weight $\frac{BC + CD}{2}$, and so on, that

is, neglecting $\left(\frac{dy}{dx} \right)^2$, simply a constant weight dx in each point. This, then, is the mass to be moved by the

moving force $c \cdot \frac{d^2y}{dx^2} dx$, and the accelerating force is, therefore, simply $c \cdot \frac{d^2y}{dx^2}$. Hence, calling t the time,

and regarding dt as constant as well as dx , x and t being independent variables, and putting $2g = 9^{\text{met}}.8088 = 32^{\text{feet}}.18169$, or $g = 16^{\text{feet}}.090845$, we have

$$\frac{d^2y}{dt^2} = 2gc \cdot \frac{d^2y}{dx^2};$$

or, putting

$$2gc = a^2; \quad a = \sqrt{2gc},$$

$$\frac{d^2y}{dt^2} = a^2 \cdot \frac{d^2y}{dx^2}.$$

This equation is precisely similar to that above obtained for the propagation of Sound along a cylindrical pipe, and its integral will, of course, be of the same form, viz.

$$y = F(x + at) + f(x - at).$$

The determination of the arbitrary functions in this equation will depend on the conditions we may set out from. 151.

Now, first, when the cord is supposed to be of indefinite length, and the part initially disturbed to be comparatively very small; and having an indefinite undisturbed portion on either side. In this case, it is evident by the very same reasoning as that of Articles 63 and 64, that a pulse or undulation will run out both ways

Determination of the arbitrary functions.

along the cord from the point of initial disturbance, with a velocity represented by $a = \sqrt{2gc}$, every molecule of the cord being once agitated during the time the pulse runs over it, and no more. Moreover, a condition similar to that which ensures the single propagation of the pulse when once it has proceeded beyond the limits of the initial disturbance ($x = \pm a$) in the theory of Sound, holds good in the present case; for we have

$$\frac{dy}{dx} = F'(x + at) + f'(x - at)$$

$$\frac{dy}{dt} = aF'(x + at) - af'(x - at).$$

So that on the positive side of the x , when $x > a$, and therefore $x + at = a$, and $F'(x + at)$ and $F'(x + at)$

= 0, we shall have

$$a \frac{dy}{dx} = - \frac{dy}{dt};$$

Condition for the single propagation of a wave.

which expresses that the tangent of the obliquity of the cord to the axis in its disturbed state, at any point, is proportional to the absolute velocity of that point in its motion, or putting $\theta =$ angle B A a,

$$a \cdot \tan \theta = -v;$$

and when this condition ceases to hold good, as it does when the pulse encounters an obstacle either fixed or less movable than the rest of the cord, it will be either wholly reflected, or divide itself into two, one running back, and producing a species of imperfectly echoed or reflected wave, just as in the theory of Sound.

Reflexion of the wave by an obstacle, as in Sound.

Sound.

Since, in the above investigation, c represents a *force* equal to the tension on the same scale that dx represents a *weight* equal to that of the element dx , we have

152.

Velocity with which a wave runs along a stretched cord.

$$\text{weight of } dx : \text{tension} :: dx : c.$$

Hence c represents the length of a portion of the cord whose weight is equal to the tension, and $\sqrt{2gc}$ the velocity which would be acquired by a body falling freely by gravity through that length. Hence this theorem, *The velocity of a pulse, or undulation propagated along a tended cord, is equal to that which a heavy body would acquire by falling freely through the length of a portion of the cord whose weight is equal to its tension.*

153.

Case when one extremity of the cord is fixed.

Let us next suppose the cord attached at one of its extremities to an immovable point, and let the undulation be supposed to reach this point, at which suppose $x = l$, then, whatever be the value of t , $y = 0$, when $x = l$. So that we must have

$$F(l + at) + f(l - at) = 0.$$

Since at may have any positive value, and since on the positive side of the x (at which we have supposed the fixed end situated) $x < l$, therefore $l - x$ is in all cases positive, and therefore may be one of the values of at . We may substitute, then, $l - x$ for at in this equation, when we get for positive values of x less than l , and for all negative ones

$$F(2l - x) + f(x) = 0; \quad \text{or } F(2l - x) = -f(x). \quad (p)$$

Now, in general,

$$y = F(x + at) + f(x - at).$$

If, then, we make $at = x + \omega$, where ω is any quantity between $+a$ and $-a$, at which values of x both $f(x)$ and $F(x)$ may be supposed to vanish,

$$y = F(2x + \omega) + f(-\omega),$$

and if we make $at = 2l - x + \omega$, we have

$$y = F(2l + \omega) + f(2x - 2l - \omega),$$

but by (p)

$$F(2l + \omega) = -f(-\omega),$$

so that for the latter value of t we have

$$y = -f(-\omega) + f(2x - 2l - \omega).$$

Now since when $t = 0$, we have $y = F(x) + f(x)$ and $\frac{dy}{dx} = F'(x) + f'(x)$; $\frac{dy}{dt} = a \{ F'(x) - f'(x) \}$,

all these values must vanish unless x lies between the limits $+a$ and $-a$. Consequently, for all values but those comprised within such limits, we now have $F(x) = 0$ and $f(x) = 0$. From the above equations, then, supposing $x > a$, or $< -a$; and, therefore, $F(2x + \omega) = 0$, and $f(2x - 2l - \omega) = 0$, we see that for values of at between $x + a$ and $x - a$, y will have real values; and that when at attains any value between $2l - x + a$, and $2l - x - a$, y will again have real values, the same as the former, only with contrary signs. Thus the reflected pulse runs back with the same velocity as the direct, and is in all respects similar and equal to it, only that it lies on the opposite side of the axis. A reasoning precisely similar applies to the case of an aerial pulse reflected from the bottom of a stopped pipe, supposed perfectly rigid.

154.

Case when both ends are fixed.

155.

In this case the cord assumes a continued vibratory motion.

If the cord be fixed at both ends, the two pulses into which the initial pulse has separated itself, will each be totally reflected, and will run along the whole length, being reflected again at the other end, and thus run backwards and forwards for ever, at least if we neglect the effect of the stiffness of the cord and resistance of the air; crossing each other at each traverse.

Suppose the whole length of the cord to be $l + l' = L$, of which l lies on the positive, and l' on the negative side of the origin of the x . That portion of the subdivided primitive pulse which runs towards the positive side

of the x will describe the length l in a time $= \frac{l}{a}$; being then reflected it will describe the whole length $l + l'$ in a time $\frac{l + l'}{a}$; and being again reflected, it will describe l' in a time $\frac{l'}{a}$, so that after a time

$$= \frac{l}{a} + \frac{l + l'}{a} + \frac{l'}{a} = \frac{2L}{a},$$

it will reach its first starting point; and having been twice inverted by reflexion, will lie now on the same side of the axis it originally was. Similarly, the negative portion of the original pulse will describe l' , $l' + l$, and l , and reach its starting point after two reflexions in the time

$$\frac{l'}{a} + \frac{l' + l}{a} + \frac{l}{a} = \frac{2L}{a}.$$

the same as the other, and will also have recovered its original situation with respect to the axis. Thus at the end of this time the two pulses will precisely reunite, and constitute a compound pulse in all respects similar to the initial impulse. The state of the cord, then, after the lapse of the time $\frac{2L}{a}$, will (abstracting the effects of

resistance, &c.) be precisely what it was at first; and so again, after the lapse of time $\frac{4L}{a}$, $\frac{6L}{a}$, &c. the same state will recur, so that if left to itself it will continue to vibrate for ever.

Sound.

Thus we see that what in an indefinite cord was merely a pulse running along it and never returning, becomes, by the reaction of the fixed extremities of a finite one, a regular vibration, in which each molecule repeats its motion to and fro on either side of the axis, at equal intervals, for ever. In the foregoing reasoning no particular assumption has been made respecting the value of a . It has not been supposed small with respect to l, l' , and, consequently, the above conclusion applies equally to the case where the initial disturbance is confined to a minute portion of the cord, and where a large portion, or even its whole length, is disturbed at once. Only in the former case the motions of the individual molecules of the cord will be performed by starts interrupted by intervals of absolute rest in the axis. In the latter there will be no moments of rest but those when the direction of the motion changes at the extreme points of their excursions.

Hence we conclude that when a stretched cord, whose length = L , is struck, or forcibly drawn out of its straight situation into any form and let go, it will continue to vibrate to and fro, and that the time of one complete vibration, after which it resumes its initial state, is represented by $\frac{2L}{a} = \frac{2L}{\sqrt{2gc}}$, being equal to the time of a

pulse running over double the length of the cord, or to the time in which a body would describe such double length with the velocity acquired by falling down a height equal to the length of a portion of the cord whose weight is the tension.

Hence the times of vibration of different cords are, as their lengths directly, and the square roots of the tending forces inversely, and the number of vibrations, *dato tempore*, as the lengths inversely, and the square root of the tensions directly.

The equations which express the conditions arising from the immobility of the ends of the cord so far limit the arbitrary functions F and f , that when the figure of the cord between its two extremities is given it may be prolonged beyond them to any extent. To show this, let y_1 represent the ordinate $P_1 M_1$ of the curve supposed to be continued beyond B , one of the fixed extremities, at a distance, $B P_1$, beyond that end equal to $B P$, the distance from it of the ordinate y , and, for simplicity, suppose $l' = 0$, or let the origin of the x be at the other fixed extremity, A , (fig. 10.) Then we have

$$\begin{aligned} y &= F(x + at) + f(x - at), \\ y_1 &= F(2L - x + at) + f(2L - x - at). \end{aligned}$$

Now the condition of Art. 153, derived from the fixity of the point B , viz.

$$F(2L - x) + f(x) = 0,$$

gives, if we write for x successively $x - at$ and $2L - x - at$, the following equations,

$$\begin{aligned} F(2L - x + at) + f(x - at) &= 0, \\ F(x + at) + f(2L - x - at) &= 0, \end{aligned}$$

whose sum is no other than

$$y + y_1 = 0, \text{ or } y_1 = -y.$$

Thus we see that the curve $A M B$ will be continued beyond B by merely reversing it from right to left and transferring it to the other side of the axis. Again, if we put y_2 for the ordinate $P_2 M_2$ at a distance = x beyond C , we have

$$y_2 = F(2L + x + at) + f(2L + x - at),$$

and

$$\left. \begin{aligned} F(2L + x + at) + f(-x - at) &= 0 \\ F(-x + at) + f(2L + x - at) &= 0 \end{aligned} \right\}$$

On the other hand, the condition of the immobility of the point A gives, as we have seen,

$$F(x) + f(-x) = 0,$$

in which, writing successively for x , $+x + at$, and $-x + at$, we get

$$\left. \begin{aligned} F(x + at) + f(-x - at) &= 0 \\ F(-x + at) + f(x - at) &= 0 \end{aligned} \right\}$$

and subtracting the sum of these from that of the two former, we find ultimately

$$y_2 - y = 0, \text{ or } y_2 = y,$$

so that the portion of the curve $C M_2 D$ is the very same with the first portion $A M B$. And thus we may go on as far as we please, repeating the same curve alternately in a direct and reverse position, and the same manifestly holds good on the other side of the point A .

A very simple consideration will show that such ought to be the case; for if we conceive two equal and similar cords, $A M B$, $B M_1 C$, (fig. 10,) both attached to the same point, B , and vibrating simultaneously, the strain on B , from both their tensions, will be always equal and opposite, provided the curves be so related as above described, and B , therefore, will be retained in equilibrium, independently of its attachment to any extraneous body, so that were it detached, or if the two cords, instead of being fixed to one immovable point, were merely linked together at B , so as to form one cord of double the length, their vibrations would be the same.

Some curious and important consequences follow from this. And, first, a cord, although vibrating freely, may yet have any number of points, equally distributed at aliquot parts of its whole length, which never leave the axis, and between which the vibrating portions are equal and similar, and lie alternately above and below the

Part II.

156.

Passage from a transitory pulse to a permanent vibration.

157.

Time of vibration of a stretched cord.

158.

In different cords.

159.

Prolongation of the figure of the cord on either side of its fixed extremities. Fig. 10.

160.

Origin of nodal points.

See fig. 10.

161.

A cord may have any number of them.

Sound. axis, and in reversed positions as to right and left. Such points of rest are called nodes or nodal points, the intermediate portions which vibrate are termed *bellies* or *ventral segments*.

162. Secondly, if a string in the act of vibration be touched in any point so as to reduce that point to rest and retain it in the axis, then if, after the contact, it vibrate at all, it will divide itself into a certain number of ventral parts similar and equal to each other and separated by nodes, and each of these will vibrate as if the others had no existence, but instead the nodes were fixed points of attachment. Hence, if L be the whole length of a cord, n the number of ventral segments into which it divides itself, and, therefore, $n - 1$, the number of its nodes, the time of one complete vibration (going and returning) will be $\frac{2L}{n\sqrt{2gc}}$ and the number of vibrations per second will be

163. Experience confirms this. If the string of a violin, or violoncello, while maintained in vibration by the action of the bow, be lightly touched with the finger or a feather exactly in the middle, or at one-third of the length, it will not cease to vibrate, but its vibrations will be diminished in extent and increased in frequency, and a note will become audible, fainter but much more acute than the original, or, as it is termed, the *fundamental* note of the string, and corresponding in the former case to a double, in the latter to a triple rapidity of vibration. The note heard in the former case being the octave, in the latter the twelfth, above the fundamental tone (See Index, *Musical Intervals*.) If a small piece of light paper, cut into the form of an inverted V, be set astride on the string, it will be violently agitated, and, probably, thrown off when placed in the middle of a ventral segment, while at a node it will ride quietly as if the string were (as it really is at those points) at perfect rest. The Sounds thus produced are termed harmonics.

164. But, further, any number of the different modes of vibration, of which a cord is thus susceptible, may be going on *simultaneously*, or be, as it were, superposed on each other. This is a consequence of the principle in mechanics of "the superposition of small motions," which, when the excursions of the parts of a system from their places of rest are infinitely small, admits of any or all the motions of which, from any causes, they are susceptible, to go on at once without interfering with or disturbing each other. In the particular case before us it is easily shown, for since the general integral of the equation

$$\frac{d^2 y}{dt^2} = a^2 \cdot \frac{d^2 y}{dx^2} \quad \text{is} \quad y = F(x + at) + f(x - at),$$

where F and f denote arbitrary functions, we may suppose

$$F(x) = F_1(x) + F_2(x) + F_3(x) + \&c.$$

$$f(x) = f_1(x) + f_2(x) + f_3(x) + \&c.$$

where $F_1, F_2, \&c.$, and $f_1, f_2, \&c.$, denote functions equally arbitrary, and we get

$$y = \{F_1(x + at) + f_1(x - at)\} + \{F_2(x + at) + f_2(x - at)\} + \&c.$$

Now each of the expressions within brackets is the integral of an equation exactly similar to the original one. Therefore, if we put

$$\frac{d^2 y_1}{dt^2} = a^2 \cdot \frac{d^2 y_1}{dx^2}; \quad \frac{d^2 y_2}{dt^2} = a^2 \cdot \frac{d^2 y_2}{dx^2}; \quad \&c.$$

we shall have

$$y = y_1 + y_2 + y_3 + \&c.$$

Thus, if the several particular modes of vibration, $y = y_1, y = y_2, \&c.$, be possible, $y = y_1 + y_2 + \&c.$ will also be possible: the ordinate of the curve into which the cord at any moment forms itself in virtue of the compound vibration will be the sum (algebraically understood) of the ordinates it would have in virtue of each simple one, separately: the compound curve will be formed by first constructing on the abscissa, as an axis, any one of the simple ones, then on that curve, as an abscissa, any other, on the new curve thence arising any other, and so on.

165. Hence it is evident that if we suppose the curve, whose ordinate is y_1 , to be of the form, fig. 11, (a) having no node, and that, whose ordinate is y_2 , to have, for instance, one node, as fig. 11, (b) the corresponding modes of vibration, when coexisting, will produce a curve, such as (c). On these we may superpose a third mode of vibration, where the string divides itself into three ventral segments, as (d), and the result will be a curve, such as (e), and so on to any extent. The reader may exercise himself in tracing the variations of form in these curves as they go through the several phases of their periodic excursions during one complete period of a vibration of the whole string as one cord.

166. Experience again confirms this result of theory. It was long known to musicians that, besides the principal or fundamental note of a string, an experienced ear could detect in its Sound when set in vibration, especially when very lightly touched in certain points, other notes, related to the fundamental one by fixed laws of harmony, and which are called, therefore, harmonic sounds. They are the very same which, by the production of distinct nodes, may be insulated, as it were, and cleared from the confusing effect of the coexistent Sounds, as in Art. 163. They are, however, much more distinct in bells, and other sounding bodies, than in strings, in which only delicate ears can detect them.

167. The monochord is an instrument well adapted to exhibit these and all other phenomena of vibrating strings. It is nothing more than a single string of catgut fixed at one end immovably, and at the other strained over a well-defined edge, which effectually terminates its vibrations, either by a known weight or by screws. A similar well-defined edge is also interposed between its fixed end and the vibrating portion, and the interval between the two edges is graduated into aliquot parts, or in any other convenient way, and it is provided with a movable

Fig. 11.
Curves arising from superposition of several coexistent vibrations.

166.
Harmonic Sounds heard with the fundamental Sound.

167.
The monochord.

bridge, or piece of wood capable of being placed at any division of the scale, and abutting firmly against the string so as to stop its vibrations, and divide it into two of equal or unequal lengths, as the case may be.

By the aid of this instrument we may ascertain the number of vibrations which belongs to any assigned musical note, or which correspond to the notes of any musical instrument, as a piano-forte, &c. For when we have ascertained the weight of a known length of the catgut, of which the string is formed, and the weight which must be applied to stretch the cord, so as to make its fundamental tone coincide with any given note, (as the middle C of

a piano-forte,) then by the formula $\frac{\sqrt{2gc}}{2L}$ we know the number of complete vibrations going and returning, and

by the formula $\frac{\sqrt{2gc}}{L}$ the number of oscillations from rest on one side of the axis to rest on the other, that

is, the number of impulses made on the ear per second corresponding to that fundamental tone. To determine the same for any note *sharper*, *higher*, or more *acute* than the fundamental note, we have only to apply the bridge, and move it backwards and forwards till the sound of the vibrating part of the string is *in unison* with that of the note to be compared, of which the ear judges with the greatest precision; then if the length of this part, read off on the divided scale, be called l , the number of its vibrations per second will be to that of the whole string $L :: L : l$, and is therefore known.

The production of harmonic Sounds from cords, and their division into aliquot parts, was first noticed, in 1673, by Wallis, (*Opera J. Wallisii*, fol. ii. p. 466, cap. cvii.) but the subject remained unattended to till taken up by Sauveur, in a valuable Memoir, published among those of the French Academy for 1701, which first put this part of the doctrine of Acoustics in a clear point of view. The contact of a solid obstacle is not the only means of producing them. If two cords equally tended, and in all other respects similar, but one only half, one third, or other aliquot part of the length of the other, be placed side by side, and the shorter be struck or sounded, the vibration will be communicated to the longer by the intervention of the air, which will thus at once be thrown into a mode of vibration, in which the whole length is divided into ventral segments, each equal to the shorter string. To understand how this may happen, let us conceive first two strings of equal length, one at rest the other vibrating, and let them be placed parallel, and side by side, then the sonorous pulses diverging at any instant from each point of the moving string, will arrive at once at each corresponding point of the other. The aerial molecules in their progress, while condensed, will press on the string and give it a very slight motion in their own direction; in their retreat they will be followed by the string, whose vibrations by hypothesis are synchronous with their own, but it will not follow them so fast as they retreat, and it will be, therefore, urged and accelerated by those behind. It will, however, come to rest, in its furthest point of excursion, at the same time with the aerial molecules, when its elasticity will begin to urge and accelerate it in the contrary direction. But now also the direction of the motion of the air has changed, and again conspiring with that of the cord still continues to accelerate it, and so on, till, after a very great number of repetitions of this process, the cord will be set in full vibration and will become itself a source of Sound. But its Sound will always be much fainter than that of the original vibrating cord, for this reason, *viz.* that its acquired motion is perpetually dissipated, *laterally*, into the surrounding air, for no cord is so exactly uniform, or so equally tended in every part of its transverse section, that it *can* vibrate rigorously in one plane. Hence it will inevitably begin to rotate, or to describe vibrations whose plane is continually shifting, (see Art. 177,) and thus it will throw off laterally a great part of the motion it receives from the air; just as a body exposed to the radiation of a hot fire never acquires a temperature equal to that of the fire, part of the heat communicated being dissipated by lateral radiation.

Just as a small pull, repeated exactly in the time of its natural swing, will raise a great bell, or a trifling impulse a heavy pendulum, so the molecules of the air, in a state of sonorous vibration, will impress on any body capable of vibrating in their own time an actual vibratory motion, and if a body be susceptible of a number of modes of vibration performed in different times, *that mode only will be excited which is synchronous with the aerial pulsations*. All other motions, though they may be excited for a moment by one pulsation, will be extinguished by a subsequent one. Hence, if two cords have any mode of vibration in common, that mode may be excited by sympathy in either of them when the other is sounded, and that only. For example, if the length of one cord be to that of the other as 2 : 3, and if either be set vibrating, the mode of vibration, corresponding to a division of the former into two, and of the latter into three ventral segments, will, if it exist in the one, be communicated by sympathy to the other. Nay, if it do not originally exist, it will, after a while, establish itself; for all accidental circumstances which may favour such a division have their effects, however minute, continually preserved and accumulated, till at length they become sensible.

In the vibrations of cords, which from their small surface can receive but a trifling impulse from the air, the Sounds and motions excited by this sort of sympathetic communication are feeble, but in vibrating bodies, which present a large surface, they become very great. It is a pretty well authenticated feat performed by persons of clear and powerful voice, to break a drinking-glass by singing its proper fundamental note close to it. (See Chladni, *Acoust.* § 224.) Looking-glasses also are said to have been occasionally broken by music, the excursions of their molecules in the vibrations into which they are thrown being so great as to strain them beyond the limits of their cohesion.

The coincidence of the theory above stated, of the propagation of a wave along a stretched cord, with experiment, has been put to careful trial by Weber. (See his *Wellenlehre auf Experimente Gegrundet*, 8vo. Leipsig. 1825, p. 460, a most instructive work.) He stretched a very equal and flexible cotton thread, 51 feet 2 inches in length, weighing 864 grains, horizontally, by a known weight. The thread was struck at 6 inches from one end at the instant of letting go a stop-watch of peculiar and delicate construction, marking thirds, (sixtieths of

Part II.

168.

Applied to measure the vibrations of a given note.

169.

Sympathy between cords which have a mode of vibration in common.

170.

Establishment of a mode of vibration by sympathy illustrated.

171.

Remarkable effects of sympathetic communication of vibration

172.

Numerical comparison of the above theory with experiment by Weber.

S. and. seconds,) whose motion was instantaneously arrested when the wave had run a certain number of times over the length of the string backwards and forwards. The mean of a great many observations, agreeing well with each other, gave as follows:—

Tension in Grains.	Length run over by the Wave.	Time of its description in Thirds.	Time of running over the length 102 f. 4 in. in Thirds, by observation.	The same time calculated from the formula $V = \sqrt{2gc}$.
10023	102f. 4in.	46	46	46·012
10023	204 8	92	46	46·012
10023	409 4	184	46	46·012
33292	409 4	99	24·72	25·246
69408	409 4	65	16·25	17·485

A completer coincidence could not have been wished for. The slight discrepancies may, perhaps, arise from the want of uniformity in the tension of so long a thread, which would, of course, form a catenary of sensible curvature. We should observe, however, that M. Weber has reckoned here the weight (864 grains) of the thread as part of the tension, a proceeding whose legitimacy may be questioned.

173. The way in which the permanent or vibratory oscillations of a cord arise by reflexion at its fixed extremities from a wave propagated along it progressively, may be rendered a matter of ocular inspection if we take a long and pretty thick cord, fasten it at one end, and holding the other in our hands, give it a regular motion to and fro, transverse to the length of the cord. Progressive waves will thus arise, which, as soon as they reach the fixed end, and are reflected, will be observed to interfere with those still on their way, and, as it were, to arrest them, producing a series of nodes and ventral segments, whose number will depend on the tension and frequency of the alternate motion communicated to the movable end. In this arrangement the continual periodic renewal of the primary impulse by the hand supplies the place of a reflecting obstacle at that end.

174. The pitch of the Sound of a vibrating string depends only on the number of vibrations made *dato tempore*, its quality will depend partly on the nature of the string, and especially on its equality of thickness, besides which, much may depend on the form and extent of the wave excited, or of the curve into which it is thrown. In instruments, like the violin or violoncello, played with a bow, or the guitar or harp, where the string is drawn softly out of its position and suddenly let go, this curve is, probably, single, and occupies the whole length of the string; but in the piano-forte, where the strings are struck, near one extremity, with a sharp sudden blow, there can be little doubt that the vibration consists in an elevation or bulge, more or less extensive, running backwards and forwards. Fig. 12 represents the different phases of a single complete vibration of a string so struck. The first wave (1) is a single elevation, it divides in (2) into two running contrary ways; in (3) that nearest the end A is reflected and takes a reversed position; in (4) they advance the same way towards B; in (5) the unreflected portion reaches B, is there reflected and reversed, as in (6). In (7) it meets and coincides with the former reflected portion, there forming a depression equal and similar to the original elevation in (1), and as far distant from the end B as the former from A. After this the same steps are repeated in the reverse direction, till the original elevation is reproduced again, as in (1). The waves, however, must be supposed to bear a much more considerable ratio to the whole string than in the figure. It is evident that the magnitude of this ratio must influence the quality of the tone, and thus a difference of character in the tone, according as the keys are struck with quick short brilliant blows, or gently pressed, and the duration of the contact of the hammers with the strings prolonged for an instant of time, giving rise to a more moderate but sustained *tenuto* effect, by bringing a larger portion of the string, or even the whole into motion at once.

Phases of a vibration traced.

175. But whether the portion disturbed at once be large or small, whether it occupy the whole string, or run along it like a bulge in its line, whether it be a single curve, or composed of several ventral segments with intervening nodes, we must never lose sight of the fact that the motion of a string with fixed ends is no other than an undulation or pulse continually *doubled back on itself* and retained constantly within the limits of the cord, instead of running out both ways to infinity.

176. It very seldom (for the reasons mentioned in Art. 170) can happen that the vibrations of a string actually lie in one plane. Most commonly they consist of rotations more or less complicated, except when produced by the sawing of a bow across the string, when they are forcibly limited to the plane of motion of the bow. The real form of the orbit described by any molecule may be made matter of ocular inspection, by letting the sun shine through a narrow slit so as to form a thin sheet of light. Let a polished wire be placed so as to penetrate this sheet perpendicularly to its plane, and the point where it cuts the plane will, at rest, be seen as a bright speck, but when set vibrating it will form a continued luminous orbit, just as a live coal whirled round appears as a circle of fire. Fig. 13 exhibits specimens of such orbits, observed by Dr. Young.

177. A very curious case of a mode of vibration, by which a string may be made to produce a Sound graver than its fundamental tone, is mentioned by M. Biot. If an obstacle be placed below the middle point of a vibrating string so as just to touch, but not to press against it, and the string be then drawn up vertically and let go, it will strike at every oscillation upon this obstacle, and bend over it, as in fig. 14, at every blow; thus resolving itself into two, of half the length. Thus the first semi-oscillation will be performed as a whole, the

Ocular evidence of the transition from progressive pulses to a permanent state of vibration.

Difference in the quality of the tone of stringed instruments whence arising.

Fig. 12.

Vibrations out of one plane. Forms of orbits described exhibited.

Fig. 13.

Sound. next as a subdivided string. Let unity represent the time of one complete oscillation from rest to rest of the whole string; then will the times in which the different phases of the motion now in question are performed be as follows: Part II.

From the position A B C to the straight line A C . . . = $\frac{1}{2}$.

From the position A C to the position A E D, D F C = $\frac{1}{4}$.

Back to the straight line = $\frac{1}{4}$.

Back to the original position A B C = $\frac{1}{2}$.

Sum = $\frac{1}{2} + \frac{1}{4} + \frac{1}{4} + \frac{1}{2} = \frac{3}{2}$.

Thus the interval between two consecutive blows made by the string on the bridge is $\frac{3}{2}$ of the time of oscillation of the string as a whole, from rest on one side of the axis to rest on the other, or of the impulses made by it on the ear when so vibrating. Hence, the blows on the bridge will be heard as a continued note, (though extremely harsh and disagreeable,) graver than that of the string vibrating as a whole, by the musical interval called a fifth. (See Index, *Musical Intervals*.)

§ III. Of the Vibrations of a Column of Air of Definite Length.

The general equation representing the motions of the molecules of a tended cord of indefinite length is, as we have seen, precisely similar in its form, and in that of its complete integral, to that of the particles of air in a sounding column. There subsists, of course, a perfect analogy between the two cases, and, *mutatis mutandis*, all propositions which are true of a vibrating cord are also true of a vibrating cylindrical volume of air.

Thus, if such a cylindrical column be enclosed in a pipe, whose length = $l + l' = L$, stopped at both ends by perfectly immovable stoppers, and if we suppose any single impulse communicated to one of its sections at the distance l from one of its extremities (A), this will immediately divide itself into two pulses running opposite ways; they will be totally reflected at the two extremities, the one, after describing the space l before and l' after reflexion, will meet the other which has described l' before and l after reflexion, at a distance = l from the other extremity B, and produce a compound agitation in the section at that place similar to the primitive disturbance; thence the partial pulses will again diverge, and after each undergoing another reflexion will again unite in their original point of departure, constituting a repetition of the first impulse, and so on, till the motion is destroyed by friction and by the imperfect fixity and rigidity of the stoppers, allowing some of it to pass into them and be lost at each reflexion.

But if the section first set in motion be maintained in a state of vibration synchronous with the return of the reflected pulse, it will unite with and reinforce it at every return, and the result will be a clear and strong musical Sound resulting from the exact combination of the original periodic impulse with all its echos. This will be transmitted through the pipe to the outer air, and thus dissipated and lost.

For simplicity, let us suppose the section primitively set in vibration and so maintained, to be situated just in the middle of the pipe. Then, when once the regular periodic pulsation of the contained air is established, it is evident that the motion of the column will consist of a constant and regular fluctuation to and fro within the pipe of the whole mass, the air being always condensed in one half of the pipe while it is rarefied in the other. The greatest excursions from their place will be made by the molecules in the middle, while those at the extremities, being constantly abutted against the stoppers, remain unmoved, and the excursions made by each intermediate molecule will be greater the nearer it is to the middle. On the other hand, the rarefactions and condensations are greatest at the extremities, and diminish as we approach the middle of the pipe, where there is neither condensation nor rarefaction. The analogy of this case with the case of the vibrating cord will be evident if we consider that the condensation in the former is represented by the angle of inclination of the vibrating curve to its abscissa in the latter, and that the mode of vibration now contemplated in the aerial molecules is analogous to that of a cord vibrating as a whole, and having its two halves symmetrical.

In the same way as a vibrating cord is susceptible of division into its several aliquot parts all vibrating simultaneously, so may the aerial column in our stopped pipe vibrate in distinct ventral segments. The manner in which this may take place will be evident on inspection of figs. 15 and 16, where the arrows denote the directions of the motions of the vibrating molecules, and where we see the immobility of the nodal sections is secured by the equal and opposite pressures of the molecules on either side of them. At these nodal sections, too, the same thing holds good as at the stopped extremities, their molecules remain constantly at rest while yet they undergo greater vicissitudes of compression and dilatation than those in any other parts of the column.

Precisely, too, as in the vibrations of strings, any number of these modes of vibration may go on simultaneously. Such combined modes may be produced by an expert flute player, by a nice adjustment of the force of his breath; at least the octave of any note may be obtained without difficulty, and distinctly heard with the fundamental tone.

Half way between two nodes (regarding the stopped ends as nodes) the condensations and rarefactions are evanescent, and the amplitudes of the molecular excursions are at a maximum. Now at such a point let us conceive a narrow ring of the cylindrical pipe in which the vibrating column is contained to be cut away, so as to open a free communication with the outer air. There will be no tendency for air to pass in or out, because the air within is constantly, at these points, in its natural state as to density; neither will its motion be impeded, being parallel to the axis of the column and without any lateral bias. The detachment then of such a ring will no way alter the vibrations of the column, nor, *a fortiori*, will the opening of a hole in the pipe at this place affect

178.

Analogy between vibrations of air in a pipe and those of a stretched cord.

179.

Transition from a propagated pulse to a permanent vibration.

180.

Effect of a permanent vibratory initial impulse.

181.

Simplest mode of vibration of the air in a pipe closed at both ends.

Analogy with the corresponding case of a stretched cord.

182.

Subdivision of a vibrating column by nodes. Figs. 15, 16.

183.

Coexistence of several modes of vibration.

184.

Vibration of air in a pipe open at one end.

Sound. them. Suppose, now, at this hole, a vibrating body placed, whose vibrations are executed in equal times with those in which the excursions to and fro of the included aerial sections are performed in the stopped pipe. They will be communicated to them, and thus the Sound of the pipe will be excited and maintained. Such an aperture is called an *embouchure*.

185. But let us now conceive the one half (A) of the pipe entirely removed, and in its place a disc substituted exactly closing the aperture, and maintained, by some external cause, constantly in a state of vibration, such, that the performance of one complete vibration, going and returning, shall exactly occupy as much time as a sonorous pulse would take to traverse the whole length of the stopped pipe (A + B), or double that of the open one (B). Its first impulse on the air will be propagated along the pipe (B) and reflected at the stopped end, and will again reach the disc just at the moment when the latter is commencing its second impulse. But the absolute velocity of the disc in its vibrations being excessively minute compared with that of Sound, the reflected pulse will undergo a second reflexion at the disc as if it were a fixed stopper. It will, therefore, in its return exactly coincide and conspire with the second original impulse of the disc, and the same process being repeated on every impulse, each will be combined with all its echos, and a musical tone will be drawn forth from the pipe vastly superior to that which the disc vibrating alone in free air would produce. This is, in fact, the simplest instance of the *resonance* of a cavity, of which more hereafter. (See Index, *Resonance*.) Now, it is manifestly of no importance whether the pulses reflected from the closed end of the pipe (B) undergo a second reflexion at the disc, and are so returned back by the pipe, or whether we regard the disc as penetrable by the pulse, (*i. e.* a mere imaginary vibrating section,) and suppose the pulse to run on and be reflected at the extremity of the other half (A) of the bisected pipe (A + B), and on its return again to pass freely through the disc and be again reflected at the stopped extremity of (B). The Sounds produced will be the same, on the principle of the superposition of vibrations. Thus we see that the fundamental Sound of a pipe open at one end is the same with that of a pipe closed at both ends, and of double the length.

186. The mode here supposed of exciting and maintaining the vibrations of a column of air in a pipe is easily put in practice. Let any one take a common tuning-fork and on one of its branches fasten with sealing-wax a circular disc of card of the size of a small wafer, or sufficient nearly to cover the aperture of a pipe. The sliding joint of the upper end of a flute, with the mouth-hole stopped, is very fit for the purpose; it may be tuned in unison with the loaded tuning-fork (a C fork) by means of the movable stopper, or the fork may be loaded till the unison is perfect. If the fork be then set in vibration by a blow on the unloaded branch, and the disc be held close over the mouth of the pipe, as in fig. 16, a note of surprising clearness and strength will be heard. Indeed, a flute may be made to "speak" perfectly well by holding close to the embouchure a vibrating tuning-fork while the fingering proper to the note of the fork is at the same time performed. We shall have further occasion to refer to this point. (*Resonance*, Index.)

187. But the most usual means of exciting the vibrations of a column of air in a pipe is by blowing into, or rather over it, either at its open end or at an orifice made for the purpose at the side, or by introducing a small current of air into it through an aperture of a peculiar construction called a *reed*, provided with a "tongue" or flexible elastic plate which nearly stops the aperture, and which is alternately forced away by the current of air and returns by its elasticity, thus producing a continual and regularly periodic series of interruptions to the uniformity of the stream, and, of course, a Sound in the pipe corresponding to their frequency. Except, however, the reed be so constructed as to be capable of vibrating in unison, or nearly so, with, at least, one of the modes of vibration of the column of air in the pipe, the Sound of the reed only will be heard, the resonance of the pipe will not be called into play, and the pipe will not *speak*; or will speak but feebly and imperfectly, and yield a false tone.

188. But of reeds more hereafter; (see Index, *Reeds*;) at present let us consider what takes place when the vibrations of a column of air are excited by blowing over the open end of a pipe or an aperture in its side. To do it effectually the air must be directed in a small current, not *into*, but *across* the aperture, as in fig. 18, so as to graze the opposite edge. By this means a small portion will be caught and turned aside down the pipe, thus giving a first impulse to the contained air and propagating down it a pulse in which the air is slightly condensed. This will be reflected at the end as an echo, and return to the aperture where the condensation goes off, the section condensed expanding into the free atmosphere. But in so doing, it lifts up, as it were, and for a moment diverts from its course the impinging current, and thus, while it passes, suspends its impulse on the edge of the aperture. The moment it has escaped the current resumes its former course, again touches the edge of the aperture, creates there a condensation, and propagates downwards another condensed pulse, and so on. Thus the current passing over the aperture is kept in a constant state of *fluttering* agitation, alternately grazing and passing free of its edge, at regular intervals, equal to those in which a sonorous pulse can run over twice the length of the pipe; or, more generally, in which the condensations and rarefactions recur at its aperture in virtue of any of the modes of vibration of which the column of air in the pipe is susceptible.

189. In general, wherever there is a communication opened between the column of air in a pipe and the free atmosphere, that point will become a point of maximum excursion of the vibrating molecules, or the middle of a ventral segment. In such a point the rarefactions and condensations vanish, the air reducing itself constantly to an equilibrium of pressure with the free atmosphere with which it is in contact. Hence, if the pipe speak at all, it will take such a mode of vibration as to satisfy this condition, but, consistently with this, it may divide itself into any number of ventral segments. But here there is a point of practical difference between the affections of a vibrating aerial column and those of a tended cord. The tension of the cord can only be maintained steadily *in practice*, by fixing its two ends; so that the case of one extremity fixed, the other free, can have no existence but in imagination, where the cord may be conceived as of indefinite length in one direction, so that the out-running pulses may lose themselves, or, at least, never return. It is true they might be stifled by wrapping one

Case where the vibrations are excited by a vibrating disc.

How performed in practice.

Fig. 16.

Action of a reed.

Excitement of vibrations by blowing over an orifice.
Fig. 18.

Practical difference between vibrations of air in an open pipe and those of a stretched cord.

Sound. end of a very long cord in cotton, but whether, under such circumstances, any mode of producing and maintaining an initial periodical impulse sufficiently regular to produce musical Sounds could be found remains to be tried. The nearest approach to the case in question is when one end of a long cord is held in the hand and agitated while the tension is maintained, as in Art. 173. **Part II.**

In cords with fixed extremities, however, all the ventral segments must, of necessity, be complete, no half segments can exist. In pipes it is otherwise. The air in a pipe closed at one end vibrates as a half, not the whole of such a segment. It is owing to this that a pipe open at both ends can yield, if properly excited, a musical Sound. The column of air in it vibrates in the mode represented in fig. 19, where there is a node in the middle and each ventral segment is only half a complete one. In general it is easy to represent, in an algebraic formula, the time of vibration, or the number of vibrations per second corresponding to any mode of vibration. For, first in a pipe open at both ends let the number of nodes be n , then there will be $n - 1$ complete ventral segments between them, as in fig. 20, and a moiety of one at each end. If, then, we call L the whole length of such a pipe in feet, V the velocity of Sound in feet per second, the length of one complete ventral segment will be $\frac{L}{n}$. This length is traversed by a sonorous pulse in a time $\frac{1}{n} \cdot \frac{L}{V}$, and this is the time of vibration of the middle section of it to which the Sound corresponds. The pipe, then, vibrating according to this mode, will yield a Sound whose pitch is that of a cord making $n \cdot \frac{V}{L}$ vibrations per second; and the series of tones it can produce is expressed by the following series of numbers of vibrations,

$$1 \cdot \frac{V}{L}; \quad 2 \cdot \frac{V}{L}; \quad 3 \cdot \frac{V}{L}, \&c.$$

In the case of a pipe closed at one end, the stopped end must be regarded as a node. (Fig. 21.) Calling the whole number of nodes, thus included, n , the number of complete ventral segments will be $n - 1$, and one half segment will terminate at the open end. Therefore $\frac{L}{2(n-1)+1}$, or $\frac{L}{2n+1}$, will be the number of such halves contained in the length L , and $\frac{2L}{2n+1}$ will, therefore, be the length of each complete one; so that each will make $\frac{2n+1}{2} \cdot \frac{V}{L}$ vibrations in one second, and thus the series of tones such a pipe can yield will be expressed by the series of vibrations,

$$\frac{1}{2} \cdot \frac{V}{L}; \quad \frac{3}{2} \cdot \frac{V}{L}; \quad \frac{5}{2} \cdot \frac{V}{L}, \&c.$$

Lastly, the number of nodes, including the two stopped ends of a pipe closed at both ends, being n , the number of segments (all complete) into which it will be divided will be $n - 1$, and the length of each will be $\frac{1}{n-1} \cdot \frac{L}{V}$; so that the series of Sounds, of which such a pipe is susceptible, is represented by the series of vibrations,

$$1 \cdot \frac{V}{L}; \quad 2 \cdot \frac{V}{L}; \quad 3 \cdot \frac{V}{L}, \&c.$$

Taking, therefore, unity for the number of vibrations per second in the fundamental tone, the series of harmonics will run as follows:

In a pipe stopped at both ends 1, 2, 3, 4, 5, &c.
 ————— open at both ends 1, 2, 3, 4, 5, &c.
 ————— stopped at one end, open at the other... 1, 3, 5, 7, 9, &c.

It being recollected, however, that in the last series the fundamental note 1 is an octave lower than in the others, *i. e.* performs its vibrations only half as rapidly.

To produce these Sounds by blowing into a pipe, it is only requisite to begin with as gentle a blast as will make the pipe speak, and to augment its force gradually. The fundamental tone will be heard first; and as the strength of the blast increases, will grow louder, till at length the tone all at once starts up an octave, *i. e.* the interval between notes whose vibrations are as 1 : 2. By blowing still harder, the next harmonic, 1 : 3, or as it is called in Music, the octave of the fifth, or the *twelfth* of the fundamental tone, is heard; but no adaptation of the embouchure, or force of the wind, will produce any note intermediate between these. The next harmonic is 1 : 4, and corresponds to the double octave, or *fifteenth* of the fundamental tone; and the next, or 1 : 5, to the *seventeenth*, or major third above the double octave. (See the explanation of these terms in Art. 210, *et seq.*) The next, 1 : 6, corresponds to the *nineteenth*, or double octave of the fifth, and so on. All the notes here enumerated are very readily produced on the flute, without changing the fingering, from the lower C or D upwards, by merely varying the force of the blast, and a little humouring the form of the lips and their position with respect to the embouchure. The reader may consult on this subject D. Bernouilli, *Sur le Son et sur les Tons des Tuyaux d'Orgues*, *Mém. Acad. Paris*, 1762; in which the true theory of wind-instruments is

Sound. first clearly stated, though pointed out by Sauveur, in a Paper published in the *Mém. de l'Acad.* 1701. Also a very instructive Paper by Lambert, *Observations sur les Flûtes*, *Mém. Acad. Berl.* 1775. (See also Euler, *De Motu Aeris in Tubis Petrop. Comm.* xvi. and Lagrange's Memoirs already cited, *Mél. de Turin*, i. and ii.) M. Biot, by adapting an organ bellows to regulate the blast, and give it the requisite force and uniformity, succeeded in drawing from a pipe furnished with a proper embouchure, not only these, but also the notes represented in the harmonic series by 7, 8, 11, and 12, but not 9 or 10, (the reason of which vacancy does not appear.) *Traité de Physique*, ii. 126. Part I.

194. The rationale of the continual subdivision of the vibrating column, as the force of the blast increases, is very obvious. A quick sharp current of air is not so easily driven aside by an external disturbing force; and when so driven, returns more rapidly to its original course, than a slow and feeble one. A quick stream, when thrown into a ripple by an obstacle, undulates more rapidly than a slow one. Consequently, on increasing the force of the blast, a period will arrive when the current *cannot* be diverted from its course and return to it so *slowly* as is required for the production of the fundamental note. The next higher harmonic will then be excited, until, the force of the blast increasing, it becomes once more incapable of sympathizing with the excursions of the aerial molecules at the embouchure in this mode of vibration, and so on.

195. If we know the velocity of Sound in the column of air included in a pipe, the length of the pipe, and the mode of vibration, the number of vibrations may be computed, and *vice versa*, if we know the number of vibrations made in a given pipe, vibrating in a known manner, we may thence compute the velocity of Sound. This furnishes a ready and simple method of determining the velocity of Sound in any gas or vapour. We have only to fill a pipe first with air, and then with the gas or vapour in question; and having set them vibrating by any proper means, so as to draw forth their fundamental tone, to compare this with a monochord, or with any musical instrument possessing a regular scale or progression of notes where vibrations are known; and having thus ascertained the number of vibrations per second performed by a column of each medium, the velocities of Sound in the respective cases will be in the direct ratio of their numbers. It is thus that Chladni, and, more lately, Vanrees, Fraeneyer, and Moll have ascertained the velocities of Sound in the various media enumerated in Art. 82.

196. That it is really the air which is the sounding body in a flute, organ-pipe, or other wind-instrument, appears from the fact, that the materials, thickness, or other peculiarities of the pipe, are of no consequence. A pipe of paper and one of lead, glass, or wood, provided the dimensions be the same, produce, under similar circumstances, exactly the same tone as to *pitch*. If further proof were necessary, the difference of pitch produced by filling the pipe with different gases would place the point beyond a doubt. If the *qualities* of the tones produced by different pipes differ, this is to be attributed to the friction of the air within them, setting in feeble vibration their own proper materials.

197. The influence of the size and situation of the embouchure of a pipe, and still more of the manner of exciting the vibrations of the sections of the aerial column near that place, are very material in determining the pitch of the tone uttered. Were it possible to excite the aerial column to vibration by setting in motion a single section of it by a wish, we should obtain, doubtless, Sounds always strictly conformable to the length of the pipe and its harmonic subdivisions as above; but, in fact, the vibrating column of air and the extraneous body (be it reed, tuning-fork, or stream of air) which sets it in motion exercise on each other a mutual influence; they vibrate as a system, (see Index, *Vibrations of a System of Bodies*.) and the resulting tone may be made to deviate more or less from the pure fundamental tone of the pipe, according to the greater or less mass of matter, and the fixity of the vibrations of the apparatus by which the pipe is made to *speak*. When, for example, the cause of vibration is the mere passage of a stream of air over the orifice, whose motions are almost entirely commanded by the condensations and rarefactions within the pipe, (Art. 194.) but little deviation can take place. Yet, by varying the inclination of the stream, (as in the case of the flute by turning the mouth-hole more inwards or outwards with respect to the lips,) and thus giving it a greater or less obliquity to the edge against which it strikes, we may alter the note very sensibly, as is known to all flute-players, who use this means of humouring the instrument, and playing in tune in keys which would otherwise be insupportable.

198. In the diapason organ-pipe, whether open or stopped, a stream of air is admitted at the vertex of the conical lower end, but is prevented from passing through the whole length of the pipe by a plate of metal separating the cone from the pipe, and is forced to escape through a narrow slit transverse to the axis of the pipe, in doing which it strikes against the edge of a thin piece of lead, or other flexible metal. This disposition will be understood by inspection of the figure 22, in which BB is the organ-pipe, and *bcb* the conical appendage at its foot by which the air is admitted. One side of the pipe BLM is flattened and a little bent inwards, and at L a narrow slit is made, just opposite to the lower edge of which is the plate of metal *bh*, which has its edge nearest the orifice a little cut away, so as not quite to fill the whole section of the pipe, but to leave a narrow slit parallel to the slit FF in the side of the pipe. Through this the air admitted at *c* escapes, and is directed in a thin sheet against the upper lip L of the lateral slit; against which it breaks, as described in Art. 188, and sets in vibration the column of air contained in the pipe. If the stream of air be too strong the pipe will yield the octave and harmonic of its fundamental note, forming the series 1, 2, 3, 4, &c. If, on the other hand, the current of air remaining constant, the breadth of the slit through which the air escapes be diminished, according to the experiments of MM. Biot and Hamel, harmonics will also be produced, but in the progression 1, 3, 5, 7, &c. the octaves of the fundamental tone and of all the others being entirely wanting.

199. In reed-pipes, or those in which the vibrations are excited and maintained by passing a current of air into the pipe through a reed, (Art. 187.) the influence of the reed on the pipe is very great. The most perfect and pure tone is produced, of course, when the reed and the pipe separately are pitched in unison, but a considerable latitude in this respect exists; and within certain limits, depending on the mass and stiffness of the reed, as

Sound.

compared with the dimensions of the pipe, a power of mutual accommodation subsists, and a mean tone is produced, less powerful and less pure and pleasing, however, as the pipe is more forced from its natural pitch, until it ceases to sound altogether, and the note produced, if any, is that of the reed alone. In this respect there is, however, a great difference in pipes of different sizes. In large organ-pipes the reed vibrates with nearly the same freedom as in the open air, and will, therefore, *speak* when the pipe has ceased to resound; but in small and narrow pipes, as in oboes, and other similar wind-instruments, a much closer correspondence between the pitches of the reed and pipe is required, or the reed will not vibrate. Messrs. Biot and Hamel adapted to a glass pipe a reed of the ordinary construction represented in fig. 23, in which the vibrating tongue L (by whose oscillations the opening of the reed at R is alternately opened and closed) could be lengthened or shortened at pleasure by thrusting in or withdrawing a wire Ff, which bears with a slight spring against the tongue at f. The blast of wind being maintained constant, the reed was made to yield its gravest note, by withdrawing the wire as far as possible, after which, by pushing it in, the pitch of the reed was gradually raised. It was observed then that the tone of the pipe grew constantly more acute, but that after a certain point, it began to diminish in intensity, till at length no Sound could be heard. At this point, the tongue of the reed, being narrowly examined through the glass, was observed to be still in rapid vibration; but its vibrations were performed entirely in the air, so as not to strike upon and close the orifice. A constant passage then being left for the air, the vibrations of the pipe could not be excited. But this state of things continued only so long as the tongue was of that precise length. The moment the wire was pushed in by the smallest quantity, the Sound sprung forth anew of a pitch still corresponding with the shortened state of the tongue.

The influence of the air in a pipe on the reed by which it is set in vibration, causes the *quality* of the tone of a reed-pipe to depend materially on its figure. Thus it is found that a reed-pipe of the funnel-shaped form, fig. 24, composed of two cones, one more divergent than the other, set on the orifice, gives the clearest and most brilliant tone; but, on the other hand, if the upper cone be reversed, so as to contract the aperture, fig. 25, the Sound is stifled. But when two similar cones, placed base to base, are adapted to the aperture of a long conical pipe, as in fig. 26, the Sound acquires remarkable fullness and force. This belongs, however, to a most intricate part of the theory of Sound, the vibrations of masses of air in cavities of any form.

The quality of the tone produced by reed-pipes will also of course materially depend on the construction of the reed itself, and the substance of which it is composed. If the vibrating lamina be of metal, and at every vibration it strikes on a metallic orifice, these blows will be heard, and will give a harsh, rude, and screaming character to the Sound. If the edges of the aperture be covered with soft leather, this is much alleviated. But if, instead of *covering* the aperture by *striking* on it, the tongue is so constructed as merely to *obstruct* it by passing backwards and forwards *through* it at each oscillation, care being taken to make it *fit* without touching the edges of the aperture, these blows are avoided altogether; the tongue coming in contact with nothing but air during its whole motion. In consequence, its tone is remarkably soft and pure, and free from any harshness.

The invention of this reed is ascribed by Biot to M. Grenié, who has taken out a patent for it; but, without erecting a prior claim on the part of Kratzenstein, we may bring forward a very familiar instrument, the Jew's-harp, as offering, at least, an apparent analogy with M. Grenié's reed. The construction of this instrument is so well known that there is no need to describe it; and though the theory of it be somewhat obscure, there can be little doubt that its action is that of a reed which calls into play the resonance of the cavity of the mouth, and sympathizes with it in its vibrations, at least in some of their modes. The Jew's-harp is an instrument much mistaken and unjustly condemned. Nothing can exceed the softness, sweetness, and delicacy of this instrument, when carefully constructed and well played;* as might be expected from a reed in which the tongue is perfectly at liberty. That the instrument itself vibrates in unison with the note it calls forth, is evident from the fact, that when merely held before the open mouth, or lightly retained between the lips, its Sound is feeble and scarce audible; but acquires a great accession of force when brought in contact with and firmly held between the teeth; the note is still further sustained and reinforced by directing a current of air forcibly through it. It is not here meant to say, that the great oscillations to and fro of the tongue are commanded by the resonance of the cavity, or are performed in the same time with its vibrations. On the contrary the spring is far too strong and large to admit of this. It is more probably by a series of subordinate vibrations going on in the tongue while oscillating, that the sympathy is established.

The instrument called the German harmonica is a reed, on M. Grenié's principle, consisting of nothing but a very thin lamina of brass, of the form of an oblong parallelogram fixed by one of its narrow ends in a frame of its own shape, but just so much larger as to allow of its free motion. This instrument vibrates by a blast urged through it yielding a clear musical tone of a very pleasing character and fixed pitch. If placed at the end of a pipe it performs the office of a reed, and its tone commands, or is commanded, by the pipe according to circumstances, as above explained.

When the action of the embouchure of a pipe is so decided as to be incapable of being, to any sensible extent, commanded or influenced by the resonance of the pipe; as, for instance, when the column of air in a stopped pipe is set in vibration by a tuning-fork furnished with a disc, as described in Art. 119, the pipe will sound, and reinforce the Sound of the tuning-fork, but more and more feebly, as the pitch of the latter departs more from that of the pipe. The experiment is easily made by tuning the upper joint of a flute with the mouth-hole stopped exactly in unison with a fork, and then moving the piston of cork at the end of the pipe to and fro, or loading the fork with wax, so as to put it more or less out of tune. The fork and aerial column vibrate as a system, in which the former has so much the preponderance as to command the latter completely.

We may here notice a very remarkable experiment, which we do not remember to have seen elsewhere

Part II.

Mutual influence of the reed and pipe.

Fig. 23. Construction of a reed,

And its manner of vibrating.

200.

Influence of the form of the pipe on the Sound.

Fig. 24.

Fig. 25.

Fig. 26.

201.

Influence of the reed on the quality of the Sound.

202.

Grenié's reed. The Jew's harp.

203.

The German harmonica.

204.

Case where the pipe is commanded by the embouchure.

205.

* As we have heard it done by M. Eulenstein.

Sound.
Singular ex-
periment :
a double
Sound
yielded by
one pipe.

described, and which shows to what an extent the principle of the *superposition* of vibrating motions and the simultaneous coincidence of different modes of vibration in the same vibrating body, must be admitted in Acoustics. If, instead of one, two such disked tuning-forks be held over the mouth of a pipe side by side, both nearly in unison with the pipe, but purposely tuned out of unison with each other, by an interval so small (see Index, *Musical Intervals*) as to produce strong beats, (see Index, *Beats*,) both Sounds at once will be reinforced by the pipe, and the beats will be heard with the same degree of distinctness as if two pipes, each in unison with one of the forks, were sounding side by side. The same column of air, then, at the same time, is vibrating as a part of two distinct systems, and each series of vibrations, however near coincidence they may be brought, continues perfectly distinct and absolutely free from any mutual influence. To those who have not tried the experiment, the fact of a pipe actually out of tune with itself, and yielding two notes in irreconcilable discord with one another, yet both equally clear and loud, will, at first sight, appear not a little extraordinary.

206.
Vox huma-
na organ-
pipe.
Fig. 27.

One of the most singular species of pipe is that employed in the organ to imitate the human voice. It is composed of a very short conical pipe, the base upwards surmounted by a short cylinder, and the pitch is regulated entirely by the reed. (See fig. 27.) There is a circular operculum which half closes the open end of the cylinder, to imitate the lips, the reed performing the part of the larynx, and the pipe itself, of the cavity of the throat and mouth. (See Index, *Voice*.) This pipe, when well executed, imitates the human voice extremely well ; but with a peculiar nasal twang, and somewhat of a screaming tone.

207.
Chimney-
pipes.
Fig. 28.

Chimney-pipes are those which are closed at the upper end by a cover, through the centre of which a pipe of smaller diameter is passed, as a continuation of the lower one. (Fig. 28.) Their Sound is intermediate between those of open and stopped pipes of the same length. Bernouilli (*Mém. Acad. Par.* 1762) has investigated the laws of their vibrations. (See also Biot, *Traité de Phys.* ii. p. 153.) If we call L and l the lengths of the greater and smaller pipes respectively, and λ that of a pipe closed at one end capable of yielding the same fundamental note, $n : 1$, the ratio of their diameters, Bernouilli finds the following equation for determining λ ,

$$\tan\left(\frac{\pi}{2} \cdot \frac{L}{\lambda}\right) \times \tan\left(\frac{\pi}{2} \cdot \frac{l}{\lambda}\right) = \frac{1}{n^2}.$$

This equation holds good when the lower end of the great pipe is closed ; if it be open, the equation is

$$\tan\left(\frac{\pi}{2} \cdot \frac{L - \lambda}{2\lambda}\right) \cdot \tan\left(\frac{\pi}{2} \cdot \frac{l}{\lambda}\right) = \frac{1}{n}.$$

§ IV. Of Musical Intervals, of Harmony, and Temperament.

208.
t nisons.

Our appreciation of the *pitch* of a Musical Sound depends, as we have seen, entirely on the number of its vibrations performed in a given time. Two Sounds whose vibrations are performed with equal rapidity, whatever be their difference in quality or intensity, affect the ear with a sentiment of accordance which we term a *unison*, and which irresistibly impresses on us the conviction of a perfect analogy, or similarity between them, which we express by saying that their *pitch* is the same, or that they sound the same *note*. In fact, their impulses on the air, and on the ear, through its medium, occurring with equal frequency, blend, and form a compound impulse, differing in quality and intensity from either of its constituents, but not in the frequency of its recurrence ; and, therefore, the ear will judge of it as of a single note of intermediate quality.

209.
Musical
concord
and dis-
cords.

But when two notes not in unison are sounded at once, the ear distinctly perceives both, and (at least with practice, and some ears more readily than others) can separate them, in idea, and attend to one without the other. But besides this, it receives an impression from them jointly, which it does not acquire when sounded singly, even in close succession, an impression of concord, or dissonance, as the case may be ; and is irresistibly led to regard some combinations as peculiarly agreeable and satisfactory, and others as harsh and grating. Now it is invariably found that the former are those, and those only, in which the vibrations of the individual notes are in some very simple numerical proportion to each other, as 1 to 2, 1 to 3, 1 to 4, 2 to 3, &c., and that the concord is more satisfactory and more pleasing, the lower the terms of the proportion are, and the less they differ from each other. While, on the other hand, such notes as vibrate in times bearing no simple numerical ratio to each other, or in which the times of the ratio are considerable, as 8 : 15, for example, when heard together produce a sense of discord, and are extremely unpleasant. This simple remark is the natural foundation of all harmony.

210.
The octave.
Fig. 29.

Next to a *unison*, in which the vibrations of the two notes are in the ratio of 1 : 1, the most satisfactory concord is the *octave*, where the vibrations are as 1 : 2, or one note performs two vibrations to each single one of the other. The *octave* approaches in its character to a *unison*, and indeed two notes so related when played together can hardly be separated in idea ; and when singly, appear rather as the same note differently modified, than as independent Sounds. The reason of this will be evident on inspecting fig. 29, where the dots in the upper line represent the periodically recurring impulses on the ear produced by the vibrations of the acuter note, while those in the lower represent the same impulses as produced by those of the graver ; as the ear receives these all in the order they are placed, it will be the same thing as if they were produced by two Sounds both of the graver pitch, but one of a different intensity and quality from the other ; the one having its impulses (represented by :) the sum of two separate impulses of the octave Sounds ; the other consisting of the alternate impulses (.) of the acuter only.

211 In like manner the *octave* of the *octave*, or the *fifteenth*, as it is called in Music, which consists of notes whose

Sound. vibrations are as 1 : 4, is a very agreeable and perfect concord ; as are, indeed, all the scale of octaves 1 : 8, 1 : 16, &c. they all partake of the peculiar character of the octave, a sense of perfect adjustment or identity.

The next in order is the combinations 1 : 3, where the vibrations of the graver note are trisected by those of the acuter, as in fig. 30, which gives a concord called the twelfth, a very agreeable one. In this, if we substitute for the note 1 its octave 2, we shall have the concord whose vibrations are in the ratio of 2 : 3 ; or, as we shall call it for brevity, the concord 2 : 3, whose pulsations are represented in fig. 31. This concord is termed the *fifth*, and is a remarkably perfect and agreeable one, even more so than the twelfth, which although simpler in a numerical estimate, yet from the greater interval between its component notes allows them to be more readily distinguished, while the notes of the fifth blend much more perfectly.

If, instead of substituting for 1 its octave 2, we substitute its double octave 4, we get the concord 4 : 3, or the *fourth*, which may be regarded as a sort of complement of the *fifth*, and is also very agreeable.

The concords 1 : 5, 2 : 5, and 4 : 5, especially the latter, in which the tones approach pretty near to each other, are all remarkably agreeable. The last is called a *major third*, and the two former are regarded rather as varieties of it than as independent concords. The concord 8 : 5 (which is its complement in the same sense as 4 : 3 the fourth is to 2 : 3 the fifth) is called the *minor sixth*, and is almost equally agreeable with the major third, to which it is related.

The concord 3 : 5 is called the *major sixth*, and, as well as its complement 6 : 5, or the *minor third*, though pleasing, is decidedly less satisfactory than the foregoing ; and, as we see by casting our eyes on the figure, the periods of recurrence of their combined pulses in the same order is longer and more complex.

Higher primes than 5 enter into no harmonic ratios. Such combinations, for instance, as 1 : 7, 5 : 7, or 6 : 7, are altogether discordant. The same may be said of the more complicated combinations of the lower primes 1, 2, 3, 5. The ear will not endure them, and cannot rest upon them. When sounded, a sense of craving for a change is produced, and this is not satisfied but by changing one or both of the notes so as to fall as easily as the case will permit into some one of the concords above enumerated. This is called the resolution of a discord ; and such is the constitution of our minds in this respect, that a concord agreeable in itself is rendered doubly so by being thus approached through a discord. For example, let us take the ratio 5 : 9, which is called a *flat seventh*, a combination decidedly discordant. If we multiply the terms of this ratio by 5, we get 25 : 45. A small change in one of the notes will reduce this to 27 : 45, or 3 : 5, a major sixth—an agreeable concord. Now this will be done, if, retaining the lower note 5 or 25, we change the upper from 45 to $45 \times \frac{4}{3}$; that is to say, to a note whose vibrations are to its own as 25 : 27. This ratio corresponds to a musical interval called a *semitone*. Hence the discord in question will be satisfactorily resolved by holding on its lower note, and making its upper one descend a semitone.

On the proper alternation of concords and discords the whole of musical composition depends, but though the principle above stated must be satisfied in the resolution of every discord, there are other rules to be attended to by which our choice is limited to some modes rather than others ; for example, in the foregoing instance it is the upper note which must descend a semitone. The ascent of the lower by the same interval, which would equally change the ratio as above indicated, would offend against other precepts with which we have here nothing to do.

The *interval*, as it is called in Music, between the two notes of which any simple concord or discord consists, depends not on the absolute number of vibrations which either makes in a given time, but on their relative proportion. For it is no matter how slowly, or how rapidly, the vibrations take place, provided the order in which their impulses reach the ear be the same. Hence, if the vibrations 4 and 5 produce on the ear the agreeable effect of a major third ; two notes, each an octave higher, or having their vibrations respectively 8 and 10 ; or in general any two having their vibrations in this ratio, will produce the same effect. This is a matter of experience, but the inspection of the figures representing the order of succession of the individual vibrations enables us to understand its reason.

If we take any note for a fundamental Sound, and tune a string or a pipe so as to vibrate with the degree of rapidity corresponding to that Sound, and represent by unity the number of vibrations it makes per second ; and if we also tune other strings to make in the same time respectively the numbers of vibrations represented by

$\frac{5}{4}$, $\frac{4}{3}$, $\frac{3}{2}$, $\frac{5}{3}$, 2 ; and then sound all these strings in succession, beginning with the fundamental note,

we shall perceive that two of the sequences, the first and last, are much wider than the rest, and would admit the interpolation of a note between each. But it is no longer possible to choose for these interpolated notes such as will make concordant intervals with any of the rest, or their octaves. But in order to obtain as many concords as possible in the scale, so as to produce the most harmonious music, they are made to harmonize with that note which bears the nearest relation to the fundamental one, (for its octave is regarded as a mere repetition of itself,) i. e. the fifth. The vibrations of a note a fifth higher than the fifth are represented by

$\left(\frac{3}{2}\right)^2$, or $\frac{9}{4}$; and as this is greater than 2, it lies beyond the octave. We must, therefore, tune our interpo-

lated string an octave lower, or to the vibration $\frac{9}{8}$, and thus we get the *second*. Again, if we tune another to

the vibration $\frac{15}{8}$, or $\frac{5}{4} \times \frac{3}{2}$, it will form, with the fifth of the fundamental note $\left(\frac{3}{2}\right)$, a major third—the next most harmonious interval on the scale. The note thus interpolated is the *seventh*.

Part II.

The double octave, or fifteenth, &c. 212.

The twelfth. Fig. 30. The fifth. Fig. 31. 213.

The fourth. Fig. 32. 214.

The major third. Fig. 33. The minor sixth. 215.

The minor third and major sixth. 216.

Discords and their resolution. 217.

Example in the resolution of the discord of the flat seventh. 218.

219.

Musical intervals depend on ratios of vibrations, not on their absolute number. 220.

Completion of the scale by the second and seventh. 221.

Sound.

The interpolated scale, with the vibrations of its respective notes, will stand thus :

221.

Signs (1), (2), (3), (4), (5), (6), (7), (8).

Names of intervals 1st. 2d. 3d. 4th. 5th. 6th. 7th. 8th.

Ratios of vibration. 1, $\frac{9}{8}$, $\frac{5}{4}$, $\frac{4}{3}$, $\frac{3}{2}$, $\frac{5}{3}$, $\frac{15}{8}$, 2.

or multiplying all by 24, to avoid fractions,

24, 27, 30, 32, 36, 40, 45, 48.

Diatonic
scale.

This is called the natural, or diatonic scale; when all its notes are sounded in succession, whether upwards or downwards, the effect is universally acknowledged to be pleasing. The ear rests with perfect satisfaction on the fundamental note, and the intervals succeed each other gracefully, with sufficient variety to avoid monotony. Accordingly all ages and nations have agreed in adopting this scale as the foundation of their music.

222.

Continuation of the scale upwards and downwards. Limits of audibility. Sounds inaudible to certain ears.

This scale consists of seven distinct notes, for the eighth being the octave of the first is regarded as a mere repetition of it. And if we add to it on both sides the octaves of all its tones above and below, and again the octaves of these, and so on, we may continue it indefinitely upwards and downwards. Not that the ear will follow these additional tones, to an unlimited extent. When the vibrations are less numerous than about 16 per second, the ear loses the impression of a continued Sound; and perceives, first, a fluttering noise, then a quick rattle, then a succession of distinct Sounds capable of being counted. On the other hand, when the frequency of the vibrations exceeds a certain limit all sense of *pitch* is lost; a shrill squeak, or chirp, only is heard; and, what is very remarkable, many individuals, otherwise no way inclined to deafness, are altogether insensible to very acute Sounds, even such as painfully affect others. This singular observation is due to Dr. Wollaston. (See his Paper on Sounds inaudible to certain ears, *Phil. Trans.* 1820.) Nothing can be more surprising than to see two persons, neither of them deaf, the one complaining of the penetrating shrillness of a Sound, while the other maintains there is no Sound at all. Thus while one person mentioned by Dr. Wollaston could but just hear a note four octaves above the middle E of the piano-forte, others have a distinct perception of Sounds full two octaves higher. The chirp of the Sparrow is about the former limit; the cry of the Bat about one octave above it; and that of some insects, probably, more than another octave. Dr. Wollaston's sense of hearing terminated at six octaves. The whole range of human hearing comprised between the lowest notes of the organ and the highest known cry of insects, seems to include about nine octaves.

223.

Remark. Feebleness of very acute Sounds.

It is probable, however, that it is not alone the *frequency* of the vibrations which renders shrill Sounds inaudible. There is no reason why an impulse, if strong enough *singly* to affect the ear, should lose its effect if repeated many thousand times in a second. On the contrary such repetition would render the noise intolerable. But this is not the case with musical Sounds; their individual impulses would, probably, be quite inaudible singly, and only impress by repetition. Now, as vibrating bodies have only a certain degree of elasticity, extreme swiftness of vibration can only take place when their dimensions are very minute, and consequently the excursions of their molecules from rest, and their absolute velocities, excessively minute also. Thus in proportion as Sounds are more acute their *intensity* (which depends wholly on the extent and force of their vibrations) diminishes. No doubt, if by any mechanism a hundred thousand hard blows per second could be regularly struck by a hammer on an anvil, at precisely equal intervals, they would be *heard* as a most deafening shriek; but in natural Sounds the impulses lose in intensity more than they gain in number, and thus the Sound grows feebler and feebler till it ceases to be heard.

224.

Possible limits of the sense of hearing in different animals.

"As there is nothing in the nature of the atmosphere (remarks Dr. Wollaston) to prevent the existence of vibrations incomparably more frequent than any of which we are conscious, we may imagine that animals like the Grylli, whose powers appear to commence nearly where ours terminate, may have the faculty of hearing still sharper Sounds, which we do not know to exist; and that there may be other insects hearing nothing in common with us, but endued with a power of exciting, and a sense which perceives vibrations of the same nature indeed as those which constitute our ordinary Sounds, but so remote that the animals who perceive them may be said to possess another sense agreeing with our own solely in the medium by which it is excited." The same may, no doubt, be true of aquatic animals. The shrimp and the whale may have no Sound in common.

225.

Key-note of a piece of music. Its analysis into chords.

By the aid of the ascending and descending series of Sounds in the natural scale thus obtained, pieces of music perfectly pleasing, both in point of harmony and melody, may be played; and they are said to be in the *key* of that which is assumed as the fundamental note of the scale, or whose vibrations are represented by 1. If such a piece be analyzed, it will be found to consist entirely, or chiefly, of triple and quadruple combinations, or chords, such as the following :

226.

The fundamental or common chord.

1. The common, or fundamental chord, or chord of the *tonic*, or the 1st, 3d, and 5th, (1, 3, 5,) or the 3d, 5th, and octave (3, 5, 8) sounded together. This is the most harmonious and satisfactory *chord* in music, and when sounded the ear is satisfied, and requires nothing further. It is, therefore, more frequently heard than any other; and its continual recurrence in a piece of music determines the *key* it is played in.

227.

Chord of the dominant.

2. The chord of the *dominant*. The fifth of the key-note is called, by reason of its near relation to the fundamental note, the dominant. The chord in question is the common chord of the dominant, or the notes (2, 5, and 7) sounded together.

228.

Chord of the subdominant.

3. The chord of the *subdominant*, or the note 4, consisting of the notes 1, 4, 6, being the common chord of 4 as a fundamental note.

229

4. The false close, or the combination (1, 3, 6) or (3, 6, 8) which is, in fact, the common chord of the note 6, only with a minor third (6, 8) instead of a major. The term *false close* arises from this, that a piece of music, frequently before its final termination, (which is always on the fundamental chord,) comes to a momentary close

Sound. on this chord, which pleases only for a short time, but requires the strain to be taken up again and closed as usual to give full satisfaction.

5. The discord of the 7th, or (2, 4, 5, 7.) It consists of four notes; and is, in fact, the common chord of the dominant, with the note immediately below it, or the seventh *in order* above it. The interval, however, between the notes (4) and (5), or between (5) and the octave of (4) next above it, is represented by the ratio

$$\frac{2 \times 4}{3} \div \frac{3}{2} = \frac{16}{9},$$

or (taking 24 as the number of vibrations in a unit of time corresponding to the note (1)) = $42 \frac{2}{3}$. This

interval, then, is less than the seventh of the diatonic scale, and is about half-way intermediate between the sixth and seventh of that scale. It is, therefore, called the flat seventh. (See Arts. 231 and 232.) This discord resolves itself into the chord (3, 5, 8;) and unless that combination, or one equivalent to it, follows, the ear is not satisfied. The notes (4) and (5) are the essential ones of this discord, and the others are regarded as accompaniments. If played together, the ear requires that in the next chord 4 should descend or be succeeded by (3,) while the note 7 is required to rise or be succeeded by (8.)

With these chords and a few others, such as the chord of the 9th, whose essential notes are 1 and 2, or 1 and 9, may a great variety of music be played, but it would be found monotonous. The ear requires, in a long piece, a variety of *key*. The fundamental note occurs so often that it seems to pervade the whole of the composition, and must therefore be changed. But this change of key, which is called *modulation*, is not possible without introducing other notes than those already enumerated. It is true the chord (2, 5, 7) is the perfect fundamental chord in the key of (5;) but the other chords in that key corresponding to those already enumerated cannot be formed, with the exception of its sub-dominant, which is, in fact, the common chord of 1. Take, for instance, its dominant. This would be formed, if it could be formed at all, of the notes (2, 4, 6) or (4, 6, 9.) But if we come to analyze the intervals of these notes, we find that

$$\frac{(4)}{(2)} = \frac{32}{27}; \quad \frac{(6)}{(2)} = \frac{40}{27}.$$

Now these differ from the ratios $\frac{5}{4}$ and $\frac{3}{2}$ which exist between the notes (3, 1) and (5, 1) of the perfect common chord. Consequently, if we would play equally well in tune in the key (5) we must introduce these new ratios; and, in fact, we ought to have for that purpose notes corresponding to all the ratios

$$\frac{3}{2} \times \frac{9}{8}, \quad \frac{3}{2} \times \frac{5}{4}, \quad \frac{3}{2} \times \frac{4}{3}, \text{ \&c.}$$

and similarly for every other key we might choose to play in. But this would require an enormous number of notes, and would render the generality of musical instruments too complicated. It becomes necessary, then, to consider how the number may be reduced, and what are the fewest notes that will answer.

Let us take for example, as above, the dominant of the note (5.) The number of its vibrations is $36 \times \frac{3}{2}$, or 54, the half of which (because it surpasses the octave of 1) is 27. This is correctly the number corresponding to (2.) Now, taking this for a key-note, the major third of (2) has $27 \times \frac{5}{4} = 34 \frac{1}{4}$ for the number of its

vibrations in a unit of time. Now in the scale as it stands we have 32 and 36, so that the note in question is almost just half-way between them, and must therefore be interpolated. It will stand between (4) and (5) on the scale, and is denoted in Music by the sign \sharp *sharp*, or \flat *flat*; thus it is written either as (4) *sharp*, (4) \sharp , or as (5) *flat*, (5) \flat . With regard to the fifth of the new fundamental note (2) its representative number is

$$27 \times \frac{3}{2} = \frac{81}{2} = 40 \frac{1}{2}.$$

This comes so near 40, that the ear hardly perceives the difference; and though a small error of one vibration in 80 is introduced by using the note (6) as the dominant of (2), yet it is not fatal to harmony; and there is no necessity for encumbering ourselves with new names of notes, and additional pipes or strings to our instruments for its sake. In practice these errors are modified and subdued by what is called *temperament*, of which this is the origin, and of which more presently.

The interval $\frac{81}{80}$ being the difference of two notes, one of which is the octave of the perfect sixth of the fundamental note, and the other arises by reckoning upwards three perfect fifths from the same origin, is called a comma. The former note is represented by

$$2 \times \frac{5}{3} = \frac{10}{3}, \quad \text{the latter by } \left(\frac{3}{2}\right)^3 = \frac{27}{8}, \quad \text{and } \frac{27}{8} + \frac{10}{3} = \frac{81}{80}.$$

In like manner, if we would choose any other note for the fundamental one, similar changes will be required,

Part II.

The false close.
230.
Chord (or discord) of the seventh.

231.
Modulation

Necessity of introducing intermediate notes.

232.
Exemplified by a numerical example.

Flats and sharps.

Introduction of imperfect harmony.

233.
The comma

234.

Sound. and no two keys will agree in giving identically the same scale. All, however, will be nearly satisfied by the interpolation of a new note half-way between each of the larger intervals of the scale, thus

The chromatic scale.

1, $\left. \begin{matrix} 1^\sharp \\ \text{or} \\ 2^\flat \end{matrix} \right\}$, 2, $\left. \begin{matrix} 2^\sharp \\ \text{or} \\ 3^\flat \end{matrix} \right\}$, 3, 4, $\left. \begin{matrix} 4^\sharp \\ \text{or} \\ 5^\flat \end{matrix} \right\}$, 5, $\left. \begin{matrix} 5^\sharp \\ \text{or} \\ 6^\flat \end{matrix} \right\}$, 6, $\left. \begin{matrix} 6^\sharp \\ \text{or} \\ 7^\flat \end{matrix} \right\}$, 7, 8;

and the scale so interpolated is called the chromatic scale.

235. Musicians have long been at issue on the most advantageous method of executing this interpolation. If, indeed, it were intended to give such a preference to the natural scale 1, 2, 3, 4, &c. as to make it perfect, to the sacrifice of all the other keys, there would be little difficulty, as a mere bisection of the intervals would,

probably, answer every practical purpose: thus 1^\sharp or 2^\flat might be represented by $\sqrt{1 \times \frac{9}{8}}$; $2^\sharp = 3^\flat$ by

$\sqrt{\frac{9}{8} \times \frac{5}{4}}$, and so on; but as in practice no preference is given to this particular key, (which is denoted in

Music by the letter C,) but, on the contrary, variety is purposely studied, it is found necessary to depart from the pure and perfect diatonic scale, even in tuning the natural notes; and to do so with the least offence to the ear is the object of a perfect system of temperament. If the ear absolutely required perfect concords there could be no music, or but a very limited and monotonous one. But this is not the case. Perfect harmony is never heard, and if heard would probably be little valued, except by the most refined ears; and it is this fortunate circumstance which renders musical composition, in the exquisite and complicated state in which it at present exists, possible.

236. In order to judge of the limits, however, to which the ear will bear a deviation from exact consonance of musical vibrations, we must first see what takes place when two notes nearly, but not quite, in unison or concord are sounded together. Conceive two strings exactly equal and similar, and equally drawn out from the straight line, to be let go at the same instant; and suppose one to make 100 vibrations per second, the other 101; let them be placed side by side, and at the same distance from the ear. Their first vibrations will conspire in producing a Sound-wave of double force, and the impression on the ear will be double. But at the 50th vibration one has gained half a vibration on the other, so that the motions of the aerial particles, in virtue of the two coexistent waves emanating from either string, are now not in the same but in opposite directions; and the two waves being by supposition of equal intensity, they will instead of conspiring exactly destroy each other, and this will be very nearly the case for several vibrations on either side of the 50th. Consequently, in approaching the 50th vibration, the joint Sound will be enfeebled; there will be a moment of perfect silence, and then the Sound will again increase till the 100th; when the one string having gained a whole vibration on the other, the motions of the particles of air in the two waves will again completely conspire, and the Sound will attain its maximum. The effect on the ear will therefore be that of an intermitting Sound alternately loud and faint. These alternate reinforcements and subsidences of the Sound are called *beats*. The nearer the Sounds of the strings approach to exact unison, the longer is the interval between the beats. If we call n the number of vibrations in which one string gains or loses exactly one vibration on the other, and m the number of vibrations

per second made by the quicker, $\frac{n}{m}$ will be the interval between two consecutive beats. When the unison is

complete, no beats are heard. On the other hand, when it is very defective they have the effect of a rattle of a very unpleasant kind. The complete destruction of the beats affords the best means of attaining by trial a perfect harmony.

237. Beats will likewise be heard when other concords, as fifths, are imperfectly adjusted. Suppose one string to make 201 vibrations, while the other makes 300; then, at and about the 100th of one, and the 150th of the other, the former will have gained half a vibration, and those vibrations of the one which fall exactly on those of the other, (see fig. 31,) being performed with contrary motions will destroy each other; those which fall intermediate only partially. The beats then will be heard, but with less distinctness than in the case of unisons.

238. This seems the proper place to notice an effect which takes place in perfect concords, and only in those which are very perfect, *viz.* the production of a grave Sound by the mere concurrence of two acute ones. If we examine the figure 212, which represents the succession of vibrations in a perfect fifth, we shall see that every third of the one coincides exactly with every second vibration of the other. These coincidences (so delicate is the ear) are remarked by it, and a Sound is heard, besides the two actually sounded, of a pitch determined only by the frequency of the precise coincidences; that is, in this case, a precise octave below the lowest tone of the concord.

239. In general, if one note makes m vibrations and the other n , while another, of which they may both be regarded as harmonics, makes one, that one will be the resultant tone, provided m and n be prime to each other; so that the only difficulty in determining the resultant of two notes, is to determine of what they are both harmonics.

This will be done by reducing m and n , if fractions, to a common denominator $\frac{m'}{N}$ and $\frac{n'}{N}$; then, if m' and n' have no common factor, $\frac{1}{N}$ will represent the fundamental tone. If, then, m and n be integers, and without any common factor, the resultant will be represented by 1.

Sound.

Hence follows a very curious fact, viz. that if several strings, or pipes, be tuned *exactly* to be harmonics of one of them, or to have their vibrations in the ratios 1, 2, 3, 4, 5, &c., then if they be all, or any number of them, from the first onward, sounded together, there will be heard but one note, viz. the fundamental note. For they are all harmonics of the first note 1; and, moreover, if we combine them all two and two, we shall find comparatively but few which will give other resultants, so that these will be lost, as well as the individual Sounds of the strings, all but the first, in the united effect of all the resultant unit Sounds. But to produce this effect, the strings, or pipes, must be very perfectly tuned to the strict harmonics. The effect can never take place by touching the keys of a piano-forte corresponding to the harmonic notes, because they are always of necessity tempered.

To return to our temperament. If we count the semitones in the chromatic scale between (1) and (8) we shall find the number of such intervals 12. If, then, we would have a scale exactly similar to itself in all parts, and which should admit of our playing equally perfectly in every key, we have only to compute the values of the fractions

$$1 = 2^0, 2^{\frac{1}{12}}, 2^{\frac{2}{12}}, 2^{\frac{3}{12}}, \dots, 2^{\frac{11}{12}}, 2,$$

which may readily be done by logarithms, and we shall find the ratios of the vibrations which will give what may be termed the scale of equal intervals.

If we examine the chromatic scale, and consider it as approximately composed of equal, or nearly equal, intervals called semitones, the following will be the number of semitones in each interval:

In the semitone	1	In the major fifth or fifth	7
— second or tone	2	— minor sixth	8
— minor third	3	— major sixth	9
— major third	4	— flat seventh	10
— minor fourth or fourth	5	— seventh	11
— major fourth or minor fifth	6	— octave	12

241.
The scale of equal intervals, or iso-harmonic scale.
242.
System of equal temperament.

If then we reckon upwards from the note (1) by fifths, viz. from (1) to (5), from (5) to (9), (or, which comes to the same, descending an octave, to (2),) from (2) to (6), from (6) to (10), that is to (10 - 7), or (3), and so on, we shall find that after taking twelve such steps as these we shall have fallen upon every note in the scale, and come back to the fundamental note or its octave. But, since no power of 2 is exactly the same with any power of $\frac{3}{2}$, it is evident that no series of steps by *perfect* fifths can ever bring us to any one of the octaves of the fundamental note. Were the chromatic scale perfect, twelve fifths should exactly equal seven octaves, and three major thirds should precisely make one octave. Neither of these, however, can be true of *perfect* fifths or thirds, for $\left(\frac{3}{2}\right)^{12} = 129.74$, and $2^7 = 128$, giving a difference of nearly one vibration in 64 and $\left(\frac{5}{4}\right)^3 = 1.953$ instead of 2. Thus, if we reckon upwards by perfect fifths, we surpass the octaves; if by thirds, we fall continually below them. In this dilemma it has been proposed to diminish all the fifths equally, making a fifth instead of $\frac{3}{2}$, to be equal to $2^{\frac{7}{12}}$; and tuning regularly from the note (1) upwards by such fifths and from the notes so tuned downwards by perfect octaves. This constitutes what has been called the system of equal temperament.

It is evident that in this system the notes will all of necessity be represented by powers of $2^{\frac{1}{12}}$; and that therefore the scale resulting from this system is identical with that of equal intervals, or the iso-harmonic scale described in the last article. Theoretically speaking, it is the simplest that could be devised; and, practically, (though fastidious ears may profess to be offended by it,) it must produce no contemptible harmony. It has, however, one radical fault, it gives all the keys one *character*. In any other system of temperament some intervals, though of the same denomination, must differ by a minute quantity from each other; and this difference falling in one part of the scale in one key, in another in another, gives a peculiarity of quality to each key, which the ear seizes and enjoys extremely. This fact, in which, we believe, all practical musicians will agree, is alone sufficient to prove, that *perfect* harmony is not necessary for the full enjoyment of music. Most practical musicians seem to have no fixed or certain system of temperament; at least very few of them when questioned appear to have any distinct ideas on the subject.

It is a mistake to suppose, as some have done, that temperament applies only to instruments with keys and fixed tones. Singers, violin players, and all others who can pass through every gradation of tone, must all temper, or they could never keep in tune with each other or with themselves. Any one who should keep on ascending by perfect fifths, and descending by octaves or thirds, would soon find his fundamental pitch grow sharper and sharper till he could at last neither sing nor play; and two violin players accompanying each other, and arriving at the same note by different intervals, would find a continual want of agreement.

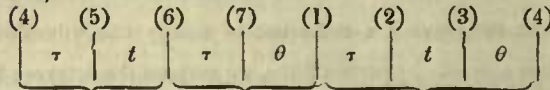
Musical intervals may be numerically represented by the logarithms of the fractions expressing the ratios of the vibrations of the notes between which the intervals are comprised; for the interval depending only on this ratio, and the sum of any two intervals corresponding to the product of their respective ratios, the logarithms of the latter are the proper measures of the magnitudes of the former. Thus an octave corresponds to a ratio of 2 : 1 of the vibrations of its extreme Sounds; two octaves to the ratio 4 : 1 or 2^2 : 1, three to 8 : 1 or 2^3 : 1, and so on; so that log. 2, 2. log. 2, 3. log. 2, &c. or any numbers in that proportion, are proper numerical representatives of these intervals. The intervals of the diatonic scale will, therefore, be represented logarithmically as follows.

	Interval.	Ratio.	Logarithm.	Approx.	Differences.
(1) to (1)	0	1	000000	0	
(1) to (2)	Major tone = τ .	$\frac{9}{8}$	005115	51	$51 = \tau$
(1) to (3)	Major 3d = III		009691	97	$46 = t$
(1) to (4)	Minor 4th = IV.		012494	125	$28 = \theta$
(1) to (5)	Major 5th = V.		017609	176	$51 = \tau$
(1) to (6)	Major 6th = VI.		022185	222	$46 = t$
(1) to (7)	Major 7th = VII.		027300	273	$51 = \tau$
(1) to (8)	Octave = VIII.		030103	301	$28 = \theta$

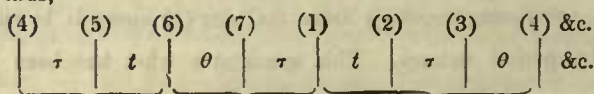
The approximate values of the intervals being all true to a 500th of a tone, an interval far too minute for the nicest ear to appreciate, may be used in all musical calculations where high multiples of them are not taken.

246. It will be observed that the diatonic scale so constructed, consists of three different intervals between consecutive notes. Thus, the interval from (1) to (2) is 51 parts of a scale on which the octave measures 301. This is called a *major tone*, τ , and the same interval occurs again between (4) and (5), and between (6) and (7), as will appear by referring to the column of differences. Again, the interval from (2) to (3) is 46 such parts only, and this, which occurs again between (5) and (6), is called a *minor tone*, (t). Lastly, the interval between (3) and (4) and between (7) and (8) is 28 such parts, and is called (but more improperly) a *semitone*, (θ), being in fact much more than the half of either a major or a minor tone. The term *limma*, which has been used by some authors, is much preferable.

247. This is the origin of what is called the *enharmonic diesis*, and of the distinction existing between the sharp of one note and the flat of that next above it; a distinction essential to perfect harmony, but which cannot be maintained in practice, except in organs and complicated instruments, which admit a great variety of keys and pedals, or in stringed instruments or the voice, where all gradations of tone can be produced, and then only when used without a fixed accompaniment. To explain this distinction, suppose, in the course of a piece of music, commenced in the key of (1), we should *modulate*, as it is called, into the key of (4), its sub-dominant; that is, change our key, and adopt a new scale, having (4) for its fundamental tone. To make the new scale perfect, the intervals should be the same, and succeed each other in the same order as in the original key (1). That is, setting out from (4) we ought to have for our sequence of intervals $\tau t \theta \tau t \theta$. Now, this sequence does not take place in the unaltered scale of (1), when we set out from any note but (1), and if we prolong it backward to (4), they will stand thus,



whereas they ought to stand thus,

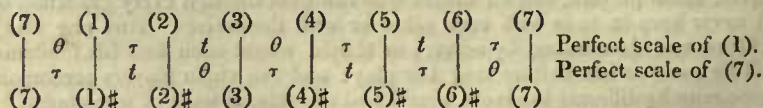


The first two intervals are the same in both. The two next will also agree if we *flatten* the note (7), so as to make $(7)^b - (6) = \theta$ and $(1) - (7)^b = \tau$, which leaves the interval $(1) - (6)$ the same as before, viz. $\tau + \theta$, or a perfect minor third. The quantity by which (7) must be flattened for this purpose, or $(7) - (7)^b$, is equal to $\tau - \theta = 51 - 28 = 23$, and this is the amount by which in this case a note differs from its flat. As to the remaining three intervals, the difference between τ and t being small, amounting only to 5, (which is the

logarithmic representative of the ratio $\frac{81}{80}$, or a *comma*), the sequence $t \tau \theta$ is hardly distinguishable from $\tau t \theta$,

and if the note (2) be *tempered flat* by an interval $= \frac{\tau - t}{2}$, or half a comma, this sequence will in both cases

be the same, and our two scales of (1) and (4) will be rendered as perfect as the nature of the case will permit, by the interpolation of only one new note. But, on the other hand, suppose we would modulate into the key (7). In this case the scales will stand thus:



This change will require the interpolation of no less than five new notes; the notes (7) and (3) being the only ones that remain unchanged. But to confine ourselves to the change from (6) to $(6)^\sharp$, we have $(7) - (6) = \tau$ and $(7) - (6)^\sharp = \theta$. Consequently $(6)^\sharp - (6) = \tau - \theta = 23 = (7) - (7)^b$, as before determined. But since the whole interval between (6) and (7), or $(7) - (6)$, which is $= \tau = 51$, is more than double of this quantity, the flattened note $(7)^b$ will lie nearer to the higher note (7), and the sharpened one $(6)^\sharp$ nearer to the lower one (6) than a note arbitrarily interpolated half way between them, to answer both purposes approximately, would do, and thus a gap, or, as it is termed, a *diesis*, would be left between $(6)^\sharp$ and $(7)^b$.

The diesis in this case amounts only to a comma ($= 5$), or the tenth part of a *major tone*, (τ) ($= 51$), in

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other cases it would be greater. But in all cases the interval between any note and *its sharp* is considered to be equal to that between the same note and *its flat*. Assuming this as a principle, a variety of systems of temperament have been devised for producing the best harmony by a system of 21 fixed Sounds, viz. each note of the seven in the scale, with its sharp and flat, (regarded as different).

The first and most celebrated is that of Huygens. He supposes the octave divided into 31 equal parts. Of these a whole tone, whether τ or t , (for he makes all his tones equal,) consists of 5, a limma (or an approximate, or tempered value of θ) = 3, the interval between each note and its sharp or flat = 2, and the diesis = 1. This gives the following scale of intervals.

(1) (1) \sharp (2) \flat (2) (2) \sharp (3) \flat (3) (4) \flat (3) \sharp (4) (4) \sharp (5) \flat (5) (5) \sharp (6) \flat (6) (6) \sharp (7) \flat (7) (1) \flat (7) \sharp (1)
2 1 2 2 1 2 1 1 2 1 2 2 1 2 2 1 2 1 1 1

and by picking our notes among these, we may obtain a scale approaching extremely near to a perfect diatonic scale, whichever we may choose for our key-note.

Instead of dividing the octave into 31 equal parts, Dr. Smith proposes to divide it into 50, of which 8 shall constitute a tempered tone, and 5 a *limma*, or tempered value of θ , and the interval between each note, and its sharp or flat, shall = 3. This will give the sequence of intervals as below.

(1) (1) \sharp (2) \flat (2) (2) \sharp (3) \flat (3) (4) \flat (3) \sharp (4) (4) \sharp (5) \flat (5) (5) \sharp (6) \flat (6) (6) \sharp (7) \flat (7) (1) \flat (7) \sharp (1).
3 2 3 3 2 3 2 1 2 3 2 3 3 2 3 3 2 3 2 1 2

This scale, he observes, approximates insensibly near to what he terms the system of equal harmony, a system, in our opinion, uselessly refined, and founded on principles for which the reader is therefore referred to his *Work on Harmonics*, (Cambridge, 1749.)

Either system, no doubt, will give very good harmony; but as on the piano-forte only 12 keys can be admitted, and as this instrument is now become an essential element in all concerts, and indeed the chief of all, a temperament *must* be devised which will accommodate itself to that condition. Of the division of the octave into 12 equal parts we have already spoken. Its fifths are all too flat, and its major thirds all too sharp; and the harmony is equally imperfect in all keys. But it has generally been considered preferable to preserve some keys more free from error, partly for variety, and partly because keys with five or six sharps or flats are comparatively little used, so that these may safely be left more imperfect, (which is called by some throwing the *wolf* into these keys.) Dr. Young recommends as a good practical temperament to tune downwards six perfect fifths from the fundamental note, and upwards six fifths equally imperfect among themselves. Or, as he observes is more easily executed, to make the third and fifth of the natural scale perfectly correct, to interpose between their octaves the second and sixth, so as to make three fifths equally tempered, and to descend from the key-note by seven perfect fifths, which will complete the scale. (*Lectures on Natural Philosophy*, vol. i. lect. 33.)

The system called by Dr. Smith that of mean tones, or the vulgar temperament, supposes the octave divided into five equal tones and two equal limmas, succeeding each other in the order $a a \beta a a \beta$ instead of $\tau t \theta \tau t \theta$ as in the diatonic scale, and such that the third shall be perfect and the fifth tempered a little flat. These conditions suffice to determine a and β , for we have

$$\begin{aligned} 5a + 2\beta &= 1 \text{ octave} = 3\tau + 2t + 2\theta, \\ 2a &= 1 \text{ third} = \tau + t, \end{aligned}$$

and consequently

$$a = \frac{\tau + t}{2}; \quad \beta = \theta + \frac{\tau - t}{4};$$

or, (since $\tau + t = 09691$, and $\tau - t = 00540$), $a = 04845$ and $\beta = 02938$. And since the interval from the first to the fifth of the scale in this system is $= 3a + \beta = 2\tau + t + \theta - \frac{\tau - t}{4}$, it appears that this is *flatter* than a perfect fifth by the quantity $\frac{1}{4}(\tau - t)$, or a quarter of a comma. In this system the sharps and flats may be inserted by bisecting the larger intervals.

Mr. Logier has lately, in a Work of great practical utility and very extensive circulation among musical students, endeavoured to place the interpolation of the intermediate notes between those of the natural scale on *à priori* grounds, by assuming the flat seventh (7) \flat as the seventh harmonic of the fundamental note (1), that is to say, the note produced by subdividing into seven equal parts the length of a string whose fundamental tone is (1), or at least one of the octaves of that note. There is something ingenious in this idea. In the first place it completes the series of the 10 first harmonics or notes, whose vibrations are multiples by 1, 2, 3, 4, 5, 6, 7, 8, 9, 10, of those of the fundamental tone, which would thus be, in their order, (1), (1), (5), (1), (3), (5), (7) \flat , (1), (2), (3), or octaves of these, and thus derives five out of the twelve notes of the octave from one uniform principle. Again, it gives something like a plausible reason for the prominent importance of the chord of the flat seventh (see Art. 230.) in music. This chord, in fact, which, if we take (1) for a fundamental note, consists of the notes (1), (3), (5), (7) \flat , becomes in this point of view a *perfect concord*, consisting entirely of harmonics of (1), and its pulses will succeed each other on the ear in a cycle comprising four vibrations of the fundamental tone (1), five of the next (3), six of the next (5), and seven of the essential note (7) \flat , as represented in fig. 35. The succession of pulses in the common chord is also represented in the same figure, and its regularity and pleasing variety, even to the eye, explains its agreeable effect on the ear. It is for musicians to say, whether they can make up their minds to regard the discord of the seventh in the light of a perfect concord or no. There is certainly nothing at all *discordant* in the vulgar sense of the word, i. e. unpleasant in its Sound, and so far it may be regarded as at least "*discordia concors*," but so far from possessing the essential character of a concord, that the ear is satisfied in hearing it, and expects and desires no more; there is no discord which calls so urgently for resolution. But, although it be true, that the seventh harmonic of the fundamental note lies between its natural

Part II.

248.

Systems of
21 fixed
Sounds in
the octave.

249.
Huygens's
system.

250.

Dr. Smith's
system.

251.

Tempera-
ments
adapted for
the piano-
forte.

252.

System of
mean tones,
or vulgar
tempera-
ment

253.

Logier's
system of
harmony

Sound,

seventh and its octave, (it must lie somewhere,) yet, in fact, it is materially too flat a Sound to be used as a good flat seventh (7)^b. Its actual Sound coincides much more nearly with the (6)[#] of Huygens and Smith; and this defect, though it might be tolerated in quick compositions, and especially in piano-forte music where the notes are not held on, but degrade rapidly in intensity, would be at once felt in a slow piece on the organ. It is still worse if we derive *from it*, by a similar process, the intermediate note between (5) and (6), or (6)^b, and thence again (5)^b, and complete the chromatic scale of twelve notes by deriving (3)^b according to the same principle from (4), and (2)^b from (3)^b, according to Mr. Logier's system as laid down by him.* The (2)^b thus derived would hardly be distinguished from (1) *natural*, or the (5)^b from (4) *natural*, as the following scale will show, where the fractions represent the ratios of the vibrations of the notes above them to those of the fundamental tone (1).

(1), (2)^b, (2), (3)^b, (3), (4), (5)^b, (5), (6)^b, (6), (7)^b, (7), (8).
 1, $\frac{4}{8}$, $\frac{9}{8}$, $\frac{7}{4}$, $\frac{5}{4}$, $\frac{4}{3}$, $\frac{3}{2}$, $\frac{3}{2}$, $\frac{4}{3}$, $\frac{5}{3}$, $\frac{7}{4}$, $\frac{15}{8}$, 2.

Sevenths, then, tuned on Mr. Logier's principle, will require a much more violent temperament than either fifths or thirds, either of which *might* be used as a means of introducing the intermediate notes; and the system must in consequence be abandoned, as must every system which professes to render musical arithmetic any thing more than a matter of convention and approximation.

We annex here, for comparison, a Table of the logarithmic values of the intervals from (1) the fundamental tone to all the other notes in the several scales of 21, or of 12 notes, according to the different systems and principles above mentioned.

The numbers marked thus (*) are what would be given by pushing the application of Mr. Logier's principle through the whole scale, and are inserted only to show the rapid progressive effect of flattening by a series of untempered harmonic sevenths.

Comparative Table of different scales of temperament, &c

Designation of Note.	Intervals in the perfect Diatonic Scale.			Diatonic Scale with its tones bisected.	System of mean tones or vulgar temperament, Art. 222.	System of equal Temperament, Art. 242.	Dr. Young's first System, Art. 251.	Dr. Young's second System, Art. 251.	Mr. Logier's Scale.	Huygens's System of 21 notes, Art. 249.	Smith's approximate System of 21 notes, Art. 250.	Mean Scale.
	Ratio.	Log.	Approx. Log.									
(1)	1	00000	0	0	0	0	0	5	0	0	0	0
(1) [#]	—	—	—	25	24	25	23	23	9*	19	18	24
(2) ^b	—	—	—	—	—	—	—	—	—	29	30	—
(2)	$\frac{9}{8}$	05115	51	51	48	50	49	49	51	49	48	49
(2) [#]	—	—	—	—	—	—	—	—	—	68	66	—
(3) ^b	—	—	—	74	73	75	74	74	67	78	78	73
(3)	$\frac{5}{4}$	09691	97	97	97	100	98	97	97	97	96	98
(4) ^b	—	—	—	—	—	—	—	—	—	107	108	—
(4)	$\frac{4}{3}$	12494	125	125	126	125	125	125	125	117	114	125
(4) [#]	—	—	—	—	—	—	—	—	—	126	126	—
(5) ^b	—	—	—	150	150	151	148	148	127*	146	144	150
(5)	$\frac{3}{2}$	17609	176	176	175	176	175	176	176	155	157	176
(5) [#]	—	—	—	—	—	—	—	—	—	175	175	—
(6) ^b	—	—	—	199	199	201	199	199	185*	194	193	199
(6)	$\frac{5}{3}$	22185	222	222	223	226	224	224	222	204	205	224
(6) [#]	—	—	—	—	—	—	—	—	—	223	223	—
(7) ^b	—	—	—	247	247	251	250	250	243	243	241	248
(7)	$\frac{15}{8}$	27300	273	273	272	276	273	273	273	252	253	273
(8) ^b	—	—	—	—	—	—	—	—	—	272	271	—
(8) = (1)	2	30103	301	301	301	301	301	301	301	282	283	301
Comma ...	$\frac{81}{80}$	00540	5 $\frac{1}{2}$	—	—	—	—	—	—	291	289	—
Limma = θ	$\frac{16}{15}$	02803	28	—	—	—	—	—	—	301	301	—
Minor tone t	$\frac{9}{8}$	04576	46	—	—	—	—	—	—	—	—	—
Major tone r	$\frac{9}{8}$	05115	51	—	—	—	—	—	—	—	—	—
Minor third	$\frac{6}{5}$	07918	79	—	—	—	—	—	—	—	—	—
Major third	$\frac{4}{3}$	09691	97	—	—	—	—	—	—	—	—	—
Fourth	$\frac{4}{3}$	12494	125	—	—	—	—	—	—	—	—	—
Fifth.	$\frac{3}{2}$	17609	176	—	—	—	—	—	—	—	—	—

* *System of the Science of Music and Practical Composition*, p. 50. We should, however, remark that the powerful descending tendency of the chord of the flat seventh is necessarily much augmented by tuning the (7)^b too flat.

Part II.

Sound.

The last column contains a scale derived by taking the mean of all those in the other columns which differ in the principle of their origin, (excepting those in the tenth column, for obvious reasons.) It approaches through its whole extent so near to the system of mean tones in col. 6, as to be quite undistinguishable from it; the deviation in no case exceeding a single unit, or a fiftieth part of a tone. This system, then, though the most inartificial, is probably as good as any which the nature of music admits, holding a sort of mean between the advantages and defects of all the rest. Consult on Temperament and on Musical Scales, Salinas, *de Musicâ*, (1577;) Zarlino, *Dimostrazione Armoniche*, and *Istituzione Armoniche*; Deschales, *Cursus Mathematicus de Progressu Musico*; Sauveur, *Mém. Acad. Par.* 1700; Smith, *Harmonics*; Pepusch, *Phil. Trans. Lond.* 1746; Farey, *Phil. Mag.* xxviii.; Young's Lectures and his Catalogue of Authors in vol. ii.

Part II.

255.

Remarks.

For one purpose, that of explaining to beginners the notes, intervals, and rules of music, the system of equal temperament, which supposes the octave divided into 12 equal parts, which in this system only are *really* semi-tones, has the advantage of avoiding all discussions and puzzling explanations on the nature of harmony, as it makes all intervals which are called by the same name strictly alike. Regarding the octave as consisting of 12 semitones, and designating its notes in succession, beginning with the fundamental note, by 0, 1, 2, 3, 4, 5, 6, 7, 8, 9, 10, 11, 12, &c. it will not be amiss if we write down in this notation the principal scales, chords, &c. which occur in music.

256.

Expression of the principal chords by chromatic numbers.

Chromatic scale	0, 1, 2, 3, 4, 5, 6, 7, 8, 9, 10, 11, 12.	
Diatonic scale	0, 2, 4, 5, 7, 9, 11, 12.	
Minor scale ascending	0, 2, 3, 5, 7, 9, 11, 12.*	
Minor scale descending ...	12, 10, 8, 7, 5, 3, 2, 0.	
Common major chord.....	0, 4, 7,	
Minor chord	0, 3, 7,	
Fundamental discord of the flat seventh .	0, 4, 7, 10.	
Chord of the added sixth (Logier, <i>Ex.</i> 117.)	0, 4, 7, 9.	
Chord of the ninth	0, 2, 4, 7.	
Minor chord with added sixth	0, 3, 7, 9.	
Diminished seventh	0, 3, 6, 9.	
Chord of the sharp sixth (Logier, <i>Ex.</i> 197.)	0, 2, 6, 8.	

Notation of chromatic numbers explained.

These are all the chords, consisting of four different notes (or tetrachords) in common use in music. As to pentachords, such as what have been called the major and minor ninth, and compound sharp sixth, whose notes are respectively 0, 2, 4, 7, 10; 0, 1, 4, 7, 10, and 0, 4, 6, 7, 10, (Logier, *Ex.* 212, 158, and 202.) they are, in fact, only chords of the seventh (0, 4, 7, 10) with a fifth note violently forced in; the effect being to distract the ear by a harsh discord, out of which it is but too glad to escape, to be very nice about its resolution. In like manner the pentachord 0, 2, 5, 7, 11, or the chord of the eleventh, is the chord of the seventh accompanied by the sub-dominant of its radical note, and thus anticipating its resolution; as is easily seen by adding 5 to each of its component numbers, when it becomes 5, 7, 10, 12, 16, or, which is the same thing, 5, 7, 10, 0, 4, or 0, 4, 5, 7, 10, (since the addition or subtraction of 12 semitones, or an octave, does not alter the character of the Sounds;) and in the same way may other pentachords be formed, as 0, 3, 4, 7, 10; 0, 4, 7, 8, 10; (*Ex.* Clementi's *Sonatas*, Op. 22. son. 1. bars 68, 69.) 0, 4, 7, 9, 10. As to such combinations as the hexachord 0, 2, 5, 7, 9, 11, or "the chord of the 13th," (Logier, *Ex.* 273.) in which only one note of the whole natural scale (4) is wanting, they are abominable jangles, as offensive to a simple and unvitiated ear, as the mixed flavours and *haut-gouts* of the palled epicure are to an appetite not spoiled by artificial excitement.

257.

Triads, tetrachords, pentachords and hexachords.

The reader who would try these chords on the piano-forte, has only to place his finger on any black or white key as a radical note, and also on the keys distant from that one by the numbers of semitones (reckoning upwards) marked in the designation of the chord as above. Thus to produce the chord of the sharp sixth having D \flat for its radical note. The note 0 corresponds to D \flat , 2 to E \flat , 6 to G \flat , and 8 to A \flat , which are, therefore, the notes to be struck together, (to whatever octaves of the instrument they may be afterwards transferred, as the rules of composition may dictate,) and so of others.

258.

Any of these chords is said to be inverted, when, instead of taking 0 for the initial note, we regard any other of its component Sounds as such. On the system of notation here employed, (which we will term the system of *Chromatic numbers* to distinguish it from those in Art. 234, to which the term *Diatonic numbers* may be applied,) nothing is easier than to represent the inversions of any chord. Take, for instance, the major concord, 0, 4, 7. The addition of 12 (the octave of 0) does not change the chord; so that it may be written thus, 0, 4, 7, 12, or, leaving out the first note, and adopting 4 for the initial note, 4, 7, 12. If, now, we choose to regard the note 4 as an initial one, and count upwards from it, we have only to subtract 4 from each of these numbers, and we get 0, 3, 8 for a first inversion. Appending 12 to this again, and rejecting the initial 0, it becomes 3, 8, 12, from each of which numbers subtracting 3 we get the second inversion, 0, 5, 9. If we repeat the same process on this we fall back on the original combination. Thus we see that this chord admits of only two inversions. Again, suppose we would find the inversions of the chord of the added sixth, or 0, 4, 7, 9. The process will stand thus:

259.

Inversions of a chord.

* 8, 11, 12, according to Logier and others. *Vide*, on this point, Weber's excellent and scientific work, *Tonsetzkunst*.

Sound.

Part II.

0, 4, 7, 9, 12	
4, 7, 9, 12	
4, 4, 4, 4	
0, 3, 5, 8	1st inversion.
3, 5, 8, 12	
3, 3, 3, 3	
0, 2, 5, 9	2d inversion.
2, 5, 9, 12	
2, 2, 2, 2	
0, 3, 7, 10	3d inversion.
3, 7, 10, 12	
3, 3, 3, 3	
Original chord again....	0, 4, 7, 9

Thus we see that this chord admits of three distinct inversions. In general, a triad admits of three forms, or one original, and two inversions, a tetrachord of 4, a pentachord of 5, and so on; though it may happen, as in the case of the triad 0, 4, 8, or the tetrachords 0, 2, 6, 8, and 0, 3, 6, 9, that some or all of the inversions reproduce the original chord.

260. If we go through the same process for other triads and tetrachords, we get their inversions as follows:

Triads.

	1st form, or radical.	2d form, or 1st inversion	3d form, or 2d inversion
Major concord.....	0, 4, 7	0, 3, 8	0, 5, 9
Minor concord.....	0, 3, 7	0, 4, 9	0, 5, 8
Equivocal triad, or double third.....	0, 4, 8	0, 4, 8	0, 4, 8

Tetrachords.

	1st form.	2d form.	3d form.	4th form.
Seventh.....	0, 4, 7, 10	0, 3, 6, 8	0, 3, 5, 9	0, 2, 6, 9
Added sixth.....	0, 4, 7, 9	0, 3, 5, 8	0, 2, 5, 9	0, 3, 7, 10
Ninth.....	0, 2, 4, 7	0, 2, 5, 10	0, 3, 8, 10	0, 5, 7, 9
Minor added sixth.....	0, 3, 7, 9	0, 4, 6, 9	0, 2, 5, 8	0, 3, 6, 10
Triple fifth.....	0, 2, 7, 9	0, 5, 7, 10	0, 2, 5, 7	0, 3, 5, 10
<i>Equivocal.</i>				
Diminished seventh.....	0, 3, 6, 9	0, 3, 6, 9	0, 3, 6, 9	0, 3, 6, 9
Sharp sixth.....	0, 4, 6, 10	0, 2, 6, 8	0, 4, 6, 10	0, 2, 6, 8

Pentachords.

	1st form.	2d form.	3d form.	4th form.	5th form.
Minor ninth (Logier, Ex. 158.)	0, 1, 4, 7, 10	0, 3, 6, 9, 11	0, 3, 6, 8, 9	0, 3, 5, 6, 9	0, 2, 3, 6, 9
Compound sharp sixth (Log. Ex. 202)	0, 4, 6, 7, 10	0, 2, 3, 6, 8	0, 1, 4, 6, 10	0, 3, 5, 9, 11	0, 2, 6, 8, 9
Major ninth (Log. Ex. 212)	0, 2, 4, 7, 10	0, 2, 5, 8, 10	0, 3, 6, 8, 10	0, 3, 5, 7, 9	0, 2, 4, 6, 9
Eleventh (Log. Ex. 267)	0, 4, 5, 7, 10	0, 1, 3, 6, 8	0, 2, 5, 7, 11	0, 3, 5, 9, 10	0, 2, 6, 7, 9

261.
Remarks.

These chords, thus figured and arranged, afford room for some remarks of importance. In the first place we observe that they all, with the exception of the triad 0, 4, 8, and the tetrachords 0, 2, 6, 8, and 0, 3, 6, 9, contain a major or minor concord, 0, 4, 7, or 0, 3, 7. This seems necessary to give any chord a decided character; for the excepted cases above specified have all an equivocal effect and leave the ear in suspense whither the modulation will lead. For with respect to the chords 0, 4, 8, and 0, 3, 6, 9, they divide the octave equally, the one into three major thirds, the other into four minor, as is immediately seen if we write them thus, 0, 4, 8, 12, and 0, 3, 6, 9, 12. In consequence, all their inversions are similar to the original chords, and they are equally related, the former to three, and the latter to four different keys, and may lead into either of them, according as a

Aliquot
division of
the octave.

Sound.

note added so as to form a dominant seventh, or anticipative sub-dominant, or some other powerful leading interval, or with either of their component notes, shall decide. This is one mode of conceiving the chord of the minor ninth, which may be either regarded as a chord of the seventh, with the first semitone 1 added, (as in its first form above,) or as a diminished seventh, 0, 3, 6, 9, with the note 11 added, as in the first inversion, with 8 as in the second, or with 2 as in the third, either of which makes a flat seventh with one or other of its notes.

Part II.

The transitions thus produced by means of the tetrachord 0, 3, 6, 9, are peculiarly graceful. It is otherwise with the equivocal triad 0, 4, 8, which is essentially harsh and displeasing, (in spite of the *perfect* harmony which, if we were to leave out the octave and tune its thirds perfect, its members must produce with each other, since it would be in that case an absolute concord.) Whether this chord, or that which we have called the triple fifth, has ever been, or can be, used in music, we know not, though perhaps, properly handled, it might become a source of modulation; which, however, is for practical musicians to consider.

262.
Equivocal
chords.

The chord of the sharp sixth 0, 2, 6, 8 is also equivocal, arising from a double aliquot division of the octave, and the two last of its inverted forms being therefore merely repetitions of the two first. Like the diminished seventh, then, it holds the ear in suspense, till the addition of another note decides the course the modulation shall take, and the chord so arising is the compound sharp sixth. (See the inversions of this latter chord compared with those of the former.)

263
Relations
between
the tetra-
chords and
pentachords

In like manner the *major* ninth contains both a ninth and a seventh, though not the other accompaniments of the seventh. The tetrachord which (for want of another name, we have called the minor added sixth, from its being a minor concord with a sixth added) is related to this compound ninth in the same way as has just been pointed out with respect to the chords of the diminished seventh and minor ninth, and to those of the sharp sixth and compound sharp sixth; the character of the tetrachord, which is undecided of itself, and admits of more than one resolution, being determined by the note added in the pentachord so as to form a dominant seventh with some one or other of its other members.

The chord of the eleventh offers room for a remark analogous to what we have before observed (Art. 262.) respecting the equivocal triad 0, 4, 8. It contains within itself three fifths and a major third; as is obvious if we take its fifth form 0, 2, 6, 7, 9, and transfer the notes 2, 6, and 9 to the next octave above, when it will become 0, 14, 18, 7, 21, or 0, 7, 14, 18, 21. The notes 0, 7, 14, 21, in this arrangement, make fifths with each other, and the note 18 forms with 14 a major third; if, then, the intervals were tuned perfect, their vibrations would succeed each other in a regular cycle, but if the cycle formed by two perfect thirds, which requires only 25 vibrations of its highest note, or 16 of its lowest to complete it, is too complex for the ear to relish, the cycle of three perfect fifths, which requires 27, will already be too complex; and if we add to this a major third, the ear will lose all sense of recurrence, and only discord will result.

264.
Chord of
the eleventh
examined.

But to place this in clearer evidence, we need go no further than the chord of the ninth, which, when written thus, 0, 4, 7, 14, manifests a major third, (0, 4,) a fifth, (0, 7,) and a double fifth, (0, 7 + 7,) of the fundamental note, and therefore, if tuned perfect, would excite a sense of perfect concord, were not the period of recurrence of the vibrations too long for the ear to seize; and a similar remark applies to the discord of the seventh, which consists of a major third, a fifth, and a double fourth, from the fundamental tone (0, 4, 7, 5 + 5.) It may be that the harshness of the triad 0, 4, 8, and of the tetrachord 0, 2, 7, 9, the former consisting, if tuned perfect, of a third and double third, the latter of a fifth, a double fifth, and a triple fifth, may arise from an imperfect, or obscure, and therefore unsatisfactory, perception of the cycles of their vibrations by the ear, the former, as before remarked, occupying 25, and the latter 27, single vibrations of the highest note. But it is time to leave these speculations.

265.
Chords of
the ninth
and seventh
analysed.

§ V. Of the Sonorous Vibrations of Bars, Rods, and Plates

The vibrations of all bodies, if of a proper degree of frequency, and of sufficient force to be communicated through the air, or any other intermedium, to our organs of hearing, produce Sounds whose pitch depends on their frequency; and their force and quality on the extent and other mechanical circumstances of the vibrations, and the nature of the vibrating body. The mathematical investigation of these vibratory motions is altogether foreign to our purpose. It is a branch, and one of the most intricate and least manageable branches, of Dynamics, and we shall, therefore, refer our readers for its theory and details to the writings of the various eminent authors who have discussed it. See Bernouilli, *Com. Petrop.* vol. xiii. On the Vibrations of Laminæ; and *Nov. Com. Petrop.* vol. xv.: Euler, *Com. Petrop.* vol. vii., *Nov. Com.* vol. xvii., and *Act. Petrop.* vol. iii. On the Vibrations of Plates; Riccati, *Soc. Ital.* vol. i. p. 444; Lexell, On the Vibrations of Rings, *Act. Petrop.* 1781; Lambert, On the Sounds of Elastic Bodies, *N. Act. Helv.* vol. i.; J. Bernouilli, On the Vibrations of Rectangular Plates, *N. Act. Petrop.* 1787; Biot, On the Vibrations of Surfaces, *Mem. Inst.* vol. iv.

266

A solid body may vibrate, either in consequence of its inherent elasticity, by which it tends to return to its own proper figure and state when forcibly deranged, or in consequence of an external tension. To the former sort of vibrations belong those of rods, tuning-forks, plates, rings, bells, gongs, and vessels of all shapes, or generally, of all solid masses which *ring* when struck. To the latter, those of vibrating strings and membranes, such as the parchment of a drum or tambourin, &c. But, further, a solid may vibrate by its own proper elasticity in two very different ways. First, an undulation may be propagated through it, as through an elastic compressible medium, and, in this case, the waves will consist of alternate strata of condensed and rarefied solid matter, precisely similar to those of an elastic fluid. If the solid be homogeneous, such as the metals, glass, &c., the elasticity being the same in all directions, the waves will be propagated from the centre of disturbance, according

References.

267
Various
ways in
which
solids may
vibrate.

Sound to exactly the same laws as in a mass of air of the same shape. But if crystallized, this may not be the case, or the vibrations instead of being in the direction of the propagated wave, may be transverse, or oblique to it, or may even not be confined to one plane, but may be performed in circles or ellipses. See Article LIGHT.

268.
Longitudi-
nal vibra-
tions of a
straight rod.

How pro-
duced.

If a straight rod of glass, or a metal, to be struck at the end in the direction of its length, or rubbed lengthways with a moistened finger, it will yield a musical Sound, which, unless its length be very great, will be of an extremely acute pitch; much more so than in the case of a column of air of the same length. The reason of this is the greater velocity with which Sound is propagated in solids than in air. Thus the velocity of propagation in cast-iron being $10\frac{1}{2}$ times that in air, a rod of cast-iron so excited will yield for its fundamental note a Sound identical with that of an organ pipe of $\frac{1}{10\frac{1}{2}}$ of its length stopped at both ends, or $\frac{1}{21}$ of its length if open at one end. See § III., all the details of which are applicable to the present case. To such vibrations, Chladni, who first noticed them in long wires, has applied the term longitudinal. (*Art. Acad. Erfurt*, 1796.) To produce the harmonics of such a rod or wire he held it lightly at the place of one of its intended nodes between the finger and thumb, and applied the friction in the middle of one of the vibrating segments. If the rod be of metal, the friction which he found to succeed, was that of a bit of cloth sprinkled with powdered rosin, if of glass, the cloth, or the finger, may be moistened and touched with some very fine sand or pumice powder. It may be observed here, that, generally speaking, a fiddle-bow well rosined is the readiest and most convenient means of setting solid bodies in vibration. To educe their gravest or fundamental tones, the bow must be pressed hard and drawn slowly, but for the higher harmonics, a short swift stroke with light pressure is most proper. In all cases the point intended to be a node must be lightly touched with the finger, and the vibration must be excited (as above said) in the middle of a ventral segment. Such is the case analysed by Chladni. In general, however, the vibrations of a cylindrical rod or tube so excited are more complex. See Art. 286, Index, *Art. Longitudinal Vibrations*.

269.
Transverse
vibrations of
a rigid rod.

Fig. 36.
Enumera-
tion of
cases.
Fig. 37.

Fig. 38.
Fig. 39.
Fig. 40.
Fig. 41.

270.
Examina-
tion of one
of them.
Figs. 39, 43,
44.

271.
Remark on
the origin
of harmony.

272.
Vibrations
of a rectan-
gular plate
in their
simplest
case.

273.
Other cases.

274.

But by far the most usual species of vibration executed by solid bodies is that in which their external form is forcibly changed, and recovered again by their spring. The simplest case is that of a rod executing vibrations to and fro in a direction transverse to its length. This case has been investigated mathematically by D. Bernouilli and Euler, as also by Riccati; (see the list of authors above cited, Art. 266;) and their results have been compared with those of experiment by Chladni, *Acoust.* sec. 5, and found correct. The cases enumerated by Chladni are six in number.

1. When one end of the rod is firmly fixed in a vice or let into a wall, the other quite free. In this case the curvature assumed by the rod in its vibrations must of necessity have its axis or position of rest for a tangent, as fig. 36.

2. One end *applied* or pressed perpendicularly against an obstacle, the other free. In this case, the *excursions* of the applied end to and fro are prevented by the friction and adhesion to the obstacle, but the axis is not of necessity a tangent. See fig. 37.

3. Both ends free. Fig. 38.

4. Both ends applied. Fig. 39

5. Both ends fixed. Fig. 40.

6. One end fixed, the other applied. Fig. 41.

All these cases have been examined by Chladni at length. We shall, however, select only the fourth case where both ends are applied, because it will afford room for an important remark. In this, then, the several modes of vibration corresponding to 1, 2, 3, 4, 5 vibrating or ventral segments of the rod will be as in figs. 39, 43, 44. Now these are similar to the curves which would be assumed by a vibrating string under the same circumstances of subdivision. But the notes produced are very different. For whereas in the case of a string the vibrations of the successive harmonics are represented by 1, 2, 3, 4, 5, &c.; in that of a rod they are represented by the *squares* of these numbers 1, 4, 9, 25, &c., which correspond to *double* the former intervals. In all the other cases the series is still less simple.

This alone suffices to shew the insufficiency of any attempt to establish, as some have wished to do, the whole theory of harmony and music on the aliquot subdivision of a vibrating string. Had vibrating rods or steel springs (which yield an exquisite tone) been always used instead of stretched chords, such an idea would never have suggested itself, yet no doubt our notions of harmony would have been what they now are. The same remark applies still more forcibly to the modes of subdivision of vibrating surfaces, which in many cases have their harmonics altogether irreducible to any musical scale.

A rectangular plate may be regarded as an assemblage of straight rods of equal length, ranged parallel to each other. Supposing such an assemblage all set in vibration similarly and at once, they will retain their parallel juxtaposition during their vibration, and may, therefore, be supposed to adhere, and form a plate. Consequently, among the possible series of vibrations of a rectangular plate will be found all those of a rigid rod. Accordingly, when fixed, (for instance,) by one of its edges in a vice, with its plane parallel to the horizon and strewed over with sand, if it be set in vibration by a fiddle-bow and touched in one of its possible nodes, its subdivisions will be rendered visible to the eye, by the sand being thrown away from the vibrating parts and accumulating on those at rest. Thus the plate will be crossed transversely by a series of *nodal lines* marked in sand, and whose distances from each other and from the ends of the plate may be measured at leisure.

But besides these, rectangular plates are susceptible of other modes of subdivision, having two sets of nodal lines, straight or curved, crossing at right angles, or otherwise, and dividing the plate into smaller plates, each vibrating in its middle, and at rest at its edges, and every two contiguous plates separated by a nodal line making their simultaneous excursions on contrary sides of their state of rest.

To produce these subdivisions, and to render them visible, take a rectangular plate (for simplicity we will

Sound. suppose it a square) of glass, or metal, of an even thickness, not too thick, and holding it firm between the points of the finger and thumb of the left hand, or between two points of a clamp-screw covered with cork or leather so as not to jar, taking care to keep the pressure confined to as small a space as possible, draw a rosined bow over the edge, which should be smoothed and a little rounded. If then the point where it is held be the centre of the plate, and the bow be applied close to one of the angles, sand strewed over it will arrange itself on the two diameters which divide it into four equal squares as in fig. 44. Each of these, in the act of vibration, becomes a surface of double curvature, and their motions are contrary to each other; those marked + making their excursions on one side of the plane of repose, while those marked - are on the other. This mode of vibration corresponds to the gravest tone produced by the plate. Part II.

If, the plate being still held in the centre, the bow be applied at the middle of one side, the sand will occupy the diagonals of the plate, which are the nodal lines corresponding to this mode. In this, as in the former case, the plate subdivides itself into four equal vibrating segments as in the fig 45, but the tone is different, being a fifth higher than in the former case, the distribution of the inertia with respect to the elastic power of the plate being such as to admit a quicker motion. Mode of producing their several subdivisions. First mode of a vibration of a square plate. Fig. 44.

If the plate be held at a , the intersection of two nodal lines fig. 46, and the bow be still applied at the middle of one side, or at the angle adjacent to a , the plate will vibrate as there represented. In this subdivision, the four small squares at the angle and the large one at the centre vibrate on one side, or negatively, while the four intermediate oblong rectangles adjacent to the sides vibrate positively. Second mode. Fig. 45. 275. Fig. 46. Other modes. 276.

These instances may serve to show the mode of proceeding in more complicated cases, and with plates of other figures. Among these, circular ones hold the chief place both for symmetry and variety. The examples, figs. 47—93, are selected from those described by Chladni, who has determined by experiment the tones corresponding to each mode of division in plates of a great variety of figures. Of these we shall only give some examples in the case of a square plate, of which we shall suppose the gravest or fundamental tone to be represented by 1. This premised, if we regard the plate as subdivided into $n \times n'$ rectangles by n nodal lines parallel to one side, and n' parallel to the other, the notes corresponding will be as in the following Table: 277. Fig. 47—93. Modes of vibration of square and circular plates observed by Chladni.

Values of n' .	Values of n .					
	0	1	2	3	4	5
0				(1) ⁱⁱⁱ # +		
1		(1)	(3) ⁱⁱ		(3) ^{iv} b -	
2		(3) ⁱⁱⁱ	(3) ⁱⁱⁱ b -	(7) ^{iv}		(1) ^v # +
3	(1) ⁱⁱⁱ b +		(7) ^{iv}	(4) ^v	(7) ^v	
4		(3) ^{iv} b -		(7) ^v	(3) ^v b	
5			(1) ^v # +			

Series of Sounds produced by a square plate.

The vibrations of triangular, hexagonal, elliptic, and semicircular plates have also been investigated by Chladni, and fig. 94—123 exhibit some out of a great variety of nodal figures, to which they give rise in their various modes of vibration. 278 Fig. 94—123.

PART III.

§ I. *Of the Communication of Vibrations and of the Vibrations of Systems*

Part III.

Sound.

279.
Savart's
researches
on the vi-
brations of
solids.

THE subject of the sonorous vibrations of solids has recently been taken up in a more general and extended point of view by M. Felix Savart, in a series of Memoirs communicated by him to the Royal Academy of Sciences of Paris, and of which copies, or copious extracts, are printed in the *Annales de Chimie*. We regret that the narrow limits which remain to us in this volume, will allow little more than a slight sketch of the contents of the principal of these most interesting papers, the whole of which are models of experimental research, and indeed, so full of new, curious, and instructive matter, that it is next to impossible either to condense or abstract them; for which reason we earnestly recommend our readers, who may be led to take an interest in the subject of this Essay, not to content themselves with the meagre statements here offered, but to procure and study diligently the original Memoirs.

280.
Method of
communi-
cating a
given vibra-
tion to a
given point
of a solid.

In order to a regular analysis of this intricate subject, it was first requisite to obtain some certain mode of communicating to any given point of a solid vibrations confined to one plane, and whose period of recurrence, as well as the plane in which they were performed, and the amplitude of their excursions, could be varied at pleasure. The vibrations of a stretched string set in motion by a fiddle-bow, afford the means of doing this. Such are necessarily confined to the plane in which the motion of the bow is performed, because any vibratory motion out of this plane is prevented, or immediately stifled by the pressure of the bow; and as the plane of its motion may be varied at pleasure, and the amplitude of excursion may be increased or diminished by a change of pressure, and velocity of stroke, all the requisite conditions are here obtained. Accordingly, if the vibrating part of such a string be brought to press on a solid not too massive, or if the end of the string be attached to a point in the solid, M. Savart has found that the regularly repeated impulses of the string are transferred to the solid with perfect fidelity.

281.
Vibrations
of a violin
string how
communi-
cated to the
wood.
Fig. 124.

A familiar example of this communication of impulses is found in the violin. In that instrument, fig. 124, the strings which are stretched from end to end of it, are divided into two unequal parts by the bridge, A, on which they all press strongly, and at the same time rest in small notches, so as not to slip laterally on it. The portion, B, of the string which lies towards the handle, C, of the instrument, is free, and is set in vibration by the bow in its own plane; but that on the other side of the bridge, D, is loaded with a mass of horn or whalebone, E, to which all the other strings are also attached, and which, being only *tied* to the wood-work, cannot propagate the vibrations of any one string sounding separately, by reason of the contradictory and unequal tensions of the other three. Thus the bridge is in fact acted on only by the vibrations of that part, A B C, of the string which is crossed by the bow, as if it terminated abruptly at its point of pressure, A. These vibrations constantly tend, therefore, to tilt the bridge laterally backwards and forwards, and to press up and down alternately the two little prominences or feet, F G, by which it rests on the belly of the violin. It, therefore, sets the wood of the upper face in a state of regular vibration, and this again is communicated to the back through a peg set up in the inside of the fiddle, and through its sides, called the *soul* of the fiddle, or its *sounding post*. In consequence, if the upper surface be strewed with sand, it will assume a regular arrangement in nodal lines when the bow is drawn; and the same subdivision is also observed in the wood of the under surface, if the sounding-post be exactly placed in the centre of symmetry of the nodal figures. The experiment can hardly be made, however, with a common fiddle, by reason of the convexity of its surface, on which sand will not rest; but if one be constructed with plane boards, or if, abandoning the fiddle, a string be stretched on a strong frame over a bridge, which is made to rest on the centre of a regularly formed plate or disc of metal or wood, strewed with sand, the surface thus set in vibration by the string will be seen to divide itself by regular nodal figures.

Vibrations
of the wood,
how ob-
served.

Plates made
to vibrate
by commu-
nication with
strings.

282.
Joint vibra-
tions of a
plate and
string as a
system.

Now M. Savart has observed this remarkable fact, *viz.* that if the tension or length of the string thus placed in vibratory communication with a plate, be changed, so as to vary the note it speaks, the nodal figures on the plate undergo a corresponding variation, and the plate *still* vibrates in unison with the string; or, which is the same thing, the two, together with the interposed bridge, form a vibrating system, in which, though the vibrations of the several parts are necessarily very different in their nature and extent, yet they have all the same periods. This experiment is very important. It shows that the Sounds of such thin plates are not like those of strings confined to certain fixed harmonics, but, according to the forms of their nodal lines, and the proportions of the vibrating areas in opposite states of excursion, may assume any assigned period; in other words, given the vibrating plate and the pitch, a nodal figure may be described on it, which shall correspond to that pitch, and the plate (with more or less readiness, however) is always susceptible of such a vibration as shall yield that note, and produce that nodal figure. How far this proposition is general, and with what limitations it is to be understood, we shall soon see. Meanwhile this remark, it will be observed, furnishes a complete explanation of the effect of sounding-boards in musical instruments. It is not, as some have supposed, that there exist in them fibres in every state of tension, some of which are therefore ready to vibrate in unison with any proposed Sound, and, therefore, reinforce it. Such a cause could at best produce but a very feeble effect. It is the *whole* board which vibrates as part of a system

with every note, and (as vibrations may be superposed to any extent) the same sounding-board may at once form a part of any number of systems, and vibrate in unison with every note of a chord. Still some modes will always be more difficult than others, and no sounding-board will be perfectly indifferent to all Sounds.

The longitudinal vibrations of a rod of glass, excited by rubbing it with a wet cloth, may also be used to excite vibrations in a given point of a solid perpendicular to its surface, by applying its end to it, or cementing it to the solid by mastic. In this way Chladni applied it to draw forth the Sounds of glass vessels, (which when hemispherical, and of sufficient size and even thickness, are remarkably rich and melodious,) in an instrument which he called the *Euphone*, exhibited by him in Paris and Brussels. The principle of this instrument was at the time concealed; but the enigma was subsequently solved by M. Blanc, who on his part independently made the same remark, and applied it to a similar purpose.

If the solid (a circular glass disc for instance) to which such a vibrating rod or tube is fastened, be of small comparative dimensions, its vibrations are commanded by those of the rod, and the Sound yielded will be that of the rod alone; and *vice versa*, if the disc be large, and the rod small, the note sounded will be that of the disc, which will entirely command the rod; but in the intermediate cases, both M. Savart and M. Blanc have observed, the note will be neither that of the disc or the rod separately, but the two will vibrate together as a system, each yielding somewhat to the other. It is a case exactly analogous to that of a reed-pipe, in which the reed and column of air mutually influence each other's note. See Art. 199. This mutual influence of propagated motion, by which two periodically recurring impulses affect each other's period, and force themselves into synchronism, extends to cases where at first sight it would hardly be suspected. Thus Ellicot observed that two clocks fastened to the same board, or even standing on the same stone pavement, beat constantly together, though when separated their rates were found to differ very considerably; and Breguet has since made the same remark on watches. Thus also two organ-pipes vibrating side by side, if very nearly in unison, will under certain circumstances force themselves into exact concord, as has been observed by Hudlestone, (*Nicholson's Journal*, i. 329.) and lately recalled to notice by some experiments made in Copenhagen. The experiment with the disked tuning-fork and pipe, related in Art. 204., may here again be referred to.

The longitudinal vibrations of a rod have also been used by M. Savart, to communicate vibrations from one solid to another; as for instance, from the upper to the under of two circular discs cemented at their centres to the two ends of the rod, at right angles to their planes, as at fig. 125. If the two discs be of the same dimensions and materials so as to yield, when separately vibrating, the same note, the vibrations of one of them, (the upper for instance,) excited by a bow, will be exactly imitated by the other, and sand strewed over both will arrange itself in precisely the same forms in both discs, and that, into whatever number of vibrating segments that immediately excited be made to subdivide itself. But if the discs separately do not agree in their tones, the system may yield a tone intermediate, and each being differently forced from its natural pitch, the nodal figures on them will then no longer correspond.

The state of vibration in which the molecules of the connecting rod are thrown in such cases, deserves a nearer examination. For simplicity let us suppose the discs equal, the rod cylindrical, and the vibration of the system such that each disc shall subdivide itself into four quadrantal segments. In this case it is clear that as the form assumed at any instant by the upper disc is undulated or wrinkled, as represented in fig. 126, the section of the rod in immediate contact with it, and which obeys all its motions, must assume a similar form, and so of all the rest. Thus if we conceive the rod split into infinitesimal columns, parallel to its axis, all the columns in two opposite quadrants will be ascending, while those in the other two are descending; and thus the two corresponding opposite quadrants of the lower plate will be drawn upwards, while the alternate ones are forced downwards, giving a similar distortion to its figure, and disposing it to a similar vibration only. It will depend on the length of the rod, and the time taken by an undulation to run over its length, compared with that of a vibration of either disc, whether the *phases* of vibration in the two discs shall be the same at the same instant or not. It may happen that, for instance, the quadrant, D B, of the upper disc shall have completed its downward motion, and begun to return before the pulsation propagated through the rod has arrived at the lower disc; and in that case the corresponding quadrants of the two discs will be always in opposite phases of their periodic motion. But the nodal lines will of necessity correspond in both.

When the two discs are unequal, the propagation of the pulses through the rod must of course cease to be uniform, and each section of it down its whole length will have its own peculiar law of form and motion, which it is beyond our power to investigate. In that case its molecules must have lateral as well as vertical motions, and its vibrations must be partly *longitudinal* and partly *twisting*, in a way easier imagined than described. If the discs be dissimilar in form as well as unequal in dimension, the vibrations of the connecting rod will of course be very complicated.

These principles have been applied by M. Savart, and apparently with success, (as appears by the very able report of M. Biot on his Paper,) to the improvement of violins, and the construction of these delicate instruments on scientific and experimental grounds. Every one is aware of the difficulty of procuring perfect violins, and the enormous prices they bear, so that fixed rules, by which any ordinary artist can with certainty produce an excellent one, are evidently highly valuable. We long to see M. Savart's construction tried in this country, but must refer to his Paper (*Annales de Chimie*, vol. xii, p. 225, &c.) for the details.

It appears from what we have said, that the motions of the molecules of a rod which communicates the vibrations of one disc to another, or, more generally, which vibrates longitudinally by any exciting cause, are not of necessity analogous to those of the air in a cylindrical pipe, at least not to that simple case of the latter vibrations, which we have heretofore considered in our 3d Section. The several transverse sections of such a rod, in the act of vibration, do not necessarily merely advance and recede longitudinally, but may become curves of double curvature; in short, such a rod may be considered as an assemblage of vibrating discs, ranged along a common axis, along which

Part III.

283.

Longitudinal vibrations of a rod employed to communicate vibrations to solids.

Chladni's *Euphone*.

284.

Mutual influence of a vibrating disc and a rod connected with it.

Of two clocks placed near together.

Of organ-pipes nearly in unison.

285.

Fig. 125. Vibrations communicated between two plates by a rod.

286.

State of vibration of the connecting rod examined. Fig. 126.

287.

Case where the discs are unequal.

288.

M. Savart's violins.

289.

Longitudinal vibrations of a rod further examined.

Sound. they may, it is true, be also carried backwards and forwards with a vibratory motion, while at the same time their flexure is changing from convex to concave, and *vice versâ*. Now it may happen that a point, or a line, (straight or curved,) in any one of such discs, may be advancing in the direction of the axis in consequence of the bodily motion of the whole disc, while, in virtue of its flexure in the act of changing its figure, it may be receding; and this advance and recess may so balance each other, that the point or line shall be at rest. If this be true at one instant, it will be so at all instants, because the vibrations have all one period, and follow the same law of increase and decrease in their phases. Thus we have a nodal point, or a nodal line; and as each disc, by reason of the law of continuity, must have a similar one, the assemblage of such lines will mark out within the rod a *nodal surface*, dividing it into separate solids whose molecules on either side of such surface are in opposite phases of their motion.

290. What is here said of rods, applies of course to solids of any figure and dimension, neither is there the slightest reason why it should not apply to vibrating masses of air, or any other elastic fluid. Any such mass may be conceived as cut up into two or more oppositely vibrating portions pervading it according to certain laws. Where these surfaces *out-crop* or intersect the external surface of the mass, there will be a *nodal line*.

291. Such nodal lines, formed on the surfaces of bodies by the longitudinal vibrations of their molecules, (*i. e.* by vibrations *parallel* to their surfaces,) may be detected and rendered visible to the eye by fine dry sand, or the powder of Lycopodium, strewed over them; and the motions of the particles in the act of forming them will easily distinguish such vibrations as are executed parallel to the surface (in which, of course, the surface is not thrown into waves) from such as take place at right angles to it, where the surface itself leaps up and down. In the latter case, the particles of sand *dance*, and are violently thrown up and down over the whole extent of the vibrating portions, till, at length, they are entirely dispersed from them. In the former, they only *glide* along close to the surface, and meet and settle on the nodal lines, and that, sometimes, with incredible swiftness. The reason why they retreat to the nodal lines is easily understood. The amplitude of the excursions of the vibrating molecules of the surface diminishes as we approach a nodal line. Hence a particle of sand anywhere situated, if thrown by an advancing vibration *towards* this line, will not be thrown quite so far back by the subsequent retreating vibration, because its *then* situation is one less agitated. Thus the motion of each particle of sand is one of alternate advance towards the node and recess from it, but the advances are always greater than the recesses. In consequence, it *creeps* along the surface, and will not rest till it has attained the node. When a large disc of glass is set vibrating vigorously by a bow, perpendicular to its plane, the grains of sand will fly up some inches from it and be scattered in all directions. M. Savart has distinguished by the name *tangential* vibrations all such motions of the superficial particles of a body as are performed parallel to the surface; while those executed at right angles to it, in virtue of which the surface itself heaves and sinks, he calls *transverse*; and to motions compounded of both these, where the surface both swells and falls and shifts laterally backwards and forwards, he gives the term "*oblique vibrations*." In this we shall follow him.

292. This acute experimenter has investigated with great minuteness the *tangential* vibrations of long flat rods or rulers of glass, as well as of cylinders and tubes. They are extremely complicated, and offer most singular phenomena, some of which we shall now describe. If we take a rectangular lamina of glass $0^m.70$ ($= 27^{in}.56$) long, $0^m.015$ ($= 0^{in}.59$) broad, and $0^m.0015$ ($= 0^{in}.06$) thick, and holding it by the edges in the middle between the finger and thumb with its flat face horizontal, strewed with sand, and, at the same time, set it in longitudinal vibration, either by rubbing its under side near either end with a bit of wet cloth, by tapping it on the end with light blows, or by rubbing lengthwise a very small cylinder of glass, cemented on to its end in the middle of its breadth, and parallel to its length; in whatever way the vibration be communicated, we shall see the sand on its upper surface arrange itself in parallel lines, at right angles to its longer dimension, and *always, in one or the other of the two systems*, represented in figs. 127 and 128. Now it is very remarkable that although the same one of these two systems will always be produced by the same plate of glass, yet among different plates of the above dimensions, *even though cut from the same sheet, side by side*, one will invariably exhibit one system, and the other the other, without any visible reason for the difference. Moreover, in the system, fig. 127, the disposition of the nodal lines is unsymmetrical, one of them, *a*, being nearer to one end, and the closer pair, *ff'*, not being situated in the middle; and this too is peculiar to the plate, for wherever it be rubbed, whichever end be struck, still the line *a* will always be formed nearest to the *same* extremity.

293. Now let the positions of the nodal lines be marked on the upper surface, and then let the plate be turned till the lower surface becomes the upper, and this being sanded, let the vibrations again be excited just as before. The nodal lines will now be formed quite differently, and will fall on the points just intermediate between those of the other surface; *i. e.* on the points of greatest excursion of its vibrating molecules. In a word, if *n, n, n, n, &c.* in fig. 129, or 130, represent the places of the nodes on the one surface, then will *n', n', &c.* be those of the other. Thus *all the motions of one half the thickness of the lamina are exactly contrary to those of the corresponding points of the other half*. This property, indeed, is general, whatever be the material, length, breadth, or thickness, of the lamina.

294. If, the other dimensions remaining, the thickness be increased, the *Sound* will remain the *same*, but the *number of nodal lines will be less*. This fact alone is sufficient to prove an essential difference between the vibrating portions of such a plate, and the ventral segments of an organ-pipe harmonically subdivided.

295. If the breadth of a plate of the above length be greater than $0^m.6$ the nodal lines cease to be straight, and ranged across the breadth at right angles to the sides. They pass into curves, and, when the breadth is increased to $0^m.04$, ($= 1^{in}.57$), they assume the forms in figs. 131, 132, the former representing the lines on the upper, the latter those on the under surface. If the breadth be enlarged to $0^m.06$, ($= 2^{in}.36$), the figures on the two faces will be as in figs. 133, 134. If the dimensions be so varied as to convert the plate into a square, the nodal figures will assume the forms in figs. 135, 136. If the form of the plate pass into the circular or triangular, the

Origin of nodal surfaces.

Of nodal lines in general.

How such nodal lines are distinguished from each other.

Motions of sand agitated by normal and by tangential vibrations.

Longitudinal-tangential vibration of rectangular rods.

Figs. 127, 128.

Different arrangement of nodes on the opposite sides. Figs. 129, 130.

Effect of a varied thickness.

Tangential vibrations of broad rectangular plates.

Sound.

same mode of vibration (longitudinal-tangential) being preserved, still the opposite sides of the plates will present different nodal figures, as in figs. 137, 138, and 139, 140.

To examine the longitudinal-tangential vibrations of cylindrical tubes or rods, as sand will not lie on their convex surfaces, M. Savart employed the ingenious artifice used by Sauveur to exhibit the harmonic nodes of a vibrating string. For this purpose the latter set astride on the string a small bit of paper cut into the form of an inverted Λ . But in this case it is found to answer better to encircle the vibrating cylinder with a narrow ring of paper, whose internal diameter is three or four times that of the cylinder, and which, therefore, hangs quite loosely on it. If a cylinder of glass about two metres ($6\frac{1}{2}$ feet) long be encircled by several such rings, or *riders*, and, being held horizontally by the middle, as lightly as possible, be rubbed in the direction of its length with a wet cloth, (it should be *very* wet,) it will yield a musical Sound, and all the riders will glide rapidly along it to their nearest nodal points on the upper surface, where they will rest. Now let all these points be marked, and then let the cylinder be turned so as to bring the opposite portion of its circumference uppermost and horizontal, and let the vibration be again excited in the same manner. Then we shall remark the very same phenomenon as in rectangular plates, *viz.* that the nodal points on this edge correspond nearly to the middles of the intervals between those of the opposite one.

If the cylinder, instead of being turned at once half round, be turned only a little at a time, and always in the same direction, the riders will come to points of rest constantly more and more towards one or the other end of the cylinder, according as it is turned to the right or to the left; and if the *locus* of all the nodal points be traced by this means, it will be found to be a species of spiral line or screw, making one or more turns round the cylinder according to its length.

But there exists here a peculiarity bearing an obvious relation to what we have observed already in the case of rectangular plates. The continuity of this spiral is interrupted near the middle of the cylinder, or rather it stops short at a point n , on one side of the central point, and recommences at N , a point equidistant on the other side; but in a *contrary* direction, so as to form on the two moieties of the length of the cylinder a right and a left-handed screw. Again, these spirals are not equally inclined to the axis in all parts of their course. They consist of portions alternately much and little inclined, having points of maximum and minimum inclination alternately at every 90° of their course round the cylinder, as in fig. 141; thus dividing the cylinder into four quadrantal portions, which are related to each other in the same manner as the upper and under faces, and the right and left sides of the vibrating parallelepipeds, examined in Arts. 292, *et seq.*

It appears then that when a cylinder is set in a state of longitudinal-tangential vibration, it assumes of itself (by reason no doubt of some casual inequality in its form or structure, giving it a bias one way or the other) four principal edges, dividing it into quadrantal portions. Of these, two opposite ones (which we will designate by the numbers 1 and 3, and call the upper and under edge) are divided by the nodal lines in points n_1, n'_1, N_1, N'_1 , and n_3, n'_3, N_3, N'_3 , where their inclination is a maximum, and the others 2, 4, which we may call the sides, at n_2, n'_2, N_2, N'_2 , and n_4, n'_4, N_4, N'_4 , where it is a minimum.

What we have said relates to the disposition of the nodal lines on the exterior surface of a tube, or of a solid cylinder. In the case of a hollow tube, the nodal lines of the internal surface may be examined by strewing in it a little fine sand, provided its diameter be so large as not to drive all the sand into a crowded line along the bottom. We shall thus detect a spiral in all respects similar to that on the external surface; only that its coils run exactly along the intervals of those of the external one. So that in all cases, those points of the internal surface are most strongly agitated by the vibration which correspond to points at rest on the outer, and *vice versa*. M. Savart has noticed a very curious phenomenon in this case. At the points of maximum inclination the sand gathers itself up in a circular heap, and remains at perfect rest; but at those of minimum inclination it forms a long ellipse, the borders of which keep constantly circulating in one direction; and if instead of sand, a small globe of ivory or wax be put into the tube; at these points it remains, it is true, without shifting its *place*, but spins constantly in one direction round a vertical axis, so long as the vibration continues.

We have all along supposed that the state of vibration into which the cylinder or tube is thrown, is that corresponding to the gravest tone it can yield by vibrations of the kind in question. M. Savart has examined its higher modes, and has pointed out other peculiarities, but for these we must refer the reader to his *Memoir, Ann. de Chim.* vol. xxv. p. 236. We will merely remark that in these modes, the threads of the screw break off, and reverse their direction at the points of union of the several ventral segments.

§ II. Of the Communication of Vibrations from one Vibrating Body to another.

We have already seen that a rod placed between two discs, one of which is set in vibration, becomes the means of communicating its vibrations to the other. But it may be announced as a general fact, that whenever a vibrating body is brought into intimate contact with another, it communicates to it its own vibrations, more or less effectually as their union is more perfect. This proposition has been carried still further by M. Savart, whose experiments show that all the particles of the body thus set in vibration by communication are agitated by motions not merely similar in their periods, but actually parallel in their directions, to those of the original source of the motion. Examples will best explain the meaning of this.

Example 1. Let A, fig. 142, be a long flat glass ruler or rod, cemented with mastic to the edge of a large bell-glass, such as is used for the *harmonica*, or musical glasses, or a large hemispherical drinking-glass, perpendicular to its circumference. Let it be very lightly supported in a horizontal position on a bit of cork at C, and then let the bell-glass be set in vibration by a bow, at a point opposite the place where the rod meets it. It will a bell-glass,

Part III.

Of still broader.
Fig. 133—140.

296.

Longitudinal-tangential vibrations of cylinders.

297.

Nodal lines spirally arranged.

298.

Two spiral nodal lines running opposite ways.

Fig. 141.

299.

Four principal edges of a vibrating cylinder

300.

Nodal lines in the interior of a cylindrical tube.

301.

Higher modes of vibration and subdivision of the cylinder into ventral segments.

302.

General law of the communication of vibratory motion.

303.

Fig. 142. Vibrations of a flat rod communicated from a bell-glass.

Sound.

vibrate *transversely*, i. e. the motions of its molecules will be perpendicular to its surface; and these motions will be communicated to the rod, without any change in their direction, whose vibrations will be longitudinal-tangential, as will be rendered evident by strewing its surface with sand, when the nodal lines will be formed as in Art 292, and, if the apparatus be inverted, and the sand strewed on the under side of the rod, the nodal lines will be seen to correspond to the points of greatest excursion on the other side, as in that article.

304. In this combination the original tone of the bell-glass is altered, and the note produced differs both from that yielded by it, or by the glass rod vibrating alone. The two vibrate as a system together and, what is singular, the sound of the glass is considerably reinforced by the combination.

Vibrate together as a system.

305. Fig. 143. Joint vibrations of two rods transverse to each other.

Example 2. Let A' be a rectangular strip of glass firmly cemented at right angles to another strip, A , across its breadth. Let the latter be lightly supported on two bits of cork, C , fastened to a wooden piece, B , so as just to touch A in the places of two of its nodes when vibrating transversely. Then, if A be placed horizontally, and strewed with sand, and A' be set in longitudinal-tangential vibration, either by rubbing with a wet cloth, or by any other means, A will vibrate transversely, as will be known by the dancing of the sand and its settling on the nodes $C C'$. On the other hand, if A be held vertically, and agitated transversely by a bow, while A' is horizontal and strewed with sand, the latter will indicate longitudinal-tangential vibrations, both by the creeping of the sand, and by the difference of the nodal figures on its two faces.

306. Fig. 144. Joint vibrations of a system of discs and rods.

Example 3. Let M be a rectangular plate (fig. 144) mounted like A in the last example, but instead of carrying a simple plate A' , let it carry a system of circular discs traversed by a lamina, as in the figure. Then, if the faces of these discs and of the lamina M be horizontally placed and strewed with sand, and the lamina M be set in longitudinal-tangential vibration, all the discs will be so too, and the sand will arrange itself in figures which, on every alternate disc, 1, 3, 5, &c. will be of one species, (such as a for instance,) but on every other, 2, 4, 6, &c. will be of a different species, as b . Now if the whole apparatus be inverted, so as to place the lamina M uppermost, and let the system of discs hang down, the *then* upper surfaces of the discs will exhibit the same system of nodal figures, but in the reverse order: i. e. the discs 1, 3, 5, &c. will give the figure b , and 2, 4, 6, &c. the figure a . In this apparatus, if the connecting piece which traverses all the discs be examined, it will be found to vibrate *transversely*, while the discs and lamina M vibrate tangentially, and *vice versâ*.

307. Fig. 145. Vibrations of a disc excited by communication from a string.

Example 4. Let A be a strong frame of wood of the form \square , across the extreme edges of which is stretched a strong ratgut or other chord, and let $L L'$ be a circular disc of glass, or metal, retained between the chord and back of the frame by the pressure of the former. Then, if the chord be set in vibration by a bow drawn transversely across it in one steady direction, the vibrations of the chord will all lie in the plane of the bow, and will be communicated in the *same* direction to the disc, which will execute tangential vibrations, each of its molecules moving to and fro in lines parallel to the bow *through the whole extent of the disc*. This is easily verified by the direction in which sand strewed on it creeps. Conceive the whole apparatus placed with the chord vertical, and projected on the plane of the horizon. If, as in fig. 145, a , FF' be the projection of the bow, the surface of the disc will be marked with nodal lines parallel to it, the sand there being left, while that in the intermediate spaces creeps along to the edges, as marked by the arrows, and runs off. If the projection of the bow FF' be oblique to the line joining the points of support of the disc, as in fig. 145, c , the nodal line will be curved, as there shown, but the motion of the molecules of sand going to form it will still be parallel to FF' . Finally, if the bow be drawn parallel to the line joining the points of support, as in fig. 145, d , the nodal line will be formed of two arcs making a cusp, but the same law of molecular motion will still hold good, as the arrows indicate.

308. Fig. 146. Passage of oblique vibrations into tangential, or into transverse. Fig. 147—150.

Example 5. Let $L L'$ be a rectangular lamina fastened at one end into a block, T , and at the other attached to a chord, $c c$, stretched parallel to its length, over a bridge, e , and put in vibration by a bow perpendicular to it, FF' . Then, if the plane of the bow and string coincide with the plane of the surface of the lamina, the latter will execute tangential vibrations *across* its breadth, and will exhibit on its upper surface a single nodal line, $n n' n''$, as in fig. 147, but on its under none, all the sand being driven off. Now incline the bow to the surface of the lamina as represented in fig. 146, c , at an angle of about 20° , still keeping it perpendicular to the string, and the nodal line will assume the curvature represented in fig. 148. If the bow be still more inclined, the curve breaks up, and at 45° of inclination becomes changed into transverse and oblique lines, as in fig. 149; and it is now observed that the sand not only runs in the direction of the arrows, but also begins to leap, indicating an oblique vibration of the surface. Lastly, when the bow is inclined 90° to the plane of the lamina, as in fig. 150, the vibration becomes altogether transverse, the nodal lines are similarly disposed on both sides of the plate, and the sand merely leaps up and down till it is danced off the vibrating parts, without any tendency to creep.

309. Vibrations of a membrane excited by communication from the air. Case 1. Transverse vibrations.

Example 6. If a very thin membrane be stretched horizontally over the orifice of a circular bowl, as a drinking-cup, or *harmonica-glass*, (extremely thin paper wetted and glued to the edges, and then suffered to become tight by drying, answers very well,) and if fine sand be strewed on it, it becomes a most delicate detector of aerial vibrations. Suppose now a circular disc of glass held concentrically over it with its plane parallel to that of the membrane, and set in transverse vibration so as to form any of Chladni's acoustic figures, as for instance fig. (99). Then will this figure be imitated exactly by the sand on the membrane. Now let the vibrating disc be shifted laterally, so as no longer to have its centre vertically over that of the membrane, but keeping its plane, as well as that of the membrane, horizontal. Still the figures marked out on the latter will be fac-similes of those on the disc, and that, whatever be the extent of lateral removal, till the vibrations become too much enfeebled by distance to have any effect at all.

310. Oblique vibrations. Tangential.

But, in place of shifting the disc laterally, let its plane be inclined to the horizon. Immediately the figures on the membrane will change though the vibrations of the disc remain unaltered, and the change will be the greater, the greater be the inclination of the plane of the disc to that of the membrane. And when the former plane is perpendicular to the horizon, the nodal figure on the membrane is found to be transformed into a system of straight lines parallel to the common intersection of the two planes, and the particles of sand

Sound. instead of dancing, creep in opposite directions to meet in these lines. One of these always passes through the centre, and the whole system is analogous to what would be produced by attaching a cord to the centre of a disc, and, having stretched it very obliquely, setting it in vibration by a bow drawn parallel to the surface. In a word, the vibrations of the membrane are now tangential, and they preserve this character unchanged, however the disc be now shifted laterally, provided its plane be not turned from the vertical position. If the disc be made to revolve about its vertical diameter, the nodal lines on the membrane will rotate, following exactly the motion of the disc. Part III.

Nothing can be more decisive or instructive than this experiment. We here see evidently, that the motions of the aerial molecules in every part of a spherical wave, propagated from a vibrating body as a centre, instead of diverging like radii in all directions so as to be always perpendicular to the surface of the wave, are all parallel to each other; in a word, they are disposed, not as in fig. 8, but as in fig. 7; and thus the hypothesis of Art. 118. is found to be completely verified. And the same thing holds good not only in air, but in liquids, as the experiments hereafter to be related (due, like all those just cited, to M. Savart) satisfactorily demonstrate. 311. Nature of the aerial motions in a Sound-wave.

This experiment is also remarkable in several other points of view. So long as the Sound of the disc, and its mode of vibration, as well as its inclination to the plane of the membrane, and the tension of the latter, continue unchanged, the nodal figure on the membrane will continue the same; but if either of these be varied, the membrane will not cease to vibrate, but the figure will be modified accordingly. Let us consider separately the effect of each of these changes. 312. Data on which the vibrations of a membrane depend.

And first, *ceteris immutatis*, let the *pitch* of the Sound whose vibrations, communicated through the air to the membrane, excite its motions, be altered, as by loading the disc, or increasing or diminishing its size, (or, if the Sound be excited by any other cause, as a pipe, the voice, &c., then by varying its pitch by any appropriate means.) The membrane will still vibrate, *differing in this respect from a rigid lamina*, which will only vibrate by sympathy with Sounds corresponding to its own subdivisions. The membrane, he it observed, will vibrate in sympathy with *any* Sound, but every particular Sound will mark out on it its own particular nodal figure, and as the pitch varies the figure varies. Thus if a slow air be played on a flute near it, each note will call up a particular form, which the next will efface, to establish its own. 313. First, the pitch of the Sound.

Secondly. Suppose the exciting cause be the vibration of a disc, or lamina of any form. If its mode of vibration be varied so as to change its nodal figures, those on the membrane will vary; and if the *same note* be produced by different subdivisions of different sized discs, the nodal figures on the membrane will be different. 314. Secondly, the nature and mode of vibration of the exciting cause.

Again, if the tension of the membrane be varied ever so little, most material changes will take place in the figures it exhibits. If paper be the substance employed, mere hygrometric changes affect it to such a degree, that if moistened by breathing on it, and allowed to dry while the exciting Sound is continued, the nodal forms will be in a constant state of fluctuation, and will not acquire permanence till the paper is so far dried as the state of the surrounding atmosphere will permit. Indeed, this fluctuation is so troublesome in experiments of this kind, that to avoid them it is necessary to coat the upper or exposed side of the paper with a thin film of varnish. Of all substances which can be employed for the exhibition of these beautiful experiments, M. Savart observes, by far the best is such a varnished paper stretched on a frame and moistened on the under side. The moisture diminishes the cohesion of the fibres, and renders them nearly independent of each other, and indifferent to all impulses. As a proof of this, he observes, that he has frequently obtained, on a circular membrane of paper so prepared, a nodal figure composed of no fewer than twenty concentric annuli, which is far beyond what can be obtained in any other way. 315. Thirdly, the tension of the membrane. Effect of hygrometric changes on paper membranes.

In some cases, a very curious and instructive phenomenon is observed in these experiments. Between the nodal lines formed by the coarser and middle-sized grains of sand, others will be occasionally observed, formed only of the very finest dust, of microscopic dimensions. This phenomenon will be seen to greater advantage if a little dust of *Lycopodon* be mixed with the sand. These intermediate lines M. Savart explains, by referring them to different and higher modes of subdivision, coexisting with that by which the principal figure is formed. The more minute particles are proportionally more resisted by the air than the coarser ones, and are thus prevented from making those great leaps which throw the coarser ones into their nodal arrangement. They, therefore, rise and fall with the surface, to which they are as it were pinned down. But they are affected by the minuter waves which have a smaller amplitude of excursion, and occur more frequently, and form their figures under the influence of these as if the greater ones did not exist. These *secondary figures* often appear as concentric rings between the primary ones, and not unfrequently the centre of the whole system is occupied by a secondary point. 316. Secondary nodal figures.

Figures 151—161 are specimens of the nodal figures thus formed on circular membranes. Of these, fig. 161 shows the modification which is apt to take place when the tension of the membrane is not quite equable. Figs. 162, 163, are figures exhibited by square membranes, and fig. 164—166 by triangular ones. 317. Fig. 151—166.

A very important application of these properties of stretched membranes has been made by M. Savart, by employing such a one as an instrument for detecting the existence and exploring the extent and limits of contiguous and oppositely vibrating portions of masses of air. For, since such a membrane is thrown into vibration by all aerial vibrations of a certain force, the fact of the existence or not of a vibratory motion in any point of the air, of a chamber for instance, or a box, or large organ-pipe, may be ascertained by observing whether sand strewn on it is set in motion, and arranged in regular forms, on holding the membrane at that point. Thus if an organ-pipe be made to sound with a constant force, and the *exploring membrane* be so far removed from it that the membrane shall just cease to be agitated visibly, the force of the Sound being increased by a quantity not sensible to the ear, the sand will recommence its motion. Nay, if two such pipes, placed close together, be made to *beat*, (see Index, *Beats*.) the membrane will be seen to be agitated at the coincidences, and at rest in the interferences of their vibrations. We shall presently return to this part of our subject. 318. Stretched membranes employed to detect sonorous vibrations.

Sound.

319.
Use of the
membrana
tympani.

Another highly interesting application of the same properties, is the view which M. Savart has taken of the use of the "membrana tympani" in the ear. Of all our organs, perhaps, the ear is one of the least understood. It is not with it as with the eye, where the known properties of light afford a complete elucidation of the whole mechanism of vision, and the use of every part of the visual apparatus. In the ear every thing is on the contrary obscure; anatomists, it is true, have scrupulously examined its construction, and many theories have been advanced of the mode in which Sounds are conveyed by it to the auditory nerve, (where of course, as with the optic nerve in the eye, all inquiry terminates, for to trace the progress of sensation along the nerve to the brain, and thence to the sentient soul, it is needless to remark, is altogether beyond our reach.) But nothing certain can be said to be known, though it is to M. Savart that we owe the most rational hypotheses hitherto proposed.

320.
Fig. 167.
Construction of the
ear.

Fig. 168.

Fig. 167 represents the auditory apparatus. It consists externally of a wide, conch-shaped opening, K L, which contracts into a narrow pipe, A B, defended from the entry of dust and insects by hairs, and a viscous exudation which is slowly secreted, and terminated by a thin elastic membrane, called the *Tympanum*, F, or drum of the ear. Behind this there is a cavity which communicates with the mouth by a small duct called the Eustachian tube, H G I. If this be stopped, deafness is said to ensue, but, as Dr. Wollaston has shown, only to Sounds within certain limits of pitch. In the cavity behind the tympanum is placed a mysterious and complicated apparatus, B C P S, represented complete, and on an enlarged scale, in fig. 168, consisting of four little bones, of which the first, S C, is called the hammer, and rests with its smaller end in contact with the tympanum, and its larger on the second bone, B P, called the *anvil*, between which, and the last, V, called the *stirrup*, a little round bone, P, forms a communication. These bones form a kind of chain, and no doubt vibrations excited in the tympanum by vibrating air, as in the experiments above detailed, are somehow or other propagated forward through these; but they are so far from being essential to hearing, that when the tympanum is destroyed, and the chain in consequence hangs loose, deafness does not follow. The last of this chain of bones, however, is attached to another membrane, p, which closes the orifice of a very extraordinary system of canals, excavated in the bony substance of the skull, called the *Labyrinth*, represented separately in fig. 169, which consists of three semicircular arcs, (1, 2, 3,) originating and terminating in a common canal, which is prolonged into a spiral cavity (4) called the cochlea. The whole cavity of the labyrinth is filled with a liquid, in which are immersed the branches of the auditory nerve, in which, no doubt, resides the immediate seat of the first impression of Sound, as that of sight does in the retina. If the membrane which closes the labyrinth be pierced, and this liquid let out, complete and irremediable deafness ensues. It appears from some most extraordinary experiments by M. Flourens on the ears of birds, (of which, however, the details are too revolting to find a place in any but works on anatomy and physiology,) that the nerves enclosed in the several canals of the labyrinth, have other uses besides their services as organs of hearing, and serve, in some unaccountable and mysterious manner, to give to animals their faculty of balancing themselves on their feet, and directing their motions. On this point we refer the reader to M. Cuvier's report on M. Flourens's Memoir, *Annales de Chimie*, vol. xxxix. p. 104, and of course to the Memoir itself, whenever and wherever it may appear; and for other not less interesting and extraordinary facts of a similar nature, to M. Majendie's Paper on the functions of the two great divisions of the spinal column, and the influence of the cerebrum and cerebellum on voluntary motion, abstracted by himself in a late volume of that collection.*

321.
Vibrations
of an un-
elastic
membrane
how pro-
duced.
Fig. 170.

To understand how the vibrations of a disc may be conceived to be communicated by the air to a membrane in M. Savart's experiments, let us take a simple case, and suppose A B C D to be a horizontal circular disc, vibrating in that mode which gives a subdivision into four quadrantal segments, A C, C B, B D, D A; and let *abcd* be an infinitely thin circular membrane placed under it, which we will suppose to be *barely coherent* so as

* From the painful subject of knowledge of the most interesting and practically useful kind, to be purchased only by the extremity of animal suffering, we turn with gladness to a pleasing duty. We have drawn largely, both in the present Essay, and in our Article on *Light*, from the *Annales de Chimie*, and we take this only opportunity distinctly to acknowledge our obligations to that most admirably conducted work. Unlike the crude and undigested scientific matter which suffices (we are ashamed to say it) for the monthly and quarterly amusement of our own countrymen, whatever is admitted into its pages has at least been taken pains with, and, with few exceptions, has sterling merit. Indeed, among the original communications which abound in it, there are few which would misbecome the first academical collections; and if any thing could diminish our regret at the long suppression of those noble Memoirs which are destined to adorn future volumes of that of the Institute, it would be the masterly abstracts of them which from time to time appear in the *Annales*, either from the hands of the authors, or from the reports rendered by the committees appointed to examine them, which latter, indeed, are universally models of their kind, and have contrived, perhaps more than any thing, to the high scientific tone of the French savans. What author indeed, but will write his best when he knows that his work, if it have merit, will immediately be reported on by a committee who will enter into all its meaning, understand it however profound, and not content with merely understanding it, pursue the trains of thought to which it leads, place its discoveries and principles in new and unexpected lights, and bring the whole of their knowledge of collateral subjects to bear upon it. Nor ought we to omit our acknowledgments to the very valuable Journals of Poggendorff and Schweigger. Less exclusively national than their Gallic compeer, they present a picture of the actual progress of Physical Science throughout Europe. Indeed, we have been often astonished to see with what celerity every thing, even moderately valuable in the scientific publications of this country, finds its way into their pages. This ought to encourage our men of science. They have a larger audience, and a wider sympathy than they are, perhaps, aware of; and however disheartening the general diffusion of smatterings of a number of subjects, and the almost equally general indifference to profound knowledge in any, among their own countrymen, may be, they may rest assured that not a fact they may discover, nor a good experiment they may make, but is instantly repeated, verified, and commented upon, in Germany, and we may add too in Italy. We wish the obligation were mutual. Here, whole branches of continental discovery are unstudied, and indeed almost unknown even by name. It is in vain to conceal the melancholy truth. We are fast dropping behind. In Mathematics we have long since drawn the rein and given over a hopeless race. In Chemistry the case is not much better. Who can tell us any thing of the Sulfo-salts? Who will explain to us the laws of Isomorphism? Nay, who among us has even verified Thenard's experiments on the oxygenated Acids—Oersted's and Berzelius's on the radicals of the Earths—Balard's and Serrulas's on the combinations of Brome—and a hundred other splendid trains of research in that fascinating science? Nor need we stop here. There are, indeed, few sciences which would not furnish matter for similar remark. The causes are at once obvious and deep seated. But this is not the place to discuss them.

Sot ad.

Part III.

to be impervious to air, but to have *no tension* of its own. Its molecules will, therefore, obey implicitly all the motions of the aerial ones adjacent to them, and its figure, at any instant, will be that assumed by a stratum of the air originally plane, and parallel to $A B C D$, in consequence of the displacement of its particles by the undulation propagated from all parts of $A B C D$, as they reach it *at once*, allowing for the time taken to traverse their respective distances from it. Let us now consider how a molecule of air, M , placed any where in the plane $a e \beta$ perpendicular to the disc, and intersecting it in $A B$, will be affected. Since the disc vibrates transversely, all its particles on one side of $A B$ (as towards C) will be at any instant in a precisely opposite *phase* of their excursion from the corresponding particle on the side towards D , and moving with equal velocity. Therefore, the undulations propagated simultaneously from both these particles will reach the molecule M in question at once, (being equidistant from it,) and being (at least in so far as their direction is not modified in their passage, and at all events as to that part of them which is at right angles to the plane $a \beta \gamma \delta$) equal, and contrary, destroy each other, so that in virtue of these the molecule M acquires no transversal vibration. And since the same is true of every other corresponding pair of molecules into which the two halves $A C B$ and $A D B$ of the vibrating disc can be divided, the molecule M will not vibrate (or at least not transversely) in virtue of the vibration of the whole disc. The same is true of every other molecule situated in the plane $a e \beta$, and also by a similar reasoning in the plane $\gamma e \delta$ at right angles to it. There will then be two *nodal planes* pervading the whole atmosphere, in which the aerial molecules have no transverse (*i. e.* vertical) motion. But if we suppose the molecule M situated any where out of these planes the case is otherwise. Suppose it, for instance, situated at f , in the quadrant $c e b$ of the membrane. This being nearer to each molecule of the quadrant $C E B$ of the disc than to the corresponding molecules of the others, the influence of the former will predominate, and the molecule f will be agitated by a transverse motion similar to that of the quadrant of the disc vertically over it. If then the membrane be strewed with sand, it will be thrown off from the vibrating quadrants, and arranged on two rectangular nodal lines $a b, c d$ parallel to those of the disc, just as if it vibrated by its own tension, while yet it is obvious that all the while it has only obeyed implicitly the motions of the adjacent air.

If, however, the membrane has tension and thickness, this will modify the effects of the direct aerial action, and that in a way far too complicated for us to enter into here in detail. We may remark, however, that in that case, each individual aerial impulse must be regarded as an arbitrary initial disturbance of its state of equilibrium, in virtue of which it will be thrown into periodic vibrations; and these again will propagate similar vibrations back through the air to the disc $A B C D$; and this being constantly repeated the result *may* be the establishment of a joint resultant periodic vibration, by the destruction of every motion not periodic, from the innumerable repetitions of the impulses and the consequent infinite superposition of plus and minus excursions. But this interchange, of course, will be the more energetic the thinner is the interposed lamina of air; for if its thickness be great, the vibrations excited in the membrane, or semi-rigid disc, $a b c d$, (as we will now suppose it,) will be feeble, and when propagated back through the air will be still further enfeebled so as to affect the motion of $A B C D$ but little. In this case then, supposing the two discs to be out of unison with each other, and to have no common mode of vibration, the disc $a b c d$ will become the seat of two distinct systems of vibration. The first, regularly periodical, being that directly communicated by sympathy. The other, the resultant of an indefinite number of vibrations kept up by means of the tension, in all phases and stages of degradation.

Now, provided the time elapsed since the commencement of the vibrations be long enough to allow of our regarding the number of previous vibrations as infinite, or which comes to the same, long enough to have allowed all traces of the initial vibrations to have been destroyed by resistance, friction, &c., these last will either exactly destroy each other, or, if they leave a residue, that residue will consist in a vibratory motion, having the same period with the primary impulse.

As this is a proposition of great importance, not only in the theory of Sound, but in many other physical theories, such as that of the Tides, for example, we must not let it rest on a vague assumption, but demonstrate it rigorously. Let then t represent the time elapsed since the commencement of the vibrations, t being so large that it may be considered as infinite in comparison of the duration of a single vibration. Then if we call T the time of one complete vibration, or one period of the primary vibrations, the impulse communicated through the air, or otherwise, to any point of the membrane, or other vibrating body, will at any instant be represented by

some periodic function of the form $F\left(\cos 2\pi \cdot \frac{t}{T}\right)$, or $F(\cos n t)$ putting $\frac{2\pi}{T} = n$, which function may always

be resolved into a series of periodical terms of the form $A \cdot \cos i n t$, i being an integer, of which each may be considered as the representative of a single vibratory motion of the simplest kind, whose superposition forms the actual vibration in question. Consequently, we may content ourselves with considering any one of them as $A \cdot \cos n t$, since all the rest are subject to the same argument.

Next, let θ be the time of one complete unforced vibration of the membrane or elastic body in virtue of its natural elasticity, and let $\nu = \frac{2\pi}{\theta} \cdot i$, so that $a \cdot \cos \nu t$ would denote the general term of a series expressing the

velocity of any one of its molecules in a state of unforced vibration, and let $F(t)$ be a function expressive of the law of diminution of the vibrating motion by friction, resistance, and imperfect elasticity. So that if t be the time since a certain velocity V was communicated to it, $V \cdot F(t) \cdot \cos \nu t$ will be its velocity after the expiration of t as it will then subsist, modified by the elastic forces and mechanical state of the membrane.

Conceive the aerial impulse to act not continuously, but at equal infinitely small intervals of time τ , (infinitely small relative not only to t but to T and θ .) Then, first, the impulse $A \cdot \cos n t$, acting during the time τ , will produce the velocity $A \cdot \tau \cdot \cos n t$.

322.

Effects of tension and thickness of the membrane.

323.

General theorem respecting forced vibrations.

324.

Demonstration.

325.

326

Sound. Secondly. The impulse $A \cdot \cos(n t - n \tau)$ which acted at the moment immediately preceding, produced in the first instance the velocity $A \cdot \tau \cdot \cos(n t - n \tau)$. But this, once produced, was immediately modified by the inherent elasticity of the membrane, and in the subsequent moment became

$$A \cdot \tau \cdot \cos n(t - \tau) \cdot F(\tau) \cdot \cos \nu \tau.$$

Similarly the impulse $A \cdot \cos n(t - 2\tau)$ acting at the instant preceding this generated the velocity $A \tau \cdot \cos(t - 2\tau)$, which, in like manner, (being regarded as an arbitrary initial disturbance,) became modified in the time 2τ to $A \tau \cdot \cos n(t - 2\tau) \cdot F(2\tau) \cdot \cos 2\nu \tau$. And so on. Thus, the whole accumulated velocity at the instant t , arising from all the preceding impulses, will be expressed by

$$A \tau \cdot \{ \cos n t + \cos n(t - \tau) \cdot \cos \nu \tau \cdot F(\tau) + \cos n(t - 2\tau) \cdot \cos 2\nu \tau \cdot F(2\tau) + \&c. \},$$

which series, since the function expressed by $F(t)$ is supposed to decrease constantly as t increases, and since the whole number of vibrations is supposed so great that the terms of the series $F(t)$, $F(2\tau)$, $F(3\tau)$, &c. shall at length become perfectly insensible, may be regarded as continued *ad infinitum*.

In fact, whatever supposition we may make as to the law of degradation of the motion within the limit of a single period, it must evidently diminish in geometrical progression in similar phases of successive periods, so that we must have

$$F(\tau + \theta) = q \cdot F(\tau); F(\tau + 2\theta) = q^2 \cdot F(\tau) \&c.$$

This premised, the series in question, by merely changing the arrangement of its terms, and grouping together those equidistant from each other by the interval θ , will become resolved into partial series thus,

$$A \tau \cdot \left\{ \begin{array}{l} \cos n t + \cos n(t - \theta) \cdot \cos \nu \theta \cdot F(\theta) + \cos n(t - 2\theta) \cdot \cos 2\nu \theta \cdot F(2\theta) + \&c. \\ + \cos n(t - \tau) \cdot F(\tau) + \cos n(t - \tau - \theta) \cdot \cos(\nu \theta + \nu \tau) \cdot F(\theta + \tau) + \&c. \\ + \&c. \end{array} \right\}$$

But we have, first, $F(\theta) = 1$, $F(\theta) = q$, $F(2\theta) = q^2$, &c.; and, moreover, since $\nu \theta = 2i\pi$, therefore $\cos \nu \theta = 1$, and $\cos 2\nu \theta = 1$, &c. Consequently the above expression becomes

$$A \tau \cdot \left\{ \begin{array}{l} (\cos n t + q \cdot \cos n(t - \theta) + q^2 \cdot \cos n(t - 2\theta) + \&c.) \\ + F(\tau) \cdot (\cos n(t - \tau) + q \cdot \cos n(t - \tau - \theta) + q^2 \cdot \&c.) \\ + F(2\tau) \cdot (\cos n(t - 2\tau) + q \cdot \cos n(t - 2\tau - \theta) + q^2 \cdot \&c. + \&c.) \end{array} \right\}.$$

Summation
of the par-
tial series.

Now, each of these series is readily summed, for we have by well-known trigonometrical formulae

$$\begin{aligned} & \cos n t + q \cdot \cos(n t - n \theta) + q^2 \cdot \cos(n t - 2n \theta) + \&c. \\ &= \cos n t \cdot \frac{1 - q \cdot \cos n \theta}{1 - 2q \cdot \cos n \theta + q^2} + \sin n t \cdot \frac{q \cdot \sin n \theta}{1 - 2q \cdot \cos n \theta + q^2}. \end{aligned}$$

Each of the fractions being constant, and independent of t , if we call them M and N , our series will become

$$A \tau \cdot \left\{ \begin{array}{l} (M \cdot \cos n t + N \cdot \sin n t) + F(\tau) \{ M \cdot \cos n(t - \tau) + N \cdot \sin n(t - \tau) \} \\ + F(2\tau) \{ M \cdot \cos n(t - 2\tau) + N \cdot \sin n(t - 2\tau) \} \\ + \&c. \end{array} \right\}.$$

328.
Summation
of the whole
by a definite
integral.

Let us now consider the area of a curve whose abscissa, x , is divided into equal elements each equal to τ , while its successive ordinates, y , are represented by $\phi(0)$, $\phi(\tau)$, $\phi(2\tau)$, &c. It is evident that its area $\int y dx$ will be equal to

$$\tau \cdot \phi(0) + \tau \cdot \phi(\tau) + \tau \cdot \phi(2\tau) + \&c.;$$

and, therefore, the sum of this series, from the term $\tau \cdot \phi(0)$ to $\tau \cdot \phi(\theta)$, will be equal to the integral $\int \phi(x) dx$, from $x = 0$ to $x = \theta$. Thus our series will assume the form of a definite integral, viz.

$$A \cdot \int_0^\theta dx \cdot F(x) \{ M \cdot \cos n(t - x) + N \cdot \sin n(t - x) \},$$

expressing in the manner now pretty general the limits of the integral by indices attached to the integral sign. Resolving now the sines and cosines of $n(t - x)$, this becomes (x and t being independent of each other)

$$\begin{aligned} & A \cdot \cos n t \cdot \int_0^\theta dx \cdot F(x) \{ M \cdot \cos n x - N \cdot \sin n x \} \\ & - A \cdot \sin n t \cdot \int_0^\theta dx \cdot F(x) \{ M \cdot \sin n x - N \cdot \cos n x \}. \end{aligned}$$

Now, whatever be the law of degradation denoted by the function F , it is clear that these definite integrals must at last reduce themselves to certain constants independent of t , which, if we call P and Q , the whole takes the simple form

$$P \cdot \cos n t - Q \cdot \sin n t,$$

which is a periodic function having the same period as the primary vibrations.*

329.
Case of per-
fect elasti-
city of the
excited
body.

In the limiting case, when the elasticity of the body on which the forced vibrations are impressed is perfect, and resistance, friction, and every other cause of loss of motion is prevented, F represents a constant, and is equal to unity. In this case both the constants P and Q in the above expressions vanish, and the whole motion

* This demonstration being general, we may here observe, that, on the undulatory theory of light, rays of one refrangibility can never excite by any combination of their own vibrations with those of the bodies they may traverse or impinge on, any resultant rays of a different refrangibility, at least so long as the exciting light continues in action. When it has ceased, the case may be otherwise.

Sound. of the body, after a great number of vibrations have elapsed, is zero. In this case then, the elastic body is completely incapable of vibrating in sympathy with any other not having a common mode. In all others P and Q have finite values, which will be greater, or less, according to the circumstances of the case.

Part III.

Thus we see that imperfect elasticity, or other equivalent causes of the gradual loss and dissipation of the impressed impulses, is the essential condition on which forced vibrations in general depend, and that in proportion as a disc or membrane is devoid of tension it should be more readily susceptible of such vibrations: precisely what M. Savart has shown to be really the case in fact.

330.

General effect of imperfect elasticity.

It may be objected to what is said in Art. 321, that it would follow from that reasoning that the Sound of a vibrating disc should be inaudible whenever the ear is situated in a plane passing through one of its nodal diameters, and at right angles to the disc. But, in the first place, what is there said applies only to such motions of the aerial particles as are performed in those planes. But, in fact, a lateral motion, or one parallel to the disc's surface, must also exist, by reason of the alternate *tilting* up and down of adjacent ventral segments, which must give the whole body of air terminated by them a small reciprocating rotatory motion about the nodal line separating them as an axis. Thus, though the transverse vibrations are here destroyed, the sensation of Sound may still be excited by tangential ones. And, secondly, though alternate motion were altogether destroyed, condensations and rarefactions still subsist.

331.

Objections considered.

But, in fact, there is observed a difference in the intensity of Sound emanating from vibrating bodies in certain cases, according to their angular position with respect to the line joining them and the ear. We have already (Art. 117.) described Dr. Young's remarkable experiment of the tuning-fork. It is precisely a case in point, and a circumstantial explanation of it will be at once interesting for its own sake, and illustrative of the general argument. Let then A, B, fig. 171, be sections of the two branches of the fork in its state of rest, and since when set in vibration they alternately approach to and recede from each other, let us consider them first in their state of approach, as at *a b*. In this state they compress the air between them, and squeeze it out laterally in the direction of the arrows P, Q, while, at the same instant, the aerial particles adjacent to the flat outward faces of the two branches, and which of necessity follow their motions, are urged inwards as indicated by the arrows R S. Thus the four quadrants of the initial circular wave propagated round the fork, are alternately in opposite states of motion, the molecules at P Q receding from the centre, while those at R, S, approach to it, and *vice versa*, when the branches of the fork having closed to the utmost begin to open again. In this case the lateral air will rush in to fill the gap, while that in contact with the broad faces will be forced outwards. If then we consider any intermediate point C, about 45° distant from Q and R, this, in virtue of both impulses, will acquire equal tendencies in opposite directions, and will rest, or at least will acquire only a small tangential motion, in consequence of the reciprocating eddies of the air round the angles of the branches. That these motions really do take place as here pointed out, any one may have ocular demonstration by imitating the opening and shutting of the branches with his hands near the flame of a candle burning steadily, taking care not to make *puffs of wind* but regular removals of the air to and fro.

332.

Phenomena of a tuning-fork explained. Fig. 171.

One of the most curious and interesting purposes to which M. Savart has applied the properties of membranes, is to explore the actual state of the air in different parts of a vibrating mass of determinate figure, as to motion or rest. For this purpose, the Sound should be excited and maintained by a constant cause at a high degree of intensity, especially if the mass of air be large, as in a chamber or gallery; and to give the membrane the greatest possible sensibility, it ought to be stretched so as to be, naturally, in unison with the note sounded, so as to act as a receiver and condenser of the small aerial motions. The greatest purity and intensity of the Sounds to be employed for this purpose, may be obtained by a *harmonica* glass, or the bell of a clock, maintained in vibration by a bow; and this may be still further augmented by adapting to it a resonant cavity, as, for instance, a large cylindrical vase of considerable diameter, closed at one end, and of such dimensions as separately to vibrate the same note. (See Art. 338.) The tones thus produced, especially when large *harmonica* glasses are used, as M. Savart remarks, are of such intensity, that no ear can long support them, and, at the same time, of such a rich and mellow quality, that all other musical Sounds appear poor and harsh in comparison. In order yet more to increase the sensibility of the membrane, the frame on which it is stretched should be fitted over the orifice of a similar resonant cavity. For convenience, and lest the tension of the membrane should vary by hygrometric changes, it is proper to have means of varying this at pleasure, a mode of which is described by M. Savart in the Memoir from which we draw our information. (*Annales de Chimie*, vol. xxiv. p. 76.)

333.

Membranes used to explore the vibrations of masses of air. Intense Sounds produced by resonance.

Mode of using a membrane.

Suppose now that, being provided with such an apparatus as here described, we shut ourselves up in an apartment of regular figure, and free from furniture or projections from the walls, recesses, &c., and place one of our resonant cylinders with its axis horizontal, and the vibrating bell or glass opposite its orifice. In the direction of its axis place the membrane horizontally, with its proper frame and resonant cylinder below it, and strew the horizontal surface with sand. If now, first, we place the membrane thus armed very near the source of the Sound, it will vibrate with great force. As we withdraw it, (keeping it still in the line of the axis of the first resonant cylinder,) its vibrations will diminish gradually, and at length cease, after which (still continuing to remove it along this line) they will recommence and reach a maximum, at a point when their intensity is nearly equal to that close to the source of Sound. Removing the membrane yet further, a new point of indifference is found, and so on till we reach the end of the chamber. If we walk along the same line, keeping the ear in the plane of the horizontal axis of the resonant cylinder, we shall perceive the Sounds to be much louder in the places where the vibrations of the membrane attain their maxima, than at the intermediate points where they are at a minimum. At these latter, a very curious phenomenon has been observed by M. Savart. When the auditor moves his head away from such a point, towards the *right*, (always supposing it to remain in the line of the axis above mentioned,) the Sound will appear to come *from the right*, and if towards the left, it will seem to come

334.

Vibrations of the air in a chamber examined.

Sound. from the left, whether the original source of Sound be to the one or the other side. This singular effect shows that the aerial molecules on either side of the point of indifference, are in opposite states of motion at any given instant. In making this experiment, the head should be so turned, that the axis of the resonant cylinder prolonged shall pass through both ears. Suppose, for instance, the Sounding apparatus to be to the observer's left, and that his head be very near it. The Sound will appear to enter at his left ear. As he removes further away, so as to pass one of the nodes, it will seem as if the Sound had changed sides, and now came from the right. When another node is passed, it will appear to have again shifted to the left, and so on.

335. But, if we quit the axis of the cylinder, and carry an exploring membrane, such as already described, about the apartment, noting all the points where it vibrates most forcibly, allowing ourselves, as it were, to be led from spot to spot by its indications, we shall trace out in the air of the room a curve of double curvature marking the maxima of the excursions of the aerial molecules. If the experiment be made in a gallery, or passage, whose length is its principal dimension, this curve will be found to be a kind of spiral, creeping round the walls, floor, and ceiling, obliquely to the axis of the gallery, thus presenting a marked analogy to the disposition of the nodal lines in a long rod vibrating tangentially, (*vide* Art. 313.); which are also, it should be remarked, imitated, with modifications more or less complicated, in square or rectangular rods.

336. A still more remarkable effect was observed by M. Savart, in thus exploring the vibrations of the air in an apartment with an open window. The spiral disposition of the vibrating portions was found to be continued out of the window into the open air, the lines of greatest intensity running out in great convolutions which seemed to grow wider, on receding from the window, and could be traced to a great distance from it.

The vibrations of the air in an organ-pipe were explored by M. Savart, by lowering into the pipe, placed vertically with its upper end open, a thin membrane stretched on a light ring, and suspended by a fine silk thread, and strewed with sand. Thus ocular demonstration of the existence of its subdivision into distinct ventral segments was obtained, the sand remaining undisturbed when the membrane occupied precisely the place of a node. By this means, too, the influence of the embouchure on the places of the nodes, a curious and delicate point in the theory of pipes, which we have not before alluded to, may be subjected to exact examination. Thus, for instance, when the column of air in the pipe vibrates in the manner described, Art. 190, fig. 19. having two half ventral segments, and one node in the middle, it is found that the node is only approximately so placed, being always, in fact, nearer to the embouchure than to the open end.

338. It is well known that if we sing near the aperture of a wide-mouthed vessel, some one note (which is in unison with the air in the vessel) will be reinforced and augmented, and sometimes to a great degree. This is what is meant by the *resonance* of the mass of air contained in the cavity of the vessel, or as it may be termed, the resonance of the cavity. This has been known from the earliest times. The ancients are said to have placed large brass jars under the seats of their immense theatres to reinforce (one does not well see how) the voices of the actors. Any vessel or cavity may be made to resound by placing opposite its orifice a vibrating body, having a surface large enough to cover the aperture, or at least to set a considerable portion of the aerial stratum adjacent to it in regular oscillation, and, at the same time, pitched in unison with the note which the cavity would of itself yield. The experiment of the disked tuning-fork, in Art. 204, is a case exactly in point. The pipe which resounds in that experiment, may be pitched precisely in unison with it by its stopper, and in proportion as it departs from a perfect unison the resonance is feebler. A series of disked tuning-forks, or vibrating steel springs, thus placed over the orifices of pipes carefully tuned, constitutes a very pretty musical instrument, capable of a fine swell and fall according as the discs are brought nearer to, or further from, the orifices of the pipes, or inclined to their axes, and of remarkable purity and sweetness of tone. A similar adaptation of resonant cavities to a series of harmonica glasses fixed on a common revolving axis, has been recommended by M. Savart as the principle of a musical instrument, whose effect, should it be found to answer the expectations his description of the tones thus drawn forth is calculated to excite, would probably surpass that of all others yet invented. See Art. 333. The cavities best adapted to this purpose are short cylinders of large diameters with movable bottoms fitting by tight friction by which they may be tuned.

339. Such cavities may be regarded as short organ-pipes. When the diameter of a pipe is greatly increased in proportion to its length, so that it becomes a box, the law of the proportionality of the time of vibration to the length ceases to hold good, and the note yielded is flatter than that of a narrow pipe of equal length, and the more so the wider the pipe. Thus M. Savart found that a cylinder of $4\frac{1}{2}$ inches in length, and 5 in diameter, resounded in unison with a narrow pipe 6 inches long, making 1024 vibrations per second. That sagacious experimenter has found, that cubical boxes *speak* with surprising promptitude and facility, and yield Sounds extremely pure, and of a peculiar quality, on which account, and by reason of the little height in which they may be *packed*, he recommends them for organ-pipes. A cube of 53 or 54 lines ($= 4\frac{1}{2}$ in.) in the side yields the same note as a pipe 10 or 11 inches long, and 2 or $2\frac{1}{2}$ inches diameter. They may be excited by an embouchure at one of their lower edges, precisely similar to that of an organ-pipe. But they will also speak if the embouchure be situated in the middle of the side. M. Savart has also examined the vibrations of a great variety of different-shaped pipes, boxes, or cavities, for which see *Annales de Chimie*, vol. xxix. p. 404.

340. There is yet another remarkable case of vibrations communicated between the different members of a system of which we have not yet spoken, though offering a good example of the verification of the general law of equality of period and parallelism of direction of the vibratory motions of all the molecules of a system laid down in Art. 302. It is when vibrations are communicated through a liquid. The following experiments of M. Savart will show the mode in which this is accomplished.

341. He took a cylindrical tinned iron vessel whose bottom was placed parallel to the horizon, and having cemented to its centre a glass rod, so as to hang perpendicularly down from it, he covered the bottom to the depth of about an inch and a half with water, on which was floated a thin disc of varnished wood, covered on its upper

Spiral form of the nodal lines in a rectangular chamber, or gallery.

Their continuation through a window into the air.

Vibrations of air in pipes explored by membranes.

Resonance of cavities.

Resonance and vibrations of box-shaped cavities. Resonance of cubical boxes.

Communication of vibrations through liquids.

Experiment.

Sound.

face with sand. The apparatus thus prepared, he impressed on the glass rod a longitudino-tangential vibration, (Art. 296.) which of course became normal when communicated to the bottom of the vessel, and observed the sand on the upper face of the disc to be also agitated with normal motions, and to assume nodal figures according to the laws of that species of motion. To show more clearly the nature of the communication, he threw out the water, and supported the wooden disc by a small solid stem perpendicular to its surface, and the bottom of the vessel, and attached to the centres of both, when it was found that the disc was affected precisely in the same way as before.

Part III.

Communi-
cation of
normal vi-
brations.

On a vessel of water, whose rim is maintained in a state of normal vibration by a bow drawn perpendicularly across it at any point, let a thin rectangular lamina of wood be floated, having its length parallel to the bow, and its extremity opposite to the point of the circumference excited. The lamina will be seen (as usual by sand strewed on its upper face) to execute longitudino-tangential vibrations, and will be crossed by nodal lines at right angles to its length. But if, instead of directing the axis or longer edges of the lamina perpendicularly towards the vibrating point of the side of the vessel, we incline it obliquely to the direction of the vibrations, still the sand on its upper face will continue to glide in the same direction as before, that is, parallel to the vibrations of the side of the vessel, so that, if the floating lamina be made to revolve slowly in a horizontal plane, the direction of the creeping motion of the sand on its surface will continually vary with respect to the position of its edges, though constant with regard to the sides of the vessel.

342.

Communi-
cation of
tangential
vibrations.

Not only are the vibrations thus faithfully transferred through the water to bodies floating on its surface, but even to such as are totally immersed in it. The experiment is easily made by suspending in such a vessel as above described under the water, and not in contact with the sides or bottom, a disc of glass, by means of fine silk threads, and strewing sand on the surface of the water which sinks and spreads evenly on the disc. This will be observed to be agitated with very decided normal, or tangential motions, according as the former or latter of the modes of excitement used in the experiments, Arts. 341, 342, is employed; and to arrange itself in nodal figures accordingly.

343.

Communi-
cation of vi-
brations to
a body en-
tirely im-
mersed.

From these and similar experiments it appears that vibratory motions are communicated through liquids precisely as through gases and solids, without change of character or *direction*. This, observes M. Savart, explains how the nerves of hearing, extended throughout the convolutions of the labyrinth and immersed in the liquid which fills it, transmit to the sensorium, not only the general impression of Sound, but of the direction in which it comes.

344.

How we
judge of the
direction of
Sound.

These remarkable and striking results all tend to confirm and strengthen the analogy between Sound and Light. The luminiferous ether, like air and liquids, transmits vibrations without altering their direction, as the phenomena of polarized light demonstrate. The additional weight of evidence thus thrown into the scale of the undulatory theory of light did not escape the penetrating mind of Dr. Young, to whom that theory was so deeply indebted. Doubtless the analogy thus ascertained would not have remained idle in his hands, had not death snatched him too from science while in the vigour of his intellect, and when so much might have yet been hoped from him. It has been our unprecedently unfortunate lot, while composing these Essays on the sister sciences of Light and Sound, to have to deplore the loss of nearly all the great modern contributors to their advancement. A Fraunhofer, a Fresnel, a Wollaston, and a Young, names forming an epoch in the history of human knowledge, have been snatched away in quick and alarming succession, not enfeebled by age or with faculties weakened by disease, but all in the meridian of their intellectual powers, or in that rich maturity when practice had only familiarized them with their resources, and perfected them in their use. To Dr. Young the theory of Sound is in many respects deeply indebted, and it richly repaid the attention he devoted to it by furnishing him with the pregnant idea of his principle of interferences; a principle which has proved the key of all the more abstruse and puzzling properties of light, and whose establishment would alone have sufficed to place him in the highest rank of scientific immortality, even were his other almost innumerable claims to such distinction disregarded.

345.

Further
analogies
between
Sound and
Light.
Of Light and
Sound

§ III. Of the Voice.

Almost every animal has a voice or cry peculiar to itself, originating in an apparatus destined for that purpose of more or less complexity. The voice is most perfect and varied in man and in birds, which, however, differ extremely in the degree in which they possess this important gift. In quadrupeds, it is limited to a few uncouth screams, bellowings, and other noises, perfectly unmusical in their character, while in many birds it assumes the form of musical notes of great richness and power, or even of articulate speech. In the human species alone, and that only in some rare instances, we find the power of imitating with the voice every imaginable kind of noise, with a perfect resemblance, and of uttering musical tones of a sweetness and delicacy attainable by no instrument. But in all, without exception, (unless, perhaps, the chirp of the grasshopper, or cricket, be one,) the Sounds of the voice are produced by a *wind instrument*, by the column of air contained in the mouth, throat, and anterior part of the windpipe, set in vibration by the issue of a stream of air from the lungs through a membranous slit in a kind of valve placed in the throat. In man and in quadrupeds, this organ is single, but in birds, as M. Savart has shown, it is double; a valve of the kind abovementioned being placed at the opening of each of the two great branches into which the trachea first divides itself as it enters the lungs, just before they unite into one common windpipe.

346.

The organs of the voice, in man, consist of

1. The thorax, which, by the aid of the diaphragm and the 24 intercostal muscles acting on the lungs within, and alternately compressing and dilating them, performs the office of a bellows

347.

The thorax.

- Sound. 2. The trachea, a cartilaginous and elastic pipe which terminates in the lungs by an infinity of roots, or bronchiæ, and whose upper extremity is formed into a species of head called the larynx situated in the throat, composed of five elastic cartilages, of which the uppermost is called the epiglottis, whose office is to open and shut, like a valve, the aperture of the exterior glottis, and which constitutes the orifice of the larynx
348. The trachea and larynx. 3. The epiglottis, where it adheres to the larynx, is also united to the tongue, and forms a somewhat concave valve, of a parabolic form, whose base is towards the tongue, and which, by its convexity, resists the pressure of the food and liquids as they pass over it in the act of swallowing.
349. The epiglottis. 350. 4. Within the larynx, rather above its middle, between the thyroïd and arytenoid cartilages, are two elastic ligaments like the parchment of a drum slit in the middle, and forming an aperture making a right angle with the exterior glottis, and which is called the interior, or true glottis. This slit, in adults, is about four-fifths of an inch long, and a twelfth of an inch broad. This aperture is provided with muscles which enlarge and contract it at pleasure, and otherwise modify the form of the larynx.
351. The mouth, uvula, &c. 5. The tongue, the cavity of the fauces, the lips, teeth, and palate, with its *velum pendulum*, and the uvula, a pendulous, conical, muscular body, which performs the office of a valve between the throat and nostrils, as well as, perhaps, the cavity of the nostrils themselves, are all concerned in modifying the impulse given to the breath as it issues from the larynx, and producing the various consonants and vowels, according to the different capacities and shapes of their internal cavity.
352. Glottis supposed to act as a reed. In speaking or singing, the glottis, it has been generally supposed, performs the part of a reed. The membranes of which it is composed being kept at a greater or less state of tension by the muscles with which it is provided, and its opening expanded or contracted according to the degree of gravity or acuteness of the Sound to be uttered. But the tone thus originally produced by the glottis is sustained and reinforced by the column of air in the larynx, throat, and mouth, whose dimensions and figure are susceptible of great variation by the action of the innumerable muscles which give motion to this complicated and intricate part of our frame. Thus in a general way we may conceive how the voice is produced and modified; but when we would penetrate further into particulars, the difficulties presented by the organs of the voice are even greater than those which beset the investigation of those of hearing.
353. Difficulty from the gravity of the notes produced. One material one has been lately much elucidated by the experimental researches of M. Savart. How, we may naturally ask, can tones of such gravity as we hear produced by the human voice, be excited in so short a column of air as that contained in the throat of a man? The vibrating column here hardly exceeds a few inches in length, yet the notes produced by a bass singer are those which would require a pipe of several feet in length sounded in the usual manner. That it is not a mere relaxation of the membrane of the glottis is evident; the dropping of the lower jaw, and the effort made in every possible way to increase the dimensions and diminish the tension of the throat and fauces generally, in singing the lower notes of the scale, sufficiently prove that the note of the glottis is reinforced in this case, as in that of acuter Sounds, by the resonance of the cavity in which it sounds.
354. Explained by M. Savart. From M. Savart's experiments it appears that in short pipes, and cavities whose other dimensions bear a considerable ratio to their length, the tone yielded is rendered much graver when the pipe or cavity is constructed of a flexible material capable of being agitated and set in vibration by the air, than when made of more rigid materials. He constructed a cubic box-pipe with paper stretched on slight square frames of wood, joined together at the edges, and made it speak by an embouchure at the edge. He then observed, that so long as the paper was tightly stretched the Sound yielded by the cube was nearly as acute as it would have been had the whole been rigid, but that when its tension was diminished by exposing it to moist vapour, or even by wetting it, the Sound descended in the scale by an interval proportioned to the degree of moisture the paper had imbibed. It was thus lowered even two whole octaves, when it grew so feeble as to be no longer audible; but, repeating the experiment in the still of night, it could yet be heard, and no limit indeed then seemed set to the descent of the Sound; and even when no longer audible the vibration of the paper sides could still be made sensible by sand strewed on them, which arranged itself in nodal lines, for the most part elliptic or circular.
355. Tension and relaxation of the uvula and mouth. The relaxation then, or increase of tension of the soft parts which form the cavity of the mouth and larynx, is no doubt a principal cause of the graduation of its tones. Whoever will sing open-mouthed before a looking-glass will not fail to be struck with the extraordinary contraction of the uvula (a small pendulous substance which seems to hang down from the roof of the mouth) which takes place in the higher notes. It shrinks up almost into a point, and every surrounding part seems to partake its tension.
356. Savart's objections against the reed-like action of the glottis. We have observed that the glottis has been most generally regarded as performing the functions of a reed, especially since the free reed (*anche libre*) invented by Kratzenstein, and revived by Grenié, (probably without knowledge of Kratzenstein's prior invention; *vide Willis, Phil. Trans. Camb. vol. iii.*) has been brought into general notice. This idea is strongly advocated among others by Biot. But M. Savart professes himself dissatisfied with such an explanation of its use. He remarks, and seemingly with justice, that the essential principle of a reed, the periodical opening and closing of the orifice through which the stream of air passes, is wanting in the glottis. Were the glottis a reed, the edges of the vocal ligaments which form the slit through which the air passes would require to be almost in contact, and should be alternately forced asunder by the effort of the air, and brought together by their tension. But on the contrary he found that the larynx of the dead subject, when left in its natural state, and gently blown into through the trachea, yielded Sounds approaching to those of the voice, although the opening left between the borders of the glottis was as much as one-sixth, or even one-fourth of an inch across, and more than half an inch long.
357. M. Savart's explanation of the voice. The instrument to which M. Savart attributes the greatest analogy to the larynx, is a species of whistle, common enough as a children's toy or even as a sportsman's call, in the form of a hollow cylinder about three-fourths of an inch in diameter, closed at both ends by flat, circular plates, having holes in their centres. The form is not of

Sound.

much importance, it may be made hemispherical, &c. Being held between the teeth and lips, the air is blown through it, and Sounds are produced which vary in pitch with the force of the blast. If the air be conducted to it through a *porte vent*, and cautiously graduated, all the Sounds within the compass of a double octave may readily be obtained from it; and if great precautions are taken in the management of the wind, tones even yet graver may be educed, so as to admit, in fact, no limit in this direction.

When we come to investigate the nature of articulate Sounds, and of speech, the difficulties are much greater. Conrad Amman, in his work on the Voice, first attempted to explain the manner in which the vowels and consonants are formed. With regard to the vowels, he regards them as mere modifications of the continued tone produced by the larynx, depending on the configurations of the mouth. Thus to pronounce A (the broad A in Ah!) the tongue must be laid flat in the lower jaw, and the mouth opened wide, and lips turned outwards. Any musical or continued tone produced in the throat will then have the character of the vowel A. If the tongue be gradually elevated so as to bring its middle nearer the palate, and at the same time thrust forwards, its extremity approaching the upper teeth, the Sound will deviate from the broad A into *a* (hate,) *e* (peep.) These Sounds therefore (the *a* in hate, and the *e* in peep) he calls dental vowels. On the other hand, if, the tongue remaining as before, the lips be thrust out and drawn together, preserving as great an interior cavity of the fauces as possible, we shall have the Sounds of the vowels in *all, hope, poor, wood*. These he calls labial vowels, &c. These distinctions are to a certain extent correct and reasonable, but they give us no insight into the question, *What it is which constitutes the essential distinction between vowel and vowel, and on what part of the mechanism of the voice do vowel Sounds depend?*

In 1779, the Imperial Academy of Petersburg proposed as one of their prize questions, an inquiry into the nature of the vowel sounds A E I O U, and the construction of an instrument capable of artificially imitating them. The prize was awarded to M. Kratzenstein, whose curious Memoir on the subject the reader may find in the XXIst volume of the *Journal de Physique*, p. 358. His principle consisted in the adaptation of a reed in all essential respects similar to Grenié's, where the tongue passes to and fro through the slit without contact, to a set of pipes of peculiar forms, some of them very odd ones, and for whose shapes no other reason could be given than their success on trial. This, however, was a great step. It showed the vowel quality of a Sound to be something distinct from mere *pitch*, and susceptible of being produced at pleasure by mechanical artifice. Pursuing this idea, Mr. Willis has lately entered more extensively into the subject, and, in a Paper recently printed in the IIIrd volume of the *Transactions of the Cambridge Philosophical Society*, has succeeded in educing all the vowel Sounds by a mere combination of a reed on Kratzenstein's construction with a cylindrical pipe of variable length, and investigating the laws of their production.

This may be the place to remark the extreme imperfection of our written language in its representation of vowels and consonants. We have six letters which we call vowels, each of which, however, represents a variety of Sounds quite distinct from each other, and while each encroaches on the functions of the rest, a great many very good simple vowels are represented by binary or even ternary combinations. On the other hand, some single vowel letters represent true diphthongs, (as the long sound of *i* in alike, and that of *u* in rebuke,) consisting of two distinct simple vowels pronounced in rapid succession, while, again, most of what we call diphthongs are simple vowels, as bleak, thief, laud, &c. This will render an enumeration of our English elementary Sounds, as they really exist in our language, no matter how written, not irrelevant. We have therefore assembled in the following synoptic table sufficient examples of each to render evident their nature, accompanied with occasional instances of the corresponding Sounds in other languages. The syllables which contain the Sounds intended to be instanced are printed in italics where words of more than one syllable are instanced.

1. { Rook; Julius; Rude; Poor; Womb; Wound; *Ouvrir*, (Fr.)
2. { Good; Cushion; Cuckoo; Rund, (Germ.); Gusto, (Ital.)
3. Spurt; Assert; Dirt; *Virtue*; Dove; Double; Blood.
4. { Hole; Toad.
5. { All; Caught; *Organ*; Sought; Broth; Broad.
6. { Hot; *Comical*; *Kommen*, (Germ.)
7. Hard; *Braten*, (Germ.); *Charlatan*, (Fr.)
8. Laugh; Task.
9. Lamb; Fan; That.
10. Hang; Bang; Twang.
11. Hare; Hair; Heir; Were; Pear; Hier, (Fr.); *Lehren*, (Germ.)
12. Iame; Tame; Crane; Faint; *Layman*; Mème, (Fr.); *Stüdchen*, (Germ.)
13. Lemon; Dead; Said; Any; Every; Friend; *Besser* (Germ.); *Eloigner*, (Fr.)
14. Liver; Diminish; Persevere; Believe.
15. Peep; Leave; Believe; *Sieben*, (Germ.); *Coquille*, (Fr.)
16. s; sibilus; cipher; the last vowel and the first consonant.

True Diphthongs.

1. Life; The Sounds No. 5 and No. 13, slurred as rapidly as possible, produce our English *i*, which is a real diphthong.
2. Brow; Plough; *Laufen*, (Germ.) The vowel Sound No. 5 quickly followed by No. 1.
3. Oil; *Küen*, (Germ.); No. 4 succeeded by No. 13.
4. Rebuke; Yew; You; No. 13 succeeded by No. 1.
5. Yoke; No. 13 succeeded by No. 3.
6. Young; Yearn; Hear; Here; No. 13 succeeded by No. 2 more or less rapidly

Part III.

357.
Amman's
work on the
voice.

358.
Kratzen-
stein's inves-
tigation of
the vowels.

359.
Remarks on
written lan-
guage

360.
Synoptic
table of
English
vowel
Sounds

361.
Diphthongs.

Sound. The consonants present equal confusion. They may be generally arranged in three classes: sharp Sounds, flat ones, and indifferent or neutral. The former two having a constant relationship or parallelism to each other, thus:

362. SHARP CONSONANTS. S. *sell, cell*; σ. (as we will here denote it) *shame, sure, schirm*, (Germ.); θ. *thing*; F *fright, enough, phantom*; K. *king, coin, quiver*; T. *talk*; P. *papa*.
363. FLAT CONSONANTS. Z. *zenith; casement*; ζ. *pleasure, jardin*, (French); ð. the *th* in the words *the, that, thou*; V. *vile*; G. *good*; D. *duke*; B. *babe*.
364. NEUTRAL CONSONANTS. L. *lily*; M. *mamma*; N. *Nanny*; ν. *hang*; to which we may add the nasal N in *gnu, Ætna, Dnieper*, which, however, is not properly an English Sound. R. *rattle*; H. *hard*.
365. COMPOUND CONSONANTS. C, or Tσ. *church, cicerone*, (Ital.) and its corresponding flat sound J. or D ζ. *jest, gender*; X. *extreme, Xerxes*; ξ. *exasperate, exalt, Xerxes*; &c. &c.

367. We have here a scale of 13 simple vowels and 21 simple consonants, 33 in all, which are the fewest letters with which it is possible to write English. But on the other hand, with the addition of two or three more vowels, and as many consonants, making about 40 characters in all, every known language might probably be effectually reduced to writing, so as to preserve an exact correspondence between the writing and pronunciation; which would be one of the most valuable acquisitions not only to philologists but to mankind, facilitating the intercourse between nations, and laying the foundation of the first step towards a universal language, one of the great desiderata at which mankind ought to aim by common consent.

368. This enumeration will serve to show what are the difficulties which any one must contend with in constructing, what has been often attempted, a *talking engine*. Still the partial success obtained by Kratzenstein, and about the same time by Kempelen, who has given a very curious account of his experiments in *Mécanisme de la Parole*, ought to encourage further trials.

369. To return, however, to Mr. Willis's curious and novel researches. He relates that, having provided an apparatus consisting of a wind-chest, or reservoir, connected with a pair of double bellows, and opening into a *port-vent*, having a free reed, on Kratzenstein's, or Grenié's construction, at its termination, his first object was to verify Kempelen's account of the vowels. He therefore adapted his reed to the bottom of a funnel-shaped circular cavity, open at top, as in fig. 172, which represents a section of the apparatus, and on making the reed speak, and placing his hand in various positions pointed out by Kempelen within the funnel, he obtained the vowels A (No. 5.), E (No. 10.), I (No. 13.), O (No. 3.), U (No. 1.) very distinctly. On using, however, a shallower cavity these positions became unnecessary, and the hand might, he found, be replaced by a flat board slid over the mouth of the cavity; and by using a very shallow funnel, as represented in fig. 173, he succeeded in obtaining the whole series in the order U (No. 1.), O (No. 3.), A (No. 5.), E (No. 10.), I (No. 13.)

370. Being thus led away from Kempelen's experiment, he proceeded to try the effect of adapting to the reed cylindrical tubes, whose length could be varied at pleasure by sliding joints. This was easily accomplished by fixing the reed with its *port-vent* into the end of a pretty long horizontal pipe coming off from the wind-chest, over which on its outside a tube, open at both ends, was made to slide on leather wrapped round it in the manner of a piston, and capable of being lengthened, by the attachment of pieces of similar tube of its own length, to any extent. He thus describes the results so obtained. Let *abcd* represent the length of the outer, or sounding pipe, projecting beyond the reed, and take *ab, bc, cd*, &c. equal to the length of a stopped pipe in unison with the reed employed, that is equal to half the length of the sonorous wave of the reed. If, now, the pipe be drawn out gradually, the tone of the reed, retaining its pitch, first puts on in succession the vowel qualities I E A O U. As the length approaches *ac* the same series makes its appearance in an inverted order, as represented in the diagram, then on passing the length *ac* in direct order again, and so on in cycles, each cycle being merely a repetition of the foregoing, but the vowels becoming less and less distinct in each successive cycle, and the distance of any given vowel from its respective central points *a, c*, &c. being the same in all the cycles.

371. If another reed be adapted to the same pipes having a different fundamental Sound or sonorous wave, the same phenomena will be produced, only that the central points of the new cycles will now be at a distance from each other equal to the sonorous wave of the new reed, but the *distances of the several vowel points from the centres of the respective cycles will be the same as before*; so that, generally, if the reed wave $ac = 2a$, and the length of the pipe which first produces any given vowel, from *a*, be equal to *u*, the same vowel will be constantly reproduced by a pipe whose length $= 2na \pm v$, *n* being any whole number.

372. When the pitch of the reed is high, so that the length *ac* of its wave is less than twice the distance *aU* corresponding to any vowel, all the vowels beyond that distance become impossible. If, for instance *ac* be less than $2aU$, but greater than $2aO$, the series will never extend so far as U, but on lengthening the pipe indefinitely the succession of vowels I E A O A E I will be repeated. If, in like manner, still higher notes be taken for the reed, more vowels will be cut off. This, Mr. Willis remarks, is exactly the case with the human voice: female singers being unable to pronounce U and O on the higher notes of their voices. For example, the proper length for a pipe to produce O is that which corresponds to the note C'' two octaves above the middle C of a piano-forte, and beyond this note in singing it will be found impossible to pronounce a distinct O.

373. Cylinders of the same length, or more generally cavities of any figure resounding the same note, give the same vowel when applied to one and the same reed.

374. The following table is given by Mr. Willis as expressing the distances from the central points of the cycles at which the several vowels are produced in inches.

Sound.

Part III.

	Vowel Sound according to its place in the Scale. Art. 360.	Example.	Length of Vowel Pipe in inches and decimals.	Piano-forte note cor- responding.
I	No. 13.	See	0.38 ?	G ^v
E	11	Pet (? Pay)	0.6	C ^v
A	10	Pay (? Pet)	1.0	D ^{iv}
	6 or 7	Paa	1.8	F ⁱⁱⁱ
	5	Part	2.2	D ⁱⁱ b
A ^o	4	Paw	3.05	D ⁱⁱⁱ b
O	4	Nought	3.8	G ⁱⁱ
	3	No	4.7	F ⁱⁱ b
U	2	But	Indefinite.	C ⁱⁱ
	1	Boot		

On this Table Mr. Willis observes that he does not despair of its completion and extension by future experiments, eventually furnishing Philologists with a correct measure for the shades of difference in the pronunciation of the vowels by different nations. One source of fallacious decision, however, it must be remarked, will subsist in its application, in the effect of contrast, on which much of the difference between vowels depends. Its influence indeed may be traced in the above Table itself. Thus Mr. Willis, assisted, no doubt, by the contrast arising from rapid and frequent transition, has been able to discriminate between the vowel Sounds yielded by pipes of the lengths 3.05 and 3.8, though the Sounds in the exemplifying words Paw and Nought, which he has chosen, are so closely allied that we confess our own inability to detect any shade of difference, for which reason we have designated them by the same number.*

Mr. Willis terminates this highly interesting Paper with some curious experiments and remarks on the mutual influence of a reed and a pipe with which it is connected, as also of the port-vent, which conducts the air to it. If a reed he made to sound in a pipe of variable length (l), the Sound yielded by it will remain constant till the length (l) (beginning we will suppose from o) becomes nearly equal to $\frac{a}{4}$, or one quarter the length of the Sound

wave of the reed; here it begins to flatten, and as l is still increased, continues to do so till the length somewhat exceeds $\frac{1}{2}a$, when it suddenly jumps back to a note about a quarter of a tone sharper than the original Sound of the reed, to which it, however, soon again descends, and continues stationary till the length l becomes nearly equal to $2a + \frac{1}{4}a$, when the flattening again commences, and continues till l exceeds $2a + \frac{1}{2}a$, and so on periodically, but less decidedly. The total amount of flattening is usually a whole tone. A jerk of the bellows, or a too hasty lengthening of the pipe, will make the pitch spring back much sooner than it would do with cautious management, nay, with proper dexterity it may be made to yield, just about the point of junction, a double note, composed of one flatter and one sharper than the reed would yield alone. Mr. Willis seems to think that in this case, however, the two Sounds are only quickly alternated so as to seem to go on at once. Examining the reed in a glass pipe with a magnifier, he found its excursions diminished when the note was flattened or sharpened; but when the double Sound was educed they were no longer well defined, but the tongue of the reed seemed thrown into strange convulsions. This recalls the experiment of Biot and Hamel described in Art 199, 202.

Being thus brought back to the subject of reeds and forced vibrations, we must not omit to recommend to our reader's attention the curious and elaborate dissertation of MM. Weber and Floss, entitled *Leges oscillationis oriundæ si duo corpora diversâ celeritate oscillantia ita conjungantur ut oscillare non possint nisi simul et synchronicè exemplo illustratæ TUBORUM LINGUATORUM*. A detailed comparison of their results with those of Mr. Willis, which the necessity of bringing this Essay to a close forbids us to enter into, would be very interesting. MM Weber and Floss agree with him in the periodical recurrence of the note of the reed at equal intervals, and in its flattening up to a certain point, &c.; while in other points there is diversity of result enough to make a careful revision of the whole subject well worth while; though, perhaps, it is not more than may be accounted for by the different constructions of their reeds; in the one set of experiments the oscillations of the tongue of the reed having been executed parallel, in the other at right angles to the axis of the cylinder. It is somewhat curious, that they seem to have entirely overlooked the vowel quality of the Sounds educed, perhaps from not having employed sliding tubes, and thus missing the effect of contrast.

We had proposed to have devoted a section to M. Savart's recent elegant application of his delicate methods of detecting and exploring sonorous vibrations to the determination of the law of elasticity in different directions with respect to the axes of crystalized bodies; but it would lead us too far, and we must be content to refer our readers to the XLIIId volume of the *Annales de Chimie* for information. The field is a wide one, and it will, we doubt not, be long before it is fully explored.

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Mr. Willis's experiments on mutual influence of reed and pipe.

Production of a double note.

377.
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* Let the reader pronounce slowly, and distinctly, the words Paw, Goaw, Naughty, Nought, for his own satisfaction.

Sound.

379.
Sound of
earthquakes.

Neither shall we devote a separate section to the description and explanation of acoustic phenomena which occur in Nature. Many such, indeed, have been sufficiently noticed already. In Art. 23 we have explained satisfactorily the origin of thunder, and we shall here only remark that the subterraneous thunder which accompanies earthquakes may (at least in some cases) be ascribed to a general cause not very dissimilar, the successive arrival at the ear of undulations propagated at the same instant from nearer and remoter points, or if from the same points, arriving by different routes, through strata of different elasticities.

380. The concise and unblunted propagation of Sound through water, remarked by Messrs. Colladon and Sturm, is curiously exemplified by the shock of an earthquake felt and heard at sea. The sensation is always described as that of striking on a rock; the Sound as that of grating on a gravelly bottom; none of the hard, rough Sounds of the first impulse being at all softened or rounded by the distance.

There is, however, one natural phenomenon so very surprising, and to us, we confess, so utterly inexplicable, though resting on the authority of ear-witnesses of such credit that it is impossible to disbelieve the facts, that we cannot forbear inserting a short description of it, with which we shall conclude.

Description
of the place
called Na-
kous.

There is a place about three leagues to the North of Tor, in the neighbourhood of Mount Sinai in Arabia Petræa, called *El Nakous*, (Nakous is the name of a sonorous metal plate used in the Greek convents in the East instead of a bell) from musical Sounds of a very singular and surprising character heard there. It has been visited by very few Europeans, two of whom, however, Mr. Seetzen and Mr. Gray of Oxford, have published accounts of it, the former in the *Monatliche Correspondenz*, (Oct. 1812;) the latter in Dr. Brewster's *Edinburgh Philosophical Journal*, where also Mr. Seetzen's account of it will be found translated, which is as follows:—

“After a quarter of an hour's walking, (from Wody El Nachel?) we reached the foot of a majestic rock of hard sandstone. The mountain was quite bare, and composed entirely of it. I found inscribed on it several Greek and Arabic names, and also some Koptic characters, which showed that the place had been visited for centuries. At noon we reached the part of the mountain called Nakous. There, at the foot of the ridge, we beheld an isolated peaked rock. Upon two sides this mountain presented two surfaces, so inclined, that the white and slightly adhering sand which covers it scarcely supports itself, and slides down with the smallest motion, or when the burning rays of the sun complete the destruction of its feeble cohesion. These two sandy declivities are about 150 feet high. They unite behind the insulated rock, and forming an acute angle, they are covered like the adjacent surfaces with steep rocks, which are mostly composed of a white and friable free-stone.

“The first Sound was heard an hour and a quarter after noon. We climbed with great difficulty as far as the sandy declivity, a height of 70 or 80 feet, and stopped under the rocks where the pilgrims are in the habit of placing themselves to listen. In climbing, I heard the Sound from beneath my knees, and this made me think that the sliding of the sand was the cause, not the effect, of the sonorous motion. At three o'clock the Sound was heard louder, and it lasted six minutes, when, having ceased for ten minutes, it began again. It appeared to me to have the greatest analogy to the humming-top; it rose and fell like the Sound of the *Æolian* harp. To ascertain the truth of my discovery, I climbed with the utmost difficulty to the highest rocks, and I slid down as fast as I could, and endeavoured, with the help of my hands and feet, to set the sand in motion. This produced an effect so great, and the sand in rolling under me made so loud a noise, that the earth seemed to tremble, and I certainly should have been afraid, had I been ignorant of the cause.

“But how can the motion of the sand produce so striking an effect, and which is, I believe, produced nowhere else? Does the rolling layer of sand act like a fiddle-bow, which, on being rubbed upon a plate of glass, raises and distributes into determinate figures the dust with which the plate is covered? Does the adherent and fixed layer of sand perform the part of the plate of glass, and the neighbouring rocks that of the sounding body? Philosophers must decide this.”

We give here M. Seetzen's account in preference to Mr. Gray's as being the earliest, and in his own words, preserving even his own conjectures (not the most plausible) on its cause, and we shall be glad if the visits of future travellers to the spot shall throw further light on this very strange phenomenon.

Slough, Feb. 3, 1830.

J. F. W. HERSCHEL.

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END OF VOLUME IV.

ERRATA IN THE ESSAY ON SOUND

<i>Art.</i>	<i>Line.</i>	<i>Error.</i>	<i>Correction.</i>
2,	5,	Hanksbee,	Hauksbee.
2,	15,	Hanksbee,	Hauksbee.
104,	6,	free,	quick and copious.
117,	10,	axes,	positions of the axis.
151,	11,	$x + a t = a$,	$x + a t > a$.
151,	16,	conditon,	condition.
178,	4,	volume,	column.
186,	7, and margin,	fig. 16,	fig. 17.
195,	7,	where,	whose.
209,	10,	times,	terms.
212,	1,	combinations,	cnmbination.
219,	4,	impulse,	impulses.
223,	8,	the extent,	the number, extent.
224,	7,	Sound	sense?
		Nevertheless spiders hear the sound of music. Vide Latreille's anecdote of Pelisson, who tamed one in the Bastille.	
235,	10,	more,	most.
236,	17,	vibrations,	vibrations of the quicker.
238,	3,	figure	figure 31.

ADDITIONAL ERRATA IN THE ESSAY ON LIGHT.

<i>Page.</i>	<i>Line.</i>	<i>Error.</i>	<i>Correction.</i>
416,	19,	α, β, γ	a, b, c .
434,	5,	organ,	origin.
434,	9,	maxima,	nature.
439,		electric attractions,	elective attractions.
452,	18,	1200,	1100.
456,	18 from bott.	on,	or.
457,	6,	two canals,	dele.
473,	33,	regularly,	irregularly.
518,	35,	mackled,	macled.
518,	11 from bott.	$(x^2 + y^2 + a^2)$,	$(x^2 + y^2 + a^2)^2$.
518,	3 from bott.	holes,	poles.
537,	2 from bott. note,	prelate,	prolate.
541,	12 from bott.	$(B'' v + C'' w) v$	$(B'' v + C'' w) u$.
544,	4,	second plane,	secant plane.
551,	34 from bott.		dele what is said about <i>camphor</i> , as that substance in a solid state <i>does</i> possess the rotatory property.
556,	7 from bott.	rays,	rings.
557,	6,	rays,	rings.
559,	17,	$\alpha \cdot \cos O A^2 +$	$y \cdot \cos O A^2 +$.
560,	17,	$y + y$	$Y + y$.
561,	17,	undulatory,	undulated.
571,	4, col. 1, }	extra,	extreme.
571,	7, col. 1, }		
574,	10,	$\mu \delta$,	$\delta \mu$.
577,	5,	hyposulphate of lime and,	dele.

Fig. 1 Art. 41

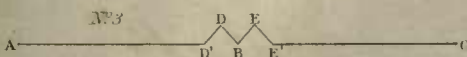
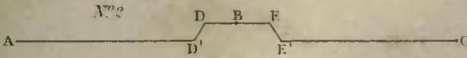
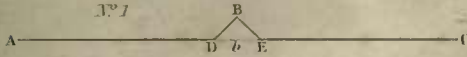
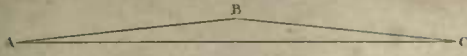


Fig. 2 Art. 41

Fig. 4 Art. 50

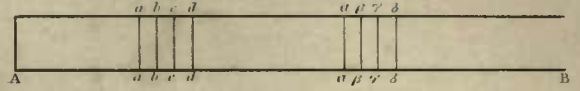


Fig. 6 Art. 117



Fig. 3 Art. 45

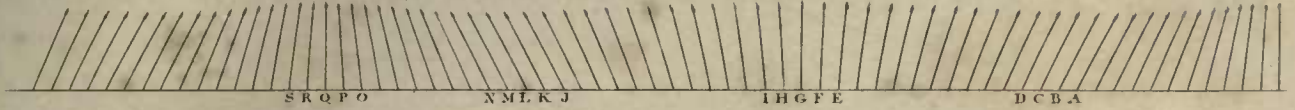
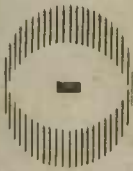


Fig. 7



Art. 113

Fig. 8



Fig. 9 Art. 140

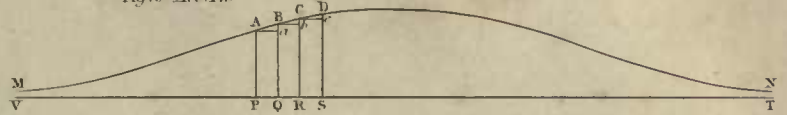


Fig. 10 Art. 150

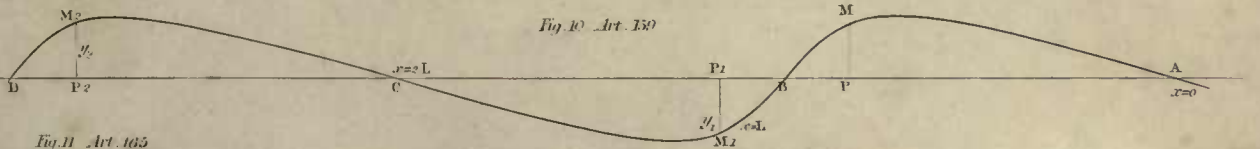


Fig. 11 Art. 165

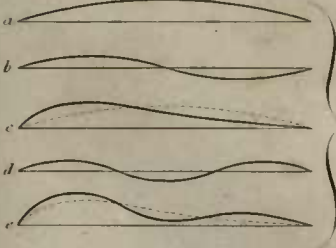


Fig. 13 Art. 176

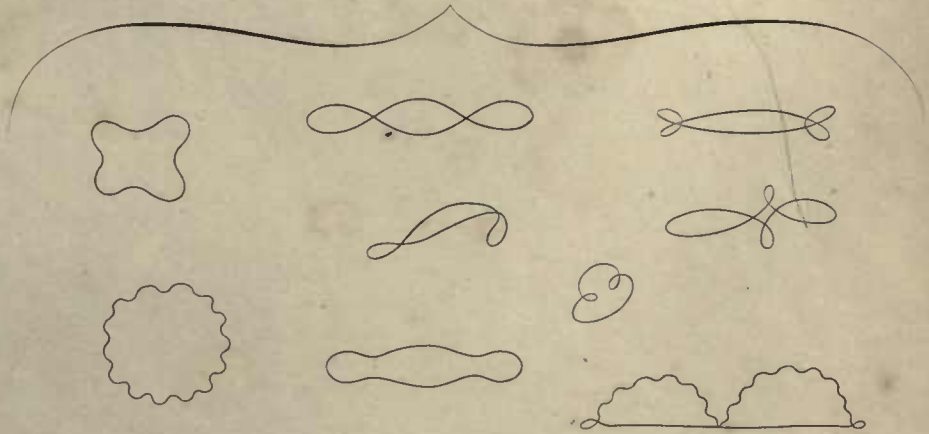


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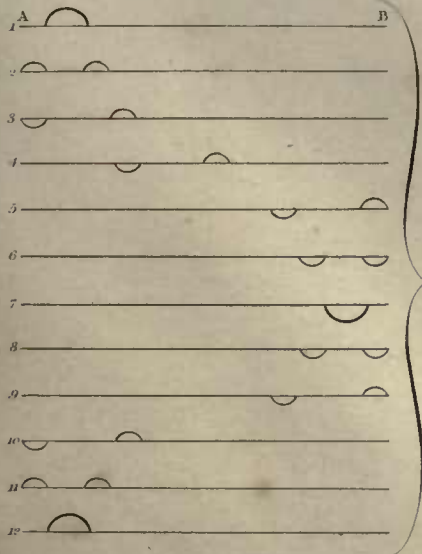


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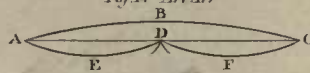
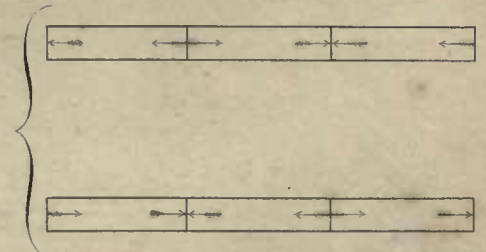


Fig. 15 Art. 182



Fig. 16 Art. 182



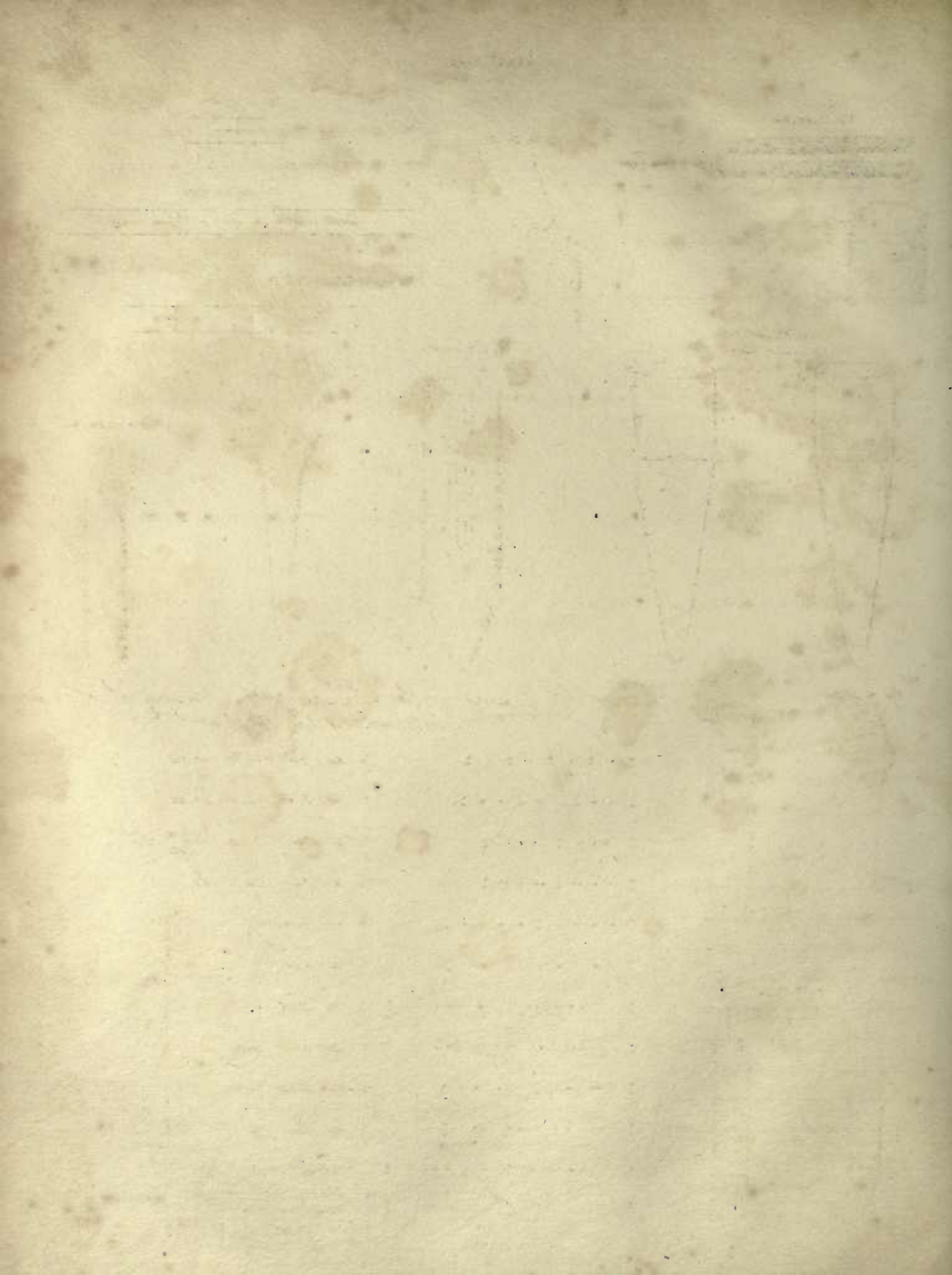


Fig. 17. Art. 186.

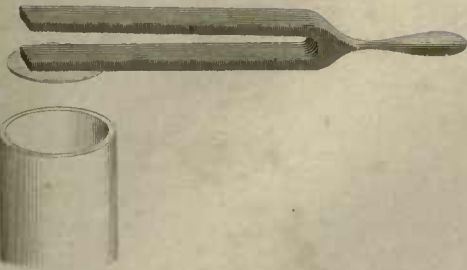


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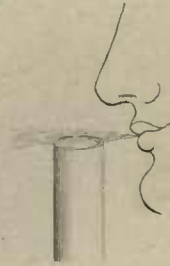


Fig. 19. Art. 190.



Fig. 20. Art. 190.



Fig. 21. Art. 191.



Fig. 22. Art. 198.

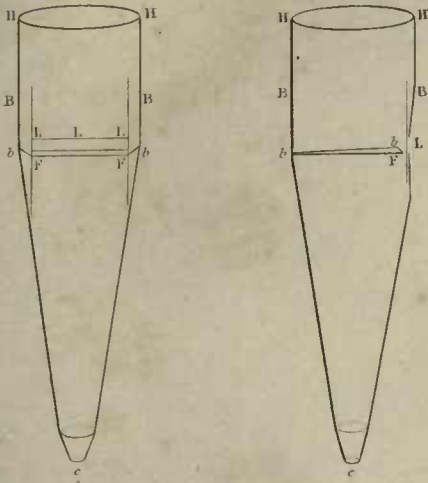


Fig. 23. Art. 199.

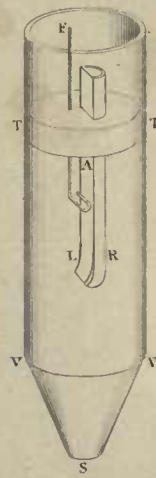


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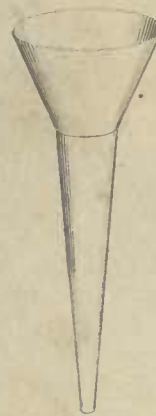


Fig. 25. Art. 200.



Fig. 26. Art. 200.

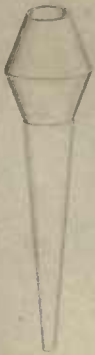


Fig. 27. Art. 206.



Order of Succession of the Pulses of the principal Harmonic intervals on the Ear.

• • • • •

$\frac{2}{1}$ the Octave, Fig. 29. Art. 210.

• • • • •

$\frac{3}{1}$ the Twelfth, Fig. 30. Art. 212.

• • • • •

$\frac{4}{1}$ the Fifth, Fig. 31. Art. 212, 238.

• • • • •

$\frac{5}{1}$ the Fourth, Fig. 32. Art. 213.

• • • • •

$\frac{7}{1}$ the Seventeenth.

• • • • •

$\frac{8}{1}$ the Tenth.

• • • • •

$\frac{9}{1}$ the Major Third.

• • • • •

$\frac{10}{1}$ the Minor Sixth.

• • • • •

$\frac{11}{1}$ the Major Sixth.

• • • • •

$\frac{12}{1}$ the Minor Third.

• • • • •

Common Chord.

• • • • •

Legner's Chord of the Fundamental Seventh, regarded as a Concord.

Fig. 33. Art. 241.

Fig. 34. Art. 213.

Fig. 35. Art. 253.

Fig. 28. Art. 207.

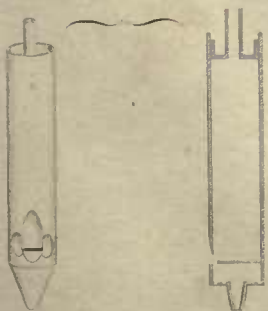


Fig. 36. Art. 269.



Fig. 37. Art. 269.

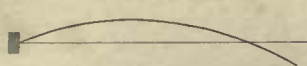


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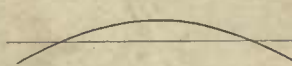


Fig. 39. Art. 269. 270.

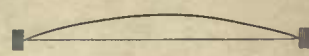


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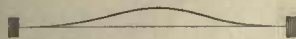


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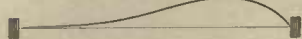


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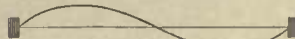


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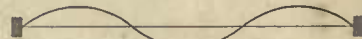


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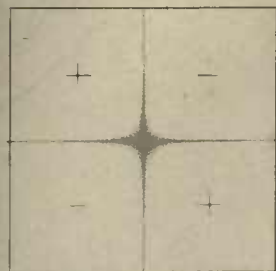


Fig. 45. Art. 275.

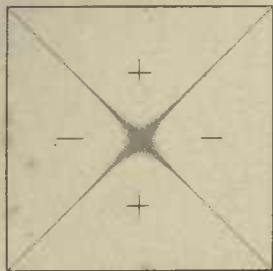


Fig. 46. Art. 276.

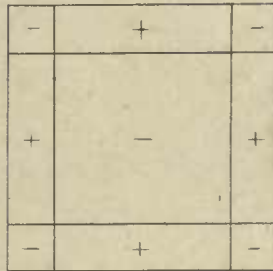


Fig. 47.

Art. 277.

Fig. 48.

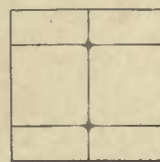


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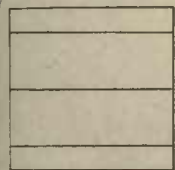


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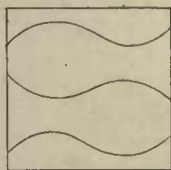


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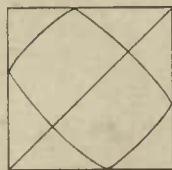


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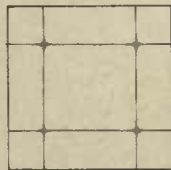


Fig. 53.



Fig. 54.



Fig. 56.



Fig. 56.



Fig. 57.

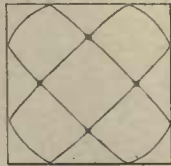


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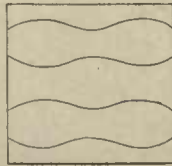


Fig. 59.



Fig. 60.

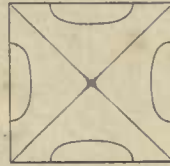


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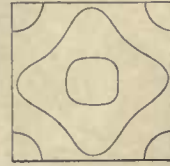


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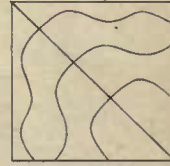


Fig. 63.



Fig. 64.

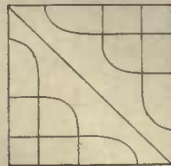


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Fig. 66.

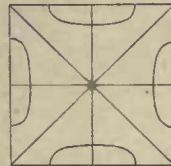


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Fig. 68.

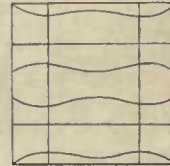


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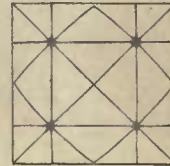


Fig. 70.

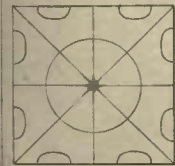


Fig. 71.

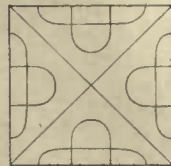


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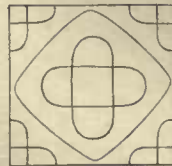


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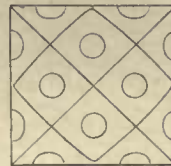


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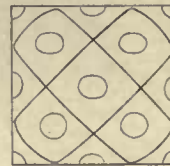


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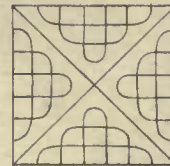
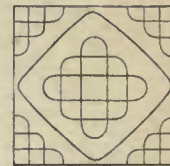


Fig. 76.



Art. 277.

SOUND.

Fig. 77.



Fig. 78.



Fig. 79.



Fig. 80.

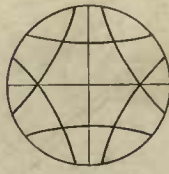


Fig. 81.

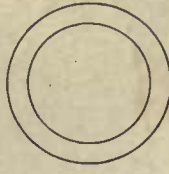


Fig. 82.



Fig. 83.



Fig. 84.

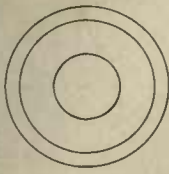


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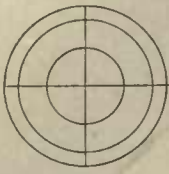


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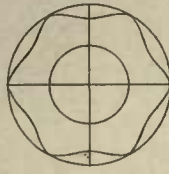


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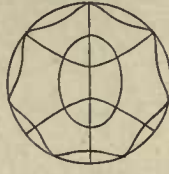


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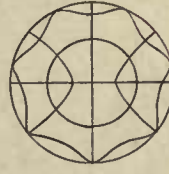


Fig. 89.



Fig. 90.



Art. 277.

Fig. 91.

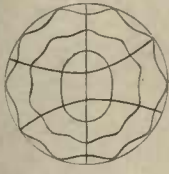


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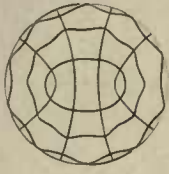


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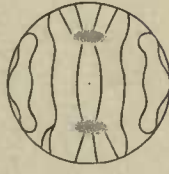


Fig. 94.

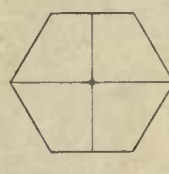


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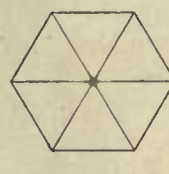


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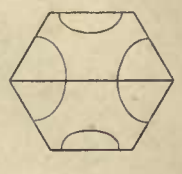


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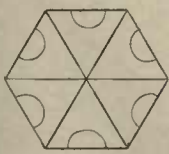


Fig. 98.



Fig. 99.

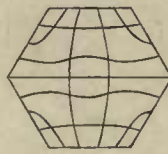


Fig. 100.



Fig. 101.



Fig. 102.



Fig. 103.



Fig. 104.



Fig. 105.



Fig. 106.



Fig. 107.



Fig. 108.



Fig. 109.



Fig. 110.

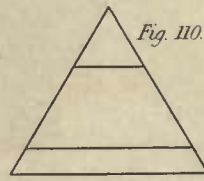


Fig. 111.

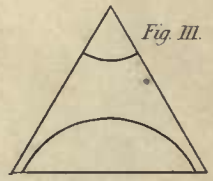


Fig. 112.

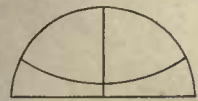


Fig. 113.



Fig. 114.



Fig. 115.

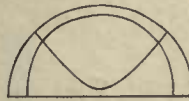


Fig. 116.



Fig. 117.

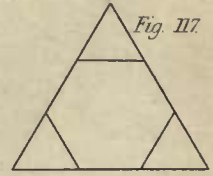


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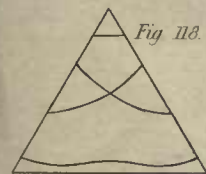


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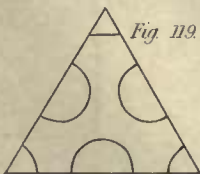


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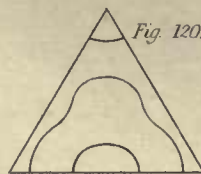


Fig. 121.



Fig. 122.

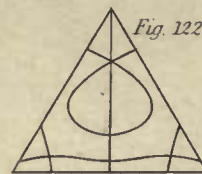
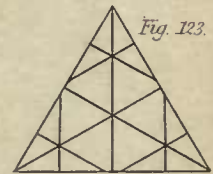


Fig. 123.



Art. 278.

Fig. 124 Art. 281

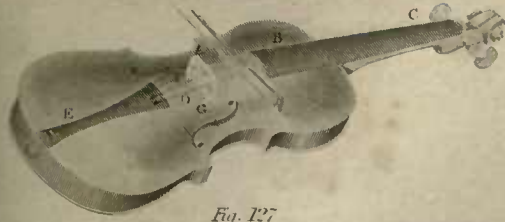


Fig. 125 Art. 284

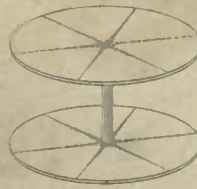


Fig. 126 Art. 285

Fig. 127

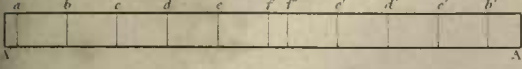


Fig. 128



Fig. 131



Fig. 132



Fig. 136



Fig. 137



Fig. 138



Fig. 139



Fig. 140



Art. 295

Fig. 141 Art. 298

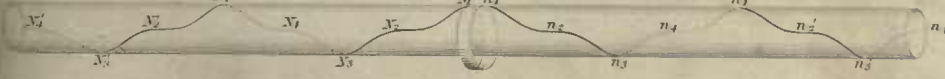


Fig. 143 Art. 305

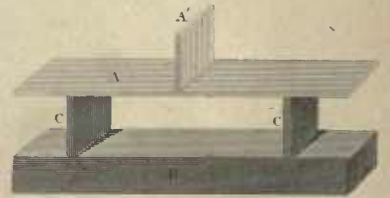


Fig. 145 (a) Art. 307



Fig. 142 Art. 303



Fig. 145 (c)

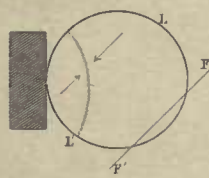


Fig. 145 (b)



Fig. 145 (d)

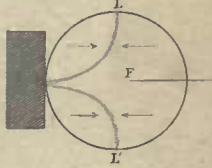


Fig. 140 Art. 308

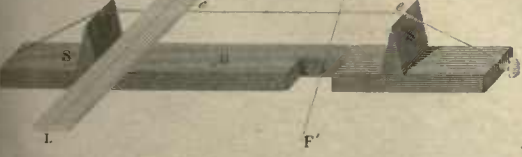
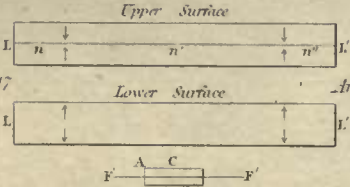


Fig. 147



Art. 308

Fig. 148
Upper Surface



Lower Surface

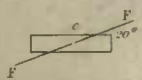


Fig. 151

Fig. 152

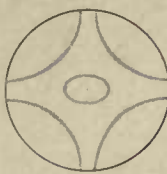
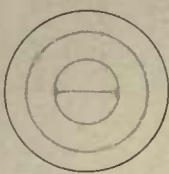
Fig. 153

Fig. 154

Fig. 155

Fig. 156

Fig. 157



Art. 317

Fig. 158

Fig. 159

Fig. 160

Fig. 161

Fig. 162

Fig. 163

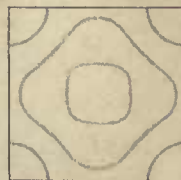
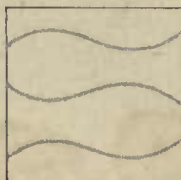
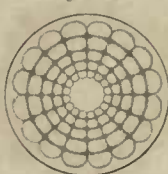
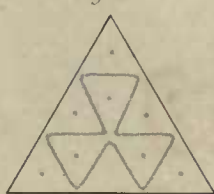
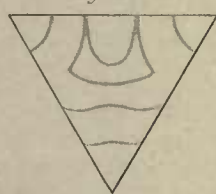
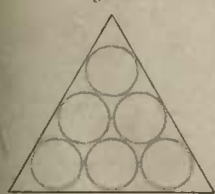


Fig. 164

Fig. 165

Fig. 166



Section of Fig. 169.



Fig. 170 Art. 321

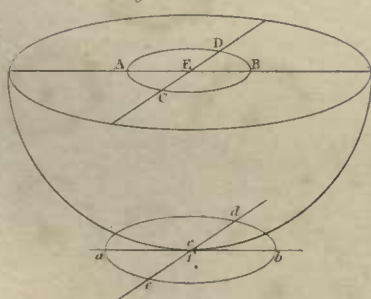
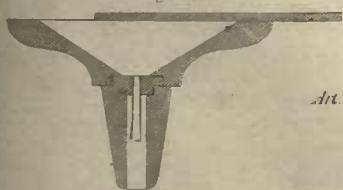


Fig. 172



Fig. 173



Art. 369

Fig. 174

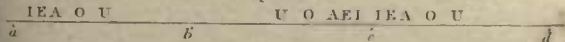


Fig. 167 Art. 320

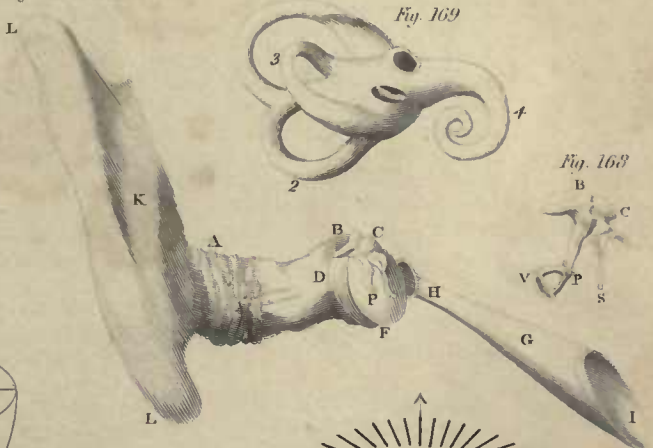


Fig. 169

Fig. 168

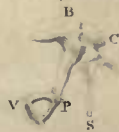
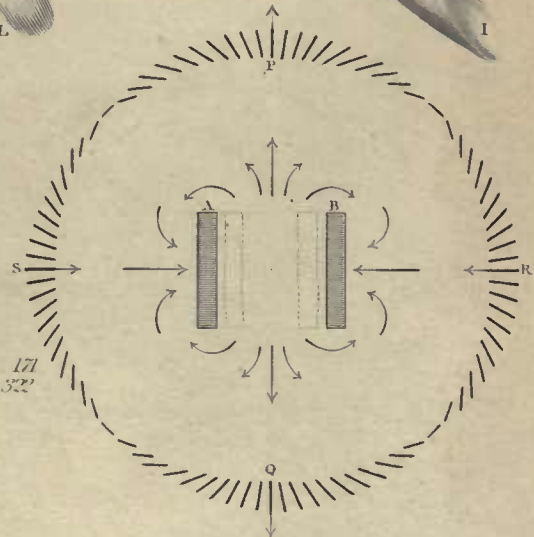


Fig. 171 Art. 322



L I G H T.

PART I.

Of unpolarized Light.

§ I. INTRODUCTION.

Light.

Part I.

1.
Subject
stated.

IN this article we propose to give an account of the properties of light; of the physico-mathematical laws which regulate the direction, intensity, state of polarization, colours, and interferences of its rays; to state the theories which have been advanced for explaining the complicated and splendid phenomena of optics; to explain the laws of vision, and their application, by the combined ingenuity of the philosopher and the artist, to the improvement of our sight; and the examination and measurement of those objects and appearances which, from their remoteness, minuteness, or refinement, would otherwise elude our senses.

2.

The sight is the most perfect of our senses; the most various and accurate in the information it affords us; and the most delightful in its exercise. Apart from all considerations of utility, the mere perception of light is in itself a source of enjoyment. Instances are not wanting of individuals debarred from infancy by a natural defect from the use of their eyes, whose highest enjoyment still consisted in that feeble glimmering a strong sunshine could excite in their obstructed organs; but when to this we join the exact perception of form and motion, the wondrous richness and variety of colour, and the ubiquity conferred upon us by just impressions of situation and distance, we are lost in amazement and gratitude.

3.

What are the means and mechanism by which we receive this inestimable benefit? Curiosity may well prompt the inquiry, but a more direct interest urges us to pursue it. Knowledge is power; and a careful examination of the means by which we see, not only may, but actually has led us to the discovery of artificial aids by which this particular sense may be strengthened and improved to a most extraordinary degree; giving to man at once the glance of the eagle, and the scrutiny of the insect—by which the infirmities of age may be deferred or remedied—nay, by which the sight itself when lost may be restored, and its blessings conferred after long years of privation and darkness, or on those who from infancy have never seen. But as we proceed in the inquiry we shall find inducements enough to pursue it from purely intellectual motives. A train of minute adaptation and wonderful contrivance is disclosed to us, in which are blended the utmost extremes of grandeur and delicacy; the one overpowering, the other eluding, our conceptions. In consequence of those peculiar and singular properties which are found to belong to light in its various states of polarization, it affords to the philosopher information respecting the intimate constitution of bodies, and the nature of the material world, totally distinct from the more general impressions of form, colour, distance, &c. which it conveys to the vulgar. Its notices, it is true, in this respect, are addressed rather to the intellect than the sense; but they are not on that account less real, or less to be depended on. Polarized light is, in the hands of the natural philosopher, not merely a medium of vision; it is an instrument by which he may be almost said to feel the ultimate molecules of natural bodies, to detect the existences and investigate the nature of powers and properties ascertainable only by this test, and connected with the more important and intricate inquiries in the study of nature.

4.

Bodies re-
garded as
self-lumi-
nous and
opaque.

The ancients imagined vision to be performed by a kind of emanation proceeding from the eye to the object seen. Were this the case, no good reason could be shown why objects should not be seen equally well in the dark. Something more, however, is necessary for seeing than the mere presence of the object. It must be in a certain state, which we express by saying that it is *luminous*. Among natural bodies some possess in themselves the property of exciting in our eyes the sensation of *brightness*, or light; as the sun, the stars, a lamp, red-hot iron, &c. Such bodies are called *self-luminous*; but by far the greater part possess no such property. Such bodies in the dark remain invisible, though our eyes are turned directly towards them; and are therefore termed dark, non-luminous, or *opaque*, though this word is also used occasionally to express want of transparency. All bodies, however, though not luminous of themselves, nor capable of exciting any sensation in our eyes, become so on being placed in the presence of a self-luminous body. When a lamp is brought into a dark room, we see, not only the lamp, but all the other bodies in the room. They are all, so long as the lamp remains, rendered luminous, and are in their turn capable of illuminating others. Thus a sunbeam passing into a darkened room renders luminous, and therefore visible, a sheet of paper on which it falls; and this, in its turn, will in like manner illuminate the whole apartment, and render visible every object it contains, so long as it continues to receive the sunbeam. The moon and planets are opaque bodies; but those parts of them on which the sun shines become for the time luminous, and perform all the offices of self-luminous bodies. Thus we see, that the communication which we call light, subsists not only between luminous bodies and our eyes, but between luminous and non-luminous bodies, or between luminous bodies and each other.

Opaque bo-
dies, become
luminous in
the presence
of a lumi-
nous body.

Light.

5. Opaque bodies intercept light.

Light emanates in straight lines,

in all directions,

6. and from every physical point in a luminous surface.

Many bodies possess the property of intercepting this peculiar intercourse between luminous bodies and our eyes, or other bodies. A screen of metal interposed between the sun and our eyes prevents our seeing it; interposed between the sun and a sheet of white paper, or other object, it *casts a shadow* on such object: *i. e.* renders it non-luminous. By this power of bodies to intercept light, we learn that the communication which constitutes it takes place in straight lines. We cannot see through a bent metallic tube, nor perceive the least glimpse of light through three small holes in as many plates of metal placed one behind the other at a distance, unless the holes be situated exactly in one straight line. Moreover, the shadows of bodies, when fairly received on smooth surfaces perpendicular to the line in which the luminous body lies, are similar in figure to the section of the body which produces them, which could not be, except the light were communicated in straight lines from their edges to the borders of the shadow. We express this property by saying that light *emanates*, or *radiates*, or is *propagated* from luminous bodies in straight lines; by which expressions nothing more is to be understood than the mere fact, without in any way prejudging the question as to the intimate nature of this emanation. Moreover, it *emanates from them in all directions*, for we see them in all situations of the eye, provided nothing intervene to intercept the light. This is the essential distinction between luminous bodies and optical images; from which, as we shall see, light emanates only in certain directions. Whether it emanates *equally* in all directions will be considered farther on.

Light also radiates from every point (at least from every *physical* point) of a luminous body. This may, perhaps, be regarded as a truism; for those points of a luminous body from which (as from the spots in the sun) no light emanates, are, in fact, non-luminous, and the body is only partially so; the figure of the spots is only seen, because it is also necessarily that of the luminous surface which surrounds them. Still it should be borne in mind, for reasons which will appear when we come to speak of the formation of images. It is possible (nay, probable) that a luminous surface, such as that of the flame of a candle, may consist only of an immense but finite number of luminous points, surrounded by non-luminous spaces; but it is not ocular demonstration this idea admits of; and it is sufficient for our purpose that, so far as our senses inform us, every physical point of a luminous surface is a *separate* and *independent* source of light. We may magnify in a telescope the sun's disc to any extent, and intercept all but the very smallest portions of it, (spots excepted,) yet the *visibility* of one part is no way impaired by the exclusion of the rest. In this sense the proposition is no truism, but an important fact, of which we shall hereafter trace the consequences.

7. When the sun shines through a small hole, and is received on a white screen behind at a considerable distance, we see a round luminous spot, which enlarges as the screen recedes from the hole. If we measure the diameter of this image at different distances from the hole, it will be found that (laying out of the question certain small causes of difference not now in contemplation) the angle subtended by the spot at the centre of the hole is constant, and equal to the apparent angular diameter of the sun. The reason of this is obvious; the light from every point in the sun's disc passes through the hole, and continues its course in a right line beyond it till it reaches the screen. Thus every point in the sun's disc has a point corresponding to it in the screen; and the whole circular spot on the screen is, in fact, an *image* or representation of the face of the sun. That this is really the case, is evidently seen by making the experiment in the time of a solar eclipse, when the image on the screen, instead of appearing round, appears horned, like the sun itself.* In like manner, if a pin-hole in a card be held between a candle and a piece of white paper in a dark room, an exact representation of the flame, but inverted, will be seen depicted on the paper, which enlarges as the paper recedes from the hole; and if in a dark room a white screen be extended at a few feet from a small round hole, an exact picture of all external objects, of their natural colours and forms, will be seen traced upon the screen; moving objects being represented in motion, and quiescent ones at rest. (See fig. 6.) To understand this, we must recollect that all objects exposed to light are luminous; that from every physical point of them light radiates in all directions, so that every point in the screen is receiving light at once from every point in the object. The same may be said of the hole; but the light that falls on the hole passes through it, and continues its course in straight lines behind. Thus the hole becomes the vertex of a conoidal solid prolonged both ways, having the object for its base at one end, and the screen at the other. The section of this solid by the screen is the picture we see projected on it, which must manifestly be exactly similar to the object, and inverted, according to the simplest rules of Geometry.

Fig. 6.

8. Now if in our screen receiving (suppose) the image of the sun we make another small hole, and behind it place another screen, the light falling on the space occupied by this hole will pass beyond it, and reach the other screen; but it is clear that it will no longer dilate itself, after passing through the second hole, and form another image of the whole sun, but only an image of that very minute portion of the sun which corresponds to the space occupied in his image on the first screen by the hole made there. The lines bounding the conoidal surface will* in this case have much less divergency, and, if the holes be small enough, and very distant from each other, will approach to physical lines, and that the nearer, as the holes are smaller and their distance greater. (See fig. 7.) If we conceive the holes reduced to mere physical points, these lines form what we call rays of light. Mathematically speaking, a ray of light is an infinitesimal pyramid, having for its vertex a luminous point, and for its base an infinitely small portion of any surface illuminated by it, and supposed to be filled with the luminous emanation, whatever that may be. This pyramid, in homogeneous media, and when the course of the ray is not interrupted, has, as we have seen,

Fig. 7.

* In the eclipse of September 7, 1820, this horned appearance was very striking in the luminous interstices between the shadows of small irregular objects, as the leaves of trees, &c. It was noticed by those who had no idea of its cause.

its sides straight lines. If cases should occur (as they will) when the course of the light is curved, or suddenly broken, we may still conceive such a pyramid having curved or broken sides to correspond; or we may (for brevity's sake) substitute for it a mere mathematical line, straight, curved, or broken, as the case may be.

9. Light requires time for its propagation. Two spectators at different distances from a luminous object of suddenly disclosed, will not begin to see it at the same mathematical instant of time. The nearer will see it sooner than the more remote; in the same way as two persons at unequal distances from a gun hear the report at different moments. In like manner, if a luminous object be suddenly extinguished, a spectator will continue to see it for a certain time afterwards, as if it still continued luminous, and this time will be greater the farther he is from it. The interval in question is, however, so excessively small in such distances as occur on the earth's surface, as to be absolutely insensible; but in the immense expanse of the celestial regions the case is different. The eclipses and emersions of Jupiter's satellites become visible much sooner (nearly a quarter of an hour) when the earth is at its least distance from Jupiter than when at its greatest. *Light then takes time to travel over space.* It has a finite, though immense velocity, viz. 192500 miles per second; and this important conclusion, deduced by calculation from the phenomenon just mentioned, and which, if it stood unsupported, might startle us with its vastness, and incline us to look out for some other mode of explanation, receives full confirmation from another astronomical phenomenon, the *aberration of light*, which (without entering into any close examination of the mode in which vision is produced) may be explained as follows:

10. Let a ray of light from a star S, at such a distance that all rays from it may be regarded as parallel, be received on a small screen A, having an extremely minute opening A in its centre; and let that ray which passes through the opening be received at any distance AB, on a screen B perpendicular to its direction; and let B be the point on which it falls, the whole apparatus being supposed at rest. If then we join the points A, B by an imaginary line, that line will be the direction in which the ray has really travelled, and will indicate to us the direction of the star; and the angle between that line and any fixed direction (that of the plumb-line, for instance) will determine the star's place as referred to that fixed direction. For simplicity, we will suppose this angle nothing, or the star directly vertical; then the point B on which the ray falls will be precisely that marked by a plumb-line let fall from A; and the direction in which we judge the star to lie will coincide precisely with the direction of gravity. Such will be the case, supposing the earth, the spectator, and the whole apparatus at rest; but now suppose them carried along in a horizontal direction AC, BD, with a uniform and equal velocity, of whose existence they will therefore be perfectly insensible, and the plumb-line will hang steadily as before, and coincide with the same point of the screen. At the moment when the ray SA from the star passes through the orifice A, let A, B be the respective places of the orifice, and the point on the screen vertically below it. When the ray has passed the orifice, it will pursue its course in the same straight line SAB as before, independent of the motion of the apparatus, and in some certain time $\left(= \frac{\text{distance AB}}{\text{velocity of light}} = t \right)$ will reach the lower screen. But in this time the aperture, screens, and plumb-line will have moved away through a space

$$Aa = Bb \left(= t \times \text{velocity of motion} = AB \times \frac{\text{earth's velocity}}{\text{velocity of light}} \right)$$

At the instant, then, that the ray impinges on the lower screen, the plumb-line will hang, not from A on B, but from a on b; and a being the *real orifice*, and B the *real point of incidence* of the light on the screen, the spectator, judging only from these facts, will necessarily be led to regard the ray as having deviated from its vertical direction, and as inclining from the vertical, in the direction of the earth's motion through an angle whose

$$\text{tangent is } \frac{Aa}{AB} \text{ or } \frac{\text{earth's velocity}}{\text{velocity of light}}.$$

11. The eye is such an apparatus. Its retina is the screen on which the light of the star or luminary falls, and we judge of its place only by the actual point on this screen where the impression is made. The pupil is the orifice. If, the eye preserving a fixed direction, the whole body be carried to one side with a velocity commensurate to that of light, before the rays can traverse the space which separates the pupil from the retina, the latter will have shifted its place; and the point which receives the impression is no longer the same which would have received it had the eye and spectator remained at rest; and this deviation is the *aberration of light*.

12. Every spectator on the earth participates in the general motion of the whole earth, which in its annual orbit about the sun is very rapid, and though far from equal to that of light, is by no means insensible, compared to it. Hence the stars, the sun, and planets, all appear removed from their true places in the direction in which the earth is moving.

13. This direction is varying every instant, as the earth describes an orbit round the sun. The direction therefore of the apparent displacement of any star from its true situation continually changes, *i. e.* the apparent place describes a small orbit about the true. This phenomenon is that alluded to. It was noticed as a fact by Bradley, while ignorant of its cause, that the stars appear to describe annually small ellipses in the heavens of about 40" in diameter. The discovery of the velocity of light by the eclipses of Jupiter's satellites, then recently made by Roemer, however, soon furnished its explanation. Later observations, especially those of Brinkley and Struve, have enabled us to assign, with great precision, the numerical amount of this inequality, and thence to deduce the velocity of light, which by this method comes out 191515 miles per second, differing

Light. from the former only by a two hundredth part of its whole quantity. This determination is certainly to be preferred.

14. But this is not the only information respecting light which astronomical observations furnish. We learn from them also, "That the light of the sun, the planets, and all the fixed stars, travels with one and the same velocity." Now as we know these bodies to be at different and variable distances from us, we hence conclude that the velocity of light is independent of the particular source from which it emanates, and the distance over which it has travelled before reaching our eye.

15. The velocity of light, therefore, in that free and perhaps void space which intervenes between us and the planets and fixed stars, cannot be supposed other than uniform; and the calculations of the eclipses of Jupiter's satellites, and the places of the distant planets made on this supposition agreeing with observation, prove it to be so. In entering such media as it traverses, when arrived within the limits of the atmospheres of the earth and other planets, we shall find reason hereafter to conclude that its velocity undergoes a change; but, at all events, we have no reason to suppose it to differ in different parts of one and the same homogeneous medium.

16. The enormous velocity here assigned to light, surprising as it may seem, is among those conclusions which rest on the best evidence that science can afford, and may serve to prepare us for other yet more amazing numerical estimates. It is when we attempt to measure the vastness of the phenomena of nature with our feeble scale of units, such as we are conversant with on this our planet, that we become sensible of its insignificance in the system of the universe. Demonstrably true as are the results, they fail to give us distinct conceptions; we are lost in the immensity of our numbers, and must have recourse to other ways of rendering them sensible. A cannon ball would require seventeen years at least to reach the sun, supposing its velocity to continue uniform from the moment of its discharge. Yet light travels over the same space in $7\frac{1}{2}$ minutes. The swiftest bird, at its utmost speed, would require nearly three weeks to make the tour of the earth. Light performs the same distance in much less time than is required for a single stroke of his wing; yet its rapidity is but commensurate to the distances it has to travel. It is demonstrable that light cannot possibly arrive at our system from the nearest of the fixed stars in less than five years, and telescopes disclose to us objects probably many thousand times more remote.

But these are considerations which belong rather to astronomy than to the present subject; and we will, therefore, return to the consideration of the phenomena of emitted light.

§ II. Of Photometry.

17. Of these, one of the most striking is certainly the diminution of the illuminating power of any source of light, arising from an increase of its distance. We see very well to read by the light of a candle at a certain distance: remove the candle twice, or ten times as far, and we can see to read no longer.

The numerical estimation of the degrees of intensity of light constitutes that branch of optics which is termed *Photometry*. (*φως, μετρω.*)

18. If light be a material emanation, a something scattered in minute particles in all directions, it is obvious that the same quantity which is diffused over the surface of a sphere concentric with the luminous points, if it continue its course, will successively be diffused over larger and larger concentric spherical surfaces; and that its intensity, or the number of rays which fall on a given space, in each will be inversely as the whole surfaces over which it is diffused; that is, inversely as the squares of their radii, or of their distances from the source of light. Without assuming this hypothesis, the same thing may be rendered evident as follows. Let a candle be placed behind an opaque screen full of small equal and similar holes; the light will shine through these, and be intercepted in all other parts, forming a pyramidal bundle of rays, having the candle in the common vertex. If a sheet of white paper be placed behind this, it will be seen dotted over with small luminous specks, disposed exactly as the holes in the screen. Suppose the holes so small, their number so great, and the eye so distant from the paper that it cannot distinguish the individual specks, it will still receive a general impression of brightness; the paper will appear illuminated, and present a mottled appearance, which, however, will grow more uniform as the holes are smaller, and closer, and the eye more distant; and if extremely so, the paper will appear uniformly bright. Now, if every alternate hole be stopped, the paper will manifestly receive only half the light, and will therefore be only half as much illuminated, and *cæteris paribus*, the degree of illumination is proportional to the number of the holes in the screen, or to the number of equally illuminated specks on its surface, *i. e.* if the specks be infinitely diminished in size, and infinitely increased in number, to the number of rays which fall on it from the original source of light.

19. Let a screen, so pierced with innumerable equal and very small holes in the manner described, be placed at a given distance (1 yard) from a candle; and in the diverging pyramid of rays behind it place a small piece of white paper of a given area, (1 square inch, for instance,) so as to be entirely included in the pyramid. It is manifest that the number of rays which fall on it will be fewer as it is placed farther from the screen, because the whole number which pass the screen are scattered continually over a larger and larger space. Thus were it close to the screen it would receive a number equal to that of the holes in a square inch of the screen, but at twice the distance (2 yards) from the candle this number will be spread over four square inches by their divergence, and the paper can therefore receive only a fourth part of that number. If, therefore, its illumination when close to the screen be represented by I, it will at twice the distance be only $\frac{I}{4}$, and

Logor,

at D times the assumed unit of distance, its illumination will be $\frac{I}{D^2}$, the areas of sections of a pyramid

being as the squares of their distances from the vertex.

20. As this reasoning is independent of the number and size of the holes, and therefore of the ratio of the sum of their areas to that of the unperforated part of the screen, we may conceive this ratio increased *ad infinitum*. The screen then disappears, and the paper is freely illuminated. Hence we conclude that when a small plane object of given area is freely and perpendicularly exposed to a luminary at different distances, the quantity of light it receives, or the degree of its illumination, is inversely as the squares of its distance from the luminary, *cæteris paribus*.

21. If a single candle be placed before a system of holes in a screen, as before, and the rays received on a screen at a given distance, (1,) the degree of illumination may be represented by a given quantity, I. Now if a second candle be placed immediately behind the other, and close to it, so as to shine through the same holes, the illumination of the screen is perceived to be increased, though the number and size of the illuminated points on it be the same. Each point is then said to be more intensely illuminated. Now, (the eye being all along supposed so distant, and the illuminated points so small as to produce only a general sense of brightness, without distinguishing the individual points,) if the one candle be shifted a little sideways, without altering its distance, the illumination of the paper will not be altered. In this case the number of illuminated points is doubled, but each is illuminated with only half the light it had before. The same holds for any number of candles. Hence we conclude that the illumination of a surface is constant when the number of rays it receives is inversely as the intensity of each, and that consequently the degree of illumination is proportional to the number and intensity of the rays jointly.

22. Now if for any number of candles placed side by side we substitute mere physical luminous points, each of these will be the vertex of a pyramid of rays, and the number of equally illuminated points in the paper, and therefore illuminations will be proportional to the number of such points. If we conceive the number of these increased, and their size diminished *ad infinitum*, so as to form a continuous luminous surface, their number will be represented by its area. Hence the illumination of the paper will be, *cæteris paribus*, as the area of the illuminating surface, (supposed of uniform brightness.)

23. Uniting all these circumstances, we see that when a given object is enlightened by a luminous surface of small but sensible size, the degree of its illumination is proportional to the

$$\frac{\text{area of the luminous surface} \times \text{intensity of its illuminating power}}{\text{square of the distance of the surface illuminated.}}$$

24. The foregoing reasoning applies only to the case when the luminous disc is a small portion of a spherical surface concentric with the illuminated object, in which case all its points are equidistant from it, and all the light falls perpendicularly on the object. When the object is exposed obliquely, conceive its surface divided into equal infinitely small portions, and regard each of them as the base of an oblique pyramid, having its vertex at any one point of the luminary; then will the perpendicular section of this pyramid at the same distance be equal to the base \times sine of inclination of the base to the axis, or the element of the illuminated surface \times by the sine of the inclination of the ray. But the number of rays which falls on the base is evidently equal to those which fall on the section, and being spread over a larger area their effect will be to illuminate it less intensely in the proportion of the area of the section to that of the base, *i. e.* in the proportion of the sine of inclination to radius. But the illumination of the section is equal to the

$$\frac{\text{area of the luminary} \times \text{intrinsic brightness}}{(\text{distance})^2} \quad (23)$$

therefore that of the elementary surface equals this fraction multiplied by the sine of the rays' inclination; or, calling A the area of the luminary, I its intrinsic brightness, D its distance, and θ the inclination of the ray to the illuminated surface $\frac{A \cdot I \sin \theta}{D^2}$ will represent the intensity of illumination

25. If L represent the absolute quantity of light emitted by the luminary in a given direction, which may be called its *absolute light*, we have $L = A \times I$, provided the surface of the luminary be perpendicular to the given direction. If not, A must represent the area of the section of a cylindroidal surface bounded by the outline of the luminary, and having its axis parallel to the given direction; consequently $\frac{L \cdot \sin \theta}{D^2}$ represents in this case the intensity of illumination of the elementary surface.

To illustrate the application of these principles we will resolve the following

PROBLEM.

26. A small white surface is laid horizontally on a table, and illuminated by a candle placed at a given (horizontal) distance: What ought to be the height of the flame, so as to give the greatest possible illumination to the surface?

Fig. 2

Let A be the surface, BC the candle. Put $AB = a$. $AC = D$; $BC = \sqrt{D^2 - a^2}$. Then, since the illumination of A is, *cæteris paribus*, as $\frac{\sin CAB}{AC^2}$, or as $\frac{CB}{AC^3} = \frac{\sqrt{D^2 - a^2}}{D^3}$ (= F) we have to make this

Light. quantity a maximum; consequently $dF = 0$, or $d \cdot F^2 = 0$, that is,

$$d \left\{ \frac{1}{D^4} - \frac{a^2}{D^6} \right\} = 0, \text{ or } -\frac{4}{D^5} + \frac{6a^2}{D^7} = 0, 2D^2 - 3a^2 = 0,$$

$$\text{or } D = a \cdot \sqrt{\frac{3}{2}} \text{ and } BC = \sqrt{D^2 - a^2} = \frac{a}{\sqrt{2}} = 0.707 \times AB.$$

27. *Definition.* The *apparent superficial magnitude*, or the apparent magnitude of any object, is a portion of a spherical surface described about the eye as a centre, with a radius equal to 1, and bounded by an outline being the intersection of this spherical surface with a conoidal surface, having the object for its base and the eye for its vertex.

28. Hence the apparent superficial magnitude of a small object is directly as the area of a section (perpendicular to the line of sight) of this conoidal surface, at the place of the object, and inversely as the square of its distance. If the object be a surface perpendicular to the line of sight, this ratio reduces itself to the area of the object divided by the square of its distance.

29. *Definition.* The *real intrinsic brightness* of a luminous object is the intensity of the light of each physical point in its surface, or the numerical measure of the degree in which such a point (of given magnitude) would illuminate a given object at a given distance, referred to some standard degree of illumination as a unit. When we speak simply of intrinsic brightness, *real intrinsic brightness* is meant.

30. *Corol. 1.* Consequently the degree of illumination of an object exposed perpendicularly to a luminary is as the apparent magnitude of the luminary and its intrinsic brightness jointly.

31. *Corol. 2.* Conversely, if these remain the same, the degree of illumination remains the same. For example, the illumination of direct sunshine is the same as would be produced by a circular portion of the surface of the sun of one inch in diameter, placed at about 10 feet from the illuminated object, and the rest of the sun annihilated; for such a circular portion would have the same apparent superficial magnitude as the sun itself. This will serve to give some idea of the intense brightness of the sun's disc.

32. *Definition.* The *apparent intrinsic brightness* of any object, or luminary, is the degree of illumination of its image or picture at the bottom of the eye. It is this illumination only by which we judge of brightness. A luminary may in reality be ever so much brighter than another; but if by any cause the illumination of its image in the eye be enfeebled, it will appear no brighter than in proportion to its diminished intensity. Thus we can gaze steadily at the sun through a dark glass, or the vapours of the horizon.

33. *Definition.* The *absolute light* of a luminary is the sum of the areas of its elementary portions, each multiplied by its own intrinsic brightness; or, if every part of the surface be equally bright, simply the area multiplied by the intrinsic brightness. It is, therefore, the same quantity as that above represented by L .

34. *Definition.* The *apparent light* of an object is the total quantity of light which enters our eyes from it, however distributed on the retina.

In common language, when we speak of the brightness of an object of considerable size, we often mean its *apparent intrinsic brightness*. When, however, the object has no sensible size, as a star, we always mean its *apparent light*, (or, as it might be termed, its *apparent absolute brightness*,) because, as we cannot distinguish such an object into parts, we can only be affected by its total light indiscriminately. The same holds good with all small objects which require attention to distinguish them into parts. Optical writers have occasionally fallen into much confusion for want of attending to these distinctions.

36. As we recede from a luminary, its *apparent light* diminishes, from two causes; first, our eyes, being of a given size, present a given area to its light, and therefore receive from it a quantity of light inversely as the square of the distance; secondly, in passing through the air, a portion of the light is stopped, and lost from its want of perfect transparency. This, however, we will not now consider. In virtue of the first cause only, then, the apparent light of a luminary is inversely as the square of its distance, and directly as its *absolute light*.

37. The apparent intrinsic brightness is equal to the apparent light divided by the area of the picture on the retina of our eye. But this area is as the apparent superficial magnitude of the luminary, that is, as its real area A divided by the square of its distance D , or as $\frac{A}{D^2}$. Moreover, the apparent light, as we have just seen, is as

$\frac{AI}{D^2}$ where I is the real intrinsic brightness. Consequently the apparent intrinsic brightness is proportional to

$\frac{AI}{D^2} + \frac{A}{D^2}$, or simply to I , and is independent on A or D . The apparent intrinsic brightness is, therefore,

the same at all distances, and is simply proportional to the real intrinsic brightness of the object. This conclusion is usually announced by optical writers by saying, that *objects appear equally bright at all distances*, which must be understood only of *apparent intrinsic brightness*, and the truth of which supposes also that no loss of light takes place in the media traversed.

38. *The angle of emanation* of a ray of light from a luminous surface is the inclination of the ray to the surface at the point from which it emanates.

A question has been agitated among optical philosophers, whether the intensity of the light of luminous bodies be the same in all directions; or whether, on the other hand, it be not dependent on the angle of emanation. Euler, in his *Réflexions sur les divers degrés de la lumière du Soleil*, &c. Berlin, Mém. 1750, p. 280, has adopted

In what sense to be understood.

Angle of emanation defined.

39.

Light. the former principle. Lambert, on the other hand, *Photometria*, p. 41, contends that the intensity of the light, or density of the rays, issuing from a luminous surface in any direction is proportional to the *sine* of the angle of emanation. If we knew the intimate nature of light, and the real mechanism by which bodies emit and reflect it, it might be possible to decide this question *à priori*. If, for instance, we were assured that light emanated strictly and solely from the molecules situated on the external surface of bodies, and that the emanation from each physical point of the surface were totally uninfluenced by the rest of the molecules of which the body consists, and dispersed itself equally in all directions, then, since every point of a plane surface is visible to an eye wherever situated above it, and each is supposed to send the same number of rays to the eye in an oblique as in a perpendicular situation, the *total light* received from a given area of the surface in the eye ought to be the same at all angles of emanation. But as the *apparent magnitude* of this area is as the sine of its inclination to the line of sight, *i. e.* of its angle of emanation, this light is distributed over a less *apparent* area; and therefore its intensity, or the apparent brightness of the surface, should be increased in the inverse ratio of the sine of the angle of emanation. On the other hand, if, as there is every reason to suppose, light emanates, not strictly from the surfaces of bodies, but from sensible depths within their substance; if the surfaces themselves be not true mathematical planes, but consist of a series of physical points retained in their places by attractive and repulsive forces, and if the intensity of emanation of each of these points depend in any way on its relation to the points adjacent, there is no reason, *à priori*, to suppose the equal emanation of light in all directions; and to find what its law really is, we must have recourse to direct observation.

Part II.

Sine

Question among opticians respecting the dependence of the emission of light on the angle of emanation.

Astronomy teaches us that the sun is a sphere. Hence the several parts of its visible disc appear to us under every possible angle of inclination. Now if we examine the surface of the sun with a telescope, the circumference certainly does not appear brighter than the centre. But if the hypothesis of equal emanation were correct, the brightness ought to increase from the centre outwards, and should become infinite at the edges, so that the disc ought to appear surrounded by an annulus of infinitely greater splendour than the central parts. To this it may, however, be justly objected, that as the surface of the sun is obviously though *generally* spherical, yet full of local irregularities, every minute portion of it may be regarded as presenting every possible variety of inclination to our eye; and the brightness of every part being thus an average of all the gradations of which it is susceptible, should be alike throughout.

40. Bouguer, in his *Traité d'Optique*, Paris, 1760, p. 90, states himself to have found, by direct comparison, that the central portions of the disc of the sun are really *much more* luminous than the borders. A result so extraordinary, however, and so apparently incompatible with all we know of the constitution of the sun and the mode of emission of light from its surface, would require to be verified by very careful and delicate reexamination. If found correct, the only way of accounting for it would be to suppose a dense and imperfectly transparent atmosphere of great extent floating above the luminous clouds which form its visible surface. This is certainly possible, but our ignorance on the subject renders it unphilosophical to resort to a body so little within our reach for the establishment of any fundamental law of emanation. The objection above advanced, it will be observed, applies with nearly the same force to all surfaces. If we examine a piece of white paper with a magnifier, we shall find its texture to be in the last degree rough and coarse, presenting no approach to a plane; and so of all surfaces rough enough to reflect light in all directions.

41. Surfaces appear equally bright at all angles. However, as it is only with such luminous surfaces as occur in nature that we have any concern, we must take their properties as we find them; and, waiving all consideration of what would be the law of emanation from a mathematical surface, it may be stated as a result of observation, that *luminous surfaces appear equally bright at all angles of inclination to the line of sight*.

This may be tried with a surface of red-hot iron; its apparent *intrinsic brightness* is not sensibly increased by inclining it obliquely to the eye.

42. Experimental proof of the law of emanation. If we take a smooth square bar of iron, or better, of silver, or a polished cylinder of either metal, heated to redness, into a dark room, the cylinder will appear equally bright in the middle of its convexity next the eye, and at the edges, and cannot be distinguished at all from a flat bar; and the square bar, when so presented as to have two of its sides at very different angles to the line of sight, will still appear of perfectly equable brightness, nor can the angle separating the adjacent sides be at all discerned; and if the whole bar be turned round on its axis, the motion can only be recognised by an alternating increase and decrease of its apparent diameter, according as it is seen alternately diagonally and laterally, its appearance being always that of a flat plate perpendicularly exposed to the eye. These and similar experiments with surfaces artificially illuminated, which the reader will have no difficulty in imagining and making, as well as those recorded by Mr. Ritchie in the *Edinburgh Philosophical Journal*, are sufficient to establish the principle announced in Article 42, to which (for the reasons already mentioned) the observation of Bouguer on the unequal brightness of the sun's disc offers no conclusive objection.

43. Law of the oblique emanation of light. Hence it follows, that the surfaces of luminous bodies, at least their ultimate molecules, do not emit light with equal copiousness in all directions; but that, on the contrary, *the copiousness of emission, in any direction, is as the sine of the angle of emanation from the surface*.

PROBLEM.

44. To determine the intensity of illumination of a small plane surface any how exposed to the rays from a luminary of any given size, figure, and distance; the luminary being supposed uniformly bright in every part. Conceive the surface of the luminary divided into infinitesimal elementary portions, of which let each be regarded as an oblique section of a pyramid, having for its vertex the centre of the infinitely small illuminated

Light. plane B, fig. 3. Let PQ be any such portion, and let the pyramid BP be continued till it meets the surface of the heavens in *p*, there projecting the surface PQ into the areola *pq*, and let the whole luminary CDEF be in like manner projected into the disc *cdef*. Let πQ be a section of the pyramid APQ, perpendicular to its axis. Then, first, the plane B will be illuminated by the element PQ, just as it would be by a surface πQ equally bright, in virtue of the principle just established. Hence PQ is equivalent to an equally bright surface πQ . Again, since the apparent magnitude of πQ seen from B is the same with that of *pq*, the area πQ is equivalent to an equally bright area *pq* placed at *pq*, (Art. 29, 30, 31, Cor. 1, 2.) PQ is, therefore, equivalent to *pq*. And since the same holds good of every other elementary portion of the surface, and the total light received by B is the sum of the lights it receives from all the elements of the luminary, the whole surface CDEF must be equivalent to its projection *cdef*.

Illumination of a plane by any luminary investigated. Fig. 3.

45. Hence the illumination of B depends, not at all on the real, but only on the apparent figure and magnitude of the luminary; and whatever the luminary be, we may always substitute for it a portion of the visible heavens, supposed of equal intrinsic brightness, and bounded by the same outline.

46. Thus, instead of the sun, we may suppose a small circle equal in apparent diameter to the sun, and equally bright; instead of a luminous rectangle perpendicular to the illuminated plane B, and of infinite height, as AGHI, fig. 3, we may substitute the spherical sector ZAG, bounded by the two vertical circles ZA, ZG, and so on.

47. Let then *pq*, any elementary rectangle infinitely small in both dimensions of the spherical surface, be represented by $d^2 A$, so that $\iint d^2 A$ shall represent the surface *cdef* itself; then if we put *z* = the zenith distance Z*p* of this portion, its illuminating power on A will be $d^2 A \cdot \cos z$, and the total illuminating power of the whole surface A will be

$$L = \iint d^2 A \cdot \cos z.$$

48. *Example 1.* To find the illuminating power of the sector ZAG confined between any two vertical circles and the horizon. (fig. 3.) Here, putting θ for the azimuth of the element $d^2 A$, if we consider it as terminated by two contiguous verticals and two contiguous parallels of altitude, we have $d^2 A = dz \times d\theta \cdot \sin z$. Hence we have

General formula for illumination of a small plane.

$$L = \iint d\theta dz \cdot \sin z \cdot \cos z = \frac{1}{2} \iint d\theta dz \cdot \sin 2z = \frac{1}{2} \int (\theta + C) dz \cdot \sin 2z;$$

and extending the integral from $\theta = 0$ to $\theta = AG$, the amplitude of the sector, (which we will call *a*), we get

$$L = \frac{a}{2} \int dz \cdot \sin 2z = \frac{a}{2} (C - \frac{1}{2} \cos 2z)$$

which extended from $Z = 0$ to $z = 90^\circ$ becomes simply $L = \frac{a}{2}$.

49. *Corol. 1.* This is a measure of the illuminating power of the sector, on the same scale that that of an infinitely small area (A) placed at the zenith would be represented by A itself. Because in this case

$$\cos z = 1, \text{ and } \iint d^2 A \cdot \cos z = A.$$

50. *Corol. 2.* On the same scale the illuminating power of the whole hemisphere is π where $\pi = 3.14159535$

51. *Example 2.* Required the illuminating power of a circular portion of the heavens whose centre is the zenith. Calling *z* the zenith distance of any element, and θ its azimuth, we shall have, as before,

$$d^2 A = d\theta dz \cdot \sin z, \text{ and therefore } L = \iint d\theta dz \cdot \sin z \cdot \cos z = \int \theta \cdot \frac{dz \cdot \sin 2z}{2} = \pi \int dz \cdot \sin 2z$$

extending the integral from $\theta = 0$ to $\theta = 2\pi$. That is $L = \pi \cdot (\text{const} - \frac{1}{2} \cos 2z)$ which being made to vanish when $z = 0$ becomes

$$L = \frac{\pi}{2} (1 - \cos 2z) = \pi \cdot (\sin z)^2$$

52. *Corol. 3.* The illuminating power of a circular luminary, whose centre is in the zenith, is as the square of the sine of its apparent semidiameter.

53. *Example 3.* Required the illuminating power of any circular portion of the heavens whatever.

Illuminating power of any circular portion of an equally bright even.

Let TKLM be the illuminating circle; conceive it composed of annuli concentric with itself, and of one of them, XYZ, (fig. 4,) let X*x* be an infinitesimal parallelogram terminated by contiguous radii SX and S*x*, S being the centre.

$$\text{Put } ZS = a; SX = x, ZX = z,$$

$$\text{Angle } ZSX = \phi, ST = r.$$

$$\text{Area } d^2 A = Xx, = dx \times d\phi \cdot \sin x$$

$$\therefore L = \iint d\phi dx \cdot \sin x \cdot \cos z.$$

but, by spherical trigonometry, $\cos z = \cos a \cdot \cos x + \sin a \cdot \sin x \cos \phi.$

Light.

Part I.

Therefore

$$L = \iint dx . d\phi . \sin x \left\{ \cos a . \cos x + \sin a . \sin x . \cos \phi . \right\}$$

The first integration performed relative to ϕ , and extended from $\phi = 0$ to $\phi = 360^\circ$, or 2π , gives

$$L = \int dx . \sin x \times 2\pi . \cos a . \cos x.$$

After which integrating, with respect to x , and extending the integral from $x = 0$ to $x = ST = r$, we find

$$L = \frac{\pi . \cos a}{2} (1 - \cos 2r) = \pi . \cos a (\sin r)^2.$$

This result is particularly elegant and remarkable. It shows, that to obtain the illuminating effect of a circular luminary (of any apparent diameter) at any altitude, on a horizontal plane, we have only to reduce its illuminating effect when in the zenith, in the ratio of radius to the cosine of the zenith distance, or sine of the altitude. For other examples, the reader may consult Lambert's *Photometria*, cap. ii. from which this is taken.

54. If the illuminating surface be not equally intrinsically bright in every part, if we call I the intrinsic brightness of the element $d^2 A$, we shall have

$$L = \iint I d^2 A . \cos z$$

General expression for illumination when the luminary is not equally bright throughout. for the general formula expressing the illuminating power of the surface A . The moon, Venus and Mercury in their phases, the sky during twilight, a white sphere illuminated by the sun, &c. afford examples of this when themselves regarded as luminaries.

PROBLEM.

55. To compare the illumination of a horizontal plane by the sun in the zenith with the illumination it would have were the whole surface of the heavens of equal brightness with the sun.

By Art. 53 we have $L = \pi . \cos a . (\sin r)^2$. If, therefore, we call L and L' the two illuminations in question, we shall have

$$L : L' :: \pi . \cos 0^\circ . (\sin \odot \text{'s semidiam.})^2 : \pi . \cos 0^\circ . (\sin 90^\circ)^2 \\ :: (\sin 16')^2 : 1 :: 1 : 46166.$$

56. Illumination at the sun's surface. The illumination of a plane in contact with the sun's surface is the same as that of a plane on the earth's surface illuminated by a whole hemisphere of equal brightness with the sun in the zenith. Hence we see that the illumination of such a plane at the sun's surface would be nearly 50,000 times greater than that of the earth's surface at noon under the equator. Such would be the effect (in point of light alone) of bringing the earth's surface in contact with the sun's!

57. Photometers For measuring the intensity of any given light, various instruments called Photometers have been contrived, many of which have little to recommend them on the score of exactness, and some are essentially defective in principle, being adapted to measure—not the illuminating—but the heating power of the rays of light; and, therefore, must be regarded as undeserving the name of photometers.

58. The eye as imperfect judge of degrees of illumination. We know of no instrument, no contrivance, as yet, by which light alone (as such) can be made to produce mechanical motion, so as to mark a point upon a scale, or in any way to give a direct reading off of its intensity, or quantity, at any moment. This obliges us to refer all our estimations of the degrees of brightness at once to our organs of vision, and to judge of their amount by the impression they produce immediately on our sense of sight. But the eye, though sensible to an astonishing range of different degrees of illumination, is (partly on that very account) but little capable of judging of their relative strength, or even of recognising their identity when presented at intervals of time, especially at distant intervals. In this manner the judgment of the eye is as little to be depended on for a measure of light, as that of the hand would be for the weight of a body casually presented. This uncertainty, too, is increased by the nature of the organ itself, which is in a constant state of fluctuation; the opening of the pupil, which admits the light, being continually expanding and contracting by the stimulus of the light itself, and the sensibility of the nerves which feel the impression varying at every instant. Let any one call to mind the blinding and overpowering effect of a flash of lightning in a dark night, compared with the sensation an equally vivid flash produces in full daylight. In the one case the eye is painfully affected, and the violent agitation into which the nerves of the retina are thrown is sensible for many seconds afterwards, in a series of imaginary alternations of light and darkness. By day no such effect is produced, and we trace the course of the flash, and the zig-zags of its motion with perfect distinctness and tranquillity, and without any of those ideas of overpowering intensity which previous and subsequent total darkness attach to it.

59. But yet more. When two unequally illuminated objects (surfaces of white paper, for instance) are presented at once to the sight, though we pronounce immediately on the existence of a difference, and see that one is brighter than the other, we are quite unable to say what is the proportion between them. Illuminate half a sheet of paper by the light of one candle, and the other half by that of several; the difference will be evident. But if ten different persons are desired, from their appearance only, to guess at the number of candles shining on each, the probability is that no two will agree. Nay, even the same person at different times will form different judgments. This throws additional difficulty in the way of photometrical estimations, and would seem to render this one of the most delicate and difficult departments of optics.

Light However, the eye, under favourable circumstances, is a tolerably exact judge of the *equality* of two degrees of illumination seen at once; and availing ourselves of this, we may by proper management obtain correct information as to the relative intensities of all lights. What these favourable circumstances are, we come now to consider.

The eye capable of judging of the equality of two degrees of illumination, and under what circumstances. Axiom in photometry.

60. 1st. The degrees of illumination compared must be of moderate intensity. If so bright as to dazzle, or so faint as to strain the eye, no correct judgment can be formed.

Hence, it is rarely adviseable to compare two luminaries directly with each other. It is generally better to let them shine on a smooth white surface, and judge of the degree in which they illuminate it; for it is an obvious axiom, *That two luminaries are equal in absolute light when, being placed at equal distances from, and in similar situations with respect to, a given smooth white surface, or two equal and precisely similar white surfaces, they illuminate it or them equally.*

2nd. The luminaries, or illuminated surfaces compared, must be of equal apparent magnitude, and similar figure, and of such small dimensions as to allow of the illumination in every part of each being regarded as the same.

63. 3rd. They must be brought close together, in apparent contact; the boundary of one cutting upon that of the other by a well-defined straight line.

64. 4th. They should be viewed at once by the same eye, and not one by one eye, and the other by the other.

65. 5th. All other light but that of the two objects whose illumination is compared should be most carefully excluded.

66. 6th. The lights which illuminate both surfaces must be of the same colour. Between very differently coloured illuminations no exact equalization can ever be obtained, and in proportion as they differ our judgment is uncertain.

67. When all these conditions obtain, we can pronounce very certainly on the equality or inequality of two illuminations. When the limit between them cannot be perceived, on passing the eye backwards and forwards across it, we may be sure that their lights are equal.

68. Bouguer, in his *Traité d'Optique*, 1760, p. 35, has applied these principles to the measure or rather the comparison of different degrees of illumination. Two surfaces of white paper, of exactly equal size and reflective power, (cut from the same piece in contact,) are illuminated, the one by the light whose illuminating power is to be measured, the other by a light whose intensity can be varied at pleasure by an increase of distance, and can therefore be exactly estimated. The variable light is to be removed, or approached, till the two surfaces are judged to be equally bright, when, the distances of the luminaries being measured, or otherwise allowed for, the measure required is ascertained.

69. Mr. Ritchie has lately made a very elegant and simple application of this principle. His photometer consists of a rectangular box, about an inch and a half or two inches square, open at both ends, of which A B C D (fig. 5) is a section. It is blackened within, to absorb extraneous light. Within, inclined at angles of 45° to its axis, are placed two rectangular pieces of plane looking-glass F C, F D, cut from one and the same rectangular strip, of twice the length of either, to ensure the exact equality of their reflecting powers, and fastened so as to meet at F, in the middle of a narrow slit E F G about an inch long, and an eighth of an inch broad, which is covered with a slip of fine tissue or oiled paper. The rectangular slit should have a slip of blackened card at F, to prevent the lights reflected from the looking-glasses mingling with each other.

70. Suppose we would compare the illuminating powers of two sources of light (two flames, for instance) P and Q. They must be placed at such a distance from each other, and from the instrument between them, that the light from every part of each shall fall on the reflector next it, and be reflected to the corresponding portion of the paper E F or F G. The instrument is then to be moved nearer to the one or the other, till the paper on either side of the division F appears equally illuminated. To judge of this, it should be viewed through a prismoidal box blackened within, one end resting on the upper part A B of the photometer; the other applied quite close to the eye. When the lights are thus exactly equalized, it is clear that the total illuminating powers of the luminaries are directly as the squares of their distances from the middle of the instrument.

71. By means of this instrument we are furnished with an easy experimental proof of the decrease of light as the inverse squares of the distances. For if we place four candles at P, and one at Q, (as nearly equal as possible, and burning with equal flames,) it is found that the portions E F, G F of the paper will be equally illuminated when the distances P F, Q F are as 2 : 1, and so for any number of candles at each side.

72. To render the comparison of the lights more exact, the equalization of the lights should be performed several times, turning the instrument end for end each time. The mean of the several determinations will then be very near the truth.

73. In some cases the looking-glasses are better dispensed with, and a slip of paper pasted over them, so as to present two oblique surfaces of white paper inclined at equal angles to the incident light. In this case the paper stretched over the slit E F G is taken away, and the white surfaces below examined and compared. One advantage of this disposition is the avoiding of a black interval between the two halves of the slit, which renders the exact comparison of their illuminations somewhat precarious.

74. If the lights compared be of different colours (as daylight, or moonlight, and candlelight,) their precise equalization is impracticable, (art. 67.) The best way of employing the instrument, in this case, is to move it till one of the sides of the slit (in spite of the difference of colours) is judged to be decidedly the brighter, and then to move it the other way, till the other becomes decidedly the brighter. The position half way between these points is to be taken as the true point of equal illumination.

75. If we would compare the degrees of illumination, or the intrinsic brightnesses of two surfaces, a given portion of each must be insulated for examination; this may be best done by the adaptation of two blackened tubes to

Light. the openings of the photometer, of equal length, and terminated by orifices of equal area, or subtending equal angles at the middle of the instrument. These, of course, cut off equal apparent magnitudes of the bright surfaces, the light of which is then to be equalized on the oiled papers of the slit E F, as in the case of candles, &c. Bouguer, *Traité*, p. 31.

Comparison of degrees of brightness of illuminated surfaces. Another method of comparing the intensity of the light from two luminaries, which is also very ready and convenient, and possesses in some cases considerable advantages, has been proposed by Count Rumford. (See *Phil. Trans.*, vol. 84, p. 67.) It consists in the equalization of the shadows cast by them on a white surface illuminated by them both at once. Suppose, for instance, we would compare the illuminating power of two flames L and l of different sizes, or from different combustibles, as of wax and tallow. Before a screen C D of white paper, in a darkened room, place a blackened cylindrical stick S, and let the flames L l be so placed as to throw the shadows A B of the stick on the screen, side by side, and with an interval between them about equal in breadth to either shadow. Moreover the inclination of the rays L S A and l S B to the surface of the screen must be adjusted to exact equality. The brighter flame must then be removed, or the feebler brought nearer to the screen, till the two shadows appear of equal intensity, when their distances (or the distance of the screen) from the lights must be measured, and their total illuminating powers will be in the direct ratio of the squares of the distances. The rationale of this is obvious, the shadow thrown by each flame is illuminated by the light of the other. The screen by the sum of the lights. The eye in this case judges of the degrees of defalcation of brightness from this sum; and if these degrees be alike, it is clear that the remaining illuminations must be equal.

77. Fig. 8. This method becomes uncertain when the lights are of considerable size and near the screen, as the penumbrae of the shadows prevent any fair comparison of the relative intensities of their central portions. It is still more so, and can hardly be used when the lights differ considerably in colour. Its convenience, however, as an extemporaneous method, requiring no apparatus but what is always at hand, (as the use of a blackened stick, though preferable, is not essential,) renders it often useful in the absence of more refined means.

78. It may happen that the lights to be compared are not movable, or not conveniently so. In this case the equalization of the shadows may be performed by inclining the screen at different angles to the directions in which it receives the light of each, and noting the angles of inclination of the rays. In this case the illuminating powers of the luminaries are as the squares of their distances directly, and the sines of the respective angles of inclination of their rays to the screen inversely.

79. When the lights to be compared are immovable. When a ray of light proceeds in empty space, or in a perfectly homogeneous medium, its course, as we have seen, is rectilinear, and its velocity uniform; but when it encounters an obstacle, or a different medium, it undergoes changes or modifications which may be stated as follows:

80. Modifications of light enumerated. It is separated into several parts, which pursue different courses, or are otherwise differently modified. One of these parts is regularly reflected, and pursues, after reflexion, a course wholly exterior to the new medium, or obstacle.

81. Regular reflexion. A second and a third portion are regularly refracted, that is, they enter the medium, and there pursue their course according to the laws of refraction. In many media these portions follow the same course precisely, and perhaps are no way distinguishable from each other. In such media (comprehending most uncrystallized substances and liquids) the refraction is said to be single. In numerous others (for instance in most crystallized media) they follow different courses, and also retain different physical characters. In these the refraction is said to be double.

82. Regular refraction. A fourth portion is scattered in all directions, one part being intromitted into the medium, and distributed over the hemisphere interior to it, while the other is in like manner scattered over the exterior hemisphere. These two portions are those which render visible the surfaces of bodies to eyes situated any how with respect to them, and are therefore of the utmost importance to vision.

83. Scattering. Of those portions which enter the medium, a part more or less considerable is absorbed, stifled, or lost, without any further change of direction; and that not at once, but progressively, as they penetrate deeper and deeper into its substance. In perfectly opaque media, such as the metals, this absorption is total, and takes place within a space less than we can appreciate; yet even here we have good reasons for believing that it does not take place *per saltum*. In crystallized bodies, those at least which are coloured, this absorption takes place differently on the two portions into which the regularly refracted ray is divided, according to laws to be explained when we come to treat of the absorption of light.

84. Absorption. The regularly refracted portions of a ray of white or solar light are (except in peculiar circumstances) separated into a multitude of rays of different colours, and otherwise differing in their physical properties, each of which rays pursues its course afterwards, independently of all the rest, according to the laws of regular refraction or reflexion. The laws of this separation, or dispersion, of the coloured rays, and their physical and sensible properties, form the subject of *Chromatics*.

85. Separation into colours, or dispersion. All those portions which are either regularly reflected, or regularly refracted, undergo, more or less, a modification termed *polarization*, in virtue of which they present, on their encountering another medium, different phenomena of reflexion and refraction from those presented by unpolarized light. Generally speaking, polarized light obeys the same laws of reflexion and refraction as unpolarized, as to the directions which the several portions, into which it is divided on encountering a new medium, take; but differs from it in the relative intensities of those portions, which vary according to the situation in which the surface of the medium and certain imaginary lines, or axes within it, are presented to the polarized ray.

86. Polarization. The rays of light under certain circumstances exercise a mutual influence on each other, increasing, diminishing, or modifying each other's effects according to peculiar laws. This mutual influence is called the *interference* of the rays of light. We shall proceed to treat of these several modifications in order; and first of the regular reflexion of light.

87. Interference. The rays of light under certain circumstances exercise a mutual influence on each other, increasing, diminishing, or modifying each other's effects according to peculiar laws. This mutual influence is called the *interference* of the rays of light. We shall proceed to treat of these several modifications in order; and first of the regular reflexion of light.

§ 3. Of the regular Reflexion of unpolarized Light from Plane Surfaces.

88. When a ray of light is incident on a smooth-polished surface, a portion of it is regularly reflected, and pursues its course after reflexion in a right line wholly exterior to the reflecting medium. The direction and intensity of this portion are the objects of inquiry in this section; the physical properties acquired by the ray in the act of reflexion being reserved for examination at a more advanced period. And first, with regard to the direction of the reflected ray. This is determined by the following laws:

Laws of Reflexion.

89. *Law 1.* When the reflecting surface is a plane. At the point on which the ray is incident raise a perpendicular. The reflected ray will lie in the same plane with this perpendicular, and with the incident ray. It will lie on the opposite side of the perpendicular, and will make an angle with it equal to that made by the incident ray.
90. The plane in which the perpendicular to any surface at the point of incidence, and the incident ray, both lie, is called the *plane of incidence*.
91. The angle included between the incident ray and the perpendicular is called the *angle of incidence*.
92. The plane in which the reflected ray and perpendicular both lie is called the *plane of reflexion*; and the angle included between the perpendicular and reflected ray is, in like manner, termed the *angle of reflexion*.
93. Adopting these definitions, the law of reflexion from a plane surface may be announced by saying, that the plane of reflexion is the same with that of incidence, and the angle of reflexion equal to that of incidence, but situated on the contrary side of the perpendicular.
94. *Corol.* The incident and reflected rays are equally inclined to the surface at the point of incidence.
- Law 2.* When the surface is a curved one, the course of a ray reflected from any point is the same as if it were reflected at the same point from a plane, a tangent to the curve surface at that point; *i. e.* if a perpendicular be raised to the curve surface at the point of incidence, the reflected ray will lie in the plane of incidence, and the angle of reflexion will equal that of incidence.
95. The demonstration of these laws is a matter of experiment. If we admit a small sunbeam through a hole in the shutter of a darkened chamber, and receive it on a polished surface of glass, or metal, we may easily with proper instruments measure the inclinations of the incident and reflected rays to the surface, which will be found equal. But this method is rude and coarse. A much more delicate verification of this law is afforded by astronomical observations. It is the practice of astronomers to observe the altitudes of the stars above the horizon by direct vision; and, at the same instant, the apparent depression below the horizon of their images reflected at the surface of Mercury, (which is necessarily exactly horizontal,) and the depression so observed is always found precisely equal to the altitude, whatever the latter may be, whether great or small. Now as these observations, when made with large instruments, are susceptible of almost mathematical accuracy, we may regard the law of reflexion, or plane surfaces, as the best established in nature.
96. Reflexion at a curved surface may be considered as taking place at that infinitely small portion of the surface which is common to it, and to its tangent plane at the point of incidence; so that if a perpendicular to the surface be erected at the point of incidence, the incident and reflected rays will make equal angles with it on opposite sides.
97. *Proposition.* To find the direction of a ray of light after reflexion at any number of plane surfaces, given in position.
- Construction.* Since the *direction* of the ray after reflexion is the same whether it be reflected at the given surfaces, or at surfaces parallel to them, conceive surfaces parallel to the given ones to pass through any point C, (fig. 9,) and from C draw the straight lines CP, CP', CP'', &c. respectively perpendicular to these respective surfaces, and lying wholly exterior to the reflecting media. Draw SC parallel to the ray when incident on the first surface, and in the plane SCP, and on the opposite side of CP, from the incident ray SC make the angle PCS' = PCS, then will CS' be the direction of the ray after reflexion at the first surface. Prolong s'C to S', then S'C will represent the ray at the moment of its incidence on the second surface, whose normal is CP'. Again, make the angle P'C s'' in the plane S'CP', but on the other side of CP', equal to the angle S'CP', then will Cs'' represent the ray at the moment of its reflexion from the second surface, and, producing s''C to S'', S''C will represent it at the moment of its incidence on the third surface, whose normal is CP''. Similarly in the plane S''CP''; but on the other side of CP'' make the angle P''C s''' = P''CS'', and Cs''' will be the direction of the ray at the moment of its quitting the third surface, and so on.
98. *Analysis.* About C as a centre conceive a spherical surface described, (fig. 10,) then will the plane PSs intersect it in a great circle PSS'p, and the plane in which CP, CP' lie, or the plane at right angles to the two first reflecting planes in another great circle P P' p, and the planes S' C s'' and S C s' in other great circles S' P' s'' and S k s'.
- Since CP and CP' are given directions, the angle PCP', or the arc PP' (which is equal to the *inclination of the two first surfaces to each other*) is given. Call this I. Again, since the direction SC of the incident ray is given, the *angle of incidence*, or the *first surface* PCS (= α) and the angle SPP', or the *inclination of the plane of the first reflection to the plane PP' perpendicular to both surfaces* (= ψ) are given. Hence in

Fig. 9.

Fig. 10

Light. the spherical triangle $PP'S'$ we have $PP' = I$; $PS' = 180^\circ - \alpha$; and the angle $P'PS' = \psi$; consequently $S'P'$, and therefore $2S'P' = S's''$ and the angle $SS'P'$ are known, as also the angle $PP'S'$, and therefore its supplement $PP's''$, which is the angle made by the second reflexion with the plane PP' . Again, in the spherical triangle $SS's''$ we have given $SS' = 180^\circ - 2\alpha$; $S's'' = 2S'P'$ and the included angle $SS's''$, whence the third side Ss'' may be found, which is the angle between the incident and twice reflected rays.

Similarly, if a third reflexion be supposed, we have given $P'S'' = 180^\circ - S'P'$; $P'P'' = I'$, and the angle $S''P'P'' = S'P'P'' = PP'P'' - P'P'S'$, whence we may compute $S''P''$ and proceed as before, and so on to any extent.

Confining ourselves however to the case of two reflexions we have, by spherical trigonometry, putting $P'S' = a'$ = the angle of incidence on the second reflecting surface, $PS'P' = \theta$; $PP'S' = \phi$, and $180^\circ - SS' = D$, the deviation of the ray after the second reflexion, the following equations: 99.

$$\left. \begin{aligned} -\cos a' &= \cos a \cdot \cos I - \sin a \cdot \sin I \cdot \cos \psi \\ \sin \theta &= \frac{\sin I}{\sin a'} \cdot \sin \psi \\ \sin \phi &= \frac{\sin a}{\sin a'} \cdot \sin \psi \\ \cos D &= \cos 2a \cdot \cos 2a' - \sin 2a \cdot \sin 2a' \cdot \cos \theta \end{aligned} \right\} \quad (A)$$

General equations of reflexion at two planes.

From these equations, any three of the seven quantities $a, a', I, \theta, \phi, \psi, D$ being given, the other four may be found. It will be observed, that ϕ is the angle between the plane of the second reflexion and the principal section of the two reflecting planes, and θ the angle between the planes of the first and second reflexion. If ϕ and D only be sought, θ must be regarded as merely an auxiliary angle; but this may not be the case, and cases may occur in which θ alone may be sought, or in which it enters as a given quantity, &c. In short, the foregoing equations contain in themselves all the conditions which can arise in any proposed case of two reflexions. 100.

Values of the symbols

Corol. If $\psi = 0$, or if the incident ray coincide with the principal section PCP , i. e. if the two reflexions both take place in the plane perpendicular to the reflecting surfaces, these formulæ take a very simple form, for we then have 101.

$$\theta = 0; \phi = 180^\circ; \cos a' = -\cos(a + I)$$

that is $(a + a') = 180^\circ - I$; and consequently $\cos(2a + 2a') = \cos(360^\circ - 2I) = \cos 2I$, or $2a + 2a' = 2I$. But since $\theta = 0$, we have by the last of the equations (A) $\cos D = \cos 2(a + a')$; consequently $D = 2a + 2a' = 2I$. That is to say, the deviation in this case, after two reflexions, is equal to twice the inclination of the reflecting planes, whatever be the original direction of the ray. This elegant property is the foundation of the common sextant and of the reflecting circle, and is commonly regarded as having been first applied to the measurement of angles by Hadley, though Newton appears also to have proposed it for the same object. See the explanation of these instruments. Case where both reflexions are in one plane. 102.

In other cases, however, D , the deviation, is essentially a function of the angles expressing the position of the incident ray, and can only be obtained from the equations above stated. 103.

Proposition. Given the angles of incidence on the two planes, and the angle made by the plane of the first reflexion with that of the second; required the positions of the incident and twice reflected rays, the deviation of the ray after both reflexions, and the angle included between the reflecting surfaces. 103.

Retaining the same notation, we have given, a, a', θ , required I, D , and ϕ, ψ .

1st, D is given at once, by the last of the general equations, (A.)

2ndly, To find the rest, put $x = \sin I$; $y = \sin \psi$; and $a = \sin a' \cdot \sin \theta$; put also $\cos a = c$; $\sin a = s$;

$\cos a' = c'$; $\sin a' = s'$. We have then $xy = a$, or $y = \frac{a}{x}$; and the first of the equations (A) then gives

$$-c' = c \sqrt{1 - x^2} - s \sqrt{x^2 - a^2}$$

which, cleared of radicals and reduced, gives

$$0 = x^4 + x^2 \{ 2c'^2 (c^2 - s^2) - 2c^2 - 2s^2 a^2 \} + (c'^2 - c^2)^2 + 2a^2 s^2 (c'^2 + c^2) + a^4 s^4$$

and this equation, which, though biquadratic, is of a quadratic form, contains the general solution of the problem.

Corol. 1. If $\theta = 90^\circ$, or if the planes of the first and second reflexions be at right angles to each other, we have simply $\sin I \cdot \sin \psi = \sin a'$, and $a = \sin a' = s'$. 104.

Case when the planes of the two reflexions are at right angles.

In this case our final equation becomes

$$0 = x^4 - 2x^2(1 - c^2 c'^2) + (1 - c^2 c'^2)^2$$

which, being a complete square, gives

$$x^2 = 1 - c^2 c'^2$$

Now $x = \sin I$, therefore $x^2 = 1 - \cos I^2$, consequently we have the following simple result,

$$\cos I (= c c') = \cos a \cdot \cos a'.$$

Light. Or the cosine of the inclination of the planes to each other is equal to the product of the cosines of the angles of incidence on each. And, *vice versa*, if this relation holds good, the planes of the two reflexions will necessarily be at right angles to each other; for, this relation being supposed, we have of course $x^2 = 1 - c^2 c'^2$, and therefore $1 - c^2 c'^2$ being put for x^2 in the general equation, the whole must vanish; now this substitution gives a biquadratic of a quadratic form for determining α , which must evidently be satisfied by taking

$$\alpha = \sin \alpha', \text{ and consequently } \theta = 90^\circ.$$

This elegant property will be useful when we come to treat of the polarization of light.

105. *Corol. 2.* In the same case if $\theta = 90$, the deviation D is given by the equation

$$\cos D = \cos 2\alpha \cdot \cos 2\alpha',$$

or, the cosine of the deviation is equal to the product of the cosines of the doubles of the angles of incidence.

106. *Problem.* A ray of light is reflected from each of two planes in such a manner that all the angles of incidence and reflexion are equal. Given the inclination of the planes, and the angles of incidence; required, first, the deviation; secondly, the inclination of the planes of the first and second reflexion to each other, and the angles made by each of these planes with the principal section of the reflecting planes.

Preserving the same notation we have $\alpha = \alpha'$, and therefore by the third of the equations (A) $\psi = \phi$, so that these equations become

$$\left. \begin{aligned} \cos \alpha (1 + \cos I) &= \sin \alpha \cdot \sin I \cdot \cos \psi \\ \sin \alpha \cdot \sin \theta &= \sin I \cdot \sin \psi \\ \cos D &= (\cos 2\alpha)^2 - (\sin 2\alpha)^2 \cdot \cos \theta \end{aligned} \right\} \quad (a)$$

107. The first of these gives (putting for $1 + \cos I$ its value $2 \left(\cos \frac{I}{2} \right)^2$ and for $\sin I$ its equal $2 \cdot \sin \frac{I}{2} \cdot \cos \frac{I}{2}$)

$$\cos \psi = \cotan \alpha \cdot \cotan \frac{I}{2}, \quad (b)$$

whence ψ is immediately known. Hence ψ is had by the equation

$$\sin \theta = \frac{\sin I}{\sin \alpha} \cdot \sin \psi. \quad (c)$$

Lastly, if we subtract each member of the third of the equations (a) from 1, divide both sides by 2, and reduce, we transform it into the following

$$\sin \frac{D}{2} = \sin 2\alpha \cdot \cos \frac{\theta}{2}. \quad (d)$$

These equations afford ready and direct means of computing ψ , θ , and D in succession, from the known values of α and I; the formulæ are adapted to logarithmic evaluation, and are in themselves not inelegant.

§ IV. Of Reflexion from Curved Surfaces.

108. The reflexion of a ray from a curved surface is performed as if it took place at a reflecting plane, a tangent to the point of incidence. The reflected ray will therefore lie in the plane which contains the incident ray and the normal or perpendicular at the point of incidence. The general expressions for the course of the ray after reflexion at surfaces of double curvature being considerably complex, and not likely to be of great service to us in the sequel, we shall confine ourselves to the particular case of a surface of revolution (comprehending the cases of a plane, and conoidal surfaces of all kinds) where the plane of incidence is supposed to pass through the axis of revolution.

109. *Proposition.* A ray being incident on any surface of revolution in a plane passing through the axis, to find the direction of the reflected ray.

General investigation of the course of a ray reflected at any curve. Fig. 11. QP (fig. 11) being a section of the surface by the plane of incidence, QN the axis, QP the incident, and Pr the reflected ray, which produced if necessary cuts the axis in q. Draw the tangent PT, the ordinate PM, and the normal PN, which produce to O, and put as follows,

$$x = QM; y = MP; p = \frac{dy}{dx}; \theta = \text{the angle MQP,}$$

or the angle made by the incident ray with the axis; then, since the angle of reflexion is equal to that of incidence, we have $\angle rPO = OPQ$, and therefore $NPq = OPQ$; consequently $QPT = TPq$. Now, $Qq = QM - Mq = QM - PM \cdot \tan MPq$

Light.

Part 1.

$$\begin{aligned} &= x - y \cdot \tan \{ \text{TPM} - \text{TPQ} \} \\ &= x - y \cdot \tan \{ \text{TPM} - \text{TPQ} \} \\ &= x - y \cdot \tan \{ \text{TPM} - \text{PTM} + \text{PQM} \} \\ &= x - y \cdot \tan \{ 90^\circ - 2 \text{PTM} + \text{PQM} \} \end{aligned}$$

But by the theory of curves we have $\tan \text{PTM} = \frac{dy}{dx} = p$, consequently $\text{PTM} = \arctan p = \tan^{-1} p$, denoting by \tan^{-1} the inverse function of that expressed by \tan ; and since $\text{PQM} = \theta$, this expression becomes

$$\begin{aligned} \text{Qq} &= x - y \cdot \cotan \{ 2 \cdot \tan^{-1} p - \theta \} \\ &= x - y \cdot \cotan \left\{ 2 \cdot \tan^{-1} \left(\frac{dy}{dx} \right) - \tan^{-1} \left(\frac{y}{x} \right) \right\} \quad (a) \\ &\quad \left(\text{Because } \tan \theta = \frac{\text{PM}}{\text{QM}} = \frac{y}{x} \right) \end{aligned}$$

This then is the general expression for the distance between the points in which the incident and reflected rays cut the axis.

Now, by Trigonometry, we have (A and B being any two quantities)

$$\begin{aligned} \cotan \{ 2 \tan^{-1} A - \tan^{-1} B \} &= \cotan \left\{ \tan^{-1} \frac{2A}{1-A^2} - \tan^{-1} B \right\} \\ &= \cotan \cdot \tan^{-1} \left(\frac{2A - (1-A^2)B}{(1-A^2) + 2AB} \right) \end{aligned}$$

that is, since $\cotan \cdot \tan^{-1} \theta = \frac{1}{\theta}$, the cotangent and tangent being reciprocals of each other, simply

$$\frac{1 - A^2 + 2AB}{2A - (1 - A^2)B}$$

Applying this to the present case, $A = \frac{dy}{dx} = p$; $B = \frac{y}{x}$, and therefore the expression above found for Qq becomes

$$\begin{aligned} \text{Qq} &= x - y \cdot \frac{(1-p^2)x + 2py}{2px - (1-p^2)y} \left\{ \right. \\ &= 2 \cdot \frac{(x+py)(px-y)}{2px - (1-p^2)y} \left. \right\} \quad (b) \end{aligned}$$

These expressions contain the whole theory of the foci and aberrations of reflecting surfaces.

Corol. 1. To find the angle made by the reflected ray with the axis, which we will call θ' .

This is the angle PqM , which is the complement of MPq . Now we have found above

$$\text{MPq} = 90^\circ - 2 \tan^{-1} p + \theta.$$

Hence

$$\theta' = 2 \cdot \tan^{-1} p - \theta$$

But $\tan \theta = \frac{y}{x}$, so that substituting we have

$$\tan \theta' = \frac{2px - (1-p^2)y}{(1-p^2)x + 2py}; \quad (c)$$

Corol. 2.

$$\text{Aq} = a' = a + 2 \frac{(x+py)(px-y)}{2px - (1-p^2)y}. \quad (d)$$

In all the foregoing formulæ we have supposed the origin of the x placed at Q the radiant point. If we would place it elsewhere, as at A , we have only to write $x - a$ for x throughout. The formulæ then become on this hypothesis,

$$\left\{ \begin{aligned} \tan \theta &= \frac{y}{x-a} \end{aligned} \right. \quad (e)$$

$$\left\{ \begin{aligned} \tan \theta' &= \frac{2p(x-a) - (1-p^2)y}{(1-p^2)(x-a) + 2py} \end{aligned} \right. \quad (f)$$

$$\left\{ \begin{aligned} \text{AQ} = a; \text{Qq} &= \frac{2(x-a+py)(px-pa-y)}{2p(x-a) - (1-p^2)y} \end{aligned} \right. \quad (g)$$

$$\left\{ \begin{aligned} \text{Aq} = a' &= \frac{2(x+py)(px-y) + \{ (1-p^2)y - 2px \} a}{2px - (1-p^2)y - 2pa} \end{aligned} \right. \quad (h)$$

General expressions for the distance of the focus from the radiant point Qy .
110.
Angle made by the reflected ray and the axis

111.

112.
Formulae when the radiant point is not in the origin of the coordinates.

Light.
113.
Formulæ
when the in-
cident rays
are parallel
to the axis.

If the incident ray be parallel to the axis, we have only to suppose the point Q infinitely distant; or placing, as in the last article, the origin of the x at a point A at a finite distance, to make $a (=AQ)$ infinite. The above expressions then give $Qq = \infty$ Part I.

$$\left. \begin{aligned} \tan \theta' &= \frac{2p}{1-p^2} \\ Aq &= x - y \cdot \frac{1-p^2}{2p} \end{aligned} \right\} \quad (i)$$

114. *Proposition.* To represent the incident and reflected rays by their equations.

The equation of any straight line is necessarily of the form $Y = aX + \beta$. Suppose we take A for the common origin of the coordinates, and, retaining the foregoing notation, representing by x and y the coordinates of the point P in the curve, let X and Y represent those of any point in the incident ray; and, Q being the point in which that ray cuts the axis, and $AQ = a$, it is evident, first, that when $X = a$, $Y = 0$; and secondly, since the ray passes through P, when $X = x$, $Y = y$. Hence we have

$$0 = a \cdot a + \beta, \quad \text{and} \quad y = ax + \beta,$$

whence we get
$$a = \frac{y}{x-a} \quad \beta = -\frac{ay}{x-a}; \quad (1)$$

therefore, the equation of the incident ray is

$$Y = \frac{y}{x-a} (X-a); \quad (2)$$

or which is the same in a different form,

$$Y - y = \frac{y}{x-a} (X-x); \quad (3)$$

or, since

$$\tan \theta = \frac{PM}{MQ} = \frac{y}{x-a},$$

$$Y = (X-a) \cdot \tan \theta; \quad (4)$$

or, again,

$$Y - y = (X-x) \cdot \tan \theta. \quad (5)$$

Similarly for the reflected ray, it is obvious that if we represent its equation by $Y = a'X + \beta'$, we shall have

$$a' = \frac{y}{x-a'} \quad \beta' = -\frac{a'y}{x-a'}; \quad (6)$$

and consequently

$$Y = \left(\frac{y}{x-a'} \right) \cdot (X-a') = (X-a') \cdot \tan \theta'; \quad (7)$$

$$Y - y = \frac{y}{x-a'} \cdot (X-x) = (X-x) \cdot \tan \theta'; \quad (8)$$

will be the corresponding forms of the equation of the reflected ray, in which a' and $\tan \theta'$ are given in terms of x , y , a , and $p = \frac{dy}{dx}$ by the equations (g) and (h) or (i).

115.
Fig. 11.

Focus.
Focal
distance.
Vertex.

If the whole figure (fig. 11) be turned about the axis AM, and Q be supposed a radiant point, the rays in the whole conical surface generated by the revolution of QP will be concentrated after reflexion in one and the same point q , which will thus become infinitely more illuminated than by any single ray from an elementary molecule of the surface. The point P will generate an annulus, having MP for its radius; and q is called the *focus* of this annulus, and the distance Aq the *focal distance* of the same annulus. This last expression is commonly understood to mean the distance of q from the *vertex*, or point where the curve meets the axis, but we shall use it at present in the more general sense.

116.

Generally speaking, then, the focus varies as the point P in the reflecting annulus varies, unless in that particular case where, by the nature of the curve, the function expressing a' is constant. Let us examine this case.

117.

Investiga-
tion of the
curves
which re-
flect all the
incident
rays to one
point.

Proposition. To find the curve which will have the same focus for every point in its surface of revolution, or on which rays diverging from or converging to any point Q, being incident, shall all after reflexion converge to or diverge from one point q .

The value of Qq assigned in Art. 109, Eq. (b) being made constant, affords the equation

$$\frac{(x+py)(px-y)}{2px - (1-p^2)y} = \text{constant} = c.$$

Light. This equation, cleared of fractions, and putting x for $x - c$, (which is merely shifting the origin of the co-ordinates to the distance c from their former origin) becomes

Part I.

$$p \{ x^2 - y^2 - c^2 \} = (1 - p^2) x y. \quad (a)$$

To integrate this equation, assume a new variable z , such that $py = xz$, and (multiplying the original equation by y) we have

$$py (x^2 - y^2 - c^2) = xy^2 - x \cdot p^2 y^2,$$

that is

$$xz (x^2 - y^2 - c^2) = xy^2 - x^3 z^2,$$

whence we find

$$y^2 = \frac{zx^2 - zc^2 + z^2 x^2}{1 + z} = x^2 z - c^2 \cdot \frac{z}{1 + z}.$$

Differentiating this equation we get

$$\begin{aligned} 2y dy \left(= 2py dx = z x z dx \text{ because } p = \frac{dy}{dx} \right) \\ = d \left(x^2 z - \frac{c^2 z}{1 + z} \right) \\ = 2xz dx + x^2 dz - c^2 d \cdot \left(\frac{z}{1 + z} \right) \end{aligned}$$

that is

$$x^2 dz - c^2 d \cdot \frac{z}{1 + z} = 0,$$

or

$$\left\{ x^2 - \frac{c^2}{(1 + z)^2} \right\} dz = 0. \quad (b)$$

This equation may obviously be satisfied in two ways; the first is, by putting the factor

$$x^2 - \frac{c^2}{(1 + z)^2} = 0, \text{ or } x = \pm \frac{c}{1 + z}$$

which gives (restoring the value of z , $z = \frac{py}{x}$) merely $x + py = c$; and, eliminating p between this and the original equation (a) we find, on reduction,

$$y^2 + (x - c)^2 = 0.$$

This is, however, (as is clear from the way in which it has been obtained,) only a *singular solution* of the differential equation, (see DIFFERENTIAL CALCULUS, *singular solutions*;) and as the value of y which results from it is always imaginary, it affords no curve satisfying the conditions of the problem.

The other way in which the equation (b) can be satisfied, is by putting $dz = 0$, or $z = \text{constant}$. Let this constant be represented by $-h$; then, since $z = \frac{py}{x}$, we have

The curve is in all cases a conic section.

$$\frac{py}{x} = \frac{y dy}{x dx} = -h,$$

which, integrated, gives

$$y^2 = h (a^2 - x^2),$$

a being another constant. This is the general equation to the conic sections, and it is obvious, from the properties of these curves, that they satisfy the conditions; because two lines drawn from their foci to any point in their periphery make equal angles with the tangent at that point, and, consequently, a ray proceeding from, or converging to, one focus, and reflected at the curve, must necessarily take a direction to or from the other. But, the foregoing analysis being direct, shows that they possess this property in common with no other curves.

Thus in the case of the ellipse, all rays, (fig. 12,) SP , SP' , &c. diverging from the focus S will after reflexion converge to the other focus H , the interior surface of the ellipse being polished; and all rays QP , QP' , &c. converging to S , will after reflexion diverge from H .

118. Ellipse. Fig. 12.

In the hyperbola, (fig. 13,) rays QP , $Q'P$, &c. converging to one focus S , and incident on the polished convex surface of the curve, will after reflexion converge to the other focus H ; and if diverging from S , and reflected on the polished concave surface PP' , will after reflexion diverge from H .

Fig. 13. 119. Hyperbola. Fig. 14.

In the case of the parabola, rays parallel to the axis, incident on the interior or concave surface, will all be reflected to the focus S , fig. 14; and if reflected at the exterior or convex surfaces, will all after reflexion diverge from S .

120. Parabola. Fig. 14.

Rays converging to, or diverging from, the centre of a sphere will all after reflexion diverge from, or converge to, the same centre.

121. Circle.

Let us now apply our general formula (b) (Art. 109) to some particular cases.

Light.

Proposition. Let the reflecting surface be a plane, or the curve PC a straight line. Required the focus of reflected rays. Part I.

122.

Focus of a
plane sur-
face.

Here we have $x = \text{constant} = a$, $p = \frac{dy}{dx} = \infty$, and the general formula becomes simply

$$Qq = a' = \frac{2xy}{y} = 2x = 2a.$$

So that the focus of reflected rays is a point on the opposite side of the reflecting plane equally distant from it with the radiant point; and as this is independent of y , or of the situation of the point P, we see that all the rays after reflexion diverge from this point, see fig. 15.

123.

Focus of a
spherical
annulus.

Proposition. To find the focus of any annulus of a spherical reflector.

Let r be the radius of the sphere, and, if we fix the origin of the coordinates at the radiant point, the equation of the generating circle will be

$$r^2 = (x - a)^2 + y^2$$

This, differentiated, gives

$$(x - a) dx + y dy = 0,$$

consequently

$$p = \frac{dy}{dx} = -\frac{x - a}{y}; \quad 1 - p^2 = \frac{2y^2 - r^2}{y^2}.$$

Hence, substituting in the general expression (b), we find for the focal distance the following value,

$$Qq = \frac{2a\{r^2 + a(x - a)\}}{r^2 + 2a(x - a)}; \quad (a)$$

which expresses in all cases the distance of the focus of reflected rays from the radiant point.

For optical purposes, however, it is more convenient to know its distance from the centre, or from the surface.

The distance from the centre (Eq, fig. 16,) is

$$= Qq - QE = \frac{2a(ax - a^2 + r^2)}{2ax + r^2 - 2a^2} - a,$$

or

$$Eq = \frac{ar^2}{2a(x - a) + r^2}; \quad (b)$$

in which positive values of Eq lie to the right of E, or the same way with those of x or of Qq .

Focus for
central rays
in a spher-
ical re-
flector.

Corol. 1 If we would find the focus of the infinitely small annulus immediately adjoining to the vertex C, or C' of the reflecting spherical surface, or, as it is termed in Optics, the *focus of central rays*, we must put in the case of the vertex C (when the reflexion takes place on a concave surface) $x = a + r$, and in the other case, viz. that where the rays are reflected on the convex surface C', $x = a - r$. The former gives

$$Eq = \frac{ra}{2a + r}; \quad Cq = \frac{r(a + r)}{2a + r}; \quad (c)$$

the latter gives the same results, writing only $-r$ for r .

124.

If we bisect the radii CE and C'E in F and F', and suppose q and q' to be the foci of central rays reflected

$$\text{at C and at C', we shall have} \quad Fq = \frac{1}{2}r - \frac{ra}{2a + r} = \frac{\left(\frac{r}{2}\right)^2}{a + \frac{r}{2}}, \quad (d)$$

which gives the following useful analogy,

$$QF : FE :: EF : Fq. \quad (e)$$

Similarly we have $QF' : F'E :: EF' : F'q$; so that the same analogy applies to both cases, and may be regarded as the fundamental proposition in the theory of the foci for central rays. For it is obvious, that if PC were any other curve than a circle, the same must hold good, taking only E the centre of curvature at the vertex.

125.

Principal
focus.

Corol. 2. If a be infinite, or the incident rays be parallel, we have $Fq = 0$, which shows that the *focus of central parallel rays* bisects the radius. This focus, for distinction's sake, is called the *principal focus* of the reflector.

126.

Conjugate
foci.

Definition. Q and q are termed *conjugate foci*. It is evident that if q be made the radiant point, Q will be its focus; for the rays will pursue the same course backwards.

127.

Corol. 3. Regarding only central rays: the conjugate foci move in opposite directions, and coincide at the centre and surface of the reflector.

For let a vary from ∞ to $-\infty$, then Fq will vary as follows: first, while a varies from ∞ to $-\frac{r}{2}$, q is

Light. positive, and increases from 0 to ∞ ; that is, as Q moves up to F, q moves through C to infinity. As the motion of q continues, Fq then becomes negative; because a is then negative and greater than $\frac{r}{2}$, and a increasing Fq diminishes; therefore q moves from the right towards F, that is in the opposite direction to Q's motion; and when Q is at an infinite distance to the right, q is again at F. Part I.
Conjugate foci move in opposite directions.

When Q comes to E, $a = 0$. $Fq = \frac{r}{2}$, or q is at E also.

When Q comes to C, $a = -r$, $Fq = -\frac{r}{2}$, or q is at C also.

It appears by the value of E q, Equation (b), that a spherical reflector ACB, fig. 17, whose chord (or aperture, as it is termed in Optics) is AB, causes the ray reflected at its exterior annulus A to converge to, or diverge from, a point q, different from the focus of central rays. Let f be this latter focus, then we shall have 128.
Longitudinal aberration, for any aperture.
Fig. 17.

$$E f = \frac{a r}{2 a + r}, C f = \frac{(a + r) r}{2 a + r}; f q = \frac{a r^2}{2 a (x - a) + r^2} - \frac{a r}{2 a + r}.$$

This quantity f q is called the *longitudinal aberration* of the spherical reflector. If the rays fall on the convex portion, we need only write $-r$ for r.

Proposition. To express approximately the longitudinal aberration of a spherical reflector whose aperture is inconsiderable with respect to its focal length. 129.

y being the semi-aperture, and $x - a$ being equal to $\sqrt{r^2 - y^2} \approx r - \frac{y^2}{2r}$, (neglecting y^4 , and higher powers of y,) we have Longitudinal aberration for small apertures.

$$f q = \text{aberration} = \frac{a r^2}{2 a r + r^2 - \frac{a y^2}{r}} - \frac{a r}{2 a + r} = \frac{a^2 y^2}{r (2 a + r)^2}; \quad (f)$$

If we put $C f = f$, we have $f = \frac{r (a + r)}{2 a + r}$, and, consequently, we may eliminate a, the distance of the radiant point, and express the aberration in terms of the aperture, radius of curvature, and distance of the focus of central rays from C, the vertex of the minor; for this gives $a = \frac{r (r - f)}{2 f - r}$, which, substituted for a in the expression (f) gives 130.
Another expression

$$\text{aberration} = \frac{(r - f)^2 \cdot y^2}{r^3} = \frac{E f^2 \cdot (\text{semi-aperture})^2}{(\text{rad.})^3}. \quad (g)$$

To express the *lateral aberration*, or the quantity by which the reflected ray A q g deviates from the axis, at the focus of central rays, or the value of f g, (fig. 17,) we have 131.

$$f g = f q \cdot \frac{A M}{M q}; \text{ but } A M = y, \text{ and } M q = E M - E q = x - a - \frac{a r^2}{2 a (x - a) + r^2} \\ = \frac{2 a (x - a)^2 + r^2 (x - 2 a)}{2 a (x - a) + r^2}; \text{ so that}$$

$$f g = \frac{2 a^2 r}{2 a + r} y \times \frac{a - x + r}{r^2 (x + 2 a) + 2 a (x - a)^2}. \quad (h)$$

When the aperture is very small, this becomes simply

$$f g = \frac{a^2 y^3}{r^2 \cdot (r + a) (r + 2 a)}. \quad (i)$$

When a is infinite, or the incident rays are parallel, we have the following,

$$\left. \begin{aligned} f q &= \text{longitudinal aberration} = \frac{y^3}{4 r} \\ f g &= \text{lateral aberration} = \frac{y^3}{2 r^2} \end{aligned} \right\} \quad (j)$$

If the rays fall on the convex side of the sphere we must make r negative, which only changes the signs of the aberrations.

132.

Lateral aberration for small apertures.

133.

Aberrations for parallel rays and small apertures.

§ V. Of Caustics by Reflexion, or Catacaustics.

134. If rays of light be incident on a medium of any other form than that of a conic section, having the radiant point in the focus, they will after reflexion no longer converge to or diverge from any one point, but will be dispersed according to a law depending on the nature of the reflecting curve; the inclination of each reflected ray to the axis varying according to the point of the curve from which it is reflected, and not being the same for any two consecutive rays. Of course each ray will intersect that immediately consecutive to it in some point or other, and the *locus* of these points of continual intersection will trace out a curve to which the reflected rays will all necessarily be tangents, and which is called a *caustic*. If these rays fall on another reflecting curve, they will be again dispersed, and another caustic will originate in the continual intersections of the consecutive rays of the former, and so on to infinity.

Caustics by
reflexion
defined.

135.
Fig. 18.

Let $QP, Q'P'$, (fig. 18,) be any two contiguous rays incident on consecutive points P, P' of a reflecting curve PP' , and after reflexion let them pursue the courses $PR, P'R'$; and since they are not necessarily parallel, let Y be their point of intersection, then will Y be the point in the caustic $YY'Y''$ corresponding to the point P in the reflecting curve; and if we determine the points $Y'Y''$, &c. from the consecutive points $P'P''$, &c. in the same manner, the locus of these, or the curve $YY'Y''$ will be the whole caustic.

136.
Coordinates
of the
caustic in-
vestigated
on any sup-
position of
the law of
divergence.

Since the reflected ray passes through P , whose coordinates are xy , its equation, as we have already seen (Art. 114), is necessarily of the form

$$Y - y = P(X - x)$$

If we regard x, y, P as variable, this will represent any one of the reflected rays PR , and the consecutive ray $P'R'$ will be represented by

$$Y - (y + dy) = (P + dP)(X - (x + dx))$$

Now since the point Y in which these two rays intersect is common to both, the coordinates X and Y at this point are the same for both; and therefore at this point both these equations coexist, and thereby determine the values of X and Y , or the situation of the point Y . Now the latter of these equations is nothing more than the former *plus* its differential, on the supposition of X and Y remaining constant. Therefore, we have to find X and Y from the two equations,

$$Y - y = P(X - x)$$

$$-dy = (X - x)dP - Pd x,$$

which gives at once

$$\left. \begin{aligned} X &= x + \frac{P - p}{dP} dx \\ Y &= y + P \cdot \frac{P - p}{dP} dx \end{aligned} \right\} \quad (k)$$

In these equations we have only to substitute for P its value $= \tan \theta'$, or $\frac{2p(x - a) - (1 - p^2)y}{(1 - p^2)(x - a) + 2py}$; and

after executing all the differentiations indicated, or implied, to eliminate x and y by the equations of the curve and the other conditions to which the quantity a may be subjected, an equation between X and Y will result which will be the equation of the caustic.

137. *Proposition.* To determine the caustic when rays diverge from one fixed point in the axis of a given reflecting curve.

Caustic
when rays
diverge from
a fixed
point.

In this case a is invariable, and the differentiation of P must be performed on this hypothesis. It will, therefore, simplify the question if we put $a = 0$; or suppose the origin of the coordinates in the radiant point, in which case

$$\left. \begin{aligned} P &= \frac{2px - (1 - p^2)y}{2py + (1 - p^2)x} \\ \frac{dP}{dx} &= (1 + p^2) \cdot \frac{(1 + p^2)(y - px) + 2q(x^2 + y^2)}{\{2py + (1 - p^2)x\}^2} \\ q &= \frac{dp}{dx} \\ P - p &= \frac{(1 + p^2)(px - y)}{2py + (1 - p^2)x} \end{aligned} \right\} \quad (l)$$

Where

Light. which substituted, we find

$$\left. \begin{aligned} X &= 2 \cdot \frac{p(p x - y)^2 - q x(x^2 + y^2)}{(1 + p^2)(p x - y) - 2 q(x^2 + y^2)} \\ Y &= 2 \cdot \frac{(p x - y)^2 + q y(x^2 + y^2)}{-(1 + p^2)(p x - y) + 2 q(x^2 + y^2)} \end{aligned} \right\} (m)$$

Corol. 1. If the incident rays be parallel, or the radiant point at an infinite distance, we may fix the origin of the coordinates where we please; and since in this case the equation of any reflected ray is, by 113 equation (i) and 114 equation (8), 138. Caustic for parallel rays.

$$Y - y = (X - x) \cdot \frac{2 p}{1 - p^2} \quad (m, 2)$$

we have $P = \frac{2 p}{1 - p^2}$; $P - p = \frac{p(1 + p^2)}{1 - p^2}$; $\frac{d x}{d P} = \frac{(1 - p^2)^2}{2 q(1 + p^2)}$

putting q for $\frac{d p}{d x}$ or $\frac{d^2 y}{d x^2}$.

These substitutions made, we get the following values for the coordinates of the caustic,

$$X = x + \frac{p}{2 q} (1 - p^2); \quad Y = y + \frac{p^2}{q}. \quad (n)$$

Corol. 2. In the general case, if we put f = the line $P y$, or the distance between the point in the curve and the corresponding point in the caustic we have 139. Distance between corresponding points in curve and caustic

$$f = \sqrt{(X - x)^2 + (Y - y)^2}$$

Which, if we write for $X - x$ and $Y - y$, their values above found become

$$f = \sqrt{1 + P^2} \cdot \frac{P - p}{d P} d x \quad (o)$$

or, writing for P its value, and executing the operations,

$$f = \frac{-(y - p x)(1 + p^2) \sqrt{x^2 + y^2}}{(y - p x)(1 + p^2) + 2 q(x^2 + y^2)}. \quad (p)$$

Corol. 3. In the case of parallel rays, when

$$P = \frac{2 p}{1 - p^2}; \quad \frac{d P}{d x} = \frac{2 q(1 + p^2)}{(1 - p^2)^2}; \quad P - p = \frac{p(1 + p^2)}{1 - p^2}; \quad \sqrt{1 + P^2} = \frac{1 + p^2}{1 - p^2}$$

we have

$$f = \frac{p(1 + p^2)}{2 q}. \quad (q)$$

Corol. 4. Call c the chord of the circle of curvature passing through the origin of the coordinates, or through the radiant point: then, by the theory of curves, 141.

$$c = \frac{2(p x - y)(1 + p^2)}{q \sqrt{x^2 + y^2}}$$

so that

$$q(x^2 + y^2) = \frac{2(p x - y)(1 + p^2) \sqrt{x^2 + y^2}}{c}$$

and substituting this for $q(x^2 + y^2)$, in the general expression for f , we eliminate q , and get

$$f = \frac{c \sqrt{x^2 + y^2}}{4 \sqrt{x^2 + y^2} - c} = \frac{r c}{4 r - c}$$

putting

$$r = \sqrt{x^2 + y^2}.$$

Hence we have

$$f - \frac{1}{4} c = \frac{(\frac{1}{4} c)^2}{r - \frac{1}{4} c}$$

which gives

$$r - \frac{1}{4} c : \frac{1}{4} c :: \frac{1}{4} c : f - \frac{1}{4} c. \quad (r)$$

Hence the following general property. (Smith's *Optics*, ed. 1738, p. 160.)

Q and q being two conjugate foci of an elementary pencil of rays reflected at any curve surface at P , fig. 19. 142.
Let VPW be the circle of curvature; (if the curve be a circle, this will be the curve itself.) Let the chords

Light. Part I.
 General relation between conjugate points or foci of reflected rays incident on any curve. 143.

PV, PW in the direction of the incident and reflected rays be divided in F, f, so that PF and Pf shall each be one quarter of the whole chords, and the relation between Q and q will be expressed by the proportion

$$QF \cdot FP :: Pf : f. \quad (s)$$

Corol. 5. Putting

$$\frac{P-p}{dP} dx = M, \text{ we have } \frac{dX}{dx} = 1 + \frac{dM}{dx};$$

$$\frac{dY}{dx} = p + P \frac{dM}{dx} + M \frac{dP}{dx} = P \left(1 + \frac{dM}{dx} \right)$$

Hence it follows that

$$P = \frac{dY}{dX};$$

P therefore is to the caustic, for the coordinates X, Y, what p is to the reflecting curve for the corresponding point whose coordinates are x, y.

144.
 Length of the caustic investigated.

Corol. 6. If we put S for the length of the caustic

$$= \text{the arc AHKY, we have } dS = \sqrt{dX^2 + dY^2}$$

or

$$dS = dX \cdot \sqrt{1 + P^2} = (dx + dM) \sqrt{1 + P^2}$$

$$dx \cdot \sqrt{1 + P^2} + df - M \cdot \frac{P dP}{\sqrt{1 + P^2}}$$

because $df = d \cdot M \cdot \sqrt{1 + P^2} + M \cdot \frac{P dP}{\sqrt{1 + P^2}}; \text{ but } M dP = (P - p) dx$

so that we have

$$dS = df + dx \left\{ \sqrt{1 + P^2} - \frac{(P - p)P}{\sqrt{1 + P^2}} \right\}$$

$$= df + dx \cdot \frac{1 + Pp}{\sqrt{1 + P^2}};$$

that is, substituting for P its value

$$\frac{2px - (1 - p^2)y}{2py + (1 - p^2)x},$$

$$dS = df + dx \cdot \frac{x + py}{\sqrt{1x^2 + y^2}} = df + d \cdot \sqrt{x^2 + y^2}$$

and integrating

$$S = \text{constant} + f + \sqrt{x^2 + y^2}.$$

Caustics always rectifiable.

Hence it follows, that the caustic is always a rectifiable curve, and its

But

$$\left. \begin{array}{l} \text{Length AKY} = QP + Py + \text{constant} \\ \text{Arc AKF} = QC + CF + \text{constant} \end{array} \right\} \text{consequently, subtracting}$$

$$\text{Arc Fy} = (QC + CF) - (QP + PY).$$

Hence it appears, that the caustic is necessarily a rectifiable curve when the reflecting curve is not itself transcendental.

145.
 Fig. 20.

If the rays PR, P'R', P''R'', &c. after reflexion at the curve PP'P'' fall on another reflector RR'R'' and are reflected in the directions RS, R'S', R''S'', &c. (fig. 20) their continual intersections will form another caustic ZZ'Z'', and so on *ad infinitum*, which may be determined by a similar analysis. In like manner, whatever be the law according to which the rays QP, Q'P', &c. are dispersed, we may conceive each to be a tangent to a curve which may be regarded as the caustic of another reflecting curve, and so on. Let VV'V'' be this curve. Since PVQ is a tangent to it, if this curve and the curve PP'P'' be given, the point Q in the axis from which the incident ray QP may be regarded as radiating, is determined in terms of the coordinates of P, and therefore the quantity *a* may be eliminated altogether. The manner of doing this is shown in the following

146.

General relation between two conjugate caustics and their intermediate reflecting curve investigated.

Proposition. To determine the relations between any two consecutive, or, as they may be termed, *conjugate* caustics VV'V'', YY'Y'', and the intermediate reflecting curve PP'P''.

Let V and Y be, as before, any two conjugate points in the caustics, P the reflecting point; then if we put

ξ and η for the coordinates of V

x and y for those of P

X and Y of Y

Light. Since the line P V Q is a tangent to the first curve at V, we must evidently have

Part I.

$$y - \eta = \frac{d\eta}{d\xi} (x - \xi)$$

and this, combined with the equation between η and ξ , which represents the curve V V' V'' suffices to determine η and ξ in terms of x , y , or *vice versa*, x and y in terms of ξ and η .

Again, we have also by Art. 114, equation (2)

$$y - \eta = \frac{y}{x - a} (x - \xi)$$

and consequently

$$x - a = y \cdot \frac{x - \xi}{y - \eta}; \quad a = \frac{\xi y - \eta x}{y - \eta}.$$

Thus a is given in terms either of x , y , or of η , ξ , whichever we may prefer. It only remains to substitute this in the value of P.

$$P = \frac{2p(x-a) - (1-p^2)y}{(1-p^2)(x-a) + 2py}$$

which thus becomes

$$P = \frac{2p(x-\xi) - (1-p^2)(y-\eta)}{(1-p^2)(x-\xi) + 2p(y-\eta)}; \quad (t)$$

and this, being free of a , may be substituted in the equations (k) Art. 136, when X and Y will be at once obtained in terms of x , y , ξ , η , the coordinates of the reflecting curve and the preceding caustic.

We shall now proceed to illustrate the theory above delivered by an example or two.

Required the caustic when the reflecting curve is a cycloid, and the incident rays are parallel to each other and to the axis of the cycloid.

147.
Caustic of
a cycloid.

The equation of the cycloid is

$$\frac{dy}{dx} = p = \frac{\sqrt{x}}{\sqrt{2-x}}$$

taking unity for the radius of the generating circle.

From this we get

$$\frac{1}{q} (2-x) \sqrt{2x-x^2}$$

and therefore

$$\frac{p}{q} = 2x - x^2;$$

consequently, by the equations (k) of Art. 136, we shall have

$$X = x + \frac{1-p^2}{2} \cdot \frac{p}{q} = 2x - x^2$$

$$Y = y + p \cdot \frac{p}{q} = y + x \sqrt{2x-x^2}$$

whence

$$\frac{dY}{dx} = p + \frac{\sqrt{x}}{\sqrt{2-x}} (3-2x) = 2\sqrt{2x-x^2} = 2\sqrt{X}$$

Now we have also

$$\frac{dX}{dx} = 2(1-x)$$

But since

$$X = 2x - x^2, \text{ we have } 1-x = \sqrt{1-X}, \quad \text{and therefore}$$

$$\frac{dX}{dx} = 2\sqrt{1-X}$$

So that we have, finally,

$$\frac{dY}{dX} = \sqrt{\frac{X}{1-X}}$$

which shows that the caustic is itself a cycloid of half the linear dimensions of the reflecting curve.

Is itself

To take one other example, let us suppose the reflecting curve a circle, and the radiant point infinitely distant. Here we have (placing the origin of the coordinates in the centre)

148.

$$x^2 + y^2 = r^2; p = -\frac{x}{\sqrt{r^2 - x^2}}; q = -\frac{r^2}{(r^2 - x^2)^{\frac{3}{2}}},$$

Caustic of a consequently, by the equations (k) of Art. 136 circle.

$$\left. \begin{aligned} X &= x + \frac{p(1-p^2)}{2q} = \frac{3r^2 - 2x^2}{2r^2} x \\ Y &= y + \frac{p^2}{q} = \frac{(r^2 - x^2)^{\frac{3}{2}}}{r^3} = \frac{y^3}{r^3} \end{aligned} \right\}; \quad (u)$$

Then since (supposing, for brevity, $r = 1$, which will not affect the result)

$$4X^2 = 9x^2 - 12x^4 + 4x^6$$

$$4Y^2 = 4 - 12x^2 + 12x^4 - 4x^6$$

Adding,

$$4(X^2 + Y^2) = 4 - 3x^2; x^2 = \frac{4}{3}(1 - X^2 - Y^2)$$

So that we get, finally, substituting this value of x^2 in that of Y , and reducing,

$$(4X^2 + 4Y^2 - 1)^2 = 27Y^2; \quad (v)$$

which is the equation of the caustic.

This equation belongs to an epicycloid generated by the revolution of a circle whose radius is $\frac{1}{4}$ that of the reflecting circle on another concentric with the latter, and whose radius is $\frac{1}{2}$ that of the reflecting circle. Fig. 21 represents the caustic in this case; QP being the incident ray, and PY the reflected. It has a cusp at F, which is the principal focus of rays reflected at the concave surface BCD, and another at F', which is that of the rays reflected from the convex surface BAD. In the latter case, it is not the rays themselves, but their prolongations backwards which touch the caustic.

149. *Corol.* When y is very small, or immediately adjacent to the cusp F, the form of the caustic approaches indefinitely to that of a semicubical parabola. For, generally,

$$X = \frac{1}{2} \sqrt{1 + 3Y^2 - 4Y^3},$$

and when Y is very small, neglecting Y^3 in comparison with Y^2

$$X = \frac{1}{2} + \frac{3}{4} Y^{\frac{5}{2}}, \text{ or } Y^2 = \left(\frac{4}{3}\right)^{\frac{2}{5}} \cdot \left(X - \frac{1}{2}\right)^{\frac{2}{5}}; \quad (w)$$

150. It is, as we have seen, only in certain very particular cases, when rays proceeding from one point and reflected at a curve proceed after reflexion all to or from one point. In general they are distributed in the manner described in Art. 145, 146, being all tangents to the caustic. The density of the rays therefore in any point of the caustic is infinitely greater than in the space surrounding it, and in the space between the caustic and the reflecting curve (PCFY, fig. 18) is greater than in the space without the caustic QYF. This is obvious, for in the latter space only the incident rays occur, while in the former are included all the reflected rays as well as the incident ones.

151. This may be easily shown experimentally, in a very satisfactory manner pointed out by Dr. Brewster, by bending a narrow strip of polished steel into any concave form, as in fig. 22, and placing it upright on a sheet of white paper. If in this state it be exposed to the rays of the sun, holding the plane of the paper so as to pass nearly but not quite through the sun, the caustic will be seen traced on the paper, and marked by a very bright well-defined line; the part within being brighter than that without, and the light graduating away from the caustic inwards by rapid gradations. If the form of the spring be varied, all the varieties of catacaustics, with their singular points, cusps, contrary flexures, &c. will be seen beautifully developed. The experiment is at once amusing and instructive.

The bright line seen on the surface of a drinking-glass full of milk, or, better still, of ink, standing in sunshine, is a familiar instance of the caustic of a circle just investigated.

152. If the figure 18 be turned round its axis, the reflecting curve will generate a surface of revolution, which, if supposed polished within or without, as the case may be, will become a *mirror*. The caustic will also generate a conoidal surface, to which all the rays reflected by the mirror will be tangents. No mirror, therefore, which is not formed by the revolution of a conic section having the radiant point in its focus, can converge all the reflected rays to one point or *focus*. There will, however, always be one point which receives the reflected rays in a more dense state than any other. This point is the cusp F, as we shall presently see. The deviation of any reflected ray from this point is what is termed its *aberration*.

153. The concentration and dispersion of rays by reflecting and refracting surfaces forming the great business of *practical optics*, it will be necessary to enter at large into this subject; and, first, it will be proper to inquire how far any given reflector will enable us to concentrate the rays which fall on it. To this end let the following problem be proposed.

154. *Proposition.* A reflector of any figure, of a given diameter or aperture AB, being proposed, to find the circle of least aberration, or the place where a screen must be placed to receive all the rays reflected from the surface, within the least possible circular space (since they cannot be all collected in one point) and the diameter of this circle.

Light. A C B (fig. 23) being the mirror, Q the radiant point, G K *f k g* the caustic, *f* the cusp or focus for central rays, *q* the focus of the extreme rays A *q*, B *q*, produce these lines till they cut the caustic in Y *y*. It is clear, then, since all the rays reflected from the portion A C B of the reflector are tangents to points of the caustic between K, *f* and *k*, *f*, that they must all pass through the line Y *y*. Retaining the notation of the foregoing propositions, (*i. e.* supposing Q *x* = X; X *y* = Y.) Let us put Q L = \hat{X} , L K = \hat{Y} , Q D = \hat{x} ; D A = \hat{y} ; and let \hat{P} , \hat{p} represent the values of P and *p* corresponding to the points K and A of the caustic and reflecting curves. The equation of the line A K *q y* will then be

$$Y - \hat{y} = \hat{P} (X - \hat{x}); \quad (x)$$

Y and X being the coordinates of any point in it. But at the point *y*, where it cuts the other branch of the caustic, these coordinates are common to the straight line, and to the caustic. At this point, therefore, the above equation, and those expressing the nature of the caustic, must subsist together. Now these are the equations (*k*) Art. 136, combined with the original equation of the reflecting curve. Eliminating, then, *x* and *y*, by the aid of two of them, and determining the values of X, Y from the rest, the problem is resolved.

Now the same equation which gives the value of *y*, or *x y*, must also give that of L K, because K is a point in both caustic and the line A K *y*, as well as *y*. But, moreover, since A K *y* is a tangent, the point K is a double point; therefore the final equation in Y must necessarily have two equal roots, besides the value of Y sought; and these being known, the other may be found from a depressed equation.

The method here followed is, apparently, different from that usually employed, which consists in making the value of *y* as determined by the intersection of the extreme reflected ray A K *y*, and any other reflected ray (from P) a maximum. But the difference is only apparent, for in the latter method we have to make Y as determined by the two equations (holding good jointly)

$$Y - \hat{y} = \hat{P} (X - \hat{x}), \text{ and } Y - y = P (X - x)$$

a maximum, or *d* Y = 0. Now in this case the former equation gives *d* X = 0 also; and therefore, differentiating the latter, we have

$$-d y = (X - x) d P - P d x,$$

whence

$$X - x = \frac{P - p}{d P} d x$$

and therefore

$$Y - y = P \cdot \frac{P - p}{d P} d x.$$

Now these are nothing more than the equations of Art. 136, expressing the general properties of the caustic; so that this consideration of the maximum only leads by a more circuitous route to the same equations as the method above stated, and is in fact nothing more than a different mode of expressing the caustic.

Let us apply this reasoning to the case when the reflector is spherical. Resuming the equations and notation of Art. 148, and putting *a* for the extreme value of *y*, or the *semi-aperture* of the mirror, and *b* for the corresponding value of *x*, that of P will be

$$\left(\frac{2 p}{1 - p^2} = \right) = \frac{2 a b}{b^2 - a^2} = \frac{2 a b}{1 - 2 a^2}.$$

Hence the equation (*m*, 2) Art. 138, of the extreme reflected ray becomes

$$Y - a = \frac{2 a b}{1 - 2 a^2} (X - b)$$

whence we get

$$2 X = \frac{1}{b} \left(1 + \frac{1 - 2 a^2}{a} \cdot Y \right)$$

Assume *z*, so that Y = *a*³ *z*³, *z* being another unknown quantity, then we have

$$4 X^2 = \frac{1}{1 - a^2} \{ 1 + (1 - 2 a^2) a^2 z^3 \}^2.$$

Substituting this for 4 X², and for Y² its value *a*⁶ *z*⁶ in the equation of the caustic (*v*) Art. 148, extracting the cube root, and reducing, we get the following equation for finding *z*,

$$a^3 z^6 + (2 - 4 a^2) z^3 + (3 a^2 - 3) z + 1 = 0.$$

Now this, according to the remark in Art. 155, must have two equal roots, viz. when *x* = *b*, or Y = *a*³, that is, when *z* = 1. Hence this equation must necessarily be divisible by (*z* - 1)². Performing the division we find it is so, and the quotient gives

$$a^2 z^4 + 2 a^2 z^3 + 3 a^2 z^2 + 2 z + 1 = 0; \quad (y)$$

for determining the remaining values of *z*.

Part I.
Fig. 23.

155.

156.
Circle of least aberration in spherical reflector.

Light.

As this investigation is rigorous, nothing having been omitted or neglected as small, we have here the complete solution of the problem, whatever be the aperture of the mirror. If this be supposed small in comparison with the radius, an approximation to the value of z will be had by the series thence derived,

157.
Case when
the aperture
is moderate.

$$z = -\frac{1}{2} - \frac{9}{32} a^2 - \frac{9}{32} a^4 - \frac{1395}{4096} a^6 - \&c.$$

and of course since $Y = a^3 z^3$,

$$Y = -\frac{a^3}{8} - \frac{27}{128} a^5 - \frac{675}{2048} a^7 - \&c. \quad (z)$$

158.
Case where
the aperture
is small
when com-
pared to
radius.

The first term of this series is sufficient for most cases which occur in practice, and gives

$$Y = -\frac{a^3}{8} \quad (a)$$

or, supposing r the radius of curvature of the reflector,

$$Y = -\frac{a^3}{8r^2} \quad (\beta)$$

The lateral aberration corresponding to the semi-aperture a is, by the equation (j), Art. 133, equal to $\frac{a^3}{2r^2}$; consequently, in the case of small apertures, the radius of the least circle of aberration is equal to $\frac{1}{4}$ of the lateral aberration (at the focus) of the exterior annulus.

159 *Corol.* The least circle of aberration is nearer the mirror than its principal focus, by $\frac{3}{4} f g$ or $\frac{3}{4}$ the longitudinal aberration $= \frac{3}{16} \cdot \frac{a^2}{r}$.

160.
Density of
reflected
rays at any
point inves-
tigated.
Fig. 24.

To complete the theory of caustics, it only remains to examine the degree of concentration of the reflected rays at any assigned point. To this end, let S (fig. 24) be any point, and through it let $PSYq$ be drawn touching the caustic in Y . Then S may be regarded as lying in a conical surface generated by the revolution of the tangent $PYsq$, about the axis; and all the rays in the annulus, generated by the revolution of the element PP' , will be contained in the hollow conoidal solid formed by the revolution of the figure $PP'Yq'q$ about the same axis. Hence at S the rays will be concentrated: first, in a plane parallel to that of the paper, in the ratio of PP' to SS' , or PY to SY ; and, secondly, in a plane perpendicular to that of the paper, or in the ratio of the circumferences of the circles generated by the revolution of P and of S , that is, in the ratio of these radii $PM : ST$. On both accounts, therefore, the concentration at S will be

represented by $\frac{PM}{ST} \times \frac{PY}{SY}$, or $\frac{Pq}{Sq} \times \frac{PY}{SY}$. If, therefore, we represent by 1 the density of the

rays immediately on their reflexion at P , their density at S corresponding, will be represented by $\frac{PY \cdot Pq}{SY \cdot Sq}$, and this is true, whatever be the situation of S .

161.
1st case.

But there are now several cases to be distinguished. First, when S is situated in any part of the spaces KHV , NDW , no such tangent can be drawn to cut the reflector within its aperture AB ; therefore these spaces receive no rays at all, and the density $= 0$ in every point.

162.
2nd case.

Secondly, when S is situated anywhere within the spaces AGB , $VHFE$, $EFDW$, only one such tangent can be drawn to cut the reflecting curve between A and B . So that in these spaces the density

is simply represented by $D = \frac{PY \cdot Pq}{SY \cdot Sq}$.

163.
3rd case.

Thirdly, within the spaces KGH and MGD two tangents can be drawn from any point S , both touching the branch Fk on the same side of the axis as the point S . If we suppose $P_1Y_1Sq_1$ and $P_2Y_2Sq_2$ to be these tangents, S will receive rays belonging to both these converging conoids, and the density will therefore be the sum of those belonging to either, or

$$D = \frac{PY_1 \cdot Pq_1}{SY_1 \cdot Sq_1} + \frac{PY_2 \cdot Pq_2}{SY_2 \cdot Sq_2}.$$

Fig. 25.

See fig. 25.

164.
4th case.

Fourthly and lastly, within the space $FHGD$ there may be drawn three tangents $q_1SY_1P_1$, $q_2SY_2P_2$, and $q_3SY_3P_3$, all falling within AB , the two first (fig. 26) touching the branch Fk on the same side as S , the

Light. third on the opposite side. The former belong to cones of rays converging to q_1, q_2 , the latter to a cone converging to q_3 , but intercepted by S after meeting at q , and again diverging. Hence, in this case, the density will be expressed by Part I.
Fig. 26.

$$D = \frac{P Y_1 \cdot P q_1}{S Y_1 \cdot S q_1} + \frac{P Y_2 \cdot P q_2}{S Y_2 \cdot S q_2} + \frac{P Y_3 \cdot P q_3}{S Y_3 \cdot S q_3}.$$

It would lead into too great complication to attempt developing the actual value of these fractions in terms of the coordinates of S, and we will therefore merely apply them to some remarkable positions of S Application to particular cases.

Case 1. S in the axis, beyond the principal focus, or between the mirror and its focus for extreme rays G. Here 165.

Y coincides with F, and q also does the same therefore in this case, $D = \left(\frac{P F}{S F} \right)^2$, which shows that the Case 1.

density is inversely as the square of the distance of S from the principal focus.

Case 2. S in the axis between the principal focus and the focus for extreme rays G, (*i. e.* in the line GF.) Here $S q_1 = 0, S q_2 = 0, S q_3 = 0$; so that here all the three several component portions of D are infinite, and of course the density is infinitely greater than on the surface of the reflector. 166.
Case 2.

Case 3. S at F. Here not only $S q = 0$, but also $S Y$; therefore at F the density is infinitely greater than in the last case, and is the greatest which exists anywhere. 167
Case 3.

Case 4. S anywhere in the caustic. Here $S Y = 0$, therefore in this case also D is infinite, or the density infinitely greater than at the surface of the reflector; and as S approaches F, this is still further multiplied by the diminution of all the values of $S q$. 168.
Case 4.

Case 5. S anywhere in H z D, the circle of least aberration. At the centre z and the circumference H the density is infinite. Between these two positions, finite, diminishing to a minimum, and again increasing according to a law too complicated to be here investigated. It will be observed, that the relations expressed in these articles (160—169) are general, and not restricted to the case where the reflecting surface is spherical. 169.
Case 5.

In all the foregoing reasoning the point S is supposed to receive the rays perpendicularly. The density of the rays therefore here intended must be understood to mean, The number of rays not incident on a given particular plane surface, but passing through a given infinitely small spherical portion of space, or received upon an infinitely small spherical body at S. 170.
Illumination of a screen exposed to the reflected rays.

In cases, however, where the aperture is small, a screen perpendicular to the axis will receive the rays from every point very nearly at a perpendicular incidence; and hence the above expressions will in this case represent the intensity of illumination of the several points in such a surface, the screen being, however, supposed to stop none of the incident light.

For further information respecting caustics, the reader is referred to Tschirnäus, Leipsic acts 1682, and *Hist. de l'Acad.*, tom. ii. p. 54, 1688; to De la Hire's *Traité des Epicycloïdes*, and *Mém. de l'Acad.*, vol. x.; to Smith's *Optics*; Carré, *Mém. de l'Acad.*, 1703; J. Bernouilli, *Opera Omnia*, vol. iii. p. 464; l'Hôpital *Analyse des Infiniment Petits*; Hayes's *Fluxions*; Petit, *Correspondence de l'Ecole Polytechnique*, ii. 553; Malus, *Journal de l'Ecole Polytech.*, vol. vi.; Gergonne, *Annales des Mathématiques*, xi. p. 229; De la Rive, *Dissertation sur les Caustiques*, &c.; Sturm, *Annales des Math.*, xvi.; Gergonne, *ditto*.

OF THE REGULAR REFRACTION OF LIGHT BY UNCRYSTALLIZED MEDIA.

§ VI. Of the Refraction of Homogeneous Light at Plane Surfaces.

When a ray of light is incident on the surface of any transparent uncrystallized medium, a portion of it is reflected; another portion is *dispersed* in all directions, and serves to render the surface visible; and the remainder enters the medium and pursues its course within it. 171.

In the reflexion of light, the law of reflexion, as far as regards the direction of the reflected ray, is the same for all reflecting media; the angle of reflexion being equal to that of incidence for all. In refraction, however, the case is otherwise, and each different medium has its own peculiar law of action on light; some turning a ray incident at a given angle more out of its course than others. Whatever be the nature of the refracting medium, the following general laws are found to hold good, and suffices (when the medium is known) to determine the direction of the refracted ray. 172

1st. The incident ray, the perpendicular to the surface at the point of incidence, and the refracted ray, all lie in the same plane. 173.

2nd. The incident and refracted rays lie on opposite sides of the perpendicular. 174.

3rd. Whatever be the inclination of the incident ray to the refracting surface, the sine of the angle included between the incident ray and the perpendicular is to the sine of that included between the refracted ray and the perpendicular in a constant ratio. 175.

These laws equally hold good for plane and for curved surfaces, and are found to be verified with perfect precision by the most delicate experiments, and all the phenomena of refracted light to take place in exact conformity with the results deduced from them by mathematical reasoning. 176.

Light. Let ACB (fig. 23) be the refracting surface, PCp the perpendicular to it at the point of incidence C , SC the incident, and Cs the refracted ray. Then we shall have

177.
Fig. 23.

$$\sin PCS : \sin pcs :: \mu : 1,$$

μ being a constant quantity; that is, constant for the same medium AB , though its value is different for different media.

178. It is usual, for brevity, to speak of the sine of incidence, and the sine of refraction, instead of the sines of the angle of incidence, and the angle of refraction.

179. The numerical value of the quantity μ , or of $\frac{\sin \text{ of incidence}}{\sin \text{ of refraction}}$ in any medium, must be ascertained before

Index of
refraction.

the law of refraction in that medium can be regarded as perfectly known. This may be done experimentally by actually measuring the angle of refraction corresponding to any one given angle of incidence, for the value of the above fraction being thus determined for one incidence holds equally for every other, or by other more easy or more refined modes to be described hereafter. This quantity μ is called the *index* of refraction of the medium AB .

180. The medium in which the ray proceeds previous to its incidence on AB is here regarded as a vacuum. If the medium AB be also a vacuum, it is clear that the ray will not change its course; so that the angle of incidence will equal that of refraction, and the value of μ will be equal to 1. This is the lowest value of μ , as no medium has yet been discovered which refracts rays from the perpendicular when incident from a vacuum. The greatest value of μ yet known is 3, when the refraction is made into chromate of lead; and between these limits almost every intermediate gradation has been found to belong to some one or other transparent body. Thus for air at its ordinary density $\mu=1.00028$, while for water it is 1.336, for ordinary crown glass 1.535, for flint glass 1.60, for oil of cassia 1.641, for diamond 2.487, and for the greatest refraction of chromate of lead 3.0.

181.
Refraction
out of any
medium into
vacuum.

It is a general law in Optics, that the visibility of two points from one another is mutual, whatever be the course pursued by the rays which proceed from one to the other. In other words, that if a ray of light proceeding from A arrives by any course at B , however often reflected, refracted, &c., a ray can also arrive at A from B by retracing precisely the same course in a contrary direction. It follows from this, that if the ray SC incident on the exterior surface of a medium AB , (fig. 23,) pursue after refraction the course Cs , then will a ray sC , incident on the exterior surface of the medium, be refracted out of it in the direction CS , being bent from the perpendicular. Consequently, since in this case the angle of incidence is the same with the angle of

refraction in the former case, and *vice versâ*, we shall have here $\frac{\sin \text{ incidence}}{\sin \text{ refraction}} = \frac{1}{\mu}$. Thus we see that the

index of refraction out of any medium into vacuum is the reciprocal of the index of refraction into the medium from the vacuum.

182. Hence it follows, that a ray can be introritted into any medium from a vacuum at any angle of incidence; for

since $\sin \text{ refr.} = \sin pcs = \frac{1}{\mu} \cdot \sin PCS$, the value of μ being greater than 1, the sine of pcs will necessarily be less than that of PCS , and of course less than unity; so that the angle of refraction can never become imaginary. Thus, as the angle of incidence PCS increases from 0 , or as the ray SC becomes more and more oblique to the surface till it barely grazes it, as at $S''C$, the refracted ray becomes also more oblique, but much less rapidly, and never attains a greater obliquity than the situation Cs'' , in which $\sin pcs'' = \frac{\sin 90^\circ}{\mu} = \frac{1}{\mu}$.

Limit of the
angle of
refraction.

This limiting angle, then, is the maximum angle of refraction from vacuum into the medium, and its value in any given medium is found by computing the angle whose sine is the reciprocal of the index of refraction. Thus in

water the angle of refraction cannot exceed arc $\sin \frac{1}{1.336}$, or $48^\circ 27' 40''$. In crown glass the limit is $40^\circ 39'$, in flint $38^\circ 41'$, in diamond $23^\circ 42'$, while for the greatest refraction of chromate of lead the limit is so low as $19^\circ 28' 20''$.

183.
Limit to the
possibility
of a ray's
egress from
any me-
dium.

Conversely, when a ray is incident on the interior surface of the medium, at any angle less than the limiting angle whose sine is $\frac{1}{\mu}$, it will be refracted and emerge according to the law laid down in Art. 181. being bent from the perpendicular. But as the angle of incidence pcs increases, the angle of refraction PCS increases more rapidly; and when the former angle has reached the limiting value pcs'' , the transmitted ray emerges in the direction CS'' , barely grazing the external surface. If the angle of incidence be still further increased, the

angle of refraction becomes imaginary: for we have $\sin PCS = \mu \times \sin pcs$, and if $\sin pcs > \frac{1}{\mu}$, the sine

When the
ray cannot
emerge it is
reflected.

of PCS must be greater than unity. This shows that the ray cannot emerge; but it does not inform us what becomes of it. To ascertain this, we must have recourse to experiment; from which we learn, that after this limit is passed, the ray, instead of being refracted out of the medium, is turned back and totally reflected within it, making the angle of reflexion $pcs''' = pcs''$.

Light.

When the ray is incident on the exterior surface of the medium, a portion is reflected (R) and the remainder (r) refracted. The ratio of R to r is smallest at a perpendicular incidence, and increases regularly till the incidence becomes 90°; but even at extreme obliquities, and when the incident ray just grazes the surface, the reflexion is never total, or nearly total, a very considerable portion being always intromitted. On the other hand, when the ray is incident on the interior surface, the reflected portion (R) increases regularly, with a very

Part I.

184.

This reflexion total.

moderate rate of increase, till the angle of incidence becomes equal to the critical angle, whose sine is $\frac{1}{\mu}$; when

it suddenly, and, as it were, *per saltum*, attains the whole amount of the incident light, and the refracted portion (r) becomes zero. This sudden change from the law of refraction to that of reflexion—this breach of continuity, as it were, is one of the most curious and interesting phenomena in Optics, and (as we shall see hereafter) is connected with the most important points in the theory of light.

The reflexion thus obtained, being total, far surpasses in brilliancy what can be obtained by any other means; from quicksilver, for instance, or from the most highly polished metals. It may be familiarly shown by filling a glass (a common drinking-glass) with water, and holding it above the level of the eye, (as in fig. 24, No. 2.) If we then look obliquely upwards in the direction E a c, we shall see the whole surface shining like polished silver, with a strong metallic reflexion; and any object, as a spoon, A C B, for instance, immersed in it will have its immersed part C B reflected on the surface as on a mirror, but with a brightness far superior to what any mirror would afford. This property of internal reflexion is employed to great advantage in the camera lucida, and might be turned to important uses in other optical instruments, especially in the Newtonian telescope, to obviate the loss of light in the second reflexion, of which more hereafter.

185.

Experiment illustrating this total reflexion. Fig. 24, No. 2.

Some curious consequences follow from this, as to vision under water. An eye placed under perfectly still water (that of a fish, or of a diver) will see external objects only through a circular aperture (as it were) of 96° 55' 20" in diameter overhead. But all objects down to the horizon will be visible in this space; and those near the horizon much distorted and contracted in dimensions, especially in height. Beyond the limits of this circle will be seen the bottom of the water, and all subaqueous objects, reflected, and as vividly depicted as by direct vision. In addition to these peculiarities, the circular space above-mentioned will appear surrounded with a perpetual rainbow, of faint but delicate colours, the cause of which we shall take occasion to explain further on. But we need not immerse ourselves in water to see, at least, a part of these phenomena. We actually live under an ocean of air, a feebly refracting medium, it is true, in comparison with water; and our vision of external objects near the horizon is modified accordingly. They are seen distorted from their true form, and contracted in their vertical dimensions; thus the sun at setting, instead of appearing circular, assumes an elliptical, or rather compressed figure, the lower half being more flattened than the upper, and this change of figure is considerable enough to be very evident to even a careless spectator. The spherical form of the atmosphere, and its decrease of density in the higher regions, however, prevent the rest of the appearances above described from being seen in it.

186.

Appearances of external objects to a spectator under water

Elliptical form of setting sun explained.

If a medium be bounded by parallel plane surfaces, a ray refracted through it will have its final direction after both refractions the same as before entering the medium.

187.

Refraction through parallel surfaces. Fig. 25, No. 2.

Let A B, D F be the parallel surfaces of the medium, and S C E T a ray refracted through it, P C p, Q E q, perpendiculars to the surfaces at C and E, then we have

$$\begin{aligned} \sin S C P : \sin p C E (= \sin C E Q) &:: \mu : 1 \\ \sin C E Q : \sin q C T &:: 1 : \mu \end{aligned} \quad \text{and, compounding these proportions,}$$

$$\sin S C P : \sin q E T, \text{ and therefore } S C P = q E T, \text{ and the ray } E T \text{ is parallel to } S C.$$

This proposition may be proved experimentally, by placing the plane glass of a sextant (unsilvered) before the object-glass of a telescope directed to a distant object, or before the naked eye, and inclining it at any angle to the visual ray. The apparent place of the object will be unchanged.

Experimental proof.

Experiment. Let a plate of glass, or any other transparent medium, be placed parallel to the horizon, and on it let any transparent fluid be poured, so as to form a compound medium consisting of two media of different refractive indices, in contact, and bounded by parallel planes; and let an object above this combination, a star, for instance, be viewed by an eye placed below it, or through a telescope. It will be found to appear precisely in the same situation as if the media were removed, whatever be the altitude of the object, or star. It follows from this, that a ray S B (fig. 26, No. 2) incident on such a combination of media, A F and D I, as described, will emerge in a direction H T parallel to the incident ray S B.

188.

Refraction at the common surface of two media in contact. Fig. 26, No. 2.

Proposition. Let there be any two media (No. 1 and 2) whose respective indices of refraction from a vacuum into each are μ and μ' . Then if these media are brought into perfect contact, (such as that of a fluid with a solid, or of two fluids with one another,) the refraction from either of them (No. 1) into the other (No. 2)

189.

Law of refraction from one medium into another.

will be the same as that from a vacuum into a medium, whose index of refraction is $\frac{\mu'}{\mu}$, the index of refraction of the second medium divided by that of the first.

Let D E F (fig. 26, No. 2) be the common surface of the two media, and let them be formed into parallel plates A F, D I, as in the experiment last described; then any ray S B incident at any angle on the surface A C will emerge at G I in a direction H T parallel to S B. Let B E H be its path within the media, and draw the perpendiculars P B p, Q E q, R H r, then

$$\sin S B P : \sin E B p = \sin B E Q :: \mu : 1$$

$$\sin R H E = \sin q E H : \sin r H T = \sin P B S :: 1 : \mu,$$

and, compounding these proportions

$$\sin H E q : \sin B E Q :: u : u'; \quad \frac{\sin B E Q}{\sin H E q} = \frac{\mu'}{u}.$$

Absolute and relative indices of refraction.

But $B E Q$ is the angle of incidence, and $H E q$ that of refraction, at the common surface of the media, consequently the *relative* index, or index of refraction from the first into the second, is equal to the quotient

$\frac{\mu'}{\mu}$ of the *absolute* indices μ' , μ , of the second and first, or their indices of refraction from vacuum.

190. This demonstration, it is true, holds good only for the case when the angles of incidence and refraction at the common surface are both less than the limits of the angles of refraction from vacuum into each medium. If they exceed these limits, the proposition however still holds good, as may be shown by direct measures of the angles of incidence and refraction in any proposed case. At present, therefore, we must receive it as an experimental truth.

191. *Example.* Required the ratio of the sine of incidence to that of refraction out of water into flint glass. The refractive index of flint glass is 1.60, and that of water 1.336, therefore the refractive ratio required is

$$\frac{1.60}{1.336} = 1.197.$$

192. If the index $\mu = -1$, the general law of refraction coincides with that of reflexion. Thus all the cases of reflexion, as far as the direction of the reflected ray is concerned, are included in those of refraction.

Of the Ordinary Refraction of Light through a System of Plane Surfaces, and of Refraction through Prisms.

193. *Definition.* In Optics, any medium having two plane surfaces, through which light may be transmitted, inclined to each other at any angle, is called a *prism*.

194. *Definition.* The *edge* of the prism is the line, real or imaginary, in which the two plane surfaces meet, or would meet if produced.

195. *Definition.* The *refracting angle* of the prism is the angle on which its two plane surfaces are inclined to each other.

196. *Definition.* The *faces* of a prism are the two plane surfaces.

197. *Definition.* The plane perpendicular to both surfaces, and therefore to the edge of a prism, is called the principal section of the prism, or of the two surfaces. This expression has been used in this general sense already, under the head of reflexion.

To determine the direction of a Ray after Refraction through any System of Plane Surfaces.

198. General problem of the refraction through any system of plane surfaces. Fig. 27.

Construction. Since the direction of the ray is the same whether refracted at the given surfaces, or at others parallel to them, conceive surfaces parallel to the given ones, all passing through one point, and from this point, but wholly exterior to the refracting media, let perpendiculars $C P$, $C P'$, $C P''$, &c. be drawn to each of the surfaces, (fig. 27.) Let $S C$ be the direction of the incident ray. Between $C P$ and $C S'$ draw $C S'$ in the plane

$S C P$, so that $\sin P C S' = \frac{1}{\mu} \cdot \sin P C S$, μ being the index of refraction of the first medium from the medium

in which the ray originally moved, which we will at present suppose a vacuum, then will $S' C$ be the direction of the ray after the first refraction. Again, let $\mu' =$ the relative refractive index of the second medium *out of the first*, or $\mu \mu' =$ its absolute refractive index from a vacuum; draw $C S''$ in the plane $S' C P'$ so as to make

$\sin P' C S'' = \frac{1}{\mu'} \cdot \sin P' C S'$, then will $S'' C$ be the direction of the ray after the second refraction, and so on.

199. *General analysis.* Let $a = S C P$ the first angle of incidence, $a' = S' C P'$ the angle of incidence on the second surface, $I = P C P'$ the inclination of the two first surfaces to each other, and putting, moreover,

$\theta = P S' P' =$ the angle which the planes of the first and second refraction make with each other.

$\psi = S P P' =$ the angle made by the plane of the first refraction with the principal section of the two first refracting surfaces.

$\phi = S' P' P =$ the angle made by the plane of the second refraction with the same principal section.

$\rho = P C S'$ the first, and $\rho' = P' C S''$ the second angle of refraction.

$D = S C S''$ the deviation after the second refraction.

Lign.

We have (conceiving $SS'S''PP'$ to be a portion of a spherical surface having C for its centre) in the spherical triangle $S'PP'$ given PS' , PP' , and the included angle, required $S'P'$ and the angles $PS'P'$, $PP'S'$; and, again, in the triangle $SS'S''$ given SS' , SS'' and the angle $SS'S''$, required SS'' the deviation. Or, in symbols, since ρ and ρ' are the angles of refraction corresponding to the angles of incidence α , α' , and the indices of refraction μ , μ' ,

Part I.

$$(B) \begin{cases} \sin \alpha = \mu \cdot \sin \rho \\ \cos \alpha' = \cos \rho \cdot \cos I + \sin \rho \cdot \sin I \cdot \cos \psi \\ \sin \alpha' = \mu' \cdot \sin \rho' \\ \sin \alpha' \cdot \sin \theta = \sin I \cdot \sin \psi \\ \sin \alpha' \cdot \sin \phi = \sin \rho \cdot \sin \psi \\ \cos D = \cos (\alpha - \rho) \cdot \cos (\alpha' - \rho') - \sin (\alpha - \rho) \cdot \sin (\alpha' - \rho') \cdot \cos \theta. \end{cases}$$

From these equations, which, however, are rather more involved than in the case of reflexion, (Art. 99, equation A,) we may determine in all circumstances the course of the ray after the second refraction; and, in like manner, as in the case of reflexion, of any of the eleven quantities α , α' , ρ , ρ' , μ , μ' , I , θ , ϕ , ψ , D , any five being given the remaining six may be found, and we may then go on to the next refraction, and so on as far as we please. It is needless to observe, however, that, except in particular cases, the complication of the formula becomes exceedingly embarrassing when more than two refractions are considered. Such is the general analysis of the problem; but the importance of it in optical researches requires an examination in some detail of a variety of particular cases.

200.

Case 1. When two plane surfaces only are concerned, at both of which the refractions are made in one plane, viz. that of the principal section of the two planes, or of the prism which they include.

201.

Let the ray SC (fig. 28) be incident from vacuum on any refracting surface AC of a prism CAD , in the plane of its principal section; draw PC perpendicular to that surface, and draw CS' so that $\sin PCS' : \sin PCS :: 1 : \mu$, then will $S'C$ be the direction of the refracted ray CD . Again, draw CP' perpendicular to AD , and take the angle $S'CP'$, such that $\sin P'CS'' : \sin P'CS' :: 1 : \mu'$, μ' being the relative index of refraction from the medium ACD into the medium ADE , then will $S''C$ be parallel to the ray after the second refraction; draw, therefore, DE parallel to $S'C$, and DE will be the twice refracted ray. As in the general case, calling SCP , α ; $S'CP$, ρ ; $S'CP'$, α' ; $S''CP'$, ρ' ; and PCP' , I , &c.

Case 1.
When both
refractions
are made in
one plane.
Fig. 28.

we have

$$\left. \begin{aligned} \sin \alpha &= \mu \cdot \sin \rho; \quad \alpha' = I + \rho; \quad \sin \alpha' = \mu' \cdot \sin \rho' \\ \text{and} \quad \pm D &= SCP'' = \alpha - \rho' + I; \quad \theta = 0; \quad \phi = 0 \end{aligned} \right\} \quad (a)$$

The first of these equations gives ρ when μ and α are known; the second gives the value of α' when ρ is found; the third gives ρ' when α' and μ' are known; and the last exhibits the deviation D .

The sign of D is ambiguous. If we regard a deviation from the original direction towards the thicker part of the prism, or from its edge as positive, which for future use will be most convenient, we must use the lower sign

202.

or take

$$D = \rho' - I - \alpha; \quad (b)$$

but if *vice versa*, then the upper sign must be used. We shall adhere to the former notation.

Case 2. If, in case 1, we suppose the medium into which the ray emerges to be the same as that from which it originally entered the prism, (a vacuum, for example,) we have $\mu' = \frac{1}{\mu}$. This is the case of refraction

203.

Case 2.
Both refractions in one plane at the faces of a prism in vacuo.

through an ordinary prism of glass, or any transparent substance. In this case, I is the refracting angle of the prism, μ its refractive index, (its absolute refractive index if the prism be placed in vacuo, its relative, if in any other medium,) and the system of equations representing the deviation and direction of the refracted ray becomes

$$\left. \begin{aligned} \sin \alpha &= \mu \cdot \sin \rho \\ \sin \alpha' &= I + \rho \\ \sin \rho' &= \mu \cdot \sin \alpha' \\ \sin D &= \rho' - \alpha - I \end{aligned} \right\} \quad (c)$$

Corol. 1. The deviation may be expressed in another form, which it will be convenient hereafter to refer to For we have

204.

$$\begin{aligned} \sin (I + D + \alpha) &= \sin \rho = \mu \cdot \sin \alpha' = \mu \cdot \sin (I + \rho) \\ &= \mu \{ \sin \rho \cdot \cos I + \cos \rho \cdot \sin I \} \\ &= \mu \left\{ \sin \rho - 2 \sin \rho \cdot \left(\sin \frac{I}{2} \right)^2 + 2 \cdot \cos \rho \cdot \cos \frac{I}{2} \cdot \sin \frac{I}{2} \right\} \end{aligned}$$

Light. because

$$\cos I = 1 - 2 \left(\sin \frac{I}{2} \right)^2 \text{ and } \sin I = 2 \cdot \sin \frac{I}{2} \cdot \cos \frac{I}{2}.$$

Now $\mu \sin \rho = \sin \alpha$ by the first of the equations (c), hence we get (equation d)

$$\sin (I + D + \alpha) = \sin \alpha + 2 \mu \cdot \sin \frac{I}{2} \cdot \cos \left(\frac{I}{2} + \rho \right); \quad (d)$$

whence, I and α being given, and ρ calculated from the equation $\sin \rho = \frac{1}{\mu} \sin \alpha$, D is easily had.

205. *Corol. 2.* If $\alpha = 0$, or if the ray be intromitted perpendicularly into the first surface, we have also $\rho = 0$, and the expression (d) becomes simply

$$\sin (I + D) = \mu \cdot \sin I, \quad (e)$$

whence also

$$\mu = \frac{\sin (I + D)}{\sin I}; \quad (f)$$

Thus we see that if $\mu \cdot \sin I > 1$, or if I , the angle of the prism, be greater than $\sin^{-1} \frac{1}{\mu}$,* the critical angle, or the least angle of total internal reflexion, the deviation becomes imaginary, and the ray cannot be transmitted at such an incidence.

206. *Corol. 3.* The equation (f) affords a direct method of determining by experiment the refractive index of any medium which can be formed into a prism. We have only to measure the angle of the prism, and the deviation of a ray intromitted perpendicularly to one of its faces. Thus I and D being given by observation, μ is known. This is not, however, the most convenient way; a better will soon appear.

Definitions. One medium in Optics is said to be *denser* or *rarer* than another, according as a ray in passing from the former into the latter is bent *towards* or *from* the perpendicular. When we speak of the *refractive density* of a medium, we mean that quality by which it turns the ray more or less from its course towards the perpendicular (from a vacuum,) and whose numerical measure is the quantity μ the index of refraction.

207. *Proposition.* Given the index of refraction of a prism, to find the limit of its refracting angle, or that which if exceeded, no ray can be directly transmitted through both its faces.

208. *This limit is evidently that value of I which just renders the angle of refraction ρ' imaginary for all angles of incidence on the first surface, or for all values of α , that is, which renders in all cases*

$$\mu \cdot \sin \{ I + \rho \} - 1 \text{ positive,}$$

or, $\sin (I + \rho) - \frac{1}{\mu}$ positive; that is, (since $I + \rho$ cannot exceed 90°) which renders in all cases $I + \rho - \sin^{-1} \left(\frac{1}{\mu} \right)$ positive. Now $\rho = \sin^{-1} \frac{\sin \alpha}{\mu}$, and consequently the value of α least favourable to a positive value of the function under consideration is -90° , which makes $\rho = -\sin^{-1} \left(\frac{1}{\mu} \right)$, its greatest negative value. Consequently, in order that no second refraction shall take place, I must at least be such that $I - 2 \sin^{-1} \left(\frac{1}{\mu} \right)$ shall be positive; that is, I , the angle of inclination of the faces of the prism to each other,

or as it is briefly expressed, the *angle of the prism*, must be at least twice the maximum angle of internal incidence.

209. For example, if $\mu = 2$, I must be at least 60° . In this case no ray can be transmitted *directly* through an equilateral prism of the medium in question.

210. *Corol. 4.* If $\mu > 1$, or if the prism be denser than the surrounding medium, $\mu \cdot \sin I$ is $> \sin I$ and $\sin^{-1} (\mu \cdot \sin I) > I$, so that the value of D (equation (d), Art. 204) is positive, or the ray is bent towards the thicker part of the prism, (see fig. 29.) If $\mu < 1$, or the prism be rarer than the medium, the contrary is the case, (see fig. 30.)

211. *Problem.* The same case being supposed, (that of a prism in vacuo, or in a medium of equal density on both sides,) required to find in what direction a ray must be incident on its first surface so as to undergo the least possible deviation.

Since $D = \rho' - \alpha - I$; (c) Art. 203, and by the condition of the minimum, $dD = 0$, we must have

$$d\rho' = d\alpha.$$

Now the equations (c) give by differentiation

$$d\alpha \cdot \cos \alpha = \mu d\rho \cdot \cos \rho; \quad d\alpha' = d\rho; \quad d\rho' \cdot \cos \rho' = \mu d\alpha \cdot \cos \alpha,$$

that is

$$d\rho' \cdot \cos \rho' = \mu d\rho \cdot \cos \alpha' = d\alpha \cdot \frac{\mu \alpha \cdot \alpha'}{\cos \rho},$$

* The reader will observe, that by the expression $\sin^{-1} \frac{1}{\mu}$ is meant what in most books would be expressed by $\arcsin \frac{1}{\mu}$.

light

or

$$\frac{d\rho}{da} (=1) = \frac{\cos a \cdot \cos a'}{\cos \rho \cdot \cos \rho'}; \text{ or } \cos a \cdot \cos a' = \cos \rho \cdot \cos \rho'.$$

That is, squaring,

$$(1 - \sin a^2) (1 - \sin a'^2) = (1 - \sin \rho^2) (1 - \sin \rho'^2)$$

in which, for $\sin a$ and $\sin \rho'$ writing their equals, $\mu' \cdot \sin \rho$ and $\mu \cdot \sin a'$, we get

$$\frac{1 - \mu^2 \cdot \sin \rho^2}{1 - \sin \rho^2} = \frac{1 - \mu'^2 \cdot \sin a'^2}{1 - \sin a'^2},$$

which gives, on reduction, simply $\sin \rho^2 = \sin a'^2$, and therefore $\rho = \pm a'$, that is $I + \rho = I \pm a'$, or $a' = I \pm a'$. The upper sign is unsatisfactory, as it would give $I = 0$. The lower therefore must be taken,

which gives $a' = \frac{I}{2}$, which satisfies the conditions of the question. We therefore have

$$a' = \frac{1}{2} I; \rho = -\frac{1}{2} I; \sin a = -\mu \cdot \sin \left(\frac{I}{2} \right); \sin \rho' = +\mu \cdot \sin \left(\frac{I}{2} \right).$$

This state of things is represented in fig. 31, for the case where $\mu > 1$, or where the prism is denser than the surrounding medium, and in fig. 32, for that in which it is rarer, or $\mu < 1$. In both cases, a , being negative, Fig. 32. indicates that the incident ray must fall on the side of the perpendicular CP, from the edge A of the prism (as SC). In both cases, the equations $\rho (= PC S') = -\frac{1}{2} I (= -\frac{1}{2} PC P')$ and $a' = P' C S' = +\frac{1}{2} PC P'$, indicate that the once refracted ray S'CD bisects the angle PCP', and therefore that the portion of it CD within the prism makes equal angles with both its faces. In both cases, also, the equality of the angles a and ρ' (without reference to their signs) shows that the incident and emergent rays make equal angles with the faces of the prism, and therefore that it is of no consequence on which face the ray is first received.

Corol. 5. In this case, also, we have the actual amount of the deviation

$$D = \rho' - a - I = 2 \sin^{-1} \left(\mu \cdot \sin \frac{I}{2} \right) - I. \quad (f)$$

Hence also

$$\sin \frac{I + D}{2} = \mu \cdot \sin \frac{I}{2}.$$

Corol. 6. In the same case, I being given by direct measurement, and D by observation, of the actual minimum deviation of a ray refracted through any prism, the value of μ , its index of refraction, is given at once, for we have

$$\mu = \frac{\sin \left(\frac{I + D}{2} \right)}{\sin \left(\frac{I}{2} \right)}. \quad (g)$$

And this affords the easiest and most exact means of ascertaining the refractive index of any substance capable of being formed into a prism.

Example. A prism of silicate of lead, consisting of silica and oxide of lead, atom to atom, had its refracting angle $21^\circ 12'$. It produced a deviation of $24^\circ 46'$ at the minimum in a ray of homogeneous extreme red light: what was the refractive index for that ray?

$$I = 21^\circ 12', \frac{I}{2} = 10^\circ 36', D = 24^\circ 46', \frac{D}{2} = 12^\circ 23'$$

$$\sin \left(\frac{I}{2} + \frac{D}{2} \right) = \sin 22^\circ 59' \quad 9.59158$$

$$\sin \frac{I}{2} = \sin 10^\circ 36' \quad 9.26470$$

$$\mu = 2.123 \quad 0.32688$$

Case 3. Let us now take a somewhat more general case, viz. to find the final direction and total deviation of a ray, after any number of refractions at plane surfaces, all the refractions being performed in one plane, and, of course, all the common sections of the surfaces being supposed parallel.

Supposing (as above) I to represent the inclination of the first surface to the second; I' that of the second to the third, &c.; and I, I', &c. to be negative when the surfaces incline the contrary way from one certain side assumed as positive, taking also $\delta, \delta', \delta'', \&c. \dots \delta^{(n-1)}$ to represent the several partial bendings of the rays at the first, second, third, n th surface respectively, and the other symbols remaining as before, we have the total deviation, $D = \delta + \delta' + \dots \delta^{(n-1)}$. Now we have, since in each case $\theta = 180^\circ$,

212.

Expression for the minimum deviation.

213.

Another mode of determining the index of refraction of a prism by experiment.

214.

Example.

215.

Deviation of a ray after several refractions is one plane.

Light.

Part I.

$$\begin{aligned}\sin a &= \mu \cdot \sin \rho; \quad a' = \rho + I; \quad \mu' \cdot \sin \rho' = \sin a'; \quad \delta = a - \rho; \\ \sin a' &= \mu' \cdot \sin \rho'; \quad a'' = \rho' + I'; \quad \mu'' \cdot \sin \rho'' = \sin a''; \quad \delta' = a' - \rho'; \quad \&c. \&c.\end{aligned}$$

Hence we get (supposing n to represent the number of surfaces)

$$\begin{aligned}\sin \rho &= \frac{1}{\mu} \cdot \sin a \\ \sin \rho' &= \frac{1}{\mu'} \cdot \sin (I + \rho) \\ \sin \rho'' &= \frac{1}{\mu''} \cdot \sin (I' + \rho') \\ &\dots \dots \dots \\ \sin \rho^{(n-1)} &= \frac{1}{\mu^{(n-1)}} \cdot \sin (I^{(n-2)} + \rho^{(n-2)})\end{aligned}$$

whence the series of values $\rho, \rho', \&c.$ may be continued to the end. These determined, we get $a, a', \&c.$ by the equations

$$a = a; \quad a' = \rho + I; \quad a'' = \rho' + I'; \quad \dots a^{(n-1)} = \rho^{(n-2)} + I^{(n-2)},$$

and finally

$$\begin{aligned}D &= \{a + a' + \dots a^{(n-1)}\} - \{\rho + \rho' + \dots \rho^{(n-1)}\} \\ &= a + \{I + I' + \dots I^{(n-2)}\} - \rho^{(n-1)}.\end{aligned}$$

Now $I + I' + \dots I^{(n-2)}$ is the inclination of the first to the last surface, or the angle (A) of the compound prism, formed of the assemblage of them all, so that we have in general

$$D = a + A - \rho^{(n-1)} \quad (h)$$

Let us now inquire, how a ray must be incident on such a system of surfaces so that its total deviation shall be a minimum.

Since $dD = 0$ and $I, I', \&c.$ are constant, we must have

$$d a = d \rho^{(n-1)},$$

216.
Case of
minimum
deviation
after any
number of
refractions.

but

$$\left. \begin{aligned}\mu \cdot \sin \rho &= \sin a \\ \mu' \cdot \sin \rho &= \sin (\rho + I) \\ \&c.\end{aligned} \right\} \therefore \begin{cases} \mu d \rho \cdot \cos \rho = d a \cdot \cos a \\ \mu' d \rho' \cdot \cos \rho' = d \rho \cdot \cos (\rho + I) \\ \mu^{(n-1)} d \rho^{(n-1)} \cdot \cos \rho^{(n-1)} = d \rho^{(n-2)} \cdot \cos (\rho^{(n-2)} + I^{(n-2)}) \end{cases}$$

and multiplying all these equations together

$$\mu \mu \dots \mu^{(n-1)} \cdot \cos \rho \cos \rho' \dots \cos \rho^{(n-1)} \cdot \frac{d \rho^{(n-1)}}{d a} = \cos a \cdot \cos (\rho + I) \dots \cos (\rho^{(n-2)} + I^{(n-2)})$$

or simply

$$\mu \mu' \dots \mu^{(n-1)} \cdot \cos \rho \cdot \cos \rho' \dots \cos \rho^{(n-1)} = \cos a \cdot \cos a' \dots \cos a^{(n-1)}; \quad (i)$$

this equation, combined with the relations already stated, between the successive values of ρ and those of a , afford a solution of the problem; but the final equations to which it leads are of great complexity and high dimensions. Thus, in the case of only three refractions, the final equation in $\sin \rho$ or $\sin \rho'$, &c. rises to the sixteenth degree; and though its form is only that of an equation of the eighth, yet there appears no obvious substitution by which it can be brought lower. The only case where it assumes a tractable form is that of two surfaces, when the equation (i) which in general may be put under the form

$$\mu^2 \mu'^2 \dots \mu^{(n)2} (1 - \sin^2 \rho^2) (1 - \sin^2 \rho'^2), \&c. = (1 - \mu^2 \cdot \sin^2 \rho^2) (1 - \mu'^2 \cdot \sin^2 \rho'^2), \&c. \quad (j)$$

reduces itself by putting

$$\sin \rho^2 = x, \text{ and } \sin \rho'^2 = y,$$

to

$$(\mu^2 \mu'^2 - 1) - \mu^2 (\mu'^2 - 1) x - \mu'^2 (\mu^2 - 1) y = 0,$$

which, combined with the equation

$$\mu' \cdot \sin \rho' = \sin (\rho + I)$$

or

$$(\mu'^2 y + x - \sin^2 I)^2 = 4 \mu'^2 \cdot \cos^2 I \cdot x y,$$

gives a final equation of a quadratic form for determining x or y , and which in the particular case of $\mu \mu' = 1$, or when the second refraction is made into the same medium in which the ray originally moved before its first incidence, gives the same result we have already found for that case by a similar process. Meanwhile, though we may not be able to resolve the final equations in the general case, the equation (j) affords a criterion of the state of minimum deviation which may prove useful in a variety of cases.

Light.

Case 4. When the planes of the first and second refraction are at right angles to each other, required the relations arising from this condition.

In this case we have $\theta = 90^\circ$, $\cos \theta = 0$, $\sin \theta = 1$, so that the general equation (B, 199) becomes

$$\left. \begin{aligned} \sin a &= \mu \cdot \sin \rho \\ \sin a' &= \mu' \cdot \sin \rho' \\ \sin a' &= \sin I \cdot \sin \psi \end{aligned} \right\} \text{and } \cos a' = \cos \rho \cdot \cos I + \sin \rho \cdot \sin I \cdot \cos \psi.$$

The last of these equations, by transposition and squaring, becomes

$$\cos a'^2 - 2 \cdot \cos a' \cdot \cos \rho \cdot \cos I + \cos \rho^2 \cdot \cos I^2 = \sin \rho^2 \cdot \sin I^2 (1 - \sin \psi^2)$$

in which, substituting for $\sin \psi$ its value $\frac{\sin a'}{\sin I}$ deduced from the third equation, and reducing as much as possible, we obtain

$$\cos a'^2 \cdot \cos \rho^2 - 2 \cdot \cos a' \cdot \cos \rho \cdot \cos I + \cos I^2 = 0,$$

which, being a complete square, gives simply

$$\cos \rho \cdot \cos a' = \cos I. \quad (k)$$

This answers to the equation $\cos a \cdot \cos a' = \cos I$, obtained, on the same hypothesis, in the case of reflexion (104); for since the latter case is included in the case of refraction, by putting $\mu = -1$ (Art. 192) we have then $a = -\rho$ and $\cos \rho = \cos a$.

Corol. 1. If i and i' be the inclinations to the first and second surfaces respectively of that part of the ray which lies between the surfaces, we have

$$i = 90^\circ - \rho \quad \text{and} \quad i' = 90^\circ - a',$$

so that the equation above found, gives

$$\sin i \cdot \sin i' = \cos I,$$

or the product of the sines of the inclination of the ray between the surfaces to either surface is equal to the cosine of the inclination of the two surfaces. The same relation may be expressed otherwise, thus: if we suppose the ray to pass both ways from within, out of the prism, the product of the cosines of its interior incidences on the two surfaces is equal to the cosine of their inclination to each other. In this way of stating it, the case of reflexion is included.

Corol. 2. We have also in the present case

$$\sin \rho = \frac{1}{\mu} \cdot \sin a; \quad \sin a' = \sqrt{\frac{\mu^2 \cdot \sin I^2 - \sin a^2}{\mu^2 - \sin a^2}}; \quad \sin \rho' = \frac{1}{\mu'} \sqrt{\frac{\mu'^2 \cdot \sin I^2 - \sin a'^2}{\mu'^2 - \sin a'^2}}$$

and

$$\cos D = \cos (a - \rho) \cdot \cos (a' - \rho')$$

so that a being given, all the rest become known. The last equation corresponds to the equation $\cos D = \cos 2a \cdot \cos 2a'$ in the case of reflexion.

§ VII. Of Ordinary Refraction at Curved Surfaces, and of Diacaustics, or Caustics by Refraction.

The refraction at a curved surface being the same as at a plane, a tangent at the point of incidence, if we know the nature of the surface, we may investigate, by the rules of refraction at plane surfaces, combined with the relations expressed by the equation of the surface, in all cases, the course of the refracted ray. We shall confine ourselves to the simple case of a surface of revolution, having the radiant point in the axis.

Proposition. Given a radiant point in the axis of any refracting surface of revolution, required the focus of any annulus of the surface.

Let CP be the curve, Q the radiant point, QqN the axis, PM any ordinate, PN a normal, and Pq or qP the direction of the refracted ray, and therefore q the focus of the annulus described by the revolution of P. Then if we put μ for the refractive index, and, assuming Q for the origin of the coordinates, put QM = x,

MP = y, $r = \sqrt{x^2 + y^2}$, $p = \frac{dy}{dx}$, we have

$$\sin QPM = \frac{x}{r}; \quad \cos QPM = \frac{y}{r};$$

Part I.

217.

Case when the planes of the first and second refraction are at right angles.

218.

219

220.

221.

General investigation of the focus of any curve surface of revolution. Fig. 33.

$$\sin NPM = \frac{p}{\sqrt{1+p^2}}; \quad \cos NPM = \frac{1}{\sqrt{1+p^2}}$$

consequently

$$\begin{aligned} \sin NPQ &= \sin QPM \cdot \cos NPM + \sin NPM \cdot \cos QPM \\ &= \frac{x + py}{r \cdot \sqrt{1+p^2}}, \end{aligned}$$

and therefore

$$\sin NPq = \frac{1}{\mu} \cdot \sin NPQ = \frac{x + py}{\mu r \sqrt{1+p^2}};$$

and

$$\cos NPq = \frac{Z}{\mu r \sqrt{1+p^2}}, \quad \text{if we put } Z = \sqrt{\mu^2 r^2 (1+p^2) - (x+py)^2}; \quad (a)$$

consequently since

$$MPq = NPq + NPM, \quad \text{we get}$$

$$\sin MPq = \frac{x + py + pZ}{\mu r (1+p^2)}; \quad \text{and } \cos MPq = \frac{-p(x+py) + Z}{\mu r (1+p^2)};$$

whence

$$\tan MPq = \frac{\sin MPq}{\cos MPq} = \frac{x + py + pZ}{-p(x+py) + Z}.$$

Now we have

$$Mq = PM \cdot \tan MPq = y \cdot \tan MPq = \frac{y(pZ + (x+py))}{Z - p(x+py)}; \quad (b)$$

consequently

$$Qq = x + y \cdot \tan MPq = (x+py) \cdot \frac{px - y - Z}{p(x+py) - Z}. \quad (c)$$

222. *Corol. 1.* If we put s for the arc CP of the curve, we have, since $r dr = x dx + y dy = dx(x+py)$,

$$Z = \sqrt{\mu^2 r^2 \cdot \left(\frac{ds}{dx}\right)^2 - \left(\frac{r dr}{dx}\right)^2} = r \cdot \sqrt{\mu^2 \left(\frac{ds}{dx}\right)^2 - \left(\frac{dr}{dx}\right)^2}; \quad (d)$$

223. *Corol. 2.* If $\mu = -1$, in which case the refraction becomes a reflexion, we have

$Z = \sqrt{r^2 (1+p^2) - (x+py)^2} = y - px$, writing for r^2 its value $x^2 + y^2$; so that the general value above found for Qq reduces itself to

$$Qq = 2 \cdot \frac{(x+py)(px-y)}{2px-y(1-p^2)},$$

which is the same as that found in (b) Art. 109, in the case of reflexion.

224. *Corol. 3.* If we put $P = \tan MqP = \cotan MPq = \frac{1}{\tan MPq}$,

we have

$$P = \frac{-p(x+py) + Z}{x + py + pZ}, \quad (e)$$

and the equation of the refracted ray, if X and Y be its coordinates, (Q being their origin) will be (since Y lies on the opposite side of the curve from Q)

$$Y - y = -P \cdot (X - x) \quad (f)$$

225 In the case of parallel rays these expressions become (by putting first $x + a$ for x , and then making a infinite)

$$\left. \begin{aligned} Z &= a \sqrt{\mu^2 (1+p^2) - 1} \\ P &= \frac{-p + \sqrt{\mu^2 (1+p^2) - 1}}{1 + p \sqrt{\mu^2 (1+p^2) - 1}} \end{aligned} \right\}, \quad (g)$$

$$Aq = x + y \cdot \frac{1 + p \sqrt{\mu^2 (1+p^2) - 1}}{\sqrt{\mu^2 (1+p^2) - 1} - p} \quad (h)$$

§ VIII. Of Caustics by Refraction, or Diacaustics.

The theory of Diacaustics is in all respects analogous to that of Catacaustics already explained. To find the coordinates X and Y of the point in the diacaustic which corresponds to the point P in the refracting curve, we have only to regard the equation (f) and its differential with respect to x, y , and p alone, as subsisting together, and we get the necessary equations for determining X and Y in terms of x, y , as in the case of reflexion, and

$$\text{these are} \quad X = x + \frac{P+p}{dP} dx; \quad Y = y - P \cdot \frac{P+p}{dP} dx; \quad (i)$$

the only difference is in the signs and in the value of P , which, instead of the formula (c , Art. 110,) is here expressed by the more complicated function (e , Art. 223,) and the equation of the diacaustic will be obtained as before by eliminating all but X, Y from these.

It is, evident, moreover, that if we suppose, as in the theory of Catacaustics, $M = \frac{P+p}{dP} dx$; and put S for the length of the caustic, and f for the line $P y$, we shall have, exactly as in that theory,

$$f = M \cdot \sqrt{1+P^2}; \quad -P = \frac{dY}{dX}; \quad \text{and } dS = df + dx \cdot \frac{1-Pp}{\sqrt{1+P^2}}.$$

See Art. 139, 143, 144.

Now we have, substituting for P its value (e),

$$\sqrt{1+P^2} = \frac{\mu r (1+p^2)}{x+py+pZ}; \quad 1-Pp = \frac{(x+py)(1+p^2)}{x+py+pZ}; \quad (k)$$

and consequently the value of dS becomes

$$dS = df + \frac{(x+py) dx}{\mu r} = df + \frac{dr}{\mu}, \text{ because } (x+py) dx = r dr,$$

and integrating

$$S = \text{const} + f + \frac{r}{\mu};$$

so that we have, finally, (fig. 34.)

$$\text{arc } Fy = (CF - Py) + \frac{1}{\mu} (QC - QP). \quad (l)$$

Fig. 34.

In the case of reflexion, $\mu = -1$, but at the same time the sign of f is negative, because in this case the reflected ray lies on the same side of the point of incidence with the incident one; thus both terms of the formula change their sign, and this expression coincides with that found Art. 144.

In the case of parallel rays, we must use the value of P found in Art. 225, equation (g). Putting $q = \frac{dp}{dx}$, and executing the operations, we find, then,

$$\left. \begin{aligned} X &= x - \frac{1}{P} \cdot \frac{\mu^2(1+p^2) - 1}{\mu^2 q} \\ Y &= y + \frac{\mu^2(1+p^2) - 1}{\mu^2 q} \end{aligned} \right\}. \quad (m)$$

Corol. If we suppose $\mu = \infty$, or the refractive power infinite, the refracted ray will coincide with the normal, and the caustic will be identical with the evolute; and it is evident that the expressions (m), when $\mu = \infty$, resolve themselves into the well-known values of the coordinates of the evolute.

If the rays incident on the refracting curve do not diverge from one point, but be all tangents to a curve $V'V''$, (fig. 35,) we must put $x-a$ for x in the value of P , (eq. (e) Art. 224;) and fix the origin of the coordinates at A , putting $AQ = a$; and if, then, we regard a as variable according to any given law, (or regard $x-a$ at once as a given function of x), and take the differential of P on this supposition, the equations (i) still hold good, and suffice to define the caustic.

Problem. The radiant point and refractive index of a medium being given, to determine the nature of the curve surface which shall refract all the rays to one point.

Here we are required to find the relation between x and y , so as to make Qq invariable. Let $Qq = c$, and we have

$$c = (x+py) \cdot \frac{px-y-Z}{p(x+py)-Z}; \quad \text{where } Z = \sqrt{\mu^2(x^2+y^2)(1+p^2) - (x+py)^2}.$$

This equation gives

$$(x+py)(p(x-c)-y) = Z(x-c+py).$$

Light. Squaring both sides, and substituting for Z its value, we get

$$(x + py)^2 \{ (p(x - c) - y)^2 + (x - c + py)^2 \} = (x - c + py)^2 \cdot \mu^2 (x^2 + y^2) (1 + p^2)$$

which, on executing the operations indicated on the left hand side, becomes totally divisible by $1 + p^2$, and reduces itself to

$$(x + py)^2 (y^2 + (x - c)^2) = \mu^2 (x - c + py)^2 (y^2 + x^2)$$

that is, putting for p its value $\frac{dy}{dx}$, multiplying by dx^2 , and extracting the square root,

$$\frac{xdx + ydy}{\sqrt{x^2 + y^2}} = \mu \cdot \frac{(x - c)dx + ydy}{\sqrt{(x - c)^2 + y^2}},$$

and integrating (each side being a complete differential)

$$\sqrt{x^2 + y^2} = b + \mu \cdot \sqrt{(x - c)^2 + y^2}, \quad (n)$$

which is the equation of the curve required, and belongs, generally, to a curve of the fourth order.

233. *Corol. 1.* About Q , (fig. 36.) with any radius QA , arbitrarily assumed, describe a circle $ABDE$, then if CP be the refracting curve, and we put $QA = b$, we have $QP = \sqrt{x^2 + y^2}$, $Pq = \sqrt{(x - c)^2 + y^2}$, and the nature of the curve is expressed by the property

$$BP = \mu \cdot Pq, \quad \text{or, } BP : Pq :: \mu : 1.$$

234. *Corol. 2.* If $b = 0$, or the circle ABE be infinitely small, we have $QP : Pq :: \mu : 1$, which is a well known property of the circle. In fact, in this case we have simply

$$x^2 + y^2 = \mu^2 \{ (x - c)^2 + y^2 \}$$

In this, if we change the origin of the coordinates by writing $x + \frac{\mu^2}{\mu^2 - 1} c$ for x , we find

$$y^2 = \left(\frac{\mu}{\mu^2 - 1} c \right)^2 - x^2.$$

The radius of the circle, therefore, is equal to $\frac{\mu}{\mu^2 - 1} \times Qq$, and the distance of its centre from the radiant

point is $\frac{\mu^2}{\mu^2 - 1} \times Qq$. Take therefore any circle HPC whose centre is E , (fig. 37,) and two points Q, q , such that $QE = \mu \times EC$ and $QC : Cq :: \mu : 1$. Then if rays diverge from Q , and fall on the surface PH beyond the centre, they will, after refraction into the medium M , all diverge from q .

235. *Corol. 3.* If $\mu = -1$, the equation (n) , when freed from radicals, is only of the second degree between x and y , and therefore belongs to a conic section. On executing the reduction we get

$$y^2 = \left(1 - \frac{c^2}{b^2} \right) \left(\left(\frac{b}{2} \right)^2 - \left(x - \frac{c}{2} \right)^2 \right),$$

which shows that the radiant point Q is in one focus and q in the other, which is the same result as that before found by a different mode of integration.

236. *Corol. 4.* When Q is infinitely distant, and the rays are parallel, we must shift the origin of the coordinates for parallel from Q to q , by putting $c - x$ for x , and afterwards supposing c infinite. This gives

$$\sqrt{c^2 - 2cx + x^2 + y^2} = b + \mu \sqrt{x^2 + y^2}.$$

Developing the first term in a descending series, we find

$$(c - b) - x + \frac{x^2 + y^2}{2c^2} + \&c. = \mu \cdot \sqrt{x^2 + y^2}.$$

Let $c - b = h$, which, since b is arbitrary, is equally general, and may represent any finite quantity, then, as c increases and at length becomes infinite, this equation becomes ultimately

$$h - x = \mu \sqrt{x^2 + y^2}.$$

Let CP be a conic section, q its focus, and AB its directrix, $qM = x$, and $PM = y$, then will $QP = h - x$ if we take $qA = h$, and the above equation we see expresses that well known property of a conic section, in virtue of which $QP : Pq$ in a constant ratio, ($\mu : 1$.)

237. *Corol. 5.* The curve is an ellipse when $QP > Pq$, or when the ray is incident from a rarer on a denser medium, and an hyperbola in the contrary case. If $QP = Pq$, the curve is a parabola; in this case $\mu = 1$, and the rays converge to the focus at an infinite distance, i. e. remain parallel.

238 To take a single example of the investigation of the diacaustic curve, from the general expressions above

Light delivered,—let the refracting surface be a plane, and we shall have, fixing the origin of the coordinates at the radiant point, and supposing the axis of the x perpendicular to the refracting plane A C B,

Part I.
Caustic of a plane refracting surface.

$$x = \text{constant} = Q C = a, \quad p = \frac{dy}{dx} = \infty. \quad \text{Thus we get}$$

$$Z = p \sqrt{(\mu^2 - 1) y^2 + \mu^2 a^2}; \quad P = - \frac{y}{\sqrt{(\mu^2 - 1) y^2 + \mu^2 a^2}};$$

$$\frac{dP}{dx} = - \frac{\mu^2 a^2 p}{\sqrt{(\mu^2 - 1) y^2 + \mu^2 a^2}};$$

and therefore by the equations (i) we get, substituting these values,

Fig. 39 and 40.

$$\mu^2 a^2 (a - X) = \{ (\mu^2 - 1) y^2 + \mu^2 a^2 \}^{\frac{3}{2}} \}$$

$$Y = \frac{1 - \mu^2}{\mu^2} \cdot \frac{y^3}{a^2}.$$

Eliminating y from these, we have the equation of the caustic

$$\left(\frac{a - X}{\mu a} \right)^{\frac{2}{3}} + \left(\frac{\sqrt{1 - \mu^2}}{\mu} \cdot \frac{Y}{a} \right)^{\frac{2}{3}} = 1$$

This is the equation of the evolute of a conic section whose centre is C, and focus the radiant point Q. If μ be greater than unity, or the refraction be made into a rarer medium from a denser, the conic section is an ellipse, (see fig. 39,) and in the contrary case an hyperbola, (fig. 40.)

§ IX Of the Foci of Spherical Surfaces for Central Rays.

Definitions. The *curvature* of any spherical surface is the reciprocal of its radius, or a fraction whose numerator is unity, and denominator the number of units of any scale of linear measure to which the radius is equal. 239. Curvature defined.

The *proximity* of one point to another is the reciprocal of their mutual distance, or the quotient of unity by the number of units of linear measure in that distance. 240. Proximity.

The *focal distance* of a spherical surface is the distance from the vertex, of the point to which rays converge, or from which rays diverge after refraction or reflexion. 241. Focal distance.

The *principal focal distance*, or *focal length*, is the distance from the vertex of the point to which *parallel* and *central* rays converge, or from which they diverge after refraction or reflexion. 242. Focal length.

The *power* of a surface is the reciprocal of its *principal focal distance*, or *focal length*, estimated as in the definitions of curvature and proximity. 243. Power.

Problem. To find the focus of a spherical refracting surface after one refraction, for central rays.

Here, putting a for the distance of the focus of incident rays Q, (fig. 41,) from the centre E, we have

$$(a - x)^2 + y^2 = r^2; \quad p = \frac{a - x}{y}; \quad 1 + p^2 = \frac{r^2}{y^2}; \quad x + p y = a;$$

and these substituted in the general expressions Art. 221, give

$$\left. \begin{aligned} y Z &= \sqrt{\mu^2 r^2 x^2 + (\mu^2 r^2 - a^2) y^2} \\ Q q &= a \left\{ 1 - \frac{r^2}{a(a - x) - y Z} \right\} \\ C q &= r \left\{ 1 - \frac{r a}{a(a - x) - y Z} \right\} \end{aligned} \right\} \quad (a)$$

These values of $Q q$ and $C q$ contain the rigorous solution of the problem, whatever be the *amplitude* (y) of the annulus whose focus q is, and we shall accordingly again have recourse to them. At present, however, our central ray concern being only with central rays, we must put $y = 0$, when we find $x = a - r$; $y Z = \mu r x = \mu r (a - r)$

$$Q q = a \cdot \frac{(a - r)(1 - \mu)}{a - \mu a + \mu r} \quad C q = \frac{\mu r (r - a)}{a - \mu a + \mu r} \quad (b)$$

Corol. 1. This latter is the focal distance for central rays. Now, since $a - r = Q C$, this gives the following proportion, 245.

$$\mu \cdot Q C - Q E : \mu \cdot Q C :: C E : C q. \quad (c)$$

Light. *Corol. 2.* If we suppose the focus of incident rays infinitely distant, or $a = \infty$, and take F the place of q for central rays, on that supposition, F will be the principal focus, and we shall have Part I.

246.
Focus for
parallel rays

whence we also find

$$\left. \begin{aligned} CF &= \frac{\mu r}{\mu - 1}; \quad \text{that is, } CE : CF :: \mu - 1 : \mu \\ CE : EF :: \mu - 1 : 1, \quad \text{and } CF : FE :: \mu : 1 \end{aligned} \right\} \quad (d)$$

247. These results will be expressed more conveniently for our future reference by adopting a different notation. Let, then,

$R = \frac{1}{r}$ = curvature of the surface, and let positive values of r and R correspond to the case where the centre E lies to the right of the vertex C, or in the direction in which the rays proceed.

$D = \frac{1}{QC}$ (fig. 42) = proximity of the focus of incident rays to the surface, D being regarded as positive when Q lies to the *right* of C, as in fig. 42, and as negative when to the left, as in fig. 41. Then, since $QE = a$, and since in the foregoing analysis a is regarded as positive when Q is to the *left* of E, we must have (fig. 42) $QE = -a$, and $QC = QE + EC = r - a$, so that

$$D = \frac{1}{r - a}; \quad a = \frac{1}{R} - \frac{1}{D}. \quad \text{Let also } m = \frac{1}{\mu} :$$

$$F = \frac{1}{CF} = \text{power of the surface} :$$

$$f = \frac{1}{Cq} = \text{proximity of the focus of refracted rays to the surface.}$$

Positive values of F and f , as well as of D and R , being supposed to indicate situations of the points F, f , Q, E, respectively, to the right of C, or in the direction towards which the light travels. This is, in fact, assuming for our positive case that of *converging* rays incident on a *convex* surface of a *denser* medium. We shall have, then,

$$r = \frac{1}{R}; \quad r' - a = \frac{1}{D}; \quad a = \frac{1}{R} - \frac{1}{D}; \quad \mu = \frac{1}{m}.$$

Fundamental equation for the foci of central rays.

But equation (b) gives $\frac{1}{Cq} = \frac{a + \mu(r - a)}{\mu r(r - a)}$, and substituting we shall get

$$f = (1 - m) R + m D. \quad (e)$$

This equation comprises the whole doctrine of the foci of spherical surfaces for central rays, and may be regarded as the fundamental equation in their theory.

248.
General expression for the power of any surface.

In the case of parallel rays, we have $D = 0$, whether the rays be incident from left to right, or from right to left. In either case, then, f has the same value, viz. $(1 - m) \cdot R$, and the principal focal distance F in either case is the same, being given by the equation

$$F = (1 - m) \cdot R, \quad (f')$$

which shows, moreover, that the power of any spherical surface is in the direct ratio of its curvature.

249. Hence also we have

$$f = F + m D. \quad (g)$$

250. In the case of reflexion, where $\mu = -1$, or $m = -1$, these equations become respectively

$$F = 2R; \quad f = 2R - D; \quad f = F - D. \quad (h)$$

Fundamental expressions for the central foci in case of reflexion.

Such are the expressions for the central foci in the case of a single surface. Let us now consider that of any system of spherical surfaces.

Problem. To find the central focus of any system of spherical surfaces.

Let C', C'', C''' , &c. be the surfaces. Q' the focus of rays incident on C' , Q'' that of refracted rays, or the focus of rays incident on C'' , and so on. Call also R', R'' , &c. the radii of the first, second, &c. surfaces μ', μ'' ,

&c. their refractive indices, or $\frac{\sin \text{inc}}{\sin \text{ref}}$ into each medium from that immediately preceding, $m' = \frac{1}{\mu'}$, $m'' = \frac{1}{\mu''}$,

251.
Central focus of a system of spherical surfaces investigated. Fig. 43.

&c. Also let $D' = \frac{1}{C'Q'}$, $D'' = \frac{1}{C''Q''}$, &c. and moreover let $C'Q' = t'$, $C''Q'' = t''$, &c. t' , t'' , &c. being regarded as positive when C'', C''' , &c. respectively lie to the right of C' , C'' , &c. or in the direction in which the light travels; and if we put $\frac{1}{C'Q'} = f'$, $\frac{1}{C''Q''} = f''$, &c. $F' = (1 - m') R'$, $F'' = (1 - m'') R''$, &c. we shall have by (249)

$$f' = F' + m' D'; \quad f'' = F'' + m'' D'', \text{ \&c.}; \quad (i)$$

Light. but we have also

Part I.

$$C'Q' = \frac{1}{D'}; \quad C''Q'' = \frac{1}{D''} = C'Q' - C'C'', = \frac{1}{f'} - t';$$

and so on; so that we have, besides, the following relations,

$$D' = D'; \quad D'' = \frac{f'}{1 - f't'}; \quad D''' = \frac{f'}{1 - f''t''}, \&c.; \quad (j)$$

and substituting these values of D' , D'' , &c. in the equations (i), and in each subsequent one, introducing the values of f' , f'' , &c. obtained from those preceding, we shall obtain explicit values of f' , f'' , &c. to the end.

The systems of equations (i) and (j) contain the general solution of the problem, whatever be the intervals between the surfaces. On executing the operations, however, for general values of t' , t'' , &c. the resulting expressions are found to become exceedingly complex, nor is there any way of simplifying them, the complication being in the subject, not in the method of treating it. For further information on this point, consult Lagrange, (*Sur la Théorie des Lunettes*, Berlin, Acad. 1778.) We shall here only examine the principal cases.

252

Problem. To find the focal distance of any system of spherical surfaces placed close together.

Here t' , t'' , &c. all vanish, and the equations (i) and (j) become simply

253.

Foci of a system of spherical surfaces placed close together.

$$D' = D'; \quad D'' = f'; \quad D''' = f'', \&c.; \\ f' = F' + m'D'; \quad f'' = F'' + m''D'', \&c.;$$

whence by substitution we obtain

$$f'' = F'' + m''F' + m'm'D' \\ f''' = F''' + m'''F'' + m'''m''F' + m'''m''m'D',$$

which it is easy to continue as far as we please.

Corol. 1. Let the number of surfaces be n , and let M' represent μ' , or the absolute refractive index out of vacuum into the first medium; $M'' = \mu'\mu''$, or the absolute refractive index from vacuum into the second medium, and so on; μ' , μ'' , &c. representing only the relative refractive indices from each medium into that succeeding it. Thus we shall have

254.

$$M^{(n)}f^{(n)} = D' + M'F' + M''F'' + \dots M^{(n)}F^{(n)}. \quad (k)$$

Cor. 2. For parallel rays, in whichever direction incident, we have $D' = 0$; and the principal focal length of the system, which we will call $\frac{1}{\phi^{(n)}}$, is given by the equation

255.

$$M^{(n)}\phi^{(n)} = M'F' + M''F'' + \dots M^{(n)}F^{(n)}. \quad (l)$$

Cor. 3. Hence it appears that $\phi^{(n)}$, the power of the system, or its reciprocal focal length for parallel rays, being found by the last equation, the focus for any converging or diverging rays is had at once by the equation

256.

$$M^{(n)}f^{(n)} = M^{(n)}\phi^{(n)} + D'.$$

For brevity and convenience, let us, however, modify our notation as follows: confining the accented letters to the several individual surfaces of which the system consists, let the unaccented ones be conceived to relate to their combined action as a system. Thus, F' , F'' , ..., $F^{(n)}$ representing the individual powers of the respective surfaces; let F , without an accent, denote the resulting power of the system. In this view D' may be used indifferently; accented, as relating to the incidence on the first surface; or unaccented, as expressing the proximity of the focus of incident rays to the vertex of the whole system. Similarly, $M^{(n)}$ may be used without an accent, if we regard the total refractive index of the system as that of a ray passing at one refraction into the last medium. This supposed, the equations (k) and (l) become

257.

Fundamental expressions for central foci of any system of spherical surfaces.

$$MF = M'F' + M''F'' + \dots M^{(n)}F^{(n)}; \quad (m)$$

$$Mf = MF + D; \quad M(F - f) + D = 0. \quad (n)$$

If the whole system be placed in vacuo, or if the last refraction be made into vacuum, we have $M = 1 = M^{(n)}$, and the equations become

258.

$$F = M'F' + M''F'' + \dots M^{(n)}F^{(n)} \} \\ f = F + D \} \quad (o)$$

Definitions. A lens in Optics is a portion of a refracting medium included between two surfaces of revolution whose axes coincide. If the surfaces do not meet, and therefore do not include space, an additional boundary is required, and this is a cylindrical surface, having its axis coincident with that of the surfaces.

259

Lenses defined and distinguished into species.

The axis of the lens is the common axis of all the bounding surfaces.

Lenses are distinguished (after the nature of their surfaces) into double-convex, with both surfaces convex, (fig. 44;) plano-convex, with one surface plane, the other convex, (fig. 45;) concavo-convex, (fig. 46;) double-concave, (fig. 47;) plano-concave, (fig. 48;) and meniscus, (fig. 49,) in which the concave surface is less curved than the convex. Also into spherical, (when the surfaces are segments of spheres;) conoidal, when portions of ellipsoids, hyperboloids, &c

Light.
260.
Species of
lenses how
distinguish-
ed alge-
braically.

These different species are distinguished, algebraically, by the equations of the surfaces, and by the signs of their radii of curvature. In the case of spherical lenses, to which our attention will be chiefly confined, if we suppose a positive value of the radius of curvature to correspond to a surface whose convexity is turned towards the *left*, or towards the incident rays, and a negative to that whose convexity is turned to the right, or from them, we shall have the following varieties of denomination :

meniscus	}	{	both radii +, as fig. 46, 49, <i>a</i> , or
concavo-convex			both radii —, as fig. 46, 49, <i>b</i> ,
plano-convex	{	{	radius of first surface +, of second infinite, fig. 45, <i>b</i> ,
			radius of first surface infinite, of second —, fig. 45, <i>a</i> ,
plano-concave	{	{	radius of first surface —, of second ∞ , fig. 48, <i>b</i> ,
			radius of first surface ∞ , of second +, fig. 48, <i>a</i> ,
double-convex			radius of first surface +, of second —, fig. 44,
double-concave			radius of first surface —, of second +, fig. 47,

the rays being supposed in all cases to pass from left to right.

A *compound lens* is a lens consisting of several lenses placed close together.

An *aplanatic lens* is one which refracts all the rays incident on it to one and the same focus.

Problem. To find the power and foci of a single thin lens in vacuo.

261.
Focus of a
single lens.

Let R' and R'' be the curvatures of its first and second surfaces respectively, μ the refractive index of the medium of which it consists, $m = \frac{1}{\mu}$; F its power: then we have, since the last refraction is made into vacuum,

$$F = \mu F' + F''; \quad f = F + D;$$

but, $F' = (1 - m') R'$, and $F'' = (1 - m'') R''$; and as $\mu' = \frac{1}{\mu}$ and $m'' = \mu$, these become respectively

$\frac{1}{\mu} (\mu - 1) R'$ and $-(\mu - 1) R''$, so that the foci of the lens are finally determined by the equations

$$\left. \begin{aligned} F &= (\mu - 1) (R' - R'') \\ f &= F + D \end{aligned} \right\} \quad (p)$$

Fundamen-
tal equations.

262. *Corol.* 1. The power of a lens is proportional to the difference of the curvatures of the surfaces in a meniscus or concavo-convex lens; and to their sum, in a double-convex or double-concave lens.

In plano-convex, or plano-concave lenses, the power is simply as the curvature of the convex or concave surface.

263. *Corol.* 2. In double-convex lenses R' is positive and R'' negative, so that when $\mu > 1$, F is positive, or the rays converge to a focus behind the lens. In plano-convex, $R'' = 0$ and R' is +; or $R' = 0$ and R'' is negative, (260); hence in both cases F is positive and the rays also converge. In meniscus lenses also, R' is +, and R'' , though +, is less than R' , (fig. 49;) therefore in these, also, the same holds good. In all these cases the *focus is said to be real*, because the rays actually meet there. In double-concave, plano-concave, or concavo-concave lenses, the reverse holds good; the focus lies on the opposite side, or towards the incident rays, and parallel rays, after refraction, diverge from it. In this case, therefore, they never meet, and the focus is called a *virtual focus*.

Fig. 49.
Real and
virtual foci.

264. *Corol.* 3. If μ be < 1 , or the lens be formed of a medium rarer than the ambient medium (which need not be vacuum, provided the whole system be immersed in it,) $\mu - 1$ is negative, and all the above cases are reversed. In this case convex lenses give virtual, and concave, real foci.

265. *Corol.* 4. For lenses of denser media, the powers of double-convex, plano-convex, and menisci are positive; and those of double plano-concave and concavo-convex lenses, negative; *vice versa* for rarer media.

Positive and
negative
powers.

266. *Corol.* 5. The focus of parallel rays is at the same distance, on whichever side of the lens the rays fall. For if the lens be turned above, R' becomes R'' , and *vice versa*; but, since they also change their signs, F remains unaltered.

267. *Corol.* 6. The equation $f = F + D$ gives $df = dD$. This shows that the foci of incident and refracted rays move always in the same direction, if the former be supposed to shift its place along the axis; and, moreover, that their proximities to the lens vary by equal increments or decrements for each.

Conjugate
foci move in
the same
direction.

Problem. To determine the central foci of any system of lenses placed close together, the lenses being supposed infinitely thin.

268.
Central foci
of a system
of thin lenses
in contact.

The general problem of a system of spherical surfaces contains this as a particular case; for we may regard the posterior surface of the first lens, and the anterior of the second, as forming a lens of vacuum interposed between the two lenses, and so for the rest. Thus the system of lenses is resolved into a system of spherical surfaces in contact throughout their whole extent; the alternate media having their refractive indices, or the alternate values of M , unity. If then we call μ', μ'', μ''' , &c. the refractive indices of the lenses, we shall have

$$M = 1; \quad M' = \mu'; \quad M'' = 1; \quad M''' = \mu''; \quad M'''' = 1, \text{ \&c.}$$

Light. The compound power F then will (258, α) be represented by

$$F = \mu' F' + F'' + \mu'' F''' + F^{iv} + \mu''' F^v + F^{vi} + \&c.$$

But

$$F' = (1 - m') R' = \frac{1}{\mu'} (\mu' - 1) R'$$

$$F'' = (1 - m'') R'' = (1 - \mu'') R'',$$

because $m' = \frac{1}{\mu'}$ and $m'' = \mu''$. Consequently,

$$\mu' F' + F'' = (\mu' - 1) (R' - R'')$$

and similarly

$$\mu'' F''' + F^{iv} = (\mu'' - 1) (R''' - R^{iv}), \&c.$$

so that we get, finally,

$$F = (\mu' - 1) (R' - R'') + (\mu'' - 1) (R''' - R^{iv}) + \&c.$$

Now, the several terms of which this consists are (by Art. 261) the respective powers of the individual lenses of which the system consists, so that if we put (according to the same principle of notation) L' , L'' , L''' , &c. for the powers of the single lenses, and L for their joint power as a system, we have

$$L = L' + L'' + L''' + \&c. \quad (q)$$

an equation which shows that the *power of any system of lenses is the sum of the powers of the individual lenses which compose it*; the word sum being taken in its algebraic sense, when any of the lenses has a negative power. Moreover it is easy to see that we also have $f = L + D$, as in the case of a single lens.

Reciprocally, we may regard a system of spherical surfaces forming the boundaries of contiguous media (as in the instance of a hollow lens of glass enclosing water) as consisting of distinct lenses, by imagining the concavity of one medium and the convexity of that in immediate contact with it separated by an infinitely thin film of vacuum, or of any medium having its surfaces equicurve, as in fig. 50; and thus a system of any number (n) of media, whose surfaces are in contact throughout their whole extent, may be conceived replaced by an equivalent system of $2n - 1$ lenses, the alternate ones being vacuum, or void of power. This way of considering the subject has often its use. It, moreover, leads to the result, that *the power of any system of spherical surfaces placed in vacuo is the sum of the powers of the several lenses into which it can be resolved, each placed in vacuo and acting alone*.

Let us now return to the case of surfaces separated by finite intervals; and, first, let us inquire the foci of a system of surfaces separated by intervals so small that their squares may be neglected. In this case the equations (j), Art. 251, become simply

$$D' = D; \quad D'' = f' + f'^2 t'; \quad D''' = f'' + f''^2 t'', \&c.;$$

and substituting these values in the equations (i), and retaining the notation of Art. 257, we find

$$Mf = M^{(n)} f^{(n)} = M' F' + M'' F'' + \dots M^{(n)} F^{(n)} + D \\ + M' f'^2 t' + M'' f''^2 t'' + \dots M^{(n-1)} f^{(n-1)2} t^{(n-1)} \}$$

Now in this we are to consider that

$$f' = F' + m' D, \quad f'' = F'' + m'' F' + m' m'' D, \&c.$$

and the values of f' , f'' , &c. so expressed, being substituted in the foregoing equation, we find

$$Mf = M' F' + M'' F'' + M''' F''' + \&c. \dots + D \quad (r) \\ + M' (F' + m' D)^2 t' + M'' (F'' + m'' F' + m' m'' D)^2 t'' + \&c.$$

Corol. In the case of two surfaces, supposing $M = 1$, or in the case of a single lens in vacuo, this gives

$$f = (\mu - 1) (R' - R'') + D + \frac{1}{\mu} \{ (\mu - 1) R' + D \}^2 t. \quad (s)$$

For parallel rays, this becomes

$$F = (\mu - 1) (R' - R'') + \frac{(\mu - 1)^2}{\mu} R'^2 t; \quad (t)$$

t being here put for t' , the interval between the surfaces or total thickness of the lens.

Problem. To determine the foci of a lens, whose thickness t is too considerable to allow of any of its powers being neglected.

Here we must take the strict formulæ

$$D' = D; \quad D'' = \frac{f}{1 - f' t}; \quad f' = (1 - m') R' + m' D; \quad \text{and } f'' = (1 - m'') R'' + m'' D''$$

The latter equation gives, on substitution, and recollecting that $m' = \frac{1}{\mu} = m$ and $m'' = \mu$,

Part I.

Superposition of powers. Power of a system of lenses is the sum of the powers of the component individuals.

269

Fig. 50.

Power of a system of spherical surfaces expressed.

270.

Foci of a system of surfaces separated by small finite intervals.

271.

Case of a single lens of small but finite thickness.

272.

Focus of a lens of any thickness.

Light

Part I.

$$f = f'' = \frac{(\mu - 1)(R' - R'') + D + \frac{\mu - 1}{\mu} \{(\mu - 1)R' + D\}R''t}{1 - \frac{1}{\mu} \{(\mu - 1)R' + D\}t}; \quad (u)$$

and for parallel rays

$$F = \frac{\mu(\mu - 1)(R' - R'') + (\mu - 1)^2 R' R'' t}{\mu - t \cdot (\mu - 1) R'} \quad (v)$$

273. *Example 1. To determine the foci of a sphere.*

Foci of a sphere.

Here $R'' = -R' = -R$; $t = \frac{2}{R}$; and the equations (u) and (v) become

$$f = \frac{(2\mu - 2)R + (2 - \mu)D}{(2 - \mu)R - 2D} R; \quad F = \frac{2\mu - 2}{2 - \mu} R. \quad (w)$$

274. *Corol. 1. If $\mu = 2$, for instance, these values become*

$$f = \frac{R^2}{D}; \quad F = \infty$$

In this case, then, since f and F express the proximities of the foci to the posterior surface of the sphere, we see that the focus for parallel rays falls on this surface, and that in any other case (as in fig. 51 and 52) q is given

by the proportion

$$QC : CE :: EH : Hq.$$

275. *Corol. 2. Whatever be the value of μ , the focus for parallel rays after the second refraction bisects the distance between the posterior surface of the sphere, and the focus after the first refraction.*276. *Example 2. To determine the foci of a hemisphere, in the two cases; first, when the convex, secondly, when the plane surface receives the incident light.*

Foci of a hemisphere.

In the first case, $R' = R$; $R'' = 0$; $t = \frac{1}{R}$: therefore we find

$$f = \frac{(\mu - 1)R + D}{R - D} R; \quad F = (\mu - 1)R.$$

277. In the other case, when the rays fall first on the plane side, $R' = 0$, $R'' = -R$, and $t = \frac{1}{R}$, so that

$$f = \frac{\mu(\mu - 1)R + D}{\mu R - D} R; \quad F = (\mu - 1)R.$$

278. If the thickness of a spherical segment exposed with its convex side to the incident rays be to the radius as

 μ to $\mu - 1$, or if $t = \frac{\mu}{\mu - 1} \cdot \frac{1}{R} = \frac{1}{(1 - m)R}$, and $R'' = 0$, the expressions (u) and (v) become

$$f = -(\mu - 1) \cdot \frac{R}{D} \{(\mu - 1)R + D\}; \quad F = \infty$$

In this case the focus for parallel rays falls on the posterior surface of the segment.

279. In general, for any spherical segment, if exposed with its convex side to the rays, $R'' = 0$, andFocus of any spherical segment, convex side first.
Plane side first

$$f = \mu \cdot \frac{(\mu - 1)R + D}{\mu + \{(\mu - 1)R + D\}t}; \quad F = \frac{\mu(\mu - 1)R}{\mu + (\mu - 1)Rt}$$

If the plane side be exposed to the rays

$$f = (\mu - 1)R + \frac{\mu D}{\mu - t \cdot D}; \quad F = (\mu - 1)R.$$

280. If $R' = R$ or if the lens be a spherical lamina of equal curvatures, the one convex, the other concave,

Focus of a spherical shell of equal curvatures.

$$f = \frac{\mu D + (\mu - 1)\{(\mu - 1)R + D\}Rt}{\mu - \{(\mu - 1)R + D\}t}; \quad F = \frac{(\mu - 1)^2 R^2 t}{\mu - (\mu - 1)Rt}.$$

§ X. Of the Aberration of a System of Spherical Surfaces.

Problem. To determine the focus of any annulus of a spherical refracting or reflecting surface.

The equations (a) of Art. 244, of the last section, in fact, contain a general solution of this problem; but the applications of practical Optics require an approximate solution for annuli of small diameter, or in which y is small compared with r . Conceiving y , then, so small that its fourth and higher powers may be neglected, the expressions in the article cited give

$$x = a - \sqrt{r^2 - y^2} = a - r + \frac{y^2}{2r}; \quad a - x = r - \frac{y^2}{2r}$$

$$yZ = \mu r (a - r) + \frac{a(\mu^2 r - a)}{2\mu r(a - r)} y^2;$$

and substituting these in the value of \overline{Cq} , found in the same article, we get for the distance of the focus of refracted rays from the vertex

$$\overline{Cq} = \frac{\mu r (r - a)}{a - \mu a + \mu r} - \frac{\mu - 1}{2\mu} \cdot \frac{a^2 (a + \mu r)}{(a - r)(a - \mu a + \mu r)^2} \cdot \frac{y^2}{r} \quad (a)$$

In conformity, however, with the system of notation adopted in the last section, instead of expressing directly \overline{Cq} , we will take its reciprocal. As we have hitherto represented the value of this reciprocal for central rays by f , we will continue to do so; and for rays incident at the distance y from the vertex, we will represent the same reciprocal by $f + \Delta f$; Δf then will be that part of f due to the deviation of the point of incidence from the vertex. Now, neglecting y^4 , we have

$$\frac{1}{\overline{Cq}} = \frac{a - \mu a + \mu r}{\mu r (r - a)} + \frac{\mu - 1}{2\mu^3} \cdot \frac{a^2 (a + \mu r)}{r^3 (a - r)^3} y^2 \quad (b)$$

Now if we put, as we have hitherto done, $\mu = \frac{1}{m}$, $r = \frac{1}{R}$, $a = \frac{1}{R} - \frac{1}{D}$, and substitute these in the above, we shall get the value of $\frac{1}{\overline{Cq}}$, or of $f + \Delta f$, in terms of m , R , and D ; and from this, subtracting the term independent of y^2 , which is the value of f , we shall get Δf as follows,

$$\Delta f = \frac{m(1 - m)}{2} (R - D)^2 \{ mR - (1 + m)D \} y^2. \quad (c)$$

Definition. The *longitudinal aberration*, is the distance between the focus for central rays and the focus q of the annulus, whose semidiameter, or *aperture*, is $y = MP$.

The *lateral aberration* at the focus, is the deviation from the axis of the refracted ray, or the portion fk , intercepted by the extreme ray, of a perpendicular to the axis drawn through the central focus.

Corol. These aberrations are readily found from the value of Δf above given; for since $\overline{Cq} = \frac{1}{f}$, we

have $\Delta \overline{Cq}$ (= longitudinal aberration) = $\Delta \frac{1}{f} = -\frac{\Delta f}{f^2}$; or, calling ω this aberration,

$$\omega = -\frac{\Delta f}{f^2}; \quad (d)$$

and since

$$\overline{Cq} : \overline{qk} :: y : \overline{fk}, \text{ or } \frac{1}{f} : \omega :: y : \overline{fk},$$

we have \overline{fk} , or the

$$\text{lateral aberration} = f \cdot y \cdot \omega = -\frac{\Delta f}{f} \cdot y; \quad (e)$$

where

$$f = (1 - m)R + mD.$$

Thus the whole theory of aberration is made to depend on the value of Δf , and we come therefore to consider the various cases of this which present themselves.

Case 1. For parallel rays $D = 0$; and, therefore,

$$\left. \begin{aligned} \Delta f &= \frac{m^2(1 - m)}{2} R^3 y^2; & \omega &= -\frac{m^2}{2(1 - m)} R y^2 \\ \text{lateral aberration} &= -\frac{m^2}{2} R^2 y^3 \end{aligned} \right\} \quad (f)$$

281.
Focus of a small annulus of a spherical surface investigated.

282

283.
Longitudinal and lateral aberration defined.
284.
Relation between them and Δf .

285
Case of parallel rays

Light.

Case 2. In reflectors, $m = \mu = -1$, and286.
Case of
reflectors.

$$\left. \begin{aligned} \Delta f &= R(R-D)^2 y^2; & w &= -\frac{R(R-D)^2}{(2R-D)^2} y^2; \\ \text{lateral aberration} &= -\frac{1}{2}(R-D)^2 y^3 \end{aligned} \right\}; \quad (g)$$

which, for parallel rays, become

$$\Delta f = R^3 y^2; \quad w = -\frac{1}{2} R y^2; \quad \text{lateral aberration} = -\frac{1}{2} R^2 y^2. \quad (h)$$

287.

In the general case, if we put either $D = R$, orAplanatic
foci defined
and inves-
tigated.

$$mR - (1+m)D = 0, \text{ which gives } D = \frac{m}{m+1} R; \quad \frac{1}{D} = (\mu+1) \cdot \frac{1}{R},$$

the value of Δf , and therefore of the aberration, vanishes. The former case is that of rays converging to the centre of curvature, in which, of course, they undergo no refraction. In the latter, the point is the same with that already determined, Art. 234. It is evident, from what was there demonstrated, that every spherical surface, CP, has two points Q, q in its axis, so related, that all rays converging to or diverging from one of them, shall after refraction *rigorously* converge to or diverge from the other. These points may be called the *aplanatic foci* of the surface; and, to distinguish them, Q may be called the aplanatic focus for incident, and q for refracted rays. To find them in any proposed case, in the axis of any proposed surface C, and on the

concave side of the surface, take CQ = $(\mu+1) \times$ radius CE of the surface, and Cq = $\left(\frac{1}{\mu} + 1\right) \times$ radius.

Then will Q and q be the aplanatic foci required. In the case of reflexion, when $\mu = -1$, CQ = Cq = 0, and both the aplanatic foci coincide with the vertex of the reflector.

288.

Aberration
shortens the
focus for
parallel rays

Let us next trace the effect of aberration in lengthening or shortening the focus, for all the varieties of position of the focus of incident rays; and, first, when $D = 0$, or for parallel rays, Δf is of the *same*, and therefore w of the *contrary* sign with R, and therefore with F, which is equal to $(1-m) \cdot R$. Hence it is evident, that the effect of aberration in this case must be to *shorten the focus of exterior rays*.

289.

Effect of
aberration
in other
cases.

Fig. 54.

290.

Q in this case is infinitely distant. As it approaches the surface, or as the rays from being parallel become more and more *convergent*, or *divergent*, the aberration diminishes; but the focus of exterior rays is still always *nearer* the surface than that of central, till Q comes up to the *aplanatic focus* Δ for incident rays on the *concave*, or to the focus F of parallel rays on the *convex side*. When Q is at the former of these points, the aberration is 0; at the latter, infinite.

When Q is situated anywhere between these points, however, the reverse is the case, and the effect of aberration is to throw the focus for exterior rays farther from the surface than that for central ones. These results are easily deduced from the consideration of all the particular cases, and hold good for all varieties of curvature, and for refracting media of all kinds. In reflectors, the aplanatic foci coincide with the vertex. In these, the focus for exterior rays is shorter than for interior in every case, except when the radiant point is situated between the surface and the principal focus on the *concave side* of the reflecting surface; but between these points, longer.

Problem. To determine the aberrations of any system of spherical refracting surfaces placed close together.

291.

Aoerration
of any
system of
spherical
surfaces in
contact.

Retaining the notation of Art. 257, let us suppose the ray, after passing through the first surface, to be incident on the second. Its aberration at this will arise from two distinct causes: first, that after traversing the first surface, instead of converging to or diverging from the focus for central rays, its direction was really to or from a point in the axis distant from that focus by the total aberration of the first surface; and, secondly, that being incident at a distance from the vertex of the second surface, a new aberration will be produced here, which (being, as well as the other, of small amount) the principles of the differential calculus allow us to regard as independent of it, and which being computed separately, and added to it, gives the whole aberration of the two surfaces regarded as a system. The same is true of the small alterations in the values of f' , f'' , &c. produced by the aberrations. If then we denote by $\delta f''$ the change in the value of f'' , produced by the action of the first surface, and by $\delta' f''$, that arising immediately from the action of the second, and by $\Delta f''$, the total alteration produced by both causes, we shall have

$$\Delta f'' = \delta f'' + \delta' f''$$

Now, first, to investigate the partial alteration $\delta f''$ arising from the total alteration $\Delta f'$ in the value of f' , or from the aberration of the first surface, we have

$$f'' = (1-m) \cdot R'' + m'' f', \text{ and therefore } \delta f'' = m'' \Delta f',$$

since, in this case,

$$D' = D, \quad D'' = f', \quad D''' = f'', \text{ \&c.}$$

Again, to discover the partial variation $\delta' f''$ in f'' , arising immediately from the action of the second surface, we have, by the equation (c) at once, putting f' for D'' , and neglecting y^4 , &c.

$$\delta' f'' = \frac{m''(1-m'')}{2} (R'' - f')^2 \{ m'' R'' - (1+m'') f' \} y^2;$$

but we have, by the same equation, also

$$\delta f'' = m'' \Delta f' = \frac{m'' m' (1-m')}{2} (R' - D)^2 \{ m' R' - (1+m') D \} y^2.$$

Light. Consequently, uniting the two, we have the value of $\Delta f''$. Similarly, the value of $\Delta f'''$ may be derived from that of $\Delta f''$, by a process exactly the same, and which gives Part I.

$$\Delta f''' = m''' \Delta f'' + \frac{m'''(1-m''')}{2} (R''' - f'')^2 \{ m''' R''' - (1 + m''') f'' \} y^2,$$

and so on. Calling, then, as in Art. 257, $M, M', M'', \dots M^{(n)}$ the absolute refracting indices of the several media into which the successive refractions are made, and putting $M^{(n)} = M$, we shall have no difficulty in arriving at the following general expression, where Δf denotes the total effect of aberration on the value of f , the reciprocal focal distance of the system,

General expression for Δf .

$$M \cdot \Delta f = \left\{ \begin{array}{l} M' \cdot \frac{m'(1-m')}{2} (R' - D)^2 \{ m' R' - (1 + m') D \} \\ + M'' \cdot \frac{m''(1-m'')}{2} (R'' - f')^2 \{ m'' R'' - (1 + m'') f' \} \\ + M''' \cdot \frac{m'''(1-m''')}{2} (R''' - f'')^2 \{ m''' R''' - (1 + m''') f'' \} \\ + \&c. \end{array} \right\} \cdot y^2; \quad (i)$$

in which it will be recollected that

$$\left. \begin{array}{l} f' = (1 - m') R' + m' D \\ f'' = (1 - m'') R'' + m'' (1 - m') R' + m' m'' D \\ f''' = (1 - m''') R''' + m''' (1 - m'') R'' + m''' m'' (1 - m') R' + m''' m'' m' D \\ \&c. \end{array} \right\}; \quad (j)$$

Successive values of f .

and these values being substituted give, if required, an explicit resulting value of Δf in terms of the radii and refractive indices, or their reciprocals, of the surfaces.

If the system of surfaces be placed in vacuo, or the last refraction be made into vacuum, $M = 1$, and the second member of the equation (i) exhibits simply the value of Δf . In all cases, the aberration ω is given as before by the equation

292.

$$\omega = - \frac{\Delta f}{f^2}, \text{ and the lateral aberration is } - \frac{\Delta f}{f} y.$$

To express the aberration of any infinitely thin lens in vacuo, let the terms of the general equation be denoted respectively by $Q', Q'', \&c.$, so as to make

293.

Aberration of a single infinitely thin lens.

$$M \cdot \Delta f = \{ Q' + Q'' + Q''' +, \&c. \} y^2. \quad (k)$$

Then, for the case of a single lens in vacuo, when $m'' = \frac{1}{m'}$, $M' = \frac{1}{m'}$, $M'' = 1$, $M = 1$, we have

$\Delta f = Q' + Q''$; and putting, for a moment, $R' - D = B$, $R' - R'' = C$, we find

$$Q' = \frac{1-m'}{2} y^2 B^2 \{ m' B - D \}$$

$$Q'' = - \frac{1-m'}{2 m'^3} y^2 (m' B - C)^2 \{ m'^2 B - m' D - C \}$$

whence
$$Q' + Q'' = \frac{1-m'}{2 m'^3} y^2 C \{ (2 m' B - C) (m'^2 B - m' D) + (C - m' B)^2 \}$$

The expression in brackets, putting for B and C their values, and $\frac{1}{\mu}$ for m' , will become

$$\frac{1}{\mu^3} \{ (2 - \mu) R' + \mu R'' - 2 D \} (R' - (1 + \mu) D) + \mu \{ (\mu - 1) R' - \mu R'' + D \}^2 \}.$$

If now we multiply out, arranging according to powers of D , and substitute the result, as also the value of m' ,

$(= \frac{1}{\mu})$, and of C , $(= R' - R'')$, in $Q' + Q''$, or Δf ; we get

$$\Delta f = (\mu - 1) (R' - R'') \cdot \frac{y^2}{2 \mu} \{ \alpha - \beta D + \gamma D^2 \}$$

General expression for it.

where

$$\left. \begin{array}{l} \alpha = (2 - 2 \mu^2 + \mu^3) R'^2 + (\mu + 2 \mu^2 - 2 \mu^3) R' R'' + \mu^3 R''^2 \\ \beta = (4 + 3 \mu - 3 \mu^2) R' + (\mu + 3 \mu^2) R'' \\ \gamma = 2 + 3 \mu \end{array} \right\}. \quad (l)$$

Light. Now it has been shown, (Art. 261,) that $(\mu - 1) (R' - R'')$ expresses the power of the lens, so that, putting L Part I. for this, we have

$$\Delta f = \frac{L}{2\mu} (a - \beta D + \gamma D^2) y^2. \quad (m)$$

Such then is the general expression for Δf , the fundamental quantity, from which the aberration w may be had in any lens by the equation $w = -\frac{\Delta f}{f^2}$.

Corol. 1. The aberration of a lens vanishes when D is so related to R' , R'' and μ , as to give

$$a - \beta D + \gamma D^2 = 0; \quad D = \frac{\beta \pm \sqrt{\beta^2 - 4a\gamma}}{2\gamma}. \quad (n)$$

Now we find, by substitution and reduction,

$$\beta^2 - 4a\gamma = \mu^2 \{ (R' + R'')^2 - (2\mu + 3\mu^2) (R' - R'')^2 \}$$

and unless this quantity be positive, that is, unless

$$\left(\frac{R' + R''}{R' - R''} \right)^2 > 2\mu + 3\mu^2; \quad (o)$$

the focus of incident rays cannot be so situated as to render the aberration nothing. But, if the curvatures R' and R'' of the surfaces be such as to satisfy this condition, the value of D may be calculated at once from the equation (k.)

295. Corol. 2. Whenever, in meniscus or concavo-convex lenses, the difference of the curvatures of the surfaces is small in comparison with their sum, that is, whenever a moderate focal length is produced by great curvatures, the aberration admits of being rendered evanescent by properly placing the focus of incident rays. In a lens of crown glass where $\mu = 1.52$, we have $\sqrt{2\mu + 3\mu^2} = 3.16$; therefore the sum of the curvatures must be at least 3.16 times their difference, to satisfy the condition of possibility. In double-convex or double-concave lenses, R' and R'' having opposite signs, the condition can never be satisfied.

Corol. 3. If $a = 0$, the aberration vanishes for parallel rays. This condition is, however, only to be satisfied by real values of R' and R'' when μ is equal to or less than $\frac{1}{2}$, and no such media are known to exist.

Corol. 4. The effect of aberration will be to shorten or lengthen the focus for exterior rays, according as the sign of Δf is the same as, or the opposite to, that of f . In particular cases it will, of course, however, depend on the values of μ , R , R' , and D which shall take place. The principal case is that of parallel rays, in which $D = 0$, and

$$\Delta f = \frac{y^2}{2\mu} \cdot L \{ (2 - 2\mu^2 + \mu^3) R^{1/2} + (\mu + 2\mu^2 - 2\mu^3) R' R'' + \mu^3 R'^{1/2} \}$$

and the focus of external rays will be shorter or longer than that of central ones, according as this quantity has the same, or opposite sign with L , that is, according as

$$(2 - 2\mu^2 + \mu^3) R^{1/2} + (\mu + 2\mu^2 - 2\mu^3) R' R'' + \mu^3 R'^{1/2}$$

is positive or negative. Now, from what we have already seen in the last corollary, this quantity never can be rendered negative by any real values of R' and R'' , unless μ be less than $\frac{1}{2}$. For all other media, therefore, (comprehending all yet known to exist in nature,) every lens, whatever be the curvatures of its surfaces, has the exterior focal length for parallel rays shorter than the central.

298. Corol. 5. In a glass meniscus, when the radiant point is on the convex side, and the rays diverge, we have $4 + 3\mu - 3\mu^2$ a positive quantity; and, R' and R'' being both positive, β is so; hence (D being negative in this case) the term $-\beta D$, and therefore the whole factor $a - \beta D + \gamma D^2$ is positive; and L being also positive, Δf is so; and, therefore, w , the aberration, negative. Hence, when Q is beyond F , the focus for parallel rays incident the other way, the exterior focus is the shorter; but when between F and C , the longer.

Corol. 6. Unless $\left(\frac{R' + R''}{R' - R''} \right)^2 > 2\mu + 3\mu^2$, no real value of D can render $a - \beta D + \gamma D^2$ negative.

It appears, therefore, that in all double-convex or concave lenses, as well as in all meniscus and concavo-convex ones, in which the sum of the curvatures of the surfaces is greater than $\sqrt{2\mu + 3\mu^2}$ times their difference, the factor $a - \beta D + \gamma D^2$ is positive for all values of D , and therefore the aberration w has in all such lenses the sign opposite to that of L . Hence, for all such lenses, we have the following simple and general rule: the effect of aberration will be to throw the focus of exterior rays more towards the incident light than that of central ones, when the lens is of a positive character, or makes parallel rays CONVERGE, but more FROM the incident light if of a negative, or if it cause parallel rays to DIVERGE.

Corol. 7. All other lenses have, as in the case of single surfaces, aplanatic foci, corresponding to the roots of the equation $a - \beta D + \gamma D^2 = 0$. In general there are two such foci of incident and two of refracted rays: and

rules might easily be laid down for determining in what positions of the radiant point, with respect to these foci and the lens, the aberration tends to shorten or lengthen the exterior focus; but it is simpler and readier to have recourse at once to the algebraic expressions.

Corol. 8. In the case of reflexion, as when rays are reflected between the surfaces of thin lenses of transparent media, we have $m = m'' = \&c. = \mu' = \mu'' = \&c. = -1$; $M' = -1$, $M'' = +1$, &c., and $M = \pm 1$, according as the number of reflexions is even or odd; therefore for n reflexions we have

$$\left. \begin{aligned} f' &= 2 R' - D \\ f'' &= 2 R'' - 2 R' + D \\ f''' &= 2 R''' - 2 R'' + 2 R' - D \\ &\&c. \end{aligned} \right\}; \quad (p)$$

and

$$\Delta f = (-1)^{n+1} \left\{ \begin{aligned} &R' (R' - D)^2 \\ &- R'' (R'' - 2 R' + D)^2 \\ &+ R''' (R''' - 2 R'' + 2 R' - D)^2 \\ &- \&c. \end{aligned} \right\} y^2; \quad (q)$$

which formulæ serve to determine, in all cases of internal reflexion between spherical surfaces, both the places of the successive foci and the aberrations.

Corol. 9. If the reflexions take place between equicurve surfaces, having their concavities turned opposite ways, f' , f'' , &c. are in arithmetical, and therefore their reciprocals, or the focal distances, in harmonic progression.

Problem. To construct an aplanatic lens, or one which shall refract all rays, for a given refractive index, and converging to or diverging from any one given point, to or from any other.

Let Q and q be the points, the former being the focus of incident, the latter of refracted rays. Let μ = index of refraction; and putting $Qq = 2c$, and assuming b any arbitrary quantity, construct the curve whose equation is (u) , Art. 232. Let HPC , (fig. 36,) be this curve; and with centre q , and any radius qN less than qP , any one of the refracted rays describe the circle HNK . Then since the ray QP , by the nature of the curve HPC , is after refraction directed to or from q , and, being incident perpendicularly on the second surface, suffers there no flexure, it will, if supposed to emerge from the medium, here continue its course to or from q . If then we suppose the figure $CPHNK$ to revolve round Qq , it will generate a solid, which, being composed of the proposed medium, is the lens required. If the rays be parallel, as in fig. 38, the curve HPC , as we have seen, is a conic section, which, if the lens be denser than the ambient medium, is an ellipse. Thus, a glass meniscus lens, whose anterior convex surface is elliptic, and posterior spherical, having its centre in the focus of rays refracted by the first surface, is *aplanatic*.

But, without having recourse to the conic sections, the same thing may, in certain cases, be accomplished by spherical surfaces only. For if Q and q (fig. 53) be the *aplanatic foci* of the spherical refracting surface, and if with the centre q and any radius greater than qC , when the incident rays diverge from Q , as in the lower portion of the figure, but less if they converge to Q as in the upper, we describe a circle KL , or kl , and turn the whole figure about Qq as an axis, the surfaces $CPKL$, or $cpkl$, will generate the aplanatic lens in question. This also follows evidently from the general formula, (*i*, Art. 291,) for if $R'' = f'$, the expression of Δf for the lens becomes simply

$$\frac{1 - m'}{2} (R' - D)^2 \{ m' R' - (1 + m') D \} y^2,$$

which vanishes when $D = \frac{m'}{1 + m'} R'$, or when Q is the aplanatic focus of incident rays for the first surface.

More generally, however, the equation $\alpha - \beta D + \gamma D^2 = 0$, assigns the universal relation between μ , D , R' , R'' , which constitutes the lens aplanatic. See *Cor. 1*, Art. 294.

Problem. To assign the most advantageous form for a single lens, or that which, with a given power, has the least possible aberration for parallel rays.

Since the aberration cannot be rigorously made to vanish for parallel rays, when $\mu > \frac{1}{2}$ (Art. 296) we have to

make it a minimum. Now $w = -\frac{\Delta f}{f^2} = -\frac{\Delta f}{L^2}$ for parallel rays, or

$$w = -\frac{y^2}{2\mu} \cdot \frac{a}{L}; \text{ and, in general, } dw = -\frac{y^2}{2\mu^2} \{ L da - a dL \}$$

In the present case L is given, therefore we must put $da = 0$, which gives

$$0 = 2(2 - 2\mu^2 + \mu^3) R' dR' + (\mu + 2\mu^2 - 2\mu^3) (R' dR'' + R'' dR') + 2\mu^3 R'' dR''.$$

But the condition $dL = 0$ gives $dR' = dR''$; so that our equation becomes, on substitution and reduction,

$$0 = (4 + \mu - 2\mu^2) R' + (\mu + 2\mu^2) R'';$$

Part I.

How to proceed in other cases. 301.

Case of reflexion between any system of transparent surfaces.

302.

303.

General construction of an aplanatic lens. Fig. 36.

Fig. 38.

304.

Case when the surfaces of an aplanatic lens are all spherical. Fig. 53.

305.

Most advantageous form for a single lens for parallel rays determined.

Light. that is to say

Part I.

$$\frac{R''}{R'} = \frac{2\mu^2 - \mu - 4}{2\mu^2 + \mu} \quad (r)$$

In the case of a glass lens, taking $\mu = 1.5$, this fraction becomes equal to $-\frac{1}{6}$, which shows that the lens must be double-convex, having the curvature of the posterior surface only $\frac{1}{6}$ that of the anterior, or its radius six times as great. Artists sometimes call such a lens a "*crossed lens*."

306. *Corol. 1.* If $\mu = 1.6861$, as is nearly the case with several of the precious stones and the more refractive glasses, $R'' = 0$; and the most advantageous figure for collecting all the light in one place is plano-convex, having its convex side turned to the incident rays.

307. *Corol. 2.* Calling the aberration of a lens of the best figure w , we shall have $w = -\frac{15}{14} y^2 \cdot L$, for glass whose refractive index is 1.5, and the proportional aberrations of other forms will be as follows:

Aberrations of various species of lenses determined for parallel rays

Plano-convex, plane side first (or towards the light)	$4.2 \times w$
Plano-convex, curved surface first	$1.081 \times w$
Double equi-convex, or concave	$1.567 \times w$

308. *Problem, To investigate a general expression for the aberration of any system of infinitely thin lenses placed close together in vacuo.*

Aberration of a system of lenses.

The general expression for $M \Delta f$, or, since $M = 1$ in the case before us, of Δf , is

$$(Q' + Q'' + Q''' + Q^{iv} + \&c.) y^2,$$

which divides itself into terms originating with the successive lenses in the following manner,

$$\Delta f = (Q' + Q'') y^2 + (Q''' + Q^{iv}) y^2 + \&c.$$

The first of these quantities we have already considered; let us now, therefore, examine the constitution of the rest. Let then μ' be the refractive index of the first lens, μ'' of the second, μ''' of the third; and let α', β', γ' represent the values of α, β, γ for the first lens, or the expressions in (l, 292,) writing only μ' for μ ; also let $\alpha'', \beta'', \gamma''$ represent their values for the second lens, or what the same expressions become when μ'' is put for μ , and R''' and R^{iv} respectively for R' and R'' , and so on for the rest of the lenses.

309. Now if we consider the values of Q''' and Q^{iv} , it will be seen that they are composed of the quantities $m''', m^{iv}, M''', M^{iv}, R''', R^{iv}, f''$ and f''' , precisely in the same manner that Q' and Q'' are of $m', m'', M', M'', R', R'', D$ and f' .

Moreover, since by Art. 251 we have

$$\begin{aligned} f' &= (1 - m') R' + m' D \\ f'' &= (1 - m'') R'' + m'' f' \\ &= (1 - m'') R'' + m'' (1 - m') R' + m'' m' D \\ &= (\mu - 1) (R' - R'') + D, \text{ since } m' = \frac{1}{\mu}, m'' = \mu. \\ &= L + D; \text{ call this } D''; (L \text{ is the power of the first lens}) \\ f''' &= (1 - m''') R''' + m''' D'' \\ f^{iv} &= (1 - m^{iv}) R^{iv} + m^{iv} f''' = L'' + D'' \text{ as before; } (L'' \text{ is the power of the second lens}) \\ &= L + L' + D; \text{ and so on.} \end{aligned}$$

And it is clear that $Q''' + Q^{iv}$ will be the same *function* of, *i. e.* similarly composed of, the refractive index and curvatures of the surfaces of the second lens, and of the quantities D'' and f''' , that $Q' + Q''$ is of the refractive index and curvatures of the first lens, and of D and f' . It follows, therefore, that the very same system of reductions which led to the equation

$$Q' + Q'' = \frac{L}{2\mu} (\alpha - \beta D + \gamma D^2)$$

being pursued in the case of $Q' + Q^{iv}$, must lead to the precisely similar equation

$$Q''' + Q^{iv} = \frac{L''}{2\mu''} (\alpha'' - \beta'' D'' + \gamma'' D''^2)$$

General expression for it.

and so on for the remaining lenses; so that we shall have, ultimately, for the whole system (writing L', D', μ for L, D, μ)

$$\Delta f = \frac{y^2}{2} \left\{ \frac{L'}{\mu'} (\alpha' - \beta' D' + \gamma' D'^2) + \frac{L''}{\mu''} (\alpha'' - \beta'' D'' + \gamma'' D''^2) + \&c. \right\}; \quad (s)$$

in which there are as many terms as lenses.

Light.

Corol. For parallel rays, therefore

$$D' = 0; \quad D'' = L'; \quad D''' = L' + L'', \text{ \&c.}$$

Part I.

$$\Delta f = \frac{y^2}{2} \left\{ \begin{aligned} & -\frac{L'}{\mu'} - a' + \frac{L''}{\mu''} (a'' - \beta'' L' + \gamma'' L'^2) \\ & + \frac{L'''}{\mu'''} (a''' - \beta''' (L' + L'') + \gamma''' (L' + L'')^2) \\ & + \text{\&c.} \end{aligned} \right\}. \quad (t)$$

310.
Case of
parallel
rays.

Although the aberration of a single lens for parallel rays admits of being destroyed only on a certain hypothesis of the refractive index, which has no place in nature, yet, by combining two or more lenses, it may be destroyed in a variety of ways. Thus, in the case of two lenses, the expression (t) being put equal to zero, gives an equation involving $\mu', \mu'', L', L'', R', R'', R''', R^{iv}$; or (since L' and L'' are given in terms of μ', μ'' and $R', R'', \text{\&c.}$ and since μ', μ'' are given quantities) only the four unknown quantities R', R'', R''', R^{iv} . Now as there are four of these, and only one equation, it may be satisfied in an infinite variety of ways, and the problem of the destruction of the *spherical aberration* (as it is termed) becomes indeterminate.

311.

The equation in the case of two lenses for parallel rays is

$$0 = -\frac{L'}{\mu'} \left\{ (2 - 2\mu'^2 + \mu'^3) R'^2 + (\mu' + 2\mu'^2 - 2\mu'^3) R' R'' + \mu'^3 R''^2 \right\}; \quad (u)$$

$$+ \frac{L''}{\mu''} \left\{ (2 - 2\mu''^2 + \mu''^3) R''^2 + (\mu'' + 2\mu''^2 - 2\mu''^3) R'' R''' + \mu''^3 R'''^2 \right\}$$

$$- \frac{L' L''}{\mu' \mu''} \left\{ (4 + 3\mu'' - 3\mu''^2) R''' + (\mu'' + 3\mu''^2) R^{iv} \right\} + \frac{L'^2 L''}{\mu'^2 \mu''} \left\{ 2 + 3\mu'' \right\}$$

312.
General
equation for
the destruc-
tion of aber-
ration in a
double lens
for parallel
rays.

This equation, if L' and L'' , the powers of the separate lenses, be assigned, is of a quadratic form in either R', R'', R''' , or R^{iv} ; it will therefore depend on the supposition adopted to limit the problem, whether these quantities admit real corresponding values. Now the equations $L' = (\mu' - 1) (R' - R'')$ and $L'' = (\mu'' - 1) (R'' - R''')$ afford the means of eliminating two of them, and the resulting equation (in R' and R''' for instance) is

313.

Another
form of the
same equa-
tion.

$$0 = L' \left\{ \frac{2 + \mu'}{\mu'} R'^2 - \frac{2\mu' + 1}{\mu' - 1} L' R' \right\}$$

$$+ L'' \left\{ \frac{2 + \mu''}{\mu''} R''^2 - \left(\frac{4(\mu'' + 1)}{\mu''} L' + \frac{2\mu'' + 1}{\mu'' - 1} L'' \right) R'' \right\}$$

$$+ \frac{\mu'^2 L'^3}{(\mu' - 1)^2} + \frac{\mu''^2 L''^3}{(\mu'' - 1)^2} + \frac{3\mu' + 1}{\mu' - 1} L' L''^2 + \frac{2 + 3\mu'}{\mu''} L' L'';$$

and, as the unknown quantities R', R'' are not combined by multiplication, the equation when L' and L'' are given is of an ordinary quadratic form with respect to each. This equation will be of use to us hereafter, when we come to treat of the theory of refracting telescopes.

If L' and L'' be not given, since either of them is of the first degree in terms of $R', R'', \text{\&c.}$, the equation (u) is of the third degree in either of the quantities $R', R'', \text{\&c.}$, or in L', L'' , if either R'' or R^{iv} be eliminated. Now as an equation of the third degree must necessarily have at least one real root, we conclude, first, that in a double lens, if the curvatures of three of the surfaces be given, that of the fourth may be found, so as to destroy the spherical aberration.

314.

Secondly. That if the curvature of one surface of each lens, and the power of either, or that of the two combined, be given, the power of the other may be found so as to destroy the spherical aberration. This is evident; for, supposing R' and R'' given, and either L' or L'' , or $L' + L''$, also given, the equation (v) becomes an ordinary cubic in which L' or L'' , as the case may be, is the only unknown quantity, and therefore necessarily admits a real value.

315.

As examples of aplanatic combinations, we may set down the following cases, in which a lens of glass of the refraction 1.50, and of the best form, having the radii of its surfaces respectively + 5.833 and - 35.000 inches, and its focal length 10.000 inches, has its aberration corrected by applying behind it another lens of similar glass, as in fig. 55. This lens is a meniscus. If its curvatures be determined by the condition of giving the maximum of power to the combination, the radii of its surfaces and its focal length will be as follows: radius of first surface, = + 2.054 inches; radius of second surface, = + 8.128; focal length of correcting lens, = + 5.497; focal length of the two combined, = + 3.474. On the other hand, if we determine the second lens by the condition of the resulting combination, having a focal length as nearly 10.000 as is consistent with perfect aplanaticity, we shall find radius of first surface, = + 3.688; radius of second, = + 6.291; focal length of correcting lens, = + 17.829; focal length of the combination, = + 6.407.

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Fig. 55.

The effect of aberration may be very prettily exhibited by covering a large convex lens with a paper

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Light.

screen full of small round holes, regularly disposed, and, exposing it to the sun, receiving the converged rays on a white paper behind the lens, which should be first placed very near it, and then gradually withdrawn. The pencils which pass through the holes will form spots on the screen, and their disposition will become more and more unequal over the surface, as the screen is further removed; those at the circumference becoming crowded together before the central ones. The manner in which the several spots corresponding to central rays blend together into one image at the focus, and those formed by the exterior ones are scattered round it, gives us a very good idea of the variation of density of the rays in the circle of aberration at or near the principal focus; and if the white screen be waved rapidly to and fro in the cone of rays, so as to pass over the focus at each oscillation, the whole cone will be seen as a solid figure in the air; and the place of the circle of least aberration will become evident to the eye, forming altogether a very pleasing and instructive experiment.

§ XI. Of the Foci for Oblique Rays, and of the Formation of Images.

318.

We have hitherto considered rays as converging to, or diverging from, a single point; but as this is not the case with luminous bodies of a sensible diameter, we now proceed to examine the cases of refraction at spherical surfaces, where more than one radiant point is concerned, or where several pencils are incident at once on the surface. We shall take for our positive, or fundamental case, as we have done all along, that of *converging* rays incident on the convex side of a more refractive medium than the ambient one, and derive all others from it by the changes in the sign and relative magnitudes of R , D , &c.

Foci of oblique pencils.

In fig. 56, then, let Q and Q' be the foci of two pencils of convergent rays incident on the spherical surface CC' , whose centre is E . Draw QEC , $Q'E C'$, cutting the surface in C and C' , and, regarding CEQ as the axis of the pencil RQ , SEQ , TEQ , the focus of refracted rays will be found by taking q , such as that $\frac{1}{Cq}$, or f , shall be equal to $(1 - m)R + mD$, (247, e.) Similarly, regarding $C'E Q'$ as the axis of the pencil converging to Q' , the focus q' will be had by the equation

$$\frac{1}{C'q'} = f' = (1 - m)R + mD'.$$

Thus when $C'Q' = CQ$, Cq' will also equal Cq , and, in general, when the locus of the point Q is given, that of q may be found.

319. Images in Optics defined.

Definition. The image of an object, in Optics, is the locus of the focus of a pencil of rays diverging from, or converging to, every point of it, and received on a refracting surface. Thus, supposing CQ' to be a line, or surface, every point of which may be regarded as a focus of incident rays, $q q'$ is its image.

320. Form of the image of a straight line

Problem. To find the form of the image of a straight line formed by a spherical refracting or reflecting surface.

Put $CE = r$; $CQ = a$; $EM = x$; $MQ' = y$; $EQ' = \sqrt{x^2 + y^2}$; $C'Q' = a'$.

Then we have

$$\frac{1}{C'q'} = \frac{1 - m}{r} + \frac{m}{a'} = \frac{(1 - m)a' + mr}{ra'};$$

and therefore

$$C'q' = \frac{ra'}{(1 - m)a' + mr}; \quad EQ' = \frac{mr(a' - r)}{(1 - m)a' + mr};$$

we have, consequently,

$$x^2 + y^2 = \frac{m^2 r^2 \cdot (a' - r)^2}{\{(1 - m)a' + mr\}^2}.$$

But, by similar triangles, $EQ' : EM :: EQ' : EQ$, or

$$x^2 + y^2 = \frac{(a' - r)^2 \cdot x^2}{a^2},$$

equating these two values we get

$$\frac{a}{x} = \frac{(1 - m)a' + mr}{mr}; \quad a' = \frac{m}{1 - m} \cdot \frac{r(a - x)}{x};$$

Is a conic section.

so, that eliminating a' , by substituting this value for it, we get for a final equation between x and y , or for the equation of the image

$$(1 - m)^2 (x^2 + y^2) = \left(\frac{r}{a}\right)^2 \cdot (ma - x)^2$$

which belongs to a conic section.

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Problem. When an oblique pencil is incident on any system of spherical surfaces, to find the focus of refracted rays.

Light.

Take E' , (fig. 57,) the centre of the first surface, and let Q' be the focus of incident rays. Join $Q'E'$ and produce it to C' , then will C' be the vertex of the surface corresponding to the pencil whose focus is Q' ; and taking

$$\frac{1}{C'Q'} = \frac{1 - m'}{C'E'} + \frac{m'}{C'Q'}$$

Q'' will be the focus of refracted rays. Again, join Q'' and E'' , the centre of the second surface, produce to C'' , and take

$$\frac{1}{C''Q''} = \frac{1 - m''}{C''E''} + \frac{m''}{C''Q''}$$

and Q''' will be the focus after refraction at the second surface, and so on.

Corol. In the case of an infinitely thin lens, when the obliquity is small, it is evident, from this construction, that the focus of oblique rays will lie at the same distance from the lens with that of rays convergent to, or divergent from, a point in the axis at the same distance with the focus of incident rays, but instead of lying in the axis, will deviate from it.

Definition. The centre of a lens is a point in its axis where a line joining the extremities of two parallel radii of its two surfaces cuts the axis. Thus, in the various lenses represented in fig. 58, 59, 60, and 61, $E'A$ and $E''B$ being two parallel radii; join BA , and produce, if necessary, till it meets the axis in X , and X is the centre.

Corol. 1. The centre is a fixed point; for, since AE' and BE'' are parallel, we have $E'X : E'E'' :: AE' : BE'' - AE'$, in which proportion three terms being invariable, the other is so also.

Corol. 2. If $C'C''$, the interval of the surfaces or thickness of the lens, be put equal to t (t being always positive) and the curvatures be respectively R' and R'' , we have, for the distance of the centre from the first surface or for $C'X$, the following value.

$$C'X = \frac{R''}{R' - R''} \cdot t.$$

Corol. 3. If a ray be so incident on a lens that its direction after the first refraction shall pass through its centre, it will suffer no deviation. This is evident, because its course within the lens will be AB , and the radii $E'A$ and $E''B$ being parallel, the internal angles of incidence on the surfaces are equal, and, therefore, the angles of refraction both ways out of the lens; consequently the two portions of the ray without the lens are parallel.

Corol. 4. If the thickness of a lens be very small, the ray passing through its centre may be regarded as undergoing no refraction whatever; for the portion AB within the lens being very small, the two portions exterior to the lens (being parallel) may be regarded as one ray. This is, *à fortiori*, still nearer the truth when the obliquity of the ray to the axis is small; because then the portion AB is very nearly coincident in direction with either of the two exterior portions.

Corol. 5. Hence, to find the focus of refracted rays in the case of a *very thin lens* and for a *pencil of small obliquity*, take X , the centre of the lens, and the focus will lie in the line QX , at the same distance from the lens as if the axis of the incident pencil were coincident with that of the lens.

Proposition. When a luminary, or illuminated object, is placed before a double or plano-convex, or meniscus lens, at a distance from it greater than its focal length, there will be formed behind the lens an image, similar to the object, but inverted; and the object and image subtend the same angle at the centre of the lens.

For the pencil of rays which emanates (either by direct radiation or by reflexion) from any point, as P , of the object, will after refraction be all made to converge to a point p behind the lens, or at least very nearly so. Were the aberration of the lens evanescent, the convergence would be mathematically exact; and since, whenever the aperture of the lens and the obliquity of the pencil are small, the aberration is so very minute, that the space over which the rays are spread may be regarded as a physical point, and every physical point in the object will have a corresponding point in the image. Now, C being the centre of the lens, the line joining Pp passes through C ; and the same being true of the line joining any other corresponding points of the object and image, it follows, by similar triangles, that the object and image are similar in figure; and as the rays cross at C , the image is inverted, and subtends the same angle pCq at C that the object does on the other side.

If a screen of white paper be placed at qp , this image will be rendered visible as a picture of the object. The experiment may be tried with any magnifier or spectacle-glass at a window, when the forms of external objects, the houses, trees, landscape, &c. will be painted on the paper screen with perfect fidelity, forming a miniature of the utmost delicacy and beauty. This is the principle of the common camera obscura, in which the rays from external objects are thrown by an inclined looking-glass downwards, and being received on a convex lens, are brought to their focus on a white horizontal table, in a room where no other light is admitted. On this table a moving picture of all external objects, in their proper forms, colours, and motions, is seen, infinitely more correct and beautiful than the most elaborate painting. See fig. 63, in which P is the object, AB the reflector, BC the lens, and p the image on the table D .

If the rays, instead of being received on white paper, be received on a plate of glass emiered on one side, the picture may be seen by an eye placed at the other side of the glass, as well as by one in front of it; for it is a property of such roughened transparent surfaces to scatter the rays which fall on them, not only by reflexion outwards, but by refraction inwards. If the surface be but slightly roughened, however, the picture will appear much less vivid when looked at obliquely than when the eye is placed immediately behind it; and in this

Part I.
Foci of oblique pencils incident on a system of spherical surfaces.
Fig. 57.

322.

323.
Centre of a lens.

324.

325.

326.

Rays through the centre pass undeviated.

327.

328.

Focus of a slightly oblique pencil through a thin lens.

329.

Fig. 62.
An inverted image of an object is formed behind a convex lens.

330.

Camera obscura explained.

331.

Light. latter situation the emerald glass may even be removed altogether, and the image will still be seen, and even more distinctly, as if a real object stood in the place in all respects similar to the picture.

332. We may examine the image on the roughened glass with a magnifying glass, or microscope. It will then appear as a delicate painting, accommodating itself to all the inequalities of the surface. But if, in the act of so examining it, the rough glass be removed, the painting remains as if suspended in air, and the objects it represents are seen brought nearer to the eye, and enlarged in their dimensions. In short, we have formed a telescope.

333. If the lens used to form the image be a concave one, or if a convex reflector be used, as in fig. 64 and 65, the rays, after refraction or reflexion, diverge, not from any actual points in which they cross, but from points in which they would cross if produced backwards. There is in this case, then, no real image formed capable of being received on a screen, but what is called a *virtual one*, visible to the eye if properly situated, either unassisted or aided by a magnifier, and situated on the same side of the lens, or on the contrary side of the reflector with the object, and therefore *erect*.

334. The perfection of the image formed by a lens or reflector, its exact resemblance to the object, and the distinctness of its parts, will depend on the exact convergence of all the rays of pencils emanating from every physical point of the object in strict mathematical points, or in as near an approach to such points as may be. If, therefore, a lens of considerable diameter be used, especially if the curvatures of its surfaces be improperly chosen so as to produce much aberration, the image will be confused; for each point of the object will form, not a point, but a small circular spot in the image, over which the rays are diffused; and as these spots overlap and encroach on each other, distinctness is destroyed. For the formation, therefore, of perfect images, the destruction of aberration is the essential condition; and whatever imperfections, either in the figures of the reflecting or refracting surfaces used, or in the materials of which they are composed, tends to throw the rays aside from their strict geometrical direction, must, of course, confound the images. Hence, in the formation of optical images, there are three great points to be attended to: first, perfect polish of the surfaces; secondly, perfect homogeneity in the material employed; thirdly, strict conformity in the figures of the reflecting and refracting surfaces to geometrical rules, and the results of analysis.

335. There is one case where the aberrations of all kinds are rigorously destroyed, and in which the image is perfect. It is when the rays are reflected at a plane surface. For (fig. 66) if PQ be an object placed before a plane reflector AB , and if perpendiculars be let fall from every point of the object to the surface, and on the other side points in these be taken at the same distances respectively behind the surface as p, q , these points will form the image. Now we have seen, that all rays from any point P , reflected at AB , will after reflexion diverge strictly from p its image. Thus, the image is as perfect and free from aberration as the object; and will appear, to an eye placed so as to receive the rays, like a real object placed behind the reflector.

336. *Corol.* The image formed by a plane reflecting surface is similar and equal to the object, and any corresponding lines in both are equally inclined to the reflecting surface. A common looking-glass is the best illustration of this case.

337. *Proposition.* To determine the image of any object formed by a plane refracting surface. Let BC be the surface, PQ the object. From any point Q draw QC perpendicular to the surface, and, μ being the index of refraction, if we regard the surface as a sphere of infinite radius, we have R its curvature $= 0$, and the equation

$$f = (1 - m) R + m D \text{ becomes simply } f = m D. \text{ Now } f = \frac{1}{Cq}; D = \frac{1}{CQ}; \text{ and } m = \frac{1}{\mu}. \text{ Hence}$$

this equation, translated into geometrical language, gives $Cq = \mu \times CQ$.

338. In the case represented in the figure, the refraction is made out of a denser medium into a rarer, the object being immersed in the denser (as under water), and the eye of a spectator in the rarer (as in air): the image q of the point Q is therefore nearer the surface than Q , (because in this case μ is less than unity.) The same holds good of all other points of the image; so that the whole object will appear raised by refraction, as in the familiar experiment where a shilling is laid in an empty vessel, and the eye withdrawn till the shilling is hidden by the edge, but reappears again, as if raised up, when the vessel is filled with water. On the other hand, to an eye placed under water, external objects would appear farther removed by the effect of refraction.

339. *Corol. 1.* The image of a straight line PQ in the object is a straight line pq in the image, less inclined to the surface if the refraction be made from a denser into a rarer medium. Thus, if a stick $DAPQ$ be partly plunged into water, the immersed portion AQ forms the image Aq less inclined; so that to a spectator in air, the stick appears broken and bent upwards at A . The appearance is familiar to every one.

340. In refraction at a plane surface, however, the rays do not rigorously diverge from, or converge to, a single point. Therefore the above result is only approximately correct, and supposes the rays to be incident nearly at right angles to the surface. And this leads us to the consideration of oblique vision through refracting surfaces, or in reflectors of any figure.

341. The eye sees by the rays which enter it, and judges of the existence of an object, by the fact of rays diverging sensibly from some point in space. If, then, rays diverge rigorously from a point, the eye which receives them is irresistibly led to the belief (unless corrected by experience and judgment) of an object being there; the illusion is complete, and vision perfect. But if such divergence be only approximate, as when the density of the rays which reach the eye in any one direction is very much greater than in directions adjacent on either side, vision is still produced, only less distinct, in proportion to the degree of deviation from strict mathematical divergence of the rays which produce it. Suppose, now, Q to be a radiant point placed anywhere with respect to the refracting or reflecting surface ACB , (fig. 68,) and let $AqFB$ be the caustic formed by the intersection of all the refracted or reflected rays. Let us suppose an eye placed at E , and from thence draw Eq a tangent

Light.

Part I.

to the caustic, which continue to the surface C, and join Q C. Then it is obvious, that any small pencil Q C, Q C' diverging from Q, will form a focus at q (Art. 134, &c.) from which it will afterwards diverge, and fall on the eye at E, nearly as if the rays came from a mathematical point; and from what was said in Art. 161 and 162, it appears that the density of rays in the cone q E is infinitely greater than in any adjacent cone having the eye for its base; so that q will appear as an image of Q, more or less confused, in proportion to the degree of curvature of the caustic at q; for it is evident, that if the curvature be great, the assumed concentration of any small finite pencil Q C C' in one mathematical point q, will deviate more from truth than if the caustic approach nearly to a straight line.

Corol. As the eye shifts its place, the apparent position of an object seen in a reflecting or refracting surface shifts also, for as E varies, the tangent E q shifts its place on the caustic, and the point of contact q, or the place of the image shifts.

This doctrine may be illustrated by a very familiar instance. If we look through a surface of still water, not very deep, but having a level horizontal bottom, the bottom will not appear a plane, but will seem to rise on all sides, and approach nearer the surface the more obliquely we look. To explain this, let Q be a point in the bottom, and let Q P e be the course of the pencil of rays by which an eye at e sees it (fig. 39) on the visual ray. The point in the caustic to which e P produced is a tangent, is Y; and from the form of the caustic D Y B (see Art. 238) it is obvious, that Y is nearer the surface the more oblique e P is to it. The apparent figure of the bottom will therefore be thus determined. From the eye E (fig. 69) draw any line E g to the point G of the surface; and having drawn P Y parallel to E G, touching the branch D Y B of the caustic having Q, vertically below E for a radiant point in Y, prolong E G to H, making G H = P Y, then will H be the image of the point Q' in the bottom, belonging to the caustic D' H B'; and the locus of H, or the apparent form of the bottom, will be the curve D F H, having a basin-shaped curvature at D, a point of contrary flexure at F, and an asymptote C G K coinciding with the surface.

But, to return to the case of images formed by rays incident at very small obliquities and nearly central, the following rules for determining their places, magnitudes, and apparent situations in all cases of spherical surfaces, will be convenient to bear in memory, and will need no express demonstration to the reader of the foregoing pages.

Rule 1. Any image formed, or about to be formed, by converging rays, or from which rays diverge, may be regarded as an object.

Rule 2. In spherical reflectors the object and its image lie on the same side of the principal focus. They move in contrary directions, and meet at the centre and surface of the reflector. The distance of the image from the principal focus and centre is had by the proportion

$$Q F : F E :: E F : F q :: Q E : E q,$$

and the image is erect when the object and surface lie on the same side of the principal focus; but inverted when on contrary sides. The relative magnitudes of the object and image (being as their distances from the centre) are given by the proportion

$$\text{object : image} :: Q F : F E :: \text{distance of the object from the principal focus : focal length of reflector.}$$

Rule 3. In thin lenses, of all species, if Q be the place of the object, q of its image, E the centre of the lens, F the principal focus of rays incident in a contrary direction, then will the object and image lie on the same, or opposite side of the lens, according as the object and lens lie on the same or opposite sides of the principal focus F. In the former case the image is erect, in the latter inverted, with respect to the object. The distance of the image from the lens, or from the object, is had by the proportions

$$Q F : F E :: Q E : E q; \quad Q F : Q E :: Q E : Q q;$$

and the magnitude of the object is to that of the image as the distance of the object from F is to the focal length, or as

$$Q F : F E.$$

Rule 4. In all combinations of reflectors and lenses, the image formed by one is to be regarded as the object, whose image is to be formed by the next, and so on, till we come to the last.

It has been already remarked (Art. 6) that visible objects are distinguished from optical images by this, that from the former light emanates in all directions, whereas in the latter it emanates only in certain directions. This is an important limitation in practical optics. A real object can be seen whenever nothing opaque is interposed between it and the eye. An image can only be seen when the eye is placed in the pencil of rays which goes to form it, or diverges from it. Thus in the case represented in fig. 62, except the eye be placed somewhere in the space D q p H, it will see no part of the image, B q D and A p H being the extreme rays refracted by the lens from the extremities of the object.

The brightness of an image is, of course, proportional to the quantity of light which is concentrated in each point of it; and, therefore, supposing no aberration, as the apparent magnitude of the lens or mirror which forms

it, seen from the object $\times \frac{\text{area of object}}{\text{area of image}}$. Or, since the area of the object : that of the image :: (distance)²

of object from lens : (distance)² of image; and since the apparent magnitude of the lens seen from the object

is as its $\left(\frac{\text{diameter}}{\text{distance from object}} \right)^2$, the brightness or degree of illumination of the image is as the apparent

342.

343.

Apparent figure of the horizontal bottom of still water. Fig. 39. Fig. 69.

344.

Rules for finding the place, &c. of an image. 345.

346.

Rule for reflectors. Fig. 16

347.

Rule for lenses.

34 .

349.

Brightness of images.

Light. magnitude of the lens seen from the image, alone, whatever be the distance of the object. Now the apparent magnitude of the lens seen from the image is always much less than a hemisphere. Therefore (even supposing no light lost by reflection or refraction) the illumination of the image is always much less than that of the object. This is the case when the image is received on a screen which reflects all the rays, or when viewed by an eye behind it having a pupil large enough to receive all the rays which have crossed at the image, *à fortiori*, then, when the eye does not receive all the rays, must the apparent intrinsic brightness be less than that of the object. This supposes the object to have a sensible magnitude; but when both the object and its image are physical points, the eye judges only of absolute light; and the light of the image is therefore proportional to the apparent magnitude of the lens, as seen from the object. In the case of a star, for instance, whose distance is constant, the absolute light of the image is simply as the square of the aperture, and this is the reason why stars can be seen in large telescopes which are too faint to be seen in small ones.

Images are never so bright as their objects

§ XII. Of the Structure of the Eye, and of Vision.

350. It is by means of optical images that vision is performed, that we see. The eye is an assemblage of lenses which concentrate the rays emanating from each point of external objects on a delicate tissue of nerves, called the retina, there forming an image, or exact representation of every object, which is the thing immediately perceived or *felt* by the retina.

Description of the eye. Fig. 70.

Aqueous humour. Its composition. Refractive power. Cornea. Its figure an ellipsoid of revolution.

Fig. 70 is a section of the human eye through its axis in a horizontal plane. Its figure is, generally speaking, spherical, but considerably more prominent in front. It consists of three principal chambers, filled with media of perfect transparency and of refractive powers, differing sensibly *inter se*, but none of them greatly different from that of pure water. The first of these media, A, occupying the anterior chamber, is called the *aqueous humour*, and consists, in fact, chiefly of pure water, holding a little muriate of soda and gelatine in solution, with a trace of albumen; the whole not exceeding eight per cent.* Its refractive index, according to the experiments of M. Chossat,† and those of Dr. Brewster and Dr. Gordon,‡ is almost precisely that of water, viz. 1.337, that of water being 1.336. The cell in which it is contained is bounded, on its anterior side, by a strong, horny, and delicately transparent coat *a*, called the *cornea*, the figure of which, according to the delicate experiments and measures of M. Chossat,§ is an ellipsoid of revolution about the major axis; this axis, of course, determines the *axis of the eye*; but it is remarkable, that in the eyes of oxen, measured by M. Chossat, its vertex was never found to be coincident with the central point of the aperture of the cornea, but to lie always about 10° (reckoned on the surface) inwardly, or towards the nose, in a horizontal plane. The ratio of the semi-axis of this ellipse to the excentricity, he determines at 1.3; and this being nearly the same with 1.337, the index of refraction, it is evident, from what was demonstrated in Art. 236, that parallel rays incident on the cornea in the direction of its axis, will be made to converge to a focus situated behind it, almost with mathematical exactness, the aberration which would have subsisted, had the external surface a spherical figure, being almost completely destroyed.

351. The posterior surface of the chamber A of the aqueous humour is limited by the *iris* $\beta\gamma$, which is a kind of circular opaque screen, or diaphragm, consisting of muscular fibres, by whose contraction or expansion an aperture in its centre, called the *pupil*, is diminished or dilated, according to the intensity of the light. In very strong lights the opening of the pupil is greatly contracted, so as not to exceed twelve hundredths of an inch in the human eye, while in feebler illuminations it dilates to an opening not exceeding twenty-five hundredths,|| or double its former diameter. The use of this is evidently to moderate and equalize the illumination of the image on the retina, which might otherwise injure its sensibility. In animals (as the cat) which see well in the dark, the pupil is almost totally closed in the daytime, and reduced to a very narrow line; but in the human eye, the form of the aperture is always circular. The contraction of the pupil is involuntary, and takes place by the effect of the stimulus of the light itself; a beautiful piece of self-adjusting mechanism, the play of which may be easily seen by approaching a candle to the eye while directed to its own image in a looking-glass.

352. Immediately behind the opening of the *iris* lies the *crystalline lens*, B, enclosed in its *capsule*, which forms the posterior boundary of the chamber A. Its figure is a solid of revolution, having its anterior surface much less curved than the posterior. Both surfaces, according to M. Chossat, are ellipsoids of revolution about their *lesser* axes; but it would seem from his measures, that the axes of the two surfaces are neither exactly coincident in direction with each other, nor with that of the cornea. This deviation would be fatal to distinct vision were the crystalline lens very much denser than the others, or were the whole refraction performed by it. This, however, is not the case; for the mean refractive index of this lens is only 1.384, while that of the aqueous humour, as we have seen, is 1.337; and that of the *vitreous* C, which occupies the third chamber, is 1.339; so that the whole amount of bending which the rays undergo at the surface of the crystalline is small, in comparison with the inclination of the surface at the point where the bending takes place, and, since near the vertex, a

Refraction. Non-coincidence of the axes of its surfaces.

* Chenevix, *Philosophical Transactions*, vol. xciii. p. 195.

† *Bulletin de la Soc. Philomatique*, 1818, p. 94.

‡ *Edinburgh Philosophical Journal*, vol. i. p. 42.

§ *Sur la Courbure des Milieux Réfringens de l'Œil chez le Bouf. Annales de Chim.* x. p. 337.

|| Dr. Young's Lectures on the Mechanism of the Eye, *Philosophical Transactions*, vol. xci.

Light. material deviation in the direction of the axis can produce but a very minute change in the inclination of the ray to the surface, this cause of error is so weakened in its effect, as, probably, to produce no appreciable aberration.

Part I.

Why not injurious to vision.

353.

Composition of crystalline.

Denser towards centre.

The crystalline is composed of a much larger proportion of albumen and gelatine than the other humours of the eye, so much so as to be entirely coagulable by the heat of boiling water. It is somewhat denser towards the centre than at the outside. According to Dr. Brewster and Dr. Gordon, the refractive indices of its centre middle of its thickness, from the centre to the outside, and the outside itself, are respectively 1.3999, 1.3786, and 1.3767, that of pure water being 1.3358. This increase of density is obviously useful in correcting the aberration, by shortening the focus of rays near the centre, according to the rule laid down in Art. 299 for finding the effect of aberration. The effect of the elliptic figure of the surfaces is, however, a matter of pretty complex calculation, and cannot be entered upon in the limits of this essay. Its use is, probably, to correct the aberration of oblique pencils.

The posterior chamber C of the eye is filled with the *vitreous humour*, a fluid differing (according to Chenevix) neither in specific gravity nor in chemical composition in any sensible respect from the aqueous; and, as we have already seen, having a refractive index but very little superior to it.

354.

The refractive density of the crystalline being superior to that of either the aqueous or vitreous humour, the rays which are incident on it in a state of convergence from the cornea, are made to converge more, and exactly in their final focus is the posterior surface of the cell of the vitreous humour covered by the retina *d*, a network (as its name imports) of inconceivably delicate nerves, all branching from one great nerve O, called the optic nerve, which enters the eye obliquely at the inner side of the orbit, next the nose. The retina lines the whole of the cavity C up to *i*, where the capsule of the crystalline commences. Its nerves are in contact with, or immersed in, the *pigmentum nigrum*, a very black velvety matter, which covers the *choroid membrane* *g*, and whose office is to absorb and stifle all the light which enters the eye as soon as it has done its office of exciting the retina; thus preventing internal reflexions, and consequent confusion of vision. The whole of these humours and membranes are contained in a thick tough coat, called the *sclerotica*, which unites with the cornea, and forms what is commonly called the *white of the eye*.

355.

Retina.

Sclerotica.

Such is the structure by which parallel rays, or those emanating from very distant objects, are brought to a focus on the retina. But as we require to see objects near, as well as at a distance, and as the focus of a lens or system of lenses for near objects is longer than for distant ones, it is evident that a power of adjustment must reside somewhere in the eye; by which either the retina can be removed farther from the cornea, and the eye lengthened in the direction of its axis, or the curvature of the lenses themselves altered so as to give greater convergency to the rays. We know that such a power exists, and can be called into action by a voluntary effort; and, evidently, by a muscular action, producing fatigue if long continued, and not capable of being strained beyond a certain point. Anatomists, however, as well as theoretical opticians, differ as to the mechanism by which this is effected. Some assert, that the action of the muscles which move the eye in its orbit, called the *recti*, or straight' muscles, when all contracted at once, producing a pressure on the fluids within, forces out the cornea, rendering it at once more convex, and more distant from the retina. This opinion, however, which has been advocated by Dr. Olbers, and even attempted to be made a matter of ocular demonstration by Ramsden and Sir E. Home, has been combated by Dr. Young, by experiments which show, at least, very decisively, that the increase of convexity in the cornea has little if any share in producing the effect. An elongation of the whole eye, spherical as it is and full of fluid, to the considerable extent required, is difficult to conceive as the result of any pressure which could be safely applied, as to give distinct vision at the distance of three inches from the eye, (the nearest at which ordinary eyes can see well,) the sphere must be reduced to an ellipsoid, having its axis nearly one-seventh longer than in its natural state; and the extension of the sclerotica thus produced, would hardly seem compatible with its great strength and toughness. Another opinion, which has been defended with considerable success by the excellent philosopher last named, is, that the crystalline itself is susceptible of a change of figure, and becomes more convex when the eye adapts itself to near distances. His experiments, on persons deprived of this lens, go far to prove the total want of a power to change the focus of the eye in such cases, though a certain degree of adaptation is obtained by the contraction of the iris, which, limiting the diameter of the pencil, diminishes the space on the retina over which imperfectly converged rays are diffused, and thus, in some measure, obviates the effect of their insufficient convergence. When we consider that the crystalline lens has actually a regular fibrous structure, (as may be seen familiarly on tearing to pieces the lens of a boiled fish's eye,) being composed of layers laid over each other like the coats of an onion, and each layer consisting of an assemblage of fibres proceeding from two poles, like the meridians of a globe, the axis being that of the eye itself; we have, so far at least, satisfactory evidence of a muscular structure; and were it not so, the analogy of pellucid animals, in which no muscular fibres can be discerned, and which yet possess the power of motion and obedience to the nervous stimulus, though nerves no more than muscles can be seen in them, would render the idea of a muscular power resident in the crystalline easily admissible, though nerves have as yet not been traced into it. On the whole, it must be allowed, that the presumption is strongly in favour of this mechanism, though the other causes already mentioned may, perhaps, conspire to a certain extent in producing the effect, and though the subject must be regarded as still open to fuller demonstration. It is the boast of science to have been able to trace so far the refined contrivances of this most admirable organ; not its shame to find something still concealed from its scrutiny; for, however anatomists may differ on points of structure, or physiologists dispute on modes of action, there is that in what we do understand of the formation of the eye so similar, and yet so infinitely superior, to a product of human ingenuity,—such thought, such care, such refinement, such advantage taken of the properties of natural agents used as mere instruments, for accomplishing a given end, as force upon us a conviction of

356.

Change of focus of eye for near objects.

Light.

deliberate choice and premeditated design, more strongly, perhaps, than any single contrivance to be found, whether in art or nature, and render its study an object of the deepest interest.

357.

Image on
the retina
the imme-
diate object
of vision.

The images of external objects are of course formed inverted on the retina, and may be seen there, by dissecting off the posterior coats of the eye of a newly-killed animal, and exposing the retina and choroid membrane from behind, like the image on a screen of rough glass, mentioned in Art. 331. It is this image, and this only, which is *felt* by the nerves of the retina, on which the rays of light act as a stimulus; and the impressions therein produced are thence conveyed along the optic nerves to the sensorium, in a manner which we must rank at present among the profounder mysteries of physiology, but which appears to differ in no respect from that in which the impressions of the other senses are transmitted. Thus, a paralysis of the optic nerve produces, while it lasts, total blindness, though the eye remains open, and the lenses retain their transparency; and some very curious cases of half blindness have been successfully referred to an affection of one of the nerves without the other.* On the other hand, while the nerves retain their sensibility, the degree of perfection of vision is exactly commensurate to that of the image formed on the retina. In cases of cataract, where the crystalline lens loses its transparency, the light is prevented from reaching the retina, or from reaching it in a proper state of regular concentration, being stopped, confused, and scattered by the opaque or semi-opaque portions it encounters in its passage. The image, in consequence, is either altogether obliterated, or rendered dim and indistinct; and the progress of blindness is accordingly. If the opaque lens be extracted, the full perception of light returns; but one principal instrument for producing the convergence of the rays being removed, the image, instead of being formed on the retina, is formed considerably behind it, and the rays being received in their unconverged state on it, produce no regular picture, and therefore no distinct vision. But if we give to the rays, before their entry into the eye, a certain proper degree of convergence, by the application of a convex lens, so as to render the remaining lenses capable of finally effecting their exact convergence on the retina, restoration of distinct vision is the immediate result. This is the reason why persons who have undergone the operation for the cataract (which consists either in totally removing, or in putting out of the way an opaque crystalline) wear spectacles of comparatively very short focus. Such glasses perform the office of an artificial crystalline. A similar imperfection of vision to that produced by the removal of the crystalline, is the ordinary effect of old age, and its remedy is the same. In aged persons the exterior transparent surface of the eye, called the *cornea*, loses somewhat of its convexity, and becomes flatter. The *power* of the eye is therefore diminished, (Art. 248 and 255,) and a perfect image can no longer be formed on the retina. The deficient power is however supplied by a convex lens, or *spectacle-glass*, (Art. 268,) and vision rendered perfect or materially improved.

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Short-sighted persons have their eyes too convex, and this defect is, like the other, remediable by the use of proper lenses of an opposite character. There are cases, however, though rare, in which the cornea becomes so very prominent as to render it impossible to apply conveniently a lens sufficiently concave to counteract its action. Such cases would be accompanied with irremediable blindness, but for that happy boldness, justifiable only by the certainty of our knowledge of the true nature and laws of vision, which in such a case has suggested the opening of the eye and removal of the crystalline lens, though in a perfectly sound state.

359.

Malconfor-
mations of
the cornea.

But these are not the only cases of defective vision arising from the structure of the organ, which are susceptible of remedy. Malconformations of the cornea are much more common than is generally supposed, and few eyes are, in fact, free from them. They may be detected by closing one eye, and directing the other to a very narrow, well-defined luminous object, not too bright, (the horns of the moon, when a slender crescent, only two or three days old, are very proper for the purpose,) and turning the head about in various directions. The line will be doubled, tripled, or multiplied, or variously distorted; and careful observation of its appearances, under different circumstances, will lead to a knowledge of the peculiar conformation of the refracting surfaces of the eye which causes them, and may suggest their proper remedy. A remarkable and instructive instance of the kind has recently been adduced by Mr. G. B. Airy, (*Transactions of the Cambridge Philosophical Society*,) in the case of one of his own eyes; which, from a certain defect in the figure of its lenses, he ascertained to refract the rays to a nearer focus in a vertical than in a horizontal plane, so as to render the eye utterly useless. This, it is obvious, would take place if the cornea, instead of being a surface of revolution, (in which the curvature of all its sections through the axis must be equal,) were of some other form, in which the curvature in a vertical plane is greater than in a horizontal. It is obvious, that the correction of such a defect could never be accomplished by the use of spherical lenses. The strict method, applicable in all such cases, would be to adapt a lens to the eye, of nearly the same refractive power, and having its surface next the eye an exact *intaglio* fac-simile of the irregular cornea, while the external should be exactly spherical of the same general convexity as the cornea itself; for it is clear, that all the distortions of the rays at the posterior surface of such a lens would be exactly counteracted by the equal and opposite distortions at the cornea itself.† But the necessity of limiting the correcting lens to such surfaces as can be truly ground in glass, to render it of any real and everyday use, and which surfaces are only spheres, planes, and cylinders, suggested to Mr. Airy the ingenious idea of a double concave lens, in which one surface should be spherical, the other cylindrical. The use of the spherical surface was to correct the general defect of a too convex cornea. That of the cylindrical may be thus explained. Suppose parallel rays incident on a concave cylindrical surface, A B C D, in a direction perpendicular to its axis, as in fig. 71, and let S S' P P' Q Q' T T', be any laminar pencil of them contained in a parallelepiped infinitely

Remarkable
case, suc-
cessfully
remedied
by glasses.

Fig. 71.

* Wollaston, on Semi-decussation of the Optic Nerves, *Philosophical Transactions*, 1824.

† Should any very bad cases of irregular cornea be found, it is worthy of consideration, whether at least a temporary distinct vision could not be procured, by applying in contact with the surface of the eye some transparent animal jelly contained in a spherical capsule of glass; or whether an actual mould of the cornea might not be taken, and impressed on some transparent medium. The operation would, of course, be delicate, but certainly less so than that of cutting open a living eye, and taking out its contents.

thin, and having its sides parallel to the axis. Any of the rays $SP, S'P'$, of this pencil lying in a plane APS perpendicular to the axis, will after refraction converge to, or diverge from, a point X , also in this plane; and, therefore, all the rays incident on $PQ, P'Q'$, will after refraction have for their focus the line XY , in the caustic surface $AFGD$, and the principal focus of the cylinder will be the line FG , whose distance from the vertex $C C'$ of the surface, or FC , is the same with the focal length of a spherical surface, formed by the revolution of AB about the axis FC . Thus we see that a cylindrical lens produces no convergency or divergency in parallel rays, incidental in the plane of its axis; while it converges or diverges rays in a plane at right angles to the axis, as a spherical surface of equal curvature would do. If then such a cylindrical surface be conjoined with a spherical one, the focus of the spherical surface will remain unaltered in one plane, but in the other will be changed to that of a lens formed by it, and a spherical surface of equal curvature with the cylinder. Hence by properly placing such a cylindro-spheric lens across the defective eye, its defect will be (approximately, at least) counteracted. It would be wrong to conclude our account of this interesting application of mathematical knowledge to the increase of the comforts and improvement of the faculties of its possessor, in other than his own words. "After some ineffectual applications to different workmen, I at last procured a lens to these dimensions,* from an artist named Fuller, at Ipswich. It satisfies my wishes in every respect. I can now read the smallest print at a considerable distance with the left" (the defective) "eye as well as with the right. I have found that vision is most distinct when the cylindrical surface is turned from the eye: and as, when the lens is distant from the eye, it alters the apparent figure of objects by refracting differently the rays in different planes, I judged it proper to have the frame of my spectacles made so as to bring the glass pretty close to the eye. With these precautions, I find that the eye which I once feared would become quite useless, can be used in almost every respect as well as the other."

Blindness, partial or total, may be caused, not only by the opacity of the crystalline lens, but of any other part, or by anything extraneous to the materials of which they consist, interposed between the external transparent surface of the cornea and the retina. In all such cases, if the sensibility of the nerve be uninjured, the restoration of sight is never to be despaired of. In a recent most remarkable case, operated by Mr. Wardrop, and by him recorded in the *Philosophical Transactions* for 1826, blindness from infancy, accompanied with complete obliteration of the pupil, by a contraction of the iris, owing to an unskilful operation, performed at six months of age, was removed, and perfect sight restored after a lapse of forty-six years, by a simple removal of the obstruction, by breaking a hole through the closed membrane. The details of this case are in the highest degree interesting, but we must refer the reader to the volume of the *Philosophical Transactions* cited for the account.

As we have two eyes, and a separate image of every external object is formed in each, it may be asked, *why do we not see double?* and to some, the question has appeared to present much difficulty. To us it appears, that we might with equal reason ask, *why*—having two hands, and five fingers on each, all endowed with equal sensibility of touch and equal aptitude to discern objects by that sense—*we do not feel decuple?* The answer is the same in both cases: it is a matter of habit. Habit alone teaches us that the sensations of sight correspond to any thing external, and to what they correspond. An object (a small globe or wafer suppose) is before us on a table; we direct our eyes to it, *i. e.* we bring its images on both retinae to those parts which habit has ascertained to be the most sensible and best situated for seeing distinctly; and having always found that in such circumstances the object producing the sensation is one and the same, the idea of unity in the object becomes irresistibly associated with the impression. But while looking at the globe, squeeze the upper part of one eye downwards, by pressing on the eyelid with the finger, and thereby forcibly throw the image on another part of the retina of that eye, and double vision is immediately produced, two globes or two wafers being distinctly seen, which appear to recede from each other as the pressure is stronger, and approach, and finally blend into one as it is relieved. The same effect may be produced without pressure, by directing the eyes to a point nearer to, or farther from them than the wafer; the optic axes in this case being both directed away from the object seen. When the eyes are in a state of perfect rest, their axes are usually parallel, or a little diverging. In this state all near objects are seen double; but the slightest effort of attention causes their images to coalesce immediately. Those who have one eye distorted by a blow, see double, till habit has taught them anew to see single, though the distortion of the optic axis subsists.

The case is exactly the same with the sense of touch. Lay hands on the globe, and handle it. It is one, nothing can be more irresistible than this conviction. Place it between the first and second fingers of the right hand in their natural position. The right side of the first and left of the second finger feel opposite convexities; but as habit has always taught us that two convexities so felt belong to one and the same spherical surface, we never hesitate or question the identity of the globe, or the unity of the sensation. Now cross the two fingers, bringing the second over the first, and place the globe on the table in the fork between them, so as to feel the left side of the globe with the right side of the second finger, and the right with the left of the first. In this state of things the impression is equally irresistible, that we have two globes in contact with the fingers, especially if the eyes be shut, and the fingers placed on it by another person. A pea is a very proper object for this experiment. The illusion is equally strong when the two fore fingers of both hands are crossed, and the pea placed between them.

So forcible is the power of habit in producing single vision, that it will bring the two images to apparent coalescence, when the rays which form one of them are really turned far aside from their natural course. To show this, place a candle at a distance, and look at it with one eye (the left suppose) naked, the other having

360.

361.

Single vision with two eyes.

Double vision artificially produced.

Another method.

362.

Single objects felt double in certain cases

363.

Force of habit in producing single vision illustrated by experiment.

* Radius of the spherical surface $3\frac{1}{2}$ inches, of the cylindrical $4\frac{1}{2}$.

Light. before it a prism, with a variable refracting angle, (an instrument to be described hereafter, see INDEX,) and, first, let the angle be adjusted to zero, then will the prism produce no deviation, and the object will appear single. Now vary the prism, so as to produce a deviation of 2° or 3° of the rays in a horizontal plane to the right. The candle will immediately be seen double, the image deviated by the prism being seen to the left of the other; but the slightest motion, such as winking with the eyelids, blends them immediately into one. Again, vary the prism a few degrees more in the same direction; the candle will again be doubled, and again rendered single by winking, and directing the attention more strongly to it; and thus may the optic axes be, as it were, inveigled to an inclination of 20° or 30° to each other. In this state of things, if a second candle be placed exactly in the direction of the deviated image of the first, but so screened, that its rays shall not fall on the left eye, and the prism be then suddenly removed in the act of winking, the two candles appear as one. If the deviation of the image seen with the right eye be made to the apparent right, the range within which it is possible to bring them to coalesce is much more limited, as it is much more usual for us to direct by an effort the optic axes towards, than from each other. If the deviation be made but a very little out of the horizontal plane, no effort will enable us to correct it. It is probable that some cases of squinting might be cured by some such exercise in the art of directing the optic axes, if continued perseveringly.

364. Such is, undoubtedly, a sufficient explanation of single vision with two eyes; yet Dr. Wollaston has rendered it probable that a physiological cause has also some share in producing the effect, and that a semi-decussation of the optic nerves takes place immediately on their quitting the brain, half of each nerve going to each eye, the right half of each retina consisting wholly of fibres of one nerve, and the left wholly of the other, so that all images of objects out of the optic axis are perceived by one and the same nerve in both eyes, and thus a powerful sympathy and perfect unison kept up between them, independent of the mere influence of habit. Immediately in the optic axis, it is probable, that the fibres of both nerves are commingled, and this may account for the greater acuteness and certainty of vision in this part of the eye.

365. Another point, on which much more discussion has been expended than it deserves, is the fact of our seeing objects erect when their images on the retina are inverted. Erect, means nothing else than having the head farther from the ground, and the feet nearer, than any other part. Now, the earth, and the objects which on it, preserve the same relative situation in the picture on the retina that they do in nature. In the case of men, it is true, stand with their heads downwards; but then, at the same time, heavy bodies fall upwards; and the mind, or its deputy, the nerve, which is present in every part of the picture, judges only of the relations of its parts to one another. How these parts are related to external objects, is known only by experience, and judged of at the instant only by habit.

366. There is one remarkable fact which ought not to escape mention, even in so brief an abstract of the doctrine of vision as the present, it is, that the spot \bigcirc , at which the optic nerve enters the eye, is totally insensible to the stimulus of light, for which reason it is called the *punctum cæcum*. The reason is obvious: at this point the nerve is not yet divided into those almost infinitely minute fibres, which are fine enough to be either thrown into tremors, or otherwise changed in their mechanical, chemical, or other state, by a stimulus so delicate as the rays of light. The effect, however, is curious and striking. On a sheet of black paper, or other dark ground, place two white wafers, having their centres three inches distant. Vertically above that to the left, hold the right eye, at 12 inches from it, and so that when looking down on it, the line joining the two eyes shall be parallel to that joining the centre of the wafers. In this situation closing the left eye, and looking full with the right at the wafer perpendicularly below it, this only is seen, the other being completely invisible. But if removed ever so little from its place, either to the right or left, above or below, it becomes immediately visible, and starts, as it were, into existence. The distances here set down may perhaps vary slightly in different eyes.

367. It will cease to be thought singular, that this fact, of the absolute invisibility of objects in a certain point of the field of view of each eye, should be one of which not one person in ten thousand is apprized, when we learn, that it is not extremely uncommon to find persons who have for some time been totally blind with one eye without being aware of the fact. One instance has fallen under the knowledge of the writer of these pages.

368. In the eyes of fishes, the humours being nearly of the refractive density of the medium in which they live, the refraction at the cornea is small, and the work of bringing the rays to a focus on the retina is almost wholly performed by the crystalline. This lens, therefore, in fishes is almost spherical, and of small radius, in comparison with the whole diameter of the eye. Moreover, the destruction of spherical aberration not being producible in this case by mere refraction at the cornea, the crystalline itself is adapted to execute this necessary part of the process, which it does by a very great increase of density towards the centre. (Brewster, *Treatise on New Philosophical Instruments*, p. 268.) The fibrous and coated structure of the crystalline lens is beautifully shown in the eye of a fish coagulated by boiling.

369. The same scientific principles which enable us to assist natural imperfections of sight, can be employed in giving additional power to this sense, even in individuals who enjoy it naturally in the greatest perfection. It being once understood, that the image on the retina is that which we really see, it follows, that if by any means we can render this image brighter, larger, more distinct than in the natural state of the organ, we shall see objects brighter than in their natural state, enlarged in dimension, and, therefore, capable of being examined more in detail, or more sharply defined and clearly outlined. The means which the principles already detailed put in our power, for the accomplishment of such ends, are the concentration of more rays than enter the natural eye by lenses; the enlargement of the image on the retina, by substituting for the object seen an image of it, either larger than the object itself, or capable of being brought nearer to us; and the destruction of aberration, by properly adapting the figure and materials of our instruments to the end proposed.

370. *Proposition.* The apparent magnitude of a rectilinear object is measured by the angle subtended by it at

Light. the centre of the eye, or by the linear magnitude of its image on the retina, and is therefore proportional Part I.
to $\frac{\text{linear magnitude of object}}{\text{its distance from the eye}}$.

The *centre* of the eye, in its optical sense, is a point nearly in the centre of the pupil in the plane of the iris, and the image of any external object PQ , being formed at the bottom of the eye at $p q$, by rays crossing there, Fig. 72.

must subtend the same angle; so that $p q = P Q \cdot \frac{p E}{P E}$.

Corol. If the object be so distant that the rays from each point of it may be regarded as parallel, the angular diameter of the object is, measured by the inclination of rays of its extreme pencils to each other. Whenever, therefore, the eye sees by parallel, or very nearly parallel, rays, the apparent magnitude of the object seen, is measured by the inclination of its extreme pencils, and the object itself is referred to an infinite distance, or to the concave surface of the heavens. 371.

Prop. When a convex lens is placed between the eye and any object, so as to have the object at a distance from the lens equal to its focal length, it will be distinctly seen by an eye capable of converging parallel rays, and will appear enlarged beyond its natural size. 372.

Let PQ be the object, C the lens, and E the centre of the eye. Since the object is in the focus of the lens, Fig. 73. the rays of a pencil diverging from any point P in it, will emerge parallel to PC , and to each other; they will, therefore, after refraction in the eye, be brought to converge on the retina to a point p , such that Ep is parallel to PC . Similarly, rays from Q will, after refraction through the lens and eye, converge to q ; such that Eq is parallel to QC . Thus, a distinct image will be formed at $p q$ on the retina, and the apparent angular magnitude of the object seen through the lens will be the angle $q E p$. Now this is equal to PCQ , or the angle subtended by the object at the centre of the lens, and is, therefore, greater than PEQ , or that subtended by it at the centre of the eye, because the lens is between the eye and object.

Hence, the nearer the eye is to the lens, the less will be the difference between the apparent magnitudes of the object, as seen with and without the lens interposed. But if the lens be of shorter focus than the least distance at which the eye can see distinctly, there will be this essential difference between vision with and without the lens, that in the former case the object is seen distinctly, and well-defined; while in the latter, or with the naked eye, it will be indistinct and confused, and the more so the nearer it is brought. 373.

Hence, by the use of a convex lens of short focus, objects may be seen distinct, and magnified to any extent we please: for let L be the power, or reciprocal focal length of the lens, and D the greatest proximity of the object to the centre of the eye at which it can be seen distinctly without a lens. Then we shall have $L : D :: \text{angle } p E q : \text{angle subtended by the object at the proximity } D$; and, therefore, $:: \text{apparent linear magnitude of object seen through the lens} : \text{apparent linear magnitude at proximity } D, \text{ with the naked eye.}$ Therefore $\frac{L}{D}$ is 374.

the ratio of these magnitudes, or, as it is called, the *magnifying power* of the lens, beyond that of the naked eye, at its greatest proximity. Magnifying power. 375.

Corol. D being given, the magnifying power is as L , or as $(\mu - 1)(R' - R'')$. This explains the use of the word power in the foregoing sections. Whatever we have demonstrated of the powers of lenses in the foregoing pages, is true of *magnifying powers*. Thus the sum of the magnifying powers of two convex lenses is the magnifying power of the two combined. If one be concave, its magnifying power is to be regarded as negative, and instead of their sum we must take their difference. Magnifying power of a system of lenses. 375.

Prop. To express, generally, the *visual angle* under which a small object placed at any distance from a lens, and seen by an eye any how situated, appears, supposing it seen distinctly. 376.

Let PQ , fig. 74, 75, 76, 77, be the object, E the lens, O the eye, and $p q$ the image. Put $\frac{1}{E Q} = D$, $\frac{1}{E q} = \text{Visual angle.}$ Fig. 74, 75, 76, 77. Then $= f; \frac{1}{E O} = e$; e being reckoned in the same direction from the centre of the lens that D and f are. 376.

the visual angle under which the image is seen is $q O p$, and we have, therefore, visual angle ($= A$) $= \frac{q p}{O q} =$

$\frac{q p}{O E - E q}$. But, $q p = Q P \cdot \frac{E q}{E Q} = Q P \cdot \frac{D}{f} = O \cdot \frac{D}{f}$ putting O for $Q P$ the linear magnitude of the object; and, moreover, $O E - E q = \frac{1}{e} - \frac{1}{f} = \frac{f - e}{f e}$, therefore we have, Vision through convex lenses. 377.

$$A = O \cdot \frac{D}{f} \cdot \frac{e f}{f - e} = O \cdot \frac{e D}{L + D - e}$$

when L , as all along, represents the power of the lens. Now $O \cdot D$ is the visual angle of the object, as seen from the centre of the lens; therefore, putting $O \cdot D$, or $\frac{Q P}{Q E} = (A)$ we get

$$A = (A) \cdot \frac{e}{L + D - e} \quad (a)$$

Light.

In concave lenses, the images of distant objects are formed erect, and on the same side of the lens with the object.

377. If, therefore, such a lens be held between the eye and distant objects at a sufficient distance from the eye for distinct vision, the objects will be seen *erect*, and *diminished* in magnitude. In this case, e is positive, and L and D both negative; therefore $L + D - e$ is a negative quantity, greater (without regard to the sign) than e , and, consequently, A is negative, and less than (A) .

378. In reflectors, $f = 2R - D$, and, therefore,
In reflectors.

$$A = (A) \cdot \frac{e}{2R - D - e} \quad (b)$$

In a convex reflector, e is necessarily negative, at least if the mirror be made of metal, because the eye must be on the side of the surface towards the incident light; and, therefore, $2R - e$ is positive, and $\frac{e}{2R - D - e}$

will be greater or less than unity, according to the value of $2R - D - e$. In concave reflectors, R is negative, and e is also negative for the same reason as in concave; therefore the sign and magnitude of A in this, as well as the former case, may vary indefinitely, according to the place of the eye, the image, and the object. The varieties of these cases are represented in fig. 78 and 79.

379.
General principles of telescopes.

If the image, instead of being seen directly by the naked eye, be seen through the medium of another lens or reflector, so placed as to cause the pencils diverging primarily from each point of the object, to emerge finally, either exactly parallel, or within such limits of convergence or divergence as the eye can accommodate itself to, the object will be seen distinctly, and either larger or smaller than it would be seen by the unassisted eye, according to the magnitude of the image, and the power of the lens or reflector used to view it. This is the principle of all telescopes and microscopes. As most eyes can see with parallel rays, they are so constructed as to make parallel pencils emerge parallel; and a mechanical adjustment allows such a quantity of motion of the lenses or reflectors with respect to each other, as to give the rays a sufficient degree of convergence or divergence as may be required.

380.
Astronomical telescope.

In the common refracting, or, as it is sometimes called, the astronomical telescope, the image is first formed by a convex lens, and is viewed through a convex lens, placed at a distance from the other nearly equal to the sum of their focal lengths. The lens which forms the image is called the *object-glass*, and that through which it is viewed, the *eye-glass* of the telescope. If the latter be concave, the telescope is said to be of the Galilæan construction, such having been the original arrangement of Galilæo's instruments. The situation of the lenses, and the course of the rays in these two constructions, are represented in fig. 80 and 81.

Fig. 80, 81.

381.

In the former construction, let PQ be the object. Draw QOG through the centres of the object and eye-glass, and this line will be the axis of the telescope. From R any point in the object draw POp through the centre O of the object-glass, and meeting pq , a line through q , the focus of the point Q , perpendicular to the axis in p , then will pq be the image of PQ . Let PA, PB be the extreme rays of the pencil diverging from P , and incident on the object-glass, and they will be refracted to and cross at p . Hence, unless the diameter of the eye-glass bGa be such, that the ray Ap shall be received on it, the point p will be seen less illuminated than the point Q in the centre of the object, and if it be so small that the line Bp produced does not meet it, then none of the rays from P can reach the eye at all. Thus, the *field of view*, or angular dimensions of the object seen, is limited by the aperture of the eye-glass. To find its extent, then, join Bb, Aa , opposite extremities of the object and eye-glass, meeting the image in r and p , and the axis in X , then rp is the whole extent of the image which is seen at all, and the angle POr , which is equal to POR , is the angular extent of the field of view. Now we have $AB : ab :: OX : GX$, and,

Field of view.

therefore, $AB + ab : AB :: OG : OX$, whence we get $OX = \frac{AB}{AB + ab} \cdot OG$; $GX = \frac{ab}{AB + ab} \cdot$

OG . But we have, moreover, $Xq = Oq - OX$; $pr = ab \cdot \frac{Xq}{GX}$, and angle $rop = \frac{rp}{Oq}$. To express this algebraically, put

Diameter of object-glass = a ; Power of object-glass = L

Diameter of eye-glass = β ; Power of eye-glass = l .

Then

$$\left. \begin{aligned} OX &= \frac{a}{a + \beta} \left(\frac{1}{L} + \frac{1}{l} \right); & GX &= \frac{\beta}{a + \beta} \left(\frac{1}{L} + \frac{1}{l} \right) \\ QX &= \frac{1}{a + \beta} \left(\frac{\beta}{L} - \frac{a}{l} \right); & \text{and } pr &= \frac{\beta l - aL}{L + l}; \end{aligned} \right\} \quad (c)$$

This last is the linear magnitude of the visible portion of the image; and it is, as we see, symmetrical both with respect to the eye-glass and object-glass.

382. Now from this it is easy to deduce both the field of view and magnifying power of the telescope; for the former is equal to the angle subtended by pr , at the centre of the object-glass, and the latter is obtained from the former, when the angle rop subtended at the centre of the eye-glass is obtained. But we have

Light.

$$\left. \begin{aligned} rOp &= L \cdot \frac{\beta l - aL}{L + l}; & rGp &= l \cdot \frac{\beta l - aL}{L + l} \\ \text{magnifying power} &= \frac{rGp}{rOp} = \frac{l}{L} \end{aligned} \right\} \quad (d)$$

Part I.
Formulæ
for field of
view and
magnifying
power.

Hence we see, that the greater the power of the eye-glass is, compared with that of the object-glass, the greater the magnifying power of the telescope; or, in other words, the greater the focal length of the object glass compared with that of the eye-glass.

The pencils of rays after refraction at the eye-glass will emerge parallel, and therefore proper for distinct vision to an eye properly placed to receive them. Now the eye will receive both the extreme rays bR' and aP' of the pencils diverging from r and p , if it be placed at their point of concurrence E ; but since bE is parallel to rG , and aE to pG , we have

$$GE = Gq \times \frac{ab}{pr}, \quad \text{or } GE = \frac{\beta(L+l)}{\beta l - aL}. \quad (e)$$

If the eye be placed nearer to, or farther off from, the eye-glass than this distance, it will not receive the extreme rays, and the *field of view*, or visible area of the object, will be lessened. In the construction of convex single eye-pieces, therefore, care must be taken to prolong the tube which carries them, (as in the figure,) so that when the eye is applied close to its end, it shall still be at this precise distance from the glass.

If the telescope be inverted, and the eye applied behind the object-glass, it is evident that it will remain a telescope, but its magnifying power will be changed to $\frac{L}{l}$; so that, if it magnified before, it will diminish objects now, and the field of view will be proportionally increased. In this way, beautiful miniature pictures of distant objects may be seen.

If the telescope, instead of being turned on objects so distant as that the pencils flowing from them may be regarded as parallel, be directed to near objects, the distance between the object-glass and eye-glass must be lengthened so as to bring the image exactly into the focus of the latter. To accomplish this, the eye-glass is generally set in a sliding tube movable by a rack-work, or by hand. The same mechanism serves also to adjust the telescope for long or short-sighted persons. The former require parallel or slightly divergent rays, the latter very divergent; and to obtain the necessary divergence for the latter, the eye-glass must be brought nearer the object-glass.

The same theory and formulæ apply to the second, or Galilæan, construction, only recollecting that in this case L , the power of the eye-glass, is negative. In this case, therefore, the value of GE is negative, or the eye should be placed between the object-glass and eye-glass; but, as that is incompatible with the other conditions, in order to get as great a field of view as possible, the eye must be brought as near to its proper place as possible, and therefore close to the eye-glass.

In the astronomical telescope objects are seen inverted, in the Galilæan, erect; for, in the former, the rays from the extremities of the object have crossed before entering the eye, in the latter, not.

If the object be brought nearer the object-glass, the magnifying power is increased; because in this case (calling D the proximity of the object) $\frac{l}{L-D}$ expresses the magnifying power, as is easily seen from what has

been said Art. 382. Thus a telescope used for viewing very near objects becomes a *microscope*. The ordinary construction of the compound microscope is nothing more than that of the astronomical telescope modified for the use it is intended for. The object-glass has in this instrument a much greater power than the eye-glass, so that, when employed for viewing distant objects, it acts as a telescope inverted, and requires to be greatly shortened. But for near objects, as D increases, $l-D$ diminishes, and the fraction $\frac{l}{l-D}$ may be increased to any amount, by bringing the object nearer to the object-glass, and at the same time lengthening the interval between the lenses, which is expressed by $\frac{1}{L-D} + \frac{1}{l}$. But as this requires two operations, it is

usual to leave the latter distance unaltered, and vary, by a screw or rack-work, only the former. Fig. 82 is a section of such an instrument. It is, however, convenient to have the power of lengthening and shortening the distance between the glasses, as by this means any magnifying power between the limits corresponding to the extreme distances may be obtained; and if a series of object-glasses be so selected, that the greatest power attainable by one within the limits of the adjustment in question, shall just surpass the least obtainable by the next, and so on, we may command any power we please. Such a series is usually comprised in a small revolving plate containing cells, each of which can be brought in succession into the axis of the microscope by a simple mechanism.

In the reflecting telescope, of the most simple construction, the image is formed by a concave mirror, and viewed by a convex or concave eye-glass, as in refracting telescopes; but since the head of the observer would intercept the whole of the incident light in small telescopes, and a great part of it in large ones, the axis of the reflector itself is turned a little obliquely, so as to throw the image aside, by which it can be viewed with little or no loss of light. The inconvenience of this is a little distortion of the image, caused by the obliquity of the rays;

Light. but as such telescopes are only used of a great size, and for the purpose of viewing very faint celestial objects, in which the light diffused by aberration is insensible, little or no inconvenience is found to arise from this cause.

Simplest, or Herschelian construction.

391. Newtonian construction.

Such is the construction of the telescopes used by Sir William Herschel in his sweeps of the heavens. To obviate the inconvenience of the stoppage of rays by the head, Newton, the inventor of reflecting telescopes, employed a small mirror, placed obliquely, as in fig. 83, opposite the centre of the large one. Thus parallel rays PA, PB, emanating from a point in the axis of the telescope, are received, before their meeting, on a plane mirror CD inclined at 45° to the axis, and thence reflected through a tube projecting from the side of the telescope to the lens G, and by it refracted to the eye E. It is manifest, that if the image formed by the mirror AB behind CD be regarded as an object, an image equal and similar to it (Art. 335) will be formed at F, at an equal distance from the plane mirror; and this image will be seen through the glass G, just as if it were formed by an object-glass of the same focal length placed in the prolongation of the axis of the eye-tube, beyond the small mirror, (supposed away.) Hence the same propositions and formulæ will hold good in the Newtonian telescope, as in the astronomical and Galilean, for the magnifying power, field of view, and position of the eye, substituting only $2R$ for L , and $2R - D$ for $L - D$, and recollecting that R is negative, as the mirror has its concavity turned towards the light.

392. Gregorian telescope. Fig. 84.

The Gregorian telescope, instead of a small plain mirror turned obliquely, has a small convex mirror with its concavity turned towards that of the large one, as in fig. 84; but instead of being placed at a distance from the large one equal to the sum of the focal lengths, the distance is somewhat greater; hence the image $p q$, formed in the focus of the great mirror, being at a distance from the vertex of the small one greater than its focal length, another image is formed at a distance, viz. at or near the surface of the great mirror, at $r s$. In the centre of the large mirror there is a hole which lets pass the rays to an eye-lens g . The adjustment to parallel or diverging rays, or for imperfect eyes, is performed by an alteration of the distance between the mirrors made by a screw.

393. Cassegrainian.

The Cassegrainian construction differs in no respect from the Gregorian, except that the small mirror is convex and receives the rays before their convergence to form an image. The magnitude of the field, the distance of the eye and of the mirrors from each other, are easily expressed in these constructions; the latter being derived from the former by a mere change of sign in the curvature of the small mirror. Let then R' and R'' be the curvatures of the two mirrors, then in the Gregorian telescope R' is negative and R'' positive; and if we put t for the distance between their surfaces, (t being negative, because the second reflecting surface lies towards the incident light) we shall have for an object whose proximity is D

$$D' = D; \quad f' = 2R' - D = 2R' - D; \quad f'' = 2R'' - D''; \quad D'' = \frac{f'}{1 - f't},$$

adopting the formulæ and notation of Art. 251. Now these give, by substitution,

$$\begin{aligned} D'' &= \frac{2R' - D}{1 - t(2R' - D)}; \quad f'' = 2R'' - \frac{2R' - D}{1 - t(2R' - D)} = \\ &= \frac{2R'' - 2R' + D - 2t(2R' - D) \cdot R'}{1 - t(2R' - D)}. \end{aligned} \quad (f)$$

This is the reciprocal distance of the second image from the second reflecting surface. If we wish that the image to be viewed by the eye-lens should fall *just* on the surface of the large mirror, we have only to put $f'' =$

$\frac{1}{-t}$ (because f'' is positive, and t negative.) For parallel rays this gives

$$R'R'' \cdot t^2 + (4R' - 2R'')t - 1 = 0; \quad (g)$$

whence t may be found when R' and R'' are given, or *vice versa*.

394.

The description of other optical instruments, and of the more refined construction of telescopes, &c. must be deferred till we are farther advanced in our account of the physical properties of light, and especially of the different refrangibility of its rays and their colours, which will form the object of the next part.

PART II.

CHROMATICS.

§ I. *Of the Dispersion of Light.*

HITHERTO we have regarded the refractive index of a medium as a quantity absolutely given and the same for all rays refracted by the medium. In nature, however, the case is otherwise. When a ray of light falls obliquely on the surface of a refracting medium, it is not refracted entirely in one direction, but undergoes a separation into several rays, and is *dispersed* over an angle more or less considerable, according to the nature of the medium and the obliquity of incidence. Thus if a sunbeam SC be incident on the refracting surface AB, and be afterwards received on a screen RV. (fig. 85,) it will, instead of a single point on the screen as R, illuminate a space RV of a greater extent the greater is the angle of incidence. The ray SC, then, which, before refraction was single, is separated into an infinite number of rays CR, CO, CY, &c. each of which is refracted differently from all the rest.

395
General phenomenon of separation of a ray into colours.
Fig. 85.

The several rays of which the dispersed beam consists, are found to differ essentially from each other, and from the incident beam, in a most important physical character. They are of different colours. The light of the sun is white. If a sunbeam be received directly on a piece of paper, it makes on it a white spot; but if a piece of white paper (that is, such as by ordinary daylight appears white) be held in the dispersed beam, as RV, the illuminated portion will be seen to be differently coloured in different parts, according to a regular succession of tints, which is always the same, whatever be the refracting medium employed.

396.

To make the experiment in the most striking and satisfactory manner, procure a triangular prism of good flint-glass, and having darkened a room, admit a sunbeam through a small round hole OP in the window shutter. If this be received on a white screen D at a distance, there will be formed a round white spot, or image of the sun, which will be larger as the paper is farther removed. Now in the beam before the screen place the prism ABC, having one of its angles C downwards and parallel to the horizon, and at right angles to the direction of the sunbeam, and let the beam fall on one of its sides BC obliquely. It will be refracted and turned out of its course, and thrown upwards, pursuing the course FGR, and may be received on a screen E properly placed. But on this screen there will no longer be seen a white round spot, but a long streak, or, as it is called in Optics, a *spectrum* RV of most vivid colours, (provided the admitted sunbeam be not too large, and the distance of the screen from the prism considerable.) The tint of the lower or *least refracted* extremity R is a brilliant red, more full and vivid than can be produced by any other means, or than the colour of any natural substance. This dies away first into an orange, and this passes by imperceptible gradations into a fine pale straw-yellow, which is quickly succeeded by a pure and very intense green, which again passes into a blue, at first of less purity, being mixed with green, but afterwards, as we trace it upwards, deepening to the purest indigo. Meanwhile, the intensity of the illumination is diminishing, and in the upper portion of the indigo tint is very feeble, but it is continued still beyond, and the blue acquires a pallid cast of purplish red, a livid hue more easily seen than described, and which, though not to be exactly matched by any natural colour, approaches most nearly to that of a fading violet: "*tinctus violâ pallor.*"

397.
Fig. 86

If the screen on which the spectrum be received have a small hole in it, too small to allow the whole of the spectrum to pass, but only a very narrow portion of it, as X, (fig. 87,) the portion of the beam which goes to form that particular spot X may be received on another screen at any distance behind it, and will there form a spot *d* of the very same colour as the part X of the spectrum. Thus if X be placed in the red part of the spectrum the spot *d* will be red; if in the green, green; and in the blue, blue. If the eye be placed at *d*, it will see through the hole an image of the sun of dazzling brightness; not, as usually, white, but of the colour which goes to form the spot X of the spectrum. Thus we see, that the joint action of all the rays is not essential to the production of the coloured appearance of the spectrum, but that one colour may be insulated from the rest, and examined separately.

398.
Insulation of each colour.

If, instead of receiving the ray X*d*, transmitted through the hole X, on a screen immediately behind it, it be intercepted by another prism *acb*, it will be refracted, and bent from its course, as in X*fgx*; and after this second refraction may be received on a screen *e*. But it is now observed to be no longer separated into a coloured spectrum like the original one RV, of which it formed a part. A single spot *x* only is seen on the screen, the colour of which is uniform, and precisely that which the part X of the spectrum would have had, were it intercepted on the first screen. It appears, then, that the ray which goes to form any single point of the spectrum is not only independent of all the rest, but having been once insulated from them, is no longer capable of further separation into different colours by a second refraction.

399.
Second refraction producing no further change of colour

This simple, but instructive experiment, then, makes us acquainted with the following properties of light:

400.

Light.
Rays of
light differ
in refrangi-
bility.

1. A beam of white light consists of a great and almost infinite variety of rays differing from each other in colour and refrangibility.

For the ray SF from any one point of the sun's disc, which if received immediately on the screen would have occupied only a single point on it, or (supposing the hole in the screen to have a sensible diameter) only a space equal to its area, is dilated into a line VR of considerable length, every point of which (speaking loosely) is illuminated. Now the rays which go to V must necessarily have been more refracted than those which go to R , which can only have been in virtue of a peculiar quality in the rays themselves, since the refracting medium is the same for all.

401. 2. White light may be *decomposed*, *analyzed*, or separated into its elementary coloured rays by refraction. The act of such separation is called the *dispersion* of the coloured rays.

402. 3. Each elementary ray once separated and insulated from the rest, is incapable of further decomposition or analysis by the same means. For we may place a third, and a fourth, prism in the way of the twice refracted ray gr , and refract it in any way, or in any plane; it remains undispersed, and preserves its colour quite unaltered.

403. 4. The dispersion of the coloured rays takes place in the plane of the refraction; for it is found that the spectrum VR is always elongated in this plane. Its breadth is found, on the other hand, by measurement, to be precisely the same as that of the white image D , (fig. 86,) of the sun, received on a screen at a distance OD from the hole, equal to $OF + FG + GR$, the whole course of the refracted light, which shows that the beam has undergone no contraction or dilation by the effect of refraction in a plane perpendicular to the plane of refraction.

404. To explain all the phenomena of the colours produced by prismatic dispersion, or of the *prismatic* colours, as they are called, we need only suppose, with Newton, that each particular ray of light, in undergoing refraction at the surface of a given medium, has the sine of its angle of incidence to that of refraction in a constant ratio, so long as the medium and the ray are the same; but that this ratio varies not only, as we have hitherto all along assumed, with the nature of the medium, but also with that of the ray. In other words, that there are as many distinct species, or at least varieties of light, as there are distinct illuminated points in the spectrum into which a single ray of white light is dispersed. This amounts to regarding the quantity μ , for any medium, not as one and invariable, but as susceptible of all degrees of magnitude between certain limits: one, the least of which, corresponds to the extreme, or least refracted red ray; the other, the greatest value of μ , to the extreme or most refracted violet. Each of these varieties separately conforms to the laws of reflexion and refraction we have already laid down. As in Geometry we may regard a whole family of curves as comprehended under one equation, by the variation of a constant parameter; so in Optics we may include under one analysis all the doctrine of the reflexions, refractions, and other modifications of a ray of white or compound light, by regarding the refractive index μ as a variable parameter.

405. To apply this, for instance, to the experiment of the prism just related: A single ray of white light being supposed incident on the first surface, must be regarded as consisting of an infinite number of coincident rays, of all possible degrees of refrangibility between certain limits, any one of which may be indifferently expressed by the refractive index μ . Supposing the prism placed so as to receive the incident ray perpendicularly on one surface, then the deviation will be given by the equation

$$\mu \cdot \sin I = \sin (I + D)$$

I being the refracting angle of the prism. D therefore is a function of μ , and if μ vary by the infinitely small increment $\delta \mu$, *i. e.* if we pass from any one ray in the spectrum to the consecutive ray, D will vary by δD , and the relation between these simultaneous changes will be given by the equation resulting from the differentiation of the above with the characteristic δ : thus we get

$$\delta \mu \cdot \sin I = \delta D \cdot \cos (I + D); \quad \delta D = \delta \mu \cdot \frac{\sin I}{\cos (I + D)} \quad (a)$$

It is evident, then, that as μ varies, D also varies; and, therefore, that no two of the refracted and coloured rays will coincide, but will be spread over an angle, in the plane of refraction, the greater, the greater is the total variation of μ from one extreme to the other.

406. In order to justify the term *analysis*, or *decomposition*, as applied to the separation of a beam of white light into coloured rays, we must show by experiment that white light may be again produced by the *synthesis* of these elementary rays. The experiment is easy. Take two prisms ABC , abc of the same medium, and having equal refracting angles, and lay them very near together, having their edges turned opposite ways, as in fig. 87. With this disposition, a parallel beam of white light intromitted into the face AC of the first prism, will emerge from the face bc of the last, undeviated, and colourless, as if no prisms were in the way. Now the dispersion having been fully completed by the prism ABC , the rays in passing through the thin lamina of air Bca must have existed in their coloured and independent state, and been dispersed in their directions; but being refracted by the second prism so as to emerge parallel, the colour is destroyed by the mixture and confusion of the rays.

Fig. 88. To see more clearly how this takes place in fig. 88, let SR and SV be two parallel white rays, incident on the first prism, and separated by refraction; the former into the coloured pencil Rc , the latter into a pencil exactly similar to Vrc . Let Rc be the *least* refracted ray of the former pencil, and Vc the *most* refracted of the latter. These, of course, must meet; let them meet in c , and precisely at c apply the vertex of the second prism, having its side ca parallel to CB , but its edge turned in the opposite direction; then will the rays RC and Vc , each for itself, and independent of the other, be refracted so as to emerge parallel to its original direction

Light.

Part II.

S R, S V, and the emergent rays will therefore be coincident and superimposed on each other as *cs*. Thus the emergent ray *cs* will contain an extreme red and an extreme violet ray. But it will also contain every intermediate variety; for draw *cf* anywhere between *cR* and *cV*. Then, since the angle which *cf* makes with the surface *BC* is greater than that made by the extreme violet ray *CB*, but less than that made by the extreme red, there must exist some value of μ intermediate between its extreme values, which will give a deviation equal to the angle between *cf* and S Y parallel to S R. Consequently, if S Y be a white ray, separated into the pencil Y *v' r'* by refraction, the coloured ray Y *fc* of that particular refrangibility will fall on *c*, and be refracted along *cs*. Every point then of the surface *gfh* will send to *c* a ray of different refrangibility, comprehending all the values of μ from the greatest to the least. Thus all the coloured elements, though all belonging originally to different white rays, will, after the second refraction, coincide in the ray *cs*, and experience proves that so reunited they form white light. White light, then, is re-composed when all the coloured elements, even though originally belonging to separate white rays, are united in place and direction.

In the reflexion of light, regarded as a case of refraction, μ has a specific numerical value, and cannot vary without subverting the fundamental law of reflexion. Thus, there is no dispersion into colours produced by reflexion, because all the coloured rays after reflexion pursue one and the same course. There is one exception to this, more apparent than real, when light is reflected from the base of a prism internally, of which more hereafter.

The recomposition of white from coloured light may be otherwise shown, by passing a small circular beam of solar light through a prism A B C, (fig. 89,) and receiving the dispersed beam on a lens E D at some distance. If a white screen be held behind the lens, and removed to a proper distance, the whole spectrum will be reunited in a spot of white light. The way in which this happens will be evident by considering the figure, in which T E and T D represent the parallel pencils of rays of any two colours (red and violet, for instance) into which the incident white beam S T is dispersed. These will be collected after refraction, each in its own proper focus; the former at F, the latter at G; after which each pencil diverges again, the former in the cone F H, the latter in G H. If the screen then be held at H, each of these pencils will paint on it a circle of its own colour, and so of course will all the intermediate ones; but these circles all coinciding, the circle H will contain all the rays of the spectrum confounded together, and it is found (with the exception of a trifling coloured fringe about the edges, arising from a slight overlapping of the several coloured images) to be of a pure whiteness.

That the reunion of *all* the coloured rays is necessary to produce whiteness, may be shown by intercepting a portion of the spectrum before it falls on the lens. Thus, if the violet be intercepted, the white will acquire a tinge of yellow; if the blue and green be successively stopped, this yellow tinge will grow more and more ruddy, and pass through orange to scarlet and blood red. If, on the other hand, the red end of the spectrum be stopped, and more and more of the less refrangible portion thus successively abstracted from the beam, the white will pass first into pale and then to vivid green, blue-green, blue, and finally into violet. If the middle portion of the spectrum be intercepted, the remaining rays, concentrated, produce various shades of purple, crimson, or plum-colour, according to the portion by which it is thus rendered deficient from white light; and by varying the intercepted rays, any variety of colours may be produced; *nor is there any shade of colour in nature which may not thus be exactly imitated, with a brilliancy and richness surpassing that of any artificial colouring.*

Now, if we consider that all these shades are produced on white paper, which receives and reflects to our eyes whatever light happens to fall on it; and that the same paper placed successively in the red, green, and blue portion of the spectrum, will appear indifferently red, or green, or blue, we are naturally enough led to conclude, that

The colours of natural bodies are not qualities inherent in the bodies themselves, by which they immediately affect our sense, but are mere consequences of that peculiar disposition of the particles of each body, by which it is enabled more copiously to reflect the rays of one particular colour, and to transmit, or stifle, or, as it is called in Optics, absorb the others.

Such is the Newtonian doctrine of the origin of colours. Every phenomenon of optics conspires to prove its justice. Perhaps the most direct and satisfactory proof of it is to be found in the simple fact, that every body, indifferently, whatever be its colour in white light, when exposed in the prismatic spectrum, appears of the colour appropriate to that part of the spectrum in which it is placed; but that its tint is incomparably more vivid and full when laid in a ray of a tint analogous to its hue in white light, than in any other. For example, vermilion placed in the red rays appears of the most vivid red; in the orange, orange; in the yellow, yellow, but less bright. In the green rays, it is green; but from the great inaptitude of vermilion to reflect green light, it appears dark and dull; still more so in the blue; and in the indigo and violet it is almost completely black. On the other hand, a piece of dark blue paper, or Prussian blue, in the indigo rays has an extraordinary richness and depth of blue colour. In the green its hue is green, but much less intense; while in the red rays it is almost entirely black. Such are the phenomena of pure and intense colours; but bodies of mixed tints, as pink or yellow paper, or any of the lighter shades of blue, green, brown, &c., when placed in any of the prismatic rays, reflect them in abundance, and appear, for the time, of the colour of the ray in which they are placed.

Refraction by a prism affords us the means of separating a ray of white light into the rays of different refrangibility of which it consists, or of analyzing it. But to make the analysis complete, and to insulate a ray of any particular refrangibility in a state of perfect purity, several precautions are required, the chief of which are as follows: 1st. The beam of light to be analyzed must be very small, as nearly as possible approaching to a mathematical ray; for if A B, *ab* be a beam of parallel rays of any sensible breadth (fig. 89) incident on the prism P, the extreme rays A B, *ab* will each be separated by refraction into spectra G B H and *g b h*: B G, *bg* being the violet, and B H, *bh* the red rays of each respectively; and since A B, *ab* are parallel, therefore C G

407.

408.

Synthesis of white light by a lens.

409.

All the rays necessary to produce white.

All natural colours imitable by combinations of the prismatic.

410.

Colours not inherent in bodies.

411.

Proved by experiment

412.

Precautions to insure the perfect homogeneity of a ray. Fig. 89.

Light.
1st. Small-
ness of the
incident
pencil.

and cg will be so, and also DH and dh . Hence the red ray DH from B will intersect the violet cg from b , in some point F behind the prism; and a screen EFf placed at F will have the point F illuminated by a red ray from B , and a violet one from b ; and therefore (as is easily seen) by all the rays intermediate between the red and violet, from points between B and b . F therefore will be white. If the screen be placed nearer the prism than F , as at $KLkl$, it is clear that from any point between L and k lines drawn parallel to KC , DL , to any intermediate direction, will fall between C and c , D and d , &c., respectively; and therefore that every point between L and k will receive from some point or other of the surface Cd of the prism a ray of each colour, and will therefore be white. Again, any point as x between k and l can receive no violet ray, nor any ray of the spectrum whose angle of deviation is greater than $180^\circ - abx$; for such ray to reach x must come from a part of the prism below b , which is contrary to the supposition of a limited beam AB , ab ; but all rays whose angle of deviation is less than $180^\circ - abx$, will reach x from some part or other of the surface BD . Hence the colour of the portion kl of the image on the screen will be white at k , pure red at l , and intermediate between white and red, or a mixture of the *least* refrangible rays of the spectrum at any intermediate point; and, in the same manner, the portion KL will be white at L , violet at K , and at any intermediate point will have a colour formed by a mixture of a greater or less portion of the more refrangible end of the spectrum. If the screen be removed beyond F , as into the situation $GgHh$, the white portion will disappear, no point between g and H being capable of receiving any ray whose angle of deviation is between $180^\circ - abg$ and $180^\circ - abH$. We may regard the whole image Gh as consisting of an infinite number of spectra formed by every elementary ray of which the beam $ABab$ is composed, overlapping each other, so that the end of each in succession projects beyond that of the foregoing. The fewer, therefore, there are of these overlapping spectra, or the smaller the breadth of the incident beam, the less will be the mixture of rays so arising, and the purer the colours. Removal of the screen to a greater distance from the prism, evidently produces the same effect as diminution of the size of the beam; for while each colour occupies constantly the same space on the screen (for $Gg = Kk$) the whole spectrum is diffused over a larger space as the screen is removed, by the divergence of its component rays of different colours, and therefore the individual colours must of necessity be continually more and more separated from each other.

413.
2nd. Small
angular di-
vergence of
the pencil.
Fig. 90.

2ndly. Another source of confusion and want of perfect homogeneity in the colours of the spectrum is the angular diameter of the sun or other luminary, even when the aperture through which the beam is admitted is ever so much diminished. For let ST (fig. 90) be the sun, whose rays are admitted to the prism ABC through a very small hole O in a screen placed close to it. The beam will be dilated by refraction into the spectrum vr . Now, if we consider only the rays of one particular kind, as the red, and regard all the rest as suppressed, it is clear that a red image r of the sun will be formed by them alone on the screen; the rays from every point of the sun's disc crossing at O , and pursuing (after refraction) different courses. If the prism be placed in its situation of minimum deviation, which at present we will suppose, this image will be a circle, and it and the sun will subtend equal angles at O . In like manner, the violet rays (considered apart from the red) will form a circular violet image of the sun, at v , by reason of their greater refrangibility; and every species of ray, of intermediate refrangibility, will form, in like manner, a circular image between r and v . The constitution of the spectrum so arising will therefore be as in fig. 91, a , being an assemblage of images of every possible refrangibility superposed on and overlapping each other. Now, if we diminish the angular diameter of the sun or luminary, each of these images will be proportionally diminished in size; but their number, and the whole extent over which they are spread, will remain the same. They will therefore overlap less and less, (as in fig. 91, b , c ;) and if the luminary be conceived reduced to a mere point (as a star) the spectrum will consist of a line d composed of an infinite number of mathematical points, each of a perfectly pure homogeneous light.

Fig. 91.

414.
Experimen-
tal methods
of obtaining
homoge-
neous pris-
matic rays.
Fig. 7.

There are several ways by which the angular diameter, or the degree of divergence of the incident beam may be diminished. Thus, first, we may admit a sunbeam through a small hole, as A , in a screen, and receive the divergent cone of rays behind it on another screen B , (fig. 7,) at a considerable distance, having another small hole B to let pass, not the whole, but only a small portion of the sun's image. The beam BC , so transmitted, will manifestly have a degree of divergence less than that of the beam immediately transmitted from A in the proportion of the diameter of the aperture B to the diameter of the sun's image on the screen B .

415.
Fig. 92.

Another and much more commodious method is to substitute for the sun its image formed in the focus of a convex lens of short focus. This image is of very small dimensions, its diameter being equal to focal length of the lens \times sine of sun's angular diameter, (or sine of $30'$, which is about one 114th part of radius,) so that a lens of an inch focus concentrates all the rays which fall on it within a circle of about the 114th of an inch in diameter, which, for this purpose, may be regarded as a physical point. The disposition of the apparatus is as represented in fig. 92. The rays converged by the lens L to F , afterwards diverge as if they emanated from an intensely bright luminous point placed at F , and a screen with a small aperture O being placed at a distance from it, and close behind it the prism ABC , the spectrum rv may be received on a screen again placed at a considerable distance behind the prism, each of whose points will be illuminated by rays of a very high degree of purity and homogeneity, and by diminishing the focal length of the lens, and the aperture O , and increasing the distance FO , or Or , this may be carried to any extent we please. It should, however, be remarked, that the intensity of the purified ray, and the quantity of homogeneous light so obtained, are diminished in the same ratio as the purity of the ray is increased.

416.
Fig. 93.

A third method of obtaining a homogeneous beam is to repeat the process of analysis on a ray as nearly pure as can be conveniently obtained by refraction through a single prism. Thus, in fig. 93, VR , the spectrum formed by a first refraction at the prism A , is received on a screen which intercepts the whole of it, except that particular colour we wish to insulate and purify, which is allowed to pass through an aperture MN ; behind this is placed another prism B , so as to refract this beam a second time. If then the portion

Light.

Part II.

MN were already perfectly pure, it would pass the second prism without undergoing any further separation ; but if there be (as there always will) other rays mixed with it, these will be dilated by the subsequent refraction into a second spectrum *vr* of faint light, with a much brighter portion *mn* in the midst ; and if the rest of the rays be intercepted, and this portion only allowed to pass through an aperture, the emergent beam *mp* will be much more homogeneous than before its incidence on the second prism,—and in proportion as the distance between the second prism and the screen is increased, the purity of the ray obtained will be greater.

Another source of impurity in the prismatic rays is the imperfection of the materials of our ordinary prisms, which are full of striæ and veins, which disperse the light irregularly, and thus confound together in the spectrum rays which properly belong to different parts of it. Those who are not fortunate enough to possess glass prisms free from this defect (which are very rare, and indeed hardly to be procured for any price) may obviate the inconvenience by employing hollow prisms full of water, or, rather, any of the more dispersive oils. A great part of the inconvenience arising from a bad prism may, however, be avoided by transmitting the rays *as near the edge of it as possible*, so as to diminish the quantity of the material they have to pass through, and therefore their chance of encountering veins and striæ in their passage.

417.
Imperfection of prisms, how eluded in practice.

When every care is taken to obtain a pure spectrum ; when the divergence of the incident beam is extremely small, and its dimensions also greatly reduced ; when the prism is perfect, and the spectrum sufficiently elongated to allow of a minute examination of its several parts, some very extraordinary facts have been observed respecting its constitution. They were first noticed by Dr. Wollaston, in a Paper published by him in the *Phil. Trans.*, 1802 ; and have since been examined in full detail, and with every delicacy and refinement which the highest talents and the most unlimited command of instrumental aids could afford, by the admirable and ever-to-be-lamented Fraunhofer. It does not appear that the latter had any knowledge of Dr. Wollaston's previous discovery, so that he has, in this respect, the full merit of an independent inventor. The facts are these : The solar spectrum, in its utmost possible state of purity and tenuity, when received on a white screen, or when viewed by admitting it at once into the eye, is not an uninterrupted line of light, red at one end and violet at the other, and shading away by insensible gradations through every intermediate tint from one to the other, as Newton conceived it to be, and as a cursory view shows it. It is interrupted by intervals absolutely dark ; and in those parts where it is luminous, the intensity of the light is extremely irregular and capricious, and apparently subject to no law, or to one of the utmost complexity. In consequence, if we view a spectrum formed by a narrow line of light parallel to the refracting edge of the prism, (which affords a considerable breadth of spectrum without impairing the purity of the colours, being, in fact, an assemblage of infinitely narrow linear spectra arranged side by side,) instead of a luminous fascia of equable light and graduating colours, it presents the appearance of a striped riband, being crossed in the direction of its breadth by an infinite multitude of dark, and by some totally black bands, distributed irregularly throughout its whole extent. This irregularity, however, is not a consequence of any casual circumstances. The bands are constantly in the same parts of the spectrum, and preserve the same order and relations to each other ; the same proportional breadth and degree of obscurity, whenever and however they are examined, *provided solar light be used*, and provided the prisms employed be composed of the same material : for a difference in the latter particular, though it causes no change in the number, order, or intensity of the bands, or their places in the spectrum, as referred to the several colours of which it consists, yet causes a variation in their proportional distances *inter se*, of which more hereafter. By solar light must be understood, not merely the direct rays of the sun, but any rays which have the sun for their ultimate origin ; the light of the clouds, or sky, for instance ; of the rainbow ; of the moon, or of the planets. All these lights, when analyzed by the prism, are found deficient in the identical rays which are wanting in the solar spectrum ; and the deficiency is marked by the same phenomenon, viz. by the occurrence of the same dark bands in the same situations in spectra formed by these several lights. In the light of the stars, on the other hand, in electric light, and that of flames, though similar bands are observed in their spectra, yet they are differently disposed ; and the spectrum of each several star, and each flame, has a system of bands peculiar to itself, and characteristic of its light, which it preserves unalterably at all times, and under all circumstances.

418.
Fixed lines in the spectrum.

Fig. 94 is a representation of the solar spectrum as laid down minutely by Fraunhofer, from micrometrical measurement, and as formed by a prism of his own incomparable flint glass. Only the great number of small bands observed by him (upwards of 500 in number) have been omitted, to avoid confusing the figure. Of these bands, or, as he terms them, "fixed lines" in the spectrum, he has selected seven, (those marked B, C, D, E, F, G, H,) as terms of comparison, or as standard points of reference in the spectrum, on account of their distinctness, and the facility with which they may be recognised. Of these, B lies in the red portion of the spectrum, near the end ; C is farther advanced in the red ; D lies in the orange, and is a strong double line easily recognised ; E is in the green ; F in the blue ; G in the indigo ; and H in the violet. Besides these, there are others very remarkable ; thus *b* is a triple line in the green, between E and F, consisting of three strong lines, of which two are nearer each other than the third, &c.

419.
Fig. 94.

The definiteness of these lines, and their fixed position, with respect to the colours of the spectrum,—in other words, the precision of the limits of those degrees of refrangibility which belong to the *deficient* rays of solar light,—renders them invaluable in optical inquiries, and enables us to give a precision hitherto unheard of to optical measurements, and to place the determination of the refractive powers of media on the several rays almost on the same footing, with respect to exactness, with astronomical observations. Fraunhofer, in his various essays, has made excellent use of them in this respect, as we shall soon have occasion to see.

420.
Utility of the fixed lines in optical determinations.

To see these phenomena, we must place the refracting angle of a very perfect prism parallel to a very small linear opening through which a sunbeam is admitted ; or, in place of an opening, we may employ a glass cylinder, or semi-cylinder of small radius, to bring the rays to a linear focus behind and parallel to it, from

421

Light.
First method of exhibiting the fixed lines.

which the rays diverge, as from a fine luminous line, in the manner described in Art. 415 for a lens. If now the eye be applied close behind the prism, the line will be seen dilated into a broad coloured band, consisting of the prismatic colours in their order; and if the prism be good, and carefully placed in its situation of minimum deviation, and of sufficiently large refracting angle to give a broad spectrum, some of the more remarkable of the fixed lines will be seen arranged parallel to the edges of the spectrum, especially the lines D and F, the former of which appears, in this way of viewing it, to form a separation between the red and the yellow. If the light of the sun be too bright, so as to dazzle the eye, any narrow line of common daylight (as the slit between two nearly closed window-shutters) may be substituted. This was the mode in which the fixed lines were first discovered by Dr. Wollaston.

422.
Second method.

Fig. 95.

But it is difficult and requires acute sight to perceive, in this manner, any but the most conspicuous lines. The reason is, their very small angular breadth; which, in the largest of them, can scarcely, under any circumstances, exceed half a minute, and in the smaller not more than a few seconds. They require, therefore, to be magnified. This may be done by a telescope interposed between the eye and the prism, in the manner represented in fig. 95, in which Ll is the line of light, from which rays, diverging in all directions, fall on the prism ABC , are refracted by it, and after refraction are received on the object-glass D of the telescope. This object-glass, it should be observed, must be of that kind denominated *achromatic*, to be presently described, (see Index,) and of which it need only be here said, that it is so constructed as to be capable of bringing rays of all colours to foci at one and the same distance from the glass. Now, if we consider only rays of any one degree of refrangibility (the extreme red, for instance) the pencils diverging from every point of Ll will, after refraction at the two surfaces of the prism, diverge from corresponding points of an image $L'l'$ situated in the direction from the base towards the vertex of the prism. Rays of any greater refrangibility will, after refraction at the prism, diverge from a linear image $L''l''$ parallel to $L'l'$, but farther from the original line Ll . Thus the white line Ll will, after refraction at the prism, have for its image the coloured rectangle $L'L''l'l''$, which will be viewed through the telescope as if it were a real object. Now every vertical line of this parallelogram will form in the focus of the object-glass a corresponding vertical image of its own colour; and the object-glass being achromatic, all these images are equidistant from it, so that the whole image of the parallelogram $L'l'$ will be a similar coloured parallelogram, having its plane perpendicular to the axis of the telescope. This will be viewed as a real object through the eye-glass, and the spectrum will thus be magnified as any other object would be, according to the power of the telescope, (Art. 382.) With this disposition of the apparatus (which is that employed by Fraunhofer) the fixed lines are beautifully exhibited, and (if the prism be perfect) may be magnified to any extent. The slightest defect of homogeneity in the prism, however, as may be readily imagined, is fatal. With glass prisms of our manufacture it would be quite useless to attempt the experiment; and those who would repeat it in this country should employ prisms of highly refractive liquids, enclosed in hollow prisms of good plate glass. The eye-pieces of telescopes, not being usually achromatic, a slight change of focus is still required, when the lines in the red and violet portions of the spectrum are to be viewed. This (if an inconvenience) might be obviated by the use of an achromatic eye-piece.

423.
Third method.

Fig. 96.

424.
Colours of the spectrum.

That an actual image of the spectrum, with its fixed lines, is really formed in the focus of the object-glass, as described, may be easily shown, by dismounting the telescope, and receiving the rays refracted by the object-glass on a screen in its focus. This, indeed, affords a peculiarly elegant and satisfactory mode of exhibiting the phenomena to several persons at once. An achromatic object-glass of considerable focal length (6 feet, for instance) should be placed at about twice its focal length from the line of light, and (the prism being placed immediately before the glass) the image will be formed at about the same distance, 12 feet behind it, ($f = L + D$; $L = \frac{1}{2}$; $D = -\frac{1}{2}$; $f = \frac{1}{2} - \frac{1}{2} = +\frac{1}{2}$) and being received on a screen of white paper or emerald glass may be examined at leisure, and the distances of the lines from each other, &c. measured on a scale. But by far the best methods of performing these measurements are those practised by Fraunhofer, viz. the adaptation of a micrometer to the eye-end of the telescope, (see *Micrometer*, in a subsequent part of this Article,) for ascertaining the distances of the closer lines; and the giving the axis of the telescope, together with the prism which is connected with it, a motion of rotation in a horizontal plane, the extent of which is read off by verniers and microscopes on an accurately graduated circle, in the same way as in astronomical observations. The apparatus employed by him for this purpose, and which is applicable to a variety of useful purposes in optical researches, is represented in fig. 96.

The fixed lines in the spectrum do not mark any precise limits between the different colours of which it consists. According to Dr. Wollaston, (*Phil. Trans.*, 1802,) the spectrum consists of only four colours, red, green, blue, and violet; and he considers the narrow line of yellow visible in it in his mode of examination already described (looking through a prism at a narrow line of light with the naked eye) as arising from a mixture of red and green. These colours, too, he conceives to be well defined in the spaces they occupy, not graduating insensibly into each other, and of, sensibly, the same tint throughout their whole extent. We confess we have never been able quite satisfactorily to verify this last observation, and in the experiments of Fraunhofer, (which we had the good fortune to witness, as exhibited by himself at Munich,) where, from the perfect distinctness of the finest lines in the spectrum, all idea of confusion of vision, or intermixture of rays is precluded, the tints are seen to pass into each other by a perfectly insensible gradation; and the same thing may be noticed in the coloured representations of the spectrum published in the first essay of that eminent artist, and executed by himself with extraordinary pains and fidelity. The existence of a pale straw yellow, not of mere linear breadth, but occupying a very sensible space in the spectrum, is there very conspicuous, and may also be satisfactorily shown by other experiments to be hereafter described, when we come to speak of the absorption of light. In short, (with the exception of the fixed lines, which Newton's instrumental means did not enable him to see,) the spectrum is, what that illustrious philosopher originally described it, a graduated succession of tints, in which all

Light.

Part II.

the seven colours he enumerates can be distinctly recognised, but shading so far insensibly into each other that a positive limit between them can be nowhere fixed upon. Whether these colours be really compound or not, whether some other mode of analysis may not effect a separation depending on some other fundamental difference between the rays than that of the degree of their refrangibility, is quite another question, and will be considered more at large hereafter. At present it may be enough to remark, that all probability, drawn from everyday experience, is in favour of this idea, and leads us to believe that orange, green, and violet are mixed colours; and red, yellow, and blue, original ones; the former we everyday see imitated by mixtures of the latter, but never *vice versâ*. This doctrine has been accordingly maintained by Mayer, in a curious Tract published among his works. (See the Catalogue of Optical Writers at the end of this Article.) A very different doctrine has, however, been advanced by Dr. Young, (*Lectures on Natural Philosophy*, i. 441,) in which he assumes red, green, and violet, as the fundamental colours. The respective merits of these systems will be considered more at large hereafter. (See Index, Composition of Colours.)

Media, as we have seen, differ very greatly in their refractive power, or in the degree in which prisms of one and the same refracting angle composed of different substances, deflect the rays of light. This was known to the optical philosophers who preceded Newton. This great man, on establishing the general fact, that one and the same medium refracts differently the differently coloured rays, might naturally have been led to inquire experimentally whether the amount of this difference of action were the same for all media. He appears to have been misled by an accidental circumstance in the conduct of an experiment, in which the varieties of media in this respect ought to have struck him,* and in consequence adopted the mistaken idea of a *proportional action* of all media on the several homogeneous rays. Mr. Hull, a gentleman of Worcestershire, was the first to discover Newton's mistake; and having ascertained the fact, of the different *dispersive powers* of different kinds of glass, applied his discovery successfully to the construction of an achromatic telescope. His invention, however, was unaccountably suffered to fall into oblivion, (though it is said that he made several such telescopes, some of which still exist,) and the fact was re-discovered and re-applied to the same great purpose by Mr. Dollond, a celebrated optician in London, on the occasion of a discussion raised on the subject by some *à priori* and paradoxical opinions broached by Euler.

425.
Media differ in dispersive power.

If a prism of flint glass and one of crown, of equal refracting angles, be presented to two rays of white light, as A B C, *abc*, (fig. 97;) S C and *sc* being the incident rays, C R, C V the red and violet rays refracted by the flint, and *cr*, *cv* those refracted by the crown; it is observed, *first*, that the deviation produced in either the red or violet ray by the flint glass, is much greater than that produced by the crown; *secondly*, that the angle R C V, over which the coloured rays are dispersed by the flint prism, is also much greater than the angle *rcv*, over which they are dispersed by the crown; and, *thirdly*, that the angles R C V, *rcv*, or the *angles of dispersion*, are *not* to each other as Newton supposed them to be, in the same ratio with the *angles of deviation* T C R, *ter*, but in a much higher ratio; the dispersion of the flint prism being much more than in proportion to the deviation produced by it. And if, instead of taking the angles of the prism equal, the refracting angle of the crown prism be so increased as to make the deviation of the red ray equal to that produced by the flint prism, the deviation of the violet will fall considerably short of such equality. In consequence of this, if the two prisms be placed close together, with their edges turned opposite ways, as in fig. 98, so as to oppose each other's action, the red ray, being equally refracted in opposite directions, will suffer no deviation; but the violet ray, being more refracted by the flint than by the crown prism, will, on the whole, be bent towards the thicker part of the flint prism, and thus an uncorrected colour will subsist, though the refraction (for one ray, at least) is corrected. *Vice versâ*, if the dispersion be corrected, that is, if the refracting angle of the crown prism, acting in opposition to the flint, be so further increased as to make the difference of the deviations of the red and violet rays produced by it equal to the difference of their deviations produced by the flint, the deviation produced by it will now be greater than that produced by the flint; and the total deviation, produced by both prisms acting together, will now be in favour of the crown.

426.
Differences of dispersion explained. Fig. 97.

Fig. 98.

By such a combination of two prisms of different media a ray of white light may therefore be turned aside considerably from its course, without being separated into its elementary coloured rays. It is manifest, that (supposing the angles of the prisms small, and that both are placed in their positions of minimum deviation) the deviations to produce this effect must be in the inverse ratio of the dispersive powers of the two media; for supposing μ , μ' to be the refractive indices of the prisms for extreme red rays, and $\mu + \delta\mu$, $\mu' + \delta\mu'$ for extreme violet, A and A' their refracting angles, and D and D' their deviations, we have, generally, in the position of minimum deviation

427.
Refraction without separation into colours.

$$\left. \begin{aligned} \mu \cdot \sin \frac{A}{2} &= \sin \frac{A + D}{2}, \text{ whence } \delta\mu \cdot \sin \frac{A}{2} = \frac{1}{2} \delta D \cdot \cos \frac{A + D}{2} \\ \mu' \cdot \sin \frac{A'}{2} &= \sin \frac{A' + D'}{2} \quad \delta\mu' \cdot \sin \frac{A'}{2} = \frac{1}{2} \delta D' \cdot \cos \frac{A' + D'}{2} \end{aligned} \right\},$$

whence, since the prisms oppose each other,

* He counteracted the refraction of a glass, by a water prism. There ought to have been a residuum of uncorrected colour; but, unluckily, he had mixed sugar of lead with the water to increase its refraction, and the high dispersive power of the salts of lead (of which, of course, he could not have the least suspicion) thus robbed him of one of the greatest discoveries in physical optics.

Light.

$$\frac{1}{2} \delta (D - D') = \frac{\delta \mu \cdot \sin \frac{A}{2}}{\cos \left(\frac{A + D}{2} \right)} - \frac{\delta \mu' \cdot \sin \frac{A'}{2}}{\cos \left(\frac{A' + D'}{2} \right)}.$$

Putting this equal to zero, we have

$$\frac{\delta \mu}{\delta \mu'} \cdot \frac{\sin \frac{1}{2} A}{\sin \frac{1}{2} A'} = \frac{\cos \frac{1}{2} (A + D)}{\cos \frac{1}{2} (A' + D')};$$

and, eliminating $\sin \frac{1}{2} A$ and $\sin \frac{1}{2} A'$ from this, by means of the two original equations from which we set out, we get

$$\frac{\delta \mu}{\delta \mu'} \times \frac{\mu'}{\mu} = \frac{\cos \frac{1}{2} (A + D)}{\cos \frac{1}{2} (A' + D')} \times \frac{\sin \frac{1}{2} (A' + D')}{\sin \frac{1}{2} (A + D)} = \frac{\tan \frac{1}{2} (A' + D')}{\tan \frac{1}{2} (A + D)}.$$

Now if we call p, p' the dispersive powers of the media, or the proportional parts of the whole refractions of the extreme red ray, to which the dispersion is equal, we shall have

$$p = \frac{\delta \mu}{\mu - 1} \quad p' = \frac{\delta \mu'}{\mu' - 1} \quad \text{and} \quad \frac{p}{p'} = \frac{\delta \mu}{\delta \mu'} \times \frac{\mu' - 1}{\mu - 1},$$

so that

$$\frac{p}{p'} = \frac{\mu}{\mu'} \cdot \frac{\mu' - 1}{\mu - 1} \cdot \frac{\tan \frac{1}{2} (A' + D')}{\tan \frac{1}{2} (A + D)} = \frac{\mu' - 1}{\mu - 1} \cdot \frac{\sin \frac{1}{2} A'}{\sin \frac{1}{2} A} \cdot \sqrt{\frac{1 - \mu^2 \cdot (\sin \frac{1}{2} A)^2}{1 - \mu'^2 \cdot (\sin \frac{1}{2} A')^2}}. \quad (a)$$

Such is the strict formula, which, when A and A' are very small, becomes

$$\frac{p}{p'} = \frac{(\mu' - 1) A'}{(\mu - 1) A}; \quad \text{or, since } (\mu - 1) A = D, \text{ and } (\mu' - 1) A' = D'; \quad \frac{p}{p'} = \frac{D'}{D}.$$

428.
Dispersive
powers com-
pared by
experiment.

The formula just obtained, furnishes us with an experimental method of determining the ratio of the dispersive powers of two media. For if we can by any means succeed in forming them into two prisms of such refracting angles, that, when placed in their respective positions of minimum deviation, a well defined bright object, viewed through both, shall appear well defined and free from colour at its edges; then, by measuring their angles, and knowing also from other experiments their refractive indices, the equation (a) gives us immediately the ratio in question.

429.
Coloured
fringes bor-
dering ob-
jects seen
through
prisms ex-
plained.
Fig. 99.

When we view through a prism any well defined object, either much darker or much lighter than the ground against which it is seen projected, as, for instance, a window bar seen against the sky, its edges appear fringed with colours and ill defined. The reason of this may be explained as follows:

Let $A B$, fig. 99, be the section of a horizontal bar seen through the prism P held with its refracting edge downwards, and first let us consider what will be the appearance of the upper edge B of the object. Since we see by light, and not by darkness, the thing really seen is not the dark object, but the bright ground on which it stands, or the bright spaces $B C, A D$ above and below. Now the bright space $B C$ above the object being illuminated with white light, will, after refraction at the prism, form a succession of coloured images $b c, b' c', b'' c'', \&c.$, superposed on and overlapping each other. They are represented in the figure as at different distances from P , but this is only to keep them distinct. In reality, they must be supposed to lie upon and interfere with each other. The least refracted $b c$ of these is red, and the most refracted $b'' c''$ violet, and any intermediate one (as $b' c'$) of some intermediate colour, as yellow for instance. Beyond b'' no image exists, so that the whole space below b'' will appear dark to an eye situated behind the prism. On the other hand, above b the images of every colour in the spectrum coexist, the bright space $b c$ being supposed to extend indefinitely above B . Therefore the space above b in the refracted image will appear perfectly white. Between b and b'' there will be seen, first, a general diminution of light, as we proceed from b towards b'' , because the number of superposed luminous images continually decreases; secondly, an excess in all this part, of the more refrangible rays in the spectrum above what is necessary to form white light, for beyond b no red image exists, beyond b' no yellow, and so on; the last which projects beyond all, at b'' , being a pure unmixed violet. Thus the light will not only decrease in intensity, but by the successive subtraction of more and more of the less refrangible end of the spectrum will acquire a bluer and bluer tint, deepening to a pure violet, so that the upper edge of the dark object will appear fringed with a blue border, becoming paler and paler till it dies away into whiteness. The reverse will happen at the lower edge A . The bright space $A D$ forms, in like manner, a succession of coloured images, $a d, a' d', a'' d'', \&c.$, of which the least deviated $a d$ is red, the most $a'' d''$ violet, and the intermediate ones of the intermediate colours. Therefore the point a , which contains only the extreme red, will appear of a sombre red; a' , which contains all the rays from red to yellow (suppose), of a lively orange red; and in proportion as the other images belonging to the more refrangible end of the spectrum come in, this tendency to an excess of red will be neutralized, and the portion beyond a'' , containing all the colours in their natural proportions, will be purely white. Hence, the lower edge of the dark object will appear bordered with a red fringe, whose tint fades away into whiteness, in the same way as the blue fringe which borders the upper edge. These fringes, of course, destroy the distinctness of the outlines of objects, and render vision through a prism confused. The confusion ceases, and objects resume their natural well defined outlines, if illuminated with homogeneous light, or if viewed through coloured glasses which transmit only homogeneous rays.

Light.

The eye can judge pretty well, by practice, of the destruction of colour, and indistinctness in the edges of objects, when prisms are made to act in opposition to one another, as above described; but (owing to causes presently to be considered) the compensation is never perfect, and there always remains a small fringe of uncorrected purple on one side, and green on the other, when the eye is best satisfied; so that observations of dispersive powers by this method are liable to a certain extent of error, and, indeed, precision in this department of optical science is very difficult to obtain.

To determine the dispersive power of a medium, having formed it into a prism, and measured by the goniometer, or otherwise, its refracting angle, and ascertained its refractive index, the next step is to find the refracting angle of a prism of some standard medium, which shall exactly compensate its dispersion, so as to produce a refraction as nearly as possible free from colour. But as it is impossible to have a series of standard prisms with every refracting angle which may be requisite, it becomes necessary to devise some means of varying the refracting angle of one and the same prism by insensible gradations. Many contrivances may be had recourse to for this. Thus, first, we may use a prism composed of two plates of parallel glass, united by a hinge, or otherwise, and enclosing between them a fluid, which may be prevented from escaping either by capillary attraction, if in very small quantity, or by close-fitting metallic cheeks, forming a wedge-shaped vessel, if in larger. This contrivance, however, is liable to a thousand inconveniences in practice. Secondly, we may use two prisms of the same kind of glass, one of which has one of its faces ground into a convex, and the other into a concave cylinder, of equal curvatures, having their axes parallel to the refracting edges. These being applied to each other, and one of them being made to revolve round the common axes of the two cylindric surfaces upon the other, the plane faces will evidently be inclined to each other in every possible angle within the limits of the motion, (see fig. 100, *a*, *b*, exhibiting two varieties of this construction.) The idea, due, we believe, to Boscovich, is ingenious, but the execution difficult, and liable to great inaccuracies.

The following method succeeds perfectly well, and we have found it very convenient in practice. Take a prism of good flint glass, whose section is a right angled triangle, *A B C*, having the angle *A* about 30° or 35°, *C* being the right angle, and whose length is twice the breadth of the side *A C*; and, having ground and polished the side *A C*, and the hypotenuse of the prism to true planes, cut it in half, so as to form two equal prisms with one face in each a square, and whose refracting angles (*A*, *A'*) cannot, of course, be otherwise than exactly equal. Cement the square faces together very carefully with mastic, so that the edges *A*, *A'*, shall be on opposite sides of the square surface, which is common to both; and then, making the whole solid to revolve round an axis perpendicular to the common surface, and passing through its centre, grind off all the angles of the squares in the lathe, and the whole will be formed into a cylindrical solid, with oblique, parallel, elliptical, plane ends, as in fig. 101. Then separate the prisms, (by warming the cement,) and set each of them in a separate brass mounting, as in fig. 102, so as to have their circular faces in contact, and capable of revolving freely upon each other about their common centre. The lower one is fixed in the centre of the divided circle *D E*, while the mounting of the upper or moveable one carries an arm with an adjustable vernier reading off to tenths of degrees, or, if necessary, to minutes. The whole apparatus is set in a swing frame between plates, which grasp the divided plate by a groove in its edge, allowing a motion in its own plane, and a capability of adjusting it to any required position, so as to admit of the compound prism deviating an incident ray in every possible plane, and under every possible situation, with respect to the faces of the prisms. It is evident, that in the position here represented, where the prisms oppose each other, (and at which the vernier must be set to read off zero,) the refracting angle is rigorously nothing; and when turned round 180°, since the prisms then conspire, their combined angle must be double that of each. In intermediate situations, the angle between the planes of their exterior faces must, of course, pass through every intermediate state, and (by spherical trigonometry) it is readily shown, that if θ be the reading off of the vernier, or the angle of rotation of the prisms on each other from the true zero, the angle of the compound prism will be had by the equation

$$\sin \frac{A}{2} = \sin \frac{\theta}{2} \cdot \sin (A) \quad (b)$$

where (*A*) is the refracting angle of each of the simple prisms, and *A* the angle of the compound one.

To use this instrument, place the prism *A'*, whose dispersive power is to be compared with the medium of which the standard prism (*A*) is formed, with its edge downwards and horizontal, before a window, and, selecting one of the horizontal bars properly situated, fix it so that the refraction of this bar shall be a minimum, or till, on slightly inclining the prism backwards and forwards, the image of the bar appears stationary. Then take the standard compound prism, adjust it to zero, and set it vertically on its frame behind the first prism. Move its index a few degrees from zero, and turn the divided circle in its own plane, till the refraction so produced by the second prism is contrary to that produced by the first. The colour will be found less than before, continue this till the colour is nearly compensated, then, by means of the swing motion, and of the motion round the vertical axis, adjust the apparatus so that two of the window bars, a horizontal and a vertical one, seen through both prisms, shall appear to make a right angle with each other, (an adjustment, at first, rather puzzling, but which a little practice renders very easy.) Then complete the compensation of the colour; verify the position of the standard prism, (by the same test,) and finally read off the vernier, and the required angle *A* of the compensating prism is easily calculated by the equation (*b*). This calculation may be saved by tabulating the values of *A* corresponding to those of θ , (the value of (*A*) being supposed known by previous exact measures,) or, by graduating the divided circle at once, not into equal parts of θ , but according to such computed values of *A*, so as to read off at once the value of the angle required.

Part II.

430.

431.

To determine the dispersion of a medium

Prisms with variable refracting angles described.

Another construction Fig. 100.

432.

Third construction. Fig. 101, 102.

433.

How used.

Light.

431.

Another method proposed by Dr. Brewster

A simpler, perhaps, on the whole, a better, method of comparing the dispersions of two prisms, is one proposed and applied extensively by Dr. Brewster, in his ingenious *Treatise On New Philosophical Instruments*, a work abounding with curious contrivances and happy adaptations. It consists in varying, not the refracting angle of the standard prism, but the direction in which its dispersion is performed. It is manifest, that if we can produce from a line of white light, by means of a standard prism any how disposed, a coloured fringe, in which the colours occupy the same angular breadth as in that produced by a prism of unknown dispersion; then, the latter, being made to refract this fringe in a direction perpendicular to its breadth, and opposite to the order of its colours, must destroy all colour and produce a compensated refraction; and therefore if the position of the standard prism which produces such a fringe be known, the dispersion of the other may be calculated. To accomplish this, let AB be a horizontal luminous line of considerable length, and let it be refracted downwards, but obliquely in the direction Aa, Bb , by a standard prism whose dispersion is greater than that of the prism to be measured. Then it will form an oblique spectrum $ab b'a'$, ab being the red, and $a'b'$ the violet; and the angular breadth of this coloured fringe will be $am = a'd' \times \sin$ inclination of the plane of refraction to the horizon. Now, let the prism whose dispersion is to be measured be made to refract this coloured band vertically upwards; then, if the plane of the first refraction be so inclined to the horizon that the angle subtended by am at the eye shall be just equal to the angle of dispersion of the other prism, all the colours of the rectangular portion $bca'd$ will be made to coalesce in the horizontal line $A'B'$, which will appear therefore free from colour, except at its extremities $A'B'$, where the coloured triangles aca' , bdb' will produce a red termination $A'A''$ and a blue one $B'B''$ at the respective ends of the line to which they correspond. Hence, if, the second prism remaining fixed, with its edge downwards and parallel to the horizon, the other or standard prism be turned gradually round in the plane perpendicular to its principal section, a position must necessarily be found where the twice refracted line $A'B'$ will appear free from colour both above and below. In this position let it be arrested, and the angle of inclination of its edge to the horizon read off; its complement is the angle $a'd'm$, which we will call θ . Let us now suppose each prism adjusted to its position of minimum deviation, and (as it is a matter of indifference which is placed first) let the prism to be examined or the fixed prism be placed next the object.* Then, D' and D being the total deviations produced by the fixed and revolving prisms on the extreme red ray, we must have

$$\delta D' - \delta D \cdot \sin \theta = 0; \quad \text{or } \delta \mu' \cdot \sin \frac{A'}{2} \cdot \sec \frac{A' + D'}{2} = \delta \mu \cdot \sin \frac{A}{2} \cdot \sec \frac{A + D}{2} \cdot \sin \theta,$$

whence we obtain

$$\frac{p'}{p} = \frac{\delta \mu'}{\delta \mu} \cdot \frac{\mu - 1}{\mu' - 1} = \frac{\mu'}{\mu} \cdot \frac{\mu - 1}{\mu' - 1} \cdot \frac{\tan \frac{1}{2}(A + D)}{\tan \frac{1}{2}(A' + D')} \cdot \sin \theta; \quad (c)$$

where the angles $\frac{1}{2}(A + D)$ and $\frac{1}{2}(A' + D')$ are given by the equations

$$\sin \frac{1}{2}(A + D) = \mu \cdot \sin \times \frac{1}{2}A; \quad \sin \frac{1}{2}(A' + D') = \mu' \cdot \sin \frac{1}{2}A';$$

from which formula, θ being known, and also the angles and refractive indices of the two prisms, the ratio of their dispersions is found.

435.

Absolute dispersive powers, how obtained. 1st. By measuring the spectrum on a screen.

By these, or other similar methods, may the dispersions of any media be compared with those of any other taken as a standard. If the media be solid, they must be formed into prisms; if fluid, they must be enclosed in hollow prisms of truly parallel plates of glass, whose angles must be accurately determined, (and one of which will serve for any number of fluids.) But to ascertain directly the dispersion of that standard prism, we must pursue a different course. The first method which obviously presents itself, is to measure the actual length of the solar spectrum cast by a prism of given refracting angle; but the light of the spectrum dies away so indefinitely at both ends, and its visible extent varies so enormously with the brightness of the sun, and the more or less perfect exclusion of extraneous light, that nothing certain can be concluded from such measures. Yet, if the brighter rays of the spectrum be destroyed, and the eye defended from all offensive light by a glass which permits only the extreme red and violet rays to pass, (see *Index, Absorption*), some degree of accuracy may be obtained by this means. A method founded on this principle has been described by the writer of these pages in the *Transactions of the Royal Society of Edinburgh*, vol. ix. as follows: Let A and B be two vertical rectangular slits in a screen placed before an open window, the one being half the length of the other, and at a known distance from each other. The eye being guarded as above described, let the slits be refracted by the prism (in its minimum position) from the longer towards the shorter. Then will a red and violet image of each a, b , and a', b' be seen. Now let the prism be removed from the slits, (or *vice versa*), still preserving its position of minimum deviation, till the violet image of the longer slit exactly falls upon and covers the red image of the shorter, as in the position $a'b$ of the figure. Then it is obvious, that the distance between the slits, divided by their distance from the prism, is the sine of the total angle of dispersion, or is equal to δD , and this being known

$$\delta \mu = \frac{\delta D}{2} \frac{\cos \frac{1}{2}(A + D)}{\sin \frac{1}{2}A},$$

and therefore $\frac{\delta \mu}{\mu - 1}$, or p , the dispersive power, is obtained.

* Dr. Brewster has chosen a somewhat different position, (*Treatise*, &c. p. 296,) with a view to simplify the formulæ; but it does not appear to us that any advantage is gained in that respect by his arrangement.

Fig. 104. 2nd. Another method.

Light.

But all these methods are only rude approximations, as the great discrepancies of the results hitherto obtained by them abundantly prove; thus, the dispersions of various specimens of flint glass, obtained by the method last described, come out no less than one-sixth larger than those previously given by Dr. Brewster. The only method which can really be relied on is that practised by Fraunhofer, (where the media can be procured in a state of sufficient purity and quantity for its application;) and consists in determining, with astronomical precision, by direct measures, the values of μ for the several points of definite refrangibility in the spectrum, marked, either by the fixed lines, or by the phenomena of coloured flames or absorbent media. (See Index, *Flames—Absorption.*) By taking advantage of the properties of the latter, a red ray, of a refrangibility strictly definite, may be insulated with great facility; and as it lies so near the extremity of the spectrum as not to be perceptible till all the brighter rays are extinguished, it is invaluable as a fixed term in optical researches, and will always be understood by us in future, when speaking of the *commencement of the spectrum*, or the *extreme red*, even though a red ray still less refrangible should be capable of being discerned by careful management, and in favourable circumstances. In like manner, by the simple artifice of putting a little salt into a flame, a yellow ray of a character perfectly definite is obtained, which, it is very remarkable, occupies precisely the place in the scale of refrangibility where in the solar spectrum the dark line D occurs, (Art. 418, 419.) These, and the fixed lines there mentioned, leave us at no loss for rays identifiable at all times and in all circumstances, (with a good apparatus,) and enable us to place the doctrine of refractive and dispersive powers on the footing of the most accurate branches of science.

Part II.

436.
Method
employed by
Fraunhofer.

Use of the
fixed lines.

The following table, extracted from Fraunhofer's *Essay on the Determination of Refractive and Dispersive Powers*, &c. contains the absolute values of the index of refraction μ for the several rays whose places in the spectrum correspond to the seven lines B, C, D, E, F, G, H, assumed by him as standards (see Art. 419, &c.) for several different specimens of glass of his own manufacture, and for certain liquids. These values, for distinction's sake, we may designate by the signs μ (B), μ (C), μ (D), &c.

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Table of the refractive indices of various glasses and liquids for seven standard rays.

Refracting medium.	Specific gravity.	Values of						
		μ (B)	μ (C)	μ (D)	μ (E)	μ (F)	μ (G)	μ (H)
Flint glass, No. 13.	3.723	1.627749	1.629681	1.635036	1.642024	1.648266	1.660285	1.671062
Crown glass, No. 9.	2.535	1.525832	1.526849	1.529587	1.533005	1.536052	1.541657	1.546566
Water	1.000	1.330935	1.331712	1.333577	1.335851	1.337818	1.341293	1.344177
Water, another experiment	1.000	1.330977	1.331709	1.333577	1.335849	1.337788	1.341261	1.344162
Solution of potash	1.416	1.399629	1.400515	1.402805	1.405632	1.408082	1.412579	1.416368
Oil of turpentine.	0.885	1.470496	1.471530	1.474434	1.478353	1.481736	1.488198	1.493874
Flint glass, No. 3	3.512	1.602042	1.603800	1.608494	1.614532	1.620042	1.630772	1.640373
Flint glass, No. 30	3.695	1.623570	1.625477	1.630585	1.637356	1.643466	1.655406	1.666072
Crown glass, No. 13	2.535	1.524312	1.525299	1.527982	1.531372	1.534337	1.539908	1.544684
Crown glass, letter M.	2.756	1.554774	1.555933	1.559075	1.563150	1.566741	1.573535	1.579470
Flint glass, No. 23 Prism of 60° 15' 42" }	3.724	1.626596	1.628469	1.633667	1.640495	1.646756	1.658848	1.669686
Flint glass, No. 23 Prism of 45° 23' 14" }	3.724	1.626564	1.628451	1.633666	1.640544	1.646780	1.658849	1.669680

The above table renders very evident a circumstance which has long been recognised by experimental opticians, and which is of great importance in the construction of telescopes, viz. the *irrationality*, (as it has been termed,) or want of proportionality of the spaces occupied in spectra formed by different media by the several coloured rays, or by those whose refrangibilities, by any one standard medium, lie between given limits. If we fix upon water, for example, as a standard medium, (and we see no reason why it should not be generally adopted as a term of reference in this, as in other physical inquiries—of course at a given temperature—that of its maximum density, for instance,) it is obvious, that any ray may be identified by stating its index of refrangibility by water; thus, a scale of refrangibilities, which, for brevity, we shall term *the water scale*, is established; and so soon as we know the refractive index of a ray from vacuum into water, we have its place in the water spectrum, its colour,

438.

Identification of a ray by its place in a water spectrum.

Light.

Part II.

and its other physical properties (so far as they depend on the refrangibility of the ray) determined. Thus 1.333577 being known to be the refractive index for a ray in water, that ray can be no other than the particular ray D, whose colour is pale orange-yellow, and which is totally deficient in solar light, and peculiarly abundant in the light of certain flames. Now let x be the refractive index of any ray whatever for water, or its place in the water scale. Then it is evident, that its refractive index for any other medium must of necessity be a function of x , because the value of x determines this and all the other properties of the ray. Hence we must have between μ and x some equation which may be generally represented by $\mu = F(x)$; $F(x)$ denoting a function of x .

439.
Function of
refrangibility.

To determine the form of this function, we must consider, that if A be the very small angle of a prism, and D the deviation produced by it at the minimum, we have $\mu \cdot \frac{A}{2} = \frac{A + D}{2}$, or $D = (\mu - 1) A$. Hence,

supposing A the refracting angle constant, the deviation is proportional to $\mu - 1$. Now, since in all media, as well as in water, the deviations observe, at least, the same order, being always least for the red and greatest for the violet, it follows, that in all media $\mu - 1$ increases as x increases; so that, supposing x_0 to be the index of refraction in the water scale for the first visible red ray, or the commencing value of x , and μ_0 the index for the same ray in the other medium, $(\mu - 1) - (\mu_0 - 1)$, or $\mu - \mu_0$ must increase with $x - x_0$; and since they vanish together, we may represent the one in a series with indeterminate coefficients, and powers of the other, thus

$$\mu - \mu_0 = A(x - x_0) + B(x - x_0)^2 + C(x - x_0)^3 + \&c.;$$

or, which comes to the same thing, $a, b, c, \&c.$, representing other indeterminate coefficients, ($x_0 - 1$ being constant,)

$$\frac{\mu - \mu_0}{\mu_0 - 1} = a \cdot \left(\frac{x - x_0}{x_0 - 1} \right) + b \cdot \left(\frac{x - x_0}{x_0 - 1} \right)^2 + \&c. \quad (d)$$

440.
Hypothesis
of constant
dispersion
in all media.

The simplest hypothesis we can form respecting the values of $a, b, \&c.$ is that which makes $a = 1$, and b , and all the other coefficients vanish. This gives $\frac{\mu - \mu_0}{\mu_0 - 1} = \frac{x - x_0}{x_0 - 1}$.

We have before used $\delta\mu$ to denote what is here signified by $\mu - \mu_0$, viz. the difference between the refractive indices of any ray in the spectrum, and that at its commencement; and we have denoted by $\frac{\delta\mu}{\mu - 1}$ the same quantity which is here expressed by $\frac{\mu - \mu_0}{\mu_0 - 1}$. This then is the expression, in our present notation, of the

Not the law
of nature.

dispersive power of the medium; and the equation now under consideration therefore indicates, that, on the hypothesis made, the dispersive power of the medium must necessarily be the same with that of water; and of course (supposing this hypothesis to be founded in the nature of light) all media must have the same dispersive power. This, as we have already seen, is not the case.

Nor that of
proportional
dispersion.

The next simplest hypothesis is that which admits a as an arbitrary constant determined by the nature of the medium, but still makes $b, c, \&c. = 0$. This reduces the equation to

$$\frac{\mu - \mu_0}{\mu_0 - 1} = a \cdot \frac{x - x_0}{x_0 - 1};$$

consequently (if μ' and x' be any other corresponding values of μ and x) we must have also

$$\frac{\mu' - \mu_0}{\mu_0 - 1} = a \cdot \frac{x' - x_0}{x_0 - 1}, \text{ and therefore } \frac{\mu' - \mu}{\mu_0 - 1} = a \cdot \frac{x' - x}{x_0 - 1}; \text{ whence we have } \frac{\mu' - \mu}{x' - x} = a \cdot \frac{\mu_0 - 1}{x_0 - 1}.$$

Hence, if this hypothesis be correct, and μ, x and μ', x' be any two pairs of corresponding refractive indices

for rays however situated, the fraction $\frac{\mu' - \mu}{x' - x}$ must be invariable. The foregoing table, however, shows very

distinctly that this is far from being the case. Thus, if we take the flint glass, No. 13, the comparison of the two rays B and C gives for the value of the fraction in question 2.562; and if we compare in like manner the rays C and D, D and E, E and F, F and G, G and H respectively, we obtain the values 2.871, 3.073, 3.193, 3.460, 3.726; the great deviation of which from equality, and their regular progression, leaves no doubt of the incompatibility of the hypothesis in question, as a general law, with nature. If we institute the same comparison for the other media in the table, we shall find the greatest diversity prevail; and if, instead of water, we assume any other as a standard, the same incompatibility will be found. Thus if the flint glass, No. 13, be compared with oil of turpentine, we find for the values of the series of fractions in question, 1.868, 1.844, 1.783, 1.843, 1.861, 1.899, which first diminish to a minimum and then increase again, &c.

441

It follows from this, that the proportion which the several coloured spaces (or the intervals B C, C D, D E, &c.) bear to each other in spectra formed by different media, is not the same in all. Thus taking the green ray E for the middle colour, and calling all that part of the spectrum which lies on the red side of E the red,

Light.

and all on the other side the *blue* portions, the ratio of the spaces occupied by the red and blue in any spectrum will be represented by the fraction $\frac{\mu(H) - \mu(E)}{\mu(E) - \mu(B)}$. Now the values of this in the several media of the foregoing table are set down in the following list:

Part II.

Incommensurability of the coloured spaces in spectra of different media.

Flint, No. 23.	2.0922	Crown, M.	1.9484
Flint, No. 30.	2.0830	Crown, No. 9	1.8905
Flint, No. 3	2.0689	Crown, No. 13. ...	1.8855
Flint, No. 13.	2.0342	Solution of potash.	1.7884
Oil of turpentine ..	1.9754	Water.	1.6936

Here we see that the same coloured spaces which in the flint No. 23 are in the ratio of 21 : 10, in the water spectrum are only in the ratio of 17 : 10 (nearly,) so that the blue portion of the spectrum is considerably more extended in proportion to the red in the flint glass than in the water spectrum.

From this it follows, that if two prisms be formed of different media (such as flint glass and water) of such refracting angles as to give spectra of equal total lengths, and these be made to refract in opposition to each other, although the red and violet rays will, of course, be united in the emergent beam, yet the intermediate rays will still be somewhat dispersed, the water prism refracting the green, or middle rays more than in proportion to the extremes; consequently, if a white *luminous* line be the object examined through such a combination, instead of being seen after refraction colourless, it will form a coloured spectrum of small breadth compared with what either prism separately would form, and having one side of a purple and the other of a green tint. Any dark object viewed against the sky (as a window bar) will be seen fringed with purple and green borders, the green lying on the same side of the bar with the vertex of the flint prism; because in such a combination, green must be considered as the most, and purple as the least, refrangible tint; and the flint prism, of necessity, having the least refraction in this case, the *most refrangible* fringe will lie towards its vertex, that being the *least refracted* side of the bar; for the same reason that, when seen through a single prism, a dark object on a white ground appears fringed with blue on its least refracted edge. (Art. 429.)

442.
Secondary spectra.

This result accords perfectly with observation. Clairaut, and, after him, Boscovich, Dr. Blair, and Dr. Brewster, have severally drawn the attention of opticians to these coloured fringes, or, as they may be termed, *secondary spectra*, and demonstrated their existence in the most satisfactory manner. Dr. Brewster, in particular, has entered into a very extensive and highly valuable series of experiments, described in his Treatise on new philosophical instruments, and in his paper on the subject in the *Edinburgh Transactions*; from which it follows, that when a compound prism, consisting of any of the media in the following list refracting in opposition to each other, unites the red and violet rays, the green will be deviated from their united course by the combination, in the direction of the refraction of that medium which stands before the other in order:

443.

1. SULPHURIC ACID.	31. Gum juniper.	61. Oil of nutmegs.
2. Phosphoric acid.	32. Rock salt.	62. ——— carraway.
3. Sulphurous acid.	33. Calcareous spar.	63. ——— lemon.
4. Phosphorous acid.	34. Oil of ambergris.	64. Amber.
5. Super-sulphuretted hydrogen.	35. ——— juniper.	65. Oil of spearmint.
6. WATER.	36. ——— spermaceti.	66. ——— hyssop.
7. Ice.	37. ——— rape seed.	67. ——— poppy.
8. White of egg.	38. ——— olives.	68. ——— penny-royal.
9. Rock crystal.	39. Zircon.	69. ——— sage.
10. Nitric acid.	40. FLINT GLASS.	70. ——— turpentine.
11. Prussic acid.	41. Oil of rhodium.	71. Canada balsam.
12. Muriatic acid.	42. ——— rosemary.	72. Oil of lavender.
13. Nitrous acid.	43. ——— fenugreek.	73. Muriate of antimony.
14. Acetic acid.	44. Balsam of capivi.	74. Oil of cloves.
15. Malic acid.	45. Nut oil.	75. ——— sweet fennel seeds.
16. Citric acid.	46. Oil of savine.	76. Red-coloured glass.
17. Fluor spar.	47. ——— rue.	77. Orange-coloured glass.
18. Topaz, (blue.)	48. ——— beech-nut.	78. Opal-coloured glass.
19. Beryl.	49. Nitrate of potash.	79. Acetate of lead, (melted.)
20. Selenite.	50. Diamond.	80. Oil of amber.
21. Leucite.	51. Resin.	81. ——— sassafras.
22. Tourmaline.	52. Gum copal.	82. ——— cummin.
23. Borax.	53. Castor oil.	83. ——— anise seeds.
24. Borax, (glass of.)	54. Oil of chamomyle.	84. Essential oil of bitter almonds.
25. Ether.	55. ——— dill-seeds.	85. Carbonate of lead.
26. Alcohol.	56. ——— wormwood.	86. Balsam of Tolu.
27. Gum Arabic.	57. ——— marjoram.	87. Sulphuret of carbon.
28. CROWN GLASS.	58. ——— bergamot.	88. Sulphur.
29. Oil of almonds.	59. ——— peppermint.	89. Oil of cassia.
30. Tartrate of potash and soda.	60. ——— thyme.	

Dr. Brewster's table of media according to action on green light.

Light.

It is evident from this table, that (generally speaking) the more refractive a medium is, the greater is the extent of the blue portion of its spectrum compared with the red.

444.

If two prisms of the proper refracting angles, composed of media not very remote from each other in this list, be made to oppose each other, the secondary spectrum will be small, and the refraction almost perfectly colourless. Such a combination is said to be *achromatic*, ($a - \chi \rho \mu \mu a$.)

Achromatic refraction.

446.

The existence of the secondary spectrum, while it renders the attainment of perfect achromaticity impossible, by the use of two media only, shows, also, that in a theoretical point of view we are not entitled to neglect the coefficients b , c , &c. of the equation (d), Art. 439. The law of nature probably requires the series to be continued to infinity; and if, by way of uniting three rays, we employ prisms of three media, *tertiary spectra*, and after them still others in succession, would doubtless be found to arise. These, however, will be small in comparison of each other.

Dispersive powers of higher orders. Tertiary spectra.

447.

Computation of their coefficients.

The table (Art. 437) gives us the means of computing the coefficients on which they depend for the particular media there stated. If we put $\frac{\mu - \mu_0}{\mu_0 - 1} = P$, and $\frac{x - x_0}{x_0 - 1} = p$, and suppose P , P' , P'' ,

p , p' , p'' , &c. to be the values of P and p corresponding to the several values of μ and x set down in the table, we shall have, for determining a , b , c , &c. in any one of those media, the equations

$$P = ap + bp^2 + cp^3 + \&c. \quad P' = ap' + bp'^2 + cp'^3 + \&c. \quad P'' = ap'' + bp''^2 + cp''^3 + \&c.$$

and as many such equations must be used as there are coefficients to determine. Confining ourselves at present to two, we find $P = ap + bp^2$; $P' = ap' + bp'^2$, whence

$$a = \frac{Pp'^2 - P'p^2}{pp'(p' - p)}; \quad b = -\frac{Pp' - P'p}{pp'(p' - p)};$$

and, since it is desirable to select rays as far removed from each other in the spectrum as possible, we shall take μ_0 and x_0 from the column μ (B); and determine P and p by the values in the column μ (E), and P' , p' by those under μ (H). The results will be as follows:

Refracting media.	Dispersive powers of the first order, that of water being 1.000.	Dispersive powers of the second order, that of water being 0.000.
Flint glass, No. 13 ..	$a = + 1.42580$	$b = + 7.57705$
Crown glass, No. 9 ..	0.88419	2.34915
Water.....	1.00000	0.00000
Solution of potash ..	0.99626	1.13262
Oil of turpentine	1.06149	4.58639
Flint glass, No. 3....	1.29013	7.63048
Flint glass, No. 30 ..	1.37026	8.44095
Crown glass, No. 13..	0.87374	2.49199
Crown glass, letter M.	0.90131	3.49000
Flint glass, No. 23 ..	1.37578	8.66904

448.

General conditions of achromaticity.

Problem. To determine the analytical relation which must hold good in order that two prisms may form an *achromatic* combination; that is, may refract a white ray without separating the extreme colours.

Resuming the equations and notation of Art. 215, since the prisms are placed in vacuo, we have to substitute μ , $\frac{1}{\mu'}$, μ' and $\frac{1}{\mu''}$ for μ , μ' , μ'' , μ''' , in those equations respectively, and we shall have

$$\left. \begin{aligned} \mu \cdot \sin \rho &= \sin a \\ \alpha' &= I + \rho \\ \sin \rho' &= \mu \cdot \sin \alpha' \end{aligned} \right\} (1); \quad \left. \begin{aligned} \mu' \cdot \sin \alpha''' &= \sin \rho''' \\ \rho'' &= -I'' + \alpha''' \\ \sin \alpha'' &= \mu' \cdot \sin \rho'' \end{aligned} \right\} (2);$$

and

$$a'' = I' + \rho'; \quad D = a + I + I' + I'' - \rho''.$$

Now, since by hypothesis the incident and emergent rays are both colourless, we must have $\delta a = 0$, and $\delta D = 0$, that is $\delta \rho''' = 0$, the sign δ being supposed to refer to the variation of the place of the ray in the spectrum. Hence the two systems of equations (1) and (2) are exactly similar, in their form; the former as relates to ρ , α , α' , ρ' , and the latter as to α''' , ρ''' , ρ'' , α'' . Now, the first system gives

$$\delta \mu \cdot \sin \rho + \mu \delta \rho \cdot \cos \rho = 0; \quad \delta \alpha' = \delta \rho; \quad \delta \rho' \cos \rho' = \delta \mu \cdot \sin \alpha' + \mu \delta \alpha' \cdot \cos \alpha';$$

whence, by elimination and reduction, we find

$$\delta \rho' = \frac{\sin I}{\cos \rho \cdot \cos \rho'} \delta \mu; \quad (e)$$

Light. and, consequently, by reason of the analogy of the two systems of equations pointed out above,

Part II.

$$\delta a'' = - \frac{\sin I''}{\cos a''' \cdot \cos a'} \delta \mu' \quad (f)$$

But, since $a'' = I' + \rho'$, we have $\delta \rho' = \delta a''$, so that we finally get

$$\frac{\cos \rho \cdot \cos \rho'}{\cos a''' \cdot \cos a''} = - \frac{\sin I}{\sin I''} \cdot \frac{\delta \mu}{\delta \mu'}. \quad (g)$$

The property expressed by this equation may be thus stated. Conceive the ray to pass both ways outwards from a point in its course between the two prisms; then, in order that the combination may be achromatic, the products of the cosines of its incidences on the surfaces of each prism must be to each other in the ratio compounded of that of the sines of their respective refracting angles, and the differences of their refractive indices for red and violet rays; besides which, they must refract in opposition to each other, or I and I'' their refracting angles must have opposite signs.

The combination of this equation with the system of equations above stated, expressing the conditions of refraction by the prism, and their relative position with regard to each other (which is included in the equation $a'' = I' + \rho'$) suffice, algebraically speaking, to resolve every problem which can occur, of this kind; but the final equations are for the most part too involved to allow of direct solution. Nevertheless, the results we have arrived at will furnish occasion for remarks of moment; and, first, since ρ' is the angle of refraction from the second surface of the first prism, $\delta \rho'$ is the angular breadth of the spectrum produced by it; this is, therefore, proportional, *ceteris paribus*, to the product of the secants of the angles of refraction at its two surfaces. Let us trace the progress of the variation of this, as the incident ray changes its inclination to the first surface, beginning with the case when it just grazes the surface from the back towards the edge. In this case

449.
Progress of dispersion by a single prism traced

$a = 90^\circ$, $\sin \rho = \frac{1}{\mu}$, consequently ρ , and therefore $I + \rho$ or a' , and therefore ρ' are all finite, and at their

maximum. Hence $\cos \rho \cdot \cos \rho'$ is finite, and at its minimum; and therefore $\delta \rho'$, or the breadth of the spectrum, is also finite, but a maximum. As the incident ray becomes more inclined to the surface ρ , and therefore a' and ρ' diminish, and the denominator of $\delta \rho'$ increases, so that the breadth of the spectrum diminishes, and reaches a minimum when $\cos \rho \cdot \cos \rho'$ attains its maximum; that is, when $d\rho \cdot \tan \rho + d\rho' \cdot \tan \rho' = 0$. Now this equation, substituting and reducing gives, for determining the value of ρ , and therefore of a , or the incidence when the spectrum is a minimum,

Position of least dispersion determined.

$$\mu^2 \cdot \sin (I + \rho) \cdot \cos (I + 2\rho) + \sin \rho = 0. \quad (h)$$

Hence we see that the position which gives a minimum of breadth to the spectrum is very different from that which gives a minimum of deviation, being given by the above equation, which is easily resolved by a table of

logarithms, and which shows at once that ρ must be greater than $45^\circ - \frac{I}{2}$.

After attaining the position so determined, the breadth of the spectrum again increases, and continues to do so till the rays can be no longer transmitted through the prism. At this limit the emergent ray just grazes the posterior face of the prism from its thinner towards its thicker part $\rho' = 90^\circ$, $\cos \rho' = 0$. At this limit, therefore, the dispersion becomes infinite. All these stages are easily traced by turning a prism round its edge between the eye and a candle; or, better, between the eye and the narrow slit between two nearly closed window-shutters.

Hence, as the incident ray varies from the position SE (fig. 105) to S'E, and therefore the refracted from FG to F'G', the breadth of the spectrum commences at a maximum, but finite value, diminishes to a minimum and then increases to infinity. The distribution of the colours in the spectrum, or the breadths of the several coloured spaces in any state of the data, will moreover differ according to the values of ρ , ρ' and $\sin I$; for the equation (e), by assigning to $\delta \mu$ the values which correspond in succession to the intervals between red and orange, orange and yellow, yellow and green, &c. will give the corresponding values of $\delta \rho'$, or the apparent breadths of these spaces. Now the denominator $\cos \rho \cdot \cos \rho'$ is an implicit function of μ , and therefore varies when the initial ray is taken in different parts of the spectrum. The variation is trifling when the angles ρ , ρ' are considerable; but near the limit, when the ray can barely be transmitted, it becomes very great, the spectrum is violently distorted, and the violet extremity greatly lengthened in proportion to the red. The effect is the same as if the nature of the medium changed and descended lower in the order of substances in the table Art. 443.

450.
Distortion of spectrum at extreme incidences. Fig. 105.

From what has just been said, we see the possibility of achromatising any prism, however large its refracting angle, by any other of the same medium, however small may be its angle; for since, by properly presenting a prism to the incident ray, its dispersion may be increased to infinity; if made to refract in opposition to another whose dispersion has any magnitude, however great, it may be made to counteract, or even overcome it. Thus in fig. 106 the dispersion of the second prism a , of small refracting angle, being increased by the effect of its inclined position, is rendered equal and opposite to that of the prism A, whose refracting angle is large.

451.
Achromatic combinations of one medium. Fig. 106.

When the prisms differ greatly in their angles, however, the second must be very much inclined, so as to bring it near to the limit of transmission. In this case, its law of dispersion, as just shown, will be greatly disturbed, and rendered totally different from what obtains in the other prisms; so that perfect achromaticity spectra.

452.
Subordinate

Light.

Part II.

cannot be produced; but when the extreme red and violet rays are united, the green will be too little refracted by the second prism, and a purple and green spectrum will arise, as in the case of prisms of different media. To this spectrum Dr. Brewster (who was the first to place it in evidence) has given the name of a *tertiary spectrum*; but it appears to us, that this term had better be reserved for the spectra mentioned in Art. 446, and those now in question may be called *subordinate spectra*.

If a small rectangular object be viewed through such a combination as above described, in which the prism A is placed in its position of minimum deviation, and achromatised by a second α , whose angle is less than that of A, but not so small as to introduce this cause of colour, it will appear distorted in figure; for the sides parallel to the edges of the prisms will undergo no change in their apparent length, while the breadth of the rectangle will appear magnified. For the first prism, by reason of its position, does not alter the angular dimensions of objects seen through it; but the second changes their angular breadth in the ratio of $d\rho'''$ to

$d\alpha''$, that is (by differentiation) in the ratio of $\frac{\cos \alpha \cdot \cos \alpha'}{\cos \rho \cdot \cos \rho'}$ to unity, a ratio which increases rapidly as the

inclination of the prism increases, and ρ' approaches a right angle.

453.

Amici's
prismatic
telescope.

M. Amici has taken advantage of these properties to construct a species of achromatic telescope, which, at first sight, appears very paradoxical, being composed merely of four prisms of the *same* kind of glass, with plane surfaces. To understand its construction, conceive a small square object op placed with the side o parallel to the refracting edges of a pair of prisms so adjusted, and perpendicular to their principal sections, *i. e.* to the plane of the paper. Then, after refraction through both, it will be seen by an eye at E, as a real object $o'p'$, having its length o unaltered, but magnified in breadth. Now, if we add a second pair of prisms, similar to the first, and similarly disposed with respect to each other, so as to form a second achromatic combination, but having the plane of their principal sections at right angles to the former, producing a refraction perpendicular to the plane of the paper, or parallel to the *length* of the distorted square, this will be in like manner seen as a real and colourless object, but again distorted, its side $o'p'$ remaining unaltered, but o' being magnified. Thus, by the effect of the first distortion, the breadth of the square is magnified, and, by that of the second, its length, and in the same ratio; and therefore the final result will be an image undistorted, achromatic, and magnified. The writer of these pages had the pleasure of witnessing the very good performance of one of these singular telescopes, magnifying about four times in the hands of its inventor, at Modena, in 1826. It is evident, that, by superposing several such telescopes on each other, the magnifying power may be increased in geometrical progression. It is equally clear, that, by using prisms of two different media to form the several binary combinations, the *subordinate spectra* may be made to counteract the *secondary spectra*, arising from the difference in the scales of dispersion in the two media; and thus an achromaticity, almost mathematically perfect, might be obtained. It is worthy of consideration, whether, for the purpose of viewing very bright objects, as the sun, for instance, this species of telescope might not prove of considerable service. It would have the advantage of being its own darkening glass, of not bringing the rays to a focus, and therefore of requiring no extraordinary care in the figuring of the surfaces; and, in short, of being exempt from all those inconveniencies which oppose the perfection of telescopes of the usual constructions, as applied to this particular object.

454.

Conditions
of achroma-
ticity for
several
prisms of
small angles

Proposition. To find the conditions of achromaticity when several prisms of different media refract a ray of white light, supposing all their refracting angles very small, and the ray to pass nearly at right angles to the principal section of each.

The refracting angles being A, A', A'', &c., and the refractive indices $\mu, \mu', \&c.$, the several partial deviations will be $D = (\mu - 1) A$; $D' = (\mu' - 1) A'$, &c.; and their sum, or the total deviation, will be $(\mu - 1) A + (\mu' - 1) A' + (\mu'' - 1) A'' + \&c.$ In order that the emergent ray may be colourless, this must be the same for rays of all colours; and its variation, when $\mu, \mu', \&c.$ are made to vary, must vanish, or

$$A \delta \mu + A' \delta \mu' + A'' \delta \mu'' + \&c. = 0.$$

Now, by equation (d) of Art. 439, we have $\delta \mu$, (or, in the notation of that article, $\mu - \mu_0$)

$$= (\mu_0 - 1) \left\{ a \cdot \frac{\delta x}{x_0 - 1} + b \left(\frac{\delta x}{x_0 - 1} \right)^2 + \&c. \right\}$$

Therefore the above equation gives, when arranged according to powers of δx ,

$$\begin{aligned} 0 = & \left\{ A (\mu_0 - 1) a + A' (\mu'_0 - 1) a' + A'' (\mu''_0 - 1) a'' + \&c. \right\} \cdot \frac{\delta x}{x_0 - 1} \\ & + \left\{ A (\mu_0 - 1) b + A' (\mu'_0 - 1) b' + A'' (\mu''_0 - 1) b'' + \&c. \right\} \left(\frac{\delta x}{x_0 - 1} \right)^2 \\ & + \&c. \end{aligned}$$

taking $a', b', \&c.$ to represent the dispersive powers of the various orders for the second prism, $a'', b'', \&c.$ for the third, and so on. Hence, in order that this may vanish for all the rays in the spectrum, we must have (putting, for brevity, μ for μ_0 , μ' for μ'_0 , &c.)

Light.

Part II.

$$\left. \begin{aligned} (\mu - 1) \cdot A a + (\mu' - 1) A' a' + (\mu'' - 1) A'' a'' + \&c. &= 0 \\ (\mu - 1) \cdot A b + (\mu' - 1) A' b' + (\mu'' - 1) A'' b'' + \&c. &= 0 \\ (\mu - 1) \cdot A c + (\mu' - 1) A' c' + (\mu'' - 1) A'' c'' + \&c. &= 0 \end{aligned} \right\} \quad (i)$$

and so on. Generally speaking, the number of these equations being infinite, no finite number of prisms can satisfy them all; but if we attempt only to unite as many rays in the spectrum as there are prisms, which is the greatest approach to achromaticity we can attain, we have as many equations as unknown quantities, *minus* one, and the ratios of the angles to each other become known. Thus, to unite two rays two media suffice, and we can only take into consideration the first order of dispersions, which give

$$(\mu - 1) A a + (\mu' - 1) A' a' = 0; \quad \frac{A'}{A} = - \frac{\mu - 1}{\mu' - 1} \cdot \frac{a}{a'}. \quad (j)$$

To unite three rays we have

$$\begin{aligned} (\mu - 1) A a + (\mu' - 1) A' a' + (\mu'' - 1) A'' a'' &= 0 \\ (\mu - 1) A b + (\mu' - 1) A' b' + (\mu'' - 1) A'' b'' &= 0 \end{aligned}$$

whence by elimination

$$\frac{A'}{A} = - \frac{\mu - 1}{\mu' - 1} \cdot \frac{a b'' - b' a''}{a' b'' - b' a''} \quad \frac{A''}{A} = - \frac{\mu - 1}{\mu'' - 1} \cdot \frac{a b' - b a'}{a' b' - b' a'}; \quad (k)$$

and so on for any number.

In the case of two media, if the quantities *b*, *c*, &c. be not known, the dispersive powers of the first order, 453. *a*, *a'*, should be determined, not by comparison of the extreme red and violet rays, which are too little luminous to render their strict union a matter of importance; we should rather endeavour to unite those rays which are at once powerfully illuminating, and differing much in colour, such as the rays *D* and *F*. The exact union of these will insure the approximate union of all the rest better, on the whole, than if we aimed at uniting the extremes of the spectrum, and a far greater concentration of light will be produced. This should be carefully borne in mind in all experiments on the dispersions of glass to be used in the construction of telescopes. Case of two media. Best rays to unite.

If we would produce the greatest possible achromaticity by three prisms, the rays to be selected for determining the values of *a*, *b*, *a'*, *b'*, should be *C*, *E*, and *G*; or, which would, perhaps, be still better, *C*, *F*, and a ray half way between *D* and *E*; but the want of a sufficiently well marked line in that part of the spectrum throws some slight difficulty in the way of this latter combination, when solar light is used, and would oblige us to have recourse to some other method of measurement, of which a variety might be suggested. 454. Best rays to unite in case of three media.

In the case of three media, if the numerators and denominators of the expressions (*k*) vanish, or nearly so, the solutions become illusory, or at least inapplicable in practice. This happens whenever either of the fractions 455. Cases in which the formulæ become inapplicable to practice.

$\frac{a}{a'}, \frac{a}{a''}, \frac{a'}{a''}$ becomes equal to either of the corresponding fractions $\frac{b}{b'}, \frac{b}{b''}$, or $\frac{b'}{b''}$. Hence, to obtain

practicable combinations, it is necessary to employ media which differ as much as possible in their *scales* of dispersive powers, *i. e.* in which the coloured spaces differ as far as possible from proportionality; such, for instance, as flint glass, crown glass, and muriatic acid; or, still better, oil of cassia, crown glass, and sulphuric acid, &c.

§ II. Of the Achromatic Telescope.

In the refracting telescopes described in Art. 380, &c. the different refrangibility of the differently coloured rays presents an obstacle to the extension of their power beyond very moderate limits. The focus of a lens being shorter as the refractive index is greater, it follows, that one and the same lens refracts violet rays to a focus nearer to its surface than red. This is easily seen by exposing a lens to the sun's rays, and receiving the converging cone of rays on a paper placed successively at different distances behind it. At any distance nearer to the lens than its focus for mean rays, the circle on the paper will have a red border, but beyond it a blue one; for the cone of red rays whose base is the lens, envelopes that of violet *within* the focus, its vertex lying beyond the other, but is enveloped by it *without*, for the converse reason. Hence, if the paper be held in the focus for mean rays, or between the vertices of the red and violet cones, *these* will then form a distinct image, being collected in a point: but the extreme, and all the other intermediate rays, will be diffused over circles of a sensible magnitude, and form coloured borders, rendering the image indistinct and hazy. This deviation of the several coloured rays from one focus is called the "*chromatic aberration*." 456. Chromatic aberration explained.

The diameter of the least circle within which all the coloured rays are concentrated by a lens supposed free from spherical aberration is easily found. Thus, in fig. 107, if *v* be the focus for violet, and *r* for red rays, *n m o* will be the diameter of this circle. Now, by similar triangles, $n o = A B \cdot \frac{m v}{C v}$, and also $n o = A B \cdot \frac{m r}{C r}$; Least circle of chromatic aberration. Fig. 107

Light. therefore equating these $\frac{mv}{Cv} = \frac{mr}{Cr}$, and $mv = mr \cdot \frac{Cv}{Cr}$, $mv + mr = mr \cdot \frac{Cr + Cv}{Cr} = rv$, consequently $mr = rv \cdot \frac{Cr}{Cr + Cv} = rv \cdot \frac{Cr}{2Cr - rv} = \frac{rv}{2}$ very nearly, since the dispersion is small in comparison with the whole refraction. Therefore $no = \frac{AB}{2} \cdot \frac{rv}{Cr}$. Now, f being the reciprocal focal distance ($= L + D = (\mu - 1)(R' - R'') + D$) we have $rv = -\delta \frac{1}{f} = \frac{\delta f}{f^2} = \frac{\delta \mu (R' - R'')}{f^2}$
 $= \frac{\delta \mu}{\mu - 1} \cdot \frac{L}{f^2}$ and $Cr = \frac{1}{f}$, supposing μ to represent the index of refraction for extreme red rays. Hence we get diameter of least circle of chromatic aberration $=$ semi-aperture $\times \frac{L}{f} \cdot \frac{\delta \mu}{\mu - 1}$
 $=$ semi-aperture \times dispersive index $\times \frac{L}{f}$;

and for parallel rays, when $L = f$, simply semi-aperture \times dispersive index.

458. *Corol.* Hence the circle of least colour has the same absolute linear magnitude whatever be the focal length of the lens, provided the aperture be the same. Now, in the telescope, the magnifying power, or the absolute linear magnitude of the image viewed by a given eye-glass, increases in the ratio of the focal length of the object-glass, (382.) Therefore, by increasing the focal length of an object-glass without increasing its aperture, the breadth of the coloured border round the image of any object diminishes in proportion to the image itself, and thus the confusion of vision is diminished, and the telescope will possess a proportionally higher magnifying power. In consequence of this property, before the invention of the achromatic telescope, astronomers were in the habit of using refracting telescopes of enormous length, even so far as 100 or 150 feet; and Huyghens, in particular, distinguished himself by the magnitude and excellent workmanship of his glasses, and by the important astronomical discoveries made with them.

459. The achromatic object-glass, however, by enabling us to reduce the length of the telescope within more reasonable bounds, has rendered it a vastly more manageable and useful instrument. To conceive its principle, we have only to recur to what has already been said in Art. 451—454, respecting achromatic prisms. A lens is nothing more than a system of infinitely small prisms arranged in circular zones round a centre, with refracting angles increasing as their distance from the centre increases, so as to refract all the rays to one point; and if we can achromatise each elementary prism, the whole system is achromatic. The equations (i) apply at once to this view of the structure of a lens. For, suppose R' , R'' to be the curvatures of the two surfaces of the first lens, L' its power, and μ' its refractive index, then, for a given aperture, or at a given distance from the centre, $R' - R''$, the difference of the curvatures, expresses the angle made by tangents to the surfaces, or the refracting angle of the elementary prism; or $R' - R'' = A'$; and similarly for the other lenses, $A'' = R''' - R''$, and so on, so that the equations become

General equations of achromaticity. that is simply

$$\begin{aligned} &(\mu' - 1)(R' - R'') \cdot a' + (\mu'' - 1)(R''' - R'') a'' + \&c. = 0 \&c.; \\ &\left. \begin{aligned} L' \cdot a' + L'' \cdot a'' + L''' \cdot a''' + \&c. &= 0 \\ L' \cdot b' + L'' \cdot b'' + L''' \cdot b''' + \&c. &= 0 \\ L' \cdot c' + L'' \cdot c'' + L''' \cdot c''' + \&c. &= 0 \\ \&c. \end{aligned} \right\} \quad (a) \end{aligned}$$

460. These equations afford all the relations necessary to insure achromaticity; and when satisfied, since they do not contain D , they show that an object-glass which is achromatic for any one distance of the object is so for all distances. It is evident, that the same system of equations may be obtained directly from the expression in Art. 265 for the joint power of a system of lenses whose individual powers are L' , L'' , &c. For the condition of achromaticity gives $\delta L = 0$, that is

$$\delta L' + \delta L'' + \delta L''' + \&c. = 0.$$

But since $L' = (\mu' - 1)(R' - R'')$ &c. (according to the system of notation there adopted)

$$\delta L' = (R' - R'') \delta \mu' = L' \cdot \frac{\delta \mu'}{\mu' - 1}.$$

But in the equation (d) if we put in succession for μ_0 the values μ' , μ'' , &c., for $\mu - \mu_0$ respectively, $\delta \mu'$, $\delta \mu''$, &c., and for a , b , &c. the systems of coefficients a' , b' , &c.; a'' , b'' , &c.; and suppose $\frac{x - x_0}{x_0 - 1} = p$, we shall have

$$\frac{\delta \mu'}{\mu' - 1} = a' p + b' p^2 + \&c.; \quad \frac{\delta \mu''}{\mu'' - 1} = a'' p + b'' p^2 + \&c.;$$

Light. and therefore

Part II,

$$0 = L' \{ a' p + b' p^2 + \&c. \} + L'' \{ a'' p + b'' p^2 + \&c. \} + \&c.$$

which, being made to vanish independently of p , gives the very same system of equations as (a.)

To satisfy all these equations at once with any finite number of lenses being impossible, we must rest content with satisfying as many of the most important as the number of lenses will permit. Thus, if we have two lenses of different media, such as flint and crown glass, for instance, one only of them can be satisfied, and this must of course be the first, viz.

$$L' a' + L'' a'' = 0, \quad \text{or} \quad \frac{L''}{L'} = - \frac{a'}{a''}; \quad (b)$$

which shows that the powers of the lenses must oppose each other, and be to each other inversely (and of course their focal lengths directly) as the dispersive powers. In such a combination, the values of a' , a'' , the dispersive powers, however, ought not to be obtained from the relative refractions for the extreme red and violet rays of the spectrum, (according to the remark in Art. 453,) but rather from the strongest and brightest rays whose colours are in decided contrast; such, for instance, as the rays C and F in Fraunhofer's scale.

With three lenses of different media, two of the equations of achromaticity can be satisfied, and the secondary spectrum corrected, thus we have

$$\left. \begin{aligned} 0 &= L' a' + L'' a'' + L''' a''' \\ 0 &= L' b' + L'' b'' + L''' b''' \end{aligned} \right\} \left\{ \begin{aligned} \frac{L''}{L'} &= - \frac{a' b''' - b' a'''}{a'' b''' - b'' a'''} \\ \frac{L'''}{L'} &= - \frac{a' b'' - b' a''}{a'' b'' - b'' a''} \end{aligned} \right\}; \quad (c)$$

and in determining the values of a' , b' , &c. the rays to be employed should be the brightest yellow for a middle ray, and a pretty strong red and blue for the extremes. The rays B, E, H are perhaps inferior to C, E, G for this purpose.

Hence in a double object-glass having a positive focus the least dispersive lens must be of a convex or positive, and the most so of a negative, or concave character. The order in which they are placed is of no consequence, as far as achromaticity is concerned.

A single lens, as we have seen, neither admits of the destruction of the spherical, nor chromatic aberration, (Art. 296 and 457;) but if we combine two or more lenses of different media, the equations s , t , u , v of Art. 309, 310, 312, and 313, combined with the equations just derived (a), Art. 459, or so many of them as are not incompatible, afford us the means of annihilating both species of aberration at once; and what is curious, and must be regarded as singularly fortunate, the relations afforded by the destruction of the chromatic aberration, which, at first sight, would appear likely greatly to complicate the inquiry, tend, on the contrary, remarkably to simplify it, being in fact the very relations the analyst would fix upon to limit his symbols, and give his final equations the greatest simplicity their nature admits, if left at his disposal. For, it will be remarked, that in the general equation for the destruction of the spherical aberration, $\Delta f = 0$, or

$$0 = \frac{L'}{\mu'} (\alpha' - \beta' D' + \gamma' D'^2) + \frac{L''}{\mu''} (\alpha'' + \beta'' D'' + \gamma'' D''^2) + \&c.; \quad (d)$$

the expressions within the parentheses are all of the second degree when expressed in terms of the curvatures of the surfaces, and of $D' = D$ the proximity of the radiant point to the first lens; and as L' , L'' , &c. are respectively of the first degree, in terms of the curvatures, the whole is, in its general form, of the third degree, and the equation of a cubic form. But the conditions of achromaticity, which assign relations only between L' , L'' , &c. without involving R' , R'' , &c. enable us to eliminate these quantities and replace them in the above equation, by giving combinations of a' , a'' , b' , b'' , &c., so that it becomes reduced to a quadratic form, and its treatment simplified accordingly.

Let us proceed now to develop the equation (d), in which, according to the foregoing remark, when the conditions of achromaticity are introduced, L' , L'' , &c. may be regarded as given quantities; for, taking $L = L' + L'' + \&c. =$ the power of the compound lens, (which we may suppose given, or, if we please, assume equal to unity,) this, combined with the equations (a), determines the values of L' , &c. Thus, in the case of two lenses,

if we put π for the ratio of the dispersive powers, or $\pi = \frac{a'}{a''}$ we have $L' = \frac{L}{1 - \pi}$, $L'' = - \frac{\pi L}{1 - \pi}$; and

similarly for three or more lenses. Suppose then we represent by r' , r'' , r''' , &c. the respective curvatures of the first, or anterior surfaces of the first, second, third, &c. lens, in order; the first being that on which the rays first

fall. Then we have $L' = (\mu' - 1) (R' - R'') = (\mu' - 1) (r' - R'')$ so that $R'' = r' - \frac{L'}{\mu' - 1}$; and similarly

$R''' = r'' - \frac{L''}{\mu'' - 1}$, &c. We must therefore put in the foregoing expressions

$$R' = r'; \quad R'' = r' - \frac{L'}{\mu' - 1}; \quad R''' = r''; \quad R^{iv} = r'' - \frac{L''}{\mu'' - 1}, \quad \&c.$$

461.
Object glass
of two
media.

462.
Object glass
of three
media.

463.

464.
Simulta-
neous de-
struction of
both aberra-
tions.

465.
Determina-
tion of the
powers of the
several
lenses.

Develop-
ment of the
general
equation.

Light.

Hence by substitution of these in the values of a, β , &c. (Art. 293) we get

Part II.

$$a' = (2 + \mu') r'^2 - (2\mu' + 1) \cdot \frac{\mu'}{\mu' - 1} \cdot L' r' + \mu' \cdot \left(\frac{\mu'}{\mu' - 1} \right)^2 L^2$$

$$\beta' = (4 + 4\mu') r' - (3\mu' + 1) \cdot \frac{\mu'}{\mu' - 1} L'$$

$$\gamma' = 2 + 3\mu',$$

and similarly for $\alpha'', \beta'', \gamma'',$ &c. So that, substituting again these expressions, and putting for D'' its equal $L' + D'$, for D''' its equal $L' + L'' + D'$, and so on, we have, finally, for the general equation $\Delta f = 0$, as follows:

$$\begin{aligned} 0 = & \left\{ \left(\frac{2}{\mu'} + 1 \right) L' r'^2 + \left(\frac{2}{\mu''} + 1 \right) L'' r''^2 + \left(\frac{2}{\mu'''} + 1 \right) L''' r'''^2 + \&c. \right\} \\ & - \left\{ \frac{2\mu' + 1}{\mu' - 1} L'^2 r' + \frac{2\mu'' + 1}{\mu'' - 1} L''^2 r'' + \frac{2\mu''' + 1}{\mu''' - 1} L'''^2 r''' + \&c. \right\} \\ & - 4 \left\{ \left(1 + \frac{1}{\mu''} \right) L' L'' r'' + \left(1 + \frac{1}{\mu'''} \right) (L' + L'') L''' r''' + \&c. \right\} \\ & + \left\{ \left(\frac{\mu'}{\mu' - 1} \right)^2 L'^3 + \left(\frac{\mu''}{\mu'' - 1} \right)^2 L''^3 + \left(\frac{\mu'''}{\mu''' - 1} \right)^2 L'''^3 + \&c. \right\} \\ & + \left\{ \frac{3\mu' + 1}{\mu' - 1} L' L'^2 + \frac{3\mu'' + 1}{\mu'' - 1} (L' + L'') L''^2 + \&c. \right\} \\ & + \left\{ \left(\frac{2}{\mu''} + 3 \right) L'^2 L'' + \left(\frac{2}{\mu'''} + 3 \right) (L' + L'')^2 L''' + \&c. \right\} \\ & + D \cdot \left\{ \begin{aligned} & - 4 \left\{ \left(1 + \frac{1}{\mu'} \right) L' r' + \left(1 + \frac{1}{\mu''} \right) L'' r'' + \left(1 + \frac{1}{\mu'''} \right) L''' r''' + \&c. \right\} \\ & + \left\{ \frac{3\mu' + 1}{\mu' - 1} L'^2 + \frac{3\mu'' + 1}{\mu'' - 1} L''^2 + \frac{3\mu''' + 1}{\mu''' - 1} L'''^2 + \&c. \right\} \\ & + 2 \left\{ \left(\frac{2}{\mu''} + 3 \right) L' L'' + \left(\frac{2}{\mu'''} + 3 \right) (L' + L'') L''' + \&c. \right\} \\ & + D^2 \left\{ \left(\frac{2}{\mu'} + 3 \right) L' + \left(\frac{2}{\mu''} + 3 \right) L'' + \left(\frac{2}{\mu'''} + 3 \right) L''' + \&c. \right\} \end{aligned} \right\} \end{aligned}$$

466. For brevity, let us represent by X , the terms of this expression, independent of the quantity D ; by Y , the assemblage of terms multiplied by D' ; and by Z , those multiplied by D'^2 , and we shall have

$$\Delta f = \frac{y^2}{2} \{ X + Y \cdot D + Z \cdot D^2 \};$$

and if this vanish the aberration is destroyed. Now, first, if we regard only parallel rays, or suppose $D = 0$, this reduces itself to $X = 0$, so that the condition $X = 0$ being satisfied, the telescope will be perfect when used for astronomical purposes, or for viewing objects so distant that D' may be disregarded.

467. The equation $X = 0$ is of the second degree in each of the quantities $r', r'',$ &c., whose number is that of the lenses. Consequently, this condition alone is not sufficient to fix their values; and, without assuming some further relations between them, or some other limitations, the problem is indeterminate, and the aberration may be destroyed in an infinite variety of ways. Confining ourselves at present to the consideration of two lenses only, since $X = 0$ contains only two unknown quantities, one other equation only is required, and we have only to consider what other condition will be attended with the greatest practical advantages. Clairaut has proposed to adjust the two lenses so as to have their adjacent surfaces in contact throughout their whole extent, to allow of their being cemented together, and thus avoid the loss of light by reflection at these surfaces. This certainly would be a great advantage were it possible so to cement two glasses of large size together, as to bring neither of them into a state of strain as the cement cools, or otherwise fixes; and were it not for the further inconvenience, that the media being of course differently expansible by heat, every subsequent change of temperature would necessarily distort their figure, as well as strain their parts, when thus forcibly held together, just as we see a compound lamina of two differently expansible metals assume a greater or less curvature, according to the temperature it is exposed to. Meanwhile the condition in question is algebraically expressed by $L' = (\mu' - 1)(r' - r'')$; for in this case $R' = r'$, and $R'' = R''' = r''$, and this being of the first degree only in r', r'' , affords a final equation of a quadratic form by elimination with $X = 0$, which latter, in the case before us of two lenses, is the same as the equation (v), Art. 312, writing only r' for R' , and r'' for R''' .

Conditions limiting it. Clairaut's.

Light.

But this condition of Clairaut's has another and much greater inconvenience, which is, that the resulting quadratic has its roots imaginary, when the refractive and dispersive powers of the glasses are such as are by no means unlikely to occur in practice; and without the limits of refraction and dispersion, for which they are real, the resulting curvatures change so rapidly on slight variations of the data, as to make their computation delicate, and interpolation between them, so as to form a table, very troublesome. D'Alembert, in his *Opuscules*, tom. iii., has proposed a variety of other limitations, such, for instance, as annihilating the spherical aberration for rays

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D'Alembert's.

of all colours, (which comes to the same as supposing at once $X = 0$ and $\frac{\delta X}{\delta \mu'} \delta \mu' + \frac{\delta X}{\delta \mu''} \delta \mu'' = 0$, and

which leads to biquadratic equations, and affords no practical advantage,) &c. But, without going into useless refinements of this kind, the very form of the general equation $X + Y \cdot D' + Z \cdot D'^2 = 0$ points out a condition combining every advantage the case is susceptible of. This consists in putting $Y = 0$. By this supposition, the term depending on D' is destroyed, without assuming $D' = 0$; so that the telescope is not only perfect for parallel rays, but admits of as considerable a proximity of the object without losing its aplanatic character, as the nature of the case will allow. The term $Z \cdot D'^2$ indeed, or

Another proposed.

$$D'^2 \cdot \left\{ \left(\frac{2}{\mu'} + 3 \right) L' + \left(\frac{2}{\mu''} + 3 \right) L'' + \&c. \right\},$$

cannot vanish when two lenses only are used, being composed wholly of given functions of the refractive and dispersive powers, unless by D' itself vanishing, or by an accidental adjustment of the values of μ' , μ'' , L' , &c. But except the object be brought within a comparatively small distance from the telescope, (such as ten times its own length,) the square of D' is always so small as to allow of our disregarding this term, and considering the instrument as perfectly aplanatic when $Y = 0$. Now this equation, being of the first degree in r' , r'' , adds no new algebraic difficulty to the problem, but leads by elimination to a final quadratic; and, what is of most consequence, for such values of μ' , μ'' , and the dispersive ratio ω as occur in practice, the roots of this quadratic are always real, and the resulting curvatures of all the surfaces are moderate, and well adapted for practice; more so, indeed, than in any construction hitherto proposed. They are, moreover, such as to afford remarkable and peculiar facilities for interpolation, as we shall presently see. These reasons seem to leave no room for hesitation in fixing on the condition $Y = 0$, as that which ought to be introduced to limit the problem of the construction of a double object-glass, and to render it, so far as it can be rendered, *aplanatic*.

This equation, in the case in question, is

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$$0 = 4 \left(1 + \frac{1}{\mu'} \right) L' r' + 4 \left(1 + \frac{1}{\mu''} \right) L'' r'' - \frac{3\mu' + 1}{\mu' - 1} L'^2 - \left(6 + \frac{4}{\mu''} \right) L L'' - \frac{3\mu'' + 1}{\mu'' - 1} L''^2; \quad (f)$$

which is to be combined with (e), Art. 412, in which $R' = r'$ and $R'' = r''$. To reduce these to numbers, μ' , μ'' and the dispersive ratio ω must first be known. The readiest and most certain way in practice, for the use of the optician, is to form small object-glasses from specimens of the glasses intended to be employed, and by trial work them till the combination is as free from colour as possible, by the test usually had recourse to in practice. This is, to examine with a high magnifying power the image of a well defined white circle, or circular annulus on a black ground. If its edges are totally free from colour, the adjustment is perfect, but (owing to the secondary spectrum) this will seldom be the case; and there will generally be seen on the interior edge of the annulus a faint green, and on the exterior a purplish border, when the telescope is thrown a little out of focus by bringing the eye-glass too near the object-glass, and *vice versa*. The reason is, that while the great mass of orange and blue rays is collected in one focus, the red and violet are converged to a focus farther from, and the green to one nearer to the object-glass; the refraction of the green rays being in favour of the *convex* or *crown* glass, and of the red and violet (which united form purple) in favour of the flint (see table, Art. 443) or concave lens. The focal lengths of the lenses are then to be accurately determined, and the ratio of the dispersions (ω) will then be known, being the same with that of the focal lengths (454). The refractive indices will be best ascertained by direct observation, forming portions of each medium into small prisms. Now, ω being known, if we take

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unity for the power of the compound lens, we have $L' = \frac{1}{1 - \omega}$ and $L'' = -\frac{\omega}{1 - \omega}$, so that L' and L'' are

known, and we have therefore only to substitute their values and those of μ' , μ'' , in the algebraic expressions, and proceed to eliminate by the usual rules. The following compendious table contains the result of such calculations for the values of μ' , μ'' and ω therein stated, together with the amount of variation produced by varying either of the refractive indices independently of the other, for the sake of interpolation by proportional parts. Fig. 108 is a representation of the resulting object-glass.

Dimensions of an aplanatic object-glass.

Table for finding the Dimensions of an Aplanatic Object-glass.

Refractive index of crown, or convex lens = $\mu' = 1.524$.Refractive index of flint, or concave lens = $\mu'' = 1.585$.

Compound focal length = 10.000.

	CROWN LENS.					FLINT LENS.				
	First surface, convex.			Second surface, convex.	Focal length of crown lens.	Third surface, concave.	Fourth surface, convex.			Focal length of flint lens.
Dispersive ratio $\varpi =$	Radius for the above refractive indices.	Variation of radius for a change of + 0.010 in ref. index of crown glass.	Variation of radius for a change of + 0.010 in ref. index of flint glass.	Radius of convexity.		Radius of concavity.	Radius for the above refractive indices.	Variation of radius for a change of + 0.010 in ref. index of crown glass.	Variation of radius for a change of + 0.010 in ref. index of flint glass.	
0.50	6.7485	+ 0.0500	- 0.0030	4.2827	5.0	4.1575	14.3697	+ 0.9921	- 0.3962	10.0000
0.55	6.7184	+ 0.0740	- 0.0011	3.6332	4.5	3.6006	14.5353	+ 1.0080	- 0.5033	8.1818
0.60	6.7069	+ 0.0676	+ 0.0037	3.0488	4.0	3.0640	14.2937	+ 1.1049	- 0.5659	6.6667
0.65	6.7316	+ 0.0563	+ 0.0125	2.5208	3.5	2.5566	13.5709	+ 1.1614	- 0.6323	5.3846
0.70	6.8279	+ 0.0335	+ 0.0312	2.0422	3.0	2.0831	12.3154	+ 1.1613	- 0.7570	4.2858
0.75	7.0816	- 0.0174	+ 0.0568	1.6073	2.5	1.6450	10.5186	+ 1.0847	- 0.7207	3.3333

471.
Example of
the use of
the table.

To apply this table to any other proposed state of the data, we have only to consider that to compute the radius of any one of the surfaces, as the first or fourth, we have only to regard each element as varying separately, and take proportional parts for each. The following example will elucidate the process: Required the dimensions for an object-glass of 30 inches focus, the refractive index of the crown glass being 1.519, and that of the flint 1.589; the dispersive powers being as 0.567 : 1, or 0.567 being the dispersive ratio. Here $\mu' = 1.519$, $\mu'' = 1.589$, and $\varpi = 0.567$. The computation must first be instituted for a compound focus = 10.000, as in the table, and we proceed thus:

1st. Subtract the decimal (0.567) representing the dispersive ratio from 1.000, and 10 times the remainder ($= 10 \times 0.433 = 4.330$) is the focal length of the crown lens.

2nd. Divide unity by the decimal above mentioned, (0.567,) subtract 1.000 from the quotient ($\frac{1}{0.567} = 1.7635$, minus 1 = 0.7635) and the remainder multiplied by 10 (or 7.635) is the focal length of the flint lens. We must next determine by the tables the radii of the first and fourth surfaces for the dispersive ratios there set down (0.55 and 0.60) next less and next greater than the given one. For this purpose we have

Refractive powers given. 1.519 and 1.589

Refractive powers in table 1.524 .. 1.585

Differences - 0.005 + 0.004

The given refraction of the crown being less, and of the flint greater, than their average values on which the table is founded. Looking out now opposite to 0.55 in the first column for the variations in the two radii corresponding to a change of + 0.010 in the two refractions, we find as follows:

	First surface.	Fourth surface.
For a change = + 0.010 in the crown	+ 0.0740	+ 1.0080
For a change = + 0.010 in the flint	- 0.0011	- 0.5033

But the actual variation in the crown instead of + 0.010 being - 0.005, and of the flint + 0.004, we must take the proportional parts of these, changing the sign in the former case; thus we find the variations in the first and last radii to be

	First surface.	Fourth surface.
For - 0.005 variation in the crown	- 0.0370	- 0.5040
For + 0.004 variation in the flint ..	- 0.0004	- 0.2013
Total variation from both causes	- 0.0374	- 0.7053
But the radii in the table are	6.7184	14.5353
Hence the radii interpolated are	6.6810	13.8300

If we interpolate, by a process exactly similar, the same two radii for a dispersive ratio 0.60, we shall find, respectively,

	First surface.	Fourth surface.
For a variation of - 0.005 in the crown	- 0.0338	- 0.5524
For a variation of + 0.004 in the flint	+ 0.0015	- 0.2264
Total variation	- 0.0323	- 0.7788
Radii in table	6.7069	14.2937
Interpolated radii	6.6746	13.5149

Having thus got the radii corresponding to the actual refractions for the two dispersive ratios 0.55 and 0.60, it only remains to determine their values for the intermediate ratios 0.567 by proportional parts; thus

	First radius.	Fourth radius.	
For 0.600	6.6746	13.5149	
For 0.550	6.6810	13.8300	0.050 : 0.567 - 0.050 = 0.017 :: - 0.0064 : - 0.0022
Differences + 0.050	- 0.0064	- 0.3151	0.050 0.017 :: - 0.3151 : - 0.1071

So that $6.6810 - 0.0022 = 6.6788$, and $13.8300 - 0.1071 = 13.7229$, are the true radii corresponding to the given data. Thus we have, for the crown lens, focal length $= 4.330 = \frac{1}{L'}$, radius of first surface $= 6.6788$

$= \frac{1}{R'}$, index of refraction $= 1.519 = \mu'$, whence by the formula $L' = (\mu' - 1) (R' - R'') \frac{1}{R''}$ radius of the other surface is $- 3.3868$. Again, for the flint lens, the focal length $= \frac{1}{L''} = - 7.635$, radius of the posterior surface $= \frac{1}{R''} = - 13.7729$, index of refraction $\mu'' = 1.589$, whence we find $\frac{1}{R'''} = - 3.3871$

for the radius of the other surface. The four radii are thus obtained for a focal length of 10 inches, and multiplied by 3 we have for the telescope proposed

radius of first surface $= + 20.0364$; of second, $- 10.1604$; of third, $- 10.1613$; of fourth, $- 41.1687$.

Here, then, we see that the radii of the two interior surfaces of the double lens (fig. 108) differ by scarcely more than a thousandth part of an inch; so that, should it be thought desirable, they may be cemented together. This is not merely a casual coincidence, for the particular state of the data; if we cast our eyes down the table we shall find this approximate equality of the interior curvatures (those of the second and third surfaces) maintained in a singular manner throughout the whole extent of the variation of ω . Thus the construction, here proposed in reality for glasses of the ordinary materials, approaches considerably to that of Clairaut already mentioned.

In order to put these results to the test of experience, Mr. South procured an achromatic telescope to be executed on this construction by Mr. Tulley, one of the most eminent of our British artists, which is now in the possession of J. Moore, Esq. of Lincoln. Its focal length was 45 inches, and aperture $3\frac{1}{4}$, and its performance was found to be fully adequate to the expectation entertained of it, bearing a magnifying power of 300 with perfect distinctness, and separating easily a variety of double stars, &c. A more minute account of its performance will be found in the *Journal of the Royal Institution*, No. 26. Should the splendid example set by Fraunhofer be followed up, and the practice of the optician be in future directed by a rigorous adherence to theory, grounded on exact measurements of the refractive powers of his glasses on the several coloured rays, it will become necessary to develop the above table more in detail.

When three media are employed in the construction of object-glasses, it should be our object to obtain as great a difference as possible in their scales of action on the differently coloured rays. Dr. Blair, to whom we are indebted for the first extensive examination of the dispersive powers of media as a physical character, and who first perceived the necessity of destroying the secondary spectrum, and pointed out the means of doing it, is the only one hitherto who has bestowed much pains on this important part of practical optics; which, considering the extraordinary success he obtained, and the perfection of the telescopes constructed on his principles, is to be regretted. We have no idea, indeed, for the reasons already mentioned, that very large object-

light.

glasses, enclosing fluids, can ever be rendered available; but to render glasses of moderate dimensions more perfect, and capable of bearing a higher degree of magnifying power, is hardly less important as an object of practical utility. His experiments are to be found in the *Transactions of the Royal Society of Edinburgh*, 1791. We can here do little more than present a brief abstract of them.

475.

Dr. Blair's construction of an object-glass of three media.

Dr. Blair having first discovered that the secondary fringes are of unequal breadths, when binary achromatic combinations, having equal total refractions, are formed of different dispersive media, was immediately led to consider, that by employing two such different combinations to act in opposition to each other, if the total refractions were *equal*, the ray would emerge of course undeviated, and with its primary spectrum destroyed; but a secondary spectrum would remain, equal to the difference of the secondary spectra in the two combinations. Therefore, by a reasoning precisely similar to that which led to the correction of the primary spectrum itself, (Art. 426 and 427,) if we increase the total refraction of that combination A which, *ceteris paribus*, gives the *least* secondary spectrum, its secondary colour will be increased accordingly, till it becomes equal to that of the other B; so that the emergent beam will be free from the secondary spectra altogether, and will be deviated on the whole in favour of the combination A. Reasoning on these grounds, Dr. Blair formed a compound, or binary achromatic convex lens A, (fig. 109,) of two fluids *a* and *b*, (two essential oils, such as naphtha and oil of turpentine, differing considerably in dispersion,) which, when examined alone, was found to have a greater refractive power on the green rays than on the united red and violet. He also formed a second binary lens B, of a *concave* character, and also achromatic, (*i. e.* having the primary spectrum destroyed,) consisting of the more dispersive oil (*b*) and glass, and in which the green rays are also more refracted than the united red and violet, but in a *greater degree* in proportion to the whole deviation, than in the other combination; and in precisely the same degree was the focal length of this lens increased or its refraction diminished, when compared with that of the combination A. When, therefore, these two lenses were placed together, as in fig. 109, an excess of refraction remained in favour of the convex combination; but the secondary spectra of each being equal and opposite (by reason of the opposite character of the lenses) were totally destroyed. In fact, he states, that in a compound lens so constructed, he could discover no colour by the most rigid test; and thence concluded, not only the red, violet, and green to be united, but also all the rest of the rays, no outstanding colour of blue or yellow being discernible. In placing the lenses together, the intermediate plane glasses may be suppressed altogether, as in fig. 110.

476.

Remarkable property of the muriatic acid.

It was in the course of these researches that Dr. Blair was led to the knowledge of the possibility of forming binary combinations, having secondary spectra of opposite characters; that is, in which (the total refraction lying the same way) the order of the colours in the secondary spectra should be inverted. In other words, that while in some combinations the green rays are more refracted than the united red and violet, in others they are less so. He found, for instance, that while in most of the highly dispersive media, including metallic solutions, the green lay among the less refrangible rays of the spectrum, there yet exist media considerably dispersive, in which the reverse holds good. The muriatic acid, among others, is in this predicament. Hence, in binary combinations of glass with this acid, the secondary spectrum consists of colours oppositely disposed from that formed by glass and the oils, or by crown and flint glass. In consequence of this, to form an object-glass of two binary combinations, as described in the last article, they must both be of convex characters. But this affords no particular advantage. Dr. Blair, however, considered the matter in another and much more important light, as offering the means of dispensing with a third medium altogether, and producing by a single binary combination a refraction absolutely free from secondary colour. To this end he considered, that it appears to depend entirely on the chemical nature of the refracting medium, what shall be the order and distribution of the colours in the spectrum, as well as what shall be the total refraction and dispersive powers of the medium; and that therefore by varying properly the ingredients of a medium, it may be practicable, without greatly varying the total refraction and dispersion, still to produce a considerable change in the internal arrangement (if we may use the phrase) of the spectrum; and therefore, perhaps, to form a compound medium in which the seven colours shall occupy spaces regulated by any proposed law, (within certain limits.) Now if a medium could be so compounded as to have the same scale of dispersions, or the same law of distribution of the colours as crown glass with a different absolute dispersion, as we have already seen, nothing more would be required for the perfection of the *double* object-glass. The property of the muriatic acid just mentioned puts this in our power.

It is observed, that the presence of a metal (antimony, for instance) in a fluid, while it gives it a high refractive and dispersive power, at the same time tends to dilate the more refrangible part of the spectrum beyond its due proportion to the less. On the other hand, the presence of muriatic acid tends to produce a contrary effect, contracting the more refrangible part and dilating the less, beyond that proportion which they have in glass. Hence, Dr. Blair was led to conclude, that by mixing muriatic acid with metallic solutions, in proportions to be determined by experience, a fluid might be obtained with the wished for property; and this on trial he found to be the case. The metals he used were antimony and mercury; and to ensure the presence of a sufficient quantity of muriatic acid, he employed them in the state of muriates, in aqueous solution; or, in the case of mercury, in a solution of sal ammoniac, which is a compound of ammonia and muriatic acid, and which is capable of dissolving a considerably greater quantity of corrosive sublimate (muriate, or chloride of mercury) than water alone. By adding liquid muriatic acid to the compound known by the name of butter of antimony, (chloride of antimony,) or sal ammoniac to the mercurial solution, he succeeded completely in obtaining a spectrum in which the rays followed the same law of dispersion as in crown glass, and even in *over-correcting* the secondary spectrum, so as to place its exact destruction completely in his power. It only remained to form an object-glass on these principles. Fig. 111 is such an one, in which, though there are two refractions at the confines of the glass and fluid, yet the chromatic aberration, as Dr. Blair assures us, was totally destroyed, and the rays of different colours were bent from their rectilinear course with the same equality as in reflexion.

Dr. Blair's discovery of media having the same scale of dispersion as glass.

His double object-glasses formed of such media

487.

To such an extent has Dr. Blair carried these interesting experiments, that he assures us he has found it practicable to construct an object-glass of nine inches focal length, capable of bearing an aperture of three inches, a thing which assuredly no artist would ever dream of attempting with glass lenses; and we cannot close this account of his labours without joining in a wish expressed on a similar occasion by Dr. Brewster, whose researches on dispersive powers have so worthily filled up the outline sketched by his predecessor, that this branch of practical optics may be resumed with the attention it deserves, by artists who have the ready means of executing the experiments it would require. Could solid media of such properties be discovered, the telescope would become a new instrument.

Part II.

477.

These experiments of Dr. Blair lead to the remarkable conclusion, that at the common surface of two media a white ray may be refracted without separation into its coloured elements. In fact, μ and μ' being the refrac-

478.

Case of colourless single refraction at common surface of two media.

tive indices of the media for any ray as the extreme red, $\frac{\mu'}{\mu}$ will be their relative refractive index for that ray,

and $\frac{\mu' + \delta \mu'}{\mu + \delta \mu}$ will be the relative index for any other ray. If, then, the refractive and dispersive powers of

the media be such that $\frac{\mu' + \delta \mu'}{\mu + \delta \mu} = \frac{\mu'}{\mu}$, or $\mu \delta \mu' = \mu' \delta \mu$, that is, if $\frac{\delta \mu}{\mu} = \frac{\delta \mu'}{\mu'}$; and if, moreover, this

relation hold good throughout the spectrum, *i. e.* if the increments of the refractive indices, in proceeding from the red to the violet end of the spectrum, be proportional to the refractive indices themselves, then the relative index is the same for all rays, and no dispersion will take place. Now this gives a relation between the disper-

sive and refractive indices of the two media, viz. $\frac{p'}{p} = \frac{\mu'}{\mu} \cdot \frac{\mu - 1}{\mu' - 1} = \frac{1 - \frac{1}{\mu}}{1 - \frac{1}{\mu'}}$; and, in addition to this

condition, the scale of dispersions must be the same in both media. According as the dispersions differ one way or the other from this precise adjustment, the violet ray may be either more or less refracted than the red at the common surface of the two media.

We shall terminate the theory of achromatic object-glasses with a problem of considerable practical importance, as it puts it in our power, having obtained an approximate degree of achromaticity in an object-glass, to complete the destruction of the colour without making any alteration in the focal lengths or curvatures of the lenses, by merely placing them at a greater or less distance from one another.

479.

Achromatic object-glass with separated lenses.

Problem. To express the condition of achromaticity, when the two lenses of a double object-glass are placed at a distance from each other, ($= t$.)

Resuming the notation of Art. 251 and 268, we have

$$f'' = L' + D; \quad f^{iv} = L'' + \frac{f''}{1 - f''t}; \quad \delta f'' = \delta L';$$

and

$$\delta f^{iv} = \delta L'' + \frac{\delta f''}{(1 - f''t)^2} = \delta L'' + \frac{\delta L'}{\{1 - t(L' + D)\}^2}.$$

Now, that the combination may be achromatic, we must have $\delta f^{iv} = 0$; and, since t and D are constant, and L' and L'' only vary by the variations of μ' , μ'' the refractive indices, we have $\delta L' = (R' - R'') \delta \mu' =$

$\frac{\delta \mu'}{\mu' - 1} L' = p' L'$, and similarly $\delta L'' = p'' L''$, so that substituting we get

$$\{1 - t(L' + D)\}^2 + \frac{p'}{p''} \cdot \frac{L'}{L''} = 0.$$

Such is the condition of achromaticity. Since it depends on D , it appears that if the lenses of an object-glass be not close together, it will cease to be achromatic for near objects, however perfectly the colour be corrected for distant ones. The eye therefore cannot be achromatic for objects at all distances, its lenses being of great thickness compared to their focal lengths; and, therefore, although in contact at their adjacent surfaces, yet having considerable intervals between others.

480.

For parallel rays the equation becomes

481.

$$p'' L'' (1 - t L')^2 = -p' L';$$

hence, the dispersions and powers of the lenses being given their interval t may be found by the expression

$$t = \frac{1}{L'} \left\{ 1 - \sqrt{-\frac{p'}{p''} \cdot \frac{L'}{L''}} \right\}.$$

The condition of achromaticity, were the lenses placed close together, would be, as we have already shown,

482.

$\underbrace{\text{Light.}} - \frac{p'}{p''} \cdot \frac{L'}{L''} = 1$. Hence, whenever this fraction is less than unity, that is whenever L'' , the power of the

concave or flint lens (which we here suppose to be the second) is *too great*; or when, as the opticians call it, the colour is over-corrected, the object-glass may be made achromatic, or their over-correction remedied, without re-grinding the glasses, merely by separating the lenses; for in this case the quantity under the radical is less than unity, and therefore t is positive, a condition without which the rays could not be refracted as we have supposed them.

483. Moreover, this affords a practical and very easy means of ascertaining, with the greatest precision, the dispersive ratio of the two media. Let a convex lens of crown be purposely a little over-corrected by a concave of flint, and then let the colour be destroyed by separating the lenses. Measure their focal lengths $\left(\frac{1}{L'}$ and $\frac{1}{L''}\right)$ and the interval t between them in this state, and we have at once for the value of ω the dispersive ratio,

$$\omega = \frac{p'}{p''} = - \frac{L''}{L'} (1 - t L')^2.$$

§ III. Of the Absorption or Extinction of Light by uncrystallized Media.

484. All media absorb light. Transparency is the quality by which media allow rays of light freely to pass through their substance, or, it may be, between their molecules; and is said to be more or less perfect, according as a more or less considerable part of the whole light which enters them finds its way through. Among media, consisting of ponderable matter, we know of none whose transparency is perfect. Whether it be that some of the rays in their passage encounter bodily the molecules of the media, and are thereby reflected; or, if this supposition be thought too coarse and unrefined for the present state of science, be stopped or turned aside by the forces which reside in the ultimate atoms of bodies, without actual encounter, or otherwise detained or neutralized by them; certain it is, that even in the most rare and transparent media, such as air, water, and glass, a beam of light intromitted, is gradually extinguished, and becomes more and more feeble as it penetrates to a greater depth within them, and ultimately becomes too faint to affect our organs. Thus, at the tops of very high mountains, a much greater multitude of stars is visible to the naked eye than on the plains at their feet; the weak light of the smallest of them being too much reduced in its passage through the lower atmospheric strata to affect the sight. Thus, too, objects cease to be visible at great depths below water, however free from visible impurities, &c. Dr. Olbers has even supposed the same to hold good with the imponderable media (if any) of the celestial spaces, and conceives this to be the cause why *so few* stars (not more than about five or ten millions) can be seen with the most powerful telescopes. It is probable that we shall be long without means of confirming or refuting this singular doctrine.

485. On the other hand, though no body in nature be perfectly, all are to a certain degree, transparent. One of the densest of metals, gold, may actually be beaten so thin as to allow light to pass through it; and that it passes through the substance of the metal, not through cracks or holes too small to be detected by the eye, is evident from the colour of the transmitted light, which is green, even when the incident light is white. The most opaque of bodies, charcoal, in a different state of aggregation, (as diamond,) is one of the most perfectly transparent; and all coloured bodies, however deep their hues, and however seemingly opaque, must necessarily be rendered visible by rays which have entered their substance; for if reflected at their surfaces, they would all appear white alike. Were the colours of bodies strictly superficial, no variation in their thickness could affect their hue; but, so far is this from being the case, that all coloured bodies, however intense their tint, become paler by diminution of thickness. Thus the powders of all coloured bodies, or the streak they leave when rubbed on substances harder than themselves, have much paler colours than the same bodies in mass.

486. And all absorb the different colours unequally. This gradual diminution in the intensity of a transmitted ray in its progress through imperfectly transparent media, is termed its *absorption*. It is never found to affect equally rays of all colours, some being always absorbed in preference to others; and it is on this preference that the colours of all such media, as seen by transmitted light, depend. A white ray transmitted through a perfectly transparent medium, ought to contain at its emergence the same proportional quantity of all the coloured rays, because the part reflected at its anterior and posterior surfaces is colourless; but, in point of fact, such perfect want of colour in the transmitted beam is never observed. Media, then, are unequally transparent for the differently coloured rays. Each ray of the spectrum has, for every different medium in nature, its own peculiar *index of transparency*, just as the index of refraction differs for different rays and different media.

487. Experiment. The most striking way in which this different absorptive power of one and the same medium on differently coloured rays can be exhibited, is to look through a plain and polished piece of smalt-blue glass, (a rich deep blue, very common in the arts—such as sugar-basins, finger-glasses, &c. are often made of,) at the image of any narrow line of light (as the crack in a window-shutter of a darkened room) refracted through a prism whose edge is parallel to the line, and placed in its situation of minimum deviation. If the glass be extremely thin, all the colours are seen; but if of moderate thickness (as $\frac{1}{10}$ inch) the spectrum will put on a very singular and striking appearance. It will appear composed of several detached portions separated by broad and perfectly black

Light.

Part II.

intervals, the rays which correspond to those points in the perfect spectrum being entirely extinguished. If a less thickness be employed, the intervals, instead of being perfectly dark, are feebly and irregularly illuminated, some parts of them being less enfeebled than others. If the thickness, on the other hand, be increased, the black spaces become broader, till at length all the colours intermediate between the extreme red and extreme violet are totally destroyed.

The simplest hypothesis we can form of the extinction of a beam of homogeneous light in passing through a homogeneous medium, is, that for every equal thickness of the medium passed through, an equal aliquot part of the rays, which, up to that depth had escaped absorption, is extinguished. Thus, if 1000 red rays fall on and enter into a certain green glass, and if 100 be extinguished in traversing the first tenth of an inch, there will remain 900 which have penetrated so far; and of these one-tenth, or 90, will be extinguished in the next tenth of an inch, leaving 810, out of which again a tenth, or 81, will be extinguished in traversing the third tenth, leaving 729, and so on. In other words, the quantity unabsorbed, after the beam has traversed any thickness of the medium, will diminish in geometrical progression, as t increases in arithmetical. So that if 1 be taken for the whole number of intromitted rays, and y for the number that escape absorption in traversing an unit of thickness, y^t will represent the number escaping, after traversing any other thickness, $= t$. This only supposes that the rays in the act of traversing one stratum of a medium acquire no additional facility to penetrate the remainder. In this doctrine, y is necessarily a fraction smaller than unity, and depending on the nature both of the ray and the medium. Hence, if C represent the number of equally illuminating rays of the extreme red in a beam of white light, C' that of the next degree of refrangibility, and so on; the beam of white light will be represented by $C + C' + C'' + \&c.$; and the transmitted beam, after traversing the thickness t , will be properly expressed by

$$C \cdot y^t + C' \cdot y'^t + C'' \cdot y''^t + \&c.$$

Each term representing the intensity of the particular ray to which it corresponds, or its ratio to what it is in the original white beam.

It is evident from this, that, strictly speaking, total extinction can never take place by any finite thickness of the medium; but if the fraction y for any ray be at all small, a moderate increase in the thickness, (which enters as an exponent,) will reduce the fraction y^t to a quantity perfectly insensible. Thus, in the case taken above, where a tenth of an inch of green glass destroys one-tenth only of the red rays, a whole inch will allow to pass

only $\left(\frac{9}{10}\right)^{10}$, or 304 rays out of a thousand, while ten times that thickness, or 10 inches, will suffer only $\left(\frac{9}{10}\right)^{100} = 0.0000266$, or less than three rays out of 100,000 to pass, which amounts to almost absolute opacity.

If x be the index of refraction of any ray in the water spectrum, we may regard y as a function of x ; and if on the line RV , (fig. 112,) representing the whole length of the water spectrum, we erect ordinates, Rr , MN , VV equal to unity and to each other; and also other ordinates Rr , MP , Vv representing the values of y for the rays at the corresponding points; the curve rPv , the locus of P , will be, as it were, a type, or geometrical picture of the action of the medium on the spectrum, and the straight line RNV will be a similar type of a perfectly transparent medium. Now if this be supposed the case when the thickness of the medium is 1, if we take always $MP' : MP :: MP : MN$, and $MP'' : MP' :: MP' : MP$, &c. and so on, the loci of $P'P''$, &c. will be curves representing the quantities of the rays transmitted by the thicknesses 2, 3, &c. of the medium, and so for intermediate thicknesses, or for a thickness less than 1, as in the curve $\rho\pi v$.

Hence, whatever be the colour of a medium, if its thickness be infinitely diminished, it will transmit all the rays indifferently; for when $t = 0$, $y^t = 1$, whatever be y ; and the curve $\rho\pi v$ approaches infinitely near to the line RNV . Thus all coloured glasses blown into excessively thin bubbles are colourless, and so is the foam of coloured liquids.

Again, if there be any, the least, preference given by the medium to the transmission of certain rays beyond others, the thickness of the medium may be so far increased as to give it any assignable depth of tint; for if y be ever so little less than unity, and if between the values of y for different rays there be ever so little difference, t may be so increased as to make y^t as small as we please, and the ratio of y^t to y'^t as different from unity as we please.

In very deep coloured media all the values of y are small. If they were equal, the medium would merely stop light, without colouring the transmitted beam, but no such media are at present known.

If the curve rPv , or the type of an absorbent medium have a maximum in any part of the spectrum, as in the green, for instance, (fig. 113;) then, whatever be the proportion in which the other rays enter, by a sufficient increase of thickness, that colour will be rendered predominant; and the ultimate tint of the medium, or the last ray it is capable of transmitting, will be a pure homogeneous light of that particular refrangibility to which the maximum ordinate corresponds. Thus green glasses, by an increase of thickness, become greener and greener, their type being as in fig. 113; while yellow ones, whose type is as in fig. 114, change their tint by reduplication, and pass through brown to red.

This change of tint by increase of thickness is no uncommon phenomenon; and though at first sight paradoxical, yet is a necessary consequence of the doctrine here laid down. If we enclose a pretty strong solution of sap-green, or, still better, of muriate of chromium in a thin hollow glass wedge, and if we look through the edge where it is thinnest, at white paper, or at the white light of the clouds, it appears of a fine green; but if we slide the wedge before the eye gradually so as to look successively through a greater and greater thickness

488.
Law of transmission

489.

490.
Law of absorption of any medium expressed by a curve. Fig. 112.

491.

492.

493.

494.
Ultimate tint of an absorptive medium. Fig. 113.

495.

Tint changes by increase of thickness.

Light.
Case of a
green-red
medium.
Fig. 115.

Numerical
illustration.

of the liquid, the green tint grows livid, and passes through a sort of neutral, brownish hue, to a deep blood-red. To understand this, we must observe, that the curves expressing the types of different absorbent media admit the most capricious variety of form, and very frequently have several maxima and minima corresponding to as many different colours. The green liquids in question have two distinct maxima, as in fig. 115; the one corresponding to the extreme red, the other to the green, but the absolute lengths of the maximum ordinates are unequal, the red being the greater. But as the extreme red is a very feebly illuminating ray, while on the other hand the green is vivid, and affects the eye powerfully, the latter at first predominates over the former, and entirely prevents its becoming sensible; and it is not till the thickness is so far increased as to leave a very great preponderance of those obscure red rays, and subdue their rivals, as in the case represented by the lowest of the dotted curves in the figure, that we become sensible of their influence on the tint. Suppose, for instance, to illustrate this by a numerical example, the index of transparency, or value of y , in muriate of chromium, to be for extreme red rays, 0.9; for the mean red, orange, and yellow, 0.1; for green, 0.5; and for blue, indigo, and violet, 0.1 each; and suppose, moreover, in a beam of white light, consisting of 10,000 rays, *all equally illuminative*, the proportions corresponding to the different colours to be as follows:

Extreme red.	Red and orange.	Yellow.	Green.	Blue.	Indigo.	Violet.
200	1300	3000	2800	1200	1000	500.

Then, after passing through a thickness equal to 1 of the medium, the proportions in the transmitted beam would be

Extreme red.	Red and orange.	Yellow.	Green.	Blue.	Indigo.	Violet.
180	130	300	1400	120	100	50.

After traversing a second unit of thickness, they would be

Extreme red.	Red and orange.	Yellow.	Green.	Blue.	Indigo.	Violet.
162	13	30	700	12	10	5.

and after a third, a fourth, a fifth, and sixth respectively,

Extreme red.	Red and orange.	Yellow.	Green.	Blue.	Indigo.	Violet.
146	1	3	350	1	1	0
131	0	0	175	0	0	0
118	0	0	87	0	0	0
106	0	0	43	0	0	0.

Thus we see, that in the first of these transmitted beams the green greatly preponderates, after the second transmission, it is still the distinguishing colour; but after the third, the red bears a proportion to it large enough to impair materially the purity of its tint. The fourth transmission may be regarded as totally extinguishing all the other colours, and leaving a neutral tint between red and green; while, in all the tints produced by further successive transmissions, the red preponderates continually more and more, till at length the tint becomes no way distinguishable from the homogeneous red of the extremity of the spectrum.

496. Whether we suppose the obscurer parts of the spectrum to consist of fewer rays equally illuminative, or of the same number of rays of less intrinsic illuminating power with the brighter, obviously makes no difference in the conclusion, but the former supposition has the advantage of affording a hold to numerical estimation which the latter does not. In the instance here taken, the numbers are assumed at random. But Fraunhofer has made a series of experiments expressly to determine numerically the illuminating power of the different rays of the spectrum. According to which, he has constructed the curve fig. 116, whose ordinate represents the illuminative power of the ray in that part of the spectrum on which it is supposed erected, or the proportional number of equally illuminative rays of that refrangibility in white light. If we would take this into consideration in our geometrical construction, we must suppose the type of white light, instead of being a straight line, as in fig. 112. . . 114, to be a curve similar to fig. 116, and the other derivative curves to be derived from it by the same rules as above. But as the only use of such representations is to express concisely to the eye the general scale of action of a medium on the spectrum, this is rather a disadvantageous than a useful refinement.

497. To take another instance. If we examine various thicknesses of the small-blue glass above noticed, it will be found to appear purely blue in small thicknesses. As the thickness increases, a purple tinge comes on, which becomes more and more ruddy, and finally passes to a deep red; a great thickness being, however, required to produce this effect. If we examine the tints by a prism, we shall find the type of this medium to be as in fig. 117, having four maximum ordinates, the greatest corresponding to a ray at the very farthest extremity of the red, and diminishing with such rapidity as to cause an almost perfect insulation of this ray; the next corresponds to a red of mean refrangibility, the next to the mean yellow, and the last to the violet, the ordinate increasing continually to the end of the spectrum. Thus, when a piece of such glass of the thickness 0.042 inch was used, the red portion of the spectrum was separated into two, the least refracted being a well defined band of perfectly homogeneous and purely red light, separated from the other red by a band of considerable breadth, and totally black. This red was nearly homogeneous; its tint, however, differing in no respect from the former, and being free from the slightest shade of orange. Its most refracted limit came very nearly up to the dark line D in the spectrum. A small, sharp, black line separated this red from the yellow, which was a pretty well defined band of great brilliancy and purity of colour, of a breadth exceeding that of the first red, and bounded on the

Relative il-
luminative
power of
the several
prismatic
rays.
Fig. 116.

Fig. 117.

Light

green side by an obscure but not quite black interval. The green was dull and ill defined, but the violet was transmitted with very little loss. A double thickness (0.084 inch) obliterated the second red, greatly enfeebled the yellow, leaving it now sharply divided from the green, which was also extremely enfeebled. The extreme red, however, retained nearly its whole light, and the violet was very little weakened. When a great many thicknesses were laid together, the extreme red and extreme violet only passed.

Part II.

Among transparent media of most ordinary occurrence, we may distinguish, first, those whose type has its ordinate decreasing regularly, with more or less rapidly from the red to the violet end of the spectrum, or which absorb the rays with an energy more or less nearly in some direct ratio of their refrangibility. In red and scarlet media the absorbent power increases very rapidly, as we proceed from the red to the violet. In yellow, orange, and brown ones, less so; but all of them act with great energy on the violet rays, and produce a total obliteration of them. In consequence of this, by an increase of thickness, all these media finally become red. Examples: red, scarlet, brown, and yellow glasses; port wine, infusion of saffron, permuriate of iron, muriate of gold, brandy, India soy, &c.

498.
Red media.

Among green media, the generality have a single maximum of transmission corresponding to some part of the green rays, and their hue in consequence becomes more purely green by increase of thickness. Of this kind are green glasses, green solutions of copper, nickel, &c. They absorb both ends of the spectrum with great energy; the red, however, more so, if the tint verges to blue; the violet, if to yellow. Besides these, however, are to be remarked media in which the type has two maxima; such may be termed dichromatic, having really two distinct colours. In most of these, the green maximum is less than the red; and the green tint, in consequence, loses purity by increase of thickness, and passes through a livid neutral hue to red, though this is not always the case. Examples: muriate of chrome, solution of sap-green, manganate of potash, alkaline infusion of the petals of the peonia officinalis and many other red flowers, and mixtures of red and blue or green media.

499.
Simple
green media.
Dichromatic
media.

Blue media admit of great variety, and are generally dichromatic, having two or even a great many maxima and minima in their types; but their distinguishing character is a powerful absorption of the more luminous red rays and the green, and a feeble action on the more refrangible part of the spectrum. Among those whose energy of absorption appears to increase regularly and rapidly from the violet to the red end of the spectrum, we may place the blue solutions of copper. The best example is the magnificent blue liquid formed by saturating sulphate of copper with carbonate of ammonia. The extreme violet ray seems capable of passing through almost any thickness of this medium; and this property, joined to the unalterable nature of the solution, and the facility of its preparation, render it of great value in optical researches. A vessel, or tube, of some inches in length, closed at two ends with glass plates, and filled with this liquid, is the best resource for experiments on the violet rays. Ammonio-oxalate of nickel transmits the *blue* and *extreme red*, but stops the *violet*.

500.
Blue media.
Insulation of
the extreme
violet.

Purple media act by absorbing the middle of the spectrum, and are therefore necessarily always dichromatic, some of them having red and others violet for their *ultimate* or *terminal tint*. Example: solution of archil; purple, plum-coloured, and crimson glasses; acid and alkaline solutions of cobalt, &c. They may be termed red-purple and violet-purple, according to their terminal tint.

501.
Purple
media.

In combinations of media, the ray finally transmitted is the residuum of the action of each. If x, y, z be the indices of transmissibility of a given ray C in the spectrum for the several media, and r, s, t their thicknesses, the transmitted portion of this ray will be $C \cdot x^r y^s z^t$; and the residuum of a beam of white light (supposing none lost by reflexion at the surfaces) after undergoing the absorptive action of all the media, will be

502.
Combina-
tions of
media.

$$C \cdot x^r y^s z^t + C' \cdot x'^r y'^s z'^t + \&c.$$

An expression which shows that it is indifferent in what order the media are placed. They may therefore be mixed, unless a chemical action take place. Thus also, by the same construction as that by which the type 1 of the first medium is derived from the straight line representing white light, may another type 2 be derived from 1, and so on; and thus an endless variety of types will originate, having so many tints corresponding to them.

This circumstance enables us to insulate, in a state of considerable homogeneity, various rays. Thus, by combining with the small-blue glass, already mentioned, any brown or red glass of tolerable fulness and purity of colour, a combination will be formed absolutely impermeable to any but the extreme red ray, and the refrangibility of this is so strictly definite as to allow of its being used as a standard ray in all optical inquiries, which is the more valuable, as the coloured glasses by which it is insulated are the most common of any which occur in the shops, and may be had at any glazier's. If to such a combination a green glass be added, a total stoppage of all light takes place. The same kind of glass, too, enables us to insulate the yellow ray, corresponding to the maximum Y in the type fig. 117, by combining it with a brown glass to stop out the more, and a green to destroy the less, refrangible rays, and by their means the existence of a considerable breadth of yellow light, evidently not depending on a *mixture*, or mutual encroachment of red and green, may be exhibited in the solar spectrum.

503.
Insulation
of an ex-
treme homo-
geneous red
ray.

Insulation
of the
yellow rays.

It has been found by Dr. Brewster, that the proportions of the different coloured rays absorbed by media depend on their temperature. The tints of bodies generally deepen by the application of heat, as is known to all who are familiar with the use of the blow-pipe; thus minium and red oxide of mercury deepen in their hues by heat till they become almost black, but recover their red colours on cooling. Dr. Brewster has, however, produced instances, not merely among artificial glasses, but among transparent minerals, where a transition takes place from red to green on the application of a high temperature; the original tint being, however, restored on cooling, and no chemical alteration having been produced in the medium.

504.
Alteration of
absorptive
power by
heat.

The analysis of the spectrum by coloured media presents several circumstances worthy of remark. First, the irregular and singular distribution in the dark bands which cross the spectrum, when viewed through such

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Light.

media as have several maxima of transmission, obviously leads us to refer Fraunhofer's *Fixed lines*, and the analogous phenomena to be noticed in the light from other sources, to the same cause, whatever it may be, which determines the absorption of some ray in preference to others. It is no impossible supposition, that the deficient rays in the light of the sun and stars may be absorbed in passing through their own atmospheres, or, to approach still nearer to the origin of the light, we may conceive a ray stifled in the very act of emanation from a luminous molecule by an intense absorbent power residing in the molecule itself; or, in a word, the same indisposition in the molecules of an absorbent body to permit the propagation of any particular coloured ray through, or near them, may constitute an obstacle *in limine* to the production of the ray from them. At all events, the phenomena are obviously related, though we may not yet be able to trace the particular nature of their connection.

506.

The next circumstance to be observed is, that when examined through absorbent media all idea of regular gradation of colour from one end to the other of the spectrum is destroyed. Rays of widely different refrangibility, as the two reds noticed in Art. 497, have absolutely the same colour, and cannot be distinguished. On the other hand, the transition from pure red to pure yellow, in the case there described, is quite sudden, and the contrast of colours most striking, while the dark interval which separates them, by properly adjusting the thickness of the glass, may be rendered very small without any tinge of orange becoming perceptible. What then, we may ask, is become of the orange; and how is it, that its place is partly supplied with red on one side, and yellow on the other? These phenomena certainly lead us very strongly to believe that the analysis of white light by the prism is not the only analysis of which it admits, and that the connection between the refrangibility and colour of a ray is not so absolute as Newton supposed. Colour is a sensation excited by the rays of light, and since two rays of different refrangibilities are found to excite absolutely the same sensation of colour, there is no *primâ facie* absurdity in supposing the converse,—that two rays capable of exciting sensations of different colours may have identical indices of refraction. It is evident, that if this be the case, no mere change of direction by refractions through prisms, &c. could ever separate them; but should they be differently absorbable by a medium through which they pass, an analysis of the compound ray would take place by the destruction of one of its parts. This idea has been advocated by Dr. Brewster, in a Paper published in the *Edinburgh Philosophical Transactions*, vol. ix., and the same consequence appears to follow from other experiments, published in the same volume of that collection. According to this doctrine, the spectrum would consist of at least three distinct spectra of different colours, red, yellow, and blue, over-lapping each other, and each having a maximum of intensity at those points where the compound spectrum has the strongest and brightest tint of that colour.

507.

Cases of persons who see only two colours.

It must be confessed, however, that this doctrine is not without its objections; one of the most formidable of which may be drawn from the curious affection of vision occasionally (and not very rarely) met with in certain individuals, who distinguish only two colours, which (when carefully questioned and examined by presenting to them, not the ordinary compound colours of painters, but optical tints of known composition) are generally found to be yellow and blue. We have examined with some attention a very eminent *optician*, whose eyes (or rather eye, having lost the sight of one by an accident) have this curious peculiarity, and have satisfied ourselves, contrary to the received opinion, that all the prismatic rays have the power of exciting and affecting them with the sensation of *light*, and producing distinct vision, so that the defect arises from no insensibility of the retina to rays of any particular *refrangibility*, nor to any colouring matter in the humours of the eye, preventing certain rays from reaching the retina, (as has been ingeniously supposed,) but from a defect in the sensorium, by which it is rendered incapable of appreciating exactly those differences between rays on which their colour depends. The following is the result of a series of trials, in which a succession of optical tints produced by polarized light, passing through an inclined plate of mica, in a manner hereafter to be described, was submitted to his judgment. In each case, two uniformly coloured circular spaces placed side by side, and having *complementary tints* (i. e. such that the sum of their light shall be white) were presented, and the result of his judgment is here given in his own words.

Light.

Part II.

Colours according to the judgment of an ordinary eye.		Colours as named by the individual in question.		Inclination of the plate of mica to eye
Circle to the left.	Circle to the right.	Circle to the left.	Circle to the right.	
Pale green.	Pale pink.	Both alike, no more colour sky out of window.	in them than in the cloudy	89.5
Dirty white.	Ditto, both alike.	Both darker than before, but	no colour.	85.0
Fine bright pink.	Fine green, a little verging on bluish.	Very pale tinge of blue.	Very pale tinge of blue.	81.1
White.	White.	Yellow.	Blue.	76.3
The limit of Rich grass green.	pink and red. Rich crimson.	Yellow. Both more coloured	than before Blue.	74.9
Dull greenish blue.	Pale brick red.	Blue. Better, but neither	full colours.	72.8
Purple (rather pale.)	Pale yellow.	Blue. Neither so rich	Yellow, colours as the last.	71.7
Fine pink.	Fine green.	Coming up to good colours, than a gilt picture-frame.	Yellow.	
Fine yellow.	Purple.	Yellow, but has got a good deal of blue in it.	Blue, but has a good deal of yellow in it.	69.7
Yellowish green.	Fine crimson.	Good yellow.	Good blue.	68.2
Good blue, verging to indigo.	Yellow, verging to orange.	Better colours than	any yet seen.	
Red, or very ruddy pink.	Very pale greenish blue, almost white.	Yellow, but has a good deal of blue.	Blue, but has a good deal of yellow.	67.0
Rich yellow.	Full blue.	Blue.	Yellow.	65.5
White.	Fiery orange.	Both gay colours, particularly	the yellow to the right.	63.8
Dark purple.	White.	Yellow.	Blue.	
Dull orange red.	White.	Fine bright yellow.	Pretty good blue.	62.7
White.	Dull dirty olive.	Has very little colour.	Yellow, but a different yellow, it is a blood-looking yellow.	61.2
Very dark purple.	White.	A dim blue, wants light.	White, with a dash of yellow and blue.	59.5
		Yellow	White, with blue and yellow in it.	59.0
		White.	Dark.	57.1
		Dark.	White.	55.0

Instead of presenting the colours for his judgment, he was now desired to arrange the apparatus so as to make the strongest possible succession of contrasts of colour in the two circles. The results were as follow :

508.

Colours according to the judgment of an ordinary eye.		Colours as named by the individual in question.		Inclination of the plate of mica to eye.
Circle to the left.	Circle to the right.	Circle to the left.	Circle to the right.	
Pale ruddy pink.	Blue green.	Yellow.	Blue.	59.1°
Blue green.	Pale ruddy pink.	Blue.	Yellow.	65.3
Yellow.	Blue.	Yellow.	Blue.	63.1
White.	Fiery orange.	Blue.	Yellow.	61.1
Pale brick-red.	White.	Yellow.	Blue.	58.5
Indigo.	Pale yellow.	Blue.	Yellow.	54.2
Yellow.	Indigo.	Yellow.	Blue.	52.1

It appears by this, that the eyes of the individual in question are only capable of fully appreciating blue and yellow tints, and that these names uniformly correspond, in his nomenclature, to the more and less refrangible rays, generally; all which belong to the former, indifferently, exciting a sense of "blueness," and to the latter of "yellowness." Mention has been made of individuals seeing well in other respects, but devoid altogether of the sense of colour, distinguishing different tints only as brighter or darker one than another; but the case is, probably, one of extremely rare occurrence.

Mayer, in an Essay *De Affinitate Colorum*, (Opera inedita, 1775,) regards all colours as arising from three primary ones, red, yellow, and blue; regarding white as a neutral mixture of rays of all colours, and black as a mere negation of light. According to this idea, were we acquainted with any mode of mixing colours in simple numerical ratios, a scale might be formed to which any proposed colour might be at once referred. He proposes to establish such a scale in which the degrees of intensity of each simple colour shall be represented by the natural numbers 1, 2, 3... 12; 1 denoting the lowest degree of it capable of sensibly affecting a tint, and 12 the full intensity of which the colour is capable, or the total amount of it existing in white light. Thus r^{12} denotes a full red of the brightest and purest tint, y^{12} the brightest yellow, and b^{12} the brightest blue. To represent mixed tints, he combines the symbols of the separate ingredients. Thus $r^{12}y^4$, or, more conveniently, $12r + 4y$, represents a red verging strongly to orange, such as that of a coal fire.

The scale proposed is convenient and complete, so far as regards what he calls perfect colours, which arise from white light by the subtraction of one or more proportions of its elementary rays; but a very slight modification

509.

Mayer's hypothesis of three primary colours.

510

Modification of Mayer's scale.

Light.

cation of his system will render it equally applicable to all, and it may be presented as follows. Suppose we fix on 100 as a standard intensity of each primary colour; or the number of rays of that colour (all supposed equally effective) which falling on a sheet of white paper, or other surface perfectly neutral, (*i. e.* equally disposed to reflect all rays) shall produce a full tint of that particular kind, and let us denote by such an expression as $xR + yY + zB$, the tint produced by the incidence of x such rays of primary red, y such rays of yellow, and z such rays of blue on the same surface together. It is obvious then, that the different numerical values assigned to x, y, z , from 1 to 100, will give different symbols of tints, whose number will be $100 \times 100 \times 100 = 1000000$, and therefore quite sufficient in point of extent to embrace all the variety of colours the eye can distinguish. The number of tints recognised as distinct by the Roman artists in Mosaic is said to exceed 30,000; but if we suppose ten times this amount to occur in nature (and it is obvious that these must be greatly more numerous than the purposes of the painter admit) we are still much within the limits of our scale. It only remains to examine how far the tints themselves are expressible by the members of the scale proposed.

511.
Whites,
greys, and
neutral
tints.

And first, then, of whites, greys, and neutral tints. The most perfectly neutral tints, which are, in fact, only greater and less intensities of whiteness, are those we observe in the clouds in an ordinary cloudy day, with occasional gleams of sunshine. From the most sombre shadows to the snowy whiteness of those cumulus-shaped clouds on which the sun immediately shines, we have nothing but a series of whites, or greys, represented by such combinations as $R + Y + B$, $2R + 2Y + 2B$, &c.; or $n(R + Y + B)$ which, for brevity, we may represent by nW . To be satisfied of this we need only look through a tube blackened on the inside to prevent surrounding objects influencing our judgments; and any small portion thus insulated of the darkest clouds will appear to differ in no respect from a portion similarly insulated of a sheet of white paper more or less shaded.

512.
Reds, yellows, and blues.

The various intensities of pure reds, yellows, and blues are represented by nR , nY , and nB respectively. They are rare in nature; but blood, fresh gilding, or gamboge moistened, and ultramarine may be cited as examples of them. Scarlets and vivid reds, such as vermilion and minium, are not free from a mixture of yellow, and even of blue; for all the primary colours are greatly increased in splendour by a certain mixture of white, and whenever any primary colour is peculiarly glaring and vivid, we may be sure that it is in some degree diluted with white. The blue of the sky is white, with a very moderate addition of blue.

513.

The mixture of red and yellow produces all the shades of scarlet, orange, and the deeper browns, when of feeble intensity. When diluted with white, we have lemon colour, straw colour, clay colour, and all the brighter browns; the last-mentioned tints growing dusker and dingier as the coefficients are smaller.

514.
Browns.

The browns, however, are essentially sombre tints, and produce their effects chiefly by contrast with other brighter hues in their neighbourhood. To produce a brown, the painter mixes black and yellow, or black and red, (that is, such impure reds as the generality of red pigments,) or all three; his object is to stifle light, and leave only a residuum of colour. There is a brown glass very common in modern ornamental windows. If examined with a prism, it is found to transmit the red, orange, and yellow rays abundantly, little green, and no pure blue. The small quantity of blue, then, that its tint does involve, must be that which enters as a component part of its green, (in this view of the composition of colours,) and its characteristic symbol may thus be, perhaps, of some such form as $10R + 9Y + 1B$; that is to say, $(9R + 8Y) + 1(R + Y + B)$, or an orange of the character $9R + 8Y$ diluted with one ray of white. It must be confessed, however, that the composition of brown tints is the least satisfactory of all the applications of Mayer's doctrine. He himself has passed it unnoticed.

515.
Purples.

Combinations of red and blue, and their dilutions with white, form all the varieties of crimson, purple, violet, rose colour, pink, &c. The richer purples are entirely free from yellow. The prismatic violet, when compared with the indigo, produces a sensible impression of redness, and must therefore be regarded on this hypothesis as consisting of a mixture of blue and red rays.

516.
Greens.

Blue and yellow, combined, produce green. The green thus arising is vivid and rich; and, when proper proportions of the elementary colours are used, no way to be distinguished from the prismatic green. Nothing can be more striking, and even surprising, than the effect of mixing together a blue and a yellow powder, or of covering a paper with blue and yellow lines, drawn close together, and alternating with each other. The elementary tints totally disappear, and cannot even be recalled by the imagination. One of the most marked facts in favour of the idea of the existence of three primary colours, and of the possibility of an analysis of white light distinct from that afforded by the prism, is to see the prismatic green thus completely imitated by a mixture of adjacent rays totally distinct from it, both in refrangibility and colour.

517.
The same colour produced by different prismatic combinations.

The hypothesis of three primary colours, of which, in different proportions, all the colours of the spectrum are composed, affords an easy explanation of a phenomenon observed by Newton, *viz.* that tints no way distinguishable from each other may be compounded by very different mixtures of the seven colours into which he divided it. Thus we may regard white light, indifferently, as composed of

$$\left. \begin{aligned} R &= a + b + c \text{ rays of pure red} \\ Y &= d + e + f + g \text{ rays of pure yellow} \\ B &= h + i + k + l \text{ rays of pure blue} \end{aligned} \right\} \text{ or of } \left\{ \begin{aligned} b \text{ rays of pure red} &= R' \\ c + d \text{ rays of orange (c red + d yellow)} &= O' \\ e \text{ rays of pure yellow} &= Y' \\ f + h \text{ rays of green (f yellow + h blue)} &= G' \\ g + i \text{ rays of prismatic blue (g yellow + i blue)} &= B \\ k \text{ rays of indigo, or pure blue} &= I' \\ l + a \text{ rays of violet (l blue + a red)} &= V' \end{aligned} \right.$$

Light. and any tint capable of being represented by $x \cdot R + y \cdot Y + z \cdot B$, may be represented equally well by

$$m \cdot R' + n \cdot O' + p \cdot Y' + q \cdot G' + r \cdot B' + s \cdot I' + t \cdot V',$$

provided we assume m, n, p , &c., such as to satisfy the equations

$$m b + n c + t a = x; \quad n d + p e + q f + r g = y; \quad q h + r i + s k + t l = z.$$

From what has been said we shall now proceed to show, that, without departing from Mayer's doctrine, any other three prismatic rays may still be equally assumed as fundamental colours, and all the rest compounded from them, provided we attend only to the predominant tint resulting, and disregard its dilution with white. For instance, Dr. Young has assumed red, green, and violet as his fundamental colours; and states, as an experimental fact in support of this doctrine, that the perfect sensations of yellow and blue may be produced, the former by a mixture of red and green, and the latter by green and violet. (*Lectures on Natural Philosophy*, p. 439.) Now, if we mix together yellow and white in the proportion of m yellow + n white, the compound will produce a perfect sensation of yellow, unless m be small compared to n ; but, assuming white to be composed as above, this compound is equivalent to

$$n R \text{ red} + (m + n) Y \text{ yellow} + n B \text{ blue.}$$

On the other hand, if we mix together P such red rays (each of the intensity b) and Q such green rays (each consisting of yellow, of the intensity f , and blue of the intensity h) as are supposed in the foregoing article to exist in the spectrum, we have a compound of

$$P \cdot b \text{ red} + Q \cdot f \text{ yellow} + Q \cdot h \text{ blue,}$$

and these will be identical with the former, if we take

$$n R = P b; \quad (m + n) Y = Q f; \quad n B = Q h.$$

Eliminating Q from the two last of these, we get

$$\frac{m}{n} = \frac{f}{h} \cdot \frac{B}{Y} - 1$$

for the relation between M and N . Now the only conditions to be satisfied are that M shall be positive, and not much less than N ; and it is evident that these conditions may be fulfilled an infinite number of ways by a proper assumption of the ratio of f to h . In the same manner, if we suppose a mixture of M rays primary blue = B with N rays of white (= $R + Y + B$) to be equivalent to P rays of prismatic green mixed with Q of violet, we get the equation

$$\frac{m}{n} = \frac{l}{a} \cdot \frac{R}{B} + \frac{h}{f} \cdot \frac{Y}{B} - 1.$$

Suppose, for example, we regard white light as consisting of 20 rays of primary red, 30 of yellow, and 50 of blue, and the several prismatic rays to consist as follows:

Red	8 rays primary red = h .
Orange	7 red + 7 primary yellow = $c + d$.
Yellow	8 yellow = e .
Green	10 yellow + 10 primary blue = $f + h$.
Blue	6 yellow + 12 primary blue = $g + i$.
Indigo	12 blue = k .
Violet	16 blue + 5 primary red = $l + a$.

Then will the union of 15 rays of such red with 30 of such green, produce a compound ray containing $15 \times 8 = 120$ of primary red, $30 \times 10 = 300$ of primary yellow, and $30 \times 10 = 300$ of primary blue; which are the same as exist in a yellow, consisting of 6 rays of white combined with 4 of primary yellow. In like manner, if 75 such green rays be combined with 100 such violet, the result will be $100 \times 5 = 500$ rays of primary red, + $75 \times 10 = 750$ of primary yellow, + $75 \times 10 + 100 \times 16 = 2350$ of primary blue, which together compose a tint identical with that which would result from the union of 25 rays of white with 22 of primary blue; that is to say, a fine lively blue. The numbers assumed above, it must be understood, are merely taken for the sake of illustration, and are no way intended to represent the true ratios of the differently coloured rays in the spectrum.

The analogy of the fixed lines in the solar spectrum might lead us to look for similar phenomena in other sources of light. Accordingly, Fraunhofer has found, that each fixed star has its own particular system of dark and bright spaces in its spectrum; but the most curious phenomena are those presented by coloured flames, which produce spectra (when transmitted through a colourless prism) hardly less capricious than those afforded by solar light transmitted through coloured glasses. Dr. Brewster, Mr. Talbot, and others, have examined these

518.
Dr. Young's
hypothesis
of three
other prima-
ry colours.

519.
Numerical
Illustration.

520.
Phenomena
of coloured
flames

Light. phenomena with attention; but the subject is not exhausted, and promises a wide field of curious research. The following facts may be easily verified:

521. 1. Most combustible bodies consisting of hydrogen and carbon, as tallow, oil, paper, alcohol, &c. when first lighted and in a state of feeble and imperfect combustion, give blue flames. These, when examined by the prism, by letting them shine through very narrow slits parallel to its edge, as described in Art. 487, all give interrupted spectra, consisting, for the most part, of narrow lines of very definite refrangibility, either separated by broad spaces entirely dark, or much more obscure than the rest. The more prominent rays are, a very narrow definite yellow, a yellowish green, a vivid emerald green, a faint blue, and a strong and copious violet.

522. 2. In certain cases when the combustion is violent, as in the case of an oil lamp urged by a blow-pipe, (according to Fraunhofer,) or in the upper part of the flame of a spirit lamp, or when sulphur is thrown into a white-hot crucible, a very large quantity of a definite and purely homogeneous yellow light is produced; and in the latter case forms nearly the whole of the light. Dr. Brewster has also found the same yellow light to be produced when spirit of wine, diluted with water and heated, is set on fire; and has proposed this as a means of obtaining a supply of homogeneous yellow light for optical experiments.

523. 3. Most saline bodies have the power of imparting a peculiar colour to flames in which they are present, either in a solid or vaporous state. This may be shown in a manner at once the most familiar and most efficacious, by the following simple process: Take a piece of packthread, or a cotton thread, which (to free it from saline particles should have been boiled in clean water,) and having wetted it, take up on it a little of the salt to be examined in fine powder, or in solution. Then dip the wetted end of it into the cup of a burning wax candle, and apply it to the exterior of the flame, not quite in contact with the luminous part, but so as to be immersed in the cone of invisible but intensely-heated air which envelopes it. Immediately an irregular sputtering combustion of the wax on the thread will take place, and the invisible cone of heat will be rendered luminous, with that particular coloured light which characterises the saline matter employed.

524. Thus it will be found that, in general,
 Salts of soda give a copious and purely homogeneous yellow.
 Salts of potash give a beautiful pale violet.
 Salts of lime give a brick red, in whose spectrum a yellow and a bright green line are seen.
 Salts of strontia give a magnificent crimson. If analyzed by the prism two definite yellows are seen, one of which verges strongly to orange.
 Salts of magnesia give no colour.
 Salts of lithia give a red, (on the authority of Dr. Turner's experiments with the blow-pipe.)
 Salts of baryta give a fine pale apple-green. This contrast between the flames of baryta and strontia is extremely remarkable.
 Salts of copper give a superb green, or blue green.
 Salt of iron (protoxide) gave white, where the sulphate was used.

Of all salts, the muriates succeed best, from their volatility. The same colours are exhibited also when any of the salts in question are put (in powder) into the wick of a spirit lamp. If common salt be used, Mr. Talbot has shown that the light of the flame is an absolutely homogeneous yellow; and, being at the same time very copious, this property affords an invaluable resource in optical experiments, from the great ease with which it is obtained, and its identity at all times. The colours thus communicated by the different bases to flame, afford in many cases a ready and neat way of detecting extremely minute quantities of them; but this rather belongs to Chemistry than to our present subject. The pure earths, when violently heated, as has recently been practised by Lieutenant Drummond, by directing on small spheres of them the flames of several spirit lamps urged by oxygen gas, yield from their surfaces lights of extraordinary splendour, which, when examined by prismatic analysis, are found to possess the peculiar definite rays in excess, which characterise the tints of flames coloured by them; so that there can be no doubt that these tints arise from the molecules of the colouring matter reduced to vapour, and held in a state of violent ignition.

PART III.

OF THE THEORIES OF LIGHT.

Light.

AMONG the theories which philosophers have imagined to account for the phenomena of light, two principally have commanded attention; the one conceived by Newton, and called from his illustrious name, in which light is conceived to consist of excessively minute molecules of matter projected from luminous bodies with the immense velocity due to light, and acted on by attractive and repulsive forces residing in the bodies on which they impinge, which turn them aside from their rectilinear course, and reflect and refract them according to the laws observed. The other hypothesis is that of Huygens, and also called after his name; which supposes light to consist, like sound, in undulations, or pulses, propagated through an elastic medium. This medium is conceived to be of extreme elasticity and tenuity; such, indeed, that though filling all space, it shall offer no appreciable resistance to the motions of the planets, comets, &c. capable of disturbing them in their orbits. It is, moreover, imagined to penetrate all bodies; but in their interior to exist in a different state of density and elasticity from those which belong to it in a disengaged state, and hence the refraction and reflexion of light. These are the only *mechanical* theories which have been advanced. Others, indeed, have not been wanting; such as Professor Oersted's, who, in one of his works, considers light as a succession of electric sparks, or a series of decompositions and recompositions of an electric fluid filling all space in a neutral or balanced state, &c. &c. In this part, however, we propose only to give an account of the Newtonian and Huygenian theories, so far as they apply to the phenomena already described; and thus prepare ourselves for the remaining more complex branches of the History of the Properties of Light, which can hardly be understood, or even described, without a reference to some theoretical views.

Part III.

525.

§ I. *Of the Newtonian or Corpuscular Theory of Light.*

Postulata. 1. That light consists of particles of matter possessed of inertia and endowed with attractive and repulsive forces, and projected or emitted from all luminous bodies with nearly the same velocity, about 200,000 miles per second.

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2. That these particles differ from each other in the intensity of the attractive and repulsive forces which reside in them, and in their relations to the other bodies of the material world, and also in their actual masses, or inertia.

3. That these particles, impinging on the retina, stimulate it and excite vision. The particles whose inertia is greatest producing the sensation of red, those of least inertia of violet, and those in which it is intermediate the intermediate colours.

4. That the molecules of material bodies, and those of light, exert a mutual action on each other, which consists in attraction and repulsion, according to some law or function of the distance between them; that this law is such as to admit, perhaps, of several alternations, or changes from repulsive to attractive force; but that when the distance is below a certain very small limit, it is always attractive up to actual contact; and that beyond this limit resides at least one sphere of repulsion. This repulsive force is that which causes the reflexion of light at the external surfaces of dense media; and the interior attraction that which produces the refraction and interior reflexion of light.

5. That these forces have different absolute values, or intensities, not only for all different material bodies, but for every different species of the luminous molecules, being of a nature analogous to chemical affinities, or electric attractions, and that hence arises the different refrangibility of the rays of light.

6. That the motion of a particle of light under the influence of these forces and its own velocity is regulated by the same mechanical laws which govern the motions of ordinary matter, and that therefore each particle describes a trajectory capable of strict calculation so soon as the forces which act on it are assigned.

7. That the distance between the molecules of material bodies is exceedingly small in comparison with the extent of their spheres of attraction and repulsion on the particles of light. And

8. That the forces which produce the reflexion and refraction of light are, nevertheless, absolutely insensible at all measurable or appreciable distances from the molecules which exert them.

9. That every luminous molecule, during the whole of its progress through space, is continually passing through certain periodically recurring states, called by Newton fits of easy reflexion and easy transmission, in virtue of which (from whatever cause arising, whether from a rotation of the molecules on their axes, and the consequent alternate presentation of attractive and repulsive poles, or from any other conceivable cause) they are more disposed, when in the former states or phases of their periods, to obey the influence of the repulsive or reflective forces of the molecules of a medium; and when in the latter, of the attractive. This curious and delicate part of the Newtonian doctrine will be developed more at large hereafter.

Ilgnt.

527.

It is the 7th and 8th of these assumptions only which render the course pursued by a luminous molecule, under the influence of the reflective or refractive forces, capable of being reduced to mathematical calculation; for it follows immediately from the 8th, that, up to the very moment when such a molecule arrives in physical contact with the surface of any medium, it is acted on by no sensible force, and therefore not sensibly deviated from its rectilinear path; and, on the other hand, as soon as it has penetrated to any sensible depth within the surface, or among the molecules, by reason of the 7th of the above postulates, it must be equally attracted and repelled by them in all directions, and therefore will continue to move in a right line, as if under the influence of no force. It is only, therefore, within that insensible distance on either side the surface, which is measured by the diameter of the sphere of action of each molecule, that the whole flexure of the ray takes place. Its trajectory then may be regarded as a kind of hyperbolic curve, in which the right lines described by it, previous and subsequent to its arrival at the surface, are the infinite branches, and are confounded with the asymptotes, and the curvilinear portion is concentrated as it were in a physical point. Now, in explaining the phenomena of reflexion and refraction, it is not the nature of this curve that we are called on to investigate. This will depend on the laws of corpuscular action, and must necessarily be of great complexity. All we have to inquire, is the *direction* the ray will ultimately take after incidence, and the final change, if any, in its velocity.

528.

Motion of a
luminous
molecule
under the
influence of
any forces.

Let us, then, consider the motion of a molecule urged to or from the surface of a medium by the united attractions or repulsions of all its particles acting according to any conceivable mathematical law. And, first, it is evident, that supposing the surface mathematically smooth, and the number of attractive or repulsive particles of which it consists, infinite, their total resultant force on the luminous molecule will act in a direction perpendicular to the surface, and will be insensible at all sensible distances from the surface, provided the elementary forces of each molecule decrease with sufficiently great rapidity as the distances increase. This condition being supposed, let x and y be the coordinates of the molecule at any assigned instant; the plane of the x and y being supposed to coincide with that of its trajectory, out of which plane there is evidently no force to turn it, and which must of course be perpendicular to the surface of the medium in which x is supposed to lie: y then will be the perpendicular distance of the luminous molecule from this surface, and Y (some function of y decreasing with extreme rapidity) will represent the force urging it inwards, or *towards* the surface when the molecule is without, *from* when within the medium. Therefore, by the principles of Dynamics, supposing $d t$ to denote the element of the time, we shall have for the equations of the motion

$$\frac{d^2 x}{d t^2} = 0; \quad \frac{d^2 y}{d t^2} + Y = 0; \quad (a)$$

and hence, multiplying the first by dx , the second by dy , adding and integrating, we get

$$\frac{dx^2 + dy^2}{dt^2} + 2 \int Y dy = \text{constant.}$$

Now, v being the velocity of the molecule, we have $v^2 = \frac{dx^2 + dy^2}{dt^2}$, and therefore this equation becomes

$$v^2 = \text{constant} - 2 \int Y dy.$$

It is, however, only with the terminal velocity, or that attained by the light after undergoing the total action of the medium, that we are concerned, and therefore if we put V for its primitive, or initial, and V' for its terminal velocity, we shall have, by extending the integral from the value of y at the commencement of the ray's motion (y_0) to its value at the end (y_1),

$$V'^2 - V^2 = - 2 \int Y dy. \quad (b)$$

Since y_0 and y_1 are supposed infinite, and since the function Y decreases by hypothesis with such rapidity as to become absolutely insensible for all sensible values of y , it is clear that we may take $y_0 = +\infty$ for the first limit of the integral in all cases. With regard to the other, we have now to distinguish two principal cases:

529.

Case of re-
flexion.

The first is that of *reflexion*, where the ray, no matter whether before its arrival at the surface, or at reaching it, or even after passing some small distance into the medium, is turned back by the prevalence of the repulsive force, and pursues the whole of its course afterwards without the medium. Now in this case if we resolve the integral $\int Y dy$ into its elements, these, in the approach of the molecule to the surface, may be represented as follows,

$$\&c. + Y' \times - dy + Y'' \times - dy + Y''' \times - dy + \&c.$$

But in the recess of the molecule, the values of y increase again by the same steps as they before diminished and become identical with the former ones; and Y' , Y'' , $\&c.$, the values of Y corresponding to the successive values of y , remain therefore the same, both in size and magnitude; the corresponding elements of the integral generated during the recess of the molecule will be then

$$\&c. + Y' \times + dy + Y'' \times + dy + Y''' \times + dy + \&c.$$

Light.

So that, combining both, the latter exactly destroy the former, and give $\int Y dy = 0$ when extended from one end to the other of the trajectory. Thus we have, in the case of reflexion,

Part III.

$$V'^2 - V^2 = 0, \quad \text{or } V' = V.$$

The second case is that in which the whole course of the ray after incidence lies within the medium, or the case of refraction. Here the values of y before incidence are all positive, and after, all negative; and, moreover, the change of sign in dy which happened in the case of reflexion, does not here take place. Hence $\int Y dy$ must be extended from $+\infty$ to $-\infty$, and its value will not vanish, but (on account of the rapid decrease of the function Y) will have some finite value. Now this can only be dependent on the arbitrary quantities which enter into the composition of Y ; in other words, on the nature of the medium and the ray, and not at all on the constants which determine the direction of the ray with respect to the surface, (as its inclination or the position of the plane of incidence.) Hence we may suppose $\int Y dy = -\frac{1}{2}kV^2$, where k is a constant independent of the direction of the ray, and determined only by its nature and that of the medium, and we shall have

530.
Case of refraction.

$$V'^2 = V^2(1+k); \quad V' = V \cdot \sqrt{1+k} = \mu V, \quad (c)$$

putting $\sqrt{1+k} = \mu$.

Hence we see that both in refraction and reflexion, on this hypothesis, the velocity of the ray after deviation is the same in whatever direction the ray be incident, viz. in a given ratio to the velocity before incidence, this ratio being one of equality in the case of reflexion.

531.
Law of velocities.

Let us next consider the direction of the ray after flexure. To this end let θ = the angle made by its path at any moment with the perpendicular to the surface, then will $\sin \theta = \frac{dx}{ds}$, putting ds for $\sqrt{dx^2 + dy^2}$, the

532.
Direction of the ray after flexure.

element of the arc. Now if we integrate the equation $\frac{d^2x}{dt^2} = 0$ once we get $\frac{dx}{dt} = \text{constant} = c$, and $dx = c dt$, wherefore $\sin \theta = \frac{c dt}{ds}$. But $\frac{ds}{dt} = \frac{V}{\sin \theta}$, therefore $\sin \theta = \frac{c}{V}$. Let therefore θ_0 and θ_1 represent the initial and terminal values of θ , or the angles of incidence and reflexion, or refraction of the rectilinear portions of the ray, and we get

Constancy of ratio of sines of incidence and refraction.

$$\sin \theta_0 = \frac{c}{V}, \quad \text{and } \sin \theta_1 = \frac{c}{V'}.$$

and dividing one by the other

$$\frac{\sin \theta_0}{\sin \theta_1} = \frac{V'}{V} = \mu.$$

That is to say, the sines of incidence and refraction, or reflexion, are to each other in a constant ratio, viz. the inverse ratio of the velocities of the ray before and after incidence.

Thus we see the Newtonian hypothesis satisfies the fundamental conditions of refraction and reflexion without entering into any consideration respecting the laws of the refracting and reflecting forces, or even the order of their superposition. There may be as many alternations of attraction and repulsion as we please, and the reflected or refracted ray may therefore, prior to its final recess from the surface, make any variety of undulations; all that is required is the extremely rapid decrease of the function Y expressing the total force before the distance attains a sensible magnitude.

533.

Hence also, V and V' being the velocities before and after incidence, and μ the index of refraction, we have

534.

$$V' : V :: \mu : 1,$$

which shows, that when a ray passes from a rarer medium to a denser, its velocity is increased, and *vice versa*.

Moreover, we have

$$k = \frac{V'^2 - V^2}{V^2} = \left(\frac{V'}{V}\right)^2 - 1 = \mu^2 - 1 = \frac{2 \int -Y dy}{V^2}.$$

535.
Refractive power of a medium.

Now if we suppose the form of the function Y to be the same for all media, and that they differ in the energy of action only by reason, first, of a greater density, owing to which more molecules are brought within the sphere of activity; and, secondly, by reason of a greater or less affinity, or intensity of action of each molecule, we may suppose Y to be represented by $S \cdot n \cdot \phi(y)$, where S is the specific gravity, or density, n the intrinsic refractive energy of the medium, and $\phi(y)$ a function absolutely independent of the peculiarities of the medium, and the same for all natural bodies. Hence $\int -Y dy = S \cdot n \cdot \int -\phi(y) dy = S \cdot n \cdot \text{constant}$ because $\int -\phi(y) dy$ taken from $y = +\infty$ to $y = -\infty$ will now be an absolute numerical constant. We have then, according to this doctrine,

$$n = \frac{\mu^2 - 1}{S} \times \frac{V^2}{2 \cdot \text{constant}}.$$

If μ be the refractive index of a given standard ray out of a vacuum, V the velocity of that ray in vacuo is known, and is also an absolute constant; so that n , the intrinsic refractive power of the medium is proportional to

Light. (refractive index)² - 1
specific gravity

Such is Newton's idea of the *refractive power* of a medium as differing from its *refractive index*. It rests, however, on a purely hypothetical assumption, that of the similarity of *form* of the law of force for all media, respecting which we can be said to know nothing whatever. For a table of its values for different media, see the Collection of Tables at the end of this Essay.

536.
Principle of
least action
employed.

The constancy of the ratio of the sines of incidence and refraction has here been derived by direct integration of the fundamental equations. There is, however, another mode of deducing this important law, much more circuitous, it is true, in this simple case, but which offers peculiar advantages in the more complicated ones of double refraction; and which, therefore we shall here explain, to familiarize the reader beforehand with its principle and mode of application. It consists in the employment of what is called, in Dynamics, the *principle of least action*, in virtue of which the sum of each element of the trajectory described by any moving molecule multiplied by the velocity of its description (or the integral $\int v ds$) is a minimum when taken between any two fixed points in the trajectory. The trajectory described by any luminous molecule may be regarded as consisting of two rectilinear portions, or hyperbolic branches, confounded with their asymptotes, and one curvilinear one concentrated in a space of insensible magnitude, a physical point. Within this *point* the whole operation of the flexure of the ray, however complicated, is performed; and *here* the velocity is variable. In the branches it is uniform. Suppose, then, A and B to be any two fixed points in these, taken as points of departure and arrival of a ray, and let C be the point in the surface of a reflecting or refracting medium where the flexure takes place, and suppose $AC = S$, $BC = S'$ and let σ be the excessively minute curvilinear portion of the ray at C, and v the variable velocity with which it is described, V and V' being those with which S and S' are described. Then may the integral $\int v ds$ be resolved into the three portions $\int V dS + \int v d\sigma + \int V' dS'$. Of these the second is utterly insensible, by reason of the minuteness of σ , and the other two, since V and V' are constant, become merely $V \cdot S + V' \cdot S'$.

537. The position of C, then, with respect to A and B, will be determined by the condition $V \cdot S + V' \cdot S' = a$ minimum, A and B being supposed fixed, and C any how variable on the surface. Now, in the case before us, V the velocity of the light before, and V' that after incidence, are *both*, as we showed in Article 529 and 530, independent of the direction of the incident and reflected or refracted rays, or of the position of C; and, therefore, are to be considered as absolute constants in this problem of minima, which is thus reduced to a simple geometrical question. Given A and B to find C, a point in a given plane, such that $V (= \text{constant}) \times AC + V' (= \text{constant}) \times BC$ shall be a minimum. Nothing is easier than the solution. Put a, b, c, a', b', c' for the respective coordinates of A and B, and x, y, o for that of C, taking the given plane for that of the x, y . Then

$$V \cdot S + V' \cdot S' = V \cdot \sqrt{(x-a)^2 + (y-b)^2 + c^2} + V' \cdot \sqrt{(x-a')^2 + (y-b')^2 + c'^2}$$

Solution of
the geometrical
problem of the
minimum

is to be a minimum by the variation of x and y , independent of each other. This gives, by differentiation,

$$\frac{V}{S} \{ (a-x) dx + (b-y) dy \} + \frac{V'}{S'} \{ (a'-x) dx + (b'-y) dy \} = 0;$$

and this, since x and y are independent, must vanish, whatever values are assigned to dx and dy , therefore we must have separately

$$\frac{V}{S} (a-x) + \frac{V'}{S'} (a'-x) = 0; \quad \frac{V}{S} (b-y) + \frac{V'}{S'} (b'-y) = 0. \quad (d)$$

These give, respectively,

$$\frac{S'}{S} = - \frac{V'}{V} \cdot \frac{a-x}{a-a'}; \quad \frac{S'}{S} = - \frac{V'}{V} \cdot \frac{b'-y}{b-y};$$

by equating which we get

$$(a'-x)(b-y) = (b'-y)(a-x);$$

or multiplying out and reducing

$$y = x \cdot \frac{b-b'}{a-a'} + \frac{a b' - b a'}{a-a'};$$

and, consequently,

$$b'-y = \frac{b-b'}{a-a'} (a'-x)$$

This equation expresses, that the two portions S and S' of the ray before and after incidence on the surface at C both lie in one plane, and that this plane is perpendicular to the surface, or to the plane of the coordinates x, y .

538.

Constancy
of the ratio
of the sines
deduced.

Again, if we resume the equations (d) and putting them under the form

$$S' (a-x) = - \frac{V'}{V} S (a'-x); \quad S' (b-y) = - \frac{V'}{V} (b'-y) \cdot S.$$

Square and add them we get

$$S'^2 \left\{ (a-x)^2 + (b-y)^2 \right\} = \left(\frac{V'}{V} \right)^2 \cdot \left\{ (a'-x)^2 + (b'-y)^2 \right\} \cdot S^2.$$

Now if we put θ for the angle made by the portion S with a perpendicular to the surface, or the angle of incidence of the ray, and θ' for that made by the other S' with the same perpendicular, or the angle of emergence, we shall have

$$\sin \theta = \frac{\sqrt{(a-x)^2 + (b-y)^2}}{S} \quad \text{and} \quad \sin \theta' = \frac{\sqrt{(a'-x)^2 + (b'-y)^2}}{S'}.$$

So that the above equation is equivalent simply to

$$\sin \theta = \frac{V'}{V} \cdot \sin \theta',$$

which is the same with the result before obtained.

The principle of least action, then, in the case before us, has enabled us to dispense with one integration of the differential equations expressing the motion of the luminous molecule; and its applicability to this purpose depends, as we have seen, on the relation between V and V' ; the velocities of the light, before and after incidence, being known. This relation has here been deduced *a priori*; but had it been merely known, as a matter of fact, a conclusion established by experiment, it would not be on that account the less applicable to the same purpose, and the laws of refraction and reflexion would be derivable from it by the same process. There would, however, be this main difference; that, in the latter case, we should have no occasion to employ the differential equations at all, and therefore none to enter into any consideration of the forces acting on the luminous molecule, or their mode of action. The principle of least action establishes, independent of, and anterior to, all particular suppositions as to the forces which operate the flexure of the ray, (further than that they are functions of the distances from their origins or centres,) an analytical relation between the velocities before and after incidence, and the directions of its direct and deviated branches; a relation nearly as general as the laws of dynamics themselves, and expressive, in fact, of only the one condition above mentioned. And this relation, from its form, enables us, whenever the relation of the velocities is known, to determine that of the directions of the two portions of the ray, and *vice versa*, without having recourse to the differential equations at all. In the simple case before us this may seem a needless refinement, the equations being so simple. It is otherwise, however, in the theory of double refraction. There the forces in action are altogether unknown, not only in respect of their intensity, but of their directions; and so far, therefore, from being able in that theory to integrate the equations of the ray's motion, we cannot even express them analytically. The principle we are now considering is, in such a case, all the ground we have to stand upon; and has been ingeniously and elegantly applied by Laplace, in that theory, to reduce the complicated laws of double refraction under the dominion of analysis.

539.
Advantages afforded by the principle of least action.

Applicable to other cases.

In fact, suppose that the velocities of the incident and deviated portions of the rays, instead of being the same in every direction, varied with the positions of these portions with respect to the surface of the medium, or to any fixed lines or axes in space. Then will V and V' , instead of being invariable, be represented by functions of the three coordinates of the point C , either rectangular, as x, y, z ; or polar, as ϕ, θ , and γ ; and the portions S and S' of the rays intercepted between A and B respectively, and the surface at C , will, in like manner, be functions of the same coordinates. So that the condition

540.
Mode of its application in general.

$$V \cdot S + V' \cdot S' = \text{a minimum}$$

will afford, by differentiation and putting the differential equal to zero, an equation of the form $L dx + M dy + N dz = 0$, or $L d\phi + M d\theta + N d\gamma = 0$, as the case may be. The equation of the surface also being differentiated affords another relation of the same kind; and these being the only conditions to which the differentials dx, dy, dz are subject, we may eliminate one, and put the coefficients of the remaining ones separately equal to zero. Thus we get two equations between the coordinates, which, combined with that of the surface, suffice to determine them, i. e. to fix the point C at which the ray AC must meet the surface, in order that, being there deviated by the action of the medium, it may, after flexure, proceed to B ; and thus the problem of reflexion or refraction may be resolved in all its generality, so soon as the nature of the functions V, V' is known. But to return to the case of ordinary reflexion and refraction, from which this is a digression.

Let us consider, a little more in detail, what may be conceived to happen to a ray at the confines of the surface of a medium. We may suppose, then, that there exist a series of laminar spaces, or strata, within which the attractive and repulsive action of the molecules of the medium alternately predominate. Of these there may be any number, and either may be exterior to the rest. It is, in fact, the assemblage of these laminae which is to be regarded as the surface of the medium. Suppose now a ray Aa (fig. 119) to be moving towards the medium. Its course will be rectilinear up to a , where it first comes within the action of the medium. If the first stratum into which it enters be one of attraction, its course will be bent as ab into a curve concave towards the surface, and its velocity in the direction perpendicular to the surface will be increased. Arrived at b let the force change to repulsive; the trajectory will have at b a point of contrary flexure, the portion $b c$ within this lamina will be convex to the surface, and the velocity towards the surface will be diminished in the whole progress of the ray through it, and so for any number of alternations. Let us now suppose, that in passing through any repulsive lamina, as C , the repulsion should be so strong, or the original velocity of approach to the surface so small, as that the whole of it shall be destroyed. In this case the ray for a moment will be moving as at C , parallel to the surface, but the repulsive force continuing its action will turn it back; and the forces

541.
Course of a ray in the confines of a reflecting and retracting medium traced. Fig. 119.

Light.

now being all equal to what they were before, but acting in a contrary direction with respect to the motion of the molecule, it will describe a portion $Cd'c'b'a'B$ similar, and equal to the portion on the other side of C . This is the case of reflexion. But suppose, as in fig. 120, the ray to have such an initial velocity of approach, or the repulsive forces to be so feeble, compared to the attractive, that before its whole velocity perpendicular to the surface is destroyed, it shall have passed through all the strata of attraction and repulsion, and entered the region where the forces of all the molecules are in equilibrium, as at e . In this case the remainder of its course will be rectilinear, and wholly within the medium. This is the case of refraction. In both cases, it is the final course it takes, or the direction of the asymptotic branches $a'B$ or eB , about which only we have any knowledge; of the number and nature of the undulations of its course between a and a' , or e , we know nothing.

542.

Motion of a ray at common surface of two media.

The whole of this reasoning applies equally to the motion of a luminous molecule at the confines of two media, as at the surface separating one medium from a vacuum. The molecules of either medium being supposed *uniformly distributed*, and *acting equally in all directions around them*, the resultant of all their forces on the luminous particle must be perpendicular to the common surface, which is all that is required in the foregoing theory.

543.

Newtonian idea of a ray of light as composed of a succession of molecules.

In the Corpuscular doctrine, a ray of light is understood to mean a continued succession or stream of molecules, all moving with the same velocity along one right line, and following each other close enough to keep the retina in a constant state of stimulus, i. e. so fast, that before the impression produced by one can have time to subside another shall arrive. It appears, by experiment, that to produce a continued sensation of light, it is sufficient to repeat a momentary flash about 8 or 10 times in a second. If a red-hot coal on the point of a burning stick be whirled round, so as to describe a circle, and the velocity of rotation be greater than 8 or 10 circumferences per second, the eye can no longer distinguish the place of the luminous point at any instant, and the whole circle appears equally bright and entire. This shows, evidently, that the sensation excited by the light falling on any one point of the retina, must remain almost without diminution till the impression is repeated during the subsequent revolution of the luminary. Now, if uninterrupted vision can be produced by momentary impressions, repeated at intervals so distant as a tenth of a second, it is easy to conceive that the individual molecules of light in a ray will not require to follow close on each other to affect our organs with a continued sense of light. As their velocity is nearly 200,000 miles per second, if they follow each other at intervals of 1000 miles apart, 200 of them would still reach our retina per second, in every ray. This consideration frees us from all difficulties on the score of their jostling, or disturbing each other in space, and allows of infinite rays crossing at once through the same point of space without at all interfering with each other, especially when we consider the minuteness which must be attributed to them, that (moving with such swiftness) they should not injure our organs. If a molecule of light weighed but a single grain, its inertia would equal that of a cannon ball of upwards of 150 pounds weight, moving at the rate of 1000 feet per second. What then must be their tenuity, when the concentration of millions upon millions of them, by lenses or mirrors, has never been found to produce the slightest mechanical effect on the most delicately contrived mechanism, in experiments made expressly to detect it. (See Mr. Bennet's *Experiments*, *Phil. Trans.* 1792, vol. lxxxii. p. 87.)

Their distance inter se.

Their extreme tenuity illustrated.

544.

Partial reflexion explained on Newton's principles.

When a ray of light falls on a reflecting or refracting surface, since all its molecules move with equal velocity and are incident in the same line, it would seem that whatever took place with one should equally happen to all; and that, if the first underwent reflexion, all should do so; while, on the other hand, if one could penetrate the surface, and pursue its course entirely within the medium, all ought to do the same. This, however, is contrary to experience; as whenever a ray of light is incident on the exterior surface of any medium, a part only is reflected, and the rest enters the medium. No theory can be satisfactory which does not render a good account of so principal a fact. The Newtonian doctrine accounts for it by the fits of easy reflexion and transmission. To understand this explanation we must recur to the ninth postulate, (Art. 526,) and suppose two molecules to arrive at the surface under the same incidence, the one in a fit of easy reflexion, the other in one of easy transmission. The former will then be more affected by the repulsive forces, the latter by the attractive of the molecules of the medium; and hence it is evident, that the one may be reflected under circumstances of incidence, &c. in which the other will penetrate the surface and be refracted. Now it will depend entirely on the nature of the medium, and the initial velocity of a luminous molecule *towards* it, (which is as the cosine of the angle of incidence,) whether it will require the whole exertion of its repulsive forces, in their most energetic manner, to destroy that velocity and produce reflexion, or only a part of them. In the former case only such molecules as arrive in the *most favourable* disposition to be reflected, or in the most intense phase of a fit of easy reflexion, can be reflected. In the latter, such as arrive in less favourable dispositions, or in *less intense* phases of fits of reflexion, may be reflected; and if the repulsive forces of the medium be very intense, in comparison with the attractive ones, or if the obliquity of incidence be so great as to give the molecule a very small velocity perpendicular to the surface, even those molecules which arrive in the less energetic phases of fits of easy transmission may still be unable to penetrate the strata of repulsion.

545.

Reflexion more copious at great obliquities.

Hence, then, we see that the proportion of the molecules of a ray falling on the surface of a medium in every possible state or phase of their fits, which undergo reflexion, will depend, first, on the nature of the medium on whose surface they fall, or if it be the common surface of two, then on both; secondly, on the angle of incidence. At great obliquities, the reflexion will be more copious; but even at the greatest, when the incident ray just grazes the surface, it by no means follows that every molecule, or even the greater part, *must* be reflected. Those which arrive in the most favourable phases of their fits of transmission, will obey the influence of small attractive forces, in preference to strong repulsive ones; but it will depend entirely on the nature of the media whether the former or the latter shall prevail, the fits in the Newtonian doctrine being conceived only to *dispose* the luminous molecules, other circumstances being favourable, to reflexion or transmission; to exalt the forces which

Light.

tend to produce the one and to depress those which act in favour of the other, but not to determine, absolutely, its reflexion or transmission under all circumstances.

These conclusions are verified by experience. It is observed, that the reflexion from the surfaces of transparent (or indeed any) media, becomes sensibly more copious as the angle of incidence increases; but at the *external* surface of a single medium is never total, or nearly total. In glass, for instance, even at extreme obliquities, a very large portion of the light still enters the glass and undergoes refraction. In opaque media, such as polished metals, the same holds good; the reflexion increases in vividness as the incidence increases, but never becomes total, or nearly so. The only difference is, that here the portion which penetrates the surface is instantly absorbed and stifled.

The phenomena which take place when light is reflected at the common surface of two media, are such as from the above theory we might be led to expect,—with the addition, however, of some circumstances which lead us to limit the generality of our assumptions, and tend to establish a relation between the attractive and repulsive forces, to which the refraction and reflexion of light are supposed to be owing. For it is found, that when two media are placed in perfect contact, (such as that of a fluid with a solid, or of two fluids with one another,) the intensity of reflexion at their common surface is always less, the nearer the refractive indices of the media approach to equality; and when they are exactly equal, reflexion ceases altogether, and the ray pursues its course in the second medium, unchanged either in direction, velocity, or intensity. It is evident, from this fact, which is general, that the reflective or refractive forces, in all media of equal refractive densities, follow exactly the same laws, and are similarly related to one another; and that in media unequally refractive, the relation between the reflecting and refracting forces is not arbitrary, but that the one is dependent on the other, and increases and diminishes with it. This remarkable circumstance renders the supposition made in Art. 535, of the identity of form of the function Y , or $\phi(y)$ expressing the law of action of the molecules of all bodies on light indifferently, less improbable.

To show experimentally the phenomena in question, take a glass prism, or thin wedge of very small refracting angle (half a degree, for instance: almost any fragment of plate glass, indeed, will do, as it is seldom the two sides are parallel,) and placing it conveniently with the eye close to it, view the image of a candle reflected from the exterior of the face next the eye. This will be seen accompanied at a little distance by another image, reflected internally from the other face, and the two images will be nearly of equal brightness, if the incidence be not very great. Now, apply a little water, or a wet finger, or, still better, any black substance wetted, to the posterior face, at the spot where the internal reflexion takes place, and the second image will immediately lose great part of its brightness. If olive oil be applied instead of water, the defalcation of light will be much greater, and if the substance applied be pitch, softened by heat, so as to make it adhere, the second image will be totally obliterated. On the other hand, if we apply substances of a higher refractive power than glass, the second image again appears. Thus, with oil of cassia it is considerably bright; with sulphur, it cannot be distinguished from that reflected at the first surface; and if we apply mercury, or amalgam, (as in a silvered looking-glass,) the reflexion at the common surface of the glass and metal is much more vivid than that reflected from the glass alone.

The destruction of reflexion at the common surface of two media of equal refractive powers explains many curious phenomena. If we immerse an irregular fragment of a colourless transparent body (as crown glass) in a colourless fluid of precisely equal refractive power, it disappears altogether. In fact, the surface being only visible by the rays reflected from it; destroy this reflexion, and the object must cease to be seen, unless from any opacity in its substance reflecting rays from its interior, which is not here contemplated. Hence, if the powder of any such substance be moistened with a fluid of the same refractive density, all the internal and external reflexions at the surfaces of the small fragments of which it consists, which, blended and confused, present the general appearance of a white opaque mass, will be destroyed, and the powder will be rendered perfectly transparent. A familiar instance of this nature is the transparency given to paper by moistening it with water, or, still better, with oil; paper is composed of an infinity of minute transparent, or nearly transparent fibres of a ligneous substance, having a refractive power probably not very different from some of the more refractive oils. Its whiteness is caused by the confused reflexion of the incident rays at all possible angles, both internally and externally, those which have escaped reflexion at one fibre, undergoing it among those beneath. If moistened with any liquid, the intensity of these reflexions is weakened, and the more the more nearly its refractive power approaches to that of the paper itself; so that a considerable number of rays find their way through it, and emerge at the posterior surface. The transparency acquired by the hydrophane, by immersion in water, is, no doubt, owing to this cause; the water filling up the minute pores, and enfeebling the internal reflexion; and Dr. Brewster, in a very curious and interesting Paper on the tabasheer, (a siliceous concretion found in sugar-canes, and the lowest in the scale of refracting powers among solids,) has explained on this principle a number of extraordinary phenomena exhibited on moistening that substance with various liquids, (see *Philosophical Transactions*, 1819.)

The reasoning of Art. 529 applies, it is evident, equally to the case when a ray is reflected from the interior surface of a dense medium placed in air, and when from the exterior. The only difference is, that in the latter case the reflexion is performed by the action of repulsive, and in the former by that of attractive forces. The course of a ray internally reflected may be conceived, as in fig. 121 and 122; and the reflexion may take place in any of the attractive regions, or laminae, whether within or without the true surface, i. e. the last layer of molecules which constitute the medium. There is one case of internal reflexion, however, too remarkable to be passed without more particular notice. It is, that when the interior angle of incidence exceeds the *limiting angle* whose sine is $\frac{1}{\mu}$, (see Art. 183 *et seq.*;) and when, as we there stated, as a result of experiment, the

Part III.

546.

Confirmed by experiment.

547.

Reflexion at common surface of two media.

548.

Phenomena exhibited experimentally.

549.

Phenomena depending on the foregoing principles.

550.

Transparency of oiled paper.

Total internal reflexion.

Light.

internal reflexion is *total*. To see how this happens, let us consider a ray incident exactly at this angle, and in the most intense phase of its fit of transmission. Then will it be refracted; and, since the angle of refraction must be just 90° , (by reason of the generality of the demonstration of the law of refraction in Art. 529,) it will emerge, grazing the surface, exactly at the extreme boundary of the outermost region C B, (fig. 123,) where all sensible action ceases. Its initial velocity under these circumstances in the direction perpendicularly to the surface, is barely sufficient to carry it up to this extreme limit, where it is quite annihilated. If, then, we conceive another ray, also incident in the most intense phase of its fit of transmission, but at an angle more oblique by an infinitely small quantity, then, since its initial velocity at right angles to the surface is less, it will be destroyed before it has quite reached this limit, and the ray will therefore begin to move parallel to the surface, just *within* the last limit to the sphere of its action.

551.

The outermost sphere of action necessarily attractive.

Now the last action which the surface exerts, or that force which extends to the greatest distance from it, cannot be otherwise than attractive; for, first, were it repulsive, it is evident that no ray incident externally at an extreme incidence, (i. e. approaching indefinitely to 90° ;) could by possibility escape reflexion; and, secondly, no ray on that supposition could emerge from within the medium, without having at its emergence an obliquity to the surface greater than some finite angle, the last action of the medium being in this case to bend it *outwards*, both which consequences are contrary to fact. Or we may consider the point thus, Since a ray incident within, at the limiting angle, emerges, if it emerge at all, parallel to the surface; and since every point in the curve described by it previous to the instant of emergence is nearer to the medium than the line of its ultimate direction, it is geometrically impossible that the curvature immediately adjacent to the point of emergence should be otherwise than *concave* towards the medium; and must, therefore, of necessity be produced by a force directed to it, i. e. an attractive one.

552.

Hence, the luminous molecule we have been considering, will be within the attractive region at the moment when its perpendicular motion is destroyed; it will, therefore, be turned inwards, as at the dotted line fig. 123, and be reflected. *A fortiori*, therefore, will every molecule incident in a less intense phase of a fit of transmission, or in one of reflexion, as well as every one incident at a more oblique incidence, i. e. with a less initial perpendicular velocity, be reflected. Those in which the circumstances are most favourable to transmission will reach the exterior attractive region, as in fig. 123. Others in which they are less so will be reflected in some intermediate region, as in fig. 122, while those which are incident at extreme internal obliquities, or in the most intense phases of fits of reflexion, will have their courses as represented in fig. 121.

553.

Fact respecting oblique reflexion from water.

The conclusion at which we have arrived in the last Art. that the attractive force of a medium on the molecules of light extends to a greater distance than the repulsive, is, as we have seen, a necessary consequence of dynamical principles; and so far from being in opposition to Newton's doctrine of reflexion, as has been said, is in perfect accordance with it. Dr. Brewster has been led to the same conclusion by peculiar considerations grounded on his experiments on the law of polarization, (*Phil. Trans.*, 1815, p. 133,) and has applied it to explain a curious fact noticed by Bouguer, viz. that although water be much less reflective than glass at small incidences, yet at great ones (as $87\frac{1}{2}^\circ$) it is much more so. Now, supposing the light to have undergone the *whole effect* of the refracting forces, in both cases before it suffers reflexion, its incidence, when it reaches the region of the repulsive forces, will have been diminished in the case of glass, to $57^\circ 44'$, but in that of water only to $61^\circ 5'$, and therefore being incident more obliquely on the water it ought to be more copiously reflected. Whatever we may think of the validity of this explanation, it is certainly ingenious, and the fact extremely remarkable, and deserving of all attention.

554.

Experiment showing the phenomena of total reflexion.

To see the phenomena of total reflexion to the best advantage, lay down a right-angled glass prism on a black substance close to a window, with its base horizontal, as in fig. 124, and apply the eye close to the side, looking downwards. The base will be seen divided into two portions, by a beautiful coloured arch like a rainbow concave to the eye, the portion above the arch being extremely brilliant and vivid, and giving a reflexion of all external objects no way to be distinguished from reality. On the other hand, the space within the concavity of the bow is comparatively sombre, the reflexion of the clouds, &c. on that part of the base being much less vivid. If, instead of placing it on a black body, we hold it in the hand, and place a candle *below* it, this will be visible; but (wherever placed) will always appear in some part of the base within the concavity of the bow. Fig. 124 represents the course of the rays in this experiment, E being the eye, N G, O F, P D rays incident through the farther side at various angles of obliquity on the base, and reflected to the eye at E, of which O F is incident precisely at the limiting angle. It is obvious, that all the rays towards N incident on that part of the base beyond F being too oblique for transmission will be totally reflected, while those incident between F and A, being less oblique than is required for total reflexion, will be only partially so, a portion escaping through the base in the direction D Q. Again, if we place a luminary at any point as L below the base, it is manifest that to reach the eye, a ray from it must fall between A and F, as L D, and that no ray falling on any part of the base between B and F can be refracted to E.

555.

Reflected prismatic bow.

The coloured arch separating the region of total from that of partial reflexion, is thus explained. For, simplicity, let us suppose the eye within the medium, (to avoid considering the reflexion at the inclined surface A C of the prism;) and, first, considering only the extreme red rays, if we drop a perpendicular from the eye on the base of the prism, and make this the axis of a cone, the side of which is inclined to the axis at the angle

whose sine is $\frac{1}{\mu}$, (or the limiting angle for extreme red rays;) and if we conceive such rays to emanate in all directions *from the eye*, then all which fall *without* the circular base of this cone will be totally, but those within only partially reflected. Thus, were there no other than such red rays of this precise refrangibility, the

Light. region of partial reflexion would be a circle whose radius = height of the eye above the base \times tangent of the angle whose sine is $\frac{1}{\mu} = \frac{H}{\sqrt{\mu^2 - 1}}$. In like manner, the radius of the circular space, within which only a partial reflexion of violet rays takes place, is $\frac{H}{\sqrt{\mu'^2 - 1}}$, or $\frac{H}{\sqrt{(\mu + \delta\mu)^2 - 1}}$, being less than the value

of the same radius for the red rays. Hence, in the space between the two circles, the violet rays will be totally, and the red only partially reflected; and, therefore, the whole of this space will have an excess of violet light. A similar reasoning holds good for the intermediate rays; and the shading away from the bright space without, to the comparatively dark one within, will, in consequence, be performed by the abstraction first of the red, next of the orange rays, and so on through the spectrum, leaving a residual light, which continually deviates more and more from white, and verges to blue. If now we suppose each ray to be incident in the contrary direction so as to be reflected to the eye instead of emanating from it, every thing will equally hold good, and the eye will see a bright space without; separated from an obscure space within the base of the cone, the transition from one to the other being not sudden, but marked by a blue border, the colour of which is more lively towards the interior. Now such is the fact, with one difference, however, that the coloured arch appears slightly tinged with pink on its convex side. This, as it is incompatible with theory, can be owing, it should seem, to no cause but contrast; a most powerful source of illusion in all the phenomena of colours, and of which this is, perhaps, one of the most striking and curious instances. Newton (*Optics*, part ii. exp. 16) takes no notice of this part of the phenomenon, (which was first observed and described by Sir W. Herschel,) though he gives the same explanation of the rest with that here set down. The effect of refraction at the side BA of the prism will somewhat modify the figure of the bow, giving it a tendency to a conchoidal form at great obliquities of the emergent rays.

If the side BC of the prism be covered with black paper, and a bright scattered light be thrown on the base from below, (as from an emiered glass applied with its rough side close to the base,) the converse of the above described phenomena will be seen. A totally black space will be seen beyond F, and a bright one within it. The separation being marked by a bow of a vivid red colour, graduating through orange and pale yellow into white, the red being outwards. It is evident that this phenomenon is, in all its parts, complementary to that of the blue bow seen by reflexion, and therefore requires no more particular explanation. It should be noticed, however, that in this bow no appearance of blue or violet within its concavity is ever seen; so that the effect which we have above attributed to contrast in the reflected bow has nothing corresponding to it in the transmitted one.

The intensity and regularity of reflexion at the external surface of a medium, is found to depend not merely on the nature of the medium, but very essentially on the degree of smoothness and polish of its surface. But it may reasonably be asked, how any regular reflexion can take place on a surface polished by art, when we recollect that the process of polishing is, in fact, nothing more than grinding down large asperities into smaller ones by the use of hard gritty powders, which, whatever degree of mechanical comminution we may give them, are yet vast masses, in comparison with the ultimate molecules of matter, and their action can only be considered as an irregular tearing up by the roots of every projection which may occur in the surface. So that, in fact, a surface artificially polished must bear somewhat of the same kind of relation to the surface of a liquid, or a crystal, that a ploughed field does to the most delicately polished mirror, the work of human hands. Now to this question the Newtonian doctrine furnishes an answer quite satisfactory. Were the reflexion of light performed by actual impact of its molecules upon those of the reflecting medium, no regular ordinary reflexion could ever take place at all, as it would depend entirely on the shape of the molecules, or asperities of the surface, and the inclinations of their surfaces to the general surface of the medium at the point of incidence, what should be the direction ultimately taken by each particular ray. Now these must vary in every possible manner in uncrystallized bodies, so that in reflexion from the surfaces of *these* the light would be uniformly scattered in every direction. On the other hand, in crystallized media, each molecule presenting only a limited number of strictly plane surfaces, and the corresponding faces of all being mathematically parallel, reflexion would indeed be regular; but its direction would be regulated only by that of the incident ray and the position of certain fixed lines within the crystal; and would be quite independent of either the smoothness or the inclination of the polished surfaces of it, whether natural or artificial; add to which, that instead of the reflected pencil of rays being *single*, it would in most cases be multiple. All these consequences are so contrary to fact, that it is evident we must suppose the distance to which the forces producing reflexion extend much greater not only than the size of, or interval between individual molecules, but even greater than the minute inequalities, or furrows in the artificially polished surfaces of media. Granting this, the difficulty vanishes; for the average action of many molecules, or many corrugations, will present an uniformity, while individually they may offer the greatest diversity. To illustrate this, we need only cast our eyes on fig. 125, where AB represents the rough surface of a medium, and AC the radius of one of the spheres of attraction, or repulsive activity of a single molecule A. Conceiving now the summits of all the elevations *a, b, c, d* to lie in a plane, let spheres be described with their centres equal to AC. Then their intersections will generate a kind of mamillated surface *aβγδ*, which, however, if the radii of the spheres be at all considerable with respect to the distances of their centres, will approach exceedingly near to a mathematical plane, infinitely more so than the surface AB need be supposed. Hence, a ray of light impinging on the medium will come within the sphere of its action not at an irregular surface, but nearly at a plane one; and the resultant action of all the molecules in action, supposing them distributed with uniformity over AB, will be perpendicular to this surface. The same will hold good of the layer of molecules (however interrupted) immediately under the summits *b, c, d, &c.*, and of all the other

556.
Transmitted
prismatic
bow.

557.
Reflexion
at surfaces
artificially
polished
explained.

Light not
reflected by
direct im-
pact on
bodies.

But by
forces at a
distance.

Light. layers into which the whole surface can be divided. So that the essential conditions on which the Newtonian doctrine of reflexion and refraction reposes, (viz. equality of force at equal distances from the general level of the surface, and the perpendicularity of its direction to that level,) still obtain.

558. Oblique regular reflexion from rough surfaces. It is evident that the inequalities in the mamillary surfaces above described will become more considerable as their radii are diminished, or as the interval of their centres is greater, and in proportion will the regularity of reflexion and refraction be interrupted. Hence too it follows, that the more oblique the incidence of the ray, the greater may be the roughness of the surface which will give a regular reflexion; and this is perfectly consonant to fact, as may be easily tried with a piece of emiered glass, which, although so rough as to give no regular image at a perpendicular incidence, will yet give a pretty distinct one at great obliquities. The reasons are, first, that a very oblique ray will not require to penetrate so far within the sphere of repulsion, to have its motion perpendicular to the surface destroyed; and, secondly, that it cannot pass between two contiguous elevations or depressions of the imaginary surface $\alpha\beta\gamma\delta$, but by reason of its obliquity must traverse several of them, and thus undergo a more regular average exertion of the forces of the medium.

559. Regular refraction at surfaces artificially polished. Thus the reflexion of light is explained on the Newtonian doctrine. But it may still be asked, how refraction at a surface artificially polished can ever be regular. In reflexion, the ray never reaches the asperities of the surface; it undergoes their average action, equalized by distance, and mutually compensated. In refraction, it is otherwise. Here the rays must actually traverse the surface, and must therefore actually pass through all its inequalities at every possible angle of obliquity. The answer to this is equally plain. Neither refraction nor reflexion are performed close to the surface, either wholly, or in great part. The greater part by far of the flexure of the ray is performed (either internally or externally) at a distance, out of the reach of these irregularities, and by the action of a much more considerable thickness of the medium than is occupied by them. Their action must be compared to the effect of mountains on the earth's surface in disturbing the general force of gravity. A stone let fall close to one of them, from a moderate height, follows not the true vertical but the direction of the deviated plumbline, which is sensibly different. Whereas, if let fall from the moon to the earth's centre, it would pass among them, were they greater a thousand fold than they are, without experiencing any sensible perturbation or change of direction in their neighbourhood.

560. In fact, however, no regular refraction can be obtained from surfaces sensibly rough, at all comparable to the regularity of their reflexion. This may arise from the impossibility of a refracted ray penetrating the surface at a sufficient degree of obliquity. It is, however, a remarkable fact, that the regular internal reflexion from a roughened surface, even at extreme obliquities, is scarcely sensible, even in cases where the external reflexion at the same obliquities is perfectly regular and copious. This would seem to indicate, that the forces which operate the external reflexion of a ray exert their energy wholly without the medium.

561. Intensity of the forces producing refraction. Whatever be the forces by which bodies reflect and refract light, one thing is certain, that they must be incomparably more energetic than the force of gravity. The attraction of the earth on a particle near its surface produces a deflexion of only about 16 feet in a second; and, therefore, in a molecule moving with the velocity of light, would cause a curvature, or change of direction, absolutely insensible in that time. In fact, we must consider, first, that the time during which the whole action of the medium takes place, is only that within which light traverses the diameter of the sphere of sensible action of its molecules at the surface. To allow so much as a thousandth of an inch for this space is beyond all probability, and this interval is tra-

versed by light in the $\frac{1}{12,672,000,000,000}$ part of a second. Now, if we suppose the deviation produced

by refraction to be 30° , (a case which frequently happens,) and to be produced by a uniform force acting during a whole second; since this is equivalent to a linear deflexion of 200,000 miles $\times \sin 30^\circ$, or of 100,000 miles = 33,000,000 \times 16 feet, such a force must exceed gravity on the earth's surface 33,000,000 times. But, in fact, the whole effect being produced not in one second, but in the small fraction of it above mentioned, the intensity of the force operating it (see MECHANICS) must be greater in the ratio of the square of one second to the square of that fraction; so that the least improbable supposition we can make gives a mean force equal to $4,969,126,272 \times 10^{24}$ times that of terrestrial gravity. But in addition to this estimate already so enormous, we have to consider that gravity on the earth's surface is the resultant attraction of its whole mass, whereas the force deflecting light is that of only those molecules immediately adjoining to it, and within the sphere of the deflecting forces. Now a sphere of $\frac{1}{1000}$ of an inch diameter, and of the mean density of

the earth, would exert at its surface a gravitating force only $\frac{1}{1000} \times \frac{1 \text{ inch}}{\text{diameter of the earth}}$ of ordinary gravity, so that the actual intensity of the force exerted by the molecules concerned cannot be less than $\frac{1000 \times \text{earth's diameter}}{1 \text{ inch}}$ (= 46,352,000,000) times the above enormous number, or upwards of 2×10^{24}

when compared with the ordinary intensity of the gravitating power of matter. Such are the energies concerned in the phenomena of light on the Newtonian doctrine. In the undulatory hypothesis, numbers not less immense will occur; nor is there any mode of conceiving the subject which does not call upon us to admit the exertion of mechanical forces which may well be termed infinite.

562. Dr. Wollaston has proposed the observation of the angle at which total reflexion first takes place at the common surface of two media, the index of refraction of one of which is known, as a means of determining that of the other; and, in the *Philosophical Transactions* for 1802, has described an ingenious apparatus which gives a measure of the index required almost by inspection. If we lay any object under the base of a prism

Light.

of flint glass with air alone interposed, the internal angle of incidence at which the visual ray begins to be totally reflected, and at which of course the object ceases to be seen by refraction is about $39^{\circ} 10'$; but when the object has been dipped in water, and brought into contact with the glass, it continues visible (while the eye is depressed) by means of the greater refractive power of the water, as far as $57\frac{1}{2}^{\circ}$ of incidence. When any kind of oil, or any resinous cement, is interposed, this angle is still greater, according to the refractive power of the medium employed; and by cements that refract more strongly than the glass, the object may be seen through the prism at whatever angle of incidence it is viewed. All that is requisite, then, to determine the refractive index of any body less refractive than glass, is to bring the substance to be examined in optical contact with the base of a prism, and to depress the eye (or increase the angle of incidence) till it ceases to be seen as a dark spot on the silvery reflexion of the sky on the rest of the base. With fluids and soft solids, or fusible ones, the requisite contact is easily obtained; but with solids, they must be brought to smooth surfaces, and applied to the base by the intervention of some fluid or cement of higher refractive power than the glass, which (since the surfaces of the interposed stratum are parallel) will produce no change in the total deviation of a ray passing through it, and therefore no error in the result. By this method, opaque as well as transparent substances may be examined, or bodies of unhomogeneous density, as the crystalline lens of the eye. It may seem paradoxical to speak of the refractive power of an opaque body; but it will be remembered, that opacity is merely a consequence of intense absorbent power, and that before a ray can be absorbed, it must enter the medium, and of course obey the laws of refraction at its surface. By this method, Dr. Wollaston has determined the refractions of a great variety of bodies; but Dr. Brewster remarks, that the method must be liable to some source of inaccuracy, which renders it unsafe to trust entirely to it in practice. Dr. Young has remarked, that the index of refraction given by it, belongs in strictness to the extreme red rays.

Part III.
Dr. Wollaston's method of determining refractive powers.

§ II. General Statement of the Undulatory Theory of Light.

The undulatory theory, among whose chief supporters we have to number Huygens, Descartes, Hooke, and Euler, and, in later times, the illustrious names of Young and Fresnel, who have applied it with singular success and ingenuity to the explanation of those classes of phenomena which present the greatest difficulties to the Corpuscular doctrine, requires the admission of the following hypotheses or postulata:

563.

1. That an excessively rare, subtle, and elastic medium, or *ether*, as it is called, fills all space, and pervades all material bodies, occupying the intervals between their molecules; and, either by passing freely among them, or, by its extreme rarity, offering no resistance to the motions of the earth, the planets, or comets in their orbits, appreciable by the most delicate astronomical observations; and having inertia, but not gravity.

Postulata in the system of undulations

2. That the molecules of the ether are susceptible of being set in motion by the agitation of the particles of ponderable matter, and that when any one is thus set in motion it communicates a similar motion to those adjacent to it; and thus the motion is propagated further and further in all directions, according to the same mechanical laws which regulate the propagation of undulations in other elastic media, as air, water, or solids, according to their respective constitutions.

3. That in the interior of refracting media the ether exists in a state of less elasticity, compared with its density, than in vacuo, (i. e. in space empty of all other matter;) and that the more refractive the medium, the less, relatively speaking, is the elasticity of the ether in its interior.

4. That vibrations communicated to the ether in free space are propagated through refractive media by means of the ether in their interior, but with a velocity corresponding to its inferior degree of elasticity.

5. That when regular vibratory motions of a proper kind are propagated through the ether, and, passing through our eyes, reach and agitate the nerves of our retina, they produce in us the sensation of light, in a manner bearing a more or less close analogy to that in which the vibrations of the air affect our auditory nerves with that of sound.

6. That as, in the doctrine of sound, the frequency of the aerial pulses, or the number of excursions to and fro from its point of rest made by each molecule of the air, determines the pitch, or note, so, in the theory of light, the frequency of the pulses, or number of impulses made on our nerves in a given time by the ethereal molecules next in contact with them, determines the *colour* of the light; and that as the absolute extent of the motion to and fro of the particles of air determine the *loudness* of the sound, so the *amplitude*, or extent of the excursions of the ethereal molecules from their points of rest, determine the brightness or intensity of the light.

The application of these postulata to the explanation of the phenomena of light, presumes an acquaintance with the theory of the propagation of motion through elastic media. This we shall assume, referring to our article on sound for the demonstration of all the properties and laws of motions so propagated, as we shall have occasion to employ. One of the principal of these is, that supposing the elastic medium uniform and homogeneous, all motions of whatever kind are propagated through it in all directions with one and the same uniform velocity, a velocity depending solely on the elasticity of the medium as compared with its inertia, and bearing no relation to the greatness or smallness, regularity or irregularity of the original disturbance. Thus, while the intensity of light, like that of sound, diminishes as the distance from its origin increases, its velocity remains invariable, and thus, too, as sounds of every pitch, so light of every colour, travels with one and the same velocity, either in vacuo, or in a homogeneous medium.

564.

The velocity of all undulations equal.

Now here arises, *in limine*, a great difficulty; and it must not be dissembled, that it is impossible to look on

Light.
Objection
from the
phenomena
of disper-
sion.

it in any other light than as a most formidable objection to the undulatory doctrine. It will be shown presently that the deviation of light by refraction is a consequence of the difference of its velocities within and without the refracting medium, and that when these velocities are given the amount of deviation is also given. Hence it would appear to follow unavoidably, that rays of all colours must be in all cases equally refracted; and that, therefore, there could exist no such phenomenon as dispersion. Dr. Young has attempted to gloss over this difficulty, by calling in to his assistance the vibrations of the ponderable matter of the refracting medium itself, as modifying the velocity of the ethereal undulations within it, and that differently according to their frequency, and thus producing a difference in the velocity of propagation of the different colours; but to us it appears with more ingenuity than success. We hold it better to state it at once in its broadest terms, and call on the reader to suspend his condemnation of the doctrine for what it *apparently* will not explain, till he has become acquainted with the immense variety and complication of the phenomena which it will. The fact is, that neither the corpuscular nor the undulatory, nor any other system which has yet been devised, will furnish that complete and satisfactory explanation of *all* the phenomena of light which is desirable. Certain admissions must be made at every step, as to modes of mechanical action, where we are in total ignorance of the acting forces; and we are called on, where reasoning fails us, occasionally for an exercise of faith. Still, if we regard hypotheses and theories as no other way valuable than as means of classifying and grouping together phenomena, and of referring facts to laws which, though possibly empirical, are yet, so far as they are so, correct representations of nature, and as such must be deducible from real primary laws, whenever they shall be discovered, we cannot but admit their importance. The undulatory system especially is necessarily liable to considerable obscurities; as the doctrine of the propagation of motion through elastic media is one of the most abstruse and difficult branches of mathematical inquiry, and we are therefore perpetually driven to indirect and analogical reasoning, from the utter hopelessness of overcoming the mere mathematical difficulties inherent in the subject when attacked directly.

566.
Objection
from the
rectilinear
propagation
of light
answered.

It is thus that we are encountered at the very outset of its application with another objection, which, in the eyes of Newton, appeared decisive against its admission, but which has since been, in a considerable degree, overcome. How is it that *shadows* exist. Sounds make their way freely round a corner,—why does not light do so? A vibration propagated from a centre in an elastic medium, and intercepted by an immovable obstacle having a small orifice, ought to spread itself, it is said, from this orifice beyond the screen as from a new centre, and fill the space beyond with undulations propagated from it in every direction. Thus, as in Acoustics, the orifice is heard as a new source of sound; so, in Optics, it ought to be seen in all directions as a new luminary. To this the answer is, first, that it is not demonstrable that a vibratory motion communicated to one particle of an elastic medium is propagated with equal *intensity* to every surrounding molecule in whatever direction situated with respect to the line of its motion, though it is with equal *rapidity*; and therefore that we have no reason to presume, *a priori*, but rather the contrary, that the motions of the vibrating particles at the orifice should be propagated laterally with *equal* intensity in all directions; secondly, that it is not true, in fact, that sounds are propagated round the corner of an obstacle *with the same intensity* as in their original direction, as any one may convince himself by the following simple experiment. Take a common tuning fork, and, holding it (when set in vibration) about three or four inches from the ear, with its flat side towards it, when its sound is distinctly heard, let a strip of card, somewhat longer than the flat of the tuning fork, be interposed, at about half an inch from the fork. The sound will be almost entirely intercepted by it; and if the card be alternately removed and replaced in pretty quick succession, alternations of sound and silence will be perceived; proving that the undulations of the air are by no means propagated with equal intensity by the circuitous route round the edge of the card, as by the direct one. Indeed any one has only, to be convinced of the fact, to attend to the sound of a carriage in the act of turning a corner from the street in which he happens to be to an adjoining one; to which we may add, that, even when there is no obstacle in the way, sounds are by no means equally audible in all directions from the sounding body, as any one may convince himself by holding a vibrating tuning fork, or pitchpipe, near his ear, and turning it quickly on its axis. This last phenomenon was first noticed, we believe, by Dr. Young, (*Phil. Trans.*, 1802, p. 25,) and since more fully described (in *Schweiggers Jahrbuch*, 1826) by M. Weber. Now if there be any inequality *at all* in the intensity of the direct and lateral propagation of undulations in a medium, it must arise from the constitution of the medium, and the proportion of the amplitude of the excursions of the vibrating particles to their distance from each other; and may therefore easily be conceived to differ in any imaginable degree in different media, and there is, at least, no absurdity in supposing the ether so constituted as to admit of comparatively very feeble lateral propagation. Now, thirdly, in point of fact, light does spread itself in a certain small degree into the shadows of bodies, out of its strict rectilinear course, giving rise to the phenomena of *inflexion* or *diffraction*, of which more presently, and which are completely accountable for on the undulatory doctrine, and form, in fact, its strongest points. For further information on this confessedly abstruse subject, the reader must consult our article on SOUND, and the works cited at the end of this Essay. It is enough here to show, that the objection which has been urged by Newton and his followers with such force against the doctrine of undulations, is really not conclusive against it, but founded rather on inadequate conceptions of the nature of elastic fluids, and the laws of their undulations.

567.
Mode in
which the
retina is
excited by
vibrations
of ether.

Although any kind of impulse, or motions regulated by any law, may be transferred from molecule to molecule in an elastic medium, yet in the theory of light it is supposed that only such primary impulses as recur according to regular periodical laws, at equal intervals of time, and repeated many times in succession, can affect our organs with the sensation of light. To put in motion the molecules of the nerves of our retina with sufficient efficacy, it is necessary that the almost infinitely minute impulse of the adjacent ethereal molecules should be often and regularly repeated, so as to multiply, and, as it were, concentrate their effect. Thus, as a great pendulum may be set in swing by a very minute force often applied at intervals exactly equal to its time

Light.

Part III.

of oscillation, or as one elastic solid body can be set in vibration by the vibration of another at a distance, propagated through the air, if in exact unison, even so may we conceive the gross fibres of the nerves of the retina to be thrown into motion by the continual repetition of the ethereal pulses; and such only will be thus agitated, as from their size, shape, or elasticity are susceptible of vibrating in times exactly equal to those at which the impulses are repeated. Thus it is easy to conceive how the limits of visible colour may be established; for if there be no nervous fibres in unison with vibrations more or less frequent than certain limits, such vibrations, though they reach the retina, will produce no sensation. Thus, too, a single impulse, or an irregularly repeated one, produces no light; and thus also may the vibrations excited in the retina continue a sensible time after the exciting cause has ceased, prolonging the sensation of light (especially of a vivid one) for an instant in the eye in the manner described, (Art. 543.) We may thus conceive the possibility of other animals, such as insects, incapable of being affected with any of our colours, and receiving their whole stock of luminous impressions from a class of vibrations altogether beyond our limits, as Dr. Wollaston has ingeniously imagined (we may almost say proved) to be the case with their perceptions of sound.

The law of motion of every particle of the ether is regulated by that of the molecule of the luminary from which it takes its origin; and will be regular or irregular, periodical or not, according as that of the original molecule is so or otherwise. But it is only with motions which may be regarded as infinitely small that we are concerned in this theory. The displacement of each particle, either of the ether or of the luminary, is supposed to be so minute as not to detach it from, or change its order of situation among the neighbouring ones. Now when we consider only such infinitesimal displacements from the position of equilibrium, it is evident, that the tension arising from them, or the force by which the displaced molecule is urged, must be proportional in quantity to its distance from its point of rest, and must tend directly to that point, provided we suppose the medium equally elastic in all directions. Hence, by the laws of Dynamics, its trajectory must be an ellipse described in one plane about the point of equilibrium as its centre; or, if one of the axes of the ellipse vanish, a straight line having that point in its middle, in which it oscillates to and fro, performing all its excursions in the latter case, or its revolutions in the former, whether great or small, in equal times, and following the law of a vibrating pendulum. We will, for the present, consider the case of rectilinear vibrations as the most simple, and show hereafter how the more general one may be reduced to it.

Proposition. To define the motion of a vibrating molecule of a luminary, supposing its excursions to and fro to be performed in straight lines.

Putting x for its distance from its point of rest, t for the time elapsed since a given epoch, and v for its velocity, and E for the absolute elastic force, the force urging the molecule to its point of equilibrium will be $E \cdot x$, and will tend to diminish x ; hence (supposing gravity to be represented by $32\frac{1}{2}$ feet) we must have

$\frac{dv}{dt} = -\frac{d^2x}{dt^2} = E x$, and therefore $\frac{2 d^2x \cdot dx}{dt^2} = -2 E x dx$, or, integrating, $\frac{dx^2}{dt^2}$ or $v^2 = E (a^2 - x^2)$ where a is the greatest distance of excursion, or the *semiamplitude* of the vibration. Hence,

$v = \sqrt{E} \cdot \sqrt{a^2 - x^2} = \frac{-dx}{dt}$, and therefore $dt = -\frac{dx}{\sqrt{E} \cdot \sqrt{a^2 - x^2}}$; or, integrating, $t + C =$

$\frac{1}{\sqrt{E}} \arccos \frac{x}{a}$, that is

$$x = a \cdot \cos \{ \sqrt{E} \cdot (t + C) \}; \quad v = a \cdot \sqrt{E} \cdot \sin \{ \sqrt{E} (t + C) \}$$

Such are the velocity and distance from the middle point of its vibration of the molecule at any instant. If we call T the whole period in which the molecule has performed one complete evolution, consisting of a complete excursion to and fro on both sides of its point of equilibrium, we shall have at the commencement of the motion when $v = 0$, or $x = a$, $a \cdot \cos \{ \sqrt{E} \cdot (t + C) \} = a$, or $(t + C) \sqrt{E} = 0$; and when one quarter of a period has been performed, or the molecule has arrived at its greatest distance $-a$ on the opposite side of the centre $-a = a \cdot \cos \{ \sqrt{E} (t + \frac{1}{2}T + C) \}$, or $\sqrt{E} \cdot (t + C + \frac{1}{2}T) = \pi$, putting π for the semicircumference of a circle whose diameter is 1. Hence we get by subtraction

$$\frac{1}{2}T \cdot \sqrt{E} = \pi; \quad \text{or} \quad T = \frac{2\pi}{\sqrt{E}}.$$

Hence we may eliminate E , and introduce T instead of it, which will give the equations $\sqrt{E} = \frac{2\pi}{T}$,

$$x = a \cdot \cos 2\pi \cdot \frac{t + C}{T}; \quad v = a \sqrt{E} \sin 2\pi \cdot \frac{t + C}{T};$$

which equations express the laws required, and which if the time t be supposed to commence at the moment when $v = 0$, or when the molecule is at the extremity of one of its excursions, become simply

$$x = a \cdot \cos 2\pi \cdot \frac{t}{T}; \quad v = a \sqrt{E} \sin 2\pi \cdot \frac{t}{T}.$$

568.
Motion of a
vibrating
luminous
molecule.

569.
Laws of
rectilinear
vibrations.

Light.

570.

Corol. Hence the excursions of the molecule to and fro will consist of four principal phases, in each of which its motion is similar, but in contrary directions, or on contrary sides of the centre. In the first phase the molecule is to the *right* of the centre of motion, and is approaching the centre, or moving from *right to left*. In the second, it is to the left of the middle point, and moving from it, or still *from right to left*. These two phases we shall term the *positive phases*. In the third phase the molecule lies on the left side, and its motion is towards the centre, and from left to right. In the fourth, it is to the right again, receding from the centre, and moving still from left to right. These we shall term the *negative phases* of its vibration.

571.

Laws of
rectilinear
vibrations
of an
ethereal
molecule.

Proposition. To define the rectilinear vibrations of any molecule of the ether, propagated from a luminous particle vibrating as in the last proposition.

In the propagation of motions through elastic, uniform media, the same or a similar motion to that of any one molecule is communicated to every other in succession; but this communication occupies time, and the motion of a molecule at a distance from the origin of the vibrations does not commence till after the lapse of an interval of time proportional to that distance, being the time in which the propagated impulse, whether of sound or light, &c. runs over that distance with a certain uniform velocity due to the intrinsic elasticity of the medium, and which in the case of light is about 200,000 miles per second; in that of sound about 1100 feet. And when the vibration of the original source of motion has ceased, that of the ethereal molecule does not cease on the instant, but continues for a time equal to that which elapsed before its commencement. Hence, if we call V the

velocity of light, and D the distance of the molecule from the luminous point, $\frac{D}{V}$ will be the interval between the commencement of the motion of the latter and of the former; hence $-t$ being the time elapsed at any instant since the commencement of the first positive phase of the vibration of the luminous point, $t - \frac{D}{V}$

will be the corresponding time in the case of the ethereal molecule. Thus we have, for the equations of the motions of the former,

$$x = a \cdot \cos 2\pi \cdot \frac{t}{T}; \quad v = b \cdot \sin 2\pi \cdot \frac{t}{T}; \text{ where } b = a \sqrt{E}$$

and in that of the latter

$$x = a \cdot \cos 2\pi \cdot \left(\frac{t - \frac{D}{V}}{T} \right); \quad v = \beta \cdot \sin 2\pi \cdot \left(\frac{t - \frac{D}{V}}{T} \right); \text{ where } \beta = a \sqrt{E}$$

a being the semiamplitude of the vibration, or the extent of the excursion of the ethereal molecule from its point of rest.

572.

Corol. Hence it is evident that the actual velocity of the molecules of ether may be less in any proportion than that of light; for the maximum value of v depends for its numerical magnitude solely on a , or on the amplitude of excursion, and on E , and not at all on V the velocity of propagation of the wave.

573.

Waves of
light
defined.

Corol. 2. If we suppose the luminous molecule to have made, from the commencement of its motion, any number of vibrations and parts of a vibration in the time t ; then if we consider an ethereal molecule at a distance $V \cdot t$ from it in any direction, (i. e. situated in a spherical surface whose radius is $V \cdot t$), this molecule will be just beginning to be put in motion. If we suppose another spherical surface concentric with the former, but having its radius less than the former by $V \cdot T$, which in future we shall call λ , every particle situated in this surface will have just completed one vibration, and be commencing its second, and so on. The interval between these surfaces will comprehend, arranged in spherical, concentric shells, molecules in every phase of their vibrations,—those in each shell being in the same phase. This assemblage of molecules is termed a *wave*, and as the impulse continues to be propagated forwards it is evident that the wave will continue to increase in radius, and will comprehend in succession all the molecules of the medium to infinity.

574.

Undula-
tions or
pulses.

Definition. The interval between the internal and external surface of a luminous wave is called an *undulation*, or a *pulse*, and its length is evidently $= V \cdot T = \lambda$, or the space run over by light in the time T of one complete period, or vibration of the luminous molecule. It is therefore proportional to that time.

575.

Different
colours
have dif-
ferent
lengths of
their undu-
lations.

Hence the lengths of the undulations of differently coloured rays differ *inter se*. For, by Postulate 6, the number of vibrations made in any given time by the ethereal particles determines the colour. Now the more numerous the vibrations are, *dato tempore*, the shorter their duration; hence T , which represents this duration, is *less*; and therefore λ , or the length of the undulation less for the violet than for the red rays. From experiments to be presently described, it has been found, that the lengths of the undulations in air, or the values of λ for the different rays, as also the number of times they are repeated in one second, are as in the following table.

Colours.	Length of an undulation in parts of an inch in air $\lambda =$.	Number of such undulations in an inch or $\frac{1}{\lambda}$	Number of undulations per second.
Extreme	0·0000266	37640	458,000000,000000
Red	0·0000256	39180	477,000000,000000
Intermediate	0·0000246	40720	495,000000,000000
Orange	0·0000240	41610	506,000000,000000
Intermediate	0·0000235	42510	517,000000,000000
Yellow	0·0000227	44000	535,000000,000000
Intermediate	0·0000219	45600	555,000000,000000
Green	0·0000211	47460	577,000000,000000
Intermediate	0·0000203	49320	600,000000,000000
Blue	0·0000196	51110	622,000000,000000
Intermediate	0·0000189	52910	644,000000,000000
Indigo	0·0000185	54070	658,000000,000000
Intermediate	0·0000181	55240	672,000000,000000
Violet	0·0000174	57490	699,000000,000000
Extreme	0·0000167	59750	727,000000,000000

Taking the velocity of light at 192000 miles per second.

From this table we see, that the sensibility of the eye is confined within much narrower limits than that of the ear, the ratio of the extreme vibrations being nearly 1·58 : 1, and therefore less than an octave, and about equal to a minor sixth. That man should be able to measure, with certainty, such minute portions of space and time, is not a little wonderful; for it may be observed, whatever theory of light we adopt, these periods and these spaces have a *real existence*, being, in fact, deduced by Newton from direct measurements, and involving nothing hypothetical but the names here given them.

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The direction of a ray in the undulatory system is a line perpendicular to the surface of the wave at any point. When, therefore, the vibration is propagated through an uniform ether, the wave being bounded by spherical surfaces, the direction of the ray is constant, and from the centre. Thus in this system a ray of light moves in a right line in an uniform medium.

577.
Direction of a ray.

The intensity of a ray is, of course, in some certain determinate ratio of the impulse made on the retina *dato tempore* by the ethereal molecules, and therefore in some certain ratio of their amplitudes of excursion, or their absolute velocities. The principle of the conservation of living forces requires that the amplitude of excursion of a molecule, situated at any distance from the vibrating centre, should be as the distance inversely, (see ACOUSTICS.) If then we suppose the sensation created in the retina to be as the simple *vis inertiae* of the molecules producing it, light ought to decrease inversely as the distance; if as the *vis viva*, (which is as the square of the velocity,) inversely as the square of the distance. As we know nothing of the mode in which the immediate sensation of light or sound is produced in the *sensorium*, we have no reason to prefer one of these ratios to the other *a priori*. But when we consider, that in the division of a beam of light by partial reflexion, or by double refraction, or otherwise, there is neither gain nor loss of light, (supposing the perfect transparency and polish of the medium which operates the division) so that the sum of the intensities remains constant, however the absolute velocities of the vibrating molecules may change, either in quantity, or (as in the case of reflexion, where they must be conceived to rebound from each other, mediately or immediately) in sign, the agreement of this law in all cases with that of the conservation of the *vis viva*, and its opposition in the other mentioned case to that of the uniform motion of the centre of gravity, (which would make not the *sum*, but the *difference* of the intensities constant, were the simple ratio of their velocities assumed for their measure,) (see DYNAMICS,) leaves us no choice in preferring the square of the absolute velocity, or of the amplitude of excursion of a vibrating molecule, for the measure of the intensity of the ray it propagates; and thus the observed law of the diminution of light is reconciled to the undulatory doctrine.

578.
Law of intensity of light

When the medium through which the vibrations are transmitted is not uniformly elastic, the waves will make unequal progress in different directions, according to the law of elasticity. In this case the figure of the wave will not be spherical. If we suppose the elasticity to vary by insensible gradations, as when light passes through the atmosphere, whose refracting power is variable, the figure of the wave will be flattened towards that part where the elasticity is less. Thus, in fig. 126, if A B be the earth's surface, C D, E F, G H, &c. the atmospheric strata, and S a luminous point, the waves will be less curved as they approach the perpendicular S B; and the line S, 1, 2, 3, 4, 5, &c. drawn so as to intersect them all at right angles, will be a curve convex downwards, so that a ray will appear to be continually bent downwards towards the earth, as we see really happens. Let us now proceed to consider the explanation of the phenomena of reflexion and refraction on the undulatory system.

579.
Form of the wave.

The perpendicular reflexion of light may be conceived, by the analogy of an elastic ball in motion impinging directly on another at rest, and in this way it has been illustrated by Dr. Young. If the balls be equal, the whole motion of the impinging ball will be transferred to the other, no reflexion taking place; and thus the impulse may be propagated undiminished along a line of balls as far as we please. So it is with light moving in a uniform medium, or passing from one medium to another of equal elasticity. But if a less ball impinge on

580.
Perpendicular reflexion.

Light. a greater at rest, it will be reflected, and with a momentum which is greater in proportion to the difference in size of the balls.

581. But to render an account of oblique reflexion and refraction, and the other phenomena we shall have to speak of, it will be necessary to lay down the following principles, which are either self-evident or follow immediately from the elementary principles of dynamics.

582. 1. When any number of very minute impulses is communicated at once to the particles of any medium, or of any mechanical system under the influence of any forces, the motion of each particle at any instant will be the sum of all the motions which it would have at that instant, had each of the impulses been communicated to the system alone, (the word sum being understood in its algebraical sense.)

583. 2. Every vibrating molecule in an elastic medium, whether vibrating by an original impulse, or in consequence of an impulse propagated to it from others, may be regarded as a centre of vibration from which a system of secondary waves emanates in all directions, according to the laws of the propagation of waves in the medium.

Principle of secondary waves. *Proposition.* In the reflexion of light on the undulatory doctrine, the angle of incidence is equal to that of reflexion.

584. Let AB be a plane surface separating the two media, and S the luminous point propagating a series of spherical waves, of which let Aa be one. So soon as this reaches the surface at A , a partial reflexion will take place; and regarding the point A as a new centre of vibration, spherical waves will begin to be propagated from it as a centre, one of which proceeds forwards into the reflecting medium, with a velocity greater or less than that of the incident wave, as the case may be; the other backwards into the medium of incidence, with a velocity equal to that of the incident wave. It is only with the latter we are at present concerned. Conceive now the wave Aa to move forward into the position Bb ; then in the time that it has run over the space PB , the wave propagated from A will have run back over a distance $Ad = PB$, and the hemisphere whose radius is Ad will represent this wave. Between A and B take any point X , and describe the hemispheric surface Xc . Then regarding X as a centre of vibration, its vibrations will not commence till the wave has reached it. It will, therefore, begin to vibrate later than A , by the whole time the wave Aa takes to run over PQ ; but when once set in vibration, it propagates backwards a spherical wave with the same velocity, so that when the original wave has advanced into the situation Bb , the wave from X will have expanded into a hemisphere, whose radius Xc is equal to PB , $-PQ$, or AB . Now this being true of every point X , if we conceive a surface touching all these hemispheres in d, c, B , this surface will mark the points at which the reflected impulse has just arrived, and which just begins to move when the original wave has reached B , and will, therefore, be the surface of the reflected wave. Conceive now the spherical surface bB prolonged below the plane AB , as represented by the dotted line DCB , and the same of the spheres about A and X . Then the spherical surfaces DCB and Cc being both perpendicular to SXC , must touch each other in C , hence the surface touching all the hemispheres about A, X , &c. as centres, below AB is a segment of a sphere having S for a centre, and therefore the surface Bcd or the reflected wave is a segment of a sphere having its centre at s as much below the line AB as S is above it.

Now to an eye placed at X , the luminous point S will appear in the direction SX perpendicular to the incident wave, and the eye placed in c will perceive the reflected image of S at s in the direction cs , perpendicular to the reflected wave; but cs passes through X , because the spheres cC and Bb touch at c . Therefore the ray by which s is seen at c passes through X . But the surfaces BD, Bd being similar and equal, the angle $BXc = BXC = AXS$, that is, the angle of incidence is equal to that of reflexion. $Q. E. D.$

585. *Cor.* If the reflecting surface be not a plane, the reflected wave will not be spherical; its form is, however, easily determined as follows: Suppose the direct wave to have assumed the position Bb . Take any point X in the reflecting surface, and describe the sphere XQ , and with the centre X and radius $= BQ$, describe another sphere. Do this for every point in the surface AB , and the surface which is a common tangent (as Bcd) to all these spheres, is the surface of the reflected wave, because it marks the farthest limit to which the reflected impulse has reached in all directions at the instant when the direct impulse has reached B . Now take Y infinitely near to X , and, making the same construction at Y , let e, e be the points in the reflected wave to which Xc and Ye are respectively perpendicular. Draw Xr perpendicular to Ye , and Xq to SYq , then, since $Ye = SB - SY$, and $Xc = SB - SX$, we have $Ye - Xc$, or $Yr = SX - SY = Yq$, and XY being common to the right angled triangles XYr, XYq , the angle rYX must be equal to XYq or to SYA , so that the same law of reflexion holds good in curve as in plane surfaces.

586. *Proposition.* To demonstrate the law of refraction in the undulatory system.

Let S , fig. 129, be a luminous point, and let any wave propagated from it reach in succession the points Y, X, B of any curve surface YXB of a refracting medium, whereof X and Y are supposed infinitely near each other. As the wave strikes Y, X, B , each of these points will become centres of undulation, which will be propagated in the refracting medium with a velocity different from that of light in the medium of incidence, by reason of their different elasticities, (Postulate 3.) Let $V : v ::$ velocity in the first medium to that in the second, (a constant ratio by hypothesis,) and, describing the sphere BQR , take $Xc = \frac{v}{V} \cdot QX$ and $Ye =$

$\frac{v}{V} \cdot YR$, then will Xc and Ye represent the spaces run over by the refracted secondary waves propagated from X and Y respectively, when the direct wave has reached B . Hence, if about X and Y as centres, and with these radii we describe spheres, and suppose e, c to be points in the curve surface which is a tangent to all such spheres, it is clear that Xc and Ye will be perpendicular to this surface, that is, to the surface of the refracted primary wave; hence, Xc and Ye will be the directions of the refracted rays at X and Y . Draw Xq, Xr perpendicular

Light. respectively to YR and Ye , then will $Yq = SX - SY$ and $Yr = Ye - Xc = \frac{v}{V} \cdot YR - \frac{v}{V} \cdot XQ = \frac{v}{V}$ Part III.

$$(YR - XQ) = \frac{v}{V} \{ (SR - SY) - (SQ - SX) \} = \frac{v}{V} \cdot (SX - SY) = \frac{v}{V} \cdot Yq. \text{ Hence we have } Yq \cdot$$

$Yr :: V : v$. But since SX, SY are direct rays, and Xc, Ye the corresponding refracted ones, therefore SXY is the complement of the angle of incidence of SX , and, consequently, YXq is equal to the angle of incidence itself, and XYr will be the complement of the angle of refraction, and therefore $YXr (= 90^\circ - XYr) =$ the angle of refraction of SY , or, (since the points Y, X are infinitely near each other,) of SX , hence we have

$$Yq : XY :: \sin \text{ incidence} : 1, \\ XY : Yr :: 1 : \sin \text{ refraction}.$$

And compounding $Yq : Yr :: \sin \text{ incidence} : \sin \text{ refraction}$. But we proved before, that $Yq : Yr$ in the constant ratio of $V : v$; therefore the sine of incidence : that of refraction in the same constant ratio. Q. E. D.

Corollary 1. In the cases both of reflexion and refraction, the undulation is propagated from the luminous point to any other point in the least possible time. For the surface both of the reflected and refracted waves mark the extreme limits to which the impulse has been propagated by reflexion or refraction in a given time. The undulation propagated from X (fig. 127) in any other direction than Xc , as, for instance, $X\gamma$, will fall short of the surface Bcd , and the point γ therefore will have been reached, and passed by the reflected or refracted primary wave in the situation $\beta\gamma\delta$, before it can be reached by the secondary undulation propagated from X in the direction $X\gamma$. 587.

Corollary 2. This property in the undulatory system corresponds to the principle of least action in the corpuscular doctrine, and may be thus stated generally : 588.

A reflected or refracted ray will always pursue such a course as would be described in the least possible time, by a point moving from the point of its departure to that of its arrival, with the velocities corresponding to the media in which it moves, and the direction of its motion. Law of swiftest propagation.

It is evident that this is general, and applies to cases where the medium is either of variable elasticity, or has different elasticities in different directions; for the ray is by definition a perpendicular to the surface of the wave, or to a surface, the locus of all the molecules in the medium, which are just attained by the undulation, and just commencing their vibration, so that the reasoning of Corol. 1, applies equally to all cases. 589. Applies generally.

The properties of foci and Caustics flow with such elegance and simplicity from this doctrine, that it would be unpardonable not to instance its application to that part of the theory of Optics. 590.

Definition. A focus is a point at which the same wave arrives at the same instant from more than one point in a surface. Foci in the undulatory system. Defined.

It is evident, that when this is the case, the ethereal molecules in the focus will be agitated by the united force of all the undulations which reach them in the same phase at the same instant, and will be proportionally more violent as the focus is common to a greater number of points, and the light in the focus will be proportionally more intense.

Proposition. Required to determine the nature of the surface which shall refract all rays from one point rigorously to one focus. Let F (fig. 129) be the focus, then will every part of a wave propagated from S and refracted at the surface AB , reach F at the same instant; therefore time of describing SX with velocity $V +$ time of describing XF with velocity v is constant for every point in the surface. Or, 591.

$$\frac{SX}{V} + \frac{FX}{v} = \text{constant, or } SX + \mu \cdot FX = \text{constant, } \mu \text{ being the relative index of refraction.}$$

This equation then defines the nature of the curve sought, and it is easy to perceive its identity with that expressed by the equation (n) Art. 232, obtained from a direct consideration of the law of refraction, but by a much more intricate process.

The intensity of the reflected or refracted ray cannot be computed generally in the present very imperfect state of our knowledge of the theory of waves. M. Poisson, however, in the case of perpendicular incidence, and on the particular hypothesis of the luminous vibrations being performed in the direction of the ray itself, has succeeded in investigating the comparative intensities of the incident, reflected, and transmitted rays. His results are as follows: 'Taking μ, μ' for the absolute refractive indices of the media, he finds (on the supposition that the intensity of light is as the square of the absolute velocity of the vibrating molecules) : 592. Intensity of a ray reflected perpendicularly.

Intensity of reflected ray : that of incident :: $(\mu' - \mu)^2 : (\mu' + \mu)^2$. Intensity of the intromitted ray : that of the incident :: $4\mu^2 : (\mu + \mu')^2$. Intensity of the ray intromitted from a medium whose refractive index = μ into a parallel plate of one whose refractive index = μ' , in contact at its second surface with a third whose refractive index = μ'' , reflected at their common surface, and again emergent at the first surface : intensity of the ray originally incident on the first surface :: $16\mu^2\mu'^2(\mu'' - \mu)^2 : (\mu' + \mu')^4 \cdot (\mu' + \mu'')^2$. And, lastly, the intensity of the ray transmitted through the parallel plate of the second medium into the third : that of the original incident ray :: $16\mu^2\mu'^2 : (\mu + \mu')^2 \cdot (\mu' + \mu'')^2$ which (in the case where the third medium is the same as the first, becomes $16\mu^2\mu'^2 : (\mu + \mu')^4$.

These results of M. Poisson, so far as they have been hitherto satisfactorily compared with experiment, manifest at least a general accordance, and the undulatory doctrine thus furnishes a plausible explanation of the connection of the reflecting power of a medium with its refractive index, and of the diminished reflection at the common surfaces of media in contact.—They have been in great measure (it should be observed) anticipated by Dr. Young, in his Paper on Chromatics, (*Encyclop. Brit.*) by reasoning which M. Poisson terms indirect, but which, we confess, appears to us by no means to merit the epithet. 593.

Light. If photometrical experiments enable us to determine the proportion of the reflected to the incident light, we may thence conclude the index of refraction of the reflecting medium, and that in cases where no other mode will apply. Thus, M. Arago having ascertained that about half the incident light is reflected at a perpendicular incidence from mercury, we have in this case $\left(\frac{\mu' - \mu}{\mu' + \mu}\right)^2 = \frac{1}{2}$; $\frac{\mu'}{\mu} = 5.829$ for the refractive index of mer-

594.
Applied to
determine
refractive
indices.

cury out of air; and this is perfectly consonant to the general tenor of optico-chemical facts, which assign to the heavy and especially to the white metals (as indicated in their transparent combinations) enormous refractive and dispersive powers. This curious and interesting application has not been overlooked by Dr. Young in the Paper alluded to.

595. To complete the theory of reflexion and refraction on the undulatory hypothesis, it will be necessary to show what becomes of those oblique portions of the secondary waves, diverging in all directions from every point of the reflecting or refracting surfaces (as X γ, fig. 127) which do not conspire to form the principal wave. But to understand this, we must enter on the doctrine of the interference of the rays of light,—a doctrine we owe almost entirely to the ingenuity of Dr. Young, though some of its features may be pretty distinctly traced in the writings of Hooke, (the most ingenious man, perhaps, of his age,) and though Newton himself occasionally indulged in speculations bearing a certain relation to it. But the unpursued speculations of Newton, and the *apperçus* of Hooke, however distinct, must not be put in competition, and, indeed, ought scarcely to be mentioned with the elegant, simple, and comprehensive theory of Young,—a theory which, if not founded in nature, is certainly one of the happiest fictions that the genius of man has yet invented to group together natural phenomena, as well as the most fortunate in the support it has unexpectedly received from whole classes of new phenomena, which at their first discovery seemed in irreconcilable opposition to it. It is, in fact, in all its applications and details one succession of *felicities*, insomuch that we may almost be induced to say, if it be not true, it deserves to be so. The limits of this Essay, we fear, will hardly allow us to do it justice.

§ III. Of the Interference of the Rays of Light.

596.
General
principles of
interference

The principle on which this part of the theory of Light depends, is a consequence of that of the “Superposition of small motions” laid down in Art. 583. If two waves arrive at once at the same molecule of the ether, that molecule will receive at once both the motions it would have had in virtue of each separately, and its resultant motion will, therefore, be the diagonal of a parallelogram whose sides are the separate ones. If, therefore, the two component motions agree in direction or very nearly so, the resultant will be very nearly equal to their sum, and in the same direction. If they very nearly oppose each other, then to their difference. Suppose, now, two vibratory motions consisting of a series of successive undulations in an elastic medium, all similar and equal to each other, and indefinitely repeated, to arrive at the same point from the same original centre of vibration, but by different routes (owing to the interposition of obstacles or other causes) exactly, or very nearly in the same final direction; and suppose, also, that owing either to a difference in the lengths of the routes, or to a difference in the velocities with which they are traversed, the time occupied by a wave in arriving by the first route (A) is less than that of its arriving by the other (B). It is clear, then, that any ethereal molecule placed in any point common to the two routes A, B, will begin to vibrate in virtue of the undulations propagated along A, before the moment when the first wave propagated along B reached it. Up, then, to this moment its motions will be the same as if the waves along B had no existence. But after this moment, its motions will be very nearly the sum or difference of the motions it would have separately in virtue of the two undulations each subsisting alone, and the more nearly, the more nearly the two routes of arrival agree in their final direction.

597.
Case of
complete
discordance

Now it may happen, that the difference of the lengths of the routes or the difference of velocities is such, that the waves propagated along B shall reach the intersection exactly one-half an undulation behind the others, *i. e.* later by exactly half the time of a wave running over a space equal to a complete undulation. In that case, the molecule which in virtue of the vibrations propagated along A would (at any future instant) be in one phase of its excursions from its point of rest, would, in virtue of those propagated along B, if subsisting alone, be at the same instant in exactly the opposite phase, *i. e.* moving with equal velocity in the contrary direction. (See Art. 570.) Hence, when both systems of vibration coexist the motions will constantly destroy each other, and the molecule will remain at rest. The same will hold good if the difference of routes or velocities be such, that the vibrations propagated along B shall reach the intersection of the routes exactly $\frac{1}{2}$, $\frac{3}{2}$, $\frac{5}{2}$, &c. of a complete period of undulation after those propagated along A; for the similar phases of vibration recurring periodically, and being (by hypothesis) continually repeated for an indefinite time, it is no matter whether the first vibration propagated along B be superimposed on, or *interfere with* (as it is called) the first, or any subsequent one propagated along A, provided the difference of their phases be the same.

598.
Case of
complete
accordance.

On the other hand it may happen, that the waves propagated along B do not reach the intersection till exactly one, two, or more whole periods after the corresponding waves propagated along A. In this case, the molecule at the intersection will, at any instant subsequent to the time of arrival of the first wave along B, be agitated at once by both vibrations in the *same* phase, and therefore the velocity and amplitude of its excursions will, instead of being destroyed, be doubled.

Light.

Lastly, it may happen, that the difference of the times of arrival of the corresponding waves is neither an exact even, or odd multiple of half a complete period of undulation. In that case, the molecule will vibrate with a joint motion, less than double what it would have in virtue of either separately.

Part III.

599.

An apt illustration of the case of interference here described, may be had by considering the analogous case in the interference of waves on the surface of water. Conceive, for instance, two equally broad canals A and B to enter two canals at right angles into the side of a reservoir, at both whose apertures, from an origin at a great distance, a wave arrives at the same instant, and runs along the two canals with equal, uniform velocities. Let their sides be perfectly smooth, and their breadths everywhere equal, but let them be led, by a gentle curvature, to meet in a point at some distance, and, the curvature of B being supposed somewhat greater than that of A, let the distance from their intersection to the reservoir, measured along B, be greater than along A. It is obvious, that (if we consider only a single wave) the portion of it propagated along A will reach the intersection first, and after it that propagated along B, so that the water at that point will be agitated by two waves in succession. But, let the original cause of undulation be continually repeated so as to produce an indefinite series of equal and similar waves. Then, if the difference of lengths of the two canals be just equal to half the interval between the summits of two consecutive waves, it is evident that when the summit of any wave propagated along A has reached the intersection, the depression between two consecutive summits (*viz.* that corresponding to the wave propagated along A, and that of the wave immediately preceding it) will arrive at the intersection by the course B. Thus, in virtue of the wave along A the water will be raised as much above its natural level, as it will be depressed below it by that along B. Its level will, therefore, be unchanged.—Now as the wave propagated along A passes the intersection, it subsides, from its maximum, by precisely the same gradations as that along B, passing it with equal velocity, rises, from its minimum, so that the level will be preserved at the point of intersection, undisturbed so long as the original cause of undulation continues to act regularly. So soon as it ceases, however, the last half wave which runs along B will have no corresponding portion of a wave along A to interfere with, and will, therefore, create a single fluctuation at the point of concurrence.

600.
Illustration
from waves
propagated
along canals

In the theory of the interferences of light we may disregard these commencing and terminal, uncompensated undulations, and parts of undulations, as being so few in number as to excite no impression on the retina, and consider the interfering rays as of indefinite duration, or as destitute of either beginning or end.

601.

Initial and
terminal vi-
brations dis-
regarded.

According to the foregoing reasoning then it appears, that if two rays having a common origin, *i. e.* forming parts of one and the same system of luminous waves proceeding from a common centre, be conducted by different routes to one point which we will suppose to be situated on a white screen, or on the retina of the eye, they will there produce a bright point, or the sensation of light, if their difference of routes be an even multiple of the length of half an undulation and a dark one; or the sense of darkness, if an odd multiple of it; and if intermediate, then a feebler or a stronger sense of light, as the difference of routes approximates to one or the other of these limits. That two lights should in any case annihilate each other, and produce darkness, appears a strange paradox, yet experiment confirms it; and the fact was observed, and broadly stated by Grimaldi long before any plausible reason could be given of it.

602.

Mutual an-
nihilation of
two rays of
light in op-
posite
phases.

Having thus obtained a general idea of the nature of interferences, let us now endeavour to subject their effects to a more strict calculation. To this end it will be necessary to fix with precision the sense of some words hitherto used rather loosely.

603.

Definition. The *phase* of an undulation affecting any given molecule of ether at any instant of time, is numerically expressed by an arc of a circle to radius unity, increasing proportionally to the time—commencing at 0 when the molecule is at rest at its greatest positive distance of excursion, and becoming equal to one circumference when the molecule, after completing the whole of a vibration, returns again to the same state of

604.

Definition.
Phase.

rest at the same point. Thus, in the equation $v = a \cdot \sqrt{E} \sin \left(2\pi \cdot \frac{t+C}{T} \right)$, $2\pi \cdot \frac{t+C}{T}$ is the phase of the undulation at the instant t .

Definition. The *amplitude* of vibration of a ray or system of waves is the coefficient a , or the maximum excursion from rest, of each molecule of the ether in its course.

605.

Amplitude
of a ray.

Corol. The intensity of a ray of light is as the square of the amplitude of the vibrations of the waves of which it consists.

Definition. Similar rays, or systems of luminous waves, are such as have the vibratory motions of the ethereal molecules which compose them regulated by the same laws, and their vibrations performed in equal times, and the curves or straight lines they describe in virtue of them, similar and similarly situated in space, so that the motions of any two corresponding molecules in each, shall at every instant of time be parallel to each other.

606.

Similar rays.

Corol. Similar rays have the same colour.

Definition. The *origin* of a ray, or a system of waves, is the vibrating material centre from which the waves begin to be propagated, or more generally, a fixed point in its length, at which an ethereal molecule, at an assumed epoch, was in the phase 0 of its undulation.

607.

Origin of a
ray.

Corol. Two systems of interfering waves having their origins distant by an exact number of undulations, may be regarded as having a common origin.

608.

Proposition. To find the origin of a ray, having given the expression for the velocity of one of its vibrating molecules.

609.

To find the
origin of a
ray.

Let $a = a \cdot \sqrt{E}$, and let $v = a \cdot \sin \left(2\pi \cdot \frac{t+C}{T} \right)$ be the expression given for the velocity

of any assumed molecule (M) at the instant t . Let V represent the velocity of light, and λ the length of an undulation, and δ the distance run over by light in the time t . Then will $\delta = V t$ and $\lambda = V T$, and consequently $\frac{t}{T} = \frac{\delta}{\lambda}$. Suppose v_0 to represent the velocity of a vibrating molecule at the origin of the

ray at the instant t , then will $v_0 = a \cdot \sin \left(2 \pi \cdot \frac{t}{T} \right) = a \cdot \sin \left(2 \pi \cdot \frac{\delta}{\lambda} \right)$. But the molecule M moves only by an impulse communicated to it from the origin, and therefore all its motions are later than those at the origin by a constant interval equal to the time required for light to run over the distance of M from the origin. Call D that distance, then $\frac{D}{V}$ is the interval in question, and $t - \frac{D}{V}$ is the time elapsed at the instant t , since the

molecule commenced its periodic motions; therefore its velocity v must $= a \cdot \sin \left(\frac{2 \pi}{T} \left(t - \frac{D}{V} \right) \right)$, and con-

sequently $C = -\frac{D}{V}$, or $D = -V C$.

Hence we see that the distance of the molecule M from the origin of the ray, is equal to the space described by Light, in a time represented by the arbitrary constant C, and is therefore given when C is so, and *vice versa*.

610. *Corol.* Since $V T = \lambda$ the expression for the velocity becomes

$$v = a \cdot \sin 2 \pi \cdot \left(\frac{t}{T} - \frac{D}{\lambda} \right) = a \cdot \sin 2 \pi \cdot \left(\frac{\delta - D}{\lambda} \right) \text{ and similarly } x = a \cdot \cos 2 \pi \cdot \left(\frac{\delta - D}{\lambda} \right)$$

611. *Proposition.* To determine the colour, origin, and intensity of a ray resulting from the interference of two similar rays, differing in origin and intensity.

Let a^2 and a'^2 be the intensities of the rays, or a, a' their amplitudes of vibration, and take $a = a \cdot \sqrt{E}$, $a' = a' \cdot \sqrt{E}$, then, if we put θ for the phase of vibration of a molecule M at the instant t which it would be in,

in virtue of the first system of waves (A), and $\theta + k$ for its phase, in virtue of the other (B), $\frac{k}{2\pi}$. T will represent the time taken by light to run over a space equal to the interval of their origin, and the velocities and distances from rest which M would have, separately at the instant t , in virtue of the two rays, will be

$$v = a \cdot \sin \theta; v' = a' \cdot \sin (\theta + k), \text{ and } x = a \cdot \cos \theta; x' = a' \cdot \cos (\theta + k).$$

Therefore, in virtue of the resulting ray, it will have the velocity

$$v + v' = a \cdot \sin \theta + a' \cdot \sin (\theta + k), \text{ and } x + x' = a \cdot \cos \theta + a' \cdot \cos (\theta + k).$$

Let the former be put equal to $A \cdot \sin (\theta + B)$, the possibility of which assumption will be shown by our being able to determine A and B, so as to satisfy this condition. Then we have

$$(a + a' \cdot \cos k) \sin \theta + a' \cdot \sin k \cdot \cos \theta = A \cdot \cos B \cdot \sin \theta + A \cdot \sin B \cdot \cos \theta,$$

and equating like terms,

$$A \cdot \cos B = a + a' \cdot \cos k; A \cdot \sin B = a' \cdot \sin k,$$

whence we get, dividing one by the other,

$$\tan B = \frac{a' \cdot \sin k}{a + a' \cdot \cos k}; A = \frac{a' \cdot \sin k}{\sin B} = \sqrt{a^2 - 2 a a' \cdot \cos k + a'^2}$$

and these values being determined, A and B are known, and, therefore, $v + v' = A \cdot \sin (\theta + B)$. Similarly, if we put $x + x' = A' \cdot \cos (\theta + B')$ we obtain values of A' and B' precisely similar, writing only $a a'$ for a, a' respectively.

612. *Corol. 1.* Hence we conclude, 1st. that the resultant ray is similar to the component ones, and has the same period, i. e. the same colour.

613. *Corol. 2.* M. Fresnel has given the following elegant rule for determining the amplitude and origin of the resultant ray, which follows immediately from the value of A and the equation $\sin B = \frac{a'}{A} \cdot \sin K$ above found.

Construct a parallelogram, having its adjacent sides proportional to the amplitudes a, a' of the component rays, and the angle between them measured by a circular arc to radius unity, equal to the differences of their phases, then will the diagonal of this parallelogram represent on the same scale the amplitude of the resulting ray, and the angle included between it, and either side will represent the difference of phases between it and the ray corresponding; or, which comes to the same thing, the difference of their origins (when reduced to space.)

Light.

Part III.

Corol. 3. Thus in the case of complete discordance, the diagonal of the parallelogram vanishes, and the angle becomes 180° , or half a circumference, corresponding to a difference of origins of half an undulation. In that of complete *accordance*, the angle is 0 , or 360° , and the origins of the rays coincide, or (which comes to the same thing) differ by an exact undulation, and the diagonal is double of the side, so that the *intensity* of the compound ray is four times that of either ray singly.

Corol. 4. If the origins of two equally intense rays differ by one quarter of an undulation, the resultant ray will have its amplitude to that of either component one, as $\sqrt{2} : 1$, and, therefore, its intensity double, and its origin will differ one-eighth of an undulation from that of either. Thus in this particular case, the *brightness* of the compound ray is the sum of the *brightnesses* of the components, and its position exactly intermediate between them.

Corol. 5. Any ray may be resolved into two, differing in origin and amplitude, by the same rules as govern the resolution of forces in Mechanics.

Corol. 6. The sum of the *intensities* of the component rays exceeds that of the resultant, when their origins differ by less than a quarter of an undulation, falls short of it when the difference is between $\frac{1}{4}$ and $\frac{1}{2}$, again exceeds it when between $\frac{1}{2}$ and $\frac{3}{4}$, and so on. For the value of A' , above found, gives

$$a^2 + a'^2 - A^2 = 2 a a' \cos k;$$

now a^2 , a'^2 , and A^2 , represent the intensities of the respective rays whose momenta are a , a' , and A .

Corol. 7. In the same manner may any number of similar rays be *compounded*, and the resultant ray will be similar to the elementary rays, and *vice versa*.

Let us now consider the interference of waves having the same period (or colour) but in all other respects dissimilar.

The law of vibration of the molecules of the luminous bodies which agitate the ether, restricting their motions to ellipses performed in planes, the same will hold good of the motions of each molecule of the ether. Now every elliptic vibration, or rather revolution, performed under the influence of a force directed to its centre and proportional to the distance, is decomposed into three rectilinear vibrations, lying in any three planes at right angles to each other, each of which separately would be performed by the action of the same force in the same time, and according to the same laws of velocity, time, and space. Hence, every elliptic vibration may be expressed by regarding the place of the vibrating molecule at any instant t as determined by three coordinates x , y , z , such that, θ being an arc proportional to the time, we shall have

$$(1.) \left\{ \begin{array}{l} x = a \cdot \cos (\theta + p); - \frac{dx}{dt} = u = a \cdot \sin (\theta + p) \\ y = b \cdot \cos (\theta + q); - \frac{dy}{dt} = v = \beta \cdot \sin (\theta + q) \\ z = c \cdot \cos (\theta + r); - \frac{dz}{dt} = w = \gamma \cdot \sin (\theta + r) \end{array} \right\} (2.)$$

In fact, if we multiply the first of these equations by an indeterminate l , the second by m , and the third by n and add, we get

$$(3); \quad lx + my + nz = \cos \theta \{ la \cdot \cos p + mb \cdot \cos q + nc \cdot \cos r \} \\ - \sin \theta \{ la \cdot \sin p + mb \cdot \sin q + nc \cdot \sin r \}$$

and, therefore, if we determine l , m , n , so that

$$la \cdot \cos p + mb \cdot \cos q + nc \cdot \cos r = 0; \quad la \cdot \sin p + mb \cdot \sin q + nc \cdot \sin r = 0$$

which (being equations of the first degree only) is always possible, we shall have, independently of θ ,

$$lx + my + nz = 0; \quad (4.)$$

and this, being the equation of a plane, shows that the whole curve represented by the above equations lies in one plane. Again, if we eliminate θ between the equations, involving x and y only, we have

$$\cos^{-1} \frac{x}{a} - \cos^{-1} \frac{y}{b} = p - q,$$

or, taking the cosines on both sides,

$$\frac{xy}{ab} - \sqrt{1 - \frac{x^2}{a^2}} \cdot \sqrt{1 - \frac{y^2}{b^2}} = \cos (p - q)$$

and reducing, we get the equation

$$\left(\frac{x}{a}\right)^2 + \left(\frac{y}{b}\right)^2 - 2 \cdot \frac{x}{a} \cdot \frac{y}{b} \cdot \cos (p - q) = \sin^2 (p - q); \quad (5.)$$

which is the equation of an ellipse having the origin of the x and y in its centre, and the same is true *mutatis mutandis* of the equations between x and z , and between y and z . Thus the curve represented by the three equations between x , y , z , θ , has an ellipse about the centre for its projection on each of the planes at right angles to each other, and is, of course, itself an ellipse.

614. Cases of complete concord and discord.
615. Case of discordance by a quarter of an undulation.
616. Composition and resolution of rays.
617. Relative intensities of simple and compound rays.
618. General problem of interferences.

Light.

Suppose now two systems of waves, or two rays coincident in direction, to interfere with each other. If we accent the letters of the above expressions to represent corresponding quantities for the second system, we shall have

Part III.

$$\left. \begin{aligned} X &= x + x' = a \cdot \cos(\theta + p) + a' \cdot \cos(\theta + p') \\ Y &= y + y' = b \cdot \cos(\theta + q) + b' \cdot \cos(\theta + q') \\ Z &= z + z' = c \cdot \cos(\theta + r) + c' \cdot \cos(\theta + r') \end{aligned} \right\} \quad (6)$$

and similarly for the velocities $u + u', v + v', w + w'$. In the same manner, then, as we proceeded in the case of two similar rays, let us suppose

$$a \cdot \cos(\theta + p) + a' \cdot \cos(\theta + p') = A \cdot \cos(\theta + P)$$

and developing

$$(a \cdot \cos p + a' \cdot \cos p') \cos \theta - (a \cdot \sin p + a' \cdot \sin p') \sin \theta = A \cdot \cos P \cdot \cos \theta - A \cdot \sin P \cdot \sin \theta,$$

whence we get

$$\left. \begin{aligned} \tan P &= \frac{a \cdot \sin p + a' \cdot \sin p'}{a \cdot \cos p + a' \cdot \cos p'}; \quad A = \frac{a \cdot \sin p + a' \cdot \sin p'}{\sin P} \\ \text{or,} \quad A &= \sqrt{a^2 + 2aa' \cdot \cos(p - p') + a'^2} \end{aligned} \right\}; \quad (7)$$

Thus we have $X = A \cdot \cos(\theta + P)$, and, similarly, $Y = B \cdot \cos(\theta + Q)$, and $Z = C \cdot \cos(\theta + R)$, and a process exactly similar gives us the corresponding expressions for the velocities.

620.
Composition and resolution of vibrations generally.

Thus we see that the same rules of composition and resolution apply to dissimilar as to similar vibrations. Each vibration must first be resolved into three rectilinear vibrations in three fixed planes at right angles to each other. These must be separately compounded to produce new rectilinear vibrations in the coordinate planes, which together represent the resulting elliptic vibration, and will have the same period as the component ones. By inverting the process, a vibration of this kind may be resolved into any number of others we please, having the same period.

621.
Case of interference of rectilinear vibrations.

A great variety of particular cases present themselves, of which we shall examine some of the principal. And first, *when the interfering vibrations are both rectilinear.*

Since the choice of our coordinate planes is arbitrary, let us suppose that of the x, y to be that in which both the vibrations are performed. Of course the resulting one will be performed in the same. Therefore we may put $z = 0$, or $c = 0, c' = 0$, and content ourselves with making

$$\left. \begin{aligned} x &= a \cdot \cos(\theta + p); \quad y = b \cdot \cos(\theta + p) \\ x' &= a' \cdot \cos(\theta + p'); \quad y' = b' \cdot \cos(\theta + p') \end{aligned} \right\}; \quad (8)$$

The resultant vibration is the general elliptic.

because $\frac{x}{y}$ and $\frac{x'}{y'}$ are constant in this case, and X, Y, A, B, P, Q , denoting as in the general case, we have

$$X = A \cdot \cos(\theta + P); \quad Y = B \cdot \cos(\theta + Q);$$

and, by elimination of θ ,

$$\left(\frac{X}{A}\right)^2 + \left(\frac{Y}{B}\right)^2 - 2 \cos(P - Q) \cdot \frac{XY}{AB} = \sin^2(P - Q); \quad (9)$$

where A, B, P, Q , are determined as in equations, (7.) In the general case, then, the resulting vibration is elliptic.

622.
Case when the resultant is rectilinear.

The ellipse degenerates into a straight line by the diminution of its minor axis when $P = Q$. Now this gives $\tan P = \tan Q$, or

$$\frac{a \cdot \sin p + a' \cdot \sin p'}{a \cdot \cos p + a' \cdot \cos p'} = \frac{b \cdot \sin p + b' \cdot \sin p'}{b \cdot \cos p + b' \cdot \cos p'}$$

which, reduced, takes the form

$$\left(\frac{a'}{a} - \frac{b'}{b}\right) \cdot \sin(p - p') = 0.$$

There are, therefore, two cases, and two only in which the resulting vibration is rectilinear. The first, when $p - p' = 0$, or when the component vibrations have a common origin, or are in complete accordance; the

Case when their directions coincide.

other, when $\frac{a'}{a} = \frac{b'}{b}$, that is, when they are both performed in one plane, and in the same direction. For if

we call m and m' the amplitudes, and ψ, ψ' the angles they make with the axis of the x , we have

$$a = m \cdot \cos \psi; \quad b = m \cdot \sin \psi; \quad a' = m' \cdot \cos \psi'; \quad b' = m' \cdot \sin \psi',$$

so that the above equation is equivalent to $\tan \psi = \tan \psi'$, or $\psi = \psi'$.

623. The latter case we have already fully considered. In the former, we have $\cos(p - p') = 0$, and, therefore,

$$A = a + a'; \quad B = b + b'; \quad P = p; \quad Q = p,$$

Light.

Part III.

Case of complete accordance of non-coincident vibrations.

624.

Amplitude and position of resultant vibration determined.

625.

626.

627.

Case of circular vibrations.

and, finally,

$$\frac{Y}{X} = \frac{b + b'}{a + a'} = \tan \phi ; \quad (10)$$

which is the tangent of the angle made by the resulting rectilinear vibration with the axis of the x .

If we put M for the amplitude of the resulting vibration, we have $M \cdot \cos \phi = A$; $M \cdot \sin \phi = B$; therefore, $M^2 \cdot (\cos^2 \phi + \sin^2 \phi) = M^2 = A^2 + B^2$.

Now,

$$A^2 = (a + a')^2 = (m \cdot \cos \psi + m' \cdot \cos \psi')^2$$

$$B^2 = (b + b')^2 = (m \cdot \sin \psi + m' \cdot \sin \psi')^2$$

and, therefore, adding these values together, and reducing

$$M^2 = m^2 + 2 m m' \cdot \cos (\psi - \psi') + m'^2 ; \quad (11)$$

Now, $\psi - \psi'$ is the angle between the directions of the component vibrations, so that this equation expresses that the amplitude of the resultant vibration is in this case also the diagonal of a parallelogram, whose sides

are the amplitudes of the component ones; and it is easily shown, by substituting in $\tan \phi = \frac{b + b'}{a + a'}$ the above values of $a + a'$, $b + b'$, that the diagonal has also the position of the resultant line of vibration.

Corol. 1. Any rectilinear vibration may be resolved into two other rectilinear vibrations, whose amplitudes are the sides of any parallelogram, of which the amplitude of the original vibration is the diagonal, and which are in complete accordance, or have a common origin with it.

Corol. 2. Hence any rectilinear vibration may be readily reduced to the directions of two rectangular coordinates, or, if necessary, into those of three, by the rules of the resolution of forces, and the component vibrations, however numerous, will be in complete accordance with the resultant.

The ellipse degenerates into a circle when $\cos (P - Q) = 0$, or $P - Q = 90^\circ$, and, also, $A = B$. Now the former condition gives $\tan P + \cot Q = 0$, that is

$$\frac{a \cdot \sin p + a' \cdot \sin p'}{a \cdot \cos p + a' \cdot \cos p'} + \frac{b \cdot \cos p + b' \cdot \cos p'}{b \cdot \sin p + b' \cdot \sin p'} = 0$$

or reducing

$$\cos (p - p') = - \frac{a b + a' b'}{a b' + a' b} = - \frac{1}{2} \frac{m^2 \cdot \sin 2 \psi + m'^2 \sin 2 \psi'}{m m' \cdot \sin (\psi + \psi')} \quad (12)$$

The condition $A = B$, or $A^2 = B^2$, gives

$$a^2 + 2 a a' \cdot \cos (p - p') + a'^2 = b^2 + 2 b b' \cdot \cos (p - p') + b'^2$$

whence we, in like manner, obtain

$$\cos (p - p') = - \frac{(a^2 + a'^2) - (b^2 + b'^2)}{2 a a' - 2 b b'} = - \frac{1}{2} \frac{m^2 \cdot \cos 2 \psi + m'^2 \cdot \cos 2 \psi'}{\cos (\psi - \psi')} \quad (13)$$

and, equating the values of $\cos (p - p')$, we find the following relation between a, a', b, b' , which must subsist when the vibrations are circular,

$$\left(\frac{a}{b} - \frac{a'}{b'} \right) (a^2 + b^2 - a'^2 - b'^2) = 0.$$

The vanishing of the first factor gives no circular vibration, it being introduced with the negative root of the equation $A^2 = B^2$, with which we have no concern. The other gives

$$a^2 + b^2 = a'^2 + b'^2, \text{ or } m = m',$$

which shows that the component vibrations must have equal amplitudes. Now, if for a and b we write their values $m \cdot \cos \psi$ and $m \cdot \sin \psi$, and for a' and b' , respectively, $m \cdot \cos \psi'$ and $m \cdot \sin \psi'$, in either of the expressions for $\cos (p - p')$, it will reduce itself to

$$\cos (p - p') = - \cos (\psi - \psi') ; \text{ or, } p - p' = 180^\circ - (\psi - \psi').$$

Hence it appears, that the *interference of two equal rectilinear vibrations* will produce a resultant circular one, provided the difference of their phases be equal to the *supplement* of the angle their directions make with each other, so that when the molecule is just commencing its motion towards its centre, in virtue of one vibration, it shall be receding from it at an obtuse angle with this motion, in virtue of the other.

Corol. Hence, if two vibrations have equal amplitudes, but differ in their phases by a quarter of an undulation, their resultant vibration will be circular.

We are now in a condition to explain what becomes of the portions of the secondary waves which diverge obliquely from the molecules of the primary ones, as alluded to in Art. 595, and to explain the mode in which those which do not conspire with the primary wave mutually destroy each other. To this end, conceive the surface of any wave ABC to consist of vibratory molecules, *all in the same phase of their vibrations*. Then will the motion of any point X (fig. 130) be the same, whether it be regarded as arising from the original motion of S , or as the resultant of all the motions propagated to it from all the points of this surface. Conceive the surface ABC divided into an infinite number of elementary portions, such that the difference of distance of each consecutive pair from X shall be constant, or $= d f$, putting the distance of any one from that point $= f$; and let $AB, BC, CD, \&c.$, and $A b, b c, c d, \&c.$ be finite portions of the surface containing each the same number of

629.

Fig. 130

Mutual destruction of secondary waves.

Light.

Part III.

these elements, and in each of which the corresponding values of f are exactly half an undulation ($\frac{1}{2} \lambda$) greater than in the preceding, so that (for instance) $BX = AX + \frac{1}{2} \lambda$, $CX = BX + \frac{1}{2} \lambda$, &c. Then it is evident, that the vibrations which reach X simultaneously from the corresponding portions of any two consecutive ones, as of AB and BC, will be in exactly opposite phases; and, therefore, were they of equal intensity, and in precisely the same direction, would interfere with, and destroy each other. Now, first, with regard to their intensity, this depends on the magnitudes of the elements of the wave AB, from which they are derived, and on the law of lateral propagation. Of the latter, we know little, *à priori*; but all the phenomena of light indicate a very rapid diminution of intensity, as the direction in which the secondary undulations are propagated deviates from that of the primary. With respect to the former, it is evident that the elements in the immediate vicinity of the perpendicular AX, corresponding to a given increment df of the distance from X, are much larger than those remote from it; so that all the elements of the portion AB are much larger than those in BC, and these again than in those of CD, and so on. Thus the motion transmitted to X from any element in AB will be much greater than that from the corresponding one in BC, and that again greater than that from the element in CD, and so on. Thus the motion arriving at G, from the whole series of corresponding elements, will be represented by a series such as $A - B + C - D + E - F + \&c.$, in which each term is successively greater than that which follows. Now it is evident that the terms approach with great rapidity to equality; for if we consider any two corresponding elements as M, N at a distance from A at all considerable, the angles XM and XN make with the surface approach exceedingly near to equality, so that the obliquity of the secondary wave to the primary, and of course its intensity, compared with that of the direct wave, is very nearly alike in both; and the elements M, N themselves, at a distance from the perpendicular, approach rapidly to equality, for the elementary triangles Mmo, Mnp are in this case very nearly similar, and have their sides mo, np equal by hypothesis. Finally, the lines MX, NX approach nearer to each other in direction so as to produce a more complete interference, as their distance from A is greater.

629. Thus we see that the terms of the series $A - B + C - D + \&c.$, at a distance from its commencement, have on all accounts (viz. their smallness, near approach to equality, and disposition to interfere) an extremely small influence on its value; and as the same is true of every set of corresponding elements into which the portions AB, BC, &c. are divided, it is so of their joint effect, so that the motion of the molecule X is governed entirely by that of the portion of the wave ABC immediately contiguous to A, the secondary vibrations propagated from parts at a distance mutually interfering and destroying each others effect.

630. It is obvious, that in the case of refraction or reflexion, we may substitute for the wave AM the refracting or reflecting surface; and for the perpendicular XA the primary refracted ray, when the same things, *mutatis mutandis*, will hold good. See M. Fresnel's Paper entitled *Explication de la Réfraction dans le Système des Ondes*, published in the *Bulletin de la Société Philomatique*, October, 1821.

631. This is the case when the portion of the wave ABCD whose vibrations are propagated to X is unlimited, or at least so considerable, that the last term in the series $A - B + C - \&c.$ is very minute compared with the first. But if this be not the case, as, if the whole of a wave except a small part about A be intercepted by an obstacle, the case will be very different. It is easy on this supposition to express by an integral the intensity of the undulatory motion of X, compared with what it would be on the supposition of no obstacle existing. For this purpose, let d^2s be the magnitude of any vibrating element of the surface, f its distance from $X = MX$, and let $\phi(\theta)$ be the function of the angle made by a laterally-divergent vibration with the direct one, which expresses its relative intensity, and which is unity when $\theta = 0$, and diminishes with great rapidity as θ increases. Then if t be the time since a given epoch, $\lambda =$ the length of an undulation, $SA = a$, the phase of a vibration arriving

at X by the route SMX will be $2\pi \left(\frac{t}{T} - \frac{a+f}{\lambda} \right)$, and the velocity produced in X thereby will be represented by

$a \cdot d^2s \cdot \phi(\theta) \cdot \sin 2\pi \left(\frac{t}{T} - \frac{a+f}{\lambda} \right)$, so that the whole motion produced will be represented by

$$\iint a \cdot d^2s \cdot \phi(\theta) \cdot \sin 2\pi \left\{ \frac{t}{T} - \frac{a+f}{\lambda} \right\}$$

the integral being extended to the limits of the aperture.

632. *Corol. 1.* If but a very small portion of the wave be permitted to pass, as in the case of a ray transmitted through a very small hole, and received on a distant screen, θ and $\phi(\theta)$ are very nearly constant, so that the motion excited in X is in this case represented by

$$a \cdot \phi(\theta) \iint d^2s \cdot \sin 2\pi \left\{ \frac{t}{T} - \frac{a+f}{\lambda} \right\}.$$

We shall have occasion to revert to these expressions hereafter.

§ IV. Of the Colours of Thin Plates.

633. Every one is familiar with the brilliant colours which appear on soap-bubbles; with the iridescent hues produced by heat on polished steel or copper; with those fringes of beautiful and splendid colours which appear in the cracks of broken glass, or between the laminæ of fissile minerals, as Iceland spar, mica, sulphate of lime, &c. In all these, and an infinite variety of cases of the same kind, if the fringes of colour be examined

Case of a wave transmitted through a limited aperture.

General account of the phenomena.

Light.

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with care they will be found to consist of a regular succession of hues, disposed in the same order, and determined, obviously, not by any colour in the medium itself in which they are formed, or on whose surfaces they appear, but solely by its greater or less thickness. Thus a soap-bubble (defended from currents of air by being placed under a glass) at first appears uniformly white when exposed to the dispersed light of the sky at an open window; but, as it grows thinner and thinner by the subsidence of its particles, colours begin to appear at its top where thinnest, which grow more and more vivid, and (if kept perfectly still) arrange themselves in beautiful horizontal zones about the highest point as a centre. This point, when reduced to extreme tenuity, becomes black, or loses its power of reflecting light almost entirely. After which the bubble speedily bursts, its cohesion at the vertex being no longer sufficient to counteract the lateral attraction of its parts.

But as it is a matter of great delicacy to make regular observations on a thing so fluctuating and unmanageable as a soap-bubble, the following method of observing and studying the phenomena is far preferable. Let a convex lens, of a very long focus and a good polish, be laid down on a plane glass, or on a concave glass lens having a curvature somewhat less than the convex surface resting on it; so that the two shall touch in but a single point, and so that the interval separating the surfaces in the surrounding parts shall be exceedingly small. If the surfaces be very carefully cleaned from dust before placing them together, and the combination be laid down before an open window in full daylight, the point of contact will be seen as a black spot in the general reflexion of the sky on the surfaces, surrounded with rings of vivid colours. A glass of 10 or 12 feet focus laid on a plane glass, will show them very well. If one of shorter focus be used, the eye may be assisted by a magnifying glass. The following phenomena are now to be attended to:

Phenomenon 1. The colours, whatever glasses be used, provided the incident light be white, always succeed each other in the very same order; that is, beginning with the central black spot, as follows:

First ring, or first order of colours,—black, very faint blue, brilliant white, yellow, orange, red.
Second ring, or second order,—dark purple or rather violet, blue, green, (very imperfect, a yellow-green.) vivid yellow, crimson red.

Third ring, or third order,—purple, blue, rich grass green, fine yellow, pink, crimson.

Fourth ring, or fourth order,—green, (dull and bluish,) pale yellowish pink, red.

Fifth ring, or fifth order,—pale bluish green, white, pink.

Sixth ring, or sixth order,—pale blue-green, pale pink.

Seventh ring, or seventh order,—very pale bluish green, very pale pink. After these, the colours become so faint that they can scarcely be distinguished from white.

On these we may remark, that the green of the third order is the only one which is a pure and full colour, that of the second being hardly perceptible, and of the fourth comparatively dull and verging to apple green; the yellow of the second and third order are both good colours, but that of the second is especially rich and splendid; that of the first being a fiery tint passing into orange. The blue of the first order is so faint as to be scarce sensible, that of the second is rich and full, but that of the third much inferior; the red of the first order hardly deserves the name, it is a dull brick colour; that of the second is rich and full, as is also that of the third; but they all verge to crimson, nor does any pure scarlet, or prismatic red, occur in the whole series.

Phenomenon 2. The breadths of the rings are unequal. They decrease, and the colours become more crowded, as they recede from the centre. Newton (to whom we owe the accurate description and investigation of their phenomena) found by measurement the diameters of the darkest (or purple) rings, just when the central black spot began to appear by pressure, and reckoning it as one of them to be as the square roots of the even numbers 0, 2, 4, 6, &c.; and those of the brightest parts, of the several orders of colours, to be as the square roots of the odd numbers 1, 3, 5, 7, &c. Now the surfaces in contact being spherical, and their radii of curvature very great in proportion to the diameters of the rings, it follows from this that the intervals between the surfaces at the alternate points of greatest obscurity and illumination are as the natural numbers themselves 0, 1, 2, 3, 4, &c. The same measurements, when the radii of curvature of the contact surfaces are known, give the absolute magnitudes of the intervals in question. In fact, if r and r' be the curvatures of two spherical surfaces, a convex and concave, in contact, and D the diameter of any annulus surrounding their point of contact, the interval of the surfaces there will be the difference of the versed sines of the two circular arcs having a common chord D . Now (fig. 130) if AE be the diameter of the convex spherical surface AD , we have $EA : AD :: AD : DB$

$= \frac{AD^2}{AE} = \frac{D^2}{8} r$, and in like manner $BC = \frac{D^2}{8} r'$, so that $\frac{1}{8} D^2 (r - r') = DC$, the interval of the

surfaces at the point D . Thus Newton found, for the interval of the surfaces at the brightest part of the first ring, one 178000th part of an inch; and this distance, multiplied by the even natural numbers 0, 2, 4, 6, 8, &c. gives their distance at the black centre and the darkest parts of the purple rings, and by the odd ones 1, 3, 5, &c. their intervals at the brightest parts.

Phenomenon 3. If the rings be formed between spherical glasses of various curvatures, they will be found to be larger as the curvatures are smaller, and *vice versa*; and if their diameters be measured and compared with the radii of the glasses, it will be found, that, provided the eye be similarly placed, the same colour is invariably produced at that point, or that distance from the centre where the interval between the surfaces is the same. Thus, the white of the first order is invariably produced at a thickness of one 178000th of an inch; the purple, which forms the limit of the first and second orders, at twice that thickness. So that there is a constant relation between the tint seen and the interval of the surfaces where it appears. Moreover, if the glasses be distorted by violent and unequal pressure, (as is easily done if thin lenses be used,) the rings lose their circular figure, and extend themselves towards the part where the irregular pressure is applied, so as to form a species of level lines each marking out a series of points where the surfaces are equidistant. Thus, too, if a

634.
Rings
formed be-
tween
convex
glasses.

635.
Order of
succession
of the
colours.

636.

637.
Law of the
breadths of
the rings
and thick-
nesses at
which they
appear.

638.
Invariable
relation be-
tween the
colours and
thicknesses
of plates.

Light. cylinder be laid on a plane, the rings pass into straight lines arranged parallel to its line of contact, but following the same law of distance from that line as the rings from their dark centre, and if the glasses be of irregular curvature, as bits of window glass, the bands of colour will follow all their inequalities; yet more, if the pressure be very cautiously relieved, so as to lift one glass from the other, the central spot will shrink and disappear, and so on; each ring in succession contracting to a point, and then vanishing, so as to bring all the more distant colours successively to the centre, as the glasses recede from absolute contact. From all these phenomena it is evident, that it is the distance between the surfaces only at any point which determines the colour seen there.

639. *Phenomenon 4.* This supposes, however, that we observe them with the eye similarly placed, or at the same angle of obliquity. For if the obliquity be changed by elevating or depressing the eye, or the glasses, the diameters (but not the colours) of the rings will change. As the eye is depressed, the rings enlarge; and the same tint which before corresponded to an interval of the 178000th of an inch, now corresponds to a greater interval. This distance ($\frac{1}{178000}$) is determined by measures taken nearly at, and reduced by calculation exactly to, a perpendicular incidence. At extreme obliquities, however, the diameters of the several rings suffer only a certain finite dilatation, and Newton's measures led him to the following rule: viz. "*That the interval between the surfaces at which any proposed tint is produced, is proportional to the secant of an angle whose sine is the first of 106 arithmetical mean proportionals between the sines of incidence and refraction, into the glass from the air, or other medium included between the surfaces, beginning with the greater;*" or, in algebraic language, the relative index of refraction being μ , and θ the angle of incidence, and ρ that of refraction of the ray as it passes out of the rarer medium into the denser; then, if t be the interval corresponding to a given tint at the oblique incidence θ , and T at a perpendicular incidence, we shall have

$$t = T \cdot \sec u \text{ where } \sin u = \sin \theta \frac{1}{107} - (\sin \theta - \sin \rho)$$

but $\sin \rho = \frac{1}{\mu} \cdot \sin \theta$, consequently we have

$$t = T \cdot \sec u; \sin u = \frac{106 + \frac{1}{\mu}}{107} \cdot \sin \theta = \frac{106\mu + 1}{107\mu} \cdot \sin \theta.$$

640. To see the rings conveniently at extreme obliquities, a prism may be used, laid on a convex lens, as in fig. 132. If the eye be placed at K, the set of rings formed about the point of contact E will be seen in the direction K H, and as the eye is depressed towards the situation I, where the ray I G intromitted from I would just begin to suffer total reflexion, the rings are seen to dilate to a certain considerable extent. When the eye reaches I, the upper half of the rings disappears, being apparently cut off by the prismatic iris of Art. 555, which is seen in that situation, but the black central spot and the lower half of the rings remains; but when the eye is still further depressed the rings disappear, and leave the central spot, like an aperture seen in the silvery whiteness of the total reflexion on the base of the prism, and dilated very sensibly beyond the size of the same spot seen in the position K H: thus proving, that the want of reflexion on that part of the base extends beyond the limits of absolute contact of the glasses, and that, therefore, the lower surface interferes with the action of the upper, and prevents its reflexion while yet a finite interval (though an excessively minute one) intervenes between them. Euler has made this an objection to the undulatory theory, but the objection rests on no solid grounds, as it is very reasonable to conclude, that the change of density or elasticity in the ether within and without a medium is not absolutely *per saltum*, but gradual. If so, and if the change take place without the media, the approach of two media within that limit, within which the condensation of the ether takes place, will alter the law of refraction from either into the interval separating them.

641. In order, however, to see to the greatest advantage the colours reflected by a plate of air at great obliquities, the following method, first pointed out by Sir William Herschel, may be employed. On a perfectly plane glass or metallic mirror, before an open window, lay an equilateral prism, having its base next the glass or mirror very truly plane, and looking in at the side A C, fig. 133, the reflected prismatic iris, a, b, c , will be seen as usual in the direction E F, where a ray from E would just be totally reflected. Within this iris, and arranged parallel to it, are seen a number of beautiful coloured fringes, whose number and distances from each other vary with every change of the pressure; their breadths dilating as the pressure is increased, and *vice versâ*. They do not require for their formation, that the surfaces should be exceedingly near, being seen very well when the prism is separated from the lower surfaces by the thickness of thin tissue paper, or a fine fibre of cotton wool interposed, but in this case they are exceedingly close and numerous. If the pressure be moderate, they are nearly equidistant, and are lost, as it were, in the blue iris, without growing sensibly broader as they approach it. As the intervals of the surfaces is diminished, they dilate and descend towards the eye, appearing, as it were, to come down out of the iris. They do not require for their formation a perfect polish in the lower surface. An emiered glass, so rough as to reflect no regular image at any moderate incidence, shows them very well. The experiment is a very easy one, and the phenomena so extremely obvious and beautiful, that it is surprising it should not have been noticed and described by Newton, especially as it affords an excellent illustration of his law above stated. To understand this, let E H, E K, E L be any rays from E incident at angles somewhat less than that of total reflexion on the base; they will therefore be refracted, and, emerging at the base B C, will be reflected at M N, (the obliquity of the reflexion being so great, that even rough surfaces reflect copiously and regularly enough for the purpose, Art. 558,) and will pursue the courses H D P p, K F Q q, L G R r, &c. entering the prism again at P, Q, R. Reciprocally, then, rays p P, q Q, &c. incident at P, Q, &c. in these directions,

Light will enter the eye at E after traversing the interval B C N M, and being reflected at M N, and will affect the eye with the colour corresponding to that obliquity and that interval between the surfaces which is proper to each. If then we put, as above, θ for the exterior angle of incidence of the ray D H on the base of the prism, and take

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$$\sin u = \frac{106\mu + 1}{107\mu} \cdot \sin \theta = \frac{106\mu + 1}{107} \cdot \sin \rho = k \cdot \sin \rho,$$

the tint seen in the direction E H will (abstraction made of the dispersion at the surface A C) be the same with that reflected at a perpendicular incidence, by a plate of air of the thickness $T = t \cdot \cos u = t \sqrt{1 - k^2 \cdot \sin^2 \rho}$, where t = the distance between the surfaces B C, M N. There will, therefore, appear a succession of colours in the several consecutive situations of the line E H, analogous to those of the coloured rings, (except in so far as the dispersion of the side A C alters the tints by separating their component rays.)

But the whole series of colours will not be seen, because those which require greater obliquities than that at which total reflexion takes place, cannot be formed. In fact, the angle, reckoned from the vertical at which a tint corresponding to a thickness T in the rings would be formed, is given by the equation

642.

$$\sin \rho = \frac{1}{k} \cdot \sqrt{1 - \left(\frac{T}{t}\right)^2} = \frac{214}{320} \sqrt{1 - \left(\frac{T}{t}\right)^2},$$

taking $\mu = \frac{3}{2}$ for glass, which it is very nearly. Now, according to this, the central tint, or black of the first order, which is formed when $T = 0$, requires that

$$\sin \rho = \frac{1}{k} = \frac{1}{\mu - \frac{1}{107}}$$

which being greater than $\frac{1}{\mu}$ shows that this tint lies above the situation of the iris, and cannot therefore be seen. The first visible tint will be that close to the iris, where $\sin \rho = \frac{1}{\mu}$ which gives

$$T = t \sqrt{1 - \left(\frac{k}{\mu}\right)^2} = t \sqrt{1 - \left(1 - \frac{\mu - 1}{107\mu}\right)^2} = t \sqrt{\frac{2(\mu - 1)}{107\mu}} = 0.079 t$$

nearly, or $\frac{t}{12.25}$. Hence it appears, that these fringes would be seen, by an eye immersed in the prism, when

the interval between its base and the glass it rests on is more than 12 times that at which colours are formed at a perpendicular incidence, *i. e.* at $12.25 \times \frac{13}{178000}$, or about $\frac{1}{1100}$ th of an inch, which is about the thickness

of fine tissue paper. Moreover, from this value of T, we see that the first tint immediately visible below the iris ascends in the scale of the rings (*i. e.* belongs to a point nearer their centre) as the value of t diminishes, or as the prism is pressed closer to the glass; and this explains why the fringes become more numerous, and appear to come out of the iris by pressure. With regard to their angular breadth, (still to an eye immersed in the

prism.) If we put $e = \frac{1 \text{ inch}}{89000}$, we have, putting ρ_0, ρ_1 , &c. for the values of ρ , corresponding to the several orders of visible tints,

$$\sin \rho_0 = \frac{1}{\mu}; \sin \rho_1 = \frac{1}{k} \sqrt{1 - \left(\frac{T+e}{t}\right)^2} = \frac{1}{\mu} \sqrt{1 - \frac{\mu^2}{k^2} \cdot 0.079 \cdot \frac{2e}{t}} = \frac{1}{\mu} \left(1 - 0.079 \times \frac{e}{t}\right)$$

very nearly, $\sin \rho_2 = \frac{1}{\mu} \left(1 - 0.079 \cdot \frac{2e}{t}\right)$ and so on. The sines then of the incidences at which the several

orders of colours are developed, beginning at the iris, increase in arithmetical progression, so that the fringes must be disposed in circular arcs parallel to the iris, and their breadths must be nearly equal, and greater the greater the pressure or the less t is, all which is conformable to observation. The refraction of the side of the prism between the eye and the base, however, disturbs altogether the succession of colours in the fringes, and in particular multiplies the number of visible alternations to a great extent, in a manner which will be evident on consideration. We have been rather more particular in explaining the origin of these fringes, and referring them to the general phenomena observed by Newton, because up to the present time we believe no strict analysis of them has been given, as well as on account of the great beauty of the phenomenon itself. If we hold the combination up to the light, and look through the base of the prism and the glass plate, so as to see the transmitted iris of Art. 556, its concavity will, in like manner, be seen fringed with bands of colours of precisely similar origin. To return now to the rings seen between convex glasses.

643.

Phenomenon 5. If homogeneous light be used to illuminate the glasses, the rings are seen in much greater

Light. number, and the more according to the degree of homogeneity of the light. When this is as perfect as possible, as, for instance, when we use the flame of a spirit lamp with a salted wick, as proposed by Mr. Talbot, they are literally innumerable, extending to so great a distance that they become too close to each other to be counted, or even distinguished by the naked eye, yet still distinct on using a magnifier, but requiring a higher and higher power as they become closer, till we can pursue them no farther, and disappearing from their closeness, and not from any confusion or running of one into the other. Moreover, they are now no longer composed of various colours, but are wholly of the colour of the light used as an illumination, being mere alternations of light and obscurity, and the intervals between them being absolutely black.

644. *Phenomenon 6.* When the illuminating light is changed from one homogeneous ray to another, as when, for instance, the colours of the prismatic spectrum are thrown in succession on the glasses at their point of contact, at such an angle as to be reflected to the eye, then, the eye remaining at rest, the rings are seen to dilate and contract in magnitude as the illumination shifts. In red light they are largest, in violet least, and in the intermediate colours of intermediate size. Newton, by measuring their diameters, ascertained that the interval of the surfaces or thickness of the plate of air, where the violet ring of any order was seen, is to its thickness, where the corresponding red ring of the same order is formed, nearly as 9 : 14 ; and, determining by this method, the thickness of the plate of air where the brightest part of the first ring was formed, when illuminated in succession by the several rays proceeding from the extreme red to the extreme violet, he ascertained those thicknesses to be the halves of the numbers already set down in the second column of the Table, p. 453, expressed in

parts of an inch, and which answer to the values of $\frac{\lambda}{2}$, or the lengths of a semiundulation for each ray.

645. This phenomenon may be regarded as an analysis of what takes place when the rings are seen in white light ; for in that case they may be regarded as formed by the superposition one on the other of sets of rings of all the simple colours, each set having its own peculiar series of diameters. The manner in which this superposition takes place, or the *synthesis* of the several orders of colours, may be understood by reference to fig. 134, where the abscissæ or horizontal lines represent the thicknesses of a plate of air between two glasses, supposed to increase uniformly, and where $R R', R R'', \&c.$ represent the several thicknesses at which the red, in the system of rings illuminated by red rings only, vanishes, or at which the darkness between two consecutive red rings is observed to happen, while $R r, R r', R r'', \&c.$ represent those at which the brightness is a maximum. In like manner, let $0 0', 0 0'', \&c.$ be taken equal to the several thicknesses at which the orange vanishes, or at which the black intervals in the system of orange rings are seen, and so on for the yellow, green, blue, indigo, and violet rings. So that $R R', 0 0', Y Y', \&c.$ are to each other in the ratio of the numbers in column 2 of the above Table, (Art. 575.) Then if we describe a set of undulating curves as in the figure, and at any point, as C in A E, draw a line parallel to A V, cutting all these curves ; their several ordinates, or the portions of this line intercepted between the curves and their abscissæ, will represent the intensity of the light of each colour, sent to the eye by that thickness of the plate of air. Hence, the colour seen at that thickness will be that resulting from the union of the several simple rays in the proportions represented by their ordinates.

646. The figure being laid down by a scale, we may refer to it to identify the colours of particular points. Thus, first at the thickness 0, or at A the origin of the tints, all the ordinates vanish, and this point, therefore, is black. As the thickness of the plate of air increases from 0 while yet very small, it is evident, on inspection, that the ordinates of the several curves increase with unequal rapidities, those for the more refrangible rays more rapidly than those for the less, so that the first feeble light which appears at a very small thickness A 1, will have an excess of blue rays, constituting the pure but faint blue of the first order, (Art. 635.) At a greater thickness, however, as A 2, the common ordinate passes nearly through the maxima of all the curves, being a little short of that of the red, and a little beyond that of the violet. The difference, however, is so small, that the several colours will all be present nearly in the proportions to constitute whiteness, and being all nearly at their maximum, the resulting tint will be a brilliant white. This agrees with observation ; the white of the first order being, in fact, the most luminous of all ; beyond this the violet falls off rapidly, the red increases, and the yellow is nearly at its maximum, so that at the thickness A 3 the white passes into yellow, and at a still greater thickness, A 4, where the violet, indigo, blue, and green, are all nearly evanescent, the yellow falling off, and the orange and red, especially the latter, in considerable abundance, the tint resulting will be a fiery orange, growing more and more ruddy. At B is the minimum of the yellow, *i. e.* of the most luminous rays. Here then will be the most sombre tint. It will consist of very little either of orange, green, blue, or even indigo ; but a moderate portion of violet and a little red will produce a sombre violet purple, which, since the more refrangible rays are here all on the increase, while the less are diminishing, will pass rapidly to a vivid blue, as at the thickness denoted by A 5. At 6, where the ordinate passes through the maximum of the yellow, there is almost no red, very little orange, a good deal of green, very little blue, and hardly any indigo or violet. Here then the tint will be yellow verging to green, but the green is diminishing and the orange increasing, so that the yellow rapidly loses its green tinge, and becomes pure and lively. At 7 the predominant rays are orange and yellow, being so copious that the little red and violet with which they are mixed does not prevent the tint from being a rich, high-coloured yellow. At 8 a full orange and copious red are mixed with a good deal of indigo and a maximum of violet, thus producing a superb crimson. At C we have again a minimum of yellow ; but there being at the same time a maximum of red and indigo, this point, though dark in comparison of that on either side, will still be characterised by a fine ruddy purple. This completes, and as we see faithfully represents, the second order of colours. At 9, 10 we see the origin of the vivid green of the third order, in the comparative copiousness of green, yellow, and blue rays at the former point, and of yellow, green,

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and violet at the latter, while the red and orange are almost entirely absent, and thus we may pursue all the tints in the scale enumerated in Art. 635 with perfect fidelity. Part III.

As the thickness increases, however, it is clear that rays differing but little in refrangibility will differ much in intensity, as the smallest difference in the lengths of the bases of their curves being multiplied by the number of times they are repeated, will at length bring about a complete opposition, so that the maximum of one ray will fall at length on a minimum of another differing little in refrangibility, and not at all in colour. Thus, at considerable thicknesses, such as the 10th or 20th order, there will coexist both maxima and minima of every colour; since each colour, in fact, consists not of rays of one definite refrangibility, but of all gradations of refrangibility between certain limits. In consequence, the tints, as the thickness increases, will grow less and less pure, and will at length merge into undistinguishable whiteness, which, however, for this very reason, will be only half as brilliant as the white of the first order, which contains all the rays at their *maximum* of intensity.

Phenomenon 7. Such are the phenomena when a plate of air is included between two surfaces of glass. It is not however as *air*, but as *distance*, that it acts; for in the vacuum of an air-pump the rings are seen without any sensible alteration. If, however, a much more refracting medium, as water or oil, be interposed, the diameters of the rings are observed to contract, preserving, however, the same colours and the same laws of their breadths; and Newton found by exact measurements, that the thicknesses of different media interposed, at which a given tint is seen, are in the inverse ratio of their refractive indices. Thus, the white of the first order being

produced in vacuo or air at the 178000th of an inch, will be produced in water at $\frac{1}{1.336}$ part of that thickness.

He found, moreover, that the law stated in Art. 639 for the dilatation of the rings by oblique incidence, holds equally good, whatever be the nature of the interposed medium. Hence it follows, that in dense media the dilatation at great obliquities is much less than in rare ones, and that in consequence a given thickness will reflect a colour much less variable by change of obliquity when the medium has a high refractive power than when low. Thus, the colours of a soap bubble vary much less by change of incidence than those of a film of air, and much more, on the other hand, than the iridescent colours on polished steel, which arise from a film of oxide formed on the heated surface.

Phenomenon 8. Surfaces of glass, or other denser medium enclosing the thin plate of a rarer, are not however necessary to the production of the colours; they are equally, and indeed more brilliantly, visible when any very thin laminae of a denser medium is enclosed in a rarer, as in air, or in vacuo. Thus, soap bubbles, exceedingly thin films of mica, &c. exhibit the same succession of colours, arranged in fringes according to the variable thickness of the plates. The following very beautiful and satisfactory mode of exhibiting the fringes formed by plates of glass of a tangible thickness has been imagined by Mr. Talbot. If a bubble of glass be blown so thin as to burst, and the glass films which result be viewed in a dark room by the light of a spirit lamp with a salted wick, they will be seen to be completely covered with striæ, alternately bright and black, in undulating curves parallel to each other according to the varying thickness of the film. Where the thickness is tolerably uniform, the striæ are broad; where it varies rapidly, they become so crowded as to elude the unassisted sight, and require a microscope to be discerned. If the film of glass producing these fringes be supposed equal to the thousandth of an inch in thickness, they must correspond to about the 89th order of the rings, and thus serve to demonstrate the high degree of homogeneity of the light; for if the slightest difference of refrangibility existed, its effect multiplied eighty-nine times would become perceptible in a confusion and partial obliteration of the black intervals. In fact, the thickness of a plate at which alternations of light and darkness or of colour can no longer be discerned, is the best criterion of the degree of homogeneity of any proposed light, and is, in fact, a numerical measure of it. This experiment is otherwise instructive, as it shows that the property of light on which the fringes depend is not restricted to extremely minute thicknesses, but subsists while the light traverses what may be comparatively termed considerable intervals.

Phenomenon 9. When the glasses between which the reflected rings are formed are held up against the light, a set of transmitted coloured rings is seen, much fainter, however, than the reflected ones, but consisting of tints complementary to those of the latter, *i. e.* such as united with them would produce white. Thus the centre is white, which is succeeded by a yellowish tinge, passing into obscurity, or black, which is followed by violet and blue. This completes the series of the first order. Those of the second are white, yellow, red, violet, blue: of the third, green, yellow, red, bluish green, after which succeed faint alternations of red and bluish green, the degradation of tints being much more rapid than in the reflected rings.

It was to explain these phenomena that Newton devised his doctrine of the fits of easy reflexion and transmission, mentioned in the 9th postulate of Art. 526. This doctrine we shall now proceed to develop further, and apply, as he has done, to the case in question. In addition then to the general hypothesis there assumed, it will be necessary to assume as follows:

The intervals at which the fits recur, differ in different rays according to their refrangibilities, being greatest for the red and least for violet rays, and for these, and the intermediate rays, *in vacuo*, and at a perpendicular incidence, are represented in fractions of an inch by the halves of the numbers in column 2 of the Table, Art. 575. In other media, the lengths of the intervals in the course of a molecule at which its fits recur are shorter, in the ratio of the index of refraction of the medium to unity.

At oblique incidences, or when a ray traverses a medium after being intromitted obliquely, (at an angle = 0 with the internal perpendicular,) the lengths of the fits are greater than at a perpendicular incidence, in the ratio of radius to the rectangle between the cosine of θ and the cosine of an arc u , given by the equation

$$\sin u = \frac{106 \mu + 1}{107 \mu} \sin \theta.$$

647.

Degradation of the tints.

648

Colours reflected by plates of different media.

649.

Colours reflected by soap bubbles, &c.

650.

Transmitted colours.

651.

Newton's explanation of the colours of thin plates.

652.

Laws of the fits.

653.

654.

Light.

Part III

655.

Explanation
of the rings
seen by ho-
mogeneous
light.

Let us now consider what will happen to a luminous molecule, the length of whose fits in any medium is $\frac{1}{2} \lambda$, which, having been intromitted perpendicularly at the first surface, and traversing its thickness ($= t$), reaches the second. First, then, if we suppose t an exact multiple of $\frac{1}{2} \lambda$, it is evident that the molecule will arrive at the second surface in precisely the same phase of its fit of transmission as at the first. Of course it is placed in the very same circumstances in every respect, and having been transmitted before must necessarily be so again. Thus every ray which enters perpendicularly into such a lamina must pass through it, and cannot be reflected at its second surface. On the other hand, if the thickness of the lamina be supposed an exact odd multiple of $\frac{1}{4} \lambda$, &c. every molecule intromitted at its first surface will on its arrival at the second be in exactly the contrary phase of its fits, and, having been before in some phase of a fit of transmission, will now be in a similar phase of a fit of reflexion. It will, therefore, not necessarily be transmitted; but a reflexion, more or less copious, will take place at the second surface in this case, according to the nature of the medium and its general action on light. For it will be remembered, that every molecule in a fit of reflexion is not necessarily reflected. It is disposed to be so; but whether it will or no, will depend on the medium it moves in and that on which it impinges, and on the phase of its fit. Now conceive an eye placed at a distance from a lamina of unequal thickness, so as to receive rays reflected at a very nearly perpendicular incidence from it. It is evident, that in virtue of the reflexion from the first surface, which is uniform, it will receive equal quantities of light from every point. But with regard to the light reflected from the second the case is different; for in all those parts where the thickness of the lamina is an exact even multiple of $\frac{1}{4} \lambda$, none will be reflected, while in all those where it is an exact odd multiple of $\frac{1}{4} \lambda$, a reflexion will take place; and since each molecule so reflected retraces the path

by which it arrived, and therefore describes again the same multiple of $\frac{\lambda}{4}$; its total path described within the lamina, when it has reached the first surface again, will be an exact multiple of $\frac{\lambda}{2}$, and therefore it will penetrate that surface and reach the eye. In consequence, in virtue of the reflexion at the second surface alone, the lamina would appear black in every part where its thickness $= 0$, or $\frac{2\lambda}{4}$, or $\frac{4\lambda}{4}$, &c., and bright in those parts where its thickness $= \frac{1\lambda}{4}$, or $\frac{3\lambda}{4}$, $\frac{5\lambda}{4}$, &c. *ad infinitum*. In the intermediate thicknesses it would have a brightness intermediate between these and absolute obscurity; so that on the whole, the lamina would appear marked all over with dark and bright alternating fringes, just as we see it actually does in the experiment described, (Art. 649.) The uniform reflexion from the first surface superposed on these, will not prevent their inequality of illumination from being distinctly seen.

656.

Of the
rings seen
by white
light.

Hence it is evident, that if we take the abscissa of a curve equal to thickness of the lamina at any point, and the ordinate proportional to the intensity of the light reflected from the second surface, and returned through the first, this curve will be an undulating line, such as we have represented in fig. 134, touching the abscissa at equal distances equal to the length of a whole fit of a ray of the colour in question. Now these distances for rays of different colours being supposed such as we have assumed in Art. 652, the construction of Art. 645 holds good, and when white light falls on the lamina, its second surface will reflect a series of colours of the composition there demonstrated, and such as we actually observe, but diluted with the light uniformly reflected from every point of the first surface.

If the lamina instead of a vacuum be composed of any refracting medium, the tints will manifestly succeed each other in a similar series, but the thickness at which they are produced will be to that in a lamina of vacuum, in the ratio of the lengths of the fits in the two cases, that is, in the proportion of 1 : the index of refraction of the medium. Thus the rings seen between two object glasses including air, ought to contract when water, oil, &c. is admitted between them, as they are found to do, and, by measure, in that precise ratio.

657.

Of the dilata-
tion of the
rings at
oblique
incidences.

At oblique incidences, θ being the angle of intromission into the lamina, $t \cdot \sec \theta$ is the whole path of the ray between the first and second surfaces, and since $\frac{1}{2} \lambda \cdot \sec \theta$ is the length of the fits of the given ray at this obliquity, in order that the luminous molecule may arrive at the second surface in the same phase, and therefore be reflected with equal intensity, it must in this space have passed over the same number of these fits;

hence we must have $\frac{2 t \cdot \sec \theta}{\lambda \cdot \sec \theta \cdot \sec u} = \text{constant}$, or t proportional to $\sec u$, which agrees with observation.

658.

Of the
transmitted
rings.

All the light which is not reflected at the second surface passes through it, and forms the transmitted series of colours. These, therefore, consist of the whole incident light ($= 1$) minus that reflected at the first surface, (which will be a small fraction, and which we will call a), minus that reflected at the second surface. Now this last will be a periodical function whose minimum is 0, and its maximum can never exceed a , because the reflexion at the second surface of a medium cannot be stronger than at the first at a perpendicular incidence.

We may then represent it by $a \left(\sin \frac{2t}{\lambda} \right)^2$, and thus we have $1 - a \left\{ 1 + \left(\sin \frac{2t}{\lambda} \right)^2 \right\}$ for the intensity of this particular coloured ray in the transmitted series, and $a \left(\sin \frac{2t}{\lambda} \right)^2$ in the reflected. Hence it is evident, that

owing to the smallness of a , the difference between the brightest and darkest part of the transmitted series will be small in comparison with the whole light, and thus the alternations in homogeneous light ought to be (as they are) much less sensible than in the reflected rings, and the tints in white light much more pallid and dilute.

Light.

Thus we see that the Newtonian hypothesis of the fits affords a satisfactory-enough explanation, or rather represents with exactness all the phenomena above described. It has been even asserted, that this doctrine is really not an hypothesis, but nothing more than a pure statement of facts; for that, first, in point of mere fact, the second surface of the lamina does send light to the eye, in the bright parts of the fringes, and does not send it in the dark parts; and, secondly, that this is the same thing with saying that the light which has traversed a

Part III.

659.

thickness $= (2n + 1) \frac{\lambda}{4}$ is, and that which has traversed $2n \frac{\lambda}{4}$ is not susceptible of being reflected. And,

in truth, if only one ray could be regarded as being concerned, and were the light reflected at the first surface of the lamina altogether out of the question, this way of stating it would be strictly correct. But, if it can be shown, that, on any other hypothesis of the nature of light, (as the undulatory,) the second link of this argument is invalid; and that though the second surface, like the first, may reflect in every part, without regard to its thickness, its full average portion of the light that is incident on it; yet that afterwards, by reason of the interference of rays reflected from the first surface, such light does not reach the eye (being destroyed in every point of its course) from those parts where the thickness is an even multiple of $\frac{\lambda}{4}$, then it is evident, that the Newtonian doctrine is something more than a mere *aliter* statement of facts, and is open to examination as a theory.

Let us now see, therefore, what account the undulatory theory gives of these phenomena. We will begin, for a reason which will presently appear, with the *transmitted* rings. Conceive, then, a ray, the length of whose undulations in any medium is λ , to be incident perpendicularly on the first surface of a lamina of that medium whose thickness is t ; and (for simplicity) let its surfaces be supposed parallel, then it will be divided into two portions, the first ($= a$) reflected, and the second ($= 1 - a$) intromitted. Let θ be the phase of this portion at reaching the second surface. Here it will be again divided into two portions, the one reflected back into the medium and equal to $(1 - a) \cdot a$, or (a being small) very nearly to a , and the remainder $(1 - a) - a(1 - a)$, or nearly $1 - 2a$, transmitted. These portions, supposing no undulation, or part of an undulation, gained or lost in the act of transmission or reflexion, will both be in the phase θ . The reflected

660.

Explanation of the transmitted rings on the undulatory hypothesis.

portion will again encounter the first surface in the phase $\theta + 2\pi \cdot \frac{t}{\lambda}$, will there be again partially reflected, with an intensity equal to $a \times a = a^2$, and the portion so reflected will reach the second surface in the phase $\theta + 2\pi \cdot \frac{2t}{\lambda}$, and will there be transmitted with an intensity $= (1 - a) \cdot a^2$, or nearly $= a^2$. Now, the reflexions being all perpendicular, this portion will be confounded with the portion $1 - 2a$ transmitted without any reflexion; and putting $a = \sqrt{1 - 2a} = 1 - a$ nearly, and $a' = \sqrt{a^2} = a$, a and a' will represent the amplitudes of vibration of the ethereal molecule at the posterior surface, which each of these rays tend to impress on it. Hence, its total excursion from rest will be represented by

$$a \cdot \cos \theta + a' \cdot \cos \left(\theta + 2\pi \cdot \frac{2t}{\lambda} \right),$$

that is

$$\begin{aligned} & (1 - a) \cos \theta + a \cdot \cos \left(\theta + 2\pi \cdot \frac{2t}{\lambda} \right). \\ & = 1 \cos \theta + a \cdot \cos \left(\theta + 2\pi \cdot \frac{2t}{\lambda} \right) - a \cdot \cos \theta. \end{aligned}$$

The first term of this is independent of t , and represents, in fact, the incident ray in the state in which it would arrive at the second surface, had no reflexions taken place. The other two terms represent rays the former of which evidently is in complete discordance with the latter, and destroys it when t is any odd multiple of $\frac{\lambda}{4}$, (or of the half length of one of Newton's fits, a fit being, as we have seen above, equal to half an undulation,) thus leaving the ray at its emergence of the same intensity as it would have had were the lamina away; but when t is any odd multiple of half a fit, then the value of $\cos \left(\theta + 2\pi \cdot \frac{2t}{\lambda} \right) = -\cos \theta$; and the emergent ray is in this case represented by $(1 - 2a) \cdot \cos \theta$, being less than the incident ray by twice the light reflected at the first surface.

Thus if the thickness of the plate be different in different parts, the light transmitted through it to the eye will not be uniform, but will have alternate maxima and minima corresponding to the thicknesses 0

$$\frac{\lambda}{4}, \quad \frac{2\lambda}{4}, \quad \frac{3\lambda}{4}, \quad \&c.$$

661.

Origin of the bright and dark rings in homogeneous light.

If we apply to the expression above given, the general formula Art. (613) for the composition of rays in one plane, we shall find for the intensity A^2 of the ray finally emergent,

Light.
662.
General
expression
for the
transmitted
ray.

$$\begin{aligned} A^2 &= (1-a)^2 + 2a(1-a) \cdot \cos 2\pi \cdot \frac{2t}{\lambda} + a^2 \\ &= 1 - 4a(1-a) \cdot \sin^2 \left(2\pi \cdot \frac{t}{\lambda} \right) \\ &= 1 - 4a \cdot \sin^2 \left(2\pi \cdot \frac{t}{\lambda} \right) \end{aligned}$$

which shows that the several maxima are equal to the incident ray, and the minima to that ray diminished by four times the light reflected at the first surface. The difference of phase between the simple and composite emergent ray, or the value of B in the formula cited, is given by the equation,

$$\sin B = \frac{a}{A} \cdot \sin \left(2\pi \cdot \frac{2t}{\lambda} \right) = a \cdot \sin \left(2\pi \cdot \frac{2t}{\lambda} \right), \text{ neglecting } A^2,$$

so that for such media as have not a very high refractive power, this difference is always small. It is, however, periodical, and differs for different thicknesses.

663.
Transmitted
tints in
white light
expressed
algebraically.

Suppose now, instead of homogeneous light, white light to fall on the lamina, and let us represent a ray of such light, as in Art. 488, by $C + C' + C'' + \&c.$, or by $S(C)$, C , C' , &c. being the intensity of the several elementary rays of all degrees of refrangibility, then will the transmitted compound beam be represented in tint and intensity by

$$C \left\{ 1 - 4a \cdot \sin^2 \left(2\pi \cdot \frac{t}{\lambda} \right) \right\} + C' \left\{ 1 - 4a \cdot \sin^2 \left(2\pi \cdot \frac{t}{\lambda'} \right) \right\} + \&c.$$

or by

$$S \cdot C \left\{ 1 - 4a \cdot \sin^2 \left(2\pi \cdot \frac{t}{\lambda} \right) \right\}.$$

Now this is the same with

$$\begin{aligned} S \left\{ C(1-4a) + C(4a - 4a \cdot \sin^2 \left(2\pi \cdot \frac{2t}{\lambda} \right)) \right\} &= \\ = (1-4a) \cdot S(C) + 4a \cdot S \left\{ C \cdot \cos^2 \left(2\pi \cdot \frac{2t}{\lambda} \right) \right\}. \end{aligned}$$

The first term of this expression represents a beam of white light of the intensity $1-4a$. The second, a compound tint of the intensity $4a$, which, diluted with the above-mentioned white light, forms the pallid tints of the transmitted series. If we disregard this dilution, and consider only the tint in its purity as it would appear were the white light suppressed, its expression

$$4a \cdot S \left\{ C \cdot \cos^2 \left(2\pi \cdot \frac{2t}{\lambda} \right) \right\} = 4a \left\{ S(C) - S \left(C \cdot \sin^2 \left(2\pi \cdot \frac{2t}{\lambda} \right) \right) \right\}$$

indicates that it is *complementary* to the tint represented by

$$S \left\{ C \cdot \sin^2 \left(2\pi \cdot \frac{2t}{\lambda} \right) \right\}.$$

But if we conceive a curve whose abscissa $= t$, and whose ordinate is $C \cdot \sin^2 \left(2\pi \cdot \frac{2t}{\lambda} \right)$, it is evident that

this will be precisely the undulating curve represented for each prismatic ray in fig. 134; and taking the sum of all the ordinates so drawn for each colour in the spectrum, we have the identical construction from which we derived the colours of the reflected rings in Art. 645. If, then, we take the series of tints so composed, and thence deduce their complements to white light, and dilute these complementary colours with white, in the proportion of $4a$ rays of the complementary colour to $1-4a$ of white, we shall have the series of transmitted tints which ought to result from the doctrine of interferences, and which, in fact, is observed.

664.
Case of
oblique
transmission
Fig. 135.

In the case of oblique transmission, let AC , BD , fig. 135, be the surfaces of the lamina, and Aa its thickness; and let AE be the surface of a wave of which the point A has just reached the first surface of the lamina; and let SA , SC , perpendicular to it, represent rays emanating from one origin S , then will a partial reflexion take place, and its intensity will be diminished in some certain ratio $1 : 1-a$ depending on the angle of incidence. The transmitted wave will be bent aside, taking the position Ab , and advancing along AB the refracted ray; so that when it reaches the position BF , the wave without the lamina will have the corresponding position FG . Here another partial reflexion will take place depending on the interior incidence, and we may denote by $(1-a)$ the transmitted portion, and by $(1-a) \cdot a$ the reflected portion. These portions set off together, from B , the former, with the velocity V due to the exterior medium, along the line BH parallel to SA , forming a wave which (provided S be sufficiently distant) may be regarded as a plane of indefinite extent moving uniformly with that velocity along BH . The latter portion proceeds along BC , according to the law of reflexion, with the velocity V' due to the medium of which the lamina is composed till it reaches C , where it undergoes another partial reflexion, and proceeds back along the line CD with a diminished intensity $= (1-a)$

light. a^2 , but with the same velocity V' till it reaches D, having described a space $= BC + CD = 2AB$ with that velocity. At D it undergoes another partial reflexion, and only a portion $= (1-a)(1-a) \cdot a^2$ is transmitted, which sets off from D along the line DI (parallel to BH) with the velocity V, that is, with the same velocity as the wave along BH. This wave may also be regarded as a plane of indefinite extent perpendicular to DI, and therefore parallel to the former. But they are not coincident; for the former, having the start of the latter, will have come into a position IHK in advance of the position DLM taken by the latter, and both the waves moving forwards now with the same velocity V will preserve this distance for ever unaltered. The interval LH we may term the interval of retardation. To determine it, we have to consider that the space BH is described by the former wave with a velocity V, while the latter describes BC + CD with the velocity V' , and therefore

$$BH = (BC + CD) \cdot \frac{V}{V'} = 2AB \cdot \frac{V}{V'} = 2t \cdot \sec \rho \cdot \mu,$$

putting μ for the relative index of refraction of the lamina, ρ for the angle of refraction aAB , and t for the thickness AB , because

$$V : V' :: \mu : 1.$$

Again, $BL = BD \cdot \cos DBL = DB \cdot \sin \phi$ (ϕ being the angle of incidence corresponding to ρ the angle of refraction,) $= 2aB \cdot \sin \phi = 2t \cdot \tan \rho \cdot \sin \phi$. Therefore the whole interval of retardation is equal to

$$2t \{ \mu \cdot \sec \rho - \tan \phi \cdot \sin \phi \} = \frac{2t \cdot \mu}{\cos \rho} (1 - \sin^2 \rho) = 2\mu t \cdot \cos \rho$$

because $\sin \phi = \mu \cdot \sin \rho$.

Thus, in virtue of the two internal reflexions, each wave which before entering the medium was single, will after quitting it be double, being followed at the constant interval $2\mu t \cdot \cos \rho$ by a feebler wave of the intensity above determined. The same being true of every wave of the system of which the ray consists, these two systems (considered as of indefinite duration) will be superposed on, and interfere with each other, according to the general principles before laid down.

Let λ be the length of an undulation in the lamina, then will $\mu\lambda$ represent that of an undulation in the surrounding medium. This is obvious, because the velocity in the latter being to that in the former as $\mu : 1$; and the same number of undulations being propagated in the same time through a given point in both cases, they must be more crowded, and therefore occupy less space in the one than the other in the ratio of the velocities.

Hence the differences of phases between the interfering systems at any point will equal

$$2\pi \cdot \frac{\text{interval of retardation}}{\mu\lambda} = 2\pi \cdot \frac{2t \cdot \cos \rho}{\lambda} = 2\pi \cdot \frac{2t'}{\lambda}, \text{ putting } t' = t \cdot \cos \rho,$$

and therefore the final resulting wave will be expressed by the equation

$$X = \sqrt{(1-a)(1-a)} \left\{ \cos \theta + a \cdot \cos \left(\theta + 2\pi \cdot \frac{2t'}{\lambda} \right) \right\},$$

which being resolved into the fundamental form $A \cdot \cos(\theta + B)$, as before, gives

$$A^2 = (1-a)(1-a) \cdot \left\{ 1 + 2a \cdot \cos \left(2\pi \cdot \frac{2t'}{\lambda} \right) + a^2 \right\},$$

and

$$\sin B = \frac{a \cdot \sin \left(2\pi \cdot \frac{2t'}{\lambda} \right)}{\sqrt{1 + 2a \cdot \cos \left(2\pi \cdot \frac{2t'}{\lambda} \right) + a^2}}.$$

Such are the general expressions for the intensity and change of origin of the compound transmitted ray. It is evident, however, that when a and a are small, which they always necessarily are in any but extreme cases, this value of A^2 reduces itself by neglecting their powers and products to

$$(1-a+a) - 4a \cdot \sin \left(2\pi \cdot \frac{t'}{\lambda} \right)^2,$$

which is exactly analogous to the expression in Art. 662, for the case of perpendicular incidence; and shows, that with the exception of a very trifling difference in the degree of dilution, the same laws of alternation in brightness, in homogeneous light, and of tint in white light, must hold good in both cases.

But there is one essential difference. The same tints will arise in the case of oblique incidence at the thickness t , which in that of perpendicular incidence is produced at the thickness $t \cdot \cos \rho$, because $t' = t \cdot \cos \rho$. Now this is always less than t , and therefore the tint produced at oblique incidences at the given thickness will be higher in the scale (or correspond to a less thickness) than in perpendicular; and, consequently, the rings, or fringes, so seen by transmission should dilate by inclining the lamina to the eye: The law of dilatation evidently, at moderate incidences, coincides nearly with Newton's rule; for this gives, on reduction, neglecting $\sin^4 \rho$,

Part III

665.

666. Undulations shorter in denser media.

667. General expression for the transmitted ray

668

Case of moderate obliquities.

669.

Dilatation of the rings explained.

Light.

$$\sec u = \sec \rho \left\{ 1 - \frac{1}{2} \cdot \frac{106}{107} (\mu - 1) \cdot \tan \rho^2 \right\},$$

which does not deviate very greatly from $\sec \rho$ at moderate incidences.

670. At great incidences the case is different, and the noncoincidence of the results of the undulatory doctrine with experiment might be drawn into an argument against it, were we sure that the law of refraction at extreme incidences, and with very thin laminae, does not vary sensibly from that of the proportional sines. This is, indeed, highly probable, as M. Fresnel has remarked, (*Mém. sur la Diffraction*, &c.) and as we have before had occasion to observe. The inquiry into which this would lead, is, however, one of the most delicate and difficult in physical optics, and the reader must be content with this general notice of a possible explanation of one of the many difficulties which still beset the undulatory doctrine.

671. The origin of the reflected rings may be accounted for in a similar way from the partial transmission of the waves reflected from the second surface back through the first, and their interference with the waves reflected immediately from the first. The relative intensities of these waves, (in general,) are a and $(1 - a)(1 - a) \cdot a$; or, in the case where a and a are both small, nearly in the ratio of $a : a$, and at a perpendicular incidence, very nearly in the ratio of equality. Hence their mutual destruction in the case of complete discordance will be much more complete than in the transmitted rings, and the colours arising, much less dilute than those of the latter, agreeably to observation.

672. There is, however, one consideration of importance to be attended to in the application of the undulatory doctrine to the reflected rings, which at first sight appears in the light of a powerful argument against its admissibility, viz. that if we apply the same reasoning to the reflected, as we have already done to the transmitted, rings, we should arrive at the conclusion, that their tints should be *precisely the same* and in the same order, beginning with a bright white in the centre; because *here*, the path traversed by the ray within the lamina vanishing, the waves reflected from the two surfaces ought to be in exact accordance, whereas it appears, by observation, that the reverse is the case, the central spot being black instead of white. It becomes necessary, then, to suppose, that in this case, half an undulation is lost or gained either by the wave reflected from the first or second surface. If this hypothesis be made, the phenomena of the reflected rings are completely represented on the undulatory system, for the compound wave reflected by the joint action of the two surfaces should be represented by the equation,

$$X = \sqrt{a} \cdot \cos \theta + \sqrt{a(1-a)(1-a)} \cdot \cos \left\{ \theta + 2\pi \cdot \frac{2t' - \frac{1}{2}\lambda}{\lambda} \right\}$$

and if this be put equal to $A \cdot \cos(\theta + B)$ we get

$$A^2 = a + a(1-a)(1-a) - 2\sqrt{a a(1-a)(1-a)} \cdot \cos \left(2\pi \frac{2t'}{\lambda} \right)$$

and in the case of a and a both very small

$$A^2 = (\sqrt{a} - \sqrt{a})^2 + 4 \cdot \sqrt{a a} \cdot \sin \left(2\pi \frac{t'}{\lambda} \right)^2$$

and at a perpendicular incidence, where $t' = t$, and where a and a may be supposed equal

$$A^2 = 4a \cdot \sin \left(2\pi \frac{t}{\lambda} \right)^2$$

673. Thus we see, that in this case the total intensity of the compound reflected wave + that of the transmitted (Art. 662) make up 1, the intensity of the incident wave; and thus, this supposition of the loss or gain of half an undulation is in no contradiction with the law of the conservation of the *vis viva*.

In fact, however, if we consider the mode in which the undulations are propagated, at the limit between two media, we shall see nothing contrary to dynamical principles in the loss of half or any part of an undulation in the transfer—for it cannot be supposed, that the density or elasticity of the ether changes abruptly at the surfaces of media, but that there intervenes some very minute stratum in which it is variable. In this stratum, therefore, the length of an undulation is neither exactly that corresponding to the denser, nor to the rarer medium, but intermediate, and of a magnitude perpetually varying. Therefore the number of undulations to be reckoned as added to the phase of the ray in traversing this stratum, will differ from what it would be if one medium terminated, and the other commenced abruptly. Without knowing the law of density, the limits between which it undergoes its change, or the exact mode in which the partial reflexion of a wave traversing it is performed, it is impossible to subject the point to strict calculation, we must rather submit to be taught by experiment, and content ourselves with such conclusions as we can deduce from observation. In the case before us, all that observation teaches us is, that there is half an undulation more of difference in the phases of two rays that have been reflected in the manner last considered, than in those of the two whose interference forms the transmitted rays. From some curious experiments of Dr. Young, too, we may gather that it is not in all cases *strictly half* an undulation of difference to be reckoned, but rather a variable fraction depending on the nature of the contiguous media.

675. The formulæ of Art. 672 show that it is only in the case of perpendicular incidence that the tints are pure, and that in all others, and especially at great obliquities, where a and a differ considerably, there will be a

Not contrary to dynamical principles

674. Not to the undulatory doctrine.

Light.

dilution of white light, and this is also agreeable to experience. At a perpendicular incidence, however, the minima of each homogeneous colour ought to be absolutely evanescent; so that if we were to remove the reflection of the upper surface of an object glass laid down on a plate, (or use a prism, so as to prevent its reaching the eye,) the intervals between the rings in homogeneous light ought to appear absolutely black. In the Newtonian doctrine this should not be the case, because the light reflected from the upper surface of the lamina of included air should still remain even in the minima of the rings. This then affords a positive means of deciding between the two theories. M. Fresnel describes an experiment made for this purpose, and states the result to be unequivocally in favour of that of undulations. (*Diffraction de la Lumière*, p. 11.)

Part III.

Experiment crucis between the two theories

§ V. Of the Colours of Thick Plates.

Under certain circumstances rings of colours are formed by plates of transparent media of considerable thickness. The circumstances under which they appear, in one principal case, are thus described by Newton, who first observed them, and who has applied his doctrine of the fits of easy reflexion and transmission to explain them, with singular ingenuity.

676

"Admitting a bright sunbeam through a small hole of one-third of an inch in diameter into a dark room, it was received perpendicularly on a concavo-convex glass mirror one quarter of an inch thick, having each surface ground to a sphere of six feet in radius, and the back silvered. Then holding a piece of white paper in the centre of its concavity, having a small hole in the middle of it to let the sunbeam pass, and after reflexion at the speculum to repass through it, the hole was observed to be surrounded with four or five coloured concentric rings or irises, just as the rings seen between object-glasses surround their central spot—but larger and more diluted in their colours". . . . "If the paper was much more distant from the mirror, or much less than six feet, the rings became more dilute and gradually vanished". . . . "The colours of these rings succeeded each other in the order of those which are seen between two object glasses, not by reflected but by transmitted light, viz. white, tawny white, black, violet, blue, greenish yellow, yellow, red, purple," &c. . . . "The diameters of these rings preserved the same proportion as those between the object-glasses, the squares of the diameters of the alternate bright and dark rings, reckoning the central white as a ring of the diameter 0, forming an arithmetical progression, beginning at 0. And in the case described, the diameter of the bright ring measured respectively $0, 1\frac{1}{2}, 2\frac{3}{8}, 2\frac{11}{12}, 3\frac{1}{8}$ ". . . . "Lastly, in the rings so formed by reflectors of different thicknesses, their diameters were observed to be reciprocally as the square roots of the thicknesses. If the back of the mirror was silvered, the rings were only so much the more vivid."

Newton's experiment with a glass mirror.

These various phenomena, and a variety of similar ones, some of more, some of less complexity, according to the variation of the distance, and obliquity of the mirror, and the curvature of its surfaces, Newton has explained very happily, by considering the fits of easy reflexion and transmission of that faint portion of the light which is irregularly scattered in all directions at the first surface of the glass, and which serves to render it visible. But for this explanation we must refer to his *Optics*, as our object here is more particularly and distinctly to show what account the undulatory doctrine gives of this phenomenon, which has hitherto been passed over rather cursorily, not without some degree of obscurity.

677

There is no surface, however perfectly polished, so free from small scratches and inequalities as not to reflect and transmit, besides those principal rays which obey the regular laws of reflexion and refraction, as dependent on the general surface, other, very much feebler, portions scattered in all directions, by which the surface is rendered visible to an eye anywhere placed, but most copiously in and about the direction of the regularly reflected and transmitted rays. It is the interference of these portions, scattered at the first surface by the ray in passing and repassing through it, nearly in its own direction, that the rings in question are attributed in the undulatory doctrine.

678.

Principle of explanation on the undulatory system.

Let FAD, EBG be the parallel surfaces of any medium exposed perpendicularly to a homogeneous ray emanating from a luminous point C, and incident at A. The chief portion will pass straight through A, and be reflected back from B to A again. But at A a scattering takes place, and the transmitted ray AB is accompanied by a diverging cone of faint rays Aa, Ab, Ac, &c., all which set out from A in the same phase of their undulations with the principal one from which they originate, so that A may be regarded as their common origin. Take Q, the focus of rays reflected at the second surface conjugate to A (if the surfaces be plane, Q and A are equidistant from B) and the cone of scattered rays, with the regularly reflected ray in its axis, will after reflexion diverge as from Q. Again, when they pass into the air again, if we take q the focus conjugate to Q of rays refracted at the surface FD, they will after refraction diverge from q, and by the nature of foci on the undulatory hypothesis, the undulations will be propagated in the air as if they had a common origin q placed in air; because, after refraction, the waves have the form of spheres diverging from q, and therefore every portion of their surfaces are equidistant from q; had they, therefore, really emanated from q, as separate rays, they must at the moment of such emanation have been all in one phase. Now, when the reflected beam reaches A a portion of it will again be scattered in a cone, having the regularly transmitted ray AG in its axis; and the rays AO, AN, AM, &c. of this cone will all have A for their origin, and will be in the same phase at their departure from A with the ray AG; but this is in the phase it would have had as emanated from q; hence, if we consider any point M out of the directly transmitted ray AG, it will be reached at once by a wave belonging to each diverging cone, the one along qM from q and the other along AM from A, and the difference of routes is equal to qA + AM - qM. Therefore, when M is very nearly coincident with G, this is very small and at G vanishes, or the waves are in exact accordance. As M recedes from G it increases, and when it becomes

679.

Its application. Fig. 136.

Light.

half an undulation, the waves are in complete discordance and annihilate each other, and so on alternately. There fore, as this is true of all rays in conical surfaces round A G as an axis, equally inclined with A M, q M, if we place a white screen at G, it will appear marked with alternate dark and bright rings round a bright centre. To determine their diameters we need only put $q A + A M - q M = n \cdot \frac{\lambda}{2}$, or, if we take $q A = a$, $A G = r$, $G M = y$,

$$a + \sqrt{r^2 + y^2} - \sqrt{(a+r)^2 + y^2} = n \frac{\lambda}{2}$$

If we resolve this equation neglecting y^2 , we find

$$y = \sqrt{n} \cdot \sqrt{\frac{\lambda}{a} \cdot r(a+r)}$$

which, on substituting 0, 1, 2, 3, &c. in succession for n , shows that the successive diameters of the alternate dark and bright rings are in the progression of the square roots of those numbers.

680.

Law of the
diameters of
the rings.

If the thickness of the plate be small compared to the distance of the screen, a will also be small, and the value of y becomes

$$y = r \cdot \sqrt{n} \cdot \sqrt{\frac{\lambda}{a}},$$

which shows that for rays of a given refrangibility the diameters of the rings are as the distance of the screen directly, and the square root of the thickness of the plate inversely.

681.

Of their
colours.

Lastly, the diameter of a ring of the same order in different homogeneous lights, are as the square roots of the lengths of their undulations. Now, this is the very same law that governs the diameters of the rings formed between object-glasses. Consequently, if instead of homogeneous we consider white light, we ought to have a succession of coloured rings whose tints agree precisely with the transmitted series in that experiment.

682.

Concentration of the
rings from
all points of
the surface.
Fig. 137.

But the rays so formed, by rays scattered from a single point A, would be too feeble to be visible. If, however, we suppose the surfaces to be concentric spheres having G in their common centre, as in fig. 137, then any rays G A, G A' falling on any points whatever of their surfaces will depict, on screens G M, G M' respectively perpendicular to them as G, equal systems of rings having G in their common centre; and, when the arc A A' is not very great, the screens may be regarded as coincident (for in that case $B M - M A = B M' - M A'$) and the rings from every point of the surface, exactly superposed on each other, and being thus increased in intensity in proportion to the area of the exposed surface, become visible.

683.

Newton's
experiment
particularly
considered.

Now this is exactly Newton's case, for the sun being a luminary of a considerable diameter, the hole in the centre of the spheres may be regarded as a portion of the sun of that size, actually placed there. Of this, every indivisible point may be regarded as the origin of a system of waves, and as depicting on the screen its own set of rings. These, were the hole infinitely small, would be infinitely more clear and pure in their tints than the transmitted rings between object-glasses, because they are not (as in those rings) diluted with the great quantity of white light which escapes interference. But owing to the size of the hole, their centres are not exactly coincident, and therefore their tints mix and dilute each other, and that the more the larger the hole is.

684.

If c be the thickness of the glass, since Q is the conjugate focus of A, on the surface B whose radius we will call $r + c$ putting $G A = r$, we have, by Art. 249, $R Q = \frac{r+c}{r-c} \cdot c$, $A Q = \frac{2rc}{r-c}$; and, by Art. 249

$A q = a = \frac{2cr}{2c - \mu(r+c)}$, taking μ for the refractive index; and when c is small compared with r , we get

$$a = \frac{2c}{\mu}; \quad y = r \cdot \sqrt{n} \cdot \sqrt{\frac{\mu}{2} \cdot \frac{\lambda}{c}}$$

showing that the diameters of the rings are in this case in the subduplicate ratio of the refractive index of the glass directly, and of its thickness inversely.

685

If we reduce this value to numbers, taking $\mu = \frac{3}{2}$, $n = 4$, $r = 6$ feet $= 72$ inches, and $\lambda = \frac{2}{89000}$ = the length of an undulation for yellow rays $\frac{2}{90000}$ nearly, we find, for the diameter of the second bright ring in yellow light, (which corresponds to the brightest part of the same ring in white,)

$$2y = 72 \times \sqrt{4 \cdot \frac{3}{4} \cdot \frac{2}{90000}} \cdot 4 = 2.35,$$

which agrees almost precisely with Newton's measure $2\frac{3}{8}$, or 2.375.

686

Case of
oblique
incidence.

When the mirror is inclined to the incident beam the phenomena become more complicated, and have been elegantly described by Newton, (*Optics*, book ii. part iv. obs. 10.) In this case, the axes of the two interfering cones of scattered rays, which are always the incident and reflected rays, are no longer coincident. But the same principles apply equally to this case in all other respects, and the reader may exercise himself in tracing their consequences.

687.

Phenomena
observed by
the Duke of
Chaulnes
and

The Duke de Chaulnes found similar rings to be exhibited when the surface of the mirror was covered with a thin film of milk dried on it, so as to make a delicate semitransparent coating, or even when a fine gauze or muslin was stretched before it; see the account of his experiments in the *Mém. Acad. Sci. Paris*, 1705; and

Light.

Sir William Herschel (*Phil. Trans.* 1807) describes a pleasing experiment, in which rings were produced by strewing hair powder in the air before a metallic mirror on which a beam of light is incident, and intercepting the reflected ray by a screen. The explanation of these phenomena seems, however, to depend on other applications of the general principle, and will be better conceived when we come to speak of the colours produced by diffraction.

Part III.
Sir W.
Herschel.

Dr. Brewster, in the *Transactions of the Royal Society of Edinburgh*, has described a series of coloured fringes produced by thick plates of parallel glass, which afford an excellent illustration of the laws of periodicity observed by the rays of light in their progress, whether, as in the Newtonian doctrine, we consider them as subjected to alternate fits of easy reflexion and transmission, or, as in the undulatory hypothesis, as passing through a series of phases of alternately direct and retrograde motions in the particles of ether, in whose vibrations they consist. We may here remark, once for all, that the explanations which the undulatory doctrine affords of phenomena of this description, may, for the most part, be translated into the language of the rival hypothesis; so as to afford, with more or less plausibility and occasional modifications, a result corresponding with observation. It is not, therefore, among phenomena of this class that we must look for the means of deciding between them. We shall adopt, therefore, in the remainder of this essay, the undulatory system, not as being at all satisfied of its reality as a *physical fact*, but regarding it as by far the simplest means yet devised of grouping together, and representing not only all the phenomena explicable by Newton's doctrine, but a vast variety of other classes of facts to which that doctrine can hardly be applied without great violence, and much additional hypothesis of a very gratuitous kind.

The fringes in question are seen when two parallel plates of glass of exactly equal thickness (portions of the same plate) are slightly inclined to each other, (at any distance,) and through them both, at nearly a perpendicular incidence, a circular luminary of 1° or 2° in diameter (a portion of the sky, for instance) is viewed. There will in this case be seen, besides the direct image, a series of lateral images reflected between the glasses, and growing fainter and fainter in succession as they are formed by 2, 4, 6, or more internal reflexions; and of which all but the first is so faint as scarcely to be visible, except in very strong lights. The direct image is colourless; but the reflected one is observed to be crossed with fifteen or sixteen beautiful bands of colour, parallel to the common section of the surfaces of the plates. Their breadth diminishes rapidly as the inclination of the plates increases. When the plates employed were 0.121 inch in thickness, and inclined at an angle of $1^\circ 11'$ to each other, the breadth of each fringe measured $26' 50''$, and at all other inclinations their breadth was inversely as the inclination. At oblique incidences its fringes are seen when the plane of incidence is at right angles to the principal section of the plates, but are at their maximum of distinctness when parallel to it.

To understand their production, let us call the surfaces of the plates in order, reckoning from that on which the incident light first falls, A, *a*, B, *b*; and let us consider a ray, or system of waves emanating from a common origin at an infinite distance. Then, when this ray falls on the plates it will at every surface undergo a partial reflexion, and the remainder will be transmitted; each of the several portions will be again subdivided whenever it meets either surface. So that either image will, in fact, consist of several emergent rays, parallel in their final directions, but which have traversed the glasses by very different routes. Thus the direct or principal image will consist of

1. The chief portion of the whole incident light, refracted at A, at *a*, at B, and at *b*, and emergent parallel to the incident ray, which we will represent by A *a* B *b*.

2. A portion refracted at A, reflected at *a*, reflected again at A, refracted again at *a*, at B and at *b*, and emergent parallel to the incident beam. This we will denote thus, A *a'* A' *a* B *b*; the letters denoting the surfaces, the accent reflexion, and its absence refraction.

3. A portion which has undergone two similar reflexions in the interior of the second plate, and which in the same manner may be represented by A *a* B *b'* B' *b*.

4. Other portions which have undergone respectively four, six, &c. reflexions to infinity within either of the plates, and which may be represented by such combinations as A *a'* A' *a'* A' *a* B *b*, A *a* B *b'* B' *b'* B' *b*, or, for brevity, by A (*a'* A')² A B *b*, A A B (*b'* B')² B, &c.; but these latter portions are too faint to have any sensible influence on the light of the direct image with which they are confounded.

The first lateral reflected image will consist of four principal portions which have undergone two reflexions each, viz.

A *a* B' *a'* B *b*; A *a* B' *a* A' *a* B *b*; A *a* B' *b'* B' *a* B *b*; A *a* B' *b'* *a* A' *a* B *b*;

all which will emerge parallel. Besides these there are infinite others, formed by a greater number of reflexions, and by the portions A *a'* A' *a* of the incident beam reflected within the first glass; but these are all too faint materially to affect the image in question, which therefore we may regard as composed solely of the four rays just enumerated. Now if we cast our eye on the figure, (138,) we see the course pursued by each of these portions 1, 2, 3, 4; and it is evident that the first portion has traversed the thickness twice, and the interval between the glasses three times, or nearly; neglecting at present all consideration of the inclination of the plates $2t + 3i$. In like manner, the portion 2 will have traversed $4t + 3i$; the portion 3, $4t + 3i$; and the portion 4, $6t + 3i$. Hence it appears that the portions 1 and 4 differ in their routes by nearly four times the thickness of the glass, and can therefore produce no colours; but the other portions, at a perpendicular incidence, would not differ at all, and at very small inclinations of the plates, and of the incident ray, will only differ by reason of the small differences of the inclinations at which they traverse their respective thicknesses and intervals. They will, therefore, interfere so as to produce colour; and this will be dependent on the interval of retardation of one ray behind the other, arising from the varying obliquity of the ray which enters the eye.

Now when we look at a luminous image of sensible magnitude, the rays by which we see its several points

688.
Dr. Brew-
ster's
fringes seen
in thick
plates.

689.
Described

690.
Explained.

691.

Fig. 138

692.

Light.
Isochromatic lines defined.

are incident in all planes, and at all inclinations. Hence, the image seen will appear of different colours in its different points, and the disposition of these colours will follow the law, whatever it be, which regulates the interval of retardation. The colours, therefore, will be arranged in bands, circles, or other forms, according to the form of the curves arising geometrically from the consideration of equal intervals of retardation prevailing in every point of their course. Such curves, now and hereafter, we shall term *isochromatic lines*, or lines of equal tint, measuring in all cases the *tint* numerically by the number of undulations, or parts of an undulation of mean yellow light to which the interval of retardation is equal.

693.
Fig. 139.

Let us, then, first consider the case when the ray is incident in a plane perpendicular to the common section. In this case, fig. 139, let KLMN be a ray formed by the union of two rays S A a B b I K L and S C E F G H K L, whose courses through the system are similar to 2 and 3, fig. 138. Draw A D perpendicular to S C, then will the interval of retardation be equal to

$$\{DC + CE + EF + FG + GH + HK\} - \{Aa + aB + Bb + bI + IK\} \\ = DC + (EF - aB) + (FG - IK) + 2(KH - Bb),$$

the first three terms being performed in air, the last in glass. Now, without entering into a trigonometrical calculation, it is evident that this will be very small at a perpendicular incidence, and will increase rapidly as the angle of incidence varies; and that (the inclination of the plates remaining constant) it will increase by nearly equal increments, as the incidence varies by equal changes from 0 on either side of the perpendicular. Therefore, in a direction at right angles to the common section of the surfaces the tints will vary rapidly, increasing on either side of the perpendicular incidence; and at very moderate obliquities on either side, the interval of retardation will become too great for the production of colour. On the other hand, if we conceive the rays S A, S C, to be incident in a plane very nearly parallel to the principal section, then will the points K and G be situated, not, as in the figure, at different distances from P, but at very nearly the same; so that (whatever be the incidence) K I will very nearly equal G F, and for the same reason F E will very nearly equal a B. Moreover, in this case G K will be very nearly equal to F I, and the angles of internal incidence will be also very nearly equal, so that H G + G K will differ very little from B b + b I, and I B will be very nearly equal to G K, and therefore to I F, so that the point F will almost exactly coincide with B, and the rays S A a B, S C E F will coincide almost precisely, making D C = 0; and these approximate equalities and coincidences will continue for great variations in the *angle* of incidence, provided the *plane* of incidence be unaltered. The interval of retardation, then, will in this case depend very little on the angle of incidence; so that in a direction parallel to the common section of the surfaces, the tints will vary but little. Hence it appears that they will be arranged in the manner described by Dr. Brewster, viz. in fringes parallel to that line. Their general analytical expression is, however, rather too complex to be now set down, though very easily investigated from what has been said.

694.
Fig. 140.

By intercepting the principal transmitted beam in the direct image, and receiving on the eye only those portions of the rays going to form it whose curves are as in fig. 140, or the portions A a' A' a B b, and A a B b' B' b, Dr. Brewster succeeded in rendering visible a set of coloured fringes, which in general are diluted and concealed in the overpowering light of the direct beam. They originate evidently in the interference of these two rays, whose courses are each represented by $4t + i$, and would therefore be strictly equal were the plates exactly parallel. Their theory, after what has been said, will be obvious on inspection of the figure, as well as those of all the rest of the systems of fringes which he has described in that highly curious and interesting memoir.

695
Fringes between glass films.

Mr. Talbot has observed, when viewing films of blown glass in homogeneous yellow light, and even in common daylight, that when two films are superposed on each other, bright and dark stripes, or coloured bands and fringes of irregular forms, are produced between them, though presented by neither separately. These are obviously referable to the same principle, the interference taking place here between rays respectively twice reflected within the upper lamina, and once reflected at the upper surface of the lower lamina, or else between rays one of which is thrice reflected in the mode represented by A a B' a' B' a A, and the other in that represented by A a B' a A' a' A, the interval between the glasses being supposed to be exactly equal to the thickness of the upper one in both cases, a condition which is sure to obtain somewhere when the laminæ are curved. A still more curious and delicate case of the production of similar fringes has been noticed by Professor Amici, to take place when two of the blue feathers of the wing of the *Papilio Idas* (a species of butterfly) are laid on each other in the field of his powerful and exquisite microscopes. These feathers he describes as small plates of perfect transparency, and uniformly and delicately striated over their whole surface. The fringes in question are formed between them, and vary in breadth, form, and situation, according to the manner in which the feathers are superposed. Their origin seems to be independent of the striæ however, and is easily understood on the principles above explained. The same may be said of the colours observed by Mr. Nicholson in combinations of parallel glasses of unequal thickness. Suppose, for instance, that instead of the plates having exactly equal thicknesses, their thicknesses t, t' differ by a very minute quantity, then the course of the rays A a' A' a B b and A a B b' B' b will (at a perpendicular incidence) be respectively $3t + i + t'$ and $t + i + 3t'$, (supposing the plates strictly parallel,) and the difference of their routes is $2t - 2t'$; so that if this be exceedingly minute, colours will arise, or, if not, may be produced by a slight inclination of the plates to each other, and so of an infinite variety of cases which may arise.

§ VI. Of the Colours of Mixed Plates.

The colours hitherto described have been referred to the interference of rays rigorously coincident with each other throughout their whole course, after the point where they begin to be superimposed. Such interfering rays, or systems of waves, being united into a point on the retina, that point is agitated by the sum or difference of their actions, and the sensation produced is according. But if this coincidence be only approximate, as, if two systems of waves be propagated from origins so nearly coincident in angular situation from the eye, that their images formed on the retina shall be too close to be distinguished by the mind from the image of a single point, the impressions produced will still be confounded together; or rather, we ought to say, the mechanical action on one point will be propagated through the substance of the retina to the other, and a sensation corresponding to their mean or average effect will be produced. If, then, the rays concentrated on contiguous points of the retina be in exact discordance, and of equal intensity, a mutual destruction will take place, as if they fell on one mathematical point; if in exact accordance, they will increase each others effects, and so for the intermediate states.

696.

Interference of rays not strictly coincident.

To apprehend this more fully, we must consider that the impression of light appears to spread on the retina to a certain extremely minute distance all around the mathematical focus of the rays concentrated by the lenses of the eye. Thus the image of a star is never seen as a *point*, but as a disc of sensible size, and that the larger as the light is stronger. Thus, too, the bright part of the new moon is seen, as it were, larger than the faintly illuminated portion of its disc projecting beyond it as an acorn cup beyond the fruit, &c. This effect is termed *irradiation*, and is manifestly the consequence of an organic action such as we have described.

697.

Irradiation.

It follows from this, that when waves emanate from origins *undistinguishably near*, they may be regarded in their effects on the eye as emanating from origins strictly in one and the same right lines, the direction of the joint ray; and the laws of their interferences will be precisely the same, considered in their effect on vision, as if the lenses of the eye were away, and the retina were a mere screen of white paper, on a single physical point of which (*viz.* the point where the images concentrated by the lenses *would* have fallen) the interfering undulations propagated simultaneously from the two origins fell, and agitated it with a vibration equal to their resultant.

698.

This premised, we are in a condition to appreciate the explanation afforded by the undulatory doctrine of the phenomena of mixed plates. They were first noticed (says Dr. Young) by him "in looking at a candle through two pieces of plate glass with a little moisture between them. He thus observed an appearance of fringes resembling the common colours of thin plates; and upon looking for the fringes by reflexion, found that the new fringes were always in the same direction as the others, but many times larger. By examining the glasses with a magnifier, he perceived, that wherever the fringes were visible, the moisture was intermixed with portions of air producing an appearance similar to dew." "It was easy to find two portions of light sufficient for the production of these fringes; for the light transmitted through the water moving in it with a velocity different from that of light passing through the interstices filled only with air, the two portions would interfere with each other and produce effects of colour according to the general law. The ratio of the velocities in water and air is that of three to four; the fringes ought therefore to appear where the thickness is six times as great as that which corresponds to the same colour in the common case of thin plates; and upon making the experiment with a plane glass and a lens slightly convex, he found the sixth dark circle actually of the same diameter as the first in the new fringes. The colours are also easily produced when butter or tallow is substituted for water, and the rings then become smaller in consequence of the greater refractive density of the oils; but when water is added so as to fill up the interstices of the oil, the rings are very much enlarged; for here the difference of velocities in water and in oil is to be considered, and this is much smaller than the difference between air and water. All these circumstances are sufficient to satisfy us of the truth of the explanation, and is still more confirmed by the effect of inclining the plates to the direction of the light; for then, instead of dilating like the colours of thin plates, these rings contract, and this is the obvious consequence of an increase of the lengths of the paths of the light which now traverses both media obliquely, and the effect is everywhere the same as that of a thicker plate. It must, however, be observed, that the colours are not produced in the whole light that is transmitted through the media; a small portion only of each pencil passing through the water contiguous to the edges of the plate is sufficiently coincident with the light transmitted through the neighbouring portions of air to produce the necessary interference; and it is easy to show that a considerable portion of the light that is beginning to pass through the water will be dissipated laterally by reflexion at its entrance, on account of the natural concavity of the surface of each portion of the fluid adhering to the two surfaces of the glass, and that much of the light passing through the air will be scattered by refraction at the second surface. For these reasons the fringes are seen when the plates are not directly interposed between the eye and the luminous object." (Young, *Phil. Trans.* 1802; *Account of some Cases of the Production of Colours.*) To see the phenomena to advantage, we may add, it is only necessary to rub up a little froth of soap and water almost dry between two plane glasses, and hold them at a distance from the eye between it and a candle, or the reflexion of the sun on any polished convex object. If two slightly convex glasses, or a plane and a convex one be used, the colours are seen arranged in rings.

699.

Phenomena of mixed plates.

§ VII. Of the Colours of Fine Fibres and Striated Surfaces.

700.
Interference
of rays re-
flected from
points or
lines very
near each
other.
Fig. 141.

If two points supposed capable of reflecting light in all directions (as two infinitely small spheres, &c.) be so near each other as to appear to the eye as one, and if rays from a common origin reflected from them reach the eye, they will interfere; and if the light be homogeneous, its intensity will vary periodically, with an interval of retardation corresponding to the difference of their paths; if white, the colour of the mixed reflected ray will be the same as if it had been transmitted through a plate of air of a thickness equal to that difference, but deprived of its diluting white. Suppose two exceedingly fine cylindrical polished fibres to be placed at right angles to the line of sight, and parallel to each other, as in fig. 141, as ABC, abc ; and let S be a luminous point very distant with respect to the interval of the fibres, and E the eye, placed so as to receive the reflected rays BE, bE , which, by supposition, are near enough to interfere. Then the differences of phases of the rays on the

retina is evidently equal to $2\pi \times \frac{(Sb + bE) - (SB + BE)}{\lambda} = 2\pi \cdot \frac{bx + by}{\lambda}$, supposing Bx and By

perpendicular to Sb and bE . If, then, we suppose I and i to be the angles of incidence of the rays SB, EB on the plane in which the axes of the two cylinders AC, ac lie, and put Bb their distance equal to a , we have for the difference of phases

$$2\pi \cdot \frac{a}{\lambda} \cdot (\sin I + \sin i).$$

Hence, if a remain the same, this will vary with the obliquity both of the incident and reflected ray to the plane of the axes of the fibres; and, therefore, if that plane be turned about an axis parallel to the fibres, a succession of colours analogous to the transmitted series of those of their plates, but much more vivid, will be seen, as if reflected on them.

701.
Colours of
scratches on
striae.

Any extremely fine scratch on a well polished surface may be regarded as having a concave, cylindrical, or, at least, a curved surface capable of reflecting the light equally in all directions; this is evident, for it is visible in all directions. Two such scratches, then, drawn parallel to each other, and then turned round an axis parallel to both in the sunshine, ought to affect the eye in succession with a series of colours analogous to those of thin plates. This is really the case. Dr. Young found, on examining the lines drawn on glass in Mr. Coventry's micrometric scales, each of them to consist of two or more finer lines exactly parallel, and at a distance of about one 10,000th of an inch. Placing the scale so as to reflect the sun's light at a constant angle, and varying the inclination of the eye, he found the brightest red to be produced at angles whose sines were in the arithmetical progression 1, 2, 3, 4.

702.
Of systems
of equi-
distant
parallel
lines.

In the beautiful specimens of graduation on glass and steel produced by Dr. Wollaston, Mr. Barton, and M. Fraunhofer, single lines exactly parallel to each other, and distant in some cases not more than one 10,000th of an inch, and at precisely equal intervals, are drawn with a diamond point. If the eye be applied close to a reflecting or refracting surface so striated, so as to view a distant, small, bright light reflected in it, it will be seen accompanied with splendid lateral spectra, which evidently originate in this manner. They are arranged in a straight line passing through the reflected, colourless image, and at right angles to the direction of the striae. Their angular distances from each other, the succession of their colours, and all their other phenomena, are in perfect agreement with the above explanation. Their vividness depends on the exact equality of distance between the parallel lines, which causes the lateral images produced by each pair to coincide precisely in distance from the principal image, and thus to produce a multiplied effect. If the distance of the lines be unequal, the images from different pairs, not coinciding, blend their colours, and produce a streak, or ray of white light. This is the origin of the rays seen darting, as it were, from luminous objects reflected on irregularly polished surfaces. These colours may be transferred, by impression from the surface originally graduated, to sealing wax, or other soft body; or from steel, by violent pressure, to softer metals. It is in this way that those beautiful striated buttons and other ornaments are produced, which imitate the splendour and play of colours of the diamond.

703.
Alleged
analogy
between
colours of
striated
surfaces
and certain
musical
tones con-
sidered.

Dr. Young has assimilated the colour thus produced when a beam of white light strikes on a succession of parallel equidistant lines, to the musical tone heard when any sudden sound is echoed in succession by a series of equidistant bars having flat surfaces situated in a direction perpendicular to the line in which they are arranged, for instance, an iron railing. It is evident that such echoes will reach the ear in succession, at precisely equal intervals of time, each being equal to the time taken by sound to traverse twice the space separating the bars; and thus producing on the ear, if the bars be sufficiently numerous, the effect of a musical sound. (*Phil. Trans.* 1801; *On the Theory of Light and Colours*.) This explanation, however, appears to us, we confess, more ingenious than satisfactory. The pitch of the musical tone produced by the echoes is independent of the sound echoed, which may be a single blow, or a noise, (i. e. a sound consisting of non-periodic vibrations,) and requires for its production a number of echoing bars sufficient to prolong the echoes a sensible time. On the other hand, the light reflected from parallel striae depends for its colour wholly on the incident ray, being red in red light, yellow in yellow, &c.; and is produced equally well from two or from twenty, as from a million of such reflecting lines. The intensity, not the colour,—the magnitude, not the frequency of the impression made on the retina by the reflected rays, is modified by their interference. We think it necessary to point out this defect in the illustration in question, inasmuch as it has become popular for its ingenuity, and *prima facie* plausibility; while, in reality, it is calculated to give very erroneous impressions of the analogy between sound and light.

Light.

A single scratch or furrow in a surface may, as that eminent philosopher has himself remarked, produce colours by the interference of the rays reflected from its opposite edges. A spider's thread is often seen to gleam in the sunshine with the most vivid colours. These may arise either from a similar cause, or from the thread itself as spun by the animal, consisting of several, agglutinated together, and thus presenting not a cylindrical, but a furrowed surface. Part III.
704.
Colours of a spider's web, &c.
705.
Of mother of pearl

The phenomena exhibited by light reflected from and refracted through the polished surface of mother of pearl, are, no doubt, referable in great measure to the same principle, so far as they depend on the structure of the surface. Dr. Brewster has described them in a most curious and interesting Paper, (published in the *Phil. Trans.* 1814, p. 397;) and a writer in the *Edinburgh Philosophical Journal*, vol. ii. p. 117, has added some further particulars illustrative of the curious and artificial structure of this singular body. Every one knows that mother of pearl is the internal lining of the shell of a species of oyster. It is composed of extremely thin laminæ of a tough and elastic, yet at the same time hard and shelly substance, disposed parallel to the irregular concavity of the interior of the shell. When, therefore, any portion of it is ground and polished on a plane tool, the artificial surface so produced intersects the natural surfaces of the laminæ in a series of undulating curves, or level-lines, which are nearer or farther asunder, according to the varying obliquity of the artificial to the natural surfaces. As these laminæ adhere imperfectly to each other, their *feather-edges* become broken up by the action of the powders, &c. used in grinding and polishing them, so as to present a series of ridges or escarpments arranged (when any very small portion of the surface only is considered) nearly parallel to, and equidistant from each other, which are distinctly seen with a microscope, and which no polishing in the least degree obliterates or impairs. The light reflected, therefore, or dispersed on their edges, will interfere and produce coloured appearances in a direction perpendicular to that of the striæ. This is, in fact, their situation; but the phenomena are modified in a very singular manner by the peculiar form of the edges and hollows, which results, no doubt, from the crystalline structure of the pearl. That it is the configuration *only* of the surface on which they depend, is evident from the remarkable fact, that, like the colours described in Art. 702, they may be transferred, by impression, to sealing wax, gum, resin, or even metals, with little or no diminution of their brilliancy; and the impression so transferred, if examined by the microscope, is found to exhibit a faithful copy of the original striæ, though sometimes so minute as hardly to exceed one 3000th of an inch in their distance from each other. For a particular description of this very curious and beautiful class of phenomena, however, our limits oblige us to refer to the original memoirs already cited, especially as their theory is still accompanied with some obscurity.

§ VIII. Of the Diffraction of Light.

When an object is placed in a very small beam of light, or in the cone of rays diverging from an extremely small point, such as a sunbeam admitted through a small pin-hole into a dark chamber, or, still better, through an opening of greater size, behind which a lens of short focus is placed, so as to form an extremely minute and brilliant image of the sun from which the rays diverge in all directions, its shadow is observed to be bordered externally by a series of coloured fringes which are more distinct the smaller the angular diameter of the luminous point, as seen from the object. If this be much increased, the shadow and fringes formed by its several points, regarded each as an independent luminary, overlap and confuse each other, obliterating the colours, and producing what is called the *penumbra* of the object; but when the luminous point is extremely minute, the shadow is comparatively sharp, and the fringes extremely well defined. 706.
Fringes formed exterior to the shadows of bodies in a small beam of light.

These fringes (which were first described by Father Grimaldi in a work entitled *Physico-Mathesis de Lumine*, Bologna, 1665, and afterwards more minutely by Newton in the third book of his *Optics*) surround the shadows of objects of all figures, preserving the same distance from every part, like the lines along the sea-coast in a map; only, where the object forms an acute, salient angle, the fringes curve round it; and where it makes a sharp, reentering one they cross, and are carried up to the shadow at each side, without interfering or obliterating each other. In white light three only are to be seen, whose colours, reckoning from the shadow, are black, violet, deep blue, light blue, green, yellow, red; blue, yellow, red; pale blue, pale yellow, pale red. In homogeneous light they are, however, more numerous, and of different breadths, according to the colours of the light, being narrowest in violet and broadest in red light, as in the coloured rings between glasses; and it is by the mutual superposition of the different sets of fringes for all the coloured rays that their tints are produced, and their obliteration after a few of the first orders caused. 707.
Their colours, &c. described.

The fringes in question are absolutely independent of the nature of the body whose shadow they surround, and the form of its edge. Neither the density or rarity of the one, nor the sharpness or curvature of the other, having the least influence on their breadth, their colours, or their distance from the shadow; thus it is indifferent whether they are formed by the edge or back of a razor, by a mass of platina or by a bubble in a plate of glass, (which, though transparent, yet throws a shadow by dispersing away the light incident on it,) circumstances which make it clear that their origin has no connection with the ordinary refractive powers of bodies, or with any *elective* attraction or repulsions exerted by them on light; for such forces cannot be conceived as independent of the *density* of the body exerting them, however minute we might regard the sphere of their action. 708.
Are independent of the body casting the shadow.

To see the fringes in question, they may be received on a smooth, white surface, and examined and measured thereon by contrivances which readily occur; this was the mode pursued by Newton. M. Fresnel, however, having (to avoid the inconvenience of intercepting the light by the interposition of the observer) received them on an emerald glass plate, was enabled, by placing himself behind it, to approach near enough to examine and measure them. 709.
M. Fresnel's method of examining them.

Light. them with a magnifier. In so doing, however, he observed, that when thus once brought under inspection, they continued visible, and were indeed much brighter and more distinct in the focus of the lens (as if depicted in the air) even when the emerald glass was altogether withdrawn; and this fortunate observation, by enabling him to avoid the use of a screen altogether, and to perform all his measurements of their dimensions by the aid of a micrometer, put it in his power to examine them with a degree of minuteness and precision no other way attainable, and fully adequate to the delicacy of the inquiry: for it is manifest that the fringes, being seen as they would be formed if received on a screen in the focus, may be regarded as any other optical image formed in the focus of a telescope, viewed with any magnifier, and treated in all respects as such images.

710. Whatever mode of examining them we adopt, however, we shall observe the following facts:

Their phenomena. 1st. Their distances, inter se. *Phenomenon 1.* That, *cæteris paribus*, the distances from each other and from the border of the shadow diminishes as the screen on which they are received, or the plane in the focus of the lens in which they are formed, approaches the border of the opaque body, and ultimately coincides with it, so that they seem to have their origin close to the edge of the body.

711. They are propagated in curved lines. *Phenomenon 2.* That they are not, however, propagated in straight lines from the edge of that body to a distance, but in hyperbolic curves, having their vertices at that edge; and therefore that it is not one and the same light which forms one and the same fringe at all distances from the opaque body. To explain this, conceive the distances of the fringes from each other and from the shadow measured accurately at a great variety of distances from the edge of the body; then, were they propagated in straight lines, and were each fringe really the axis of a pencil of rays emanating from a point at that edge, their intervals and distances from the shadow ought to be proportional to the distances from the edge of the body; but it is not so, in fact,—the former distances increasing as we recede from the opaque body much more rapidly at first, and less so as we recede, than according to the law of proportionality; and if the locus of each fringe be laid down from such measures, it will be found to be an hyperbolic curve having its convexity outwards or from the shadow. Thus in fig. 142 O is the luminous point, A the edge of the body, and GH a screen perpendicular to the straight line OA, C the border of the visible shadow, and D, E, F the places of the successive minima of the fringes in a line at right angles to the edge of the shadow. If the screen be brought nearer to the body A as at *gh*, and if *c, d, e, f* be the points corresponding to CDEF, their loci will be the hyperbolas *AcC, AdD, &c.*

Fig. 142. 712. It will be noticed also that the border C of the visible shadow is not coincident with B, that of the geometrical one, which lies in the straight line OA, grazing the edge of the object. The deviation is difficult to perceive in the shadow of a large body, having nothing to measure from; but if we examine those of very narrow bodies, as of a hair, for instance, in such a beam of light as described, we shall find on measuring the total breadth of the shadow a full proof of this. This fact was observed by Grimaldi. The limit of the visible shadow also follows the same law of curvilinear propagation as the fringes. Thus, Newton found the shadow of a hair one 280th of an inch in diameter placed at 12 feet distance from the luminous point, to measure at 4 inches from the hair $\frac{1}{60}$ inch, or upwards of 4 diameters of the hair, at two feet, $\frac{1}{20}$ inch, or 10 diameters; while at 10 feet it measured only $\frac{1}{8}$ inch, or 35 diameters, instead of 120, which it should have done if the rays terminating the shadow had proceeded in straight lines; or rather, to speak more correctly, if the shadow were bounded by straight lines.

713. To account for these remarkable facts, Newton supposes that the rays passing at different distances from the edges of bodies are turned aside outwards, as if by a repulsive force; and that those nearest are turned more aside than those more remote, as in fig. 143, where X is a section of the hair, and AD, BE, CF, &c. rays which pass at different distances beside it, and which are turned off at angles rapidly diminishing as the distance increases in directions DG, EH, FI, &c. It is manifest that the curve WYZ, to which all these deflected rays are tangents, and within which none can enter, will be convex outwards; and its curvature will be greatest at the vertex W, and will diminish continually as it recedes from X, being, in fact, the caustic of all the deflected rays.

714. This will be the boundary of the visible shadow. To account for the fringes, he supposes (*Optics*, book iii. question 3) that each ray in its passage by the body undergoes several flexures to and fro, as in fig. 144 at *a, b, c*; and that the luminous molecules, of which that ray consists, are thrown off at one or other of the points of contrary flexure, or other determinate points of the serpentine curve described by them according to the state of their fits in which they there happen to be, or other circumstances; some outwards, as in the directions *aA, bB, cC, dD*, and others we may suppose inwards, as *aa, bβ, cγ, &c.* With the latter we have here no concern. The former, it is evident, will give rise to as many such caustics as above described, as there are deflected rays; and each caustic, when intercepted on a screen at a distance, will depict on it the maximum of a fringe. The intervals, however, between these caustics, or minima of the fringes, will not be totally black; because the rays from the other caustics, after crossing on the confines of the shadow, or interior fringes, will pursue their course, and partially illuminate all the space beyond. Thus the fringes should be less numerous and the degradation of colour more rapid than in the coloured rings.

715. This theory accounts then perfectly for the curvilinear propagation of the fringes, for their rapid degradation, for their apparently originating in the very edge of the body, (since each caustic will actually come up to that edge, as at A, fig. 142,) and for the remarkable brightness of the fringes, especially the first, which really contains in itself all the light which would have passed into the region BC between the visible and geometrical shadows. It should appear, therefore, that M. Fresnel, in the objections he has taken against these points of the Newtonian doctrine of inflexion in his excellent work *Sur la Diffraction de la Lumière*, (§ 1, p. 15, 17, 19,) must have formed a very inadequate conception of the doctrine he opposes, which, if viewed in the light he has there placed it in, would indeed deserve no other epithet than puerile, and must be looked upon as quite unworthy of its illustrious author; and were these the only difficulties to be explained, we should certainly not be justified

in passing a hasty sentence on it. Other objections advanced by the same eminent philosopher, however, are more serious, and refer to a phenomenon of which the doctrine of deflective forces seems incapable of giving any account; but of which, in justice to Newton we ought to add, it does not appear that he was aware, or its importance could not fail to have struck him.

Phenomenon 3. All other things remaining the same, let the opaque body A be brought nearer the luminous point O, (fig. 142.) The fringes then, formed at the same distance as before behind A, are observed to dilate considerably in breadth,—preserving, however, the same relative distances from each other, and from the border of the shadow. This fact is evidently incompatible with the idea of their being caused by any deflecting force emanating from the opaque body, since it is inconceivable that such a force should depend on the distance the light has travelled from another point no way related to the body.

To explain the diffracted fringes on the undulatory doctrine, Dr. Young conceived the rays passing near the edge of the opaque body to interfere with those reflected very obliquely on its edge, and which in the act of reflexion had lost half an undulation, as in the case of the reflected rings. This supposition would, in fact, lead us to conclude the existence of a series of fringes propagated hyperbolically, and perfectly resembling those really existing. M. Fresnel, however, has shown that a minute though decided difference exists between their places, as given by this theory and by direct measurement; and has, moreover, remarked, that were this the true explanation, they could hardly be supposed absolutely independent of the figure of the edge of the opaque body, which experience shows they are; and that in cases where this edge is extremely sharp, the small quantity of light which *could* be reflected from it would be insufficient to interfere with that passing by it, so as to form fringes so bright as we see them. These objections appear conclusive, especially as the supposition of a reflexion on the edge of the body is unnecessary, since a more strict application of the undulatory doctrine, assisted by the principle of interferences, will be found to afford a full and precise explanation of all the facts, regarding the opaque body as merely an obstacle *bounding* the waves propagated from the luminous point on one side.

To show this, let us consider a wave AMF propagated from O, and of which all that part to the right of A (fig. 145) is intercepted by the opaque body AG; and let us consider a point P in a screen at the distance AB behind A, as illuminated by the undulations emanating simultaneously from every point of the portion AMF, according to the theory laid down in Art. 628, *et seq.* For simplicity, let us consider only the propagation of undulations in one plane. Put AO = a, AB = b, and suppose λ = the length of an undulation; and drawing PN any line from P to a point near M, put PF = f, NM = s, PB = x; then, supposing P very near to B,

and with centre P radius PM describing the circle QM, we shall have $f = PQ + QN = \sqrt{(a+b)^2 + x^2} - a + QN = b + \frac{x^2}{2(a+b)} + QN$. Now, QN is the sum of the versed sines of the arc s to radii OM and PM,

and is therefore equal to $\frac{s^2}{2OM} + \frac{s^2}{2PM} = \frac{s^2}{2} \left(\frac{1}{a} + \frac{1}{b} \right) = \frac{a+b}{2ab} \cdot s^2$; so that, finally,

$$f = b + \frac{x^2}{2(a+b)} + \frac{(a+b)s^2}{2ab}.$$

Now, if we recur to the general expression demonstrated in Art. 632, for the motion propagated to P from any limited portion of a wave, we shall have in this case $a \cdot \phi(\theta) = 1$, because we may regard the obliquity of all the undulations from the whole of the efficacious part of the surface AMN as very trifling, when P is very distant from A in comparison with the length of an undulation. And as we are now only considering undulations propagated in one plane, that expression becomes merely $V = \int ds \cdot \sin 2\pi \left(\frac{t}{T} - \frac{f}{\lambda} \right)$, and the corresponding expression for the excursions of a vibrating molecule at P will be

$$X = \int ds \cdot \cos 2\pi \left(\frac{t}{T} - \frac{f}{\lambda} \right).$$

If then we put for f its value, and take

$$2\pi \left(\frac{t}{T} - \frac{b}{\lambda} - \frac{x^2}{2\lambda(a+b)} \right) = \theta; \quad s \cdot \sqrt{\frac{2(a+b)}{ab\lambda}} = \nu,$$

and consider that in those expressions t and x remain constant, while s only varies, the latter will take the form

$$X = \sqrt{\frac{ab\lambda}{2(a+b)}} \cdot \left\{ \cos \theta \cdot \int d\nu \cdot \cos \left(\frac{\pi}{2} \nu^2 \right) + \sin \theta \cdot \int d\nu \cdot \sin \left(\frac{\pi}{2} \nu^2 \right) \right\},$$

which shows that the total wave on arriving at P may be regarded as the resultant of two waves X' . cos θ and X'' . sin θ, differing in their origin by a quarter-undulation, and whose amplitudes X' and X'' are given by the expression

$$X' = \sqrt{\frac{ab\lambda}{2(a+b)}} \cdot \int d\nu \cos \frac{\pi}{2} \nu^2; \quad X'' = \sqrt{\frac{ab\lambda}{2(a+b)}} \cdot \int d\nu \sin \frac{\pi}{2} \nu^2,$$

the integrals being taken between limits of ν corresponding to s = -AM, and s = +∞. Consequently,

716.
Dilatation of the fringes by the approach of the radiant point.

717.
Dr Young's explanation of the fringes on the undulatory doctrine. Objections against it.

718.
Fresnel's explanation Fig. 145.

Light, since
the limits of ν must be

$$s = AM = PB \times \frac{a}{a+b} = \frac{ax}{a+b}, \text{ and } \nu = s \sqrt{\frac{2(a+b)}{ab\lambda}},$$

$$\nu = -x \sqrt{\frac{2a}{(a+b)b\lambda}} \text{ and } \nu = +\infty.$$

719.
Rule for
determining
the illumina-
tion of
any point in
the screen.

720.
Maxima
and minima
numerically
estimated.

Hence, to determine the intensity of the light at any point P on the screen, we must first of all calculate the values of these integrals; and having thus determined X' and X'' , the square root of the sum of their squares $\sqrt{X'^2 + X''^2}$ will represent the amplitude of a single vibration, the resultant of both, (Art. 615;) and the sum of their squares simply ($X'^2 + X''^2$), the intensity of the light, or the sensation produced in the eye.

M. Fresnel, in the work already cited, has given a table of the values of these integrals for limits successively increasing from 0 up to ∞ , (at which latter limit each is equal to $\frac{1}{2}$, as may readily be proved;) and, calculating on this, he finds that the intensity of the light, without the limit of the geometrical shadow, passes through a series of maxima and minima according to the following table:

Table of the Maxima and Minima for the Exterior Fringes, and of the Corresponding Intensities of the Light illuminating them.

	Values of ν .	Intensities of the light.		Values of ν .	Intensities of the light.
First maximum . . .	1.2172	2.7413	Fourth minimum . .	3.9372	1.7783
First minimum . . .	1.8726	1.5570	Fifth maximum . . .	4.1832	2.2206
Second maximum . .	2.3449	2.3990	Fifth minimum . . .	4.4160	1.8014
Second minimum . .	2.7392	1.6867	Sixth maximum . . .	4.6069	2.1985
Third maximum . .	3.0820	2.3022	Sixth minimum . . .	4.8479	1.8185
Third minimum . . .	3.3913	1.7440	Seventh maximum . .	5.0500	2.1818
Fourth maximum . .	3.6742	2.2523	Seventh minimum . .	5.2442	1.8317

In this it is to be remarked, that no minimum is zero, and that the difference between the successive maxima and minima diminishes very rapidly as the values of ν increase, which explains the rapid degradation of their tints.

721.
Illumina-
tion of the
border of
the geomet-
rical
shadow.

722.
Illumina-
tion within
the shadow.

If the point P be situated on the very edge of the geometrical shadow, its illumination should on this theory be $(\frac{1}{2})^2 + (\frac{1}{2})^2 = \frac{1}{2}$. To compare this with the illumination of the same point, were the opaque body removed, we have only to consider, that at a great distance from the shadow the light must be the same, whether the body be there or not. Now the limit to which the maxima and minima approximate is 2, which therefore represents the uniform illumination beyond the fringes; so that the light on the border of the geometrical shadow is equal to $\frac{1}{4}$ of the full illumination from the radiant point.

Within the shadow we have only to make s or ν negative. This does not alter the values of the integrals, but it does their limits, which must in that case be taken not from $\nu = -x \sqrt{\frac{2a}{(a+b)b\lambda}}$ to $+\infty$, but

from $\nu = +x \sqrt{\frac{2a}{(a+b)b\lambda}}$ to $+\infty$. The computations have been executed by M. Fresnel, who finds

that no periodical increase or decrease here takes place, but that the light degrades rapidly and constantly within the geometrical shadow to total darkness.

723.
Visible
shadow
larger than
the geomet-
rical.

724.

The actual visible shadow then is marked by no sudden defalcation of light, and it will depend on the judgment of the eye where to establish its termination. If we regard all that part as shadow which is less illuminated than the general light of the screen beyond the fringes, then the visible shadow will extend considerably beyond the geometrical one, and this explains why the shadows of small bodies are so much dilated, as we have seen they are.

If we would know the breadths of the several fringes, we have only to find the values of x in the equation

$$x = \nu \cdot \sqrt{\frac{(a+b)b}{a}} \frac{\lambda}{2}$$

where ν has in succession the several values set down in the foregoing table. If we consider the variation of x for successive values of a and b , we shall see the origin both of the curvilinear propagation of the fringes, and of their dilatation on the approach of the luminous point. In fact if we regard, first, the relation between b and x , or the locus of any fringe regarded as a curve, having the line AB for an abscissa and BP as an ordinate, we have $x^2 = \nu^2 \frac{\lambda}{2} \left(b + \frac{b^2}{a} \right)$, which is the equation of an hyperbola having its convexity outwards and passing through A. Secondly, on the other hand, if we regard a as the variable quantity and b as constant, we see that for one and the same distance from the screen, the breadths of the fringes increase as a diminishes; the increments of

Light.

their squares, as the incident rays from being parallel become more divergent, being directly as their divergence. Thirdly, for equal values of λ , a , and b , x is proportional to ν ; so that the breadths of the several fringes are always in the same ratio to each other, and form a progression the same with those of the values of ν in the foregoing table. Lastly, the breadths of the fringes for different coloured rays are as the square roots of the lengths of their undulations.

Part III.

The accordance of this theory with experiment, so far as it regards the distances of the fringes from the shadow and from each other, has been put to a severe test by M. Fresnel, and found perfect. It were to be wished, however, that he had stated somewhat more precisely the instrumental means by which he determined the place of the border of the *geometrical shadow*, from which his measures are all stated to be taken; and which, being marked by no phenomenon of maximum or minimum, might be liable to uncertainty if judged of by the eye alone. This, however, in no way invalidates the accuracy of the final conclusions, as the intervals between the fringes are distinctly marked, and susceptible of exact measurement. The dilatation of the fringes on the approach of the luminous point is, perhaps, the strongest fact in favour of the undulatory doctrine, and in opposition to that of inflection, which has yet been adduced. It seems hardly reconcilable to any received ideas of the action of corpuscular forces, to suppose the force of deflection exerted by the edge of a body on a passing ray, to depend on the distance which the ray has passed over before arriving at that edge from an arbitrarily assumed origin. M. Fresnel has placed this argument in a strong light, in his work already cited.

725.

Besides the exterior fringes above described, there are others formed in certain circumstances within the shadows of bodies which afford peculiarly apt illustrations of the principle of interferences. The first class of phenomena of this kind was noticed by Grimaldi, who found that when a long, narrow body is held in a small diverging beam of light, the shadow received on a screen at a distance will be marked in the direction of its length with alternate streaks or fringes brighter and darker than the rest. These are more or less numerous, according as the distance of the shadow from the body is smaller or greater in proportion to the breadth of the latter. To study the phenomena more minutely, Dr. Young passed a sunbeam through a hole made with a fine needle in thick paper, and brought into the diverging beam a slip of card one-thirtieth of an inch in breadth, and observed its shadow on a white screen at different distances. The shadow was divided by parallel bands, as above described, but *the central line was always white*. That these bands originated in the interference of the light passing on both sides of the card, Dr. Young demonstrated beyond all controversy, by simply intercepting the light on one side by a screen interposed between the card and the shadow, leaving the rays on the other side to pass freely, in the manner represented in fig. 146, where O is the hole, A B the card, E F its shadow, and C D the intercepting body receiving on its margin the margin of the shadow of the edge B of the body. In this arrangement all the fringes which had before existed in the shadow E F immediately disappeared, although the light inflected on the edge A was allowed to retain its course, and must have necessarily undergone any modification it was capable of receiving from the proximity of the other edge B. The same result took place when the intercepting screen was placed as at c d before the edge B of the body, so as to throw its own shadow on the margin B of the card.

726.

Fringes observed by Grimaldi within narrow shadows.

Dr. Young's fundamental interference. Fig. 146.

Without entering minutely into the rationale of this phenomenon, which, however, the formulæ of the preceding articles enable us fully to do, by considering the illumination of any point X between E and F as arising from the whole wave $a A B b$, minus the portion A B, and which M. Fresnel has done at full length, and with great success, in his Memoir already so often cited; we shall content ourselves with showing how fringes or alternations of colour must originate in such circumstances; in fact, if we join A X, B X, the difference of routes of the waves arriving at X by the paths O A X, O B X is equal to B X - A X. It is therefore nothing in the middle of the shadow E F, which ought therefore to be illuminated by double the light deflected into the shadow at that distance by either edge, Art. 722, which will be less as the object is larger, and the shadow broader. But on either side of the middle B X - A X increases; and when it attains a value equal to half an undulation, the waves are in complete discordance, and therefore the middle bright portion will be succeeded by a dark band on either side, and these again by bright ones, and so on.

727.

Explanation.

An elegant variation of this experiment of Dr. Young is afforded by a phenomenon described by Grimaldi. When a shadow is formed by an object having a rectangular termination; besides the usual external fringes there are two or three alternations of colours, beginning from the line which bisects the angle, disposed, within the shadow on each side of it, in curves which are convex towards the bisecting line, and which converge towards it as they become remote from the angular point. These fringes are the joint effect of the light spreading into the shadow from each outline of the object, and interfering as above; and that they are so, is proved by placing a screen within a few inches of the object, so as to receive only one edge of the shadow, when the whole of the fringes disappear. If, on the other hand, the rectangular point of the screen be opposed to the point of the shadow, so as barely to receive the angle of the shadow on its extremity, the fringes will remain undisturbed. (Young, Experiments and Calculations relating to Physical Optics, *Phil. Trans.*, 1803.)

728.

Grimaldi's crested fringes.

Such are some of the more remarkable appearances produced within and beyond the shadows of narrow bodies. Let us next consider the effect of transmitting a beam through a very narrow aperture. And the first case is when the aperture is circular. Suppose, for instance, we place a sheet of lead, having a small pin-hole pierced through it, in the diverging cone of rays from the image of the sun, formed by a lens of short focus, and in the line joining the centres of the hole and focus prolonged place a convex lens or eye-glass, behind which the eye is applied. The image of the hole will be seen through the lens as a brilliant spot, encircled by rings of colours of great vividness, which contract and dilate, and undergo a singular and beautiful alternation of tints, as the distance of the hole from the luminous point on the one hand, or on the eye-glass on the other, is changed. When the latter distance is considerable, the central spot is white, and the rings follow nearly the order of the colours of thin plates. Thus, when the diameter of the hole was about $\frac{1}{30}$ th of an inch, its distance

729.

Case of diffraction through a small circular aperture.

Light. (a) from the luminous point about 6 feet 6 inches, and its distance (b) from the eye-lens 24 inches, the series of colours was observed to be, Part III.

1st order. White; pale yellow; yellow; orange; dull red.

2d order. Violet; blue (broad and pure;) whitish; greenish yellow; fine yellow; orange red, very full and brilliant.

3rd order. Purple; indigo blue; greenish blue; pure, brilliant green; yellow green; red.

4th order. Good green, but rather sombre and bluish; bluish white; red.

5th order. Dull green; faint bluish white; faint red.

6th order. Very faint green; very faint red.

7th order. A trace of green and red.

730.
Table of
colours of
central spot
and sur-
rounding
rings.

When the eye-lens and hole are brought nearer together, the central white spot contracts into a point and vanishes, and the rings gradually close in upon it in succession, so that the centre assumes in succession the most surprisingly vivid and intense hues. Meanwhile the rings surrounding it undergo great and abrupt changes in their tints. The following were the tints observed in an experiment made some years ago, (July 12, 1819,) the distance between the eye-glass and luminous point ($a + b$) remaining constant, and the hole being gradually brought nearer to the former.

$b =$	Central Spot.	Surrounded by
24.00	White	Rings as in the foregoing Article.
18.00	White	{ The two first rings confused, the red of the 3rd and green of the 4th orders splendid.
13.50	Yellow	{ Interior rings much diluted, the 4th and 5th greens and 3rd, 4th and 5th reds the purest colours.
10.00	Very intense orange	All the rings are now much diluted.
9.25	Deep orange red	The rings all very dilute.
9.10	Brilliant blood red	The rings all very dilute.
8.75	Deep crimson red	The rings all very dilute.
8.36	Deep purple	The rings all very dilute.
8.00	Very sombre violet	A broad yellow ring.
7.75	Intense indigo blue	A pale yellow ring.
7.00	Pure deep blue	A rich yellow.
6.63	Sky blue	A ring of orange, from which it is separated by a narrow, sombre space.
6.00	Bluish white	{ Orange red, then a broad space of pale yellow, after which the other rings are scarcely visible.
5.85	Very pale blue	A crimson red ring.
5.50	Greenish white	Purple, beyond which yellow verging to orange.
5.00	Yellow	Blue, orange.
4.75	Orange yellow	Bright blue, orange red, pale yellow, white.
4.50	Scarlet	Pale yellow, violet, pale yellow, white.
4.00	Red	White, indigo, dull orange, white.
3.85	Blue	White, yellow, blue, dull red.
3.50	Dark blue	Orange, light blue, violet, dull orange.

731.
Fresnel's
analysis of
this case.

The series of tints exhibited by the central spot is, evidently, so far as it goes, that of the reflected rings in the colours of thin plates. The surrounding colours are very capricious, and appear subject to no law. They depend, indeed, on very complicated and unmanageable analytical expressions, with which we shall not trouble the reader, but content ourselves with presenting the explanation given by M. Fresnel of the changes of tint of the central spot in white light, and its alternations of light and total darkness observed by him in an homogeneous illumination. Let then a and b be the distances of a small hole whose radius is r from the luminous point, and a screen placed behind the hole perpendicularly to the ray passing directly through its centre. Then if we consider any infinitely narrow annulus of the hole whose radius is z , and breadth dz , this annulus will send to the central point of the screen a system of waves whose intensity is proportional to the area of the annulus, or $2\pi z dz$, but whose phase of undulation differs from that of the central ray, by reason of the difference of the paths described by them. Now, calling f the distance of each point in the annulus from the centre of the screen, we have $f^2 = b^2 + z^2$, and, in like manner, if f' be the distance of the luminous point from the same annulus, $f'^2 = a^2 + z^2$, so that $(f + f') - (a + b)$ the difference of paths, or interval of retardation, is equal to

$$\frac{z^2}{2} \left(\frac{1}{a} + \frac{1}{b} \right) = \frac{z^2 (a + b)}{2ab}.$$

Hence, the general expression in Art. 632 for the amplitude of the total wave, incident on the centre of the screen in this particular case, is equivalent to

$$X = \int 2\pi z dz \cdot \sin 2\pi \left\{ \frac{t}{T} - \frac{z^2 (a + b)}{2ab\lambda} \right\},$$

Light. or, integrating, which from the peculiar form of the differential is in this case easy,

Part III.

$$X = \frac{a b \lambda}{a+b} \left\{ \text{const} + \cos 2 \pi \cdot \left(\frac{t}{T} - \frac{z^2 (a+b)}{2 a b \lambda} \right) \right\}$$

which, extended from $z = 0$ to $z = r$, gives

$$X = \frac{a b \lambda}{a+b} \left\{ \cos \lambda \left(\frac{t}{T} - \frac{(a+b) r^2}{2 a b \lambda} \right) - \cos 2 \pi \frac{t}{T} \right\}$$

$$= \frac{a b \lambda}{a+b} \left\{ \sin \frac{\pi (a+b) r^2}{a b \lambda} \cdot \sin 2 \pi \frac{t}{T} + \left(\cos \frac{\pi (a+b) r^2}{a b \lambda} - 1 \right) \cdot \cos 2 \pi \frac{t}{T} \right\}.$$

This expresses, as we have before remarked in a similar case, (Art. 718,) two partial waves differing by a quarter-undulation, and expressing it, as in that case, by $X = X' \cdot \cos \theta + X'' \cdot \sin \theta$, where $\theta = \frac{t}{T}$, we find for the intensity A^2 of their resultant

$$A^2 = X'^2 + X''^2 = 4 \left(\frac{a b \lambda}{a+b} \right)^2 \left(\sin \frac{\pi (a+b) r^2}{2 a b \lambda} \right)^2.$$

To make use of this, however, we must compare it with what would be the direct illumination of the centre of the screen, if the aperture were infinite, *i. e.* if the direct light from the luminous point shone full upon it. To this case, however, neither our formula nor our reasoning are applicable; for if we make r infinite in this expression, it becomes illusory, and presents no satisfactory sense, and in our reasoning we have neglected to consider the law of diminution of the intensity of the oblique waves, or regarded $\phi(\theta)$ in Art. 631 as invariable, which in this extreme case is far from the truth. We must, therefore, have recourse to another method. Now, M. Fresnel has demonstrated (and our limits oblige us to take his demonstration for granted) that this total illumination is equal to one-fourth of that which the centre of the screen would receive from an opening of such a radius, that the difference of routes of a ray passing through the centre, and one diffracted at the circum-

732.
Illumination of the central spot compared with the total illumination. Fresnel's theorem.

ference, shall be an exact semi-undulation, *i. e.* in which $\frac{r^2 (a+b)}{2 a b} = \frac{\lambda}{2}$, or $r = \sqrt{\frac{a b \lambda}{a+b}}$. If then we substitute this for r in the above formula, and put C for the whole illumination, we get, on the same scale,

$$C = \left(\frac{a b \lambda}{a+b} \right)^2 \left(\sin \frac{\pi}{2} \right)^2 = \left(\frac{a b \lambda}{a+b} \right)^2$$

and, consequently,

$$A^2 = 4 C \left(\sin \frac{\pi (a+b) r^2}{2 a b \lambda} \right)^2.$$

In this expression r, a, b are independent of λ , and therefore the value of A^2 is of the form $4 C \left(\sin 2 \pi \cdot \frac{\beta}{\lambda} \right)^2$ where $B = \frac{(a+b) r^2}{4 a b}$. Hence, if we suppose light of all colours to emanate from the luminous point, the compound tint produced in the central point of the screen will be represented by $S \left\{ 4 C \cdot \left(\sin 2 \pi \frac{\beta}{\lambda} \right)^2 \right\}$ and will therefore, by Art. 673, be the same with that reflected by a plate of air whose thickness is B , or $\frac{(a+b) r^2}{4 a b}$ which increases as b diminishes when $a+b$ remains constant. Thus we see the origin of the succession of colours of the central spot in the Table above recorded, which is the more satisfactory, as that experiment was made without reference to, and indeed in ignorance of, this elegant application of M. Fresnel's general principles, the merit of which is due (as he himself states) to M. Poisson.*

733.
The colours those of the reflected rings.

Another very curious result of M. Poisson's researches is this, that the centre of the shadow of a very small circular opaque disc, exposed to light diverging from a single point, is precisely as much illuminated by the diffracted waves as it would be by the direct light, if the disc were altogether removed. We cannot spare room for the demonstration of this singular theorem. It has been put to the test of experiment by M. Arago, with a small metallic disc cemented on a very clear and homogeneous plate of glass, and with full success.

734.
Poisson's theorem for the illumination in the centre of a small circular shadow.

When the light is transmitted through two equal apertures, placed very near each other, the rings are formed about each as in the case of one; but besides these arise a set of narrower, straight, parallel fringes bisecting the interval between their centres, and at right angles to the line joining them. If the apertures be unequal, these fringes assume the form of hyperbolas, having the aperture in their common focus. Besides these also two other sets of parallel rectilinear fringes (in the case of equal apertures) go off in the form of a St. Andrew's cross from the centre at equal angles with the first set. See figures 147, 148. When the apertures are more numerous or varied in shape, the variety and beauty of the phenomena are extraordinary; but of this more presently.

735.
Case of diffraction through two apertures very near each other. Fig. 147 and 148

M. Fresnel has shown, that when the light from a single luminous point is received on two plane mirrors

* The coincidence in the higher orders of colours was, however, in our experiments less complete, and especially the green of the third order, which was wanting altogether in some cases.

Light.

very slightly inclined to each other, so as to form two almost contiguous images, if these be viewed with a lens, there will be seen between them a set of fringes perpendicular to the line joining them. These are evidently analogous to those produced by the two holes in the experiments last described. The experiment is delicate; for if the surfaces of the reflectors at the point where they meet be ever so little, the one raised above or depressed below the other, so as to render the difference of routes of the rays greater than a very few undulations, no fringes will be seen. But it is valuable, as demonstrating distinctly that the borders of the apertures in the preceding experiment have nothing to do with the production of the fringes, the rays being in this case abandoned entirely to their mutual action after quitting the luminous point. An exactly similar set of fringes is formed if, instead of two reflectors, we use a glass, plane on one side, and on the other composed of two planes, forming a very obtuse angle, as in fig. 149. This being interposed between the eye-lens E and the luminous point S, forms two images S and S' of it; and the interference of the rays SE and S'E from these images, forms the fringes in question.

Fig. 149.

737.

Effect of interposing a denser medium in one of two interfering rays.

Since the production of the fringes and their places with respect to the images of the luminous point, depends on the difference of routes of the interfering rays, it is evident, that if, without altering their paths, we alter the *velocity* of one of them with respect to the other, during the whole or a part of its course, we shall produce the same effect. Now, the velocity of a ray may be changed by changing the medium in which it moves. In the undulatory system, the velocity of a ray in a rarer medium is greater than in a denser. Hence, if in the path of one of two interfering rays we interpose a parallel plate of a transparent medium denser than air, (at right angles to the ray's course,) we shall increase its interval of retardation, or produce the same effect as if its course had been prolonged. If then a *thick* plate of a dense medium, such as glass, be interposed in one of the rays which form visible fringes, they will disappear; because the interval of retardation will be thus rendered suddenly equal to a great number of undulations, whereas the production of the fringes requires that the difference of routes shall be very small. If, however, only a very thin lamina be interposed, they will remain visible, but shift their places. Thus, in fig. 150, let SA, SB be rays transmitted through the small apertures A, B from the luminous point S, and received on the screen DCE, these forming a set of fringes of which C, the middle one, will be *white*. Let D, E be the dark fringes immediately adjacent on either side; and things being thus disposed, let a thin film of glass or mica G be interposed in *one* of the rays SA, its thickness being such that the ray in traversing it shall just be retarded half an undulation. Then will the rays AE, BE, which before were in complete discordance, be now in exact accordance, and there will be formed at E a bright fringe instead of a dark one. On the other hand, the ray AC will now be half an undulation behind BC, instead of in complete accordance with it, so that at C there will be formed a dark fringe, and so on. In other words, the whole system of fringes will be formed as before, but will have shifted its place, so as to have its middle in E instead of in C, *i. e.* will have moved *from* the side on which the plate of the dense medium is interposed. It is evident, that if the plate G be thicker, the same effect will take place in a greater degree.

Fig. 150.

Displacement of the fringes by it explained.

738.

Mode of putting this to the test of experiment.

To make the experiment, however, it must be considered that the refractive power of glass, or indeed of any but gaseous media, is so great, that any plate of manageable thickness would suffice to displace the fringes so far as to throw them wholly out of sight. But we shall succeed, if, instead of a single plate G placed over *one* aperture A, we place two plates G, g of very nearly equal thicknesses, (such as will arise from two nearly contiguous fragments of one and the same polished plate,) one over each aperture; or we may vary the thickness of the plate traversed by either ray by inclining it, so as to bring it within the requisite limits. This done, the effect observed is precisely that described; the fringes shift their places *from* the thicker plate, without sustaining any alteration in other respects. This elegant experiment affords a strong indirect argument in favour of the undulatory system, and in opposition to that of emission, since it proves that the rays of light are *retarded* in their passage through denser media, agreeably to what the undulatory system requires, and contrary to the conclusions of the corpuscular doctrine.

Argument against the corpuscular system.

739.

Arago and Fresnel's method of determining refractions of gases.

MM. Arago and Fresnel have taken advantage of this property, to measure the relative refractive powers of different gases, or of the same in different states of temperature, pressure, humidity, &c. It is manifest, that if any considerable portion of the path of *one* of the interfering rays be made to pass through a tube closed at both ends with glass plates, and the other through equal glass plates *only*, the fringes will be formed as usual. But if the tube be *exhausted*, or warmed, or cooled, or filled with a gas of different refractive density, a displacement of the fringes will take place, which (if they be received in the focus of a micrometer) may be measured with the greatest delicacy. Knowing the amount of their displacement, as compared with the breadth of the fringes, we know the number of undulations gained or lost by one ray on the other; and hence, knowing the internal length of the tube, we have the ratio of the refracting power of the medium it contains to that of air. What renders this method remarkable is, that *there is actually no conceivable limit to the precision of which it is susceptible*, since tubes of any length may be employed, and micrometers of any delicacy.

740.

Fraunhofer's experiments on diffraction and interference. His apparatus.

The phenomena of diffraction, and those arising from the mutual interference of several very minute pencils of rays emanating from a common origin, have been investigated by M. Fraunhofer with great care and extraordinary precision, by the aid of a very delicate apparatus devised and executed by himself.

This apparatus consisted of a repeating, 12-inch theodolite, reading to every 4", carrying, attached to its horizontal circle, a plane circular disc of six inches in diameter, having its axis precisely coincident with that of the theodolite, and having its own particular divisions independent of those of the theodolite. In the centre of this disc was placed vertically a metallic screen, having in it one or more narrow, vertical, rectangular slits, or other apertures, and so fixed as to have the middle of its aperture, or system of apertures, exactly coincident with the axis of the instrument. Attached to the great circle of the theodolite, horizontally, was a telescope, having its object-glass three inches and a half from the centre, and its axis directed exactly to it, and precisely parallel to the plane of the limb, and provided with a delicate micrometer, whose parallel threads were exactly vertical.

Light.

The instrument being insulated on a support of stone, a beam of solar light was directed by a heliostat, through a very narrow slit, also exactly vertical, having a breadth of one hundredth of an inch, and distant $463\frac{1}{2}$ inches from the centre of the theodolite, so as to fall on the screen, and, being transmitted through its apertures, to be received into the telescope. It is manifest that the eye-glass of the telescope will here view the fringes, &c. as they are formed in its focus. The magnifying power of the telescope used by Fraunhofer varied from 30 to 50 times.

Part III.

M. Fraunhofer first examined the effect produced by the diffraction of the light through a single slit,—the breadth of which he first determined with the greatest precision by means of a micrometer-microscope, with which he assures us that he found it practicable to appreciate so minute a quantity as $1\text{--}50,000$ th of an inch. The slit being then placed on the apparatus, and accurately adjusted before the object-glass of the telescope, which was directed exactly to the aperture in the heliostat, the image of the latter was formed in its focus, accompanied by lateral fringes, which by the effect of the magnifying power were dilated into broad and brilliant prismatic spectra. The distances of the red ends of these spectra from the middle point, or white central image, were then measured accurately by means of the micrometer. The result of a great number of experiments with apertures of all breadths from one-tenth to one-thousandth of an inch, agreed to astonishing precision with each other, and with the following laws, viz. that (under the circumstances of the experiment,)

741.
Fringes produced by a single narrow aperture.

1. *The angles of deviation of the diffracted rays, forming similar points of the systems of fringes produced by different apertures, are inversely as the breadths of the apertures.*

Their laws and dimensions.

2. *That the distances of similar rays (the extreme red, for instance,) from the middle in the several spectra, constituting the successive fringes, form in each case an arithmetical progression whose difference is equal to its first term.*

3. That calling γ the breadth of the aperture, in fractions of a Paris inch, the angular distances L' , L'' , L''' , &c. in parts of a circular arc to radius unity, of the extreme red rays in each fringe from the middle line, are respectively represented by $L' = \frac{L}{\gamma}$, $L'' = 2 \cdot \frac{L}{\gamma}$, $L''' = 3 \cdot \frac{L}{\gamma}$, &c. where $L = 0.0000211$, and a similar law

holds for all the other coloured rays, different values being assigned to L for each.

This conclusion agrees perfectly with the result of an experiment related by Newton in the IIIId Book of his *Optics*. He ground two knife edges truly straight, and placed them opposite to each other, so as to be in contact at one end, and at the other to be at a small distance, such that the angle included between them was about $1^\circ 54'$, thus forming a slit whose breadth at their intersection was evanescent, and at 4 inches from that point $\frac{1}{4}$ th of an inch, and in the intermediate points, of course, of every intermediate magnitude. Exposing this in a sunbeam emanating from a very small hole at 15 feet distance, he received their shadows on a white screen behind them, and observed that when they were received very near to the knife edges, (as at half an inch,) the fringes exterior to the shadow of each edge ran parallel to its border without sensible dilatation, till they met and joined without crossing, at angles equal to that contained between the knife edges. But when the shadows were received at a great distance from the knives, the fringes had the form of hyperbolas, having for one asymptote the shadow of the knife to which they respectively belonged, and for the other a line perpendicular to that bisecting the angle of the two shadows, each fringe becoming broader and more distinct from the shadow which it bordered, as it approached the angle. These hyperbolas crossed without interfering, as represented in fig. 151. Their points of crossing, Newton found, however, not to be at a constant distance from the angle included between the projections of the knife edges, but to vary in position with the distance from the knives, at which the shadow is received on the screen; and hence, he says, "I gather that the light which makes the fringes upon the paper, is not the same light at all distances of the paper from the knives; but when the paper is held very near the knives, the fringes are made by light which passes by their edges at a less distance, and is more bent than when the paper is held at a greater distance from the knives." Newton, however, left these curious researches, which could hardly have failed to have led in his hands to a complete knowledge of the principles of diffraction—unfinished; being, as he says, interrupted in, and unwilling to resume them: doubtless, owing to the chagrin and opposition his optical discoveries produced to him. An unmeet reward, it must be allowed, for so noble a work, but one of which, unhappily, the history of Science affords but too many parallels.

742.
Newton's experiment with two knife edges.

Fig. 151

The above were the results obtained by M. Fraunhofer when the two edges of the aperture were both in a plane perpendicular to the incident rays; but when the same effective breadth was procured, by inclining a larger aperture obliquely, so as to reduce its actual breadth in the ratio of the cosine of its incidence to radius, or by limiting the incident ray by two opaque edges at different distances from the object-glass of the telescope, the phenomena were very different. To accomplish this, two metallic plates were fixed upright on the horizontal plate of the theodolite, having their edges exactly vertical, and precisely at opposite extremities of a diameter. Then, by turning the plate round on its axis, the passage allowed to the light between them could be increased or diminished at pleasure. The phenomena, then, were as follows. When the opening allowed to the light was considerable, as 0.02 or 0.04 inch (Paris,) the fringes were exactly similar to those observed when the edges were equidistant from the object-glass; but as the opening diminished, they ceased to be symmetrical on both sides of the middle line, those on the side of that edge of the aperture nearest to the telescope becoming broader than those on the other, which, on their part, undergo no sensible alteration. As the aperture contracts, this inequality increases, till at length the dilated fringes begin to disappear in succession, the outermost first, which they do by suddenly acquiring an extraordinary magnitude, so as to fill the whole field of the telescope, and thus, as it were, losing themselves. While these are thus vanishing, those on the other side remain quite unaltered till the last is gone, when they all disappear at once, which happens at the moment that the opening is reduced to nothing by the two edges covering each other.

743
Case when the edges of the aperture were at different distances from the origin of the light.

Light. When the aperture placed before the object-glass, instead of being a straight line, was a small, circular hole, and the aperture of the heliostat, in like manner, a minute circle, the phenomena of the rings were observed, and their diameters could be accurately measured by the micrometer. The results of these measurements led M. Fraunhofer to the following laws: 1st, that for apertures of different diameters, the diameters of the rings are inversely as those of the apertures forming them; 2dly, that the distances from the centre of the maxima of extreme red rays (or of rays of any given refrangibility) in the several rings of one and the same system, form an arithmetical progression, whose difference is somewhat less than its first term. Thus, calling γ the diameter

of the aperture, and putting $L = \frac{0.0000214}{\gamma}$ and $l = \frac{0.0000257}{\gamma}$, he found $L' = l$, $L'' = l + L$, $L''' = l + 2L$,

&c., where L' L'' , &c. represent the angular semidiameters of the several rings expressed in arc of a circle to radius unity. The near coincidence of the value of L in this case, with that in the case of a linear aperture, and the small, but decided difference of the values of the first term of the progression in the two cases, are very remarkable.

745. When the aperture was a very narrow, circular annulus, such as might be traced with a steel point on a gilt disc of glass, of whatever diameter, the image was a circular spot, surrounded in like manner by coloured rings, the diameters of which depended nowise on the *diameter*, but only on the *breadth* of the annulus, being in fact (as might be expected) the very same as the intervals between similar opposite fringes, on both sides of the central line in the image produced by a linear aperture of equal breadth.

746. But the most curious parts of M. Fraunhofer's investigations are those which relate to the interference of rays transmitted through a great many narrow apertures at once. When these apertures are exactly equal, and placed at exactly equal distances from one another, phenomena of a totally different kind from those originating in a single aperture are seen. In his first experiments of this kind he formed a *grating* of wire, by stretching a very fine wire across a frame, in the form of a narrow, rectangular parallelogram, whose shorter sides were screws tapped in the same die, and therefore precisely similar; across these screws in the consecutive intervals between their threads the wires were stretched, and of course could not be otherwise than parallel and equidistant. The diameter of the wire was 0.002021 Paris inch, the intervals between them each 0.003862, and the grating consisted of 260 such wires. When this apparatus was placed precisely vertical before the object-glass of his telescope, and illuminated by a narrow line of light 0.01 inch in breadth, also exactly vertical, forming the aperture of the heliostat, the image of this was seen in the telescope, colourless, well defined, and in all respects precisely as it would have been seen without the interposition of any grate or aperture at all, occupying the centre of the field, only less bright. On either side of this was a space perfectly dark, after which succeeded a series of prismatic spectra, which he calls spectra of the second class, not consisting of tints melting into each other, according to the law of the coloured rings, or any similar succession of hues depending on a regular degradation of light, but of perfectly *homogeneous* colours; so much so, as to exhibit the same dark lines crossing them as exist in the purest and best defined prismatic spectrum. In the disposition of things already described, the first, or nearer spectrum is completely insulated, the space between it and the central image, as well as between it and the second spectrum, being quite dark. The violet ends of the spectra are inwards, and the red outwards; but the violet end of the third spectrum is superposed on the red end of the second, so as in place of a dark interval to produce a purple space; and as we proceed farther from the middle, the spectra become more and more confounded, but not less than thirteen may easily be counted on each side by the aid of a prism refracting them transversely, so as to separate their overlapping portions.

747. The measurement of the distances of similar points in the several spectra are rendered susceptible of the utmost precision by means of the dark lines which cross them. A very remarkable peculiarity of these spectra must, however, be here noticed, *viz.* that although the dark lines hold exactly the same places in the order of colours, or, in other words, correspond to precisely the same *degrees of refrangibility*, as in the prismatic spectra formed by refraction, yet the ratio of the intervals between them, or the breadths of the several coloured spaces, differ entirely in the two cases. Thus, in the diffracted spectra, the interval between the lines C and D (fig. 94) is very nearly double of that between G and H, while in a spectrum formed by a flint-glass prism of an angle of 270, the proportion is reversed, and in a water prism of the same angle $CD : GH :: 2 : 3$.

748. In the diffracted fringes formed by a single aperture, their distances (as we have seen) from the axis depends only on the breadth of the aperture, being inversely as that breadth. In the spectra formed by a great number, their distances from the central image depends neither on the breadths of the apertures nor on the intervals between them, but on the sum of these quantities, that is, on the distances between the middle points of the consecutive apertures, (or, in the case before us, on the distances between the axes of the wires.) By a series of measures performed with the utmost care and precision on wire gratings of a great variety of dimensions, M. Fraunhofer ascertained the following laws and numerical values.

749. 1. For different gratings, if we call γ the breadth of each of the interstices through which the light passes, and δ that of each of the opaque intervals between them, the magnitudes of spectra of the same order, and the distances of similar points in them from the axis, is inversely as the sum $\gamma + \delta$.

750. 2. The distances of similar points, (*i. e.* of similar colours or similar fixed lines,) in the several consecutive spectra formed by one and the same grating from the axis, constitute an arithmetical progression whose difference is equal to its first term.

751. 3. For the several refrangibilities corresponding to the fixed lines B, C, D, E, &c. the first term of this progression is numerically represented by the respective fractions which follow, being the lengths of the arcs, or their sines to radius unity.

Light.

Part III.

$$\begin{aligned} B &= \frac{0.00002541}{\gamma + \delta}; & E &= \frac{0.00001945}{\gamma + \delta}; & H &= \frac{0.00001464}{\gamma + \delta}. \\ C &= \frac{0.00002422}{\gamma + \delta}; & F &= \frac{0.00001794}{\gamma + \delta}; & & \&c. \\ D &= \frac{0.00002175}{\gamma + \delta}; & G &= \frac{0.00001587}{\gamma + \delta}; \end{aligned}$$

These results were all, however, deduced from gratings so coarse as to allow of our regarding the angles of diffraction as proportional to their sines; but when extremely fine gratings are employed, the spectra are formed at great distances from the axis, and the analogy of other similar cases, as well as theory, would lead us to substitute $\sin B$, $\sin C$, $\sin D$, &c. in the place of B , C , D , &c. This, M. Fraunhofer found by experiment to be really the case. The construction of gratings proper for these delicate purposes, however, was no easy matter. Those employed by him were nothing more than a system of parallel and equidistant lines ruled on plates of glass covered with gold-leaf, or with the thinnest possible film of grease; by the former of these methods he found, that the proximity of the lines might be carried to the extent of placing about a thousand in the inch, but when he would draw them still closer, the whole of the gold-leaf was scraped off. When the surface was covered with a film of grease so thin as to be almost imperceptible to the sight, (although the intervals were in this case transparent,) no change was produced in the optical phenomena, so far as the spectra were concerned, only the brightness of the central image being increased. By this means he was enabled to obtain a system of parallel lines at not more than half the distance from each other that could be produced on gold-leaf: but beyond this degree of proximity, he found it impossible to carry the ruling of equidistant lines on any film of grease or varnish. But this being still far short of his wishes, he had recourse to actual engraving with a diamond point on the surface of the glass itself, and by this means was enabled to rule lines so fine as to be absolutely invisible under the most powerful compound microscope, and so close that 30,000 of them lie in a single Paris inch. When so excessively near, however, no accuracy of machinery will ensure that perfect equidistance which is essential to the production of the spectra now under consideration, and he found it impossible to succeed in placing them nearer than 0.0001223, (or about 8200 to the inch,) with such a degree of precision as to enable him to distinguish the fixed lines in the spectra; and, if it be considered, that a deviation to the extent of the hundredth part of the just interval frequently occurring, is sufficient to obliterate these, and that to produce the spectra in sufficient brightness to affect the eye, some hundreds or even thousands must be ruled, we shall be enabled to form some conception of the difficulties to be encountered in researches of this kind. For a detail of some of these, and of the methods employed by him to count their number and measure their distances, we must refer to his original Memoir, (read to the Royal Bavarian Academy of Sciences, June 14, 1823.)

In the course of these researches, M. Fraunhofer met with a very singular and instructive peculiarity in one of the engraved glass-gratings used by him; which, although it produced spectra equidistant on either side of the axis, yet gave always those on one side a much greater degree of brightness than those on the other. Attributing this to the *form* of the furrows being sharper terminated on one side than on the other, owing either to the figure of the diamond point or the manner of its application, he endeavoured to produce a similar structure of the striæ in a film of grease spread on glass, by purposely applying the engraving tool obliquely, and the attempt proved successful.

When the incident rays from the opening in the heliostat fell obliquely on the grating, it might be supposed that the phenomena would be the same as those exhibited by a closer grating, having intervals less in proportion of the cosine of the angle of incidence to radius. But the analogy of the unsymmetrical fringes produced by a single aperture, whose sides lie in a plane oblique to the incident ray, may lead us to expect a different result, and experiment confirms the surmise; thus, M. Fraunhofer found, that on inclining a grating, whose intervals $(\gamma + \delta)$ were each equal to 0.0001223 inch, so as to make the angle of incidence 55° with the perpendicular, the distance of the first fixed line D from the axis on the one side of the axis was $15^\circ 6'$, and on the other no less than $30^\circ 33'$, or more than double.

The facts deduced by M. Fraunhofer in the above detailed researches are certainly extremely curious. The most interesting and remarkable point about them is the perfect homogeneity of colour in the spectra, indicating a saltus, or breach of continuity, in the law of intensity of each particular coloured ray in the diffracted beam. For it is obvious, that taking any one refrangibility (that corresponding to the fixed line C, for example,) the expression of its intensity in functions of its distance from the axis must be (analytically speaking) of such a nature as to vanish completely for every value of that distance, excepting for a certain series in arithmetical progression, or, as it is called, a *discontinuous* function; so that the curve representing such value, having the distance from the axis for its abscissa, must be a series of points arranged above the axis at equal intervals; or, at least, a curve of the figure represented in fig. 151, in which certain extremely narrow portions, equidistantly arranged, start up to considerable distances from the axis, while all the intermediate portions lie so close to that line as to be confounded with it. The manner in which such a function can be supposed to originate from the

summation of a series of the values of $\int d\nu \cdot \sin \frac{\pi}{2} \nu^2$ and $\int d\nu \cdot \cos \frac{\pi}{2} \nu^2$, (Art. 718,) taken successively be-

tween limits corresponding to the boundaries of the several interstices, involves too many complicated considerations to enter into in this place. M. Fraunhofer, meanwhile, states the following general expression, as the result of his own investigations founded on the principle of interferences. Let n indicate the order of any

752.

Case of extremely close gratings.

Methods of constructing them.

753.

The spectra modified by the form of the striæ in the gratings.

754.

Case of inclined gratings. Unsymmetrical spectra of the second class.

755.

Theoretical considerations.

Light.
Fraun-
hofer's
formula.

spectrum, reckoned from the axis; ϵ the distance from the middle of one interstices to that of the adjacent one $= \gamma + \delta$; λ the length of an undulation of an homogeneous ray; σ the angle of incidence of the ray from the luminous point on the grating; and y the length of a perpendicular let fall from the micrometer thread of the telescope, (or from the point in the focus of its object-glass, where that particular homogeneous ray in that spectrum is found,) on the plane of the grating. Then, if the angular elongation of that ray from the axis be called $\theta^{(n)}$, we shall have, in general,

$$\cotan \theta^{(n)} = \frac{\sqrt{\{\epsilon^2 - (\epsilon \cdot \sin \sigma + n\lambda)^2\} \cdot \{4y^2 + \epsilon^2 - (\epsilon \cdot \sin \sigma + n\lambda)^2\}}}{2y(\epsilon \cdot \sin \sigma + n\lambda)}.$$

In this equation, n is to be regarded as $+$ for the spectra which lie on the side of the axis on which the incident ray makes an obtuse angle with the plane of the grating, and negative for the spectra on the other side. This formula he states to be rigorous, and independent of any approximation. When y is very great (as it, in fact, always is,) compared with ϵ and λ , this reduces itself simply to

$$\cotan \theta^{(n)} = \frac{\sqrt{\epsilon^2 - (\epsilon \cdot \sin \sigma + n\lambda)^2}}{\epsilon \cdot \sin \sigma + n\lambda}, \text{ or } \sin \theta^{(n)} = \frac{\epsilon \cdot \sin \sigma + n\lambda}{\epsilon}.$$

756.
Lengths of
undulations
of the rays
B, C, D, &c.
assigned by
Fraunhofer.

This formula, applied to M. Fraunhofer's measures of the distances of the same fixed lines in successive spectra on either side of the axis, in the case of inclined gratings, represents them with perfect exactness. When the gratings are perpendicular to the ray $\sigma = 0$, and the equation becomes $\sin \theta^{(n)} = \frac{n\lambda}{\epsilon}$, which is the law before noticed for symmetrical spectra. And hence, too, it appears that the values of λ , or the lengths of the undulations for the several rays designated by C, D, E, &c., are no other than the numerators of the fractions in Art. 751, expressed in parts of a Paris inch, which thus become data of the utmost value in the theory of light, from the great care and precision with which they have been fixed, and for the possibility of identifying them at all times.

757.
Diffracted
spectra pro-
duced by
reflexion.

If the unruled surface of the glass grating be covered with black varnish, and the light reflected from the ruled surface be received in the telescope, the very same phenomena are seen as if the light had been transmitted through the glass, and the same analytical expression, according to M. Fraunhofer, applies to both cases.

758.
Alleged
limit to the
powers of micro-
scopes.

A curious consequence of this expression is, that if ϵ , the distance between the lines, be less than λ , and the light fall perpendicularly on the grating, so that $\sin \sigma = 0$, we shall have $\sin \theta^{(n)} > 1$, and therefore $\theta^{(n)}$ imaginary. It appears, therefore, that lines drawn on a surface distant from each other by a less quantity than one undulation of a ray of light, produce no coloured spectra. Hence, such scratches, or inequalities, on polished surfaces, have no effect in disturbing the regularity of reflexion or refraction, and produce no dimness or mistiness in the image; if less distant from each other than this limit. M. Fraunhofer seems inclined to conclude further, that an object of less linear magnitude than λ can in consequence never be discerned by microscopes, as consisting of parts: a conclusion which would put a natural limit to the magnifying power of microscopes, but which we cannot regard as following from the premises.

759.
Spectra
produced
by compo-
site gra-
tings.

When the intervals of the parallel interstices are unequal, and disposed with no regularity, the light of the diffracted spectra of different combinations is confounded together, and a white misty streak at right angles to the direction of the lines arises; but when they are *regularly unequal*, so that the same intervals recur in regular periods, if we call $E (= \epsilon' + \epsilon'' + \epsilon''' + \&c.)$ the interval between any two distant by a whole period, we shall have, for the law of the lateral spectra, the equation $\sin \theta^{(n)} = \frac{n\lambda}{E}$. And the spectra so formed, are

Singular
phenome-
non noticed
by Fraun-
hofer
respecting
the inten-
sity of the
spectra.

still observed to consist of homogeneous light, exhibiting the fixed lines with great distinctness. A very curious, and, as far as concerns the practical measurement of the phenomena, useful observation has been made by M. Fraunhofer on the spectra so formed by these *composite* gratings, *viz.* that although they follow the same law in respect of their distances from the axis, yet the successive spectra differ greatly in intensity, some being so faint as to be scarce perceptible, while the immediately adjacent ones will often be very intense. Owing to this cause, spectra of the higher orders, which in a simple grating the interval of whose interstices is represented by E , are confused and obliterated by the encroachment of those adjacent, are often very distinct when formed by a *composite* grating, the period of recurrence of whose similar interstices is $E = \epsilon' + \epsilon'' + \epsilon''' + \&c.$ Thus, M. Fraunhofer was never able through a simple grating to see the fixed lines C and F in the spectrum of the 12th order, reckoning from the axis, while in a composite grating, consisting of three systems of lines continually repeated, whose intervals ϵ' , ϵ'' , ϵ''' were to each other as 25 : 33 : 42, these fixed lines as well as the lines D and E, were distinctly seen in the 12th spectrum, owing to the almost total disappearance of the 10th and 11th. Nay, even the fixed line E in the 24th spectrum could be seen, and its distance from the axis measured with this grating.

760.
Various
stages of
the pheno-
mena.
Spectra of
the first class

Such are the extreme cases of the phenomena as produced by a single aperture, and by an infinite, or, at least, very great number; but the intermediate steps and gradations by which one set of phenomena pass into the other, remain to be traced. When a single interstice is left open in a grating, the spectra are formed as described in Art. 741. These, M. Fraunhofer calls *spectra of the first class*, and their colours are not homogeneous, but graduate into one another.

When two contiguous interstices are left open, the spectra of the first class appear as before; but between the axis and the first spectrum on either side appear other spectra, which M. Fraunhofer terms *imperfect spectra of the second class*, their colours being similar to those of the first class, and no fixed lines being visible in them.

761

Light. When three adjacent interstices are left open, a third set of spectra, or *spectra of the third class*, are formed Part III. between the axis and the nearest of the imperfect spectra of the second class. Besides these, no new *classes* of Spectra of the third class. spectra arise by a further increase of the number of interstices; but these undergo a series of modifications as 762. the interstices grow more numerous. These are chiefly as follows: Modifications of these spectra by increasing the number of interfering rays.

The spectra of the third class grow narrower, and approach the axis, till at last they run together and form by their union the colourless, well-defined image of the opening of the heliostat in the axis of the whole phenomenon. By a series of exact measurements, M. Fraunhofer found their breadths to be inversely as the number of interstices by which they are produced in the same grating, and inversely as the intervals of the interstices for different ones; and in general, that $\gamma + \delta = \epsilon$ representing this interval, m the number of interstices used, and n the order of the spectrum, $\theta^{(n)}$ the distance of extremity of the red rays in that spectrum is given by the equation

$$\theta^{(n)} = \frac{n}{m} \times \frac{0.0000208}{\epsilon}.$$

As the spectra of the third class contract into the axis, they leave a dark space between it and the first spectrum of the second class. This and the other spectra of that class meanwhile grow continually more vivid and homogeneous in respect of colour; till at length, when the number of interfering rays is very much increased, the fixed lines begin to appear in them, and they acquire the character of *perfect* spectra of the second class.

M. Fraunhofer next examined the phenomena produced by immersing in media of different refractive powers the gratings used, when he found all the phenomena precisely similar; but the distances at which the several spectra were formed from the axis, to be less than when in air, in the inverse ratio of the refractive indices.

A very beautiful and splendid class of optical phenomena has been investigated and described by M. Fraunhofer, which arise by substituting for the gratings used in the above experiments very small apertures of regular figures, such as circles and squares, either singly or arranged in regular forms, in great numbers; as, for instance, when two equal wire gratings are crossed at right angles. Fig. 151 is a representation of the phenomenon produced when the light is received on the object-glass of the telescope through two circular holes of the diameter 0.02227 inch, placed at a distance of 0.03831 inch centre from centre. Each compartment is a separate spectrum. In the bands $a a$, $b b$ we see here plainly the origin and minute structure of the vertical and crossed fringes described in Art. 735. The appearances vary as the number of apertures is increased, the spectra growing purer and more vivid. That which arises when two equal wire gratings are crossed, is figured in M. Fraunhofer's work, and is one of the most magnificent phenomena in Optics.

When we look at a bright star through a very good telescope with a low magnifying power, its appearance is that of a condensed, brilliant mass of light, of which it is impossible to discern the shape for the brightness; and which, let the goodness of the telescope be what it will, is seldom free from some small ragged appendages or rays. But when we apply a magnifying power from 200 to 300 or 400, the star is then seen (in favourable circumstances of tranquil atmosphere, uniform temperature, &c.) as a perfectly round, well-defined planetary disc, surrounded by two, three, or more alternately dark and bright rings, which, if examined attentively, are seen to be slightly coloured at their borders. They succeed each other nearly at equal intervals round the central disc, and are usually much better seen and more regularly and perfectly formed in refracting than in reflecting telescopes. The central disc, too, is much larger in the former than in the latter description of telescope.

These discs were first noticed by Sir William Herschel, who first applied sufficiently high magnifying powers to telescopes to render them visible. They are not the real bodies of the stars, which are infinitely too remote to be ever visible with any magnifiers we can apply; but *spurious*, or unreal images, resulting from optical causes, which are still to a certain degree obscure. It is evident, indeed, to any one who has entered into what we have said of the law of interferences, and from the explanation given in Art. 590 and 591 of the formation of *foei* on the undulatory system, that (supposing the mirror or object-glass rigorously aplanatic) the focal point in the axis will be agitated with the united undulations, in complete accordance, from every part of the surface, and must, of course, appear intensely luminous; but that as we recede from the focus in any direction in a plane at right angles to the axis, this accordance will no longer take place, but the rays from one side of the object-glass will begin to interfere with and destroy those from the other, so that at a certain distance the opposition will be total, and a dark ring will arise, which, for the same reason, will be succeeded by a bright one, and so on. Thus the origin both of the central disc and the rings is obvious, though to calculate their magnitude from the data may be difficult. But this gives no account of one of the most remarkable peculiarities in this phenomenon, *viz.* that the apparent size of the disc is different for different stars, being uniformly larger the brighter the star. This cannot be a mere illusion of judgment; because when two unequally bright stars are seen at once, as in the case of a close double star, so as to be directly compared, the inequality of their spurious diameters is striking; nor can it be owing to any real difference in the stars, as the intervention of a cloud, which reduces their brightness, reduces also their apparent discs till they become mere points. Nor can it be attributed to irradiation, or propagation of the impression from the point on the retina to a distance, as in that case the light of the central disc would encroach on the rings, and obliterate them; unless, indeed, we suppose the vibrations of the retina to be performed according to the same laws as those of the ether, and to be capable of interfering with them; in which case, the disc and rings seen on the retina will be a resultant system, originating from the interference of both species of undulations.

Not to enter further, however, on this very delicate question, we shall content ourselves with stating some of the phenomena we have observed, as produced by diaphragms, or apertures of various shapes variously applied to mirrors and object-glasses, and which form no inapt supplement to the curious observations of Fraunhofer on the effect of *very minute* apertures, of which they are in some sort the converse.

Formula for spectra of third class.

763. Transition from imperfect to perfect spectra of second class.

764. Phenomena of gratings immersed in fluids.

765. Substitution of very minute apertures for gratings.

766. Rings seen about the stars in telescopes.

767. Spurious discs of the stars.

Explanation of the rings on the principle of interferences.

768. Phenomena produced by apertures of various figures.

Light. When the whole aperture of a telescope is limited by a *circular* diaphragm, whether applied near to, or at a distance from, the mirror or object-glass, the disc and rings enlarge in the inverse proportion of the diameter of the aperture. When the aperture was much reduced (as to one inch, for a telescope of 7 feet focal length) the spurious disc was enlarged to a planetary appearance, being well defined, and surrounded by one ring only, strong enough to be clearly perceived, and faintly tinged with colour in the following order, reckoning from the centre of the disc. White, very faint red, black, very faint blue, white, extremely faint red, black. When the aperture was reduced still farther (as to half an inch) the rings were too faint to be seen, and the disc was enlarged to a great size, the graduation of light from its centre to the circumference being now very visible, giving it a hazy and cometic appearance, as in fig. 152.

Fig. 152.

770. When annular apertures were used the phenomena were extremely striking, and of great regularity. The exterior diameter of the annulus being three inches, and the interior $1\frac{1}{2}$, the appearance of Capella was as in fig. 153, and of the double star Castor, as in 154. As the breadth of the annulus is diminished, the size of the disc and breadth of the rings diminish also, (contrary to what took place in Fraunhofer's experiments with extremely narrow annuli, and obviously referring the present phenomena to different principles,) at the same time the number of visible rings increases. Fig. 155, 156, and 157 exhibit the appearance of Capella with annular apertures of 5.5 inch — 5 inch (*i. e.* whose exterior diameter = 5.5 and interior = 5) of 0.7 — 0.5, of 2.2 — 2.0. In the last case the disc was reduced to a hardly perceptible round point, and the rings were so close and numerous as scarcely to admit being counted, giving, on an inattentive view, the impression of a mere circular blot of light. When the breadth of the annulus was reduced to half this quantity, the intervals between the rings could no longer be discerned. The dimensions of the rings and disc, generally, seem to be proportional to $\frac{r' - r}{r}$.

Fig. 153 to 157.

771. Besides the rings immediately close to the central disc, however, others of much greater diameter and fainter light, like halos, are seen with annular apertures, which belong (in Fraunhofer's language) to spectra of a different class. With a single annulus they are too faint to be distinctly examined, but with an aperture composed of two annuli, as in fig. 158, they are very distinct and striking, presenting the phenomenon in fig. 159, (in which it is to be understood that light is represented in the engraving by darkness, and darkness by light.)

Another set of rings. Fig. 158 and 159.

772. When the aperture was in the form of an equilateral triangle, the phenomenon was extremely beautiful; it consisted of a perfectly regular, brilliant, six-rayed star, surrounding a well-defined circular disc of great brightness. The rays do not unite to the disc, but are separated from it by a black ring. They are very narrow, and perfectly straight; and appear particularly distinct in consequence of the *total destruction of all the diffused light* which fills the field when no diaphragm is used; a remarkable effect, and much more than in the mere proportion of the light stopped. Fig. 160 is a representation of this elegant appearance. The same arises when, in place of an equilateral triangle, the aperture is the difference of two concentric, equilateral triangles similarly situated.

Image produced by a triangular aperture.

Fig. 160.

773. As a triangle has but three sides and three angles, it seems singular that a *six-rayed* star should be produced. Supposing three to arise from the angles, and three from the sides, it might be expected that some sensible difference should exist in the alternate rays, marking their different origin. When the telescope is in perfect focus, however, all the rays are precisely alike; but if thrown out of focus, their difference of origin becomes apparent. Fig. 161 represents the phenomenon then seen, in which the alternate branches are seen to consist of a series of fringes parallel to their length, and the others of small arcs of similar fringes immediately adjacent to the vertices of the hyperbolas to which they belong, and which consequently cross the rays in a direction perpendicular to their length. As the telescope is brought better in focus, the hyperbolas approach their asymptotes, and are confounded together in undistinguishable proximity; and thus three rays arise composed of continuous lines of light, and three intermediate ones composed of an infinite number of discontinuous points placed infinitely near each other. To represent analytically the intensity of the light in one of these *discontinuous* rays would call for the use of functions of a very singular nature and delicate management.

When out of focus.

Fig. 161.

774. The phenomenon just described affords in certain cases a very perfect position-micrometer for astronomical uses. If the diaphragm be turned round, the rays turn with it; and if a brilliant star (as α Aquilæ) have near it a very small one, the diaphragm may be so placed as to make one of the rays pass through the small star, which thus remains like a bead threaded on a string, and may be examined at leisure. If then the position of the diaphragm be read off on a graduation properly contrived, the relative situations of the two stars become known. We have satisfied ourselves by trial of the practicability of this; and by proper contrivances the principle may be made available in cases which at first sight appear to present considerable difficulties.

Application to the construction of a position-micrometer.

775. When three circular apertures, having their centres at the angles of an equilateral triangle, were used, the image consisted of a bright central disc. Six fainter ones in contact with it, and a system of very faint halo-like rings surrounding the whole as in fig. 162. When, however, three equal and similar *annular* apertures were thus disposed, the appearance when *in focus* was as in fig. 153, being exactly the same as if two of them were closed. But when thrown a little out of focus, the difference was perceived. Fig. 163 represents the appearance in this case, each of the apertures then produces its own central disc and system of rings, whose intersections give rise to the system of intersectional fringes there depicted. As the telescope is brought better in focus these disappear, and the phenomenon is as in fig. 164; the centres gradually approaching, and the rings blending till the point of complete coincidence is attained.

Three circular apertures. Fig. 162.

Fig. 163.

Fig. 164.

776.

An aperture in the form of the difference between two concentric squares produced not an eight, but a four rayed star. The rays, however, were not, as in the case of the triangular aperture, uninterrupted fine lines, gradually tapering away from the centre to their extremities, but composed of distinct alternating obscure and

Lght. bright portions, as represented in fig. 165. The portions nearest the central disc (which is circular) were composed of bands transverse to the direction of the rays, and tinged with prismatic colour. Similar bands, no doubt, existed in the more distant portions, which extended to a great length. *Part III. Square apertures. Fig. 165.*

An aperture consisting of fifty squares, each of about half an inch in the side, regularly disposed at intervals so as to leave spaces between them in both directions equal in breadth to the side of each, produced an image totally different from that described by Fraunhofer as resulting from the crossing of two equal very close gratings, though the distribution and shape of the apertures were the same in both cases. It was as represented in fig. 166, consisting of a white, round, central disc, surrounded by eight vivid spectra, disposed in the circumference of a square, beyond which were arranged in the shape of a cross, triple lines of very faint spectra extending to a great distance. *Effect of very numerous square apertures Fig. 166.*

When the aperture consisted of numerous equilateral triangles regularly disposed, as in fig. 167, the image presented the very beautiful phenomenon represented in fig. 168, consisting of a series of circular discs arranged in six diverging rays from the central one, and each surrounded with a ring. The central disc was colourless and bright; the rest more and more strongly coloured and elongated into spectra, according to their degree of remoteness from the centre. These are only a few of the curious and beautiful phenomena depending on the figures of the apertures of telescopes, which afford a wide field of further inquiry, and one at least as interesting to the artist as to the philosopher. *778. Fig. 167.*

PART IV.

OF THE AFFECTIONS OF POLARIZED LIGHT.

§ I. *Of Double Refraction.*

779. WHEN a ray of light is incident on the surface of a transparent medium, a portion of it is reflected, at an angle equal to that of incidence, another small portion (so small, however, that we shall neglect its consideration) is dispersed in all directions, serving to render the surface visible, and the rest enters the medium and is refracted. The *law* of refraction, or the rule which regulates the path of this portion within the medium, has been explained in the preceding parts; and no exceptions to it, as a general law, have hitherto been noticed. It is, however, very far from general; and, in fact, obtains only where the refracting medium belongs to one or other of the following classes, *viz.*

Exceptions to the law of ordinary refraction numerous.

Classes of bodies in which it holds.

Class 1. Gases and vapours.

2. Fluids.

3. Bodies solidified from the fluid state too suddenly to allow of the regular crystalline arrangement of their particles, such as glass, jellies, &c., gums, resins, &c., being chiefly such as in the act of cooling pass through the viscous state.

4. Crystallized bodies, having the cube, the regular octohedron, or the rhomboidal dodecahedron for their primitive form, or which belong to the *tessular* system of Mohs. A very few exceptions (probably only apparent ones, arising from our imperfect knowledge of crystallography) exist to the generality of this class.

The solid bodies belonging to these classes, moreover, cease to belong to them when forcibly compressed or dilated, either by mechanical violence, or by the unequal action of heat or cold, which brings their particles into a state of strain, such as in extreme cases to produce their disruption, as is familiarly seen in the cracking of a piece of glass by heat too suddenly and partially applied. The class of fluids too admits some exceptions, at least when very minutely considered; but the deviation from the ordinary law of refraction in these cases is of so microscopic a kind, that we shall at present neglect to regard it.

780. All other bodies, comprehending all crystallized media, such as salts, gems, and crystallized minerals, not belonging to the system above mentioned; all animal and vegetable bodies in which there is any disposition to a regular arrangement of molecules, such as horn, mother of pearl, quill, &c.; and, in general, all solids when in a state of unequal compression or dilatation, act on the intromitted light according to very different laws, dividing the refracted portion into two distinct pencils, each of which pursues a rectilinear course so long as it continues within the medium, according to its own peculiar laws, but without further subdivision. This phenomenon is termed double refraction. It is best and most familiarly seen in the mineral termed Iceland spar, which is, in fact, carbonate of lime in a regular crystalline form. This is generally obtained in oblique parallel-pipeds, easily reduced by cleavage to regular, obtuse rhomboids, and is not uncommonly met with in a state of limpid transparency, on which account, as well as by reason of its remarkable optical properties, it easily attracted attention. Bartholinus, in 1669, appears to have been the first to give any account of its double refraction, which was afterwards more minutely examined by Huygens, the first proposer of the undulatory theory of light, whose researches on this phenomenon form an epoch in the history of Physical Optics little if at all less important than the great discovery of the different refrangibility of the coloured rays by Newton. To Huygens we owe the discovery of the law of double refraction in this species of medium. Newton, misled by some inaccurate measurements, (a thing most unusual with him,) proposed a different one; but the conclusions of Huygens, long and unaccountably lost sight of, were at length established by unequivocal experiments by Dr. Wollaston, since which time a new impulse has been given to this department of Optics; and the successive labours of Laplace, Malus, Brewster, Biot, Arago, and Fresnel present a picture of emulous and successful research, than which nothing prouder has adorned the annals of physical science since the developement of the true system of the universe. To enter, however, into the history of these discoveries, or to assign the share of honour which each illustrious labourer has reaped in this ample field forms no part of our plan. Of the splendid constellation of great names just enumerated, we admire the living and revere the dead far too warmly and too deeply to suffer us to sit in judgment on their respective claims to priority in this or that particular discovery; to balance the mathematical skill of one against the experimental dexterity of another, or the philosophical acumen of a third. So long as "one star differs from another in glory,"—so long as there shall exist varieties, or even incompatibilities of excellence,—so long will the admiration of mankind be found sufficient for all who truly merit it. Waving, then, all reference to the history of the subject, except in the way of incidental remark, or where the necessity of the case renders it unavoidable, we shall present the reader with as

Light. systematic an account as we are able, of the present state of knowledge with respect to the laws and theory of Part IV.
Double Refraction. The Huygenian law having been demonstrated to apply rigorously to the case for which he himself devised it, as well as to a very large class of other bodies, we shall begin with that class, and proceed afterwards to consider more complicated cases.

In all crystallized bodies, then, which possess double refraction, it is found that that portion of a ray of 781.
ordinary light incident on any natural or artificially polished surface which enters the body is separated into *two* Axes of
equal pencils which pursue rectilinear paths, making with each other an angle not of constant magnitude, but double
varying according to the position which the incident ray holds with respect to the surface, and to certain fixed refraction.
lines, or axes within the crystal, and which lines are related in an invariable manner to the planes of cleavage, or other fixed planes or lines in the primitive form of the crystal. Now, it is found that in every crystal there is at least one such fixed line, along which if one of these two pencils be transmitted the other is so also, so that in this case the two pencils coincide, the angle between them vanishing. Moreover, no crystal has yet been discovered in which more than two such lines exist. These lines are called the optic axes. All double refracting crystals, then, at present, may be divided into such as have one, and such as have two, optic axes.

When a ray penetrates the surface of a crystal so as to be transmitted undivided along the optic axis; 782.
or when, moving within the crystal along that line, it meets the surface and passes out, whatever be the Rays
inclination of the surface, its refraction is always performed according to the ordinary law of the propor- moving
tional sines. Thus, in this particular case, the crystal acts precisely as an uncrystallized medium, (some rare along the
instances excepted, of which more hereafter.) axes suffer
ordinary refraction only.

But in all other cases the law is essentially different, and (for one portion of the divided pencil, at least) 783.
of a very singular and complicated nature. This we shall first proceed to explain in the simpler case of What is
crystals with one optic axis. But, first, we must explain somewhat more distinctly, what we mean by meant by
axes and fixed lines within a crystal. Suppose a mass of brickwork, or masonry, of great magnitude, built of axes and
bricks, all laid parallel to each other. Its exterior form may be what we please; a cube, a pyramid, or any other fixed lines
figure. We may cut it (when hardened into a compact mass) into any shape, a sphere, a cone, or cylinder, &c.; within a
but the edges of the bricks within it lie still parallel to each other; and their directions, as well as those of crystal.
the diagonals of their surfaces, or of their solid figures, may all be regarded as so many axes, *i. e.* lines having (so long as the mass remains at rest) a determinate position, or rather *direction* in space, no way related to the exterior surfaces, or linear boundaries of the mass, which may cut across the edges of the bricks in any angles we please. Whenever, then, we speak of fixed lines, or axes *of*, or *within*, a crystal, we always mean directions in space parallel to each of a system of lines drawn in the several elementary molecules of the crystal, according to given geometrical laws, and related in a given manner to the sides and angles of the molecules themselves. We must conceive the axis, then, of a crystallized mass not as a single line having a given *place*, but as any line whatever having a given *direction* in space, *i. e.* parallel to the *axis of each molecule*, which is a line having a determinate place and position within it.

In the remainder of this section, when we speak of the *axis* or *axes* of a crystallized mass or surface generally, 784.
we mean the direction of the optic axis or axes of its molecules, or of a crystal similar and similarly situated to any one of them.

Of the Law of Double Refraction in Crystals with One Optic Axis.

This class of crystals comprises all such as belong to Mohs's rhombohedral system, or which have the acute or 785.
obtuse rhomboid, or regular six-sided prism, for their primitive form, as well as all which belong to his Enumeration
pyramidal system, or whose primitive form is either the octohedron with a square base, the right prism with a of crystals
square base, or the bi-pyramidal dodecahedron. All such crystals Dr. Brewster has shown to have but one having
axis, which is that to which the primitive form is symmetrical, *viz.* in the rhomboid, the axis of the figure, or a single
line joining the two angles formed by three equal plane angles; in the hexagonal prism, the geometrical axis axis in
of the prism; in the octohedron, or square based prism, a line drawn through the centre of the base at right classes.
angles to it. The cases in accordance with the rule are so numerous, and the exceptions, once believed to be so, have so often disappeared on the attainment of a more perfect knowledge of the crystalline forms of the excepted minerals, that when any case of disagreement seems to occur, we are justified in attributing it rather to our own incorrect determination of this datum, than to want of generality in the rule itself.

In all crystals of this class, one of the two equal pencils into which the refracted ray is divided follows the 786.
ordinary law of Snellius and Descartes, having a constant index of refraction (μ), or invariable ratio of the sine Refraction
of incidence to that of refraction, whatever be the inclination of the surface by which it enters; so that its of the ordi-
velocity within the medium, when once entered, is the same in whatever direction it traverses the molecules; nary ray in
and with respect to this ray the crystal comport itself as an uncrystallized medium. This, then, is called the this class of
crystals.
ordinary pencil.

To understand the law obeyed by the other, or *extraordinary* portion of the divided pencil, let us consider 787.
it as fairly immersed in the medium, and pursuing its course among the molecules. Then its velocity will not, Huygens's
as in the case of the ordinary ray, be the same in whatever direction it traverses them, but will depend on the law for the
angle it makes with the axis; being a minimum when its path within the crystal is parallel to the axis, and a velocity of
maximum when at right angles to it, or *vice versa*; and in all intermediate inclinations of an intermediate the extra-
magnitude according to the following law. Let an ellipsoid of revolution, either oblate or prolate, as the case ordinary
ray.

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may be, be conceived, having its axis of revolution coincident in direction with the axis of the crystal, and its polar to its equatorial radius in the ratio of the minimum and maximum velocities above mentioned, *i. e.* as the velocity of a ray moving parallel to that of one perpendicular to the axis. Then in all intermediate positions, the radius of this spheroid parallel to the ray will represent its velocity on the same scale that its polar and equatorial radii represent the velocities in their respective directions.

788.

Its connection with the law of extraordinary refraction.

This is the Huygenian law of velocities, in its most simple and general form. It does not at first sight appear what this has to do with the law of extraordinary refraction; but the reader who has considered with the requisite attention what has been said in Art. 539, 540, with prospective reference to this very case, will easily perceive that, the law of velocity of the ray within the medium once established, it becomes a mere matter of pure Geometry to deduce from it the law of extraordinary refraction, whether we adopt the Corpuscular theory, and employ Laplace's principle of least action, as in that Article; or whether, preferring the Undulatory hypothesis, we substitute for this principle the equivalent one of swiftest propagation, as explained in Art. 587, 588. We should observe, however, that the Huygenian law, as just stated, is worded in conformity with the undulatory doctrine, in which the velocity in a denser medium is supposed *slower* than in a rarer. But when we use the principle of least action, we must invert the use of the word, or, which comes to the same thing, suppose the velocity in the medium to be *inversely* proportional to the radius of the ellipsoid. The results being necessarily the same in both cases, we shall use at present the language of the Corpuscular system.

789.

Investigation of the latter from the former law

Retaining, then, the notation of Art. 540, the law of refraction will be derived from the equation $V \cdot S + V' \cdot S' = a$ minimum, where V is the velocity *without*, and V' that *within* the medium, and where S and S' are the spaces described without and within it, in the passage of a ray from point to point. Let a and b be the polar and equatorial semiaxes of the ellipsoid above spoken of, (which we shall call the ellipsoid of double refraction,) and let α, β, γ be the coordinates of the point (A) without the crystal, and α', β', γ' those of one (B) within it, through which the ray is supposed to pass, and x, y, z the coordinates of a point in the surface of the crystal, on which it must be incident, so as to be capable of passing from A to B in the manner required by the law of extraordinary refraction; and let ϕ be the angle which the interior portion S' makes with the axis of the crystal. Then will the radius of the spheroid parallel to this portion (by conic sections) be expressed by

Expression for the radius of the spheroid of refraction.

$$r = \frac{ab}{\sqrt{b^2 \sin^2 \phi + a^2 \cos^2 \phi}} = \frac{ab}{W}; \quad (1)$$

where a is the equatorial, and b the polar radius of the spheroid. Now, if we take μ to represent the index of ordinary refraction, since we have, generally, $V' = \frac{\text{const}}{r}$, and since, when $r = b$ the extraordinary and ordinary rays coincide, and therefore $V' = \mu V$, consequently we must have $\mu V = \frac{\text{const}}{b}$, and $\text{const} = b \mu V$, so that we shall get

$$V' = \mu V \cdot \frac{b}{r}, \text{ and } \frac{V'}{V} = \mu \cdot \frac{b}{r}.$$

790.

Introduction of the principle of least action or swiftest propagation. Fig. 169.

In general, as we have already seen, the condition of least action affords the equation

$$d\{VS + V'S'\} = 0, \text{ or } V \cdot dS + V' \cdot dS' + S' \cdot dV' = 0; \quad (2)$$

But to make use of this, we must express V' , S , and S' , in terms of variable quantities relating to a point any how taken in the surface of the crystal. Whether this point be expressed by rectangular or polar coordinates is no matter: it will be more convenient, however, to use polar. Let, then, C (fig. 169) be the point of incidence of the ray AC on the surface H a O b, and about C as a centre describe a sphere. Let Z C z be the perpendicular to the surface at C, and let P C p be the position of the axis of the crystal. The plane Z P H z p O Z perpendicular to the surface, and passing through the axis, is called the principal section of the surface. Let Z A a, z B b be vertical planes, containing the incident and refracted rays, and join B p by the arc of a great circle. Then it is evident, that this arc will be equal to ϕ .

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Suppose, now, the axis of the x to be parallel to H C the projection of the axis of the crystal, and since we may choose the plane of the x, y , as we please, let it coincide with the refracting surface, so that $z = 0$. Then dropping the perpendiculars A M, M m, B N, N n, and putting $\lambda = Z P = z p =$ angle between the axis and perpendiculars.

$\omega = O a =$ inclination of the plane of incidence to the principal section.

$\omega' = O b =$ inclination of plane of refraction to ditto.

$\theta =$ angle Z C A = Z A = angle of incidence

$\theta' = z C B = z B =$ angle of refraction.

We shall have as follows:

$$AC = S; AM = \gamma; C m^2 = (a - x)^2; M m^2 = (\beta - y)^2;$$

consequently,

$$\left. \begin{aligned} a - x &= \gamma \cdot \tan \theta \cdot \cos \omega; \quad \beta - y = \gamma \cdot \tan \theta \cdot \sin \omega; \quad S = \frac{\gamma}{\cos \theta} \\ \text{and, similarly,} \\ a' - x &= \gamma' \cdot \tan \theta' \cdot \cos \omega'; \quad \beta' - y = \gamma' \cdot \tan \theta' \cdot \sin \omega'; \quad S' = \frac{\gamma'}{\cos \theta'} \end{aligned} \right\} \quad (3)$$

Light. Now, differentiating these equations, and considering that $d(a - x) = d(a' - x)$ and $d(\beta - y) = d(\beta' - y)$ Part IV.
we get

$$\begin{aligned} d(\tan \theta \cdot \cos \varpi) &= \frac{\gamma'}{\gamma} \cdot d(\tan \theta' \cdot \cos \varpi'); \\ d(\tan \theta \cdot \sin \varpi) &= \frac{\gamma'}{\gamma} \cdot d(\tan \theta' \cdot \sin \varpi'); \end{aligned}$$

which equations, developed and reduced, afford the following,

$$\left. \begin{aligned} \frac{d\theta}{d\theta'} &= \frac{\gamma'}{\gamma} \cdot \left(\frac{\cos \theta}{\cos \theta'} \right)^2 \cdot \cos(\varpi - \varpi'); & \frac{d\theta}{d\varpi'} &= \frac{\gamma'}{\gamma} \cdot \cos \theta^2 \cdot \tan \theta' \cdot \sin(\varpi - \varpi'); \\ \frac{d\varpi}{d\theta'} &= \frac{\gamma'}{\gamma} \cdot \frac{\sin(\varpi' - \varpi)}{\tan \theta \cdot \cos \theta'^2}; & \frac{d\varpi}{d\varpi'} &= \frac{\gamma'}{\gamma} \cdot \frac{\tan \theta'}{\tan \theta} \cos(\varpi' - \varpi); \end{aligned} \right\} \quad (4)$$

which are necessary conditions, in order that the point C may remain on the surface.

But since S, S', V' may be regarded as functions of θ' and ϖ' , which are the polar coordinates we propose to use as independent variables, we shall have 792.

$$dV' = \frac{dV'}{d\theta'} d\theta' + \frac{dV'}{d\varpi'} d\varpi';$$

and, moreover,

$$dS = \frac{\gamma \cdot \sin \theta}{\cos \theta^2} \left(\frac{d\theta}{d\theta'} d\theta' + \frac{d\theta}{d\varpi'} d\varpi' \right) dS' = \frac{\gamma' \cdot \sin \theta'}{\cos \theta'^2} d\theta',$$

so that, substituting their values in the equation (2,) we get

$$\begin{aligned} 0 &= \left\{ V \cdot \frac{\gamma \cdot \sin \theta}{\cos \theta^2} \cdot \frac{d\theta}{d\theta'} + V' \cdot \frac{\gamma' \cdot \sin \theta'}{\cos \theta'^2} + \frac{\gamma'}{\cos \theta'} \cdot \frac{dV'}{d\theta'} \right\} d\theta' \\ &+ \left\{ V \cdot \frac{\gamma \cdot \sin \theta}{\cos \theta^2} \cdot \frac{d\theta}{d\varpi'} + \frac{\gamma'}{\cos \theta'} \cdot \frac{dV'}{d\varpi'} \right\} d\varpi'; \end{aligned} \quad (5)$$

in which the coefficients of each of the two independent differentials being separately made to vanish, we get

$$\left. \begin{aligned} \frac{dV'}{d\theta'} &= -V \cdot \frac{\gamma}{\gamma'} \cdot \frac{\sin \theta \cdot \cos \theta'}{\cos \theta^2} \cdot \frac{d\theta}{d\theta'} - V' \cdot \tan \theta' \\ \frac{dV'}{d\varpi'} &= -V \cdot \frac{\gamma}{\gamma'} \cdot \frac{\sin \theta \cdot \cos \theta'}{\cos \theta^2} \cdot \frac{d\theta}{d\varpi'} \end{aligned} \right\} \quad (6)$$

In these, substituting the values of $\frac{d\theta}{d\theta'}$ and $\frac{d\theta}{d\varpi'}$ found in equation (4,) we obtain the following

$$\left. \begin{aligned} \frac{dV'}{d\theta'} &= -V \cdot \frac{\sin \theta}{\cos \theta'} \cdot \cos(\varpi - \varpi') - V' \cdot \tan \theta' \\ \frac{dV'}{d\varpi'} &= -V \cdot \sin \theta \cdot \sin \theta' \cdot \sin(\varpi - \varpi') \end{aligned} \right\}; \quad (7)$$

These are the very same equations with those deduced by Laplace and Malus, by a more abstruse and complicated calculus, from the primary dynamical relations of the problem, and from them it is easy to express, in general, the law of refraction corresponding to any given law of velocities, for we have only to put them under the form

$$V \cdot \sin \theta \cdot \cos \varpi \cdot \cos \varpi' + V \cdot \sin \theta \cdot \sin \varpi \cdot \sin \varpi' = -V' \cdot \sin \theta' - \cos \theta' \cdot \frac{dV'}{d\theta'}$$

$$V \cdot \sin \theta \cdot \cos \varpi \cdot \sin \varpi' - V \cdot \sin \theta \cdot \sin \varpi \cdot \cos \varpi' = \frac{1}{\sin \theta'} \frac{dV'}{d\varpi'};$$

and multiplying the first by $\cos \varpi'$, and the second by $\sin \varpi'$, and adding, we get

$$V \cdot \sin \theta \cdot \cos \varpi = \frac{\sin \varpi'}{\sin \theta'} \cdot \frac{dV'}{d\varpi'} - \cos \theta' \cdot \cos \varpi' \cdot \frac{dV'}{d\theta'} - \sin \theta' \cdot \cos \varpi' \cdot V; \quad (b)$$

and, again, multiplying the first by $\sin \varpi'$, and the second by $-\cos \varpi'$, and adding, we find

$$V \cdot \sin \theta \cdot \sin \varpi = -\frac{\cos \varpi'}{\sin \theta'} \frac{dV'}{d\varpi'} - \cos \theta' \cdot \sin \varpi' \cdot \frac{dV'}{d\theta'} - \sin \theta' \cdot \sin \varpi' \cdot V'; \quad (9)$$

Light. Now, the second members of these equations, (when V' the velocity of the extraordinary ray is any function of ϕ Part IV.
the angle it makes with the axis, or of its position within the crystal,) is always explicitly given in terms of θ'
and ϖ , so that, calling P and Q their values so expressed, we have at once

$$\tan \varpi = \frac{Q}{P}; \quad \cos \varpi = \frac{P}{\sqrt{P^2 + Q^2}}; \quad \sin \theta = \sqrt{P^2 + Q^2};$$

so that ϖ and θ are directly expressed in terms of ϖ' and θ' ; and, therefore, the direction in which a ray, moving anyhow within the crystal will emerge, is known, and *vice versa*.

793. It only remains to execute these processes in the case before us. To this end (for simplicity) we shall put $V = 1$, and suppose (since a and b , the semiaxes of the spheroid, are arbitrary) $b = \frac{1}{\mu}$, or $\mu = \frac{1}{b}$, and put W for the radical $\sqrt{a^2 \cdot \cos \phi^2 + b^2 \cdot \sin \phi^2}$, when we shall have

$$V' = \frac{W}{a b}; \quad d V' = \frac{a^2 - b^2}{a b} \cdot \frac{\cos \phi}{W} \cdot d(\cos \phi).$$

Now in the spherical triangle $Z B p$ we have, the side $Z p = \lambda$; $Z B = \theta'$, angle $p Z B = \varpi'$, and side $p B = \phi$, therefore, by spherical trigonometry,

$$\cos \phi = \cos \lambda \cdot \cos \theta' + \sin \lambda \cdot \sin \theta' \cdot \cos \varpi', \quad (10)$$

and differentiating separately with respect to θ' and ϖ' ,

$$\frac{d \cdot \cos \phi}{d \theta'} = -\cos \lambda \cdot \sin \theta' + \sin \lambda \cdot \cos \theta' \cdot \cos \varpi'$$

$$\frac{d \cdot \cos \phi}{d \varpi'} = -\sin \lambda \cdot \sin \theta' \cdot \sin \varpi'.$$

If, then, we write these values in the partial differences of V' in the equations (8) and (9,) they will become

$$\sin \theta \cdot \cos \varpi = -\frac{1}{a b W} \left\{ W^2 \cdot \sin \theta' \cdot \cos \varpi' + (a^2 - b^2) \cos \phi [\sin \lambda (1 - \cos \varpi'^2 \cdot \sin \theta'^2) - \cos \lambda \cdot \sin \theta' \cdot \cos \theta' \cdot \cos \varpi'] \right\}$$

$$\sin \theta \cdot \sin \varpi = -\frac{1}{a b W} \left\{ W^2 \cdot \sin \theta' \cdot \sin \varpi' - (a^2 - b^2) \cos \phi [\sin \lambda \cdot \sin \varpi' \cdot \cos \varpi' \cdot \sin \theta'^2 + \cos \lambda \cdot \sin \theta' \cdot \cos \theta' \cdot \sin \varpi'] \right\}.$$

In these, let $b^2 + (a^2 - b^2) \cos \phi^2$ be put for W^2 , and, bearing in mind that the value of $\cos \phi$ is as given in the equation (10,) we shall see that they will reduce themselves respectively to

$$\sin \theta \cdot \cos \varpi = -\frac{1}{a b W} \left\{ b^2 \cdot \sin \theta' \cdot \cos \varpi' + (a^2 - b^2) \cdot \sin \lambda \cdot \cos \phi \right\}$$

that is, by reason of (10,)

$$\left. \begin{aligned} -\sin \theta \cdot \cos \varpi &= \frac{(a^2 - b^2) \cdot \cos \lambda \cdot \sin \lambda \cdot \cos \theta' + (a^2 \cdot \sin \lambda^2 + b^2 \cdot \cos \lambda^2) \cdot \cos \varpi' \cdot \sin \theta'}{a b W} \\ \text{and} \\ -\sin \theta \cdot \sin \varpi &= \frac{b^2 \cdot \sin \theta' \cdot \sin \varpi'}{a b W} \end{aligned} \right\} \quad (11)$$

794. These equations, conjointly with the equations expressing the value of W in terms of $\cos \phi$, and of $\cos \phi$ in terms of θ' and ϖ' , afford a complete solution of the problem in the case when a ray passes out of a crystal into air, and suffice to determine both the inclination of the refracted ray to the surface, and the inclination of the plane in which it lies to the principal section.

For brevity, let us put

$$a^2 \cdot \sin \lambda^2 + b^2 \cdot \cos \lambda^2 = A; \quad a^2 \cdot \cos \lambda^2 + b^2 \cdot \sin \lambda^2 = B; \quad (a^2 - b^2) \cdot \sin \lambda \cdot \cos \lambda = C; \quad (12)$$

and, dividing the second of the equations (11) by the first, we find

$$\tan \varpi = \frac{b^2 \cdot \tan \theta' \cdot \sin \varpi'}{A \cdot \tan \theta' \cdot \cos \varpi' + C} \quad (13)$$

which gives immediately the inclination of the plane of emergence to the principal section, or, as it is sometimes termed, the *azimuth* of the emergent ray.

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Reciprocally, if having given the angle of incidence and azimuth of a ray incident externally on the crystal, we would find the angle of refraction and azimuth of the intromitted ray, we must find θ' and ϖ' from the above equations in terms of θ and ϖ . This may be thus accomplished :

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Given the path of the ray without, required that within the crystal.

Take $x = \tan \theta' \cdot \cos \varpi'$, and $y = \tan \theta' \cdot \sin \varpi'$,

then $x^2 + y^2 = \tan^2 \theta'$, and $\cos \theta' = \frac{1}{1 + x^2 + y^2}$;

and, moreover, $\tan \varpi = \frac{b^2 y}{A x + C}$;

now, since $W^2 = b^2 + (a^2 - b^2) \cdot \cos \phi^2$,

$$= \cos \theta' \left\{ \frac{b^2}{\cos \theta'^2} + (a^2 - b^2) (\cos \lambda + \sin \lambda \cdot \tan \theta' \cdot \cos \varpi')^2 \right\}$$

the second of the equations (11) becomes, by squaring,

$$(\sin \theta \cdot \sin \varpi)^2 \left\{ \frac{b^2}{\cos \theta'^2} + (a^2 - b^2) \cdot (\cos \lambda + \sin \lambda \cdot \tan \theta' \cdot \cos \varpi')^2 \right\} = \frac{b^2}{a^2} (\tan \theta' \cdot \sin \varpi')^2,$$

that is,

$$a^2 (\sin \theta \cdot \sin \varpi)^2 \{ b^2 (1 + x^2 + y^2) + (a^2 - b^2) (\cos \lambda + x \cdot \sin \lambda)^2 \} = b^2 y^2,$$

that is, developing

$$a^2 \cdot (\sin \theta \cdot \sin \varpi)^2 \{ A x^2 + 2 C x + B + b^2 y^2 \} = b^2 y^2.$$

Now we have

$$A x + C = \frac{b^2 y}{\tan \varpi}, \text{ and } x = \frac{b^2 y}{A \cdot \tan \varpi} - \frac{C}{A}.$$

And, on substitution, this equation will be found to take the form $p y^2 + q = 0$, and being resolved to give

$$y = \tan \theta' \cdot \sin \varpi' = \frac{a^2 \cdot \sin \theta \cdot \sin \varpi}{\sqrt{A - a^2 \cdot \sin^2 \theta} (A \cdot \sin \varpi^2 + b^2 \cdot \cos \varpi^2)}, \quad (14)$$

and substituting this in the value of x , we find

$$x = \tan \theta' \cdot \cos \varpi' = \frac{a^2 b^2}{A} \cdot \frac{\sin \theta \cdot \cos \varpi}{\sqrt{A - a^2 \cdot \sin^2 \theta} \{ A \cdot \sin \varpi^2 + b^2 \cdot \cos \varpi^2 \}} - \frac{C}{A}. \quad (15)$$

These equations are identical with those demonstrated by Malus in his *Théorie de la Double Réfraction*, with some slight differences of notation only, arising from our having reckoned ϖ and ϖ' from the opposite point of the circle.

The values of A , B , and C depend only on a , b , and λ , that is, on the peculiar nature of the crystal, which determines the ratio of the axes of the spheroid of double refraction, and on the inclination of the axis to the surface on which the ray is incident. The former are constant for one and the same crystal, however the surface be placed; the latter is constant for any given surface. Hence it appears, that the general law of extraordinary refraction, when we confine ourselves to the consideration of a surface given in position with respect to the axis, resolves itself into an infinite variety of particular laws, some of which we shall now consider.

796.

Particular applications.

Case 1. $\lambda = 0$, the surface perpendicular to the axis; $A = b^2$; $B = a^2$; $C = 0$, and the equations (14) and (15) become

797.

1st. When the axis is perpendicular to the surface.

$$\tan \theta' \cdot \sin \varpi' = \frac{a^2}{b} \cdot \frac{\sin \theta \cdot \sin \varpi}{\sqrt{1 - a^2 \cdot \sin^2 \theta}}; \tan \theta' \cdot \cos \varpi' = \frac{a^2}{b} \cdot \frac{\sin \theta \cdot \cos \varpi}{\sqrt{1 - a^2 \cdot \sin^2 \theta}};$$

these equations (as well as Equation 13) give $\varpi' = \varpi$, so that in this case the plane of refraction is the same with that of incidence, and the extraordinary ray is not deviated out of the vertical plane. Hence, we get simply

$$\tan \theta' = \frac{a^2}{b} \cdot \frac{\sin \theta}{\sqrt{1 - a^2 \cdot \sin^2 \theta}}; \quad (16)$$

which expresses the law of extraordinary refraction in this case. If $\theta = 0$, $\theta' = 0$, or the ray incident perpendicularly passes unrefracted along the axis. If $\theta = 90^\circ$, $\tan \theta' = \frac{a^2}{b \sqrt{1 - a^2}}$. Now if we put $b = \frac{1}{\mu}$ and

$a = \frac{1}{\mu'}$, this becomes

$$\tan \theta' = \frac{\mu}{\mu' \sqrt{\mu'^2 - 1}}; \quad (17)$$

which, μ and μ' being each greater than unity, is always real, so that the ray can enter the crystal however oblique its incidence.

Light.

Case 2. When the axis lies in the surface, or $\lambda = 90^\circ$; $A = a^2$; $B = b^2$; $C = 0$, and the equations become

Part IV.

798.
2d. When
the axis lies
in the sur-
face.

$$\tan \theta' \cdot \sin \varpi' = \frac{a \cdot \sin \theta \cdot \sin \varpi}{\sqrt{1 - \sin^2 \theta \{a^2 \cdot \sin^2 \varpi + b^2 \cdot \cos^2 \varpi\}}}; \quad (18)$$

$$\tan \theta' \cdot \cos \varpi' = \frac{b^2}{a} \cdot \frac{\sin \theta \cdot \cos \varpi}{\sqrt{1 - \sin^2 \theta \{a^2 \cdot \sin^2 \varpi + b^2 \cdot \cos^2 \varpi\}}}; \quad (19)$$

$$\tan \varpi' = \frac{a^2}{b^2} \cdot \tan \varpi = \left(\frac{\mu'}{\mu} \right)^2 \cdot \tan \varpi. \quad (20)$$

The latter of these equations shows that the extraordinary ray deviates from the plane of incidence. The amount of this deviation is nothing when the plane of incidence coincides with the principal section, but increases on either side of it till it attains a certain magnitude, the deviation being *from* the axis, or the plane of refraction making a greater angle with the axis than that of incidence. The two planes then approach each other, and when $\varpi = 90^\circ$, $\tan \varpi = \infty$, $\tan \varpi' = \infty$, and, consequently, $\varpi' = 90^\circ$, or the plane of refraction coincides with that of incidence.

799.
Case of
refraction in
the princi-
pal section.

The equations (18) and (19) show that in the present case, the refracted ray does not describe a conical surface about the perpendicular when the incident one does so, and therefore that the law of refraction varies in every different azimuth. Two cases deserve express notice, *viz.* those in which the plane of incidence is coincident with the principal section, and when perpendicular to it. In the former, $\varpi = 0$ and $\varpi' = 0$, so that we have

$$\tan \theta' = \frac{b^2}{a} \cdot \frac{\sin \theta}{\sqrt{1 - b^2} \cdot \sin \theta^2}. \quad (21)$$

A remarkable relation holds good in this case between the angles of refraction of the ordinary and extraordinary ray, their tangents being to each other in a given ratio. In fact, if we find (θ') = the angle of refraction for the

ordinary ray, we have $\sin(\theta') = \frac{1}{\mu} \cdot \sin \theta = b \cdot \sin \theta$, and, consequently,

$$\tan \theta' = \frac{b}{a} \cdot \frac{\sin(\theta')}{\sqrt{1 - \sin^2(\theta')}} = \frac{b}{a} \cdot \tan(\theta). \quad (22)$$

In the latter case, when the plane of refraction is at right angles to the axis, $\varpi = \varpi' = 90^\circ$, and we get

$$\tan \theta' = \frac{a \cdot \sin \theta}{\sqrt{1 - a^2 \cdot \sin^2 \theta}}; \sin \theta' = a \cdot \sin \theta. \quad (23)$$

800.
Case of re-
fraction at
right angles
to the prin-
cipal sec-
tion.

In this case, therefore, the sine of incidence is in a given ratio to that of refraction, and the extraordinary refraction is performed according to the same law as the ordinary, only with a different index, *viz.* μ' , or $\frac{1}{a}$, instead of μ , or $\frac{1}{b}$. Hence, if we consider only this particular case, the medium will appear to have two indices of refraction, an ordinary and an extraordinary one.

801.
Experimen-
tal method
of deter-
mining the
spheroid of
double re-
fraction.

It was by a careful examination of these cases, that Dr. Wollaston was enabled to verify the Huygenian law. The circumstance last mentioned puts it in our power to determine in the case of any particular crystal the axes of its spheroid of double refraction. We have only to cut a prism of it, having its refracting angle parallel to the axis, and ascertain its indices of refraction according to the principles laid down in the former part of this Essay,

and calling them μ and μ' , the semiaxes of the spheroid will be respectively $\frac{1}{\mu}$ and $\frac{1}{\mu'}$. Thus, in the instance

of carbonate of lime, which Malus examined with the utmost care, he found the two values of a and b to be respectively equal to the numbers 0.67417 and 0.60449, having determined $\mu' = 1.4833$, and $\mu = 1.6543$. (*Théorie de la Double Réfraction*, p. 199.)

802. In this arrangement, however, it is not possible to decide simply from the phenomena of refraction, which is the ordinary, and which the extraordinary ray. There are, however, infallible and easy criteria, as we shall speedily show. Meanwhile, we may for the present content ourselves with observing, that as a moderate deviation from the exact azimuth $\varpi = 90^\circ$ imparts to the extraordinary ray a deviation from the plane of incidence which does not happen to the ordinary one, this may serve for a criterion to distinguish them in certain cases.

803. The square of the velocity of the ordinary ray within the medium is $\mu^2 V^2$, or μ^2 , that is, $\frac{1}{b^2}$, and is constant. That of the extraordinary is V'^2 , or $\frac{W^2}{a^2 b^2}$, that is to say, $\frac{\cos^2 \phi^2}{b^2} + \frac{\sin^2 \phi^2}{a^2}$,

or,

$$V'^2 = \frac{1}{b^2} - \left(\frac{1}{b^2} - \frac{1}{a^2} \right) \cdot \sin^2 \phi^2.$$

Light. The square of the velocity of the extraordinary ray is therefore (in the corpuscular doctrine) diminished by a quantity proportional to the square of the sine of the inclination of the ray within the crystal, to the axis. We say diminished, in the algebraical sense of the word, supposing $a > b$, this agrees with common parlance; but if $a < b$, then it will be increased. This gives rise to the subdivision of the crystallized bodies now treated of into two classes, which have by some been termed attractive and repulsive: by others, positive and negative, which seems preferable, as the former phrases involve theoretical considerations. Positive crystals are, then, such as have a

Part IV.
Division of
crystals
into posi-
tive and
negative.

less than b , or in which the spheroid of double refraction is *prolate*. In these the coefficient $-\left(\frac{1}{b^2} - \frac{1}{a^2}\right)$

which we call k is positive, and the square of the velocity, or $v^2 + k \cdot \sin^2 \theta$, (where $v = \frac{1}{b} =$ velocity of the

ordinary ray within the medium,) is increased by the action of the medium, and is a minimum in the axis. In the negative class the coefficient k is negative, $a > b$, or the spheroid of double refraction is *oblate*, and the velocity of the extraordinary ray is a maximum along the axis. In positive crystals, therefore, the index of ordinary refraction (μ) is less than that of extraordinary; in negative, greater. To the former class belong quartz, ice, zircon, apophyllite, (when uniaxal;) and to the latter, Iceland spar, tourmaline, beryl, emerald, apatite, &c. The negative class, as far as our present knowledge extends, far out-numbers the positive among natural and artificial crystals. They were first distinguished by M. Biot.

In the undulatory doctrine the velocity is the reciprocal of what it is in the corpuscular doctrine, and is therefore *directly* as the radius of the spheroid of double refraction. Hence a wave propagated within the crystal from any point will run over in the same time in different directions, distances proportional to the radii of the spheroid parallel to those directions; and therefore at any instant the surface of the whole wave will be itself a spheroid similar to the spheroid of double refraction. This is Huygens's conception of the subject. It requires us to regard the crystal, or the ether within the crystal through which the undulation is propagated, as having different elasticities in different directions. As far as regards the molecules of a solid body there is no apparent impossibility or improbability in such an idea, but the contrary; but if we regard the propagation of the light within the medium to take place by the elasticity of the ether only, we must then suppose its molecules in crystallized bodies to be in a very different physical state from what they are in free space, and either to be in some manner connected with the solid particles, (forming atmospheres, for instance, about them,) or as subjected to laws of mutual action which approximate to those governing the molecules of solid bodies; and partaking, themselves, of a regular crystalline arrangement and mutual dependency.

804.
Undulations
propagated
in spher-
oidal
surfaces

To pursue the particular applications of the general formulæ (13,) 14,) and (15) farther, would be far beyond our limits. The reader who is curious on this very interesting part of Physical Optics, and who wishes to be delighted and instructed by a combination of consummate mathematical skill with sound experimental research, which may deservedly be cited as a model of the kind, will find every thing which relates to the subject in its best form in the work, already so often cited, of Malus, *Théorie de la Double Réfraction*, which gained the mathematical prize of the French Institute in 1810. To the theory of the internal reflexion of the extraordinary ray which offers many remarkable particularities, as there delivered, we must especially refer him, as well as to his investigation of the foci of lenses formed of doubly refracting crystals, of which we shall here only extract the results, in the single case of a double convex lens having the axis of double refraction in the direction of the axis of the lens.

805.
Malus's
further
researches.

Foci of a
doubly
refracting
lens.

Let r, r' be the radii of the anterior and posterior surfaces of the lens, both supposed convex.

d = distance of the radiant point in the axis.

a, b = the equatorial and polar radii of the spheroid of double refraction, as above.

D = distance of the conjugate focus behind the lens for extraordinary rays.

Δ = extraordinary focal length for parallel rays.

F = ordinary focal length for parallel rays.

Then shall we have for the general expression of D ,

$$D = -\frac{a^2 b d r r'}{d(r+r')(2b^2 - a^2 - a^2 b) - a^2 b r r'}; \quad F = \frac{-b r r'}{(r+r')(1-b)}.$$

If the lens be equi-convex, or $r = r'$,

$$D = -\frac{a^2 b r d}{2(2b^2 - a^2 - a^2 b)d - a^2 b r}; \quad \Delta = -\frac{a^2 b r}{2(2b^2 - a^2 - a^2 b)};$$

$$F = -\frac{b r}{2(1-b)}; \quad \Delta - F = -2F \cdot \frac{a^2 - b^2}{2b^2 - a^2 - a^2 b}.$$

In the case of Iceland spar, these last equations become

$$D = -r \cdot 88,2286; \quad F = -r \cdot 0,7642; \quad D - F = -F \cdot 114,4546;$$

and in the case of rock crystal (quartz)

$$D = -r \cdot 0,9628; \quad F = -r \cdot 0,8958; \quad D - F = -F \cdot 0,0748.$$

To represent, in general, the course of any extraordinarily refracted ray, Huygens has giving the following construction, (fig. 170.) Let HED be the elliptic section of the spheroid of double refraction by the surface, and RC the incident ray falling on C its centre, and BCK the orthographic projection of the ray RC on the

806.

Light. surface. Let HME be the portion of the spheroid within the crystal, whose axis passes through C, and may be anyhow inclined to the surface. Then will the surface of this spheroid be the boundary of the wave propagated from C as a centre, after the lapse of a given time. Draw CO in the plane RCK at right angles to RC, and make OK (perpendicular to CK, or parallel to RC) equal to the space described by light in the medium exterior to the crystal in the same given time. This will determine the point K in the line BCK. Through K draw KT perpendicular to BK, and about KT as an axis let a plane revolve passing through K T, till it touches the surface of the spheroid in I. Join CI, and CI is the extraordinary refracted ray.

807. The demonstration of this construction (granting the principle of spheroidal undulations) is evident, if we consider the manner in which the general wave, a perpendicular to whose surface forms what we term a ray of light, (at least in singly refracting media,) arises from the reunion of all the elementary waves propagated from every part of the surface, (Art. 566.) In this construction, if we conceive a plane wave from an infinitely distant luminary perpendicular to RC to move along RC, every point in the line CK will become *in succession*, and every point in the line CD perpendicular to CK, or parallel to KT *simultaneously*, a centre of vibration. The general wave, therefore, will be a surface touching all ellipsoids described about each point of the surface, having their axes parallel, their generating ellipses similar, and their linear dimensions proportional to the distance of their centre from the line KT. Of course it can be no other than the tangent plane IKT drawn as above.

808. This then will be the form and position of the general wave within the crystal. Now if we consider only that very minute portion of it which emanates from C, it is evident that I is the corresponding point in it; and therefore CI is necessarily the direction of the ray, because I is the point on which that portion of the general wave transmitted through a very small aperture at C would fall.

809. Thus we see, that in the case of the extraordinary ray, we are no longer to regard the *ray* as a perpendicular to the surface of the wave. It is propagated obliquely to that surface. So soon, however, as the wave emerges into the ambient medium, the usual law of perpendicular propagation is restored.

Oblique propagation of the extraordinary ray. 810. To show the identity of the law of extraordinary refraction resulting from this construction with that expressed by the general equations (13,) (14,) and (15,) we have only to translate it into analytical language. This has been done by Malus, in his work above referred to; and the reader may also consult Biot's *Traité Général de Physique*, for a more elementary exposition of the process, which is one of considerable complexity, for which reason we shall not embarrass ourselves with it here.

811. Some very remarkable and important consequences follow from this mode of viewing the subject. It appears that when a plane wave is incident on a doubly refracting surface, the transmitted *extraordinary wave* is also plane, and advances with a uniform velocity in a direction oblique to itself. Consequently the velocity is also uniform in a direction perpendicular to itself. Moreover, its common section with the surface is always parallel to KT, or to the common section of the incident wave with the same surface. Hence, it is evident, that it moves in the same way as an ordinarily transmitted wave would do, and at any instant has the same position that such a wave would have, provided the index of refraction in the latter case were properly assumed. The only difference is, that the motions of the vibrating molecules, of which they respectively consist, are executed in different planes. Now, when this wave emerges from the medium, it obeys the same laws as on its entry, only reversed; so that it still continues a plane wave, and its common section with the surface of *emergence* remains unaltered.

812. Hence it follows, that if we cut a prism of any doubly refracting crystal with one axis, and transmit through it a ray incident in a plane at right angles to the edge of the prism, the ordinary and extraordinary ray will both emerge in that plane, and their separation will take place in a plane containing the incident and ordinarily-refracted ray, and will therefore be, apparently, such as would arise from attributing two ordinary refractive powers to the medium. It is only when the edge of the prism is oblique to the plane of incidence, that the extraordinary ray can deviate from the plane containing the incident and ordinarily refracted rays.

813. We see, then, that in the theory of extraordinary refraction, it is necessary to consider, as distinct, two things, which, in that of ordinary, are one and the same, *viz.* the velocity of the *luminous waves*, and the *velocity of the rays of light*. This distinction will require to be very carefully kept in view hereafter, when we come to treat of the law of refraction in crystals with two axes of double refraction. For this, however, we are not yet prepared, as the knowledge of this law presupposes an acquaintance with a multitude of facts relative to the polarization of light, of which we have yet said nothing. It will suffice here to mention, that the whole doctrine of double refraction has recently undergone a great revolution; one, indeed, which may be said to have changed the face of Physical Optics, in consequence of the researches of M. Fresnel. It had all along been taken for granted, that in crystals possessed of double refraction, one of the pencils followed the ordinary law of proportional sines. It had, moreover, been ascertained, by experiments hereafter to be related, that the difference of the squares of the velocities between the two pencils is in all cases proportional to the product of the sines of the angles contained between the extraordinary ray (as it was termed) and the two axes, or directions in which the refraction is single. It was hence concluded, that the velocity of the *extraordinary pencil* was in all cases

represented by $\sqrt{v^2 + k \cdot \sin \phi \cdot \sin \phi'}$, v being that of the ordinary one, and k a constant depending on the nature of the crystal, and ϕ, ϕ' the angles in question. This granted, there would be no difficulty in determining the form of the surface of double curvature, which should be substituted for the Huygenian spheroid; so as to render the same construction with that described in Art. 806, or the general formulæ in Art. 792, applicable to this case. In fact, if we call α the semi-angle between the two axes, and conceive three coordinates x, y, z , of which x bisects that angle, the plane of the x, y containing both axes, it is easy to see, by spherical trigonometry, that we must have

$$\cos \phi = \frac{x \cdot \cos a + y \cdot \sin a}{\sqrt{x^2 + y^2 + z^2}}; \quad \cos \phi' = \frac{x \cdot \cos a - y \cdot \sin a}{\sqrt{x^2 + y^2 + z^2}}.$$

Hence, since $r (\sqrt{x^2 + y^2 + z^2})$ the radius of the surface of the wave, is always equal to

$$\frac{1}{\sqrt{v}}, \text{ or } \frac{1}{\sqrt{v^2 + k^2 \cdot \sin \phi \cdot \sin \phi'}},$$

a simple substitution would give at once the equation of its surface as referred to the three coordinates x, y, z ; namely,

$$0 = (k^2 - v^2) (x^2 + y^2 + z^2)^2 + 2 (x^2 + y^2 + z^2) (v^2 - k^2 x^2 \cdot \cos a^2 - k^2 y^2 \cdot \sin a^2) + k^2 (x^2 \cdot \cos a^2 + y^2 \cdot \sin a^2)^2 - 1,$$

which it would be easy then to transform into functions of r, ω , and θ , as required for the application of the general analytical formulæ by the usual substitutions

$$z = r \cdot \sin \theta; \quad y = r \cdot \sin \theta \cdot \sin \omega; \quad x = r \cdot \sin \theta \cdot \cos \omega.$$

The researches of M. Fresnel, however, as before remarked, have destroyed the basis on which this theory rested, by demonstrating the non-existence of an ordinarily refracted ray in the case of crystals with two axes. The theory which he has substituted in its place, however, and which it is impossible to regard otherwise than as one of the finest generalizations of modern science, we must reserve for a more advanced place in this essay. We shall now proceed to treat

Of the Polarization of Light.

The phenomena which belong to this division of our subject are so singular and various, that to one who has only studied the subject of Physical Optics under the relations presented in the foregoing pages, it is like entering into a new world,—so splendid as to render it one of the most delightful branches of experimental inquiry; and so fertile in the views it lays open of the constitution of natural bodies, and the minuter mechanism of the universe, as to place it in the very first rank of the physico-mathematical sciences, which it maintains, by the rigorous application of geometrical reasoning its nature admits and requires. The intricacy as well as variety of its phenomena, and the unexampled rapidity with which discoveries have succeeded each other in it, have hitherto prevented the possibility of embodying it satisfactorily in a systematic form; but, after the rejection of numberless imperfect generalizations, it seems at length to have acquired that degree of consistency as to enable us—not, indeed, to deduce every phenomenon, by distinct steps, from one general cause—but to present them, at least, in something like a regular succession; to show a mutual dependence between their several classes, which is a main step to a complete generalization; and to dispense with the bewildering detail of an immense multitude of individual facts, which, having served their purpose in the inductive process, must in future be considered as having their interest merged in that of the laws from which they flow.

814.

§ II. General Ideas of the Distinction between Polarized and Unpolarized Light.

In all the properties and affections of light which we have hitherto considered, we have regarded it as presenting the same phenomena of reflexion and transmission, both as respects the direction and intensity of the reflected or transmitted beam, however it may be presented to the reflecting or refracting surface, provided the angle of incidence, and the plane in which it lies, be not varied. And this is true of light in the state in which it is emitted immediately from the sun, or from other self-luminous sources. A ray of such light, incident at a given angle on a given surface, may be conceived to revolve round an axis coincident with its own direction; or, which comes to the same thing, the reflecting or refracting surface may be actually made to revolve round the ray as an axis, preserving the same relative situation to it in all other respects, and no change in the phenomena will be perceived. For instance, if in a long cylindrical tube we fix a plate of glass, or any other medium at any angle of inclination to the axis; and then, directing the tube to the sun, turn the whole apparatus round on its axis, the intensity of the reflected or refracted ray will suffer no variation, and its direction (if deviated) will revolve uniformly round with the apparatus, so that if received on a screen connected invariably with the tube, it will continue to fall on the very same point in all parts of its rotation. Or we may receive the light from a piece of white hot iron at any angle on any medium, and its phenomena will be precisely the same, whether the iron be at rest, or be made to revolve round an axis coincident with the direction of the ray.

815.

But, if instead of employing a ray immediately emitted from a self-luminous source, we subject to the same examination a ray that has undergone some reflexions, refractions, or been in any one of a great variety of ways subjected to the action of material bodies, we find this perfect uniformity of result no longer to hold good. It is no longer indifferent in what plane, with respect to the ray itself, the reflecting or refracting surface is presented to it. It seems to have acquired *sides*; a right and left, a front and back; and the *intensity*, though not the direction of the reflected or transmitted portion, depends materially on the position with respect to these

816.

Polarized rays have acquired fixed relations to external space.

Light. sides, in which the plane of incidence lies, though every thing else remains precisely the same. In this state it is said to be polarized. The difference between a polarized and an ordinary ray of light can hardly be more readily conceived than by assimilating the latter to a cylindrical, and the former to a four-sided prismatic rod, such as a lath or a ruler, or other long, flat, straight stick. It is evident that the cylinder, if inclined to any surface at a given angle in a given plane, may be turned round its own axis without altering its relations to the plane, while those of the prism will vary essentially according to the position of its sides. Let us suppose, for instance, (it is but a simile, which we do not wish the reader to dwell on for a moment, or to imagine that any analogy is hereafter intended to be established,) that we had occasion to thrust such a rod into a surface composed of detached fibres, all lying in one direction, or of scales or laminae arranged parallel to one another, we should find a much greater facility of penetration on presenting its broad side in the direction of the laminae or fibres, than transverse to them. A thin sheet may be slipped between the bars of a grating, which would present an insuperable obstacle to it if presented cross-wise.

817. But, to be more particular, and to give a more clear conception of the marked distinction which exists between a polarized and an unpolarized ray. There are many crystallized minerals, which when cut into parallel plates are sufficiently transparent, and let pass abundance of light with perfect regularity, but which, nevertheless, at its emergence is found to have acquired that peculiar modification here in question. One of the most remarkable of these is the tourmaline. This mineral crystallizes in long prisms, whose primitive form is the obtuse rhomboid, having its axis parallel to the axis of the prism. The lateral faces of these prisms are frequently so numerous as to give them an approach to a cylindrical or cylindroidal form. Now if we take one of these crystals, and slit it (by the aid of a lapidary's wheel) into plates parallel to the axis of the prism of moderate and uniform thickness, (about $\frac{1}{20}$ of an inch,) which must be well-polished, luminous objects may be seen through them, as through plates of coloured glass. Let one of these plates be interposed perpendicularly between the eye and a candle, the latter will be seen with equal distinctness in every position of the axis of the plate with respect to the horizon, (by the axis of the plate is meant any line in it parallel to the axes of its molecules, or to the axis of the prism from which it was cut.) And if the plate be turned round on its own plane, no change will be perceived in the image of the candle. Now, holding this first plate in a fixed position, (with its axis vertical, for instance,) let a second be interposed between it and the eye, and turned round slowly in its own plane, and a very remarkable phenomenon will be seen. The candle will appear and disappear alternately at every quarter revolution of the plate, passing through all gradations of brightness, from a maximum down to a total, or almost total, evanescence, and then increasing again by the same degrees as it diminished before. If now we attend to the position of the second plate with respect to the first, we shall find that the *maxima* of illumination take place when the axis of the second plate is parallel to that of the first, so that the two plates have either the same positions with respect to each other that they had in the original crystal, or positions differing by 180° , while the *minima*, or evanescences of the image, take place exactly 90° from this parallelism, or when the axes of the two plates are exactly crossed. In tourmalines of a good colour, the stoppage of the light in this situation is total, and the combined plate (though composed of elements separately very transparent and of the same colour) is perfectly opaque. In others it is only partial; but however the specimens be chosen, a very marked defalcation of light in the crossed position takes place. We shall at present suppose that the specimens employed possess the property in question in its greatest perfection. Now it is evident that the light which has passed through the first plate has acquired in so doing a property totally distinct from those of the original light of the candle. The latter would have penetrated the second plate equally well in all its positions; the former is incapable altogether of penetrating it in some positions, while in others it passes through readily, and these positions correspond to certain *sides* which the ray has acquired, and which are parallel and perpendicular respectively to the axis of the first plate. Moreover, these *sides* once acquired, are retained by the ray in all its future course, (provided it be not again otherwise modified by contact with other bodies,) for it matters not how great the distance between the two plates, whether they be in contact or many inches, yards, or miles asunder, not the least variation is perceived in the phenomenon in question. If the position of the first plate be shifted, the sides of the transmitted ray shift with it, through an equal angle, and the second will no longer extinguish it in the position it at first did, but must be brought into a position removed therefrom, by an angle equal to that through which the first plate has been made to revolve.

818. A great many other crystallized bodies besides the tourmaline possess this curious property, and several in great perfection. The tourmaline, however, is one easily procured, and being exceedingly useful in optical experiments, we would recommend the reader who has any desire to familiarize himself with the practical manipulations of this branch of optical science, to provide himself with a good pair of corresponding plates of this mineral, cut and polished as above directed. The colour is a point of great moment. Those of a blue or green colour possess the property in question very imperfectly; the yellow varieties, unless when verging to greenish brown, are equally improper, the best colour is a hair-brown, or purplish brown, and they may be slit and polished by any lapidary.

819. But it is not only by such means that the *polarization* of a pencil of light may be operated, nor is this the only character which distinguishes polarized from ordinary light. We shall, therefore, describe in order, the principal means by which the polarization of light may be performed, and the assemblage of characters which are invariably found to coexist in a ray when polarized.

The chief modes by which the polarization of light may be effected, are

1st. By reflexion at a proper angle from the surfaces of transparent media.

2d. By transmission through a regularly crystallized medium possessed of the property of double refraction.

3d. By transmission through transparent, uncrystallized plates in sufficient number, and at proper angles.

Property of the tourmaline and other crystals.

Selection of proper tourmalines.

Various modes of polarizing light.

Light.

4th. By transmission through a variety of bodies, such as agate, mother-of-pearl, &c. which have an approach to a laminated structure, and an imperfect state of crystallization.

The characters which are invariably found to coexist in a polarized ray, being the chief of those by which it may be most easily recognised as polarized, are—

1. Incapability of being transmitted by a plate of tourmaline, as above described, when incident perpendicularly on it, in certain positions of the plate; and ready transmission in others, at right angles to the former.

2. Incapability of being reflected by polished transparent media at certain angles of incidence, and in certain positions of the plane of incidence.

3. Incapability of undergoing division into two equal pencils by double refraction, in positions of the doubly refracting bodies, in which a ray of ordinary light would be so divided.

Besides which, there might be enumerated a vast variety of other characters, which, however, it will be better to regard as *properties* at once of polarized light, and of the various media which affect it. It cannot fail to be remarked, that all these characters are of the *negative kind*, and consist in denying to polarized light properties which ordinary light possesses, and that they are such as affect the intensity of the ray, not its direction. Thus, the direction which a polarized ray will take under any circumstances of the action of media, is never *different* from what an unpolarized ray might take, and from what a portion of it at least actually does. For instance, when an unpolarized ray is separated by double refraction into two equal pencils, a polarized ray will be divided into two unequal ones, one of which may even be altogether evanescent, but their directions are precisely the same as those of the pencils into which the unpolarized ray is divided. Hence we may lay it down as a general principle, that the *direction* taken by a polarized ray, or by the parts into which it may be divided by any reflexions, refractions, or other modifying causes, may always be determined by the same rules as apply to unpolarized light; but that the relative *intensities* of these portions differ from those of similar portions of unpolarized light, according to certain laws which it is the business of the optical inquirer to ascertain.

Part IV.

820.

Characters of a polarized ray of light.

Affect the intensity and not the direction of the ray.

§ III. Of the Polarization of Light by Reflexion.

When a ray of direct solar light is received on a plate of polished glass or other medium, a portion more or less considerable is always reflected. The intensity of this portion depends only on the nature of the medium and on the angle of incidence, being greater as the refractive power of the former is greater, and as the ray falls more obliquely on the surface. But it is, moreover, found, that at a certain angle of incidence, (which is therefore called the *polarizing angle*;) the reflected ray possesses all the characters above enumerated, and is therefore polarized.

821.

Light polarized by reflexion.

This remarkable fact was discovered by Malus in 1808, when accidentally viewing, through a doubly refracting prism, the light of the setting sun reflected from the glass windows of the Luxembourg Palace in Paris. On turning round the prism, he was surprised to observe a remarkable difference in the intensity of the two images; the most refracted alternately surpassing and falling short of the least in brightness, at each quadrant of the revolution. This phenomenon connecting itself in his mind with similar phenomena produced by rays which had undergone double refraction, and with which, from the researches he was then engaged in, he was familiar, led him to investigate the circumstances of the case with all possible attention, and the result was the creation of a new department of Physical Optics. So true it is, that a thousand indications pass daily before our eyes which might lead to the most important conclusions. The seeds of great discoveries are everywhere present and floating around us, but they fall in vain on the unprepared mind, and germinate only where previous inquiry has elaborated the soil for their reception, and awakened the attention to a perception of their value.

822.

Discovery by Malus.

To make this new property acquired by the reflected ray evident by experiment, let any one lay down a large plate of glass on a black cloth, on a table before an open window, and placing himself conveniently so as to look obliquely at it, let him view the reflected light of the sky, (or, which is better, of the clouds if not too dark,) from the whole surface, which will thus appear pretty uniformly bright. Then let him close one eye, and apply before the other a plate of tourmaline, cut as above directed, so as to have *its axis in a vertical plane*. He will then observe the surface of the glass, instead of being as before equally illuminated, to have on it, as it were, an obscure cloud, or a large blot, the middle of which is totally dark. If this be not seen at first, it will come into view on elevating or depressing the eye. If the inclination of a line drawn from the centre of the dark spot to the eye be measured, it will be found to make an angle of about 33° with the surface of the glass. If now, keeping the eye fixed on the spot, the tourmaline plate (which it is convenient to have set in a small circular frame for such experiments) be turned slowly round in its own plane, the spot will grow less and less obscure, and when the *axis of the tourmaline is parallel to the reflecting surface*, (or horizontal,) will have disappeared completely, so as to leave the surface equally illuminated, and, on continuing the rotation of the tourmaline, will appear and vanish alternately.

823.

Experiment

It appears from this experiment, that the ray which has been reflected from the surface of the glass at an inclination of 33° , or an incidence of 57° , has thereby been deprived of its power to penetrate a tourmaline plate whose axis lies in the plane of incidence. It has therefore acquired the same character, or (so far as this goes, at least) undergone the same modification as if, instead of being reflected on glass, it had been transmitted through a tourmaline plate, whose axis was perpendicular to the plane of reflexion.

824

It has, moreover, acquired all the other enumerated characters of a polarized ray. And, first, it has become

Light.

825.
Experiment.
The polarized ray
incapable of
a second reflexion, &c.
Fig. 171.

incapable of reflexion at the surface of glass, or other transparent media at certain definite angles, and in certain positions of the plane of incidence. To show this experimentally, let a piece of polished glass have one of its surfaces roughened, and blackened with melted pitch or black varnish, so as to destroy its internal reflexion, and let this be fixed on a stand, so as to be capable of varying at will the inclination of its polished surface to the horizon, and of turning it round a vertical axis in any azimuth. A very convenient stand of this kind is figured in fig. 171, consisting of a cylindrical support A sliding in a vertical tube B, attached to a round base F like a candlestick, and carrying an arm C, which can be set to any angle of inclination to the horizon by means of a stiff shoulder joint D. To this arm the blackened glass E is fixed, having its plane parallel to the axis of the joint D. Let this apparatus be set on a table, so that the rays reflected from a pretty large plate of glass G, at an angle of about 57° (of incidence) shall be received on the glass E, which ought to be inclined with its polished surface looking downwards, and making an angle of about 73° with the horizon, see Art. 842. Then let the observer apply his eye near the glass E, so as to see the glass G reflected in it, and slowly turn the stand F round in a horizontal plane, keeping always the reflected image of G in view. He will then perceive, that at a certain point of the rotation of the stand, the illumination of this image, which in other situations is very bright, will undergo a rapid diminution, and at last wholly disappear, and (if the glass G be large enough) the same appearance of a cloud or large dark spot will then be visible upon it. If the inclination of the arm C D be correct, it will be easy to find such a position by turning the stand a little backwards and forwards, as shall make the centre of this spot totally black; if not, bring it to as great a degree of obscurity as possible by the horizontal motion, then, holding fast the stand, vary a little one way or another the inclination of the reflector E, and a very complete obscurity will readily be attained.

826.
Another
mode of
making the
experiment.
Fig. 172.

Another, and, for some experimental purposes, a better way of exhibiting the same phenomenon, is to take two metallic or pasteboard tubes, open at both ends, and fitting into each other so as to turn stiffly. Into each of these, at the end remote from their junction, fix with wax, or in a frame, a plate of glass, blackened at the back as above described, so as to make an angle of 33° with the axis of the tube, as represented in fig. 172. Then having placed the tube containing one of the plates (A) so that the light from any luminary, reflected at the plate shall traverse the axis of the tube, fix it there, and the reflected ray will be again reflected at B, and on its emergence may be received on a screen or on the eye. Now make the tube containing the reflector B revolve within the other, so that that reflector shall revolve round the ray A B as an axis, preserving the same inclination. Then will the twice reflected ray revolve with equal angular motion, and describe a conical surface. But in so doing, it will be observed to vary in intensity, and at two points of the revolution of the tube B will disappear altogether. Now if we attend to the position of the reflectors at this moment, it will be found that the planes of the first and second reflexion make a right angle.

827. By repeating these experiments with all sorts of reflecting media, and determining by exact measurement the angles at which the original ray must be incident that polarization shall take place, and those at which a polarized ray ceases to be reflected, the following laws have been ascertained to hold good, previous to announcing which a definition will be necessary.

828.
Plane of
polarization
defined.

Definition. The plane of polarization of a polarized ray is the plane in which it must have undergone reflexion, to have acquired its character of polarization; or that plane passing through the course of the ray perpendicular to which it cannot be reflected at the polarizing angle from a transparent medium; or, again, that plane in which, if the axis of a tourmaline plate exposed perpendicularly to the ray be situated, no portion of the ray will be transmitted. Also, a polarized ray is said to be polarized in its plane of polarization, as just defined.

829.
Sides of a
polarized
ray.

The plane of polarization of any polarized ray is to be considered as one of the sides of the ray which thus, in all its future progress, carries with it certain relations to surrounding fixed space, which must be regarded, while they continue unchanged, as inherent in the ray itself, and as having no further any relation to the particular mode in which they originated.

830.
Laws of po-
larization by
reflexion.
Law 1.

The laws of polarization by reflexion are these:
Law 1. All reflecting surfaces are capable of polarizing light if incident at proper angles; only, metallic bodies, or bodies of very high refractive powers, appear to do so but imperfectly, the reflected ray not entirely disappearing in circumstances when a perfectly polarized ray would be completely extinguished. Of this more hereafter.

831.
Law 2.
Brewster's
law of the
tangents.

Law 2. Different media differ in the angles of incidence at which they polarize light; and it is found, that these angles may always be determined from the following simple and elegant relation, discovered by Dr. Brewster after a laborious examination of an infinite variety of substances.

The tangent of the polarizing angle for any medium is the index of refraction belonging to that medium.

Thus, the indices of refraction of water, crown-glass, and diamond, being respectively 1.336, 1.535, and 2.487, their respective polarizing angles will be $53^\circ 11'$, $56^\circ 55'$, and $68^\circ 6'$. For diamond, however, or bodies of very high refractive powers, we must understand by the polarizing angle, that angle of incidence at which the reflected ray approximates most nearly to the character of a ray completely polarized.

832.
All the
colours not
polarized at
one inci-
dence.

It follows from this law, that one and the same medium does not polarize all the coloured rays at the same angle, and that therefore the disappearance of the reflected pencil can never be total, except where the incident ray is homogeneous. This will account in some degree for the want of complete polarization of a white ray, reflected at any angle from highly refractive media, which are generally also highly dispersive. Of the reality of the fact, it is easy to satisfy oneself by a very simple experiment, which we have often made. Receive a sun-beam on a plane glass, with the back roughened and blackened, at an incidence (θ) nearly equal to the polarizing angle (α), and let the reflected ray pass into a darkened room, and fall on another similar glass, which may be held in the hand, so as to reflect the ray in a plane at right angles to that of the first reflection, and

Light. also at an angle (θ') nearly equal to the polarizing angle (a') of the second plate. It will be easy to find a position where the reflected ray (which must be received on a white screen) very nearly vanishes; but no adjustment of the angles of incidence θ and θ' will produce a total disappearance. When the disappearance is most nearly total, the reflected light is coloured of a neutral purple; the yellow, or most luminous rays, being now totally extinguished. In this position, if θ remain constant, and θ' the incidence on the second plate be varied a little on one side or the other of the polarizing angle a' , the reflected ray assumes on the one hand a pretty intense blue-green, and on the other a ruddy plum colour or amethyst red. The several changes of tint, arising from variations of incidence on both plates, were observed to be as follows:

Part IV.
Proved by experiment.

1st. $\theta < a$;	$\theta' < a'$;	Reflected ray,	Strong green.
	Intermediate,	_____	White.
	$\theta' > a'$;	_____	Pale red or amethyst.
2d. $\theta = a$;	$\theta' < a'$;	_____	Strong blue green.
	$\theta' = a'$;	_____	Neutral purple.
	$\theta' > a'$;	_____	Strong plum colour.
3d. $\theta > a$;	$\theta' < a'$;	_____	Light greenish blue.
	Intermediate,	_____	White.
	$\theta' > a'$;	_____	Strong red, or plum colour.

The rationale of these changes of colour will be more evident when we have announced the following law, which expresses one of the most general and distinguishing characters of polarized light.

Law 3. When a polarized ray (no matter how it acquired its polarization) is incident on a reflecting surface of a transparent, or other medium capable of completely polarizing light, in a plane perpendicular to that of the ray's polarization, and at an angle of incidence equal to the polarizing angle of the medium, no portion of the ray will be reflected. If the medium be of such a nature as to be capable only of incompletely polarizing light, a portion will be reflected, but much less intense than if the incident ray were unpolarized.

833.
Law 3.
Non-reflexibility of polarized light in certain, and what cases.

It is evident that this property may be employed to distinguish polarized from common light, as well as that of extinction by a plate of tourmaline. It is, however, much less convenient though better adapted for delicate inquiries.

The polarizing angle for white light is, in fact, the angle for the most luminous or mean yellow rays; and when the two reflexions, in planes at right angles to each other, are made at this angle, the yellow rays only totally escape reflexion, but a very small portion both of the red and blue end of the spectrum are reflected, and form a feeble purple beam, such as above described. The polarizing angle for red rays being less than for violet, it is evident that when either θ or θ' is equal to the polarizing angle for red, it will be less than that of yellow, and still less than that of blue and violet rays; thus, the red disappears most completely from the reflected beam in those cases when θ or θ' are less than a or a' , leaving an excess of the green and blue rays, and *vice versa* in the converse cases. Thus, too, if θ be $< a$, and at the same time $\theta' < a'$, the colour produced will be a more intense green than if the incidences deviated opposite ways from the polarizing angles; and it is evident, that a compensation may arise from the effect of such opposite deviations giving an intermediate white ray, exactly as we see to have happened.

834.
Explanation of the colours in the last experiment.

Some very remarkable consequences follow from the law announced by Dr. Brewster for finding the polarizing angle, which may be presented in the form of distinct propositions. Thus,

835.

Prop. 1. When a ray is incident on a transparent surface, so that the reflected portion shall be completely polarized, the reflected and refracted portions make a right angle. For θ being the angle of incidence, we have

836.
Consequences of the law of polarization.

$\tan \rho = \mu$ and ρ , being the angle of refraction, $\sin \rho = \frac{\sin \theta}{\mu} = \frac{\sin \theta}{\tan \theta} = \cos \theta$. Therefore $\rho = 90^\circ - \theta$, but θ

being the angle of incidence is also that of reflexion, and $\rho + \theta$ is therefore equal to the supplement of the angle between the reflected and refracted rays, which is therefore a right angle. Q. E. D.

Prop. 2. When a beam of common light is incident at the polarizing angle on a parallel plate of a transparent medium, not only the portion reflected at the first surface, but also that reflected internally at the second, and the compound reflected ray, consisting of both united, are polarized.

837.
Polarization by internal reflexion.

Since $\sin \rho = \cos \theta$, and since ρ is also the angle of incidence on the second surface, we shall have $\tan \rho =$

$\cotan \theta = \frac{1}{\tan \theta} = \frac{1}{\mu} =$ index of refraction out of the medium. Hence, ρ is the angle of polarization for rays

internally incident, and therefore that portion of the beam which, having penetrated the first surface, falls on the second, being incident at its polarizing angle, the portion reflected here will also be polarized, and being again incident on the first surface, in the plane of its polarization, that part of it which is transmitted will (as we shall see hereafter) suffer no change in its plane of polarization, so that both it and the first reflected ray will come off polarized in the same plane. Q. E. D.

Corol. 1. Hence, to obtain a stronger polarized ray, we may dispense with roughening or blackening the posterior surface, provided we are sure that the surfaces are truly parallel.

838.

If a series of parallel plates be laid one on the other so as to form a pile, the portions reflected from the several surfaces all come off polarized in the same plane, and by this means a very intense polarized ray may be obtained. It can never, however, for a reason we shall presently state, contain more than half the incident light, whatever be the number of plates employed.

839.
Mode of obtaining an intense polarized beam.

Light. For a great variety of optical experiments, a pile consisting of ten or a dozen panes of common window-glass set in a frame, is of great use and very convenient. Such a pile laid down before an open window affords a dispersed beam, each ray of which is polarized at the proper angle, and of great intensity and very proper for the exhibition of many of the phenomena hereafter to be described.

841. *Prop. 3.* If a ray be completely polarized by reflexion at the surface of one medium, and the reflected ray completely transmitted or absorbed at that of a second, Required the inclination of the two surfaces to each other?

Let a and a' be the polarizing angles of the respective media; then, since the planes of reflexion are at right angles to each other, and a, a' are the angles of incidence, if we call I the inclination required, we shall have by Art. 104, $\cos I = \cos a \cdot \cos a'$. Now, if μ, μ' be the refractive indices of the media, we have $\tan a = \mu$, $\tan a' = \mu'$, and therefore

$$\tan I = \sqrt{\mu^2 + \mu'^2 + \mu^2 \mu'^2}.$$

842. *Corol. 1.* If the media be both alike,

$$\tan I = \mu \cdot \sqrt{2 + \mu^2}; \text{ or } \cos I = \frac{1}{1 + \mu^2}.$$

Thus, in the case of crown-glass, $\mu = 1.535$ and $I = 72^\circ 40'$, as in Art. 825.

843. By the help of this law, connecting the angle of polarization with the refractive index, we may easily deduce the one from the other. This affords a valuable and ready resource in cases to which other methods can hardly be applied, for ascertaining the refractive powers of media, which are either opaque, or in such small or irregularly shaped masses, that they cannot be used as prisms. For ascertaining the angle of polarization, only one polished surface, however small, is necessary, and we have only to receive a ray reflected from it on a blackened glass, or other similar medium of known refractive index, at the polarizing angle, and in a plane perpendicular to that at which it is reflected by the surface under examination. For this purpose it is convenient to have the glass plate (or, which is better, a polished plate of obsidian or dark coloured quartz) set in a tube diagonally, so as to reflect laterally the ray which traverses the axis of the tube. At the other end, the substance to be examined must be fixed on a revolving axis perpendicular to the axis of the tube, and having its plane adjusted so as to be parallel to the former, which must then be turned round till the dispersed light of the clouds, reflected by it, is entirely extinguished by the obsidian plate, and the inclination of the reflecting surface to the axis of the tube in this situation may be measured by a divided circle, connecting with the axis of rotation. By this means we may ascertain the polarizing angles, and therefore the refractive indices of the smallest crystals, or of polished stones, gems, &c., set in such a manner as not to admit of other modes of examination. To insure a fixed zero point on the graduated circle, the following mode (among many others) may be resorted to. Let a polished metallic reflector or small piece of looking-glass be permanently attached to the revolving axis, so that its plane shall be perpendicular to the axis of the tube, when the index of the divided circle marks $0^\circ 0'$. This adjustment being made once for all, let the surface to be examined be attached by wax or otherwise, not to the axis itself, but to a ring turning stiffly on it. Then, bringing the image of the sun, or any very distant object, sufficiently bright or well defined, seen in the reflector, to coincide with any other equally well defined, and also at a great distance, alter the attachment of the substance by pressure on the wax, and by turning round the ring, till a similar coincidence is obtained when the eye is transferred to it. Then we are assured that the two surfaces are parallel, and that therefore the reading off on the circle measures the true angle between the axis of the tube and the perpendicular, or the angle of reflexion, or at least differs from it only by a constant quantity, which may be ascertained at leisure, and applied as index error. (This mode of bringing a movable surface to a fixed position with respect to the divisions of an instrument, is applicable to a great variety of cases, and is at once convenient and delicate.)

844. Dr. Brewster has remarked, that glass surfaces frequently exhibit remarkable, and apparently unaccountable, deviations from the general law; but on minute examination he found that this substance is liable to a superficial tarnish, or formation of infinitely thin films of a different refractive power from the mass of glass beneath. As the polarized ray never penetrates the surface, its angle of polarization is determined solely by this film, which is too thin to admit of any direct measure of its refractive index. When this tarnish has gone to a great extent, scales of glass detach themselves, as is seen in very old windows, (especially those of stables,) and even in green glass bottles which have long lain in damp situations, and which acquire a coat actually capable of being mistaken for gilding.

845. In metallic or adamantine bodies, which polarize light but imperfectly, that angle at which the reflected beam approaches nearer in its character to those described as of polarized light, is to be taken for the angle of polarization, and from it the refractive power may still be found. The results deduced by this means for metallic bodies, agree with those obtained from the quantity of light reflected, in assigning very high refractive powers to them. Thus, for steel the polarizing angle is found to be above 71° , and for mercury $76\frac{1}{2}^\circ$, and their indices of refraction are, therefore, respectively 2.85 and 4.16. This latter result, indeed, differs greatly from that of Art. 594, but the observations are so uncertain, and the angle of greatest polarization so indefinitely marked, (not to mention the errors to which a determination of the reflective power itself is liable to,) that we cannot expect coincidence in such determinations. Perhaps 5.0 may be taken as a probable index.

846. The law of polarization announced by Dr. Brewster is general, and applies as well to the polarization of light at the separating surfaces of two media in contact, as at the external or internal surface of one and the same medium. He has attempted to deduce from it several theoretical conclusions, as to the extent and mode of action of the reflecting and refracting forces, for which we must refer the reader to his Paper on the subject *Philosophical Transactions*, 1816

Light.

If a ray be reflected at an angle greater or less than the polarizing angle, it is *partially polarized*, that is to say, when received at the polarizing angle on another reflecting surface, which is made to revolve round the reflected ray without altering its inclination to it, the twice reflected ray never vanishes entirely, but undergoes alternations of brightness, and passes through states of maxima and minima which are more distinctly marked according as the angle of the first reflexion approaches more nearly to that of complete polarization. The same is observed when a ray so *partially* polarized is received on a tourmaline plate, revolving (as above described) in its own plane. It never undergoes complete extinction, but the transmitted portion passes through alternate maxima and minima of intensity, and the approach to complete extinction is the nearer the nearer the angle of reflexion has been to the polarizing angle. We may conceive a partially polarized ray to consist of two unequally intense portions; one completely polarized, the other not at all. It is evident that the former, periodically passing from evanescence to its total brightness, during the rotation of the tourmaline or reflector, while the latter remains constant in all positions, will give rise to the phenomenon in question. And all the other characters of a *partially polarized ray* agreeing with this explanation, we may receive it as a principle, that when a surface does not completely polarize a ray, its action is such as to leave a certain portion completely unchanged, and to impress on the remaining portion the character of complete polarization. Thus we must conceive polarization as a property or character not susceptible of *degree*, not capable of existing sometimes in a more, sometimes in a less, intense state. A single elementary ray is either wholly polarized or not at all. A *beam* composed of many *coincident rays* may be partially polarized, inasmuch as some of its component rays only may be polarized, and the rest not so. This distinction once understood, however, we shall continue to speak of a ray as wholly or partially polarized, in conformity with common language. We shall presently, however, obtain clearer notions on the subject of *unpolarized* light, and see reason for discarding the term altogether.

Part IV.
847.
Partial polarization.

How conceived.

If a ray be partially polarized by reflexion, Dr. Brewster has stated that a second reflexion *in the same plane* renders this polarization more complete, or diminishes the ratio of the unpolarized light in the reflected beam; and that by repeating the reflexion, the ray may be completely polarized, although none of the angles of reflexion be the polarizing angle. Thus he found, that one reflexion from glass at $56^{\circ} 45'$ of incidence, two at incidences of $62^{\circ} 30'$ or at $50^{\circ} 20'$, three at $65^{\circ} 33'$ or at $46^{\circ} 30'$, four at $67^{\circ} 33'$ or $43^{\circ} 51'$, and so on, alike sufficed to operate the complete polarization of the ray finally reflected, provided all the reflexions were made in one plane. At angles above 82° , or below 18° , more than 100 reflexions were required to produce complete polarization.

848.
Polarization by several reflexions in one plane.

§ IV. Of the Laws of Reflexion of Polarized Light.

When polarized light is reflected at any surface, transparent or otherwise, the *direction* of the reflected portion is precisely the same as in the case of natural light, the angle of reflexion being equal to that of incidence; the laws we are now to consider are those of the intensity of the reflected light, and of the nature of its polarization after reflexion.

849.

One essential character of a polarized ray is, its insusceptibility of reflexion in a plane at right angles to that of its polarization when incident at a particular angle, *viz.* the polarizing angle of the reflecting surface. In this case, the intensity I of the reflected ray is 0. In all other cases it has a certain value, which we are now to inquire. Let us suppose, then, to begin with the simplest case, that the polarized ray falls on the reflecting surface at a constant angle of incidence, equal to its polarizing angle, and that the reflecting surface is turned round the incident ray as an axis, so that the plane of incidence shall make an angle ($= \alpha$) of any variable magnitude with the plane of polarization. It is then observed, as we have seen, that when $\alpha = 90^{\circ}$ or 270° , we have $I = 0$, and when $\alpha = 0^{\circ}$, or 180° , I is a maximum. Hence, it is clear that I is a periodic function of α , and the simplest form which can be assigned to it (since negative values are inadmissible) is $I = A \cdot (\cos \alpha)^2$. This value, which was adopted by Malus on no other grounds than those we have stated, is however found to represent the variation of intensity throughout the quadrant, with as much precision as the nature of photometrical experiments admits, and we must therefore receive it as an empirical law at present, for which any good theory of polarization ought to be capable of assigning a reason *a priori*.

850.
Intensity of reflection of a polarized ray incident at the polarizing angle in any plane.

A remarkable consequence follows from this law. It is that, so far as the intensity of the reflected ray is concerned, an ordinary or unpolarized ray may be regarded as composed of two polarized rays, of equal intensity, having their planes of polarization at right angles to each other. For such a compound ray being incident on a reflecting surface, as above supposed, if α be the inclination of the plane of polarization of one portion to that of incidence, $90 - \alpha$ will be that of the other, and, therefore, since

851.
An unpolarized ray equivalent to two polarized ones.

$$A \cdot (\cos \alpha)^2 + A \cdot (\cos . 90 - \alpha)^2 = A, \quad (a)$$

the reflected ray will be independent of α , and therefore no variation of intensity will be perceived on turning the reflecting surface round the incident ray as an axis, which is the distinguishing character of unpolarized light. Any such pair of rays as here described are said to be *oppositely polarized*.

When the polarized ray is not incident at the polarizing angle, but at any angle of incidence, the law of intensity of the reflected ray is more complicated. M. Fresnel has stated the following as the general expression for it. Let the intensity of the incident ray be represented by unity, and calling, as before, α the inclination of the plane of incidence to that of primitive polarization, and i the angle of incidence, i' the corresponding angle of refraction. Then will the intensity of the reflected ray be represented by

852.
Fresnel's general law for the intensity of a reflected ray

Light.

$$I = \frac{\sin^2 (i - i')}{\sin^2 (i + i')} \cdot \cos^2 a + \frac{\tan^2 (i - i')}{\tan^2 (i + i')} \cdot \sin^2 a. \quad (b)$$

Part IV.

This formula is in some degree empirical, resulting partly from theoretical views, of which more hereafter, and being not yet verified, or indeed compared with experiment, except in particular cases, by M. Arago, whose results, so far as they go, are consonant with it.

853.
Particular
cases
examined.

It will be well to examine some of these. And first, then, when $a = 90^\circ$, and $i =$ the polarizing angle of the reflecting surface, we have by (835 and 836) $i + i' = 90^\circ$, and therefore $\tan (i + i') = \infty$, so that $I = 0$. In these circumstances, then, the reflected ray is completely extinguished, which agrees with fact.

854.
Perpendicu-
lar inci-
dence.

2dly. When the incidence is perpendicular, we have, in this case, both i and i' vanishing, and each term of I takes the form $\frac{0}{0}$. Now at the limit we have (μ being the refractive index) $i = \mu \cdot i'$, and very small arcs being equal to their sines or tangents, we have $\sin (i - i') = i' (\mu - 1)$; $\sin (i + i') = i' (\mu + 1)$, and so for the tangents. Consequently,

$$I = \left(\frac{\mu - 1}{\mu + 1} \right)^2 \cdot (\cos^2 a + \sin^2 a) = \left(\frac{\mu - 1}{\mu + 1} \right)^2,$$

which agrees with the expression deduced by Dr. Young and M. Poisson, (Art. 592,) for the intensity of the reflected ray in the case of unpolarized light. And if we regard the unpolarized ray as composed of two rays, each of the same intensity, ($= \frac{1}{2}$) polarized in opposite planes, the reason of the coincidence will be evident.

855. 3d. When $a = 0$, or the plane of polarization coincides with the plane of incidence, we have, in general,

$$I = \frac{\sin^2 (i - i')}{\sin^2 (i + i')}. \quad (c)$$

856. 4th. When $a = 90^\circ$, or when the plane of polarization is at right angles to the plane of incidence,

$$I = \frac{\tan^2 (i - i')}{\tan^2 (i + i')}. \quad (d)$$

857. 5th. When $a = 45^\circ$,

Intensity of
reflexion of
natural light

$$I = \frac{1}{2} \left\{ \frac{\sin^2 (i - i')}{\sin^2 (i + i')} + \frac{\tan^2 (i - i')}{\tan^2 (i + i')} \right\}. \quad (e)$$

This last is the same result with that which would result from the supposition of two equal rays polarized, the one *in*, the other *perpendicularly to*, the plane of incidence, and each of half the intensity with the incident beam. It is therefore the general expression for the intensity of a ray of natural or unpolarized light reflected at an incidence $= i$ from the surface. The expressions in Art. 592 apply only to perpendicular incidences. We are thus furnished very unexpectedly with a solution of one of the most difficult and delicate problems of experimental Optics. Bouguer is the only one who has made any extensive series of photometrical experiments on the intensity of light reflected from polished surfaces at various angles, but his results are declared by M. Arago to be very erroneous, which is not surprising, as the polarization of light was unknown to him, and its laws might affect the circumstances of his experiments in a variety of ways.

858.
Polarization
of the light
of the sky.

One only need be mentioned, as every optical experimentalist ought to be aware of, and on his guard against it, it is that the light of clear, blue sky, is always partially polarized in a plane passing through the sun, and the part from which the light is received. The polarization is most complete in a small circle, having the sun for its pole, and its radius about 78° , (according to an experiment not very carefully made.) Now the semi-supplement of this (which is the polarizing angle) is 51° , which coincides nearly with the polarizing angle of water, ($52^\circ 45'$). Thus strongly corroborating Newton's theory of the blue colour of the sky, which he conceives to be the blue of the first order, reflected from particles of water suspended in the air. Dr. Brewster is the first, we believe, who noticed this curious fact. But to return to our subject.

859.
Case of a
ray partially
polarized.

When the incident ray is only partially polarized, we may regard it as consisting of two portions: the one, which we shall represent by a , completely polarized in a plane, making the angle a with that of incidence; the other $= 1 - a$ in its natural state, or, if we please, composed of two portions $\left(\frac{1 - a}{2} \right)$, one polarized in the plane of incidence, and one at right angles to it. The intensity of the reflected portion of the former is equal to

$$a \cdot \frac{\sin^2 (i - i')}{\sin^2 (i + i')} \cos^2 a + a \cdot \frac{\tan^2 (i - i')}{\tan^2 (i + i')} \cdot \sin^2 a,$$

and that of the latter will be represented by

$$\frac{1 - a}{2} \left\{ \frac{\sin^2 (i - i')}{\sin^2 (i + i')} + \frac{\tan^2 (i - i')}{\tan^2 (i + i')} \right\}$$

therefore, their sum, or the total reflected light, will be

$$\frac{\sin^2 (i - i')}{\sin^2 (i + i')} \cdot \frac{1 + a \cdot \cos 2a}{2} + \frac{\tan^2 (i - i')}{\tan^2 (i + i')} \cdot \frac{1 - a \cdot \cos 2a}{2}.$$

The above formulæ, it must be observed, apply only to the case of reflexion from the surfaces of uncrystallized media. The consideration of those where crystallized surfaces are concerned, cannot be introduced in this part of the subject.

Light.

When the plane of reflexion coincides with that of the primitive polarization of the ray, the polarization is not changed by reflexion. Hence, at a perpendicular incidence it is unchanged. But in other relative situations of the two planes above-mentioned, the case is different, and it becomes necessary to inquire what change reflexion produces in the state and plane of polarization of the ray. Now it is found, as we have already seen, that when the reflection takes place in the plane of primitive polarization, if the incident ray be only partially polarized, the reflected one will be more so, in that plane. But if the incident ray be completely polarized, it retains this character after reflexion, (except in one remarkable case,) and only the *plane* of polarization is changed. Now, according to M. Fresnel, the new plane of polarization will make an angle with the plane of reflexion, represented by β , such that

$$\tan \beta = \frac{\cos (i + i')}{\cos (i - i')} \cdot \tan \alpha.$$

According to this formula, the plane of polarization coincides with the plane of incidence when $i + i' = 90^\circ$. Now this is precisely the case when the ray falls at the polarizing angle on the reflecting surface. If $\alpha = 90^\circ$, or the ray before incidence be polarized in a plane perpendicular to the plane of incidence, it will continue to be so after reflexion, since in that case we have $\tan \beta = \infty$, or $\beta = 90^\circ$.

The formula has been compared by M. Arago with experiment only in one intermediate case, viz. when $\alpha = 45^\circ$, and the coincidence of the results with experiment at a great variety of incidences, and over a range of values of β from $+38^\circ$ to -44° , both in the case of glass and water, is as satisfactory as can be desired. The particulars of this interesting comparison will be found in *Annales de Chimie*, xvii. p. 314. It may be observed also, that these results of M. Fresnel support one another, the latter being concluded from the former by considerations purely theoretical, so that every verification of the one is also a verification of the other.

When the polarized ray is reflected from a crystallized surface, the intensity of the reflected portion is no longer the same, but depends on the laws of double refraction, in a manner of which more hereafter. Whether, or how far, the laws above stated hold good for metallic surfaces, remains open to inquiry.

Part IV.

860.

Position of the plane of polarization of the reflected ray.

861.

862.

Reflexion from crystallized surfaces.

§ V. Of the Polarization of Light by ordinary Refraction, and of the Laws of the Refraction of Polarized Light.

When a ray of natural or unpolarized light is transmitted through a plate of glass at a perpendicular incidence, it exhibits at its emergence no signs of polarization; but if the plate be inclined to the incident ray, the transmitted ray is found to be partially polarized in a plane at right angles to the plane of refraction, and therefore at right angles to the plane of polarization of the portion of the reflected ray which has undergone that modification. The connection between the polarized portions of the reflected and refracted pencils is, however, still more intimate, since M. Arago has shown by a very elegant and ingenious experiment that these portions are always of equal intensity. This law may be stated thus: *When an unpolarized ray is partly reflected at, and partly transmitted through, a transparent surface, the reflected and transmitted pencils contain equal quantities of polarized light, and their planes of polarization are at right angles to each other.*

863.

Polarization by refraction.

Arago's law.

Hence it appears, that the transmitted ray contains a maximum of polarized light, when the light is incident at the polarizing angle of the medium, and this maximum is equal to the quantity of light the surface is capable of completely polarizing by reflexion. Now in all media known, this is much less than half the incident light, consequently the transmitted portion can never be wholly polarized by a single transmission.

864.

When a ray is totally reflected at the inner surface of a medium, there is no transmitted portion, and it is a remarkable coincidence with the above law, that in this case the reflected beam contains no polarized portion whatever.

865

With regard to the portion of light which has passed through the surface, and has not acquired polarization, M. Arago maintains that it remains in the state of natural or totally unpolarized light. Dr. Brewster, on the other hand, concludes from his experiments, that, although not polarized, it has undergone a physical change, rendering it more largely susceptible of polarization by subsequent transmission at the same angle. The question, in a theoretical point of view, is a material one, and *apparently* very easily decided.* The facility, however, is only apparent, and as we have no title to decide it on the grounds of our own experience, we shall content ourselves with reasoning on the conclusions to which the two doctrines lead. Let I be the light incident on the first surface of a glass plate at the polarizing angle, and, after transmission through both surfaces, let $a + b$ be the intensity of the transmitted beam, (and of course $1 - a - b$ that of the reflected,) and let a be the polarized portion and b the unpolarized. When $a + b$ falls on another plate at the same angle, the portion a being polarized in a plane perpendicular to that of incidence, and incident at the polarizing angle, will be totally transmitted, and its *plane of polarization* (as may be proved by direct experiment) *in this case undergoes no change*. Hence the portion a will be transmitted (supposing no absorption) undiminished through any number of subsequent plates. With regard to the portion b , if this be to all intents and purposes similar to natural light, it will be divided by reflexion at the second plate into two portions, the first of which $= b \cdot (1 - a - b)$ being reflected wholly polarized, and the other $= b(a + b)$ will be transmitted. Of this, the portion $b a$ will be polarized in a plane at right angles to that of refraction, and will therefore be afterwards transmitted undiminished through all the subsequent plates. But the portion b^2 will be unpolarized light, and will be again divided by the third plate, and so on. Thus, there will be ultimately transmitted a pencil, consisting of a polarized portion

866.

Polarization by several oblique transmissions.

Light. $= a + b a + b^2 a + \dots b^{n-1} a = a \cdot \frac{1 - b^n}{1 - b}$, and an unpolarized portion $= b^n$, so that no finite number of

plates could ever *completely* polarize the whole transmitted beam.

867. On the other hand, if the unpolarized portion b of the transmitted beam $a + b$ be more disposed than before, as Dr. Brewster conceives, to subsequent polarization, the progression above stated, instead of converging according to the law of a geometric progression, will converge more rapidly, or may even suddenly terminate under certain physical conditions. Now, Dr. Brewster states it as a general law, deduced from his own experiments, that *If a pencil of light be incident on a number of uncrystallized plates, inclined at the same or different angles, but all their surfaces being perpendicular to the plane of the first incidence, the total polarization of the transmitted pencil will commence when the sum of the tangents of the angles of incidence on each plate is equal to a certain "constant quantity due to the refractive power of the plates, and the intensity of the incident pencil."* This last phrase, which makes the number and position of the plates necessary to operate total polarization, depend on the intensity of the incident light, shows evidently that the *total polarization* here understood, is not mathematically, but only approximatively total. In fact, he states, this constant quantity for crown glass plates, and for the flame of a wax candle at 10 feet distance, to be equal to the number 41.84. In other words, the remainder of unpolarized light for this intensity of illumination, becomes insensible. Considered in this light, we regard Dr. Brewster's experiments as by no means incompatible with the law of decrease indicated by the geometric progression above-mentioned and the contrary sense which has been put upon this expression by M. Arago, or his commentator, (*Encyclop. Brit. Supp.*, vol. vi. part 2, *Polarization of Light*,) appears to us strained beyond what strict criticism authorizes.

Conceiving, then, as we do, that no decided incompatibility in matter of fact exists between the statements of these distinguished philosophers, we cannot but regard as most simple, that doctrine which recognises no change of physical character in the unpolarized portion of either the transmitted or reflected beam. (See Art. 848.)

868. In what has been above said of the polarization of the transmitted ray, we have not taken into consideration that part of the light reflected at each surface which is reflected back again, and traversing (partially at least) all the plates, mixes with the transmitted beam, and, being in an opposite plane, destroys a part of its polarization.

869. If a pile of parallel glass plates be exposed to a polarized ray, so that the angle of incidence be equal to the polarizing angle, and then turned round the ray as an axis preserving the same inclination, the following phenomena take place:

1. When the plane of incidence is at right angles to that of the ray's polarization, *the whole of the incident light is transmitted*, (except what is destroyed by absorption within the substance of the glass, or lost by irregular reflexion from the inequalities in the surface arising from defective polish,) and this holds good whatever be the number of the plates. The polarization of the transmitted ray is unaltered.

2. As the pile revolves round the incident ray as an axis, a portion of the light is reflected, and this increases till the plane of incidence is coincident with the plane of primitive polarization, when the reflected light is a maximum. Now, M. Arago assures us, that the quantity of *polarized* light reflected from each plate is greater in proportion to the intensity of the incident beam than if natural light had been employed; and the same proportion holding good at each plate, the transmitted ray, however intense it may have been at first, will be weakened in geometrical progression with the number of plates, and at length will become insensible; so that in this situation the pile will present the phenomenon of an opaque body. In this reasoning, the light reflected backwards and forwards between the plates is neglected; but as it is all polarized in the same plane, and as in this situation the reflexions, however frequent, produce no change in its plane of polarization, all the reflected rays are in the same predicament; and, supposing the number of plates very great, the total extinction of the transmitted light will ultimately (though less rapidly) take place.

870. Hence, a pile of a great number of glass plates inclined at an angle equal to the complement of the polarizing angle ($35^\circ \pm$) to a polarized ray ought to present the same phenomenon with a plate of tourmaline cut parallel to the axis of its primitive rhomboid, alternately transmitting and extinguishing the whole of the light in the successive quadrants of its rotation, and being thus either opaque or transparent, according to its position. The analogy, however, cannot fairly be pushed farther, so as to deduce from this principle an explanation of the phenomena of the tourmaline; for, although it be true that a plate of tourmaline so cut, *is* composed of laminæ inclined to its surface, these laminæ are in optical contact; and, moreover, their position with respect to the surface is not the same in plates cut in all directions around the axis, because although an infinite number of plates may be cut containing *the axis* of a rhomboid in their planes, only three can have the same relation to its several *faces*, parallel to which the component laminæ must be supposed to lie. Moreover, the phenomena are not produced, unless the tourmaline be coloured. The analogy between piles of glass plates and laminæ of agate (of which more presently) is also, we are inclined to think, more apparent than real.

871. A pile of plates such as described above presents, moreover, the same difference of phenomena when exposed to polarized and unpolarized light, that a plate of tourmaline does; since in the latter case, supposing the pile sufficiently numerous, one half the incident light is transmitted, completely polarized in a plane perpendicular to that of incidence.

872. The laws which regulate the polarization of a pencil transmitted by a transparent surface, inclined at any proposed angle to the incident ray, and in any plane to that of its primitive polarization (supposing it polarized) remain open to experimental investigation.

§ VI. Of the Polarization of Light by Double Refraction.

When a ray of natural light is divided into two by double refraction, in such a manner that the two pencils at 873. their final emergence remain distinct and susceptible of separate examination, they are *both found completely polarized, in different planes*, exactly, or nearly, at right angles to each other. To show this, take a pretty thick rhomboid of Iceland spar, and, covering one side of it with a blackened card, or other opaque thin substance, having a small pinhole through it, hold it against the direct light of a window or a candle, with the covered surface from the eye. Two images of the pinhole will then be seen: one, undeviated from the line joining the eye and the real hole, by the ordinarily refracted rays; and the other, deviating from that line, in a plane parallel to the principal section of the surface of incidence, by the extraordinary. These images will appear, to the naked eye, of equal brightness; but, if we interpose a plate of tourmaline, (as already described,) and turn the latter about in its own plane, they will be rendered unequal, and will appear and vanish alternately at every quarter revolution of the tourmaline; the *ordinary* image being always at its maximum of brightness, and the *extraordinary* one extinct, when the axis of the tourmaline plate is perpendicular to the principal section of the surface of incidence, and *vice versa* when parallel to it. Light polarized by double refraction oppositely in the two pencils. Experiments in proof thereof.

The same thing happens, when, instead of examining the two images through a tourmaline plate, we receive 874. their light on a glass plate inclined at the polarizing angle to it, and turn this plate round the ordinary ray as an axis. The images will appear and disappear alternately, as the reflector performs successive quadrants of its revolution. Experiment varied.

Hence, we see that the two pencils are *completely* and oppositely polarized; the ordinary pencil in a plane 875. passing through the axis of the rhomboid; the extraordinary one in a plane at right angles to it.

The same phenomenon is much better seen by using a *prism* of any double refracting crystal, having such a 876. refracting angle as to give two distinctly separated images of a distant object, (as a candle.) These appear and disappear alternately at quarter revolutions of a tourmaline plate or glass reflector, and are of equal brightness at the intermediate half-quarters. Another form of the experiment.

Double refraction, then, polarizes the two refracted pencils oppositely, into which an unpolarized incident ray 877. is separated. Let us now see what happens to a *polarized ray*. For this purpose let a plate of glass be laid down before an open window, so as to polarize the reflected light, and hold the rhomboid of Iceland spar (covered as before) with the covered side from the eye, not (as in the former experiment) against the direct light, but inclined downwards, against the reflected light from the glass. Then, generally speaking, two images of the pinhole will be seen, but of *unequal* intensities; and, *if we turn round the rhomboid, in the plane of the covered side, these images will be seen to vary perpetually in their relative brightness, the one increasing to a maximum, while the other vanishes entirely, and so on reciprocally*. When the principal section of the rhomboid is in the plane of reflexion (*i. e.* of polarization) of the incident ray, the ordinary image is a maximum; the extraordinary is extinct, and *vice versa* when these two planes make a right angle. The experiment may be advantageously varied by using a doubly refracting prism; and, while looking through it at the polarized image of a candle, turning it round slowly in the plane bisecting its refracting angle. Transmission of polarized light through doubly refracting media.

This experiment leads us to the following remarkable law, *viz.* that if a ray, at its incidence on a doubly 878. refracting surface, be polarized in the plane parallel to the principal section, it will not suffer bifurcation, but will pass wholly into the ordinary image; if, on the other hand, its plane of primitive polarization be perpendicular to the principal section, it will pass entirely into the extraordinary image. In intermediate positions of the plane of primitive polarization, bifurcation takes place, and the ray is *unequally* divided between the two refracted pencils, in every case except when the plane of primitive polarization makes an angle of 45° with the principal section. In general, if a be the angle last mentioned, and A the incident light, (supposing none lost by reflexion,) $A \cdot \cos^2 a$ will be the intensity of the ordinary, and $A \cdot \sin^2 a$ of the extraordinary pencil, their sum being A . Unequal division of the light between the two refracted pencils.

All these changes and combinations are exhibited in the following remarkable experiment of Huygens, which, 879. reasoned on by himself and Newton, first gave rise to the conception of a polarity, or distinction of *sides*, in the rays of light when modified by certain processes. Take two pretty thick rhomboids of Iceland spar, (which should be very transparent, as they are easily procured,) and lay them down one upon the other, so as to have their homologous sides parallel, or so that the molecules of each shall have the same relations of situation as if the two rhomboids were contiguous parts of one larger crystal. They should be laid on a sheet of white paper having a small, very distinct, and well-defined black spot on it. This spot then will be seen double through the combined crystals, as if they were one, (*a*, fig. 173,) and the line joining the images will be parallel to the principal section of either. Now, let the upper crystal be turned slowly round in a horizontal plane on the lower, and two new images will make their appearance between the two first seen, which, at first, are very faint, as at *b*, fig. 173, and form a very elongated rhombus with the two former. They increase, however, in intensity, while the other pair diminishes, till the angle of rotation of the upper crystal is 45° , where the appearance of the images is as at *c*. Continuing the rotation, the rhomb approaches to a square, as at *d*, and the two original images have become extremely faint; and when the rotation is just 90° , they will have disappeared altogether, leaving the others diagonally placed, as at *e*. As the rotation still proceeds, they reappear and increase in brightness, till the angle of revolution $= 90^\circ + 45^\circ = 135^\circ$, when the images are all equal, as at *f*; after which the original images still increasing, and the others diminishing, the appearance *g* is produced, which, on the completion of a precise half revolution, passes into *h* by the union of both the original images into one, and the total evanes-

Huygens's experiment

Fig. 13f.

Light.

cence of the other pair. In this case only single refraction (apparently) happens; or, rather, the double refractions of the two rhomboids taking place in opposite directions, and being equal in amount, compensate each other. Unless, however, the rhomboids be of exactly equal thickness, this precise compensation will not take place, and the images will remain distinct, though at a minimum of distance. We may express the four images thus:

O o, the image ordinarily refracted by both rhomboids.

O e, the image refracted ordinarily by the first, and extraordinarily by the second.

E o, the image refracted extraordinarily by the first, and ordinarily by the second.

E e, the image refracted extraordinarily by both.

Then, if A be the intensity of the incident light, supposing none lost by reflexion or absorption,

$$O o = \frac{1}{2} A \cdot \cos^2 a = E e; \quad O e = \frac{1}{2} A \cdot \sin^2 a = E o,$$

and the sum of all the four images = A.

880. The same phenomena (with some unimportant variations) take place when we apply two doubly refracting prisms one behind the other close to the eye, and view a distant object through them, turning one round on the other. The rationale of these phenomena follows so evidently from the laws stated in Art. 875 and 878, that it will not be necessary to enlarge on it.

881. The property of a double refraction, in virtue of which a polarized ray is unequally divided between the two images, furnishes us with a most convenient and useful instrument for the detection of polarization in a beam of light, and for a variety of optical experiments. It is nothing more than a prism of a doubly refracting medium rendered achromatic by one of glass, or still better, by another prism of the same medium properly disposed, so as to increase the separation of the two pencils. The former method is simple; and, when large refracting angles are not wanted, the uncorrected colour in one of the images is so small as not to be troublesome. It is most convenient to make the refracting angle such as to produce an angular separation of about 2° between the images. Thus, in fig. 174, let ABCGF be a prism of Iceland spar, cut in such a manner (we will at present suppose) that the refracting edge CG shall contain the axis of the crystal; and let it be achromatized as much as possible by a prism of glass CDEFG. Then, if Q be a small, colourless, luminous circle of about a degree or two in apparent diameter, as seen by an eye at O, the interposition of the combined prisms will divide it into two, Q and q. Now, if the light of Q be completely unpolarized, these two will remain exactly of equal intensity while the prism ABCG is turned round in a plane at right angles to the line of vision. But if any polarity exist in the original light, the two images Q, q will, in turning round the prism, appear alternately more and less bright one than the other; and being always seen immediately side by side, the least inequality, and consequently the least admixture of polarized light in the incident beam, will be detected.

Fig. 174.
First achromatized by glass.

882. Iceland spar, from its very great double refraction, is commonly used for these prisms; but it is so soft, and its structure so lamellar, as to be difficult to polish, and still more so to preserve polished. We have found quartz and limpid topaz to answer extremely well. The following ingenious mode of rendering available the low double refraction of the former, due to Dr. Wollaston, is here eminently useful. Let ABCDabcd and EFGHefgh (fig. 175) be two halves of a hexagonal prism of quartz (the form it affects) produced by a section parallel to two of the sides. In the vertical face ADda draw any line LK parallel to the sides, and therefore to the axis of the prism, (which is also that of double refraction,) and join CL, ck. Then a plane CLkc will cut off a prism CLKdcD, having Lk, Dd, or Cc, for its refracting edges, either of which is parallel to the axis. Again, in the other half of the prism join Ef and Hg, and cut the prism by a plane passing through these lines; then, regarding either portion as a double refracting prism, having for refracting edges the lines EH, fg, these will have the axis of double refraction perpendicular to their refracting edges; and, in particular, the axis will lie in the faces HEEh, or FGgf at right angles to HE or fg. If, then, we take care to make the refracting angle CLD of the prism CLKdcD equal to that of the edge HE of the prism HEEfgh; and if we make these two prisms act in opposition to each other, placing the edge HE opposite to Dd, and the edge he opposite to KL; and having thus brought the two surfaces DLkd and HEEh in contact, cement them together with mastic, or Canada balsam, it is evident, that their principal sections will be at right angles to each other; and therefore only two images will be formed, the whole of the extraordinary ray of the one prism passing into the ordinary image of the other, and *vice versa*. Now, to see how this acts to double the separation of the images, let us conceive mn to be a luminous line viewed through one of the prisms with its edge downwards and horizontal. It will be separated into two images, e and o, the one more raised than the other. Suppose the ordinary image to be most refracted. Then, if we interpose the other prism with its edge upwards, both these images will be refracted downwards; but the ordinary image o, which was before *most* raised, now undergoing extraordinary refraction, is *least depressed*, and comes into the position oe, while the extraordinary one e, which was before *least raised* is now *most depressed*, and comes into the situation eo; and it is evident that (the refracting angles being equal, and the double refraction of the two prisms the same) the line oe will fall as far short of the original line mn, as eo surpasses it, viz. by a quantity equal to the distance between the two first images o and e; so that the distance between the twice refracted images is double that of those which have undergone only one refraction. We have found this combination extremely advantageous, as quartz takes a very perfect polish, and from its hardness is not liable to injury from scratches.

883.
Action of crystals possessing no double refraction.

Crystals which have no double refraction may be regarded as limits of those which have, or as crystals in which the two rays are propagated with equal velocity, and therefore undergo no bifurcation; or, in other words, in which the images formed coincide. In this case we should expect to find no polarization of the emergent light, because the two pencils, being polarized at right angles to each other, form together a single ray having the characters of unpolarized light. This is verified by experiment. The light transmitted by fluor spar, for

Light. instance, exhibits no signs of polarization, unless so far as the ordinary action of the surface goes. We are aware of no experiments indicating how far the action of the *surfaces* of feebly double refracting crystals may modify their polarizing forces, or rather their effects on a ray which has penetrated below the surface; or, in other words, how far piles of crystallized laminae may have an analogous or different action from those of uncrystallized. Dr. Brewster, indeed, found piles of mica films to polarize light by transmission, like glass piles, but the subject is open to further inquiry. Part IV.

§ VII. *Of the Colours exhibited by Crystallized Plates when exposed to Polarized Light, and of the Polarized Rings which surround their Optic Axes.*

This splendid department of Optics is entirely of modern and, indeed, of recent origin. The first account of the colours of crystallized plates was communicated by M. Arago to the French Institute in 1811, since which period, by the researches of himself, Dr. Brewster, M. Biot, M. Fresnel, and, latterly, also of M. Mitscherlich, and others, it has acquired a development placing it among the most important as well as the most complete and systematic branches of optical knowledge. As might be expected, under such circumstances, as well as from the state of political relations, and the consequent limited intercourse between Britain and the Continent at the period mentioned, an immense variety of results could not but be obtained independently, and simultaneously, or nearly simultaneously, on both sides of the channel. To the lover of knowledge, for its own sake,—the *philosopher*, in the strict original sense of the word,—this ought to be matter of pure congratulation; but to such as are fond of discussing rival claims, and settling points of scientific precedence, such a rapid succession of interesting discoveries must, of course, afford a welcome and ample supply of critical points, the seeds of an abundant harvest of dispute and recrimination. Regarding, as we do, all such discussions, when carried on in a spirit of rivalry or nationality, as utterly derogatory to the interests and dignity of science, and as little short, indeed, of sacrilegious profanation of regions which we have always been accustomed to regard only as a delightful and honourable refuge from the miserable turmoils and contentions of interested life, we shall avoid taking any part in them; and, taking up the subject (to the best of our abilities and knowledge) as it is, and avoiding, as far as possible, all reference to misconceived facts and over-hasty generalizations, which in this as in all other departments of science, have not failed (like mists at daybreak) to spread a temporary obscurity over a subject imperfectly understood, shall make it our aim to state, in as condensed a form as is consistent with distinctness, such general facts and laws as seem well enough established to run no hazard of being upset by further inquiry, however they may merge hereafter in others yet more general;—a consummation devoutly to be wished.

884.

The general phenomenon of the coloured appearances to which this section is devoted, may be most readily and familiarly shown as follows. Place a polished surface of considerable extent (such as a smooth mahogany table, or, what is much better, a pile of ten or a dozen large panes of glass laid horizontally) close to a large open window, from which a full and uninterrupted view of the sky is obtained; and having procured a plate of mica, of moderate thickness, (about a thirtieth of an inch, such as may easily be obtained, being sold in considerable quantity for the manufacture of lanterns,) hold it between the eye and the table, or pile, so as to receive and transmit the light reflected from the latter as nearly as may be judged at the polarizing angle. In this situation of things, nothing remarkable will be perceived, however the plate of mica be inclined; but if instead of the naked eye we look through a tourmaline plate, having its axis vertical, the case will be very different. When the mica plate is away, the tourmaline will destroy the reflected beam, and the surface of the table, or pile, will appear dark and non-reflective; at least in one point, on which we will suppose the eye to be kept steadfastly fixed. No sooner is the mica interposed, however, than the reflective power of the surface appears to be suddenly restored; and on inclining the mica at various angles, and turning it about in its own plane, positions will readily be found in which it becomes illuminated with the most vivid and magnificent colours, which shift their tints at the least change of position of the mica, passing rapidly from the most gorgeous reds to the richest greens, blues, and purples. If the mica plate be held perpendicular to the reflected beam, and turned about in its own plane, two positions will be found in which all colour and light disappears; and the reflected ray is extinguished, as if no mica was interposed. Now, if we draw on the plate with a steel point two lines corresponding to the intersection of the mica with a vertical plane passing through the eye in either of these two positions, we shall find that they make an exact right angle. For the moment, let us call these lines A and B; and let a plane drawn through the line A, perpendicular to the plate, be called the *section A*; and one similarly drawn through the line B, the *section B*. Then we shall observe further, that when we turn the plate from either of these positions, 45° round, in its own plane, so that the sections A and B shall make angles of 45° with the plane of reflexion, (*i. e.* of polarization of the incident ray,) the transmitted light will be a maximum.

885.
First method of exhibiting the colours of crystallized plates. Instantaneous in mica.

Two remarkable sections of the crystallized plate.

If the thickness of the mica do not exceed $\frac{1}{30}$ th of an inch, it will be coloured in this position; if materially greater, colourless; and if less, more and more vividly coloured, and with tints following closely the succession of the reflected series of the colours of thin plates, and, like them, rising in the scale, or approaching the central tint (black) as the thickness is less. The analogy in this respect, in short, is complete, with the exception of the enormous difference of thickness between the mica plate producing the tints in question, and those required to produce the Newtonian rings. It appears by measures made in the manner hereafter to be described, that the tint exhibited by a plate of mica exposed perpendicularly to the reflected ray, as above described, is the same with that reflected by a plate of air of $\frac{1}{100}$ th part of the thickness of the mica employed.

886.

Law of the tints exhibited at a perpendicular incidence.

Light.

887.
Tints exhibited in the two sections above mentioned.

If the mica (still exposed perpendicularly to the ray) be turned round in its own plane, the tint does not change, but only diminishes in intensity as its section A or B approaches the plane of polarization of the incident light. When, however, the plate is not exposed perpendicularly, this invariability no longer obtains; and the changes of tint appear in the last degree capricious and irreducible to regular laws. In two situations, however, the phenomena admit a simple view. These are when the sections A and B are both 45° from the plane of polarization, and the mica plate is inclined backwards and forwards in the plane of one or the other of these sections. This condition is easily attained by first holding the plate perpendicularly to the reflected ray; then turning it in its own plane till the lines A, B are each 45° inclined to the vertical plane, then finally causing it to revolve about either of these lines as an axis. It will then be seen that when made to revolve round one of them (as A) or in the plane of the section B, the tint, if white, will continue white at all angles of inclination; but if coloured, will descend in the scale of the coloured rings, growing continually less highly coloured, till it passes, after more or fewer alternations, into white; after which, further inclination of the plate will produce no change. On the other hand, if made to revolve round B, or in the plane of A, the tints will rise in the scale of the rings; and when the mica plate is inclined either way, so as to make the angle of incidence about $35^\circ 3'$, will have attained its maximum, corresponding to the black spot in the centre of Newton's rings. In this position of the plate, the reflected beam is totally extinguished by the tourmaline, as if the sections A or B had been vertical. But if the angle of incidence be still further increased the colours reappear, and descend again in the scale of the rings, passing through their whole series to final whiteness. We take no notice here of a slight deviation from the strict succession of the Newtonian colours, which is observed in the higher orders of the tints, as we shall have more to say respecting it hereafter.

888.
Characters of the two most remarkable sections.

We see, then, that the sections A and B, though agreeing in their characters in the case of a perpendicular exposure of the mica, yet differ entirely in the phenomena they exhibit at oblique incidences. If the incidence take place in the plane of the section B, the tint descends, on both sides of the perpendicular, *ad infinitum*. While, if the incidence be in the section A, it rises to the central black, which it attains at equal incidences on either side of the perpendicular ($35^\circ 3'$), and then descends again *ad infinitum*, or to the composite white at the other extreme of the scale.

889.
The principal section defined. Contains the two optic axes. Characters of these axes.

The section A, then, (which, for this reason, we will call the *principal section* of the mica plate,) is characterised by containing two remarkable lines inclined at equal angles to the surface of the plate, along either of which, if a polarized ray be incident, its polarization will not be disturbed by the action of the plate. To satisfy ourselves of this, we have only to fix the mica to the extremity of a tube, so as to have the axis of the tube inclined at an angle of $35^\circ 3'$ to the perpendicular (or $54^\circ 57'$ to the plate) in the plane of the section A; then directing the axis of the tube to the centre of the dark spot, or the reflecting surface, it will be seen to continue dark, and remain so while the tube makes a complete revolution on its axis. Now, this could not be if the mica exercised any disturbing power on the plane of polarization. Hence, we conclude, that the two lines in question possess this remarkable property, viz. that whatever be the plane of polarization of a ray incident along either of them, it remains unaltered after transmission. For, although in the experiment above described, the plane of polarization remained fixed, and that of incidence was made to revolve, it is obvious that the reverse process would come to the very same thing.

890. Now, this character belongs to no other lines, however chosen, with respect to the plate. If we fix the plate on the end of the tube at any other angle, or in any other plane with respect to the axis of the latter, although two positions in the rotation of the tube will always be found where the disappearance of the transmitted ray takes place, in no other case but that of the two lines in question will this disappearance be total, or nearly so, in all points of its revolution.

891.
Position of the optic axes in mi

The refracting index of mica being 1.500, an angle of incidence of $35^\circ 3'$ corresponds to one of refraction $= 22^\circ 31'$. Hence, the position of the lines within the mica corresponding to these external lines is $22\frac{1}{2}^\circ$ inclined to the perpendicular, and the angle included between them 45° . These, then, are axes within the crystal, bearing a determinate relation to its molecules. Dr. Brewster has termed them axes of no polarization, a long name. M. Fresnel, and others, have used the phrase optic axes, to which we shall adhere. As this term has before been applied to the "axes of no double refraction," we must anticipate so far as to advertise the reader that these, and the "axes of no polarization," are in all cases identical.

892
The polarized rings about the optic axes. General description of their phenomena. Fig. 176.

Having, by the criteria above described, determined the principal section, and ascertained the situation of the optic axes of the mica plate under examination, let the plate be inclined to the polarized beam, so that the latter shall be transmitted along the optic axes, the principal section A making an angle of 45° with the plane of polarization; and let the eye (still armed with the tourmaline plate, with its axis vertical) be applied close to the mica. A splendid phenomenon will then be seen. The black point corresponding to the direction of the optic axis will be seen to be surrounded with a set of broad, vivid, coloured rings, of an elliptic, or, at least, oval form, divided into two unequal portions by a black band somewhat curved, as represented in fig. 176. This band passes through the pole, or angular situation of the optic axis, about which the rings are formed as a centre. Its convexity is turned towards the direction of the other axis, and on that side the rings are also broader. If, now, the other axis be brought into a similar position, a phenomenon exactly similar will be seen surrounding its place, as a pole. If the mica plate be very thick, these two systems of rings appear wholly detached from, and independent of, each other, and the rings themselves are narrow and close; but if thin (as a 30th or 40th of an inch) the individual rings are much broader, and especially so in the interval between the poles, so as to unite and run together, losing altogether their elliptic appearance, and dilating towards the middle (or in the direction of a perpendicular to the plate) into a broad coloured space, beyond which the rings are no longer formed about each pole separately, but assume the form of reentering curves, embracing and including both poles. Their nature will presently be stated more at large.

Light.

If preserving the same inclination of the mica plate to the visual ray, it be turned about it as an axis, the back band passing through the pole will shift its place, and revolve as it were on the pole as a centre with double the angular velocity, so as to obliterate in succession every part of the rings. When the plate has made 45° of its revolution, so as to bring its principal section into the plane of polarization of the incident beam, this band also coincides in direction with that plane, and is then visibly prolonged, so as to meet that belonging to the set of rings about the other pole; and is crossed at the middle point between the poles by another dark space perpendicular to it, or in the plane of the section B, presenting the appearance in fig. 177.

Fig. 177.

These phenomena, if a tourmaline be not at hand, may be viewed, (somewhat less commodiously, unless the mica plate be of considerable size,) by using in its place the reflector figured in fig. 170, or by a pile of glass plates interposed obliquely between the eye and the mica. In this manner of observing them, the colours are surprisingly vivid, no part of the red and violet rays being absorbed more than the rest; whereas the tourmalines generally exert a considerable absorbing energy on these rays in preference to the rest, and thus the contrast of colours is materially impaired. On the other hand, however, from the greater homogeneity of the transmitted light, the rings are more numerous and better defined; and in this respect the phenomenon is greatly improved by the use of homogeneous light.

894. Other modes of exhibiting these phenomena.

We have taken mica as being a crystallized body very easily obtained of large size, and presenting its axes readily, and without the necessity of artificial sections. It is thus admirably adapted for obtaining a general rough view of the phenomena, preparatory to a nicer examination. From the wide interval between its axes, however, and the considerable breadth of its rings, it is less adapted, when employed as above stated, to give a clear conception of the complicated changes which the rings undergo, on a variation of circumstances. For this reason we shall now describe another and much more commodious mode of examining the systems of polarized rings presented by crystals in general, which has the advantage of bringing the laws of their phenomena so evidently under our eyes as to make their investigation almost a matter of inspection.

895.

It is evident, that when we apply the eye close to, or very near a plate of mica, or other body, and view, beyond it, a considerable extent of illuminated surface, each point of that surface will be seen by means of a ray which has penetrated the plate in a different direction with respect to the axes of its molecules; so that we may consider the eye as in the centre of a spherical surface from all points of which rays are sent to it, modified according to the state of primitive polarization, and the influence of the peculiar energies of the medium, corresponding to the direction in which they traverse it, and the thickness of the plate in that direction.

896. General principle of methods of viewing the rings.

Any means, therefore, by which we can admit into the eye through the plate and tourmaline a cone of rays nearly or completely polarized in one general direction, or according to any regular law, will afford a sight of the rings; and therefore exhibit, at a single view, a synopsis, as it were, of the modifications impressed on an infinite number of rays so polarized traversing the plate in all directions. The property of the tourmaline so often referred to puts it in our power to perform this in a very elegant and convenient manner, by the aid of the little apparatus of which fig. 178 is a section. ABCD is a short cylinder of brass tube, the end of which, AC, is terminated by a brass plate, having an aperture *ab*, into which is set a tourmaline plate cut parallel to the axis: *hgik* is another similar brass cylinder, provided with a similar aperture and a similar tourmaline plate *G*, and fitted into the former so as to allow of the one being freely turned round within the other by the milled edges *BD*, *hk*. A lens *H* of short focus, set in a proper cell, is screwed on in front of the tourmaline *G*, so as to have its focus a little behind its posterior surface, (that next the eye, *O*.) Between the two surfaces *AC*, *gi* is another short cylinder of thin tube *cd*, carrying a brass plate with an aperture somewhat narrower than those in which the tourmalines are set, and on which any crystallized plate *F* to be examined may be cemented with a little wax. This, with the cylinder to which it is fixed, is capable of being turned smoothly round within the cylinder *ABCD* by means of a small pin *e* passing through a slit *f* made in the side, and extended round so as to occupy about 120° of the circumference; by which a rotation to that extent may be communicated to the crystallized plate *F* in its own plane between the tourmaline plates. The pin *e* should screw into the ring *cd*, that it may be easily detached, and admit the ring and plate to be taken out for the convenience of fixing on it other crystals at pleasure.

Periscopic tourmaline apparatus described.

Fig. 178.

The use of the lens *H* is to disperse the incident light, and thus equalize the field of view when illuminated by any source of light, whether natural or artificial, as well as to prevent external objects being distinctly seen through it, which would distract the attention and otherwise interfere with the phenomena. The rays converged by the lens to a focus within the crystallized plate *F*, afterwards diverge and fall on the eye *O*, after traversing the plate in all directions within the limit of the field of view. As by this contrivance they pass through a very small portion of the crystal, there is the less chance of accidental irregularities in its structure disturbing the regular formation of the rings, since we have it in our power to select the most uniform portion of a large crystal. The rays, after passing through the lens, are all polarized by the tourmaline *G*, in planes parallel to its axis; and passing through the eye in this state, if the crystal *F* be not interposed, the rays will, or will not, penetrate the second tourmaline, according as its axis is parallel or perpendicular to that of the first. In consequence, when the cylinder carrying the former is turned round within that carrying the latter, the field of view is seen alternately bright and dark.

897. Mode of action of this apparatus.

When the crystallized substance *F* is interposed, provided it be so disposed that one or other of its optic axes is situated any where in the cone of rays refracted by the lens, so that one of them shall reach the eye by traversing the axis, the polarized rings are seen. If both the axes of the crystal (supposing it to have more than one) fall within the field, a set of rings will be seen round both, and may be studied at leisure. In order to bring the whole of their phenomena distinctly under view, it is requisite to select such crystals as have their axes not much inclined to each other, so as to allow the rings about both to be seen without the necessity of looking very obliquely into the apparatus. In mica the axes are rather too far removed for this. The best crystal we can select for the purpose is nitre.

898. Selection of crystals.

Part IV. Further particulars

Light. Nitre crystallizes in long, six-sided prisms, whose section, perpendicular to their sides, is the regular hexagon. They are generally very much interrupted in their structure; but by turning over a considerable quantity of the ordinary saltpetre of the shops, specimens are readily found which have perfectly transparent portions of some extent. Selecting one of these, cut it with a knife into a plate above a quarter of an inch thick, directly across the axis of the prism, and then grind it down on a broad, wet file, till it is reduced to about $\frac{1}{4}$ th or $\frac{1}{8}$ th inch in thickness; smooth the surfaces on a wet piece of emiered glass, and polish them on a piece of silk strained very tight over a strip of plate glass, and rubbed with a mixture of tallow and colcothar of vitriol. This operation requires practice. It cannot be effected unless the nitre be applied *wet*, and rubbed till quite dry, increasing the rapidity of the friction as the moisture evaporates. It must be performed in gloves, as the vapour from the fingers, as well as the slightest breath, dims the polished surface effectually. With these precautions a perfect vitreous polish is easily obtained. We may here remark, that hardly any two salts can be polished by the same process. Thus, Rochelle salt must be finished wet on the silk, and instantly transferred to soft bibulous linen, and rapidly rubbed dry. Experience alone can teach these peculiarities, and the contrivances (sometimes very strange ones) it is necessary to resort to for the purpose of obtaining good polished sections of soft crystals, especially of those easily soluble in water.

899. Nitre, method of preparing and polishing it.

900. The nitre thus polished on both its surfaces (which should be brought as near as possible to exact parallelism) is to be placed on the plate at F; and the tourmaline plates being then brought to have their axes at right angles to each other (which position should be marked by an index line on the cylinders) the eye applied at O, and the whole held up to a clear light, a double system of interrupted rings of the utmost neatness and beauty will be seen, as represented in fig. 179. If the crystallized plate be made to revolve in its own plane between the tourmalines (which both remain unmoved) the phenomena pass through a certain series of changes periodically, returning, at every 90° of rotation, to their original state. Fig. 180 represents their appearance when the rotation is just commenced; fig. 181, when the angle of rotation is $22\frac{1}{2}^\circ$, or $67\frac{1}{2}^\circ$; and fig. 182, when it equals 45° . When the tourmalines are also made to revolve on each other, other more complicated appearances are produced, of which more presently. We shall now, however, suppose them retained in the situation above mentioned, *i. e.* with their axes crossed at right angles, and proceed to study the following particulars:

1. The form and situation of the rings.
2. Their magnitudes in the same and different plates.
3. Their colours.
4. The intensity of the illumination in different parts of their periphery.

901. Situation of the axes in the crystal. The situation of the rings is determined by the position of the principal section of the crystal, or by that of the optic axes within its substance. These in nitre lie in a plane parallel to the axis of the prisms, and perpendicular to one or other of its sides. It is no unusual thing to find crystals of this salt whose transverse section consists of distinct portions, in which the principal sections make angles of 60° with each other; indicating a composite or maced structure in the crystal itself. These portions are divided from each other by thin films, which exhibit the most singular phenomena by internal reflexion, on which this is not the place to enlarge. In an uninterrupted portion, however, the forms of the rings are as represented in the figures above referred to, their poles subtending at the eye an angle of about 8° . Now, it is to be remarked, that as the plate is turned round between the tourmalines, although the black hyperbolic curves passing through the poles shift their places upon the coloured lines, and in succession obliterate every part of them; forming, first, the black cross in fig. 179, by their union; then breaking up and separating laterally, as in fig. 180, and so on. Yet the rings themselves retain the same form and disposition about their poles, and, except in point of intensity, remain perfectly unaltered; their whole system turning uniformly round as the crystallized plate revolves, so as to preserve the same relations to the axes of its molecules. Hence we conclude, that the coloured rings are related to the optic axes of the crystal, according to laws dependent only on the nature of the crystal, and not at all on external circumstances, such as the plane of polarization of the incident light, &c.

902. Form of the rings, that of lemniscates. Fig. 183. The general form of the rings, abstraction made of the black cross, is as represented in fig. 183. If we regard them all as varieties of one and the same geometrical curve, arising from the variation of a parameter in its equation, it will be evident that this equation must, in its most general form, represent a re-entering symmetrical oval, which at first is uniformly concave, and surrounds both poles, as A; then flattens at the sides, and acquires points of contrary flexure, as B; then acquires a multiple point, as C; after which it breaks into two conjugate ovals DD, each surrounding one pole. This variation of form, as well as the general figure of the curves, bears a perfect resemblance to what obtains in the curve well known to geometers under the name of the *lemniscate*, whose general equation is

$$(x^2 + y^2 + a^2)^2 = a^2(b^2 + 4x^2),$$

when the parameter b gradually diminishes from infinity to zero; $2a$ representing the constant distance between the poles.

903. Verified by experiment. The apparatus just described affords a ready and very accurate method of comparing the real form of the rings with this or any other proposed hypothesis. If fixed against an opening in the shutter of a darkened room, with the lens H outwards, and a beam of solar light be thrown on the latter, parallel to the axis of the apparatus, the whole system of rings will be seen finely projected against a screen held at a moderate distance from E. Now, if this screen be of good smooth paper tightly stretched on a frame, the outlines of the several rings may easily be traced with a pencil on it, and the poles being in like manner marked, we have a faithful representation of the rings, which may be compared at leisure with a system of lemniscates, or any other curve graphically constructed, so as to pass through points in them chosen where the tint is most decided. This has

Light

accordingly been done, and it has been found that lemniscates so constructed coincide throughout their whole extent, to minute precision, with the outlines of the rings so traced, the points graphically laid down falling on the pencilled outlines. The graphical construction of these curves is rendered easy by the well-known property of the lemniscate, in which the rectangle under two lines $PA \times P'A$ drawn from the poles to any point A in the periphery is invariable throughout the whole curve. This is easily shown from the above equation, and the value of this constant rectangle in any one curve is represented by $a \times b$.

Part IV.

When we shift from one ring to another, a remains the same, because the poles are the same for all. To determine the variation of b , let the rings be illuminated with homogeneous light, (or viewed through a red glass,) and outlined by projection, as above. Then, if we determine the actual value of ab by measuring the lengths of two lines $PA, P'A$ drawn from P, P' to any point of each curve; and, calculating their product, (to which ab is equal,) it will be found that this product, and therefore the parameter b , increases in the arithmetical progression 0, 1, 2, 3, 4, &c. for the several dark intervals of the rings beginning at the pole, and in the progression $\frac{1}{2}, \frac{3}{2}, \frac{5}{2}$, &c. for the brightest intermediate spaces. To ensure accuracy, the mean of a number of values of $PA \times P'A$, at different points of the periphery, may be taken to obviate the effect of any imperfection in the crystal.

904.
Variation of the parameter b in arithmetic progression from ring to ring.

This, then, is the law of the magnitudes of the successive rings formed by one and the same plate. But if we determine the value of the same product for plates of nitre similarly cut, but of different thicknesses, or of the same reduced in thickness by grinding, it will be found to vary inversely as the thickness of the plate, *cæteris paribus*.

905.
Effect of varying the thickness of the plate.

The colours of the polarized rings bear a great analogy to those reflected by thin plates of air, and in most crystals would be precisely similar to them but for a cause presently to be noticed. In the situation of the tourmaline plates here supposed (crossed at right angles) they are those of the reflected rings, beginning with a black centre, at the pole. If examined in the situation of fig. 179, and traced in a line from either pole cutting across the whole system, at right angles to the line joining the poles, they will almost precisely follow the Newtonian scale of tints. For the present we will suppose that they do so in all directions. It is evident, then, that each particular tint (as the bright green of the third order, for instance) will be disposed in the form of a lemniscate, and will have its own particular value of the product ab . The tint, then, may be said to be corresponding to,—dependent on,—or, if we will, *measured by* ab . In conformity with this language the coloured curves have been termed, and not inaptly, *isochromatic lines*. Now, in the colours of thin plates, we have seen that these tints arise from a law of periodicity to which each homogeneous ray is subject; and that (without entering at this moment into the cause of such periods) the successive maxima and minima of each particular coloured ray passed through, in the scale of tints, correspond to successive multiples by $\frac{1}{2}, \frac{3}{2}, \frac{5}{2}, \frac{7}{2}$, &c. of the period peculiar to that colour. In the colours of thin plates, the quantity which determines the number of periods is the thickness of the plate of air, or other medium traversed; and the number of times a certain standard thickness peculiar to each ray is contained therein, determines the number of periods, or parts of a period, passed through. In the colours, and in the case now under consideration, the number of periods is proportional to the product $(\theta \times \theta')$ of the distances from either pole, for one and the same thickness of plate,—and for different plates to t the thickness,—and, therefore, generally, to $\theta \times \theta' \times t$, provided we neglect the effect of the inclination of the ray in increasing the length of the path of the rays within the crystal, or regard the whole system of rings as confined within very narrow limits of incidence.

906.
The colours of the rings.

Numerical measure of the tint of any isochromatic line.

Law of periodicity.

This condition obtains in the case here considered, because of the proximity of the axes in nitre to each other and to the perpendicular to the surfaces of the plate. But in crystals such as mica, or others where they are still wider asunder, it is not so; and the projection of the isochromatic curves on a plane surface will deviate materially from their true form, which ought to be regarded as delineated on a sphere having the eye, or rather a point within the crystal, for a centre. In such a case, it might be expected that the usual transition from the arc to its sine should take place; and that, instead of supposing the tint, or value of ab , to be proportional simply to $\theta \times \theta' \times t$, (putting $\theta = AP$, and $\theta' = AP'$.) we ought to have it proportional to $\sin \theta \times \sin \theta' \times$ length of the path of the ray within the crystal. Now (putting ρ for the angle of refraction, and t for the thickness of the plate) we have $t \cdot \sec \rho =$ length of the ray's path within the crystal. If, then, we put n for

907.
Transition from nitre to other crystals whose axes are farther asunder.

the number of periods corresponding to the tint ab for the ray in question, and suppose $h = \frac{ab}{n}$, or the unit whose multiples determine the order of the rings, we shall have

$$n = \frac{ab}{h} = \frac{t}{h} \cdot \sin \theta \cdot \sin \theta' \cdot \sec \rho, \quad (a)$$

General expression for the tint polarized by any crystallized plate.

and

$$h = \frac{t}{n \cdot \cos \rho} \cdot \sin \theta \cdot \sin \theta'. \quad (b)$$

If, then, the suppositions made be correct, we ought to have the function on the right hand side of this last equation invariable, in whatever direction the ray penetrates the crystallized plate, and whatever be the order of the tint denoted by n . We shall here relate only one experiment, to show how very precisely the agreement of this conclusion with fact is sustained.

A ray of light was polarized by reflexion at a plate of perfectly plané glass, and transmitted through a plate of mica, having its principal section 45° inclined to the plane of primitive polarization, and the mica plate made to revolve in the plane of its principal section about an axis at right angles thereto, (or about the axis B, Art. 885.) In this state of things, if viewed through a tourmaline as above described, or by other more refined

908.
Experiment verifying this law.

Light.

Part IV.

means presently to be noticed, the succession of tints exhibited by the mica was that of a section of the rings in fig. 182, made by a line drawn through both the poles. To render the observation definite, a red glass was interposed so as to reduce the rings to a succession of red and black bands, and the angles of incidence corresponding to the maxima and minima of the several rings very accurately measured. These are set down in Col. 2 of the following table. Col. 1 contains the values of n , 0 corresponding to the pole, $\frac{1}{2}$ to the first maximum, 1 to the first minimum, $1\frac{1}{2}$ to the second maximum, and so on. The third column contains the angles of refraction computed for an index 1.560; the fourth and fifth, those of θ and θ' ; the sixth, those of h deduced from the above equation, and which ought to be constant. The excesses above the mean are stated in the last column, and show how very closely that equation represents the fact. The thickness of the mica was 0.023078 inches = t .

Values of n .	Angles of incidence.	Angles of refraction = ρ .	Values of θ .	Values of θ' .	Values of h .	Excesses above the mean.
0.0	35° 3' 30"	22° 31' 0"	0° 0' 0"	45° 2' 0"		
0.5	32 55 20	21 14 40	1 16 20	43 45 40	0.032952	- 0.000195
1.0	30 34 40	19 49 30	2 41 30	42 20 30	0.033622	+ 0.000475
1.5	28 15 40	18 24 0	4 7 0	40 55 0	0.033035	- 0.000112
2.0	25 34 20	16 43 30	5 47 30	39 14 30	0.033327	+ 0.000180
2.5	22 46 20	14 57 15	7 33 45	37 28 15	0.033148	+ 0.000001
3.0	19 35 40	12 55 10	9 35 50	35 26 10	0.033058	- 0.000089
3.5	15 48 40	10 27 50	12 3 10	32 58 50	0.033026	- 0.000121
4.0	10 48 50	7 11 10	15 19 53	29 42 10	0.033010	- 0.000137

909.
General
establish-
ment of the
law.

Proceeding thus, and measuring across the system of rings in all directions for plates of various crystals and of all thicknesses, it has been ascertained, as a general fact, that in all substances which possess the property of developing periodical colours by exposure to polarized light in the manner described, the tint (n), or rather the number of periods and parts of a period corresponding, in the case of a ray of given refrangibility, to a thickness t , an angle of refraction ρ , and a position within the crystal, making angles θ and θ' with the optic axes, is represented by the equation

$$n = \frac{t \cdot \sec \rho}{h} \times \sin \theta \cdot \sin \theta',$$

Case of a
crystal
formed into
a sphere.

h being a constant depending only on the nature of the crystal and the ray. Were the crystal of a spherical form, instead of a parallel plate, $t \cdot \sec \rho$, which represents the path traversed by the ray within it, must be replaced by a constant equal to the diameter of the sphere, and in that case the tint would be simply proportional to the product of the sines of θ and θ' . This elegant law is due to M. Biot, though it is to Dr. Brewster's indefatigable and widely extended research that we owe the general development of the splendid phenomena of the polarized rings in biaxial crystals. It appears, then, from this, that if, on the surface of a sphere formed of any crystal, curves analogous to the lemniscate, or having $\sin \theta \times \sin \theta'$ constant for each curve, and varying in arithmetical progression from curve to curve, be described,—then, if the sphere be turned about its centre in a polarized beam, as above described, the tint polarized at every point of each curve will be the same, and in passing from curve to curve will obey the law of periodicity proper to the crystal.

910.
Methods of
viewing the
rings at
great obli-
quities.

There is hardly any character in which crystals differ more widely than in the angular separation of their optic axes, as the table annexed to the end of this article will show. This, while it affords most valuable criteria to the chemist and mineralogist, in discriminating substances and pointing out differences of structure and composition which would otherwise have passed unnoticed, renders the investigation of their phenomena difficult, since it is frequently impossible, by any contrivance, to bring both the axes under view at once; and necessitates a variety of artifices to obtain a sight of the rings about both. It is often very easy to cut and polish crystallized bodies in some directions, and very difficult in others. However, by immersing plates of them in oil, and turning them round on different axes, or by cementing on their opposite sides prisms of equal refracting angles oppositely placed, as in fig. 184, we may look through them at much greater obliquities than without such aid; and thus, by increasing the range of vision to nearly a hemisphere, avoid in most instances the necessity of cutting them in different directions.

Fig. 184.

911.
Rings in
crystals
with one
axis.
Fig. 185.

When the two axes coalesce, or the crystal becomes uniaxial, the lemniscates become circles; and the black hyperbolic lines, passing through the poles, resolve themselves into straight lines at right angles to each other, forming a black cross passing through the centre of the rings, as in fig. 185. In this case the tint is represented by $t \cdot \sin \theta^2$; and in the case of plates, where t , the thickness, is considerable, or where, from the otherwise peculiar nature of the substance the rings are of small dimensions, θ is small, and therefore proportional to its sine; so that in passing from ring to ring θ^2 increases in arithmetical progression. Hence the diameters of the rings are as the square roots of the numbers 0, 1, 2, 3, &c.; and therefore their system is similar, with the exception of the black cross, to the rings seen between object-glasses. Carbonate of lime cut into a plate at right angles to the axis of its primitive rhomboid, exhibits this phenomenon with the utmost beauty. The most familiar instance, however, may be found in a sheet of clear ice about an inch thick frozen in still weather. A pane of window-glass, or a polished table to polarize the light, a sheet of ice freshly taken up in winter

Light.

produce the rings, and a broken fragment of plate glass to place near the eye as a reflector, are all the apparatus required to produce one of the most splendid of optical exhibitions.

If θ be not very small, the measure of the tint, instead of $t \cdot \sin \theta^2$, is $t \cdot \sec \rho \cdot \sin \theta^2$. We have seen that in uniaxal crystals, $\sin \theta^2$ is proportional to the difference of the squares of the velocities v and v' of the ordinary and extraordinary ray, or to $v^2 - v'^2$. Now, if we denote by τ and τ' the times taken by these two rays to traverse the plate, we have $v = \frac{t \cdot \sec \rho}{\tau}$ and $v' = \frac{t \cdot \sec \rho}{\tau'}$; therefore $t \cdot \sec \rho \cdot \sin \theta^2$ is proportional to

$$(t \cdot \sec \rho)^3 \times \left(\frac{1}{\tau'^2} - \frac{1}{\tau^2} \right), \text{ that is, to } \frac{(\tau + \tau')(\tau - \tau')}{(\tau \tau')^2} \cdot (t \sec \rho)^3,$$

or (which is the same thing) to $(v + v') \cdot v v' (\tau - \tau')$. But, neglecting the squares of very small quantities, of the order $v' - v$ and $\tau - \tau'$, for such they are in the immediate neighbourhood of the axis, the factors $v + v'$ and $v v'$ are constant; so that the tint is simply proportional to $\tau - \tau'$, the difference of times occupied by the two rays in traversing the plate; or the *interval of retardation* of the slower ray on the quicker. This very remarkable analogy between the tints in question and those arising from the law of interferences, was first perceived by Dr. Young; and, assisted by a property of polarized light soon to be mentioned, discovered by Messrs. Arago and Fresnel, leads to a simple and beautiful explanation of all the phenomena which form the subject of this section, and of which more in its proper place.

The forms of the rings are such as we have described, only in regular and perfect crystals; every thing which disturbs this regularity, distorts their form. Some crystals are very liable to such disturbances, either arising from an imperfect state of equilibrium, or a state of strain in which the molecules are retained, or to actual interruptions in their structure. Thus, specimens of quartz and beryl are occasionally met with, in which the single axis usually seen is very distinctly separated into two, the rings instead of circles have oval forms, and the black cross (which in cases of a well developed single axis remains quite unchanged during the rotation of the crystallized plate in its own plane) breaks into curves convex towards each other, but almost in contact at their vertices, at every quarter revolution. Cases of interruption occur in carbonate of lime very commonly, and in muriacite perpetually; and the effects produced by them on the configurations of the rings rank among the most curious and beautiful of optical phenomena. They have not, however, been anywhere described, and our limits will not allow us to make this article a vehicle for their description.

The form of the rings being, then, considered, let us next inquire more minutely into their colours. These being all composite, and arising from the superposition on each other of systems of rings formed by each homogeneous ray, we can obtain a knowledge of their constitution only by examining the rings in homogeneous light. This is easy, for we have only to illuminate the apparatus described above by homogeneous light of all degrees of refrangibility from red to violet, by passing a prismatic spectrum from one end to the other over the illuminating lens H, the eye being applied as usual at O, and observe the changes which take place in the rings, in passing from one coloured illumination to another; and, if necessary, measure their dimensions. This is readily done, either by projecting them on a screen in a darkened room, as described in Art. 903, or by detaching the lens H, fig. 178, and simply looking through the apparatus at a sheet of white paper strongly illuminated by the rays of a prismatic spectrum, where the rings will appear as if depicted on the paper, and their outlines easily marked, or their diameters measured. The following are the general facts which may thus be readily verified.

First, in the case of crystals with a single axis, the rings remain circular, and their centres are coincident for all the coloured rays, but their dimensions vary. In the generality of such crystals, their diameters for different refrangibilities follow nearly the law of the Newtonian rings, when viewed in similar illuminations; their squares (or rather the squares of their sines) being proportional, or nearly so, to the lengths of the fits, or of the undulations of the rays forming them. This law, however, is very far from universal; and in certain crystals is altogether subverted. Thus, in the most common variety of apophyllite, (from Cipit, in the Tyrol,—not from Passa, as is commonly stated,) the diameters of the rings are nearly alike for all colours, those of the green rings being a very little less; those formed by rays at the confines of the blue and indigo exactly equal, and those of violet rays a little greater than the red rings. It is obvious, that were the rings of all colours exactly equal, the system resulting from their superposition would be simple alternations of perfect black and white, continued *ad infinitum*. In the case in question, so near an approach to equality subsists, that the rings in a tourmaline apparatus appear merely black and white, and are extremely numerous, no less than thirty-five having been counted, and many of those too close for counting being visible in a thick specimen.

When examined more delicately, colours are, however, distinguished, and are in perfect conformity with the law stated, being for the first four orders as follow:

First order. Black, greenish white, bright white, purplish white, sombre violet blue.

Second order. Violet almost black, pale yellow green, greenish white, white, purplish white, obscure indigo inclining to purple.

Third order. Sombre violet, tolerable yellow green, yellowish white, white, pale purple, sombre indigo.

Fourth order. Sombre violet, livid grey, yellow green, pale yellowish white, white, purple, very sombre indigo, &c.

Carbonate of lime, beryl, ice, and tourmaline (when limpid) are instances of uniaxal crystals, in whose rings the Newtonian scale of tints is almost exactly imitated; and, consequently, the intervals of retardation of the ordinary and extraordinary rays of any colour on one another, are proportional to the lengths of their undulations. On the other hand, in the hypsulphate of lime, we are furnished with an instance of more rapid

Part IV.

912.

Analogy between the colours of the polarized rings and those produced by the law of interference

913

Circumstances which distort the rings.

914.

Colours of the rays.

915.

In crystals with one axis. Deviations from Newton's scale In the apophyllite.

916.

917

Light. degradation of tints, and therefore of a more rapid variation of the interval just mentioned. The following was the scale of colour of the rings observed in this remarkable crystal :

In hyposulphate of lime.

First order. Black, very faint sky blue, pretty strong sky blue, very light bluish white, white, yellowish white, bright straw colour, yellow, orange yellow, fine pink, sombre pink.

Second order. Purple, blue, bright greenish blue, splendid green, light green, greenish white, ruddy white, pink, fine rose red.

Third order. Dull purple, pale blue, green blue, white, pink.

Fourth order. Very pale purple, very light blue, white, almost imperceptible pink.

After which the succession of colours was no longer distinguishable.

918. Other remarkable cases of deviation.

A degradation still more rapid has been observed in certain rare varieties of uniaxal apophyllite, accompanied with remarkable and instructive phenomena. In these, the diameters of the rings (instead of diminishing as the refrangibility of the light of which they are formed increases) increase with great rapidity, and actually become infinite for rays of intermediate refrangibility ; after which they again become finite, and continue to contract up to the violet end of the spectrum, where, however, they are still considerably larger than in the red rays. In consequence of this singularity, their colours when illuminated with white light furnish examples of a complete inversion of Newton's scale of tints. The following were the tints exhibited by two varieties of the mineral in question, in one of which the critical point where the rings become infinite took place in the indigo, and in the other in the yellow rays. In the former they were

First order. Black, sombre red, orange, yellow, green, greenish blue, sombre and dirty blue.

Second order. Dull purple, pink, ruddy pink, pink yellow, pale yellow (almost white,) bluish green, dull pale blue.

Third order. Very dilute purple, pale pink, white, very pale blue.

In the latter variety, the tints were

First and only order. Black, sombre indigo, indigo inclining to purple, pale lilac purple, very pale reddish purple, pale rose red, white, white with a hardly perceptible tinge of green.

919. Relation between the diameters of the rings and the doubly refractive energy.

The doubly refracting energy of a crystal may be not improperly measured by the difference of the squares of the velocities of an ordinary and extraordinary ray similarly situated with respect to the axes ; but as this difference, for rays variously situated in one and the same crystal, is proportional to $\sin^2 \theta$, or in biaxial crystals to $\sin \theta \cdot \sin \theta'$, the intrinsic double refractive energy of any crystal may be represented by

$$e = \frac{v^2 - v'^2}{\sin \theta \cdot \sin \theta'} ; \quad (c)$$

regarding this henceforth as the definition of this energy, we have, in uniaxal crystals, $e = \frac{v^2 - v'^2}{\sin^2 \theta}$, and this will evidently measure the actual amount of separation of two such rays when emergent from the crystal.

If in this we put for v and v' their equals $\frac{t \cdot \sec \rho}{\tau}$ and $\frac{t \cdot \sec \rho}{\tau'}$, we shall have, after reduction,

$$v^2 - v'^2 = v v' (v + v') \cdot \frac{\tau' - \tau}{t \cdot \sec \rho} . \quad (c)$$

In a parallel plate, perpendicular to the axis and in the immediate vicinity of the axis, v' and $\sec \rho$ may be regarded as constant, and $v^2 - v'^2$ is proportional to $\tau' - \tau$, the interval of retardation of one ray on the other, to which the tint in white light and the number of periods and parts of a period in homogeneous light (to which, for brevity, we will continue to extend the term *tint*) are proportional. We see, then, that in such cases the intrinsic double refracting energy is directly as the tint polarized, and inversely as $\sin^2 \theta$, and therefore also inversely as the squares of the diameters of the rings. As the rings increase in magnitude, then, *ceteris paribus*, the double refractive energy diminishes ; and hence a very curious consequence follows, *viz.* that in the two cases last mentioned it vanishes altogether for those colours where the rings are infinite ; in other words, that although the crystal be doubly refractive for all the other coloured rays, there is one particular ray in the spectrum (*viz.* the indigo in the former, and the yellow in the latter case) with respect to which its refraction is single. In the passage through infinity, there is generally a change of sign. In the instances in question this change takes place in the value of e or $v^2 - v'^2$, which passes from negative to positive. And the spheroid of double refraction changes its character accordingly from oblate to prolate, passing through the sphere as its intermediate state. The manner in which this may be recognised, without actually measuring, or even perceiving its double refraction, will be explained further on.

Case of crystals at once attractive, repulsive, and neutral.

920. Application to biaxial crystals.

For crystals with two axes we have only, at present, the ground of analogy to go upon in applying the above formula and phraseology to their phenomena. The general fact of an intimate connection of the double refracting energy with the dimensions of the rings, is indeed easily made out ; for it is a fact easily verified by experiment, that *all crystals, whether with one or two axes, in which the rings or lemniscates formed are of small magnitude in respect of the thickness of the plate producing them, are powerfully double refractive, and vice versa* ; and that, generally speaking, the separation of the ordinary and extraordinary pencils is, *ceteris paribus*, greater in proportion as the rings are more close and crowded round their poles. In uniaxal crystals, in which the laws of double refraction are comparatively simple, there is little difficulty in submitting the point to the test of direct experiment and exact measurement, and it is found to be completely verified. In biaxial, however, such precise and direct comparison is more difficult, and calls for a knowledge of the general laws of double refraction. The analogy, however, supported by the general coincidence above mentioned, is too strong to be refused ; and, as we advance, will be found to gain strength with every step.

Light.

In biaxial crystals, similar deviations from exact proportionality between the lengths of the periods of the several coloured rays and those of their undulations, or fits, exist; but their effect in disturbing the colours of the rings is interfered with, and frequently masked by, another cause, which has no existence in uniaxial crystals, viz. that the optic axes differ in situation, within one and the same crystal for the differently refrangible homogeneous rays; and, therefore, that the elementary lemniscates, whose superposition forms the composite fringes seen in a white illumination, differ not only in magnitude but in the places of their poles and the interval between them. To make this evident to ocular inspection, take a crystal of Rochelle salt, (tartrate of soda and potash,) and having cut it into a plate perpendicular to one of its optic axes, or nearly so, and placed it in a tourmaline apparatus, let the lens H be illuminated with the rays of a prismatic spectrum, in succession, beginning with the red and passing gradually to the violet. The eye being all the time fixed on the rings, they will appear for each colour of perfect regularity of form, remarkably well defined, and contracting rapidly in size as the illumination is made with more refrangible light; but in addition to this, it will be observed, that the whole system appears to shift its place bodily, and advance regularly in one direction as the illumination changes; and if it be alternately altered from red to violet, and back again, the pole, with the rings about it, will also move backwards and forwards, vibrating, as it were, over a considerable space. If homogeneous rays of two colours be thrown at once on the lens, two sets of rings will be seen, having their centres more or less distant, and their magnitudes more or less different, according to the difference of refrangibility of the two species of light employed.

Since the plate in this experiment is supposed to have its surfaces perpendicular to the mean position of the optic axis, the cause of these appearances cannot be found in a mere apparent displacement of the rings by refraction at the surface, existing to a greater extent for the violet than the red rays, add to which, that the angle which their poles describe, is neither the same in magnitude nor direction for different crystals. In some, the optic axes approach each other in violet light, and recede in red; while in others the reverse is the case. In all, however, so far as we are aware, the optic axes for all the coloured rays lie in one plane, viz. the principal section of the crystal. This is rendered matter of inspection by cutting any crystal so that both axes shall be visible in the same plate, and placing it with its principal section in the plane of primitive polarization. In this state of things the first ring about each pole, as in fig. 179, is seen divided into two halves, and puts on, if the plate be pretty thick, the appearance of two semi-elliptic spots, one on each side of the principal section. These spots are observed to be differently coloured at their two extremities: in some crystals the ends of the spots, as well as the segments of the rings adjacent to them, which are turned towards each other, being coloured red, and the other, or more distant ends, with blue; and in others, the reverse. In some crystals this coloration is slight, and in a very few, imperceptible; but in others it is so great, that the spots are drawn out into long spectra, or tails of red, green, and violet light; and the ends of the rings are in like manner distorted and highly coloured, presenting the appearance in fig. 186. This is the case with Rochelle salt, above mentioned. If these spectra be examined with coloured glasses, or with homogeneous light, they will be seen to be composed as in fig. 187, by the superposition of well defined spots of the several simple colours arranged in lines on each side of the principal section. In the case of Rochelle salt, the angular extent of these spectra, within the medium, which measures the interval between the optic axes for violet and red rays, amounts to no less than 10° .

Dr. Brewster has given the following list of crystals presenting these phenomena, which he has divided into two classes, according to his peculiar and ingenious views.

Class I.

Nitre.
Sulphate of baryta.
Sulphate of strontia.
Phosphate of soda.
Tartrate of potash and soda.
Supertartrate of potash and soda.
Arragonite.
Carbonate of lead. (?)
Sulphato-carbonate of lead.

Class II.

Topaz.
Mica.
Anhydrite.
Native borax.
Sulphate of magnesia.

Unclassed.

Chromate of lead.
Muriate of mercury.
Muriate of copper.
Oxynitrate of silver.
Sugar.
Crystallized Cheltenham salts.
Nitrate of mercury.
Nitrate of zinc.
Nitrate of lime.
Soperoxalate of potash.
Oxalic acid.
Sulphate of iron.
Carbonate of lead. (?)
Cymophane.
Felspar
Benzoic acid.
Chromic acid.
Nadelstein (Faröe.)

Part IV.

921.

Separation of the optic axes of differently refrangible rays in biaxial crystals.

922.

All the axes lie in the plane of the principal section.

Fig. 186.

Fig. 187.

923.

Dr. Brewster's list of crystals exhibiting deviations of tint from this cause.

924.

Phenomena of the virtual poles explained.

To which list a great many more might be added. Bicarbonate of ammonia, indeed, is the only biaxial crystal we have examined in which the optic axes for all colours appear to be strictly coincident.

This separation of the axes of different colours explains a remarkable appearance presented by the rings of all biaxial crystals, when placed with their principal section 45° inclined to the plane of polarization of the incident light. It is universally observed that, in traversing the whole system of rings in the plane of the principal

Light.

section, the nearest approximation to Newton's scale of colours is obtained by assuming, for the origin of the scale, not the poles themselves, but other points (which have been called *virtual poles*, though improperly) lying either between or beyond them, according to the crystal examined, and at a distance from them, invariable for each species of crystal, whatever be the thickness of the plate. In consequence, the poles themselves are not absolutely black, but tinged with colour; and their tint descends in the scale as the thickness of the plate increases, and as, in consequence, one, two, or more orders of rings intervene between them and the points from which the scale originates. These points are observed to lie between the poles in all crystals which have the blue axes nearer than the red, such as Rochelle salt, borax, mica, sulphate of magnesia, topaz; and beyond them for those in which the red axes include a less angle than the blue, as sulphate of baryta, nitre, arragonite, sugar, hyposulphite of strontia; and this fact, as well as the constancy of their distance from the poles when the thickness of the plate is varied, renders their origin evident. In fact, since the violet rings are smaller than the red, if the centre about which the former are described, instead of being coincident with that of the latter, be shifted in either direction, carrying its rings with it, some one of the violet rings will necessarily be brought up to, and fall upon a red ring of the same order; and the same holding good with the intermediate rays, provided the law which determines the separation of the different coloured axes be not very different from that which regulates the dimensions of the rings of corresponding colours, the point of coincidence of a red and violet ring of the same order will be nearly that of a red and green, or any intermediate colour. The tint, then, at this point will be either absolutely black, (if they be dark rings which are thus brought to coincidence,) or white, if bright; and from this point the tints will reckon either way with more or less exactness, according to the same scale which would have held good had the points of coincidence been the poles themselves. Should, however, the two laws above mentioned differ very widely, an uncorrected colour will be left at the point of nearest compensation, just as happens when two prisms whose scales of dispersion are dissimilar are employed to achromatise each other. To what an extent the disturbance of the Newtonian scale of tints may be carried by this and the other causes already explained, the reader may see by turning to the table of tints exhibited by Rochelle salt in *Phil. Trans.* 1820, part i.

925.

Two suppositions as to the mode of action of crystals in forming the rings.

Doctrine of polarization considered.

We come next to consider the law of the intensity of the illumination of the rings in different parts of their periphery; but this part of their theory will require us to enter more fundamentally into the mode in which their formation is effected, and to examine what modifications the polarized ray incident on the crystallized plate undergoes in its passage through it, so as to present phenomena so totally different from those which it would have offered without such intervention. It is evident then, first, that since the ray, if not acted on by the plate, would have been entirely stopped by the second tourmaline, but, when so acted on, is partially transmitted so as to exhibit coloured appearances of certain regular forms; that the crystallized plate must have either destroyed altogether the polarization of that part of the light which is thereby enabled to penetrate the second tourmaline, or, if not, must have altered its plane of polarization, so as to allow of a partial transmission. Between these two suppositions it is not difficult to decide. Were the portion of light which passes through the second tourmaline and forms the rings wholly *depolarized*, that is, restored to its original state of natural light, since the remainder, its complement to unity, which continues to be stopped by the tourmaline, retains its state of polarization unaltered, it is evident, that each ray at leaving the crystallized plate would be composed of two portions, one unpolarized ($= A$), the other ($= 1 - A$) polarized. Of these, the half only of the first ($\frac{1}{2} A$) would be transmitted by the second tourmaline. Now, suppose this to be turned round in its own plane through any angle ($= a$) from its original position, then the unpolarized portion will continue to be half transmitted; and the polarized, being now partially also transmitted, (in the ratio of $\sin^2 a : 1$.) will mix with it, so that the compound beam will be represented by

$$\frac{1}{2} A + (1 - A) \cdot \sin^2 a = \sin^2 a + \frac{A}{2} \cdot \cos 2 a.$$

Now, if we suppose a to pass in succession through the values $0, 45^\circ, 90^\circ, 135^\circ, 180^\circ$, &c., this will become respectively $\frac{1}{2} A, \frac{1}{2}, 1 - \frac{1}{2} A, \frac{1}{2}, \frac{1}{2} A$, &c. Hence, at every quarter revolution the tints ought to change from those of the reflected rings to those of the transmitted, the complements of the former to white light; and at every half quarter revolution no rings at all should be seen, but merely an uniformly bright field illuminated with half the intensity of light which would be seen were the second tourmaline altogether removed.

926.

Phenomena of the complementary rings.
Fig. 188.

But the phenomena which actually take place are very different. At the alternate quadrants, it is true, the complementary rings are produced, and the appearance is as represented in fig. 188. The black cross is seen changed into a white one; the dark parts of the rings are become the bright ones; the green is changed into red, and the red into green, &c.; so that if we were to examine no farther, the fact would appear to agree with the hypothesis. But in the intermediate half quadrants, this agreement no longer subsists. Instead of a uniformly illuminated field, a compound set of rings, consisting of eight compartments, alternately occupied by the primary and complementary set, is seen, presenting the appearance of fig. 191, and which is further described in Art. 935.

927.

Hypothesis of a change of polarization.

The phenomena then are incompatible with the idea of *depolarization*. It remains to examine what account can be given of them on the supposition of a *change* of polarization operated by the plate; and here we must remark *in limine*, that this cause is what in Newton's language would be termed a *vera causa*, a cause actually in existence; for we have already seen that every ray, whether polarized or not, traversing a double refracting medium in any direction, except precisely along its axis, is resolved into two, polarized in opposite planes. When the incident ray is polarized, these portions (generally speaking) differ in intensity, and though, owing to the parallelism of the plate they emerge superposed, their polarization is not the less real, and either of them may be suppressed, and the other suffered to pass, by receiving them on a tourmaline properly situated. This is so far

Light. agreeable to the observed fact, when the tourmaline plate next the eye is removed, the rays of which the two sets of rings consist, coexist in the transmitted cone of rays whose apex is the eye, but, being complementary to each other, produce whiteness. This may be made matter of ocular demonstration, by employing, instead of a tourmaline, which absorbs one image, a doubly refracting achromatic prism, of sufficiently large refracting angle to separate the two pencils by an angle greater than the apparent diameter of the system of rings, when the primary set will appear in one image, and its complementary set in the other; meanwhile, to return to our tourmalines, since the two sets of rings seen in the two positions of the posterior tourmaline are complementary, it follows, that *all the* rays suppressed in one position are transmitted in that at right angles to it, and *vice versa*; and, as a necessary consequence, that every pair of corresponding rays in the primary and complementary set are polarized in opposite planes. Part IV.
Both sets of rings shown at once.

The only thing, then, which appears mysterious in the phenomena thus conceived, is the production of colour. 928. A doubly refracting crystal, which receives a polarized ray of whatever colour, divides it between its two pencils, according to a ratio dependent only on the situation of the planes of polarization and of incidence, and of the axes of the crystal, and not at all on its refrangibility. How then happens it, that at certain angles of incidence the red rays pass wholly into one image, and the green or violet into the other, while at other incidences the reverse takes place: whence, in short, arises the law of periodicity observed. To answer this question, M. Biot imagined his theory of alternate, or as he terms it *movable polarization*, according to which, as soon as a polarized ray enters into a thin crystallized lamina, its plane of polarization commences a series of oscillations, or rather alternate assumptions *per saltum* of two different positions, one in its original plane, the other in a plane making with that plane double the angle which the principal section of the crystal makes with it. These alternations he supposes to be more frequent for the more refrangible rays, and to recur periodically, like Newton's fits of easy reflexion and transmission, at equal intervals all the time the ray is traversing the crystal, which intervals are shorter the more inclined its path is to the axis or axes. This theory is remarkably ingenious in its details; and in its application to the phenomena of the rings, though open (as stated by its author) to certain obvious criticisms, is yet, we conceive, capable of being regarded as a faithful representation of most of their leading features. There is, however, one objection against it of too formidable a nature to allow of its being received unless explained away, if any other can be devised not open to the same or greater. It is, that it requires us to consider the action of a *thin* crystal on light as totally different, not merely in degree, but in kind, from that of a *thick* one, while yet it marks no limit by which we are to determine where its action as a *thin* crystal ceases, and that proper to a *thick* one commences, nor establishes any gradations by which one mode of action passes into the other. A thick crystal, as we know, polarizes the rays ultimately emergent from it in two planes, dependent only on the position of the crystal and that of the ray, while M. Biot's theory makes the position of the plane of polarization of the incident ray an element in determining their ultimate polarization by a *thin* one. Nor are we in this theory to regard as thin crystals only *films* or delicate laminæ. A plate of a tenth of an inch thick or more may be a thin plate in some cases of feebly polarizing bodies, such as apophyllite, &c. M. Biot's doctrine of movable polarization.
Objection against it.

As the apparatus employed by M. Biot for studying the phenomena of the colours of thin crystallized plates offers great conveniences for the measurement of the angles at which different tints are produced, and for their exhibition in their state of greatest purity and contrast, we shall here describe it, and state some of the chief results at which he has arrived. A (fig. 189) is a plane glass blackened at the posterior surface, or a plate of obsidian inclined at the polarizing angle to the axis of a tube AB, so as to reflect along it a polarized ray; (if greater intensity be required, we may use a pile of glass plates, taking care that they be of truly parallel surfaces, and placed exactly parallel to each other.) BC is a tube, stiffly movable round AB as an axis, having a graduated ring at B, read off by a vernier attached to the tube AB, and carrying two arms, G and H, through which the axis of a swing frame E passes, which can thus be inclined at any angle to the common axis of the tubes, its inclination, or the angle of incidence of the ray reflected along the axis on the plane of the frame being read off by an index on the divided lateral circle D. In this frame is an aperture F, in which turns a circular plate of brass having a hole in its centre, over which is fastened with wax the crystallized plate to be examined, and which can thus be turned round in its own plane, independently of any motion of the rest of the apparatus, so as to place its principal section in any *azimuth* with respect to the plane of incidence. We have found it convenient to have this part of the apparatus constructed as in fig. 190, where *a* is the square plate of the frame; *b* a divided circle movable in it and read off by an index; *c*, *d* is a circular plate movable within the divided circle to admit of adjustment, after which it is fastened in its place by a little clamp, so as to turn with the circle; this carries in its centre another swinging circle *e*, moving stiffly on its axis, and having in the middle an aperture, over which the crystal is cemented, thus giving room for an adjustment of the plane of the surface of incidence, in case it be not exactly at right angles to the principal section of the crystal, an adjustment very useful when artificial surfaces are under examination, which it is hardly possible to cut and polish with perfect precision. It is also convenient for some experiments to have a second frame similar to the first, placed on the prolongation of the arms G, H. M is a doubly refracting prism, rendered achromatic either by a prism of flint glass, or, still better, by another prism of the same doubly refracting medium. Two prisms of quartz, arranged as in Art. 882, are very convenient. Their angles should be such, that when placed at M the two images of a small aperture P, in a diaphragm near the end of the tube, should appear almost in contact. The prisms so adjusted are mounted on a stand N, independent of the other apparatus, and capable of being turned round by an arm K, carrying a vernier, by whose aid the angle of rotation, or position of the plane in which the double refraction takes place, can be read off on a divided circle L. The prism should be so adjusted in its cell, that when the vernier reads off zero, the extraordinary image should be extinguished; and when 90°, the ordinary. Occasionally a tourmaline plate or a glass reflector may be substituted for the prism. 929.
M. Biot's general apparatus described, Fig. 189, 190.

To use this apparatus, the crystallized lamina (which we will at present suppose to be a parallel plate of any 930.

Light.
Use of this
apparatus.

uniaxial crystal, having its axis perpendicular to the plane of the plate,) is to be placed on the swing frame across the aperture, and being adjusted so as to have its axis directed precisely along the axis of the tube when the vernier of D reads off zero, which is readily performed by the various adjustments belonging to the frame, as above described, the instrument is ready for use. The attainment of this condition may be known by turning the tube C on the tube A B as an axis, when the extraordinary image of the aperture P, seen through a doubly refracting prism, ought to vanish in the zero position of the vernier K, and not be restored in any part of the rotation of the tube; for it is manifest, that the axis is the only line to which this property belongs, or to which all the rings are symmetrical. It is then evident, that, however the parts of the apparatus be disposed, 1st, the reading off of the vernier D will give the angle of incidence on the plate; 2d, that of the vernier B, the angle made by the *plane of incidence* with the *plane of primitive polarization*; 3d, that of the vernier c will indicate the angle included by any assumed section of the crystallized plate perpendicular to its plane with the plane of incidence; and, lastly, that the reading of the vernier K will give the angle between the plane of primitive polarization and the principal section of the doubly refracting prism.

931.
Its applica-
tion to the
phenomena
of the rings
of one axis.
Fig. 188.

Suppose now we adjust the vernier B to zero, it will then be found, that however the plate E be situated, or whatever be the incidence of the ray, only the ordinary image will be seen (being white,) the extraordinary being extinguished (or black.) In this case we traverse the system of rings in the direction of the vertical arm of the black cross, fig. 185, of the primary, and the white one of the complementary set, see fig. 188. The phenomena are the same if we set the vernier B to 90° , and then turn the frame E on its axis, thus varying the incidence in a plane at right angles to that of primitive polarization, or, which comes to the same thing, traversing the rings along the horizontal arm of the black and white crosses. In intermediate positions of the vernier B, we traverse the rings along a diameter, making an angle with vertical equal to the reading of the vernier. In this case the two images of P are both visible, and finely coloured; the *extraordinary* image presenting the tint of the primary rings due to the particular angle of incidence indicated by the vernier D; the *ordinary*, that of the complementary system corresponding to the same angle. The colours of the two images are thus seen in circumstances the most favourable, being finely contrasted and brought side by side, so as to be capable of the nicest comparison. It is when the vernier D reads 45° , or the plane of incidence is 45° , inclined to that of primitive polarization, that the contrast of the two images is at its maximum, the tints in the extraordinary image being then most vivid, and those in the ordinary free from any mixture of white light. In general, if A represent the light of the extraordinary image in the position above mentioned, and α the angle read off on the vernier B, in any other position of the plane of incidence, the two images in this new position (for the same angle of incidence) will be represented respectively by

$$A \cdot (\sin 2\alpha)^2, \text{ and } 1 - A (\sin 2\alpha)^2$$

that is, by

$$A \cdot (\sin 2\alpha)^2, \text{ and } (\cos 2\alpha)^2 + (1 - A) \cdot (\sin 2\alpha)^2.$$

The former of these expressions indicates a ray whose tint is represented by A, and its intensity by $(\sin 2\alpha)^2$; the latter, a complementary tint $1 - A$ of the same intensity, diluted with a quantity of white light, whose intensity is represented by $(\cos 2\alpha)^2$.

932.
Agreement
of the for-
mulae with
M. Biot's
hypothesis.

These expressions represent with great fidelity the tints of both images, the intensity of the extraordinary, and the apparent degree of dilution of the ordinary one; and since a ray A polarized in a plane making an angle 2α with the principal section of the doubly refracting prism, would be divided between the extraordinary and ordinary image in the ratio of $(\sin 2\alpha)^2 : (\cos 2\alpha)^2$, it follows, that if we regard the pencil at its emergence from the crystallized plate as composed of two portions, one ($= A$) polarized in the above named plane, the other ($= 1 - A$) preserving its primitive polarization, the two pencils formed by the doubly refracting prism will be composed as follows:

	Extraordinary image.	Ordinary image.
1st. From the pencil A	$A (\sin 2\alpha)^2$	$A (\cos 2\alpha)^2$
2d. From the pencil $(1 - A)$	0	$1 - A$
Sum	$A (\sin 2\alpha)^2$	$1 - A + A \cdot \cos 2\alpha^2$ $= 1 - A \cdot (\sin 2\alpha)^2$

Office of the
doubly
refracting
prism or
tourmaline.

which are identical with those above. Thus we see, that the facts are so far perfectly conformable to M. Biot's hypothesis of movable polarization, and that we are even necessitated to admit it, *provided we take it for granted, that the rings exist actually formed and superposed in the pencil emergent from the crystallized lamina, and that the office of the doubly refracting prism is merely to analyze the emergent pencil, and separate the two sets from each other.* But if the objection mentioned above against that doctrine be really well founded, this assumption cannot be correct, and we are then driven to conclude, that the doubly refracting prism, or tourmaline, or glass reflector, interposed between the eye and the crystallized plate, performs a more important office than merely to separate the tints already formed; and that, in fact, they are actually produced by its action,—the crystallized plate only preparing the rays for the process they are here finally to undergo.

933.

To explain how this may be conceived to happen will form the object of another Section. Meanwhile we will here only add, that the transition from uniaxial to biaxial crystals is readily made. We have only to consider, that by varying the angle of incidence, (the line bisecting the angle between the optic axes being supposed perpendicular to the plane of the plate,) we cross the rings in a line passing through their centre of symmetry O, fig. 183, and making an angle with their principal diameter PP', equal to the angle read off on the vernier B, and that by turning the plate in its own plane, or varying the angle read off by the vernier c, we in effect make the system traversed pass through the successive states represented in fig. 179, 180, 181, 182, changing, not the tint, but the intensity of the extraordinary image.

Light.

When the doubly refracting prism is turned in its cell, the tints grow more dilute, and when placed in an azimuth a , that is, when its principal section is placed in the plane of incidence, both images are colourless, but of unequal brightness. This accords with M. Biot's doctrine of movable polarization; for if we grant that the pencil A is polarized in a plane making an angle $2a$ with that of primitive polarization, it will make, now, an angle $= a$ with that of the principal section of the prism, and A. $(\sin a)^2$ will be that part of the extraordinary image arising from the pencil A; on the other hand, the pencil $1 - A$ retaining its original polarization, $(1 - A) \cdot \sin a^2$ will be the portion of the extraordinary image produced by it in the new position of the prism, and the sum, or the whole image, will be simply $1 \times \sin a^2$, which being independent of A, or of the tint, indicates that the image is colourless. In the same manner it may be shown, that the ordinary image will equal $1 \times \cos a^2$, and their intensities will, therefore, be to each other as $\sin a^2$ to $\cos a^2$, and will be equal at 45° of azimuth; all which is conformable to fact.

The motion of the prism in its cell corresponds to a rotation of the posterior tourmaline in its own plane in the tourmaline apparatus. The general appearance presented by the rings of a single axis, when this rotation is not a precise quadrant, is represented in fig. 191, and the succession of changes being as follows: At the first commencement of the rotation the arms of the black cross appear to dilate; they grow at the same time fainter, and segments of the complementary rings appear in them, whose bright intervals correspond to the dark ones of the primary set, their red to the green portions of that set, and *vice versé*. The junction of the two sets is marked by a faint white or undecided tint. As the rotation proceeds, the primary segments contract in extent, and become more diluted with white, while the secondary extend, and grow more decided; at the same time the centre of the system grows gradually bright, and when the rotation has attained 90° , the whole has assumed the appearance in fig. 188. The phenomena are precisely analogous in the rings of biaxal crystals. The least deviation from exact rectangularity in the tourmalines gives rise to complementary segments in the dark hyperbolic curves answering to the arms of the black cross, and to a corresponding dilution and contraction of the primary segments, which at last disappear altogether in the undistinguishable whiteness of a pair of white hyperbolas precisely similar to the black ones of the primary rings in their perfect state.

Hitherto we have considered the rings as so narrowed by the thickness of the plate, as to be all contracted within a compass round the poles which the eye can take in at once; but if the thickness be greatly diminished, this will no longer be the case; and, instead of rings of a distinguishable form, we shall see only broad bands of colour extending to great distances from the poles, and even visible when the axes themselves are so much inclined to the surfaces of the plate as to be quite out of sight; or even when the axes actually lie in the plane of the plate. This is the case with the laminae into which sulphate of lime readily splits; the axes lie in their plane, so that to see the rings in them, we must form artificial surfaces perpendicular to the lamina, a difficult and troublesome operation, from the extreme softness and fissile nature of the substance. The phenomena of the colours of this crystal were early studied, and almost of necessity misconceived, till Dr. Brewster, by exhibiting the real axes, showed that they form only a particular case of the general phenomenon we have already dwelt on.

Adhering to the denominations employed in Art. 885—888, let us call the plane containing the two axes, *the section A*; that perpendicular to it, and passing through the line which bisects their *lesser* included angle, *the section B*; and that which similarly passes through the line bisecting their *greater* included angle, and is perpendicular to both the others, *the section C*. If the crystal have but one axis, the sections A and B pass through it, and C is at right angles to it. Then if the lamina contains both axes, its plane will be that of the section A, and the other two sections will intersect it in two lines (B and C) at right angles to each other. Conceive, now, a polarized ray to pass through such a lamina at a perpendicular incidence. Then if the plane of polarization coincide with either of the sections B and C, its polarization will be undisturbed, and the whole of the transmitted light will pass into the ordinary image. But if the plate be turned round in its own plane, the extraordinary image will reappear and become a maximum at every 45° of the plate's rotation; and if it be sufficiently thin, will exhibit some one of the colours of the rings, and the tints will descend regularly in the scale as the thickness is increased, the thickness being a measure of the tint, conformably to the general law in Art. 907, of which this is only a particular case.

When two such plates are laid together, with their sections B and C corresponding, it is evident that they are in the same relation as if they formed part of one and the same crystal; and we might therefore expect to find what really happens, *viz.* that such a compound plate polarizes the same tint that a single plate equal to the sum of the thicknesses would do. But if they be crossed, *i. e.* laid so together that the section B of the one shall coincide with the section C of the other, M. Biot has shown that the tint polarized is that due to the difference of their thicknesses. If, therefore, this difference be exactly nothing, the crossed plates will be exactly neutralized, at least at a perpendicular incidence, and that whatever be their thickness. (To procure two plates of exactly the same thickness, we have only to choose a clear and truly parallel plate terminated by fresh surfaces of fissure, and break it across.)

When, however, the incidence is not perpendicular, such a compound plate as described will still exhibit colours which vary in, apparently, a very irregular manner as the incidence changes, and with different degrees of rapidity in different planes. The tourmaline apparatus here renders signal service in rendering the law of these tints, at first sight extremely puzzling, a matter of inspection. When such a crossed plate is placed between the tourmalines, crossed at right angles, it exhibits the singularly beautiful and striking phenomenon represented in fig. 192, in which the tints are those of the reflected scale of Newton, the origin being in the black cross. If the tourmalines be parallel, the complementary colours are produced with equal regularity, as in fig. 193. If the compound crystal be turned round in its own plane, the figures turn with it, but undergo no change other than an alternation of intensity, being at a maximum of brightness when the arms of the cross are parallel and

Part IV.

934.

Effect of turning the prism in its cell.

935.

Effect of turning the tourmalines about on each other. Fig. 191.

936.

Tints produced by very thin plates at great distances from the axes.

937.

Phenomena of a single thin plate.

938.

Phenomena of crossed plates at a perpendicular incidence.

939.

Phenomena at oblique incidences.

Fig. 192.

Fig. 193.

Light.

perpendicular to the plane of original polarization, and vanishing altogether when they make angles of 45° with that plane. If the plates be not crossed exactly at right angles, or be not precisely of equal thickness, other phenomena arise which it is easier for the reader to produce for himself than to read a detailed account of. The same may be said of the very splendid but complicated phenomena produced by crossing two equally thick plates of biaxial crystals, such as mica, topaz, &c. having the section A at right angles to their surfaces.

940.

Law of tints formed by the superposition of similar plates.

Regarding, however, at present only the tint produced at a perpendicular incidence, it is found that when any number of plates of one and the same medium, of any thicknesses, are superposed with their homologous sections corresponding, the tint polarized is that due to the sum of their thicknesses; but when any one or more of them have their sections B and C at right angles to the homologous sections of the others, the tint is that due to the sum of the thicknesses of those placed one way, *minus* the sum of those of the plates placed the other way. In algebraical language, if we call $t, t', t'',$ &c. the thicknesses, and regard as negative those of the plates

941.

Law of tints produced by dissimilar plates

laid crosswise, the tint T polarized by the system will be that due to the thickness $t + t' + t'' + \&c.$ When the ray is made to traverse a plate of quartz, zircon, carbonate of lime, or any other uniaxial crystal cut so as to contain the axis of double refraction, the same law of the tints holds good, the tint T being proportional to the thickness t of the plate, and for any given plate we have $T = kt$, k being a constant depending on the nature of the plate. Now, if several plates of different uniaxial crystals be superposed, of which $t, t',$ &c. are the thicknesses, and if a negative value of t be supposed to denote a transverse position of the axis of the plate, the resultant tint will be represented by

$$T = kt + k't' + k''t'' + \&c.$$

942.

Opposite action of plates of positive and negative crystals.

In this equation, if the plates be all of one substance, $k, k',$ &c. are all alike; but if they be different, k is to be regarded as a negative quantity for all such crystals as belong to M. Biot's repulsive class, (Art. 803,) such as carbonate of lime; and positive for all such (quartz, for instance) which belong to his attractive class. Thus, each term in the above equation may change its sign from two causes, either from a change in the nature of the crystal, or from a change of 90° in its azimuth.

943.

General law.

The above is only a particular case of a more general law which may be thus announced,—*The tint ultimately produced is proportional to the interval of acceleration or retardation of the ordinary ray on the extraordinary, after traversing the whole system; the partial acceleration or retardation in each plate being proportional to the length of the path described within the plate, multiplied by the square of the sine of the angle which the transmitted ray makes, internally, with the optic axis of the plate, if it have but one axis, or to the product of the sines of its inclination to either, if it have two; and this law holds good for all positions of the plates, and all arrangements of them one among the other.* Thus (to instance its application) in the case of two similar and equal plates crossed at right angles; by the laws of polarization, the ray which, after its transmission through the first plate is ordinary, is refracted extraordinarily by the second, and *vice versa*; thus the two rays, on entering the second plate exchange velocities; and, therefore, when finally emergent, since the thickness of the second is equal to that of the first, the one ray will have lost ground on the other in its second transmission just as much as it gained it in its first; and thus the interval of retardation and the tint will be reduced to nothing.

944.

Superposition of plates cut at right angles to their axes.

From this it appears, that if two uniaxial plates cut at right angles to the axis be superposed, and adjusted so as to have their axes precisely coincident, the system of rings will have their diameters diminished if the plates be both attractive or both repulsive; but enlarged, if their characters be opposite. The experiment is rather delicate; but if made with care, placing the plates on one another with soft wax, and adjusting their surfaces by pressure to the exact position, it succeeded perfectly in the hands of Dr. Brewster.

945.

Method of ascertaining whether a crystal be positive or negative.

This affords a means, independent of any measurement of the separation of the ordinary and extraordinary pencils, of ascertaining whether an uniaxial crystal be attractive or repulsive; for if its rings be *dilated* by combining it with a thin plate of carbonate of lime, cut at right angles to the axis, it is positive; if contracted, negative. A simpler and readier method still is to fasten on a plate of the substance under examination, so cut as to show the rings, a plate of sulphate of lime of moderate thickness, and then, interposing it between the tourmalines, to turn it about in its own plane. A position will be found where the rings are unaltered. In this situation the section B or C of the sulphate of lime is in the plane of primitive polarization. If the compound plate be turned 45° from this situation, it will now be observed (if the thicknesses of the two plates be properly proportioned) that the rings in two opposite quadrants are entirely obliterated; and that in the other two they are removed to a much greater distance from the centre, forming segments of larger circles, much closer together; and in which the tints, instead of commencing from the centre, commence from a black interval between two adjacent white rings in the midst of the system, and thence descend in the scale both inwards and outwards. In this state of things, the position of the sulphate of lime, with respect to the tourmalines, must be carefully noted; and the crystallized plate being detached, a plate of carbonate of lime, (perpendicular to its axis,) or of any other known uniaxial crystal, must be substituted for it; and the sulphate of lime replaced in the same position. If, then, it be found, that the same two quadrants of the rings are obliterated in this, as in the former case, and the new set of rings in the other quadrants be also similarly situated,—then the crystal examined is of the same character as the carbonate of lime, or other crystal used as a standard of comparison; but if, on the other hand, the quadrants where the rings were obliterated in the former case be those where the new rings are formed in the latter, then the characters of the two substances are opposite. If the crystallized plate be too thin, or of too feeble polarizing power to exhibit these phenomena with necessary distinctness, we must place it in azimuth 45° on the divided apparatus described in a former article (929;) and, fixing conveniently in the polarized beam a *very* thin plate of sulphate of lime also in azimuth 45° , ascertain, by making the crystal revolve, whether its tints have been raised or depressed in this plane by the action of the sulphate; then, removing the crystal, replace it with a standard one, and repeat the observation without touching

Light. the sulphate. If both crystals have their tints raised, or both depressed, their characters are similar; if they be contrarily affected, dissimilar. An analogous mode of observation applies to biaxial crystals.

Part IV.

§ VIII. On the Interferences of Polarized Rays.

In repeating the experiments of Dr. Young on the law of interference it occurred to M. Arago, that it would be worth while to examine whether the state of polarization of the interfering rays would cause any modification in the phenomena. The experiment was easy in the case where both rays had the same polarization, being, in fact, the ordinary case; but when the interfering rays were required to have a different state of polarization, it will easily be conceived that it must be a matter of great delicacy and difficulty to superadd this condition to the others called for by the nature of the case, which requires that the interfering rays should emanate *at the same instant* from a common origin, and should have executed the same precise number of undulations or periods (within a very few units) between their origin and the point where their interference is observed. For it is not possible to change the state of polarization of a ray without either altering its course, or transmitting it through some medium in which more or fewer undulations are executed in the same space. The joint ingenuity of himself and M. Fresnel, who was associated with him in this interesting inquiry, however, soon found means of obviating the difficulties and delicacies of the subject, and the results of their experiments have been embodied by them in the following laws:

946.
Origin of the subject

1. *That two rays polarized in one and the same plane act on or interfere with each other just as natural rays, so that the phenomena of interference in the two species of light are absolutely the same.*

947.
Laws of interference of polarized light.

2. *That two rays polarized in opposite planes (i. e. at right angles to each other) have no appreciable action on each other, in the very same circumstances where rays of natural light would interfere so as to destroy each other.*

948
949

3. *That two rays primitively polarized in opposite planes may be afterwards reduced to the same plane of polarization, without acquiring thereby the power of interfering with each other.*

4. *That two rays polarized in opposite planes, and then reduced to similar states of polarization, interfere like natural rays, provided they belong to a pencil the whole of which was primitively polarized in one and the same plane.*

950

5. *In the phenomena of interference produced by rays which have undergone double refraction, the place of the coloured fringes is not alone determined by the difference of routes or velocities, but that in certain circumstances a difference of half an undulation must be allowed for.*

951.

Such are the laws of interference of polarized pencils, as stated by Messrs. Arago and Fresnel. We use in their enunciation, and indeed throughout the sequel of this part of the doctrine of Light, the language of the undulatory system, as really the most natural, and adapting itself with the least violence and obscurity to the facts. The reader may, if he please, substitute that of the corpuscular hypothesis and the Newtonian fits, superadding that of a rotation of the luminous molecules about their axes, with M. Biot; or simply content himself with a bare enunciation of facts, and with general terms expressive of the existing conditions of periodicity, without much trouble, and only a little circumlocution, but with a great sacrifice of clearness of conception. With respect to the laws themselves, the first is easily verified; we have only to repeat any of the experiments on the interference of rays emanating from a common origin, described in our section on that subject, substituting polarized instead of natural light, and the results will be precisely similar, and that in whatever plane the light be polarized. Rays, then, polarized in the same plane, interfere as natural rays under similar circumstances.

952.

Experimental verification of the first law.

The verification of the second law is more difficult and delicate. The conditions of the production of colours by interference require that the interfering rays should emanate simultaneously from a common origin, or form parts of one and the same wave proceeding therefrom as a centre; and should have performed, at the point where their interference is examined, the same number of undulations in their respective routes, within a very few units. Now at their leaving their origin they *could not be otherwise* than in the same state of polarization; and as they are required to arrive at the point of interference in opposite states, a change of polarization must be operated on one or both rays, either by reflexion, transmission, or double refraction, after leaving their origin, and that without altering, more than by a few undulations, the difference of their routes. Now, when we consider how minute a quantity an undulation is, it is easy to conceive the delicacy required in adjusting the parts of any apparatus constructed for this purpose, or the peculiar contrivances which must be resorted to to render such extreme and almost impracticable nicety unnecessary.

953.
Difficulties peculiar to the inquiry

Several ingenious and elegant methods of making the experiment have been devised by the authors last named, of which we shall content ourselves with stating one or two of the easiest and most satisfactory. And, first, the origin of the interfering rays being the image of the sun at the focus of a small lens, as we shall suppose it throughout this section, (unless the contrary be expressly said,) it is evident that if we interpose between the eye and this image a rhomboid of Iceland spar, there will be formed two images separated from each other by a space which will be greater the thicker is the rhomboid; but which will always (unless extremely thick rhomboids be used) be very small; so that the single luminous point will now be resolved into two, very near each other, and which, by the laws of polarization, send to the eye rays polarized in opposite planes. But in this disposition of things, the condition of near equality of routes is subverted; for the ordinary and extraordinary pencils pursue different paths within the crystal, and with very different velocities; so that a difference will thus arise in the total number of undulations executed by each, sufficient to destroy all evidence

954.
Verification of the second law.

Light.
M. Fresnel's
experiment
with a
bisected
rhomboid.

of interference by the production of coloured fringes. To obviate this difficulty, M. Fresnel sawed in half a rhomboid of Iceland spar, the two halves of which must of necessity have, at their line of separation and its immediate confines, precisely equal thicknesses. These halves he placed one on the other, only turning one 90° round in azimuth, so as to have their principal sections at right angles. In this state, a pencil entering them nearly at the intersection of the planes of separation would at its final emergence be divided, not into four, but into two only, (see Art. 879,) the ray ordinarily refracted in the first half having undergone extraordinary refraction in the second, and *vice versa*. The two rays, therefore, have exchanged velocities and directions, in the second transmission; and, therefore, when emergent, will have described exactly equal paths with equal velocities in each respectively, and will differ only in their states of polarization, which will be at right angles to each other. We have here, then, a case in which pencils diverge from two points side by side, and in a state in all other respects proper for interfering; nevertheless, when we look for the fringes which ought to be formed under such circumstances, (and which with natural light would be seen, see Art. 735 and 736,) none are visible. Their absence, then, must be owing to the opposite state of polarization of the interfering rays.

955.
M. Arago's
experiments
with mica
piles.

M. Arago, to make the same experiment, employed a process independent of double refraction. Two fine slits were made in a thin plate of copper, through which rays from the common origin were transmitted, and formed fringes (in their natural state) when viewed by an eye lens in the manner described, (Art. 709.) He now prepared two piles of pieces of very thin mica, or films of blown glass laid one on the other, fifteen in number, and then divided this compound plate in half by a sharp instrument, so that the halves, in the immediate neighbourhood of the line of division, could not be otherwise than of almost exactly equal thickness. These piles, when exposed at an incidence of 30° to a ray, were found to polarize the portion transmitted almost completely. They were then placed before the slits so as to receive and transmit the rays from the luminous point at precisely that incidence, and through spots which were very near each other in the undivided state of the pile. They were, moreover, so arranged, (being set on revolving frames,) that the plane of incidence could be varied (and therefore that of polarization) by turning either round in azimuth without altering its inclination to the ray, or varying the spot through which the ray passed. And it was then found, that when both piles were placed so as to polarize the rays in parallel planes, as, for instance, when both were inclined directly downwards, or one directly down and the other directly up—the fringes were formed as if the piles were away; but where one of the piles was turned round the incident ray as an axis through 90° , and so placed as to polarize the rays transmitted by it at right angles to the other, the fringes totally disappeared, nor could they be restored by inclining either pile a little more or less to the incident ray in the plane of incidence, the effect of which would be to alter gradually the length of the ray's path within the pile without changing its polarization, and thus, to compensate any slight inequality which might still subsist in their thicknesses. In intermediate positions the fringes appeared, but always the more vividly the nearer the planes of polarization approached to exact parallelism, thus attaining their maximum, and undergoing total obliteration at each quadrant of the rotation of either pile, (the other being at rest.)

956.
Tourmaline
plates sub-
stituted for
the piles.

A plate of tourmaline carefully worked to exact parallelism, and bisected, would answer equally well with the transparent piles to polarize the rays; but the tourmaline should be selected of very homogeneous texture, such are not easy to meet with, though they may be found; and in this manner the experiment is perfectly easy and satisfactory. One half the tourmaline is fixed over one aperture, the other movable in a cell in its own plane over the other. The same phenomena will then be observed by turning round the movable tourmaline as with the oblique pile in the last experiment.

957.
M. Fresnel's
fundamen-
tal experi-
ment.
Analysis of
the pola-
rized tints.

An experiment still more simple, and equally conclusive, is the following, of M. Fresnel. He placed before the sheet of copper (having, as before, two narrow slits in it very near each other) a single thin parallel lamina of sulphate of lime. Now, as this body possesses double refraction, each pencil would be divided into two—an ordinary and an extraordinary one—which, according as they emanate from the right or left hand slit, we will term *R o*, *R e*, and *L o*, *L e*. If natural light be used to illuminate the slits, these pencils will be of equal intensity, but those marked *e* will be polarized oppositely from those marked *o*. We may then form four combinations: 1. *R o* may interfere with *L o*; 2. *R e* may interfere with *L e*; 3. *R o* with *L e*; 4. *R e* with *L o*. Now of these, *R o* and *L o* are similarly polarized, and they have described equal paths with equal velocities; therefore, supposing them capable of interference, they will give rise to a set of fringes corresponding exactly to the middle of the line joining the two slits, or, as we may express it, in the axis of the apparatus. The same may be said of *R e* and *L e*. These two sets of fringes will therefore be superposed, and appear as one of double intensity. Again, *R o* may be combined with *L e*; but as these two rays have traversed the sulphate in different directions and with different velocities, those rays of each pencil which meet *in the axis* will differ by too many undulations to produce colour; and if the pencils interfere, the place of the fringes will, instead of the axis, be shifted towards the side where the pencil has the greatest velocity, (Art. 737,) and that the more, the thicker the lamina of sulphate, so that if taken of a proper thickness, this set of fringes may be removed entirely out of the reach of the middle set, and should be seen independent of it. In like manner, the pencil *R e* may interfere with *L o*, and give rise to another set of lateral fringes; but as the ray which in the former combination was the swifter, in this is the slower, this set will lie on the opposite side of the middle set, supposing it produced at all; and thus there should be seen three sets of fringes, one bright, in the middle, and two fainter on either side. But, in fact, only one set is seen, *viz.* the middle set. Therefore the combination of the rays *R o* and *L e*, *L o* and *R e*, which are polarized oppositely, produce no fringes, *i. e.* they do not interfere.

But if we cut the lamina in half, and turn one half a quadrant round in its own plane, these rays are *then* reduced to the same polarization; and the rays *R o* and *L o*, *R e* and *L e*, which in the former case gave rise to

Light. the central fringes, are now placed in opposite states of polarization; and it is accordingly found that the central fringes have disappeared entirely, and that two lateral sets formed respectively by R_o and L_e , R_e and L_o , have started into existence. If we turn the lamina slowly round, these will gradually fade away, and the central reappear and become brighter, and so on alternately; thus affording a convincing proof of the truth of the second of the laws above enunciated. Part IV
Experiment varied.

The experiment related by Messrs. Arago and Fresnel in support of their third law is as follows: Resuming the arrangement of Art. 955 or 956, and placing the piles or tourmalines so as to polarize the two pencils oppositely, let a doubly refracting crystal be placed *between the eye and the sheet of copper*, with its principal section 45° inclined to either of the planes of polarization of the interfering rays. Each pencil will then divide itself by double refraction into two of equal intensity, and polarized in two planes at right angles, one of which is the principal section itself. We ought, therefore, to expect to see two systems of fringes, one produced by the interference of the ordinary ray from the right hand aperture (R_o) with that of the left (L_o), and the other by that of R_e with L_e ; yet no fringes are seen. The experiment may be varied by substituting for the doubly refracting prism a tourmaline, or pile, with its principal section in azimuth 45° . This must reduce to a common polarization all the rays which traverse it, *viz.* the half of each pencil, yet no fringes are seen, and therefore no interference takes place. 959.
Verification of the third law.

The following experiment is adduced in the Memoir cited in support of the fourth and fifth of the above laws. A lamina of sulphate of lime is perpendicularly exposed to a polarized pencil diverging from a minute point, and immediately behind it is placed a plate of brass pierced with two very small holes near together. The principal section of the lamina is to be placed at an angle of 45° with the plane of primitive polarization. In consequence, from each of the holes (right, R ,—and left, L) will emerge a ray composed of two equal rays, R_o and R_e , and L_o , L_e oppositely polarized, *viz.* at angles $+45^\circ$ and -45° with the plane of primitive polarization, which we will suppose vertical. In this situation of things a rhomboid of Iceland spar is placed between the two holes, and the focus of the eye lens employed to view the fringes, with its principal section vertical, *i. e.* making again with that of the lamina angles of 45° either way. Each of the four rays then above mentioned will be divided into two equal rays, an ordinary and an extraordinary, thus giving rise in all to the eight rays 960.
Experiments in proof of the fourth and fifth laws.

$R_o o$, $R_e o$; $L_o o$, $L_e o$; $R_o e$, $R_e e$; $L_o e$, $L_e e$.

These rays are received on the eye lens, and conveyed into the eye. Let us now examine their respective route and states of polarization.

First, then, the rays R_o and R_e , after quitting the lamina, are parallel; and by reason of the very small thickness of it, may be regarded as superposed, being undistinguishable from each other; but they have described within the lamina different paths by different velocities, so that on emerging they will differ in phase, by an interval of retardation proportioned to the thickness of the lamina, and which we will call d , so that a being the phase of the ray R_o , $x + d$ will be that of R_e . The very same may be said of L_o and L_e . Moreover, the two rays of either of these pairs respectively are oppositely polarized, *viz.* in planes $+45^\circ$ and -45° from the vertical. This we may represent at once thus: 961.

Ray.	Phase.	Plane of Polarization.
R_o	x	$+45^\circ$
R_e	$x + d$	-45°
L_o	x	$+45^\circ$
L_e	$x + d$	-45°

Next, the portions into which either of these rays is subdivided, in traversing the rhomboid, follow in their passage through it different paths, and have different velocities; but all which are refracted *ordinarily* have one common direction and velocity; and so of those refracted *extraordinarily*; hence, between the ordinary and extraordinary rays here produced, will arise a difference of phase which we shall call δ , so that if x be the phase of any ordinary ray, $x + \delta$ will be that of the corresponding extraordinary one; and their planes of polarization will be opposed, and will form angles respectively $= 0$ and 90° with the vertical. Thus the circumstances will stand thus: 962.

A.			B.		
Ray.	Phase.	Plane of Polarization.	Ray.	Phase.	Plane of Polarization.
$R_o o$	x	0°	$R_o e$	$x + \delta$	90°
$R_e o$	$x + d$	0°	$R_e e$	$x + d + \delta$	90°
$L_o o$	x	0°	$L_o e$	$x + \delta$	90°
$L_e o$	$x + d$	0°	$L_e e$	$x + d + \delta$	90°

These eight pencils are all equal in intensity, and all those contained in the first set (marked A) will meet in one part of the field of view, while those marked B (on account of the thickness of the rhomboid, which we here suppose considerable, so as to produce a sensible, and even a large separation of the ordinary and extraordinary pencils) will meet in another, distant from the point of concurrence of (A) by an interval proportional to the thickness of the rhomboid, and which we will here suppose so large as to throw the fringes (if any) there produced, entirely out of the way of mixing with those produced at the concurrence of A. Let us then consider separately, the pencils of rays of the parcel A, and see what interferences can take place. And first, $R_o o$ may 963.

Light. combine with Loo , and since their difference of phase is zero, they will interfere in the axis of the apparatus; and their planes of polarization being coincident, there is no reason why fringes should not there be produced by their concurrence. The same holds good of the combination Reo and Leo , and, consequently, there will be superposed on each other in the axis two sets of fringes, producing one of double brilliancy.

964. Next, Roo may interfere with Leo ; but there being a constant difference of phases d in favour of the latter, the fringes produced by their concurrence will lie to the left of the axis, by an interval proportional to the thickness of the lamina of sulphate, and will be seen separately. Similarly, the concurrence of the pencils Reo and Loo will determine the production of another set of lateral fringes; but the difference of phases d being in this case in favour of the right hand pencil, this system will be situated as much to the right of the axis as the other was to the left.

965. Thus in the ordinary image three sets of fringes ought to be seen, and in the extraordinary, by a similar reasoning, as many. Now, in fact, this is the case, and the phenomena are seen on making the experiment precisely as here described. But it is evident that the rays which form the lateral fringes, by their interferences, are precisely those which, at their leaving the sulphate, had opposite polarizations, but have been afterwards reduced to similar polarization by the action of the rhomboid.

966. If instead of a rhomboid of sensible double refraction we substitute a plate of sulphate of lime, or of rock crystal, so thin as to produce no visible separation of the pencils, the fringes produced by the pencils B will be superposed on those arising from the interference of the pencils A, and we should expect therefore, instead of six, to see three sets of fringes, the middle one being still the brightest. But, in fact, we see but one set, and the lateral fringes vanish altogether. This remarkable result proves that the colours resulting from the concurrence of the rays ordinarily refracted by the rhomboid, are complementary to those resulting from that of the extraordinary rays; and therefore that we must allow half an undulation to be gained or lost when we would pass from one set to the other, precisely as in the phenomena of the reflected and transmitted colours of thin plates.

967. One of the most important consequences of these laws, is that they supply the defective link in the chain which connects the doctrine of undulations with the colours of crystallized laminae as described in the last section. It had been already remarked (as we have seen) by Dr. Young, that the passage of the ordinary and extraordinary rays with different velocities through the crystallized plate, would give rise to that difference of physical condition of the rays at their emergence which would lead to the production of colours; but the difficulty remained to explain, not why colours were produced in certain circumstances, but why they were not produced in all, in short, what share the polarization of the incident, and the analysis of the emergent rays, had in the production of the phenomena.

968. To see the nature of this difficulty more clearly, imagine a wave proceeding from a distant radiant point to be incident on a very thin crystallized lamina. It will be subdivided into two, each traversing the plate in a different direction and with its own proper velocity, and each of them emerging parallel to its original direction. The incident wave will, therefore, after emergence be resolved into two parallel to each other, but separated by a small interval equal to the interval of retardation. Now the hindmost of these ought, according to the law of interferences, to interfere with a subsequent wave of the system to which the foremost belongs, and thus periodical colours should arise on merely looking against the sky through such a lamina without any other apparatus. Why then are none seen? To this the law of Messrs. Arago and Fresnel afford a satisfactory answer. The two systems of waves into which the incident system is resolved are *oppositely polarized*, and therefore, though all other conditions be satisfied, incapable of interfering.

969. To understand how the colours of the polarized rings must be conceived to be produced by interference, let us take the simplest case when a polarized ray, A B, fig. 194, is incident on any thin crystallized plate B, whose principal section is 45° inclined to the plane of primitive polarization. Let A be the system of waves which constitutes the incident ray; then in its passage through the crystallized lamina it will be divided into systems O and E of equal intensities, polarized in planes $+45^\circ$ and -45° inclined to that of primitive polarization, and the one lagging a few undulations behind the other, so as to interfere, as represented in the figure, and constituting the parallel rays C F and D G. Let these now be received on, and transmitted through, a doubly refracting prism F G H L placed with its principal section in the plane of primitive polarization, or 45° inclined to that of the crystallized lamina. Then will each of the incident rays be again subdivided, C F into H M and I P, and D G into K N and L Q, all of equal intensity. Of these, H M and K N emerge parallel, as also K N and L Q respectively. Now the systems of waves O and E which follow each other at a certain interval d will continue to do so in both the refracted rays, as if they formed one compound system; so that each of the pencils H M K N and I P L Q will consist of a double system of waves O e and E e, O o and E o respectively. The former pair following each other at the interval d , and the latter at the interval $d \pm \frac{1}{2}$ undulation, (by reason of the demonstrated fact, that in passing from the ordinary to the extraordinary system half an undulation must be allowed. See Art. 966.) Now as each ray of these pairs respectively have similar polarizations, *viz.* those of the pair ordinarily refracted (O o and E o) in the plane of the principal section of the prism, and those of the extraordinary pair O e and E e in a plane at right angles to it, there is no reason why interference should not take place, and the consequence must be, the production of complementary colours in the two pencils finally emergent

corresponding to the intervals of retardation d and $d + \frac{\lambda}{2}$, which is just what really happens.

970. Conceive now another ray incident on B in the direction A B, but polarized in a plane at right angles to that of the ray considered in the last paragraph. Then this will undergo precisely the same series of divisions and subdivisions as the former. But the intervals of retardation will be different; for its plane of polarization when incident on B being *now* related to the plane of ordinary refraction, as that of the other ray at its incidence was

Light.

Part IV.

to the extraordinary, and *vice versa*, a difference of half an undulation must (as already explained) be admitted in the relative position of the two systems of waves O, E, at their emergence, from this cause, independent of the interval of retardation within the plate; so that if d were the interval in the former case, $d - \frac{1}{2}\lambda$ will be the difference now, and, after passing through the prism, we shall have for the intervals of retardation in the two binary pencils, instead of d and $d + \frac{1}{2}\lambda$ which they were before, $d - \frac{1}{2}\lambda$ and d . Hence the two pencils will exchange colours when the polarization of the incident light is varied by a quadrant, and this is also conformable to fact. If this reasoning be not thought conclusive, the reader is referred forwards to Art. 983 and 984.

Next, let the incident ray be unpolarized. This case, as we have seen Art. 851, is the same with that of a ray consisting of two equal rays oppositely polarized, and therefore in each pencil will coexist, superposed on each other, the primary and complementary colour arising from either portion, which being of equal intensity will neutralize each other's colours and the emergent pencils will be white, and each of half the intensity of the incident beams. This then is the reason (on this doctrine) why we see no colours when the light originally incident on the crystallized plate is unpolarized

971.
Why colours are not produced by unpolarized light.

Thus, the theory of interferences, modified by the principles above stated, affords, as we see, an explanation of the colours of crystallized plates totally distinct from that of movable polarization. The only delicacy in its application to all cases, lies in the determination which of the emergent pencils must be regarded as having its interval of retardation increased by half an undulation. M. Fresnel gives the following rule for this essential point. (Note on M. Arago's Report to the Institute on a Memoir of M. Fresnel relative to the colours of doubly refracting laminae, *Annales de Chimie*, vol. xvii. p. 80.)* *The image whose tint corresponds precisely to the difference of routes, is that in which the planes of polarization of its constituent pencils after having been separated from each other, are brought together by a contrary motion, while, on the other hand, the pencils whose planes of polarization are brought to coincidence by a continuance of the same motion by which they were separated, produce by their reunion the complementary image.* To understand this better, let PC be the plane of primitive polarization projected on that of the paper, to which let us suppose the ray perpendicular, CO that of the principal section of the crystallized lamina, and CS that of the principal section of the doubly refracting prism; then the incident pencil polarized in the plane PP' will after penetrating the lamina be divided into two, one O polarized in the plane CO, the other E in the plane CE perpendicular to it. Now, CO may always be so taken as to make an angle not greater than a right angle with CP, and CE so as to have CP between CE and CO; so that the plane CP may be conceived to open or unfold itself like the covers of a book, into CO and CE, one on either side. Again, CS may always be regarded as making an angle not greater than a right angle with CO, and when the ray O resolves itself into two (Oo and Oe) by refraction at the prism, its plane of polarization CO may be conceived to open out into the two CS and CT at right angles to each other, including CO between them; and in like manner the ray E will resolve itself into two Eo and Ee, and its plane of polarization CE will open out into the two CS and CT', having CE between them in the case of fig. 195 (a), and into CS' and CE in that of fig. 195 (b); in the former case CT' is a prolongation of CT, in the latter CS' is a prolongation of CS. The rays Oo and Eo then which make up the ordinary pencil, have, in the case of fig. (a), been each brought to a coincident plane of polarization CS by two motions in contrary directions, as represented by the arrows, and the extraordinary ones Oe and Ee have been separated and brought back to a coincident plane by motions continued in the same direction for each respectively. The reverse is the case in fig. b. In the case then of fig. a the colours of the ordinary pencil Oo + Eo will be those which correspond precisely to the difference of routes, and those of the extraordinary one Oe + Ee will correspond to that difference plus half an undulation, while in that of fig. b the reverse happens. This rule is empirical, *i. e.* is merely a result of observation. It is clear that the principle of the conservation of the *vis viva* in this, as in the colours of uncrystallized plates, requires that the two images should be complementary to each other, and therefore half an undulation *must* be gained or lost by one or the other pencil, but which of the two is to be so modified we have no means of knowing *a priori*.

972.
M. Fresnel's general rule for determining how to allow for the half undulation gained or lost.
Fig. 195.

This once determined, however, we have no difficulty in deducing the formulæ of intensity and other circumstances of the phenomena when the azimuth of the crystallized plate is arbitrary, instead of being, as we have hitherto supposed, limited to 45°. The analytical expressions of the intensity of the pencils we must reserve for our next section.

973.

§ IX. Of the application of the Undulatory Doctrine to the explanation of the phenomena of Polarized Light and of Double Refraction.

The phenomena of double refraction and polarization, as exhibited in the experiments of Huygens on Iceland spar, were regarded by Newton and his followers as insuperable objections to the undulatory doctrine, inasmuch as it appeared to them impossible, by reason of the *quaquaversum* pressure of an elastic fluid, to conceive an undulation as having a different relation to different regions of space, or as possessing *sides*. "Are not," says Newton, "all hypotheses erroneous in which light is supposed to consist in pressure or motion propagated

974.
Newton's objections against the undulatory theory.

* This Memoir was read to the Institute, Oct. 7, 1816. A Supplement was received Jan. 19, 1818. M. Arago's report on it was read June 4, 1821. And while every optical philosopher in Europe has been impatiently expecting its appearance for seven years, it lies as yet unpublished, and is known to us only by meagre notices in a periodical Journal.

Light. through a fluid medium?" "for pressures or motions propagated from a shining body through an uniform medium, must be on all sides alike, whereas it appears that the rays of light have different properties in their different sides." "To me, this seems inexplicable, if light be nothing else than pressure or motion propagated through ether." *Opticks*, book iii. quest. 28. And, again, quest. 29; "Are not rays of light very small bodies emitted from shining substances?" "The unusual refraction of Iceland crystal looks very much as if it were performed by some kind of attractive virtue lodged in certain sides both of the rays and of the particles of the crystal." "I do not say this virtue is magnetical.—It seems to be of another kind. I only say, that, whatever it be, it is difficult to conceive how the rays of light, unless they be bodies, can have a permanent virtue in two of their sides which is not in their other sides, and this, without any regard to their position as to the space or medium through which they pass."

975. Although we have no knowledge of the intimate constitution of elastic media, or the manner in which their contiguous particles are related to each other and affect each other's motion, yet it is certain that the mode and laws of the propagation of motion through them by undulation cannot but depend very materially on this connection. The only analogies we have to guide us into any inquiry into these laws, are those of the propagation of sound in air or water, and of tremors through elastic solids, and along tended chords and surfaces; and such is the extreme difficulty of the subject when taken up in a purely mathematical point of view, that we are forced to have recourse to these analogies, and, dismissing in the present state of science the vain hope of embracing the whole subject in analytical formulæ, suffer ourselves to be instructed by experience, as to what modifications the peculiar constitution of vibrating media may produce in the propagation of motion through them. Now, when sound is propagated through air or water, in which the molecules are at least *supposed* to have no mutual connection but to be capable of moving with equal facility, and to be restored to their places with equal elastic forces, in whatever direction they are displaced, and in which, moreover, it is (at least theoretically) taken for granted, that the motion of any molecule has an equal tendency to set in motion those adjacent to it, in whatever direction these may be situated with respect to it; it is difficult to conceive that the motion of a molecule in the surface of a wave, at some distance from the centre whence the sound emanates, can be performed otherwise than in the direction of the radius, or at right angles to the surface of the wave; so that in this case the motion of the vibrating molecules must coincide with the direction of the rays of sound, and there appears, therefore, no reason why such rays should bear different relations to the different regions of space surrounding them, whether right or left, above or below; for the ray being regarded as an axis, all parts of the sphere round it are similarly related to it.

976. But if we conceive a connection of any kind, such as may possibly be established by repulsive and attractive forces, or magnetic or other polarities subsisting between the molecules of the vibrating medium, the case is altered. It will no longer then follow of necessity, that the individual motion of each molecule is performed in the direction in which the general wave advances, but it may be conceived to form any angle with that direction, even a right angle. A familiar instance of such a mode of propagation may be seen in the wave which runs along a long stretched cord, struck, shaken, or otherwise disturbed at one end. The direction of the wave is the length of the cord, and that of the motion of each molecule lies in a plane perpendicular to it. Now this is precisely the kind of propagation which M. Fresnel conceives to obtain in the case of light. He supposes the eye to be affected only by such vibrating motions of the ethereal molecules as are performed in planes perpendicular to the directions of the rays. According to this doctrine, a polarized ray is one in which the vibration is constantly performed in one plane, owing either to a regular motion originally impressed on the luminous molecule, or to some subsequent cause acting on the waves themselves, which disposes the planes of vibration of their molecules all one way. An unpolarized ray may be regarded as one in which the plane of vibration is perpetually varying, or in which the vibrating molecules of the luminary are perpetually shifting their planes of motion, and in which no cause has subsequently acted to bring the vibrations thus excited in the ether to coincident planes.

977. The analogy of the tended cord (which appears to have suggested itself to Dr. Young on considering the optical properties of biaxial crystals in 1818) will help our conception greatly. Suppose such a cord of indefinite length, stretched horizontally, and one end of it being held in the hand, let it be agitated to and fro with a motion perpendicular to the length of the cord. Then will a wave or succession of waves be propagated along it, and every molecule of the cord will, after the lapse of a time proportional to its distance from the hand, begin to describe a line or curve similar and similarly situated to that described by the extremity at which the agitation originates. If the original agitation be regularly repeated and constantly confined to one plane, the same will be true of the motion of each molecule, and the whole extent of the cord will be thrown into the form of an undulating curve lying in one plane, so far as the motion has reached. In this case it will represent a polarized ray or system of waves. If, after a few vibrations in one plane, the extremity be made to execute a few in another, and then again in another, and so on, so that the plane of vibration shall assume in rapid succession all possible situations, since each molecule obeys exactly the same law of motion with the extremity, the curve will consist of portions lying in all possible planes, and since by reason of the propagation of the undulation along it, every point of it is in succession agitated by the motion of every other, all these varied vibrations will run through any given point of it, and were a sentient organ like the human retina stationed there, the impression it would receive would be analogous to that excited in the eye by an unpolarized ray of light.

978. It may be objected to this mode of conceiving the luminiferous undulations, that the molecules of the ether, if it be a fluid, such as we have hitherto all along regarded it, cannot be supposed connected in strings, or chains like those of a tended cord, but must exist separate and independent of each other. But it is sufficient for our purpose to admit such a degree of lateral adhesion (we hesitate to term it *viscosity*) as may enable each molecule in its motion not merely to *push before it* those which lie directly in the line of its motion, but to *drag along*

Fresnel's
theory of
transverse
vibrations.

Propagation
of light
assimilated
to that of
waves along
a stretched
cord.

Objects
considered.

Light. *with it* those which lie on either side, in the same direction with itself. Or, acknowledging at once the difficulty, since light is a real phenomenon, we are not to expect it to be produced without a mechanism adequate to so wonderful an effect. We do not hesitate to attribute to the fluids which are imagined to account for the phenomena of heat, electricity, magnetism, &c. properties altogether repugnant to our ordinary notions of fluids, and why should we deny ourselves the same latitude when light is to be accounted for. It is true the properties we must attribute to the ether appear characteristic of a solid than of a fluid, and may be regarded as reviving the antiquated doctrine of a plenum. But if the phenomena can be thereby accounted for, *i. e.* reduced to uniform and general principles, we see no reason why that, or any still wilder doctrine, should not be admitted, not indeed to all the privileges of a demonstrated fact, but to those of its representative, or *locum tenens*, till the real truth shall be discovered. Assuming it, then, with M. Fresnel, as a postulatam, that the vibrations of the ethereal molecules which constitute light are performed in planes at right angles to the direction of the ray's progress, let us see what account can be given of the phenomena of polarized light.

Part IV.

And first, then, of the interference of two polarized rays, whether polarized in the same, or different planes. The plane of polarization in this doctrine may be assumed to be either that in which the vibrations are executed, (*i. e.* a plane passing through the direction of the ray and the line described by each of the vibrating molecules in its excursion,) or one perpendicular to it, which we please. Reasons, presently to be stated, render the latter preferable, but at present it is a matter of indifference which we assume. Now, in § 3, Part III. we have investigated at length, with a view to the present inquiry, the modes of vibration which result from the combination of any assigned vibrations, whether executed in the same or different planes; and it follows from the purely mechanical principles there laid down, 1st, That the combination of two vibrations executed in the same plane, produces a resultant vibration in the same plane, which may be of any degree of intensity from the sum to the difference of the intensities of its component vibrations, according to the difference of their phases. Now, each of these systems of vibration represents a polarized ray; so that rays polarized in the same plane ought, on these principles, to be capable of destroying or reinforcing each other by interference, as we see they do. But the case is otherwise when the component vibrations are executed in different planes, for in that case it is obvious that they never can destroy each other completely so as to produce rest. The general case of non-coincident planes of vibration is analyzed in Art. 618; and in Art. 621 we see, that even when each of the component vibrations is rectilinear, the resultant is elliptic; so that each molecule of the ether performs continual gyrations in one direction, and never can be totally quiescent.

979.
Explanation of the phenomena of interference on this doctrine.

Thus we see that the interference of rays similarly polarized, and the non-interference of those dissimilarly, is a necessary consequence of the hypothesis we are considering; and indeed was the phenomenon which first suggested it. It may be familiarly explained by the analogy of our tended cord. Conceive such a cord to have its extremity agitated at equal regular intervals with a vibratory motion performed in one plane, then it will be thrown, as we have seen, into an undulatory curve, all lying in the same plane. Now, if we superadd to this motion another, similar and equal, but commencing exactly half an undulation later, it is evident that the direct motion every molecule would assume, in consequence of the first system, will at every instant be exactly neutralized by the retrograde motion it would take in virtue of the other; and, therefore, each molecule will remain at rest, and the cord itself be quiescent. But if the second system of motions be performed in a plane at right angles to the first, the effect will evidently only be to distort the figure of the cord into a curve of double curvature, which, in the general case, will be an elliptic helix, and will pass into the ordinary circular one when the two component vibrations differ in phase by a quarter of an undulation, or 90°. (See Art. 627. *Corol.*)

980.
Analogy of the stretched cord.

In this case the extremity of the cord describes a circle with a continuous motion, and this motion is imitated by each molecule along its whole length. It is easy to make this a matter of experiment; we have only to hold in our hands the end of a long stretched cord, or grasp it firmly in any part of its extent, and work the part held round and round, with a regular circular motion, and we shall see the cord thrown into a helicoidal curve, each portion of which circulates in imitation of the original source of the motion.

981.
Case of a rotatory or spiral motion.

But experience shows, not merely that two equal rays polarized at right angles do not destroy each other for any assignable difference of origins, but, that whatever be this difference, the intensity of the resultant ray remains absolutely the same. Now this is also a necessary consequence of the theory of transverse vibrations. To show this, we need only refer to the expressions for A, B, C in equation (7.) Art. 619, resuming at the same time the notation and reasoning of that article. The intensity of the impression made on the eye by any ray being proportional to the *vis viva*, is represented by the sum of the several *vires vivæ* in the three rectangular directions, or by $A^2 + B^2 + C^2$, that is, by

982
Resultant of two rays oppositely polarized investigated

$$a^2 + b^2 + c^2 + a'^2 + b'^2 + c'^2 + 2aa'. \cos(p - p') + 2bb'. \cos(q - q') + 2cc'. \cos(r - r').$$

Now if we assume the directions of the coordinates x and y to be those transverse to that of the ray, and the one in the plane of polarization of one ray, the other in that of the other, at right angles to it, and that of z in the direction of the ray itself, we have

$$a' = 0, \quad b = 0 \quad c = 0, \quad c' = 0;$$

and therefore the above expression for the intensity becomes

$$A^2 + B^2 + C^2 = a^2 + b^2,$$

which is independent of $p - p', q - q', r - r'$, the difference of phases, and is equal to the sum of the inten-

Ligh. sities of the separate rays. And we may remark, by the way, that no other supposable mode of vibration but that in question, in which c and c' , the amplitudes of vibration in the direction of the ray vanish, could produce the same result. (Fresnel's *Considerations Théoriques sur la Polarization de la Lumière*. *Bulletin de la Société Philomatique*, October, 1824.)

983. Let us now consider what will happen when a ray polarized in any plane is resolved into two polarized in any other two planes at right angles to each other, and these again reduced to two others also at right angles to each other, by a second resolution. Suppose C , (fig. 195, a) to be the course of a ray projected on a plane perpendicular to its direction, (that of the paper,) and in which, consequently, the vibrations of the molecule C are performed. Let PCP' be the line of vibration of this molecule, and therefore (according to the hypothesis assumed) at right angles to the plane of primitive polarization. When this ray is divided into two others oppositely polarized, the vibrations are of course resolved into two others performed in planes at right angles to each other. Let CO and CE be the projections of these planes, which are therefore perpendicular to the planes of polarization of the two new rays respectively. Suppose that at any instant the molecule C of the primitive ray is moving from C in the direction CP ; then this motion, if resolved into two, will give rise to two motions, one in the direction from C towards O , the other from C towards E . If each of these motions be again resolved into two, in planes whose projections are SCS' and TCT' , at right angles to each other, that in the direction CO will produce two motions, one in the direction CS , and the other in the direction CT ; and on the other hand the motion in the direction CE will produce one in the direction CS , and the other (in the case of fig. 195, a) in the direction CT' opposite to CT . Thus the two resolved motions in the plane SS' will conspire, but those in the plane TT' will oppose, each other. In the case of fig. 195, b , the reverse will happen; the motions in the plane TT' conspiring, and those in the plane SS' opposing, each other. For simplicity of conception, however, we will confine ourselves to the former case. If, now, we pass from the consideration of the vibrations to that of the rays, it will appear that we have, in fact, resolved the original ray polarized in the plane PP' into two, polarized in planes perpendicular respectively to CO and CE ; and these again, finally, each into two, *viz.* one polarized in the perpendicular to SS' , and one perpendicular to TT' . The two portions polarized perpendicular to SS' form one ray, and those perpendicular to TT' another; but in the former, the component portions tend to strengthen,—in the latter, to destroy each other. Hence, if we consider the two former portions as having a common origin, we must regard the latter as differing by half an undulation.

984. Hitherto we have supposed the second resolution of the rays to take place at the same point C in the course of the ray as the first, but this may not be the case, and several cases may be imagined; first, we may suppose the two portions into which the ray is first resolved to run on in the same line with equal velocities; and after describing any given space, to be then resolved, at another point C' (whose projection in the figure will coincide with C) into the final rays SS' and TT' . It is evident that this will make no difference in the result, for the phases in which each ray arrives at C' will be alike; and after the second resolution the conspiring vibrations in the direction SS' will still be in the same phase, and the opposing ones in the plane TT' must still be regarded as in opposite phases, *i. e.* as differing by half an undulation. Or, secondly, we may suppose, that, owing to any cause, the two resolved rays do not travel with equal velocity, (as in the case where the resolution is performed by double refraction.) In this case, if i be the interval of retardation of the one ray on the other when they arrive at C' , i will represent the difference of phases of the two rays at the instant of their second resolution. Consequently, when resolved, the final ray, whose vibrations are performed in SS' , will be the *sum*; and that whose vibrations are performed in TT' , the *difference* of two rays, one in a certain phase (θ), the other in the phase $\theta + i$; or, which is the same thing, the former will be the *sum* of two components in the phases θ and $\theta + i$; the latter, the *sum* of two in the phases θ and $\theta + i + 180^\circ$, so that still the difference of half an undulation is to be applied. In the case of fig. 195, b , if we pursue the same reasoning, it will appear that this difference still subsists, but must be applied conversely, *viz.* to the compound ray whose vibrations are performed in CS .

985. We have here, then, the theoretical origin of the allowance of half an undulation, in those cases where it is required to account for the polarized tints, Art. 966, and of the rule laid down in Art. 972 for its correct application. However arbitrary the assumption may have appeared as there presented, and however singular it may have seemed to make the affections of a ray at one point of its course dependent on those which it had at a former instant, we now see that the whole is a direct and very simple consequence of the ordinary elementary rules for the composition and resolution of motions. It is worthy of notice, that the fact was ascertained before the theory of transverse vibrations was devised, so that this theory has the merit of affording an *a priori* explanation of what had previously all the appearance of a mere gratuitous hypothesis.

986. In conceiving the resolution of a ray into two others polarized in different planes, we may be aided by the analogy of the tended cord, which we have before had occasion to refer to. In fig. 196 let AB be a stretched cord, branching at B into the two BC and BD , making a small angle with each other at B , and having either equal or unequal tensions. Suppose the plane in which the two branches lie to be (for illustration's sake) horizontal, and let the extremity A of the single cord be made to vibrate regularly in a vertical plane; or, at least, let the vibrations of the cord, before arriving at B , be reduced to a vertical plane by means of a small polished vertical guide IK , against which the cord shall press lightly, and on which it may slide freely without friction. Beyond the point of bifurcation B , and at such a distance that the excursions of the molecule B shall subtend no sensible angle from them, let two other such polished *guiding planes* be placed, inclined at different angles to the horizon, and making a right angle with each other. Suppose now B to make any excursion from its point of rest, then were the plane EF parallel to IK , the molecule of the branch BC contiguous to EF would slide on EF through a space equal to the whole excursion of B ; but since it is inclined to IK at an angle

Rationale of the rule for allowing half an undulation.
Fig. 195.

Application of the analogy of the stretched cord.
Fig. 196.

Light.

(= θ) a part only of the motion of B will be employed in causing this molecule to glide on EF, and the remainder will cause the cord to bend over and press on the obstacle; but by reason of the minuteness of the excursions of B, this bending and the resistance of the obstacle and consequent loss of force will be very minute and may be neglected. Now, since the pressure of the obstacle removes the cord from the position it would have taken had no obstacle existed, in a direction perpendicular to its surface, it is easy to see that the amplitude of excursion of the contiguous molecule on the plane EF must be to that of B as $\cos \theta$ to radius; and, therefore, calling a the amplitude of B's excursions, $a \cdot \cos \theta$ will be that of the molecule contiguous to EF, and of course that of every subsequent molecule of the branch BC. Here the part of B's motion, which is perpendicular to EF, is not expended or destroyed in bending the cord BC over the obstacle, but remains in activity, and exerts itself on the branch BD, causing it to glide on the plane GH; and the amplitude of the excursions of the molecule in contact with this plane will in like manner be represented by $a \cdot \cos$ (inclination of GH to IK,) that is, by $a \cdot \cos (90 - \theta)$, or by $a \cdot \sin \theta$. The *vis viva*, then, in each of these respective planes is represented by $a^2 \cdot \cos^2 \theta$ and $a^2 \cdot \sin^2 \theta$, whose sum is equal to a^2 , the initial *vis viva*.

If we decompose, in like manner, the maximum velocity a of the ethereal molecule C (fig. 195) in the direction CP into two in the respective directions CO and CE, we get $a \cdot \cos \theta$ and $a \cdot \sin \theta$ for the elementary velocities; and since the amplitudes, *ceteris paribus*, are as the velocities, (Art. 610,) the amplitudes of the component rays will be respectively $a \cdot \cos \theta$ and $a \cdot \sin \theta$; and their intensities, which are as the squares of the amplitudes, (Art. 605,) will be $a^2 \cdot \cos^2 \theta$ and $a^2 \cdot \sin^2 \theta$. Now this is the very law propounded by Malus for the intensities of the two portions into which a polarized ray is divided by double refraction, and of which the theory of transverse vibrations gives, as we see, a simple and rational *a priori* account, thus raising it from a mere empirical law to the rank of a legitimate theoretical deduction.

We have not done with the analogy of the tended cord. What we have shown in Art. 986 is independent of the tensions of the branches into which the cord is divided, and relates only to the amplitudes of their excursions from rest when thrown into vibration. But the velocity with which the waves, once produced, will be propagated along either branch depends solely on its tension. Nothing, however, prevents the tensions of the two branches from being very different; for, whatever be the ratio of two forces applied in the directions BC and BD, they may be balanced at B by a proper force applied along any other line as BA. Hence the waves will run along BC and BD with different velocities. Similarly, if we conceive, that owing to the peculiar constitution of crystallized bodies, and the relation of their particles to the ether which pervades them, its molecules are more easily displaced, or yield to a less force in certain planes than in others; or, in other words, that it possesses different elasticities in different directions; then will the planes of polarization assumed by the resolved portions of the rays determine the elasticities brought into action, and, by consequence, the velocities of their propagation. Now we have, in a former section, shown that the bending of a ray at the confines of a medium depends essentially on its velocity within as compared with that without, by the analytical relations deduced from the "principle of swiftest propagation." A difference of velocity, therefore, draws with it, as a necessary consequence, a diversity of path; and thus the bifurcation, or double refraction of a ray incident on a crystallized surface, presents no longer any difficulty in theory, provided we can find an adequate reason for the resolution of its vibrations into two determinate planes at the moment of its entering the crystal.

Let us take (with M. Fresnel, *Annales de Chimie*, xvii. p. 179 *et seq.*) the case of a crystal with one axis. We may regard this, or rather the ether within it, modified in its action by the molecular forces of the crystal, as an elastic medium in which the elasticity in a direction perpendicular to the axis is different from that in a direction parallel to it, that is, in which the molecules are more easily compressible in the one than in the other direction; but, equally so in all directions perpendicularly to the axis, on whatever side the pressure be applied. To aid our conceptions in imagining such a property, we may assimilate an uniformly elastic medium to an assemblage of thin, elastic, hollow, spherical shells in contact; and such a medium as we are considering, to a similar assemblage of oblate or prolate hollow ellipsoids, arranged with all their axes parallel to one common direction, which is that of the axis of the crystal.* It is evident that the resistance of the spherical assemblage to pressure must be the same in all directions, but that of the spheroidal must differ according as the pressure is applied perpendicularly or parallel to the axis. Thus, it is easy to crush an egg by a force applied in the direction of its shorter diameter, which will yet sustain a violent pressure applied at the extremities of its longer. It is, moreover, evident, if any molecule of such an assemblage were disturbed, so as to throw it into vibration, that, provided always the amplitude of its excursions were extremely small compared to the diameter of each ellipsoid, the immediate tendency of the vibration will be to communicate motion to two strata only of molecules, *viz.* that in which the axis and equator of the disturbed molecule lie respectively, since it is only at the poles and equator that they touch, and therefore only through these points that motion can be communicated from one to the other. Consequently, any motion communicated to a molecule of such a mass could only be propagated by vibrations performed in planes parallel and perpendicular to the axis. Hence, if a vibratory motion in any plane be propagated into such an assemblage of particles from without, it will immediately, on its reaching it,

Part IV.

987.

Rationale of Malus's rule for the intensity of the complementary rays.

988.

Case of the two resolved undulations propagated with different velocities.

989.

Explanation of the phenomena of double refraction in crystals with one axis.

* The idea of spheroidal molecules in Iceland spar suggested itself to Huygens (rather fancifully, perhaps) as a means by which spheroidal undulations might be propagated through it, (*Op. Reliq.* tom. i. *Tractatus de Lumine*, p. 70, cited by Wollaston, *Phil. Trans.* ciii. p. 58;) and the last-named eminent Philosopher, in the Bakerian Lecture for 1813, has most ingeniously shown how such molecules may be combined to build up crystals, having the primitive forms and cleavages of acute and obtuse rhomboids. It is true, that in all this there is much hypothesis; and it should be observed, too, that the crystallographic structure would require oblate spheroids, where in the text we have employed prolate, and *vice versa*. But we intend there only an *analogy*, not a *theory*. It would be easy to devise hypothetical modes of action where these forms might be reversed if needful.

light.

be resolved into two, in the planes above named; and these, by reason of the different elasticities, will be propagated with different velocities.

990.

Bifurcation
of the re-
fracted ray
explained.

The reader must not suppose that this is intended for an account of the real mechanism of crystallized bodies. It is merely intended to show that it is not absurd, or contradictory to sound mechanical principles, to assume that such *may* be their constitution, that vibrations can only be propagated through them by molecular excursions executed in planes parallel and perpendicular to their axes. Assuming, then, that such is the case, the vibrations of a ray incident on such a crystal will be resolved into two, performed in these respective planes, and their velocities of propagation being different, the rays so arising will follow different courses when bent by refraction. Let us first consider that whose vibrations are executed in planes perpendicular to the axis. Since the crystal is symmetrical with respect to its axis, and equally elastic in all directions perpendicular to it, the velocity of propagation of this portion will be the same in all directions. Its index of refraction, therefore, will be constant, and the refraction of this portion will follow the ordinary law. Moreover, its plane of polarization being that *perpendicular* to which the vibrations are performed, will necessarily pass through the axis, in which respect it also agrees with the ordinary ray, as actually observed.

Properties
of the ordi-
nary ray.

991.

Properties
of the extra-
ordinary ray
explained.

The extraordinary ray arises from the other resolved portion of the original vibration, which is performed in a plane parallel to the axis. By the principle of transverse vibrations, it is also performed in a plane perpendicular to the ray. If, then, we suppose a plane to pass through the extraordinary ray and the axis, it will cut a plane perpendicular to the ray in a straight line, which will be the direction of the vibratory motion. This direction, then, is inclined to the axis in an angle equal to the complement of that made by the extraordinary ray with the latter line, and therefore, when the extraordinary ray is parallel to the axis, the line of vibration is perpendicular to it, and *vice versa*. In the former case, the elastic force resisting the displacement of the molecules is the same as in the case of the ordinary ray, and therefore the velocities of both rays are equal, and their directions coincide, and thus along the axis there is no separation of the rays. In the latter, the elasticity is that parallel to the axis, and therefore differing from the former by the greatest possible quantity. Here, then, the difference of velocities, and therefore of directions is at its maximum. In intermediate situations of the extraordinary ray, the elasticity developed is intermediate, and therefore also the velocity and double refraction. Thus we see, that according to this doctrine the difference of velocities, and consequent separation of the pencils should be nothing in the axis, and go on increasing till the extraordinary ray is at right angles to it, which is conformable to fact. Lastly, the plane of polarization of the extraordinary ray being at right angles to the plane of vibration, must also be at right angles to a plane passing through the ray and the axis, which is also conformable to fact.

992.

The theory of M. Fresnel gives then, as we see, at least a plausible account of the phenomena of double refraction in the case of uniaxal crystals; and when we consider the profound mystery which, on every other hypothesis, was admitted to hang over this part of the subject, we must allow that this is a great and important step. But the same principles are equally applicable to biaxal crystals with proper modifications, and (which is a strong argument for their reality) lead, when so applied, to conclusions which, though totally at variance with all that had been taken for granted before, on the grounds of imperfect analogy and insufficient experiment, have been since verified by accurate and careful experiments, and have thus opened a new and curious field of optical inquiry. Nothing stronger can be said in favour of an hypothesis, than that it enables us to anticipate the results of experiment, and to predict facts opposed to received notions, and mistaken or imperfect experience.

993.

Explana-
tion of the
phenomena
of movable
polarization.

But before we enter on this, it may be right to show how the phenomenon on which the theory of movable polarization is founded, is accounted for by the doctrine of transverse vibrations. According to this theory, as soon as a polarized ray enters a crystal, it commences a series of alternate assumptions of one or other of two planes of polarization, in the azimuths 0° and $2i$, i being the inclination of the principal section to the plane of primitive polarization: the plane assumed being in azimuth 0° , when the thickness traversed is such as to render the interval of retardation of the ordinary on the extraordinary ray 0, or any whole number of undulations, and in azimuth $2i$ when it is any whole odd number of semi-undulations. Suppose a ray polarized in the azimuth 0 to be incident perpendicularly on a crystallized lamina, having its principal section in the azimuth i , then it will be resolved into two, the vibrations of which are respectively performed in the principal section, and perpendicular to it. Consequently, if we represent by unity the amplitude of the original vibrations, those of the two resolved vibrations will be equal respectively to $\sin i$ and $\cos i$. Now, the thickness of the plate being first supposed such as to render the interval of retardation an exact number of undulations, these rays will emerge from the lamina in exact accordance, and being parallel, the systems of waves of which they consist will run on together. Being polarized, however, in opposite planes they will neither destroy each other, nor produce a compound ray equal to their sum, but their resultant must be determined as in Art. 623. For we have here the case of rectilinear vibrations, in complete accordance, of given amplitudes, and making a given angle (90°) so that the result there obtained is immediately applicable to this case, and the resultant vibration will be, first, rectilinear, so that the compound ray will appear wholly polarized in one plane: and, secondly, its amplitude will be, both in quantity and direction, the diagonal of a parallelogram whose sides are the amplitudes of the component vibrations. Consequently, it will be identical with that by whose resolution these were produced, and therefore the resultant, or emergent compound ray will be, in respect both of its polarization and intensity, precisely similar to the original incident one.

Case of
complete
accordance.

994.

Fig. 197.
Case of
complete
discordance.

When the difference of paths within the crystal is an exact odd multiple of half an undulation, the waves at their egress from the posterior surface will be in complete discordance. But their resultant may still be determined by the same rule, regarding either of the rays as negative, *i. e.* as having its vibrations executed in the opposite direction. For suppose the molecule C moving in the direction CP, with the velocity CP

Light.

(fig. 197) at the entry of the ray, then the resolved velocities in the planes CO and CE will be represented in quantity and direction by CO and CE. But at their egress, the vibrations in the direction CE having gained or lost a half undulation on those in CO, if CO represent the quantity and direction of motion of the molecule C in that plane, CE' equal and opposite to CE will represent its motion in the other plane, and this, combined with CO will compose, not the original motion CP, as in the former case, but CQ, making an equal angle with CO on the other side. The resultant ray, then, instead of being polarized in the plane of the incident one, (*i. e.* perpendicular to CP) will be polarized in a plane perpendicular to CQ, making an angle equal to PCQ ($= 2 PCO = 2 i$) with CO.

When the difference of routes is neither an exact number of whole, or half undulations, the vibrations of the resultant ray (by Art. 621) will no longer be rectilinear, but elliptic; and in the particular case when the interval of retardation is a quarter or an odd number of quarter undulations, it will be circular. In this case, the emergent ray, varying its plane of vibration every instant, will appear wholly depolarized, so as to give two equal images by double refraction in all positions of the analysing prism.

These several consequences may be rendered strikingly evident by a delicate and curious experiment related by M. Arago. Let a polarized pencil, emanating from a single radiant point, be incident on a double rhomboid of Iceland spar, composed of two halves of one and the same rhomboid, superposed so as to have their principal sections at right angles to each other. Then the emergent rays will emanate as if from two points (see Art. 879) near each other, and polarized in opposite planes. Let these two cones of rays be received on an emerald glass, or in the focus of an eye-lens, so that the glass or field of view shall be illuminated at once by the light of both, which being oppositely polarized will exhibit no fringes or coloured phenomena, but merely a uniform illumination; and let all the light but that which falls on a single very small point of the field of view be stopped by a plate of metal, with a small hole in it, so as to allow of examining the state of polarization of the compound ray illuminating this point, separately from all the rest. Then it will be seen, on analysing its light by a tourmaline or double refracting prism, that, when the spot examined is distant from both radiants by the same number of undulations, although in fact composed of two rays oppositely polarized, (as may be proved by stopping one of them, and examining the other singly,) yet it presents the phenomenon of a ray completely polarized in one plane, which is neither that of the one or the other of its component rays, but the original plane of polarization of the incident light. Suppose now, by a fine screw we shift gradually the place of the metal plate so as to bring the hole a little to one or the other side of its former place. The ray which illuminates it will appear to lose its polarized character as the motion of the plate proceeds, and at length will offer no trace of polarization; continuing the motion, and bringing in succession other points of the field of view under examination, the light which passes through the hole will again appear polarized, at first partially, and at length totally; not, however, as before, in the plane of primitive polarization, but in a plane making with it twice the angle included between it and the principal section of the first rhomboid, and so on alternately. Thus we are presented with the singular phenomenon of two rays polarized in planes at right angles, which produce by their concurrence a ray either wholly polarized in one or the other of two planes, or not polarized at all, according to the difference of routes of the rays before their union.

In 1821, M. Fresnel presented to the Academy of Sciences of Paris a Memoir, containing the general application of the principle of transverse vibrations to the phenomena of double refraction and polarization as exhibited in biaxial crystals, which was read in November of that year. A brief report on the experimental parts of this Memoir by the Committee of the Academy appointed to examine it, about half a dozen pages, was published in the *Annales de Chimie*, vol. xx. p. 337, recommending it to be printed as speedily as possible in the collection of the *Mémoires des Savans Etrangers*. We are sorry to observe, that this recommendation has not yet been acted upon, and that this important Memoir, to the regret and disappointment of men of science throughout Europe, remains yet unpublished; though we trust (from the activity recently displayed by the Academy in the publication of their Memoirs in arrears) this will not long continue to be the case.* An abstract by the author himself, which appeared in the *Bulletin de la Société Philomatique* of 1822, and was subsequently reprinted in the *Annales de Chimie*, 1825, enables us, however, to present a sketch, though an imperfect one, of its contents, supplying to the best of our ability the demonstration of the fundamental propositions, and reaping a melancholy gratification from the inadequate tribute, which, in thus introducing for the first time to the English reader a knowledge of these profound and interesting researches, we are enabled to pay to departed merit. *His saltem accumulæ donis—et fungar inani munere*. For even at the moment when we are recording his discoveries, their author has been snatched from science in the midst of his brilliant career by a premature death, like his hardly less illustrious contemporary, Fraunhofer, the early victim of a weakly constitution and emaciated frame, unfit receptacles for minds so powerful and active.

M. Fresnel assumes, as a postulatam, that the displacement of a molecule of the vibrating medium in a crystallized body (whether that medium be the ether, or the crystal itself, or both together, in virtue of some mutual action exercised by them on each other,) is resisted by different elastic forces, according to the different directions in which the displacement takes place. Now it is easy to conceive, that in general the resultant of

Part IV.

995

996.

Experiment exhibiting these several cases of interference

997.

Fresnel's general theory of double refraction.

998.

General expression for the elastic forces of a medium in vestigated.

* This delay has been productive of a singular consequence, which will suffice to show the small degree of publicity which labours, even the most important, can acquire by the circulation of such notices as those mentioned in the text. So lately as December 1826, the Imperial Academy of Sciences of Petersburg proposed as one of their prize questions for the two years 1827 and 1828, the following, "To deliver the optical system of waves from all the objections which have (as it appears) with justice, been urged against it, and to apply it to the polarization and double refraction of light." In the programme announcing this prize, M. Fresnel's researches on the subject are not alluded to (though his *Memoir on Diffraction* is noticed,) and it is fair to conclude, were not then known to the Academy. Precisely one month before the publication of this programme, the Royal Society of London awarded their Rumford Medal to M. Fresnel, "for his application of the undulatory theory to the phenomena of polarized light, and for his important experimental researches and discoveries in physical optics." Our readers will be gratified to know, that the valuable mark of this high distinction reached him a few days before his death.

L.gnt.

all the molecular forces which act on a displaced molecule, is not necessarily parallel to the direction of its displacements when the partial forces are unsymmetrically related to this direction, but the proposition may be demonstrated *à priori*, as follows. Suppose three coordinates x, y , and z , to represent the partial displacements of any molecule M in their respective directions, and $r (= \sqrt{x^2 + y^2 + z^2})$ the total displacement, making angles α, β, γ , respectively with the axes of the x, y, z , so that $x = r \cdot \cos \alpha, y = r \cdot \cos \beta, z = r \cdot \cos \gamma$. Now, since in this theory we assume that the displacements of the molecules are infinitely, or at least extremely small compared with the distances of the molecules *inter se*, it is evident that whatever be the law of molecular action, the force resulting from any displacement must (*cæteris paribus*) be proportional to the linear magnitude of that displacement, and can, therefore, be only of the form $r \cdot \phi$, where ϕ is some unknown function of the angles α, β, γ , or their cosines. And, moreover, since such infinitely small displacements, in whatever direction made, neither alter the angular position, nor distance of the displaced molecule among the rest, by any sensible quantity, all their forces will act on it in its displaced position in the same manner as before. Hence the total force developed by the simultaneous displacements x, y, z , or by the single displacement r must be equivalent to (or the statical resultant of) the three which would be developed independently by the several partial displacements x, y, z . Now the force originating in the partial displacement x alone will result from $r \phi$ by making $r = x$ and ϕ equal to α , where α is the same function of $1, 0, 0$, that ϕ is of $\cos \alpha, \cos \beta, \cos \gamma$. α therefore is a constant depending only on the position of the axes of the x, y, z with respect to the molecules of the crystal. And when this partial force $= \alpha x$ is resolved into the directions of these several axes, since its direction (whatever it be) is determinate, the resolved portions can only be of the form $Ax, A'x, A''x$, where A, A', A'' are in like manner dependent only on the position of the coordinates x, y, z with respect to the molecules, and not at all on α, β, γ , which are arbitrary, and where $A^2 + A'^2 + A''^2 = \alpha^2$. The same being true of the partial forces brought into play by the displacements y and z , it follows that the total force arising from the displacement r must be the resultant of the three forces

$$f = Ax + By + Cz, \quad f' = A'x + B'y + C'z, \quad f'' = A''x + B''y + C''z,$$

respectively parallel to the axes of the x, y, z , where the coefficients are independent of α, β, γ , and where, in like manner, $B^2 + B'^2 + B''^2 = b^2, C^2 + C'^2 + C''^2 = c^2$. But we have $x = r \cdot \cos \alpha, y = r \cdot \cos \beta, z = r \cdot \cos \gamma$, so that if we put

$$\begin{aligned} f &= r \{ A \cdot \cos \alpha + B \cdot \cos \beta + C \cdot \cos \gamma \}, \\ f' &= r \{ A' \cdot \cos \alpha + B' \cdot \cos \beta + C' \cdot \cos \gamma \}, \\ f'' &= r \{ A'' \cdot \cos \alpha + B'' \cdot \cos \beta + C'' \cdot \cos \gamma \}, \end{aligned}$$

the resultant of f, f', f'' will be the force urging the displaced molecule.

999.

Expression of the elasticity in any assigned direction.

Acts obliquely to the direction of displacement.

Now these forces acting in the directions of the coordinates may each be decomposed into two, one in the direction of the displacement r , and the other at right angles to it in the planes respectively of r and x, r and y, r and z , the sum of the former will be

$$F = f \cdot \cos \alpha + f' \cdot \cos \beta + f'' \cdot \cos \gamma,$$

which is the whole force tending to urge the displaced molecule directly to its position of equilibrium. The latter will be respectively equal to $f \cdot \sin \alpha, f' \cdot \sin \beta$, and $f'' \cdot \sin \gamma$; but as they act, although in one plane, yet not in the same direction, they will not destroy each other, unless they be to each other in the ratio of the sines of the angles they make with each other's direction. But it is evident, that since α, β, γ are arbitrary, this condition cannot hold good in general, because it furnishes two equations, which, taken in conjunction with the relation $\cos \alpha^2 + \cos \beta^2 + \cos \gamma^2 = 1$, suffice to determine α, β, γ . Hence it follows, that the displaced molecule is, except in certain cases, urged by the elastic forces of the medium obliquely to the direction of its displacement.

1000.

Axes of elasticity defined and investigated

Mr. Fresnel next goes on to observe, that in general every elastic medium has three rectangular axes, in the direction of which, if a molecule be displaced, the resultant of the molecular forces urging it will act in the direction of its displacement. These are the excepted cases just alluded to, and to the axes possessing this property, (which he regards as the true fundamental axes of the crystal,) he gives the name of *Axes of Elasticity*.

To demonstrate this proposition we must observe, that, by mechanics, in order that the resultant of three rectangular forces f, f', f'' shall make angles α, β, γ with their three directions, and therefore be coincident in direction with r , they must be to each other in the ratio of the cosines of these angles, and therefore we must have the following equations expressive of this condition,

$$\frac{f}{f'} = \frac{\cos \alpha}{\cos \beta}; \quad \frac{f}{f''} = \frac{\cos \alpha}{\cos \gamma}; \quad \frac{f'}{f''} = \frac{\cos \beta}{\cos \gamma}.$$

These three equations are in general equivalent to two only, but when combined with the equation $\cos \alpha^2 + \cos \beta^2 + \cos \gamma^2 = 1$ resulting from the geometrical conditions of the case, they suffice to determine α, β , and γ ; and if we put u, v, w for the cosines of these angles, furnish the following system of equations which every axis of elasticity must satisfy.

$$\begin{aligned} (Au + Bv + Cw)v &= (A'u + B'v + C'w)u, \\ (Au + Bv + Cw)w &= (A'u + B'v + C'w)u, \\ (A'u + B'v + C'w)w &= (A''u + B''v + C''w)u, \\ u^2 + v^2 + w^2 &= 1. \end{aligned}$$

Light.

Suppose by elimination we have derived from these equations the position of one axis of elasticity, then it will follow of necessity, that two others must exist, at right angles to it and to each other. To prove this, we must consider the connection between the partial forces developed by any displacement of the molecule M, and the molecular attractions and repulsions of the medium. Let ϕ be the action of any molecule dm on M, which we suppose to be exerted in the direction of their line of junction, and to be a function of their mutual distance ρ . Then, if we suppose M displaced by any arbitrary quantities $\delta x, \delta y, \delta z$ (infinitely small in comparison with ρ) in the direction of the three coordinates, we have

$$\delta \phi = \left(\frac{x}{\rho} \delta x + \frac{y}{\rho} \delta y + \frac{z}{\rho} \delta z \right) \cdot \frac{d\phi}{d\rho},$$

and putting

$$\phi' = \frac{d\phi}{d\rho}, \quad \text{and} \quad \frac{x}{\rho} = \cos \lambda, \quad \frac{y}{\rho} = \cos \mu, \quad \frac{z}{\rho} = \cos \nu,$$

we have

$$\delta \phi = \phi' \cdot \{ \delta x \cdot \cos \lambda + \delta y \cdot \cos \mu + \delta z \cdot \cos \nu \}.$$

Consequently, since the force of the molecule dm , resolved into the directions of the coordinates, is respectively equal to

$$(\phi + \delta \phi) dm \cdot \frac{x}{\rho}, \quad (\phi + \delta \phi) dm \cdot \frac{y}{\rho}, \quad \text{and} \quad (\phi + \delta \phi) dm \cdot \frac{z}{\rho},$$

the sum of all these throughout the medium will be the total action on M; but since in the original position of the molecule M it is in equilibrium, we have

$$\int \phi dm \cdot \frac{x}{\rho} = 0, \quad \int \phi dm \cdot \frac{y}{\rho} = 0, \quad \text{and} \quad \int \phi dm \cdot \frac{z}{\rho} = 0,$$

so that the whole action of the medium on M in its displaced situation will be, in the three directions

$$\int \frac{x}{\rho} dm \cdot \delta \phi, \quad \int \frac{y}{\rho} dm \cdot \delta \phi, \quad \int \frac{z}{\rho} dm \cdot \delta \phi;$$

that is, in the direction of the x ,

$$\int \phi' dm \cdot \{ \cos \lambda^2 \delta x + \cos \mu^2 \delta y + \cos \nu^2 \delta z \};$$

$\delta x, \delta y, \delta z$, are the partial displacements of M in the directions of the coordinates, and are, therefore, the same we denoted in Art. 998 by x, y, z . Restoring these denominations, we see that, on this hypothesis, (the most natural which can be formed respecting the mode of molecular action) the coefficients A, B, C, can be no other than the following,

$$A = \int \phi' dm \cdot \cos \lambda^2, \quad B = \int \phi' dm \cdot \cos \lambda \cdot \cos \mu, \quad C = \int \phi' dm \cdot \cos \lambda \cdot \cos \nu;$$

and by similar reasoning we find

$$\begin{aligned} A' &= \int \phi' dm \cdot \cos \lambda \cdot \cos \mu, & B' &= \int \phi' dm \cdot \cos \mu^2, & C' &= \int \phi' dm \cdot \cos \mu \cdot \cos \nu; \\ A'' &= \int \phi' dm \cdot \cos \lambda \cdot \cos \nu, & B'' &= \int \phi' dm \cdot \cos \mu \cdot \cos \nu, & C'' &= \int \phi' dm \cdot \cos \nu^2; \end{aligned}$$

and, consequently, the following relations must necessarily subsist between these coefficients

$$B = A', \quad C = A'', \quad C' = B''.$$

This premised, suppose we have determined one axis of elasticity of the medium by the foregoing equations. Since the positions of the axes of the coordinates are arbitrary, we are at liberty to suppose that of the x coincident with the axis so determined, which renders $A' = A'' = 0$, and consequently $B = 0$ and $C = 0$, and $B' = C'$, because the relations above demonstrated are general and independent of any particular situation of the axes. The equations of Art. 1000 then become

$$\begin{aligned} A u v &= (B' v + C' w) u, & A u w &= (B'' v + C'' w) v, \\ (B' v + C' w) w &= (C' v + C'' w) v, & u^2 + v^2 + w^2 &= 1. \end{aligned}$$

Now if we put $u = 0$, or $a = 90^\circ$, the two former of these are satisfied without any relation supposed between v and w , so that if we determine these from the two latter only, the whole system will be satisfied. These (making $u = 0$) give at once by elimination

$$w = \sqrt{\frac{1}{2} \left(1 \pm \frac{1}{\sqrt{4m^2 + 1}} \right)}; \quad v = -\sqrt{\frac{1}{2} \left(1 \mp \frac{1}{\sqrt{4m^2 + 1}} \right)},$$

where $m = \frac{C'}{B' - C'}$. Now since m^2 is necessarily positive, $4m^2 + 1$ is so, and is > 1 ; therefore $\frac{1}{\sqrt{4m^2 + 1}}$

is real and < 1 , consequently w^2 and v^2 are both positive, and therefore v and w both real, and less than unity. Hence it follows, that there are necessarily two axes at right angles to the x which satisfy the conditions of axes of elasticity, and the opposite signs of v and w show that they are at right angles to each other.

For simplicity, therefore, we will in future suppose the directions of the coordinates to be coincident with those of the axes of elasticity, so as to make

Part IV.

1001.

Three exist in any crystal at right angles to each other.

General relation between the partial elasticities.

1002.

One axis being given the situation of the other determined.

1003.

Light.

$$\Lambda = a, A' = A'' = 0; \quad B' = b, B = B'' = 0; \quad C'' = c, C = C' = 0;$$

then we have by Art. 998 for the partial forces,

$$f = ax = ar \cdot \cos \alpha, \quad f' = by = br \cdot \cos \beta, \quad f'' = cz = cr \cdot \cos \gamma,$$

and by 999,

$$F = r \{ a \cdot \cos \alpha^2 + b \cdot \cos \beta^2 + c \cdot \cos \gamma^2 \}$$

for the whole force urging the molecule M in the direction of the r , generally assumed, in which it will be observed that

$$a = \int \phi' \cdot \cos \lambda^2 dm, \quad b = \int \phi' \cdot \cos \mu^2 dm, \quad c = \int \phi' \cdot \cos \nu^2 dm.$$

1004. M. Fresnel next conceives a surface, which he terms the "Surface of Elasticity," constructed according to the following law:—on each of the *axes of elasticity*, and on every radius r drawn in all directions, take a length proportional to the square root of the elasticity exerted on the displaced molecule by the medium in the direction of the radius, or to \sqrt{F} . Then if we call R this length, or the radius vector of the surface of elasticity, we shall have

$$R^2 = \{ ar \cdot \cos \alpha^2 + br \cdot \cos \beta^2 + cr \cdot \cos \gamma^2 \} \times \text{const.}$$

Its radius vector expressed.

The values of R parallel to the axes are then had by the equation

$$R^2 = \text{const } ar, \quad R^2 = \text{const } \times br, \quad R^2 = \text{const } \times cr$$

which (for brevity, as we shall have no further occasion to recur to our former denominations) we shall express simply by a^2, b^2, c^2 , so that the equation of the surface of elasticity will be of the form

$$R^2 = a^2 \cdot \cos X^2 + b^2 \cdot \cos Y^2 + c^2 \cdot \cos Z^2,$$

where X, Y, Z , now stand for α, β, γ , the angles made by R with the axes of the coordinates.

1005. Velocity and plane of polarization of an interior wave determined.

Let us now imagine a molecule displaced and allowed to vibrate in the direction of the radius R , and retained in that line, or at least let us neglect all that part of its motion which takes place at right angles to the radius vector. Then the force of elasticity by which its vibrations are governed will be proportional to R^2 , and the velocity of the luminous wave propagated by means of them, in a direction transverse to them (or at right angles to R) will be proportional to R , so that the surface of elasticity being known, the velocity of a wave transmitted through the medium in a given direction, and with a given plane of polarization will be had at once as follows. Parallel to the surface of the wave, and at right angles to its plane of polarization draw a straight line. This will be the direction of the vibrations by which the wave is propagated. Parallel to this line draw a radius vector to the surface of elasticity, and it will represent the wave's velocity.

1006. Equation of the surface of elasticity.

The equation of the surface of elasticity, if we put for $R, \cos X, \cos Y, \cos Z$, their values in terms of three coordinates will become

$$(x^2 + y^2 + z^2)^2 = a^2 x^2 + b^2 y^2 + c^2 z^2.$$

It is, therefore, in general a surface of the fourth order. If we suppose it cut by a plane passing through its centre, whose equation must therefore in general be of the form $mx + ny + pz = 0$, the curve of intersection will be a species of oval whose diameters are not necessarily all equal.

1007. Resolution of an incident wave into two.

Suppose now any molecule set in vibration in this plane, then at any period of its motion it will not be urged directly to its point of rest but obliquely, so that it will not describe a straight line, but will circulate in a curve more or less complicated; its motion in this, however, will always be resolvable into two vibratory rectilinear ones at right angles to each other, one parallel to the greatest, and the other to the least diameter of the section. Each of these vibratory motions will, by the laws of motion, be performed independently of the other, and therefore the motion propagated through the crystal will affect every molecule of it in the same way as if two separate and independent rectilinear vibrations (at right angles as above) were propagated through it, with different velocities. Consequently every system of waves propagated from without into the crystal, will necessarily on entering it be resolved into two propagated with different velocities, and polarized in planes at right angles to each other, viz. those parallel respectively to the greatest and least diameter of a section of the surface of elasticity parallel to the plane of either wave. And as every difference in the velocities of two waves propagated parallel to each other through a medium, gives rise to a corresponding difference in their planes at their emergence from it into another, where they assume a common velocity, these waves will at their egress no longer be parallel, and the rays which are perpendicular to them will be inclined to each other, thus producing the phenomena of double refraction; and it is evident that the waves at their egress must retain the planes of polarization they received in the crystal, because any molecule of the exterior medium at the junction of the media will begin to move only in the plane in which it was displaced by the contiguous molecule in the medium.

Polarized in opposite planes.

1008.

This theory then accounts perfectly both for the bifurcation of the emergent ray, and the opposite polarizations of the two portions into which it is divided. These portions will coincide in direction, and there will be no double refraction when the section of the surface of elasticity above mentioned is (if such can ever be the case) a circle, because all its radii being then equal, the elasticity is the same in all directions, and all vibrations performed in it will have equal periods, so that in this case the resolution of the incident wave into two no longer takes place, nor is its plane of polarization changed. Now the section in question becomes a circle, when $x^2 + y^2 + z^2 = \text{const} = r^2$, or when $a^2 x^2 + b^2 y^2 + c^2 z^2 = r^4$. Combining these with $mx + ny + pz = 0$, we get

$$r^4 = r^2 (x^2 + y^2 + z^2),$$

$$p^2 r^4 = r^2 (p^2 x^2 + p^2 y^2 + (m x + n y)^2),$$

and

$$p^2 r^4 = p^2 a^2 x^2 + p^2 b^2 y^2 + c^2 (m x + n y)^2,$$

and equating these, and considering that the equation thence resulting ought to be verified independently of any particular values of x, y , we get

$$r^2 (m^2 + p^2) = a^2 p^2 + m^2 c^2,$$

$$m n r^2 = m n c^2,$$

$$r^2 (p^2 + n^2) = b^2 p^2 + n^2 c^2.$$

Investigation of the optic axes.

These equations cannot be satisfied except by supposing either m, n , or p to vanish, or the section in question to pass through one or other of the axes. If we suppose $m = 0$, we have $r = a, \left(\frac{n}{p}\right)^2 = \frac{a^2 - b^2}{c^2 - a^2}$, which shows that $\left(\frac{n}{p}\right)^2$ cannot be positive, and of course $\frac{n}{p}$ not real, unless a , the semiaxis of the surface through which the section passes, be that intermediate in length between b and c , the other two semiaxes.

It appears then, that the surface of elasticity admits of two circular sections and no more, formed by diametral planes passing through the mean axis of the surface, and (since $\frac{n}{p}$ has two values equal but of opposite signs)

that these sections are both equally inclined to each of the other two axes. The normals to these sections are the directions of no double refraction, or the *optic axes* of the crystal. Of these, then, there will be two and two only, in all crystals which possess three unequal axes of elasticity, and rays propagated along them will suffer neither double refraction, nor change of polarization.

The position of these axes depends wholly on the values of a, b, c , the semiaxes of the surface of elasticity. We have, however, no other measure of the elasticity of the medium than the velocity with which the rays are propagated through it; and if, as the phenomena of ordinary dispersion indicate, the rays of different colours be propagated in one and the same medium with velocities somewhat different, (an effect which might result from certain suppositions relative to the extent of the sphere of action of its molecules compared with the lengths of an undulation,) the semiaxes a, b, c , which must be taken proportional to the velocities of propagation, must be supposed to vary a little for waves of different lengths. Now this variation may not be in the same ratio for all

the three semiaxes, and thus a variation in the values of $\frac{n}{p}$ will arise. But $\frac{n}{p}$ is the tangent of the inclination

of the plane of section to the plane of the xy , or of half the angle the two circular sections make with each other, *i. e.* the cotangent of half the angle between the optic axes, which will thus vary, and give rise to that separation of axes of different colours, and their distribution over a certain angle, in the plane containing any two of the same colour, which observation shows to exist, (Art. 921 and 922.)

The general laws of double refraction flow with great facility from these principles. We have only to resume the construction and reasoning of Art. 806 and 807, *et seq.*, substituting for the ellipsoid of revolution, which the Huygenian theory assumes as the figure of a wave originating in any molecule of the crystal, the surface, whatever it be, which, in the general case, terminates a wave so propagated, and investigating the point of contact I (fig. 170) of this surface with a plane IKT passing through the line KT drawn as there described. There is this difference, however, in the two cases, or, at least, in the method of treating them, that in the theory there stated the form of the wave is made a matter of arbitrary assumption, in the present case it is to be determined *à priori*. This will render it necessary to depart in some respects from the course before adopted. If we know, *à priori*, the form of the wave, the position of the tangent plane is given; *vice versa*, if we can determine the position of this plane in all cases, *à priori*, the figure of the wave, which must be such as to touch all such planes, under the conditions of the case, becomes known.

Now, in Art. 807, it is shown that the tangent plane is in all cases coincident with the position assumed within the crystal, by the surface of a plane indefinite wave propagated from an infinitely distant luminary, perpendicular to the line of incidence RC. It follows, moreover, from Art. 811, that if we know the velocity with which such a plane wave advances within the crystal in a direction perpendicular to its surface, we may calculate its inclination to the surface of incidence by the law of ordinary refraction, assuming an index of refraction which is to that of the ambient medium as the velocity of the wave before incidence is to its velocity within the medium perpendicular to its own surface. The reader will here keep in view the distinction noticed in Art. 813 between the velocity of the wave and that of the ray conveyed by it, whose direction, generally speaking, is oblique to its surface. Now the velocity of a wave within the medium in any direction is given by the equation of the surface of elasticity, whose radius vector expresses it in all cases. But it has been shown, that every vibration impressed on the molecules of the crystal is resolved into two rectilinear ones propagated with velocities proportional to the greatest and least diameters of that section of the surface of elasticity which is parallel to the plane in which they are performed. Now it is the same thing, (as far as the law of double refraction is concerned,) whether we regard the bifurcation to take place by the separation of a single exterior ray into two interior ones, or a single interior into two exterior. We will take the latter case, and suppose the

1010.
Dispersion of the axes of different colours explained.

1011.
Application of the Huygenian construction to the general case.

1012
Direction and velocity of a plane wave.

Light.
Velocities
of an ordi-
nary and
extraordi-
nary plane
wave inves-
tigated.

ordinary and extraordinary plane waves to be parallel *within* the medium. Their velocities may then be investigated as follows: the equation of the surface of elasticity being Part IV.

$$R^4 = a^2 x^2 + b^2 y^2 + c^2 z^2,$$

if we take, for the equation of the second plane,

$$z = m x + n y,$$

and put V for the maximum or minimum radius vector of the surface in the section in question, V will be the value of R , which makes $dR = 0$, and therefore will be given by elimination from the following system of equations

$$V^2 = x^2 + y^2 + z^2,$$

$$V^4 = a^2 x^2 + b^2 y^2 + c^2 z^2,$$

$$z = m x + n y,$$

and their differentials, regarding V as constant. This elimination, which is complicated enough, must be conducted as follows: first, if among the differential equations we eliminate dx, dy, dz ; and for z in the whole system substitute its value, we shall get, putting $p = a^2 - b^2$; $q = a^2 - c^2$; $r = b^2 - c^2$;

$$V^4 = (a^2 + m^2 c^2) x^2 + (b^2 + n^2 c^2) y^2 + 2 m n c^2 x y,$$

$$V^2 = (1 + m^2) x^2 + (1 + n^2) y^2 + 2 m n x y,$$

$$0 = m n q x^2 - m n r y^2 + k x y,$$

where

$$k = p + n^2 q - m^2 r = (1 + n^2) q - (1 + m^2) r.$$

These, by elimination, give the following, in which

$$M = k^2 + 4 m^2 n^2 q r;$$

$$M x^2 = V^2 (V^2 - c^2) \{ (1 + n^2) k + 2 m^2 n^2 r \} - r k V^2,$$

$$M y^2 = - V^2 (V^2 - c^2) \{ (1 + m^2) k - 2 m^2 n^2 q \} + r q V^2,$$

$$M x y = - m n \{ (1 + n^2) q + (1 + m^2) r \} V^2 (V^2 - c^2) + 2 m n q r V^2;$$

and by equating the square of the last of these to the product of the two first, we find, after all reductions, the following equation for determining V :

$$(V^2 - a^2) (V^2 - b^2) + m^2 (V^2 - b^2) (V^2 - c^2) + n^2 (V^2 - a^2) (V^2 - c^2) = 0.$$

1013.
General
equation of
a wave pro-
pagated
from a
point in the
medium.

The roots of this equation determine the maximum and minimum values of the radius vector in the plane of section, and therefore the velocities of ordinary and extraordinary plane waves moving parallel to each other within the crystal, and these found, the figure of the wave becomes known, from the condition that its surface must always be a tangent to a plane distant by the quantity V from the secant plane whose equation is $z = m x + n y$; and that, whatever be the values of m and n . Its investigation is therefore reduced to a purely geometrical problem. Required the equation of a curve surface, which shall touch every plane parallel to a plane whose equation is $z = m x + n y$; and distant from it by a quantity V , a function of m and n given by the above equation, which, being resolved, will be found to lead to the following equation

$$\begin{aligned} (a^2 x^2 + b^2 y^2 + c^2 z^2) (x^2 + y^2 + z^2) - a^2 (b^2 + c^2) x^2 - b^2 (a^2 + c^2) y^2 \\ - c^2 (a^2 + b^2) z^2 + a^2 b^2 c^2 \end{aligned} \Big\} = 0.$$

1014.
Nonexist-
ence of the
Cartesian
law of re-
fraction in
biaxial
crystals.

The surface represented by this equation is, generally speaking, of the fourth order, and consists of two distinct surfaces, or sheets, (*nappes*.) One of these, by its contact with the plane in question, determines the direction of the ordinary, and the other of the extraordinary ray. Now, it is important to remark, that this equation, so long as particular values are not assigned to a, b, c , is not decomposable into quadratic factors, so that neither of the sheets of which it consists is spherical, or ellipsoidal; and, consequently, neither the ordinary nor the extraordinary ray follows either the Cartesian or Huygenian law of refraction. This is a consequence too remarkable not to have been put to the test of experiment. Two methods have been put in practice by M. Fresnel for this purpose. The first consisted in measuring directly the velocities of the two rays in plates of topaz cut in different directions with respect to their axes by the method explained under the head of interferences, (Art. 738 and 739.) Since a difference of velocity of the interfering rays displaces the diffracted fringes as a difference of thickness would do, it is manifest that if, in two plates differently cut, but of precisely the same thickness, the fringes formed by the ordinary rays are differently displaced when the plates are combined successively with one and the same equivalent plate of glass, or any other standard medium, their velocity cannot be the same in both plates; and if such difference be observed to take place, both in the fringes formed by the interference of the ordinary and of the extraordinary rays severally, with a compensated pencil, it is clear that neither can have a constant velocity. Now the condition of equal thickness is secured by cementing the two plates edge to edge, and grinding and polishing them together, and carefully examining the surfaces after the operation, to be satisfied of their precise continuity, which may be done by the reflected image of a distant object, and yet more delicately by pressing slightly on them a convex lens of long focus, over their line of junction. If the coloured rings formed between the surfaces be uninterrupted, we are sure that this condition

Light

is rigorously satisfied. The experiment so made, M. Fresnel found to confirm the conclusion to which the above theory leads. But in corroboration of this important result, the following method was also used.

In topaz the extraordinary refraction is stronger than the ordinary; so that the ordinary ray, when the two are separated by a prism of that medium, may be at once recognised, by being the least deviated. M. Fresnel procured two prisms to be cut from one topaz, in both of which the base was parallel to the cleavage planes, and therefore perpendicular to a line bisecting the angle between the optic axes and to the principal section of the crystal, *i. e.* to the mean axis of elasticity; but in one the plane of the refracting angle was coincident with, and in the other perpendicular to, that section, these being the planes in which the difference between the velocities of the ordinary ray is the greatest, as is easily seen from what has above been said. These prisms were cemented side by side, so as to have their bases in one plane and their refracting edges in one straight line; and were then very carefully ground and polished to plane surfaces, so that the refracting angles in both could not be otherwise than precisely equal. In this situation the compound prism ABC, fig. 199, 1, (which is seen in perspective in fig. 199, 2,) whose refracting angle ABC was about 92° , was achromatised by two prisms CBA and DCA of crown glass, in which circumstances a slight, uncompensated refraction remained in favour of the topaz prism. Looking now through the side EB, the whole combination was turned round the refracting edge as an axis, till the image of a distant object, a black line on a white ground, appeared stationary; so that the refracted rays, both ordinary and extraordinary, must have traversed the prisms very nearly parallel to the base, or at right angles to the mean axis, but in the different planes above mentioned in each. Now it was observed, that the least refracted image of the black line so seen, that is the ordinary one, was broken at the junction of the two prisms, being more deviated by one than by the other, while the most refracted or extraordinary image formed a continuous line in both. This latter fact (which, at first sight, would lead us to suspect that the extraordinary image had been mistaken for the ordinary one) is a consequence of the theory above explained, and is an additional confirmation of it.

When two of the axes of elasticity (as *b* and *c*, for instance) are equal, the general equation of the surface of the wave becomes decomposable into two factors, and may be put under the form

$$(x^2 + y^2 + z^2 - b^2) \{ a^2 x^2 + b^2 \cdot (y^2 + z^2) - a^2 b^2 \} = 0;$$

which is the product of the equation of a sphere with that of an ellipsoid of revolution. In this case the two circular sections coincide with the plane of the *yz*, and the two optic axes with the axis of the *x*. We have here then the case of uniaxial crystals, and are thus furnished with an *à priori* demonstration, both of the Huygenian law of elliptic undulations, in the case of the extraordinary wave in such crystals, and of the constancy of the index of refraction in that of the ordinary. The manner in which this results as a corollary from the general case is at once elegant and satisfactory.

M. Fresnel gives the following simple construction for the curve surface bounding the wave in the case of unequal axes, which establishes an immediate relation between the length and direction of its radii. Conceive an ellipsoid having the same semiaxes *a*, *b*, *c*; and having cut it by any diametral plane, draw perpendicular to this plane from the centre two lines, one equal to the greatest, and the other to the least, radius vector of the section. The loci of the extremities of these perpendiculars will be the surfaces of the ordinary and extraordinary waves; or, in other words, their lengths will be the lengths of the radii of the waves in those directions, and will therefore measure the velocity of the two rays propagated in those directions, in the same way as the radii of the Huygenian ellipsoid are proportional to the velocities of the extraordinary ray in their direction.

Finally, if we divide unity by the squares of the two semiaxes of a diametral section of the ellipsoid, the difference of these quotients will be found to be proportional to the product of the sines of the angles which the perpendicular to this section makes with the two normals to the planes of the circular sections of the ellipsoid. Now, in all the crystals hitherto known, these sections differ very little from the circular sections of the surface of elasticity, and may, without sensible error, be supposed to coincide with them; consequently, the two normals in question may be taken for this purpose as the optic axes of the crystal. We have thus the origin of that law, deduced from the phenomena of the coloured lemniscates, which makes the difference of the squares of the reciprocal velocities proportional to the product of the sines made by the ray with the optic axes; and thus the phenomena of the polarized rings are all made to depend on the same general principles.

Such is the beautiful theory of Fresnel and Young, (for we must not in our regard for one great name forget the justice due to the other, and to separate them and assign to each his share would be as impracticable as invidious, so intimately are they blended throughout every part of the system; early, acute, and pregnant suggestion characterising the one,—and maturity of thought, fulness of systematic developement, and decisive experimental illustration, equally distinguishing the other. If the deduction in succession of phenomena of the greatest variety and complication from a distinctly stated hypothesis, by strict geometrical reasoning, through a series of intermediate steps, in which the powers of analysis alone are relied on, and whose length and complexity is such as to prevent all possibility of foreseeing the conclusions from the premises, be a characteristic of the truth of the hypothesis,—it cannot be denied that it possesses that character in no ordinary degree; but, however that may be, as a generalization the reader will now be enabled to judge whether the encomium we passed on it in a former Article be merited. We can only regret that the necessary limits of this Essay, which is already extended greatly beyond our original design, forbid our entering farther into its details.

The axes of elasticity are those which M. Fresnel regards as the fundamental axes of a doubly refractive medium. The optic axes can in no view of the subject be regarded as such, for several obvious reasons. First, they are seldom symmetrically situated relative to fundamental lines in the crystalline form; secondly, because they vary in position according to the colour of the incident light; thirdly, because it is found that for one and the same coloured illumination, and in the same crystal, their situation varies by a variation of temperature.

Part IV.

1015.

Another experiment to prove the same.

1016.

Case of uniaxial crystals.

1017.

Construction of the wave by an ellipsoid.

1018.

Origin of the rule of the product of the two sines.

1019

1020.

Dr. Brewster's theory of polarizing axes.

Light. This important fact has been lately ascertained by M. Mitscherlich, and we shall presently have occasion to speak further of it. From all these reasons it follows, that we can regard them only as resultant lines, to which no *à priori* properties can be supposed to belong, but which simply satisfy the condition $v - v' = 0$, according to the laws which regulate the constitutions of the functions v, v' , the velocities of the two rays, in terms of those quantities which we may regard as fundamental data, and the situation of the ray within the medium. The axes of elasticity themselves may, perhaps, be regarded as mere resultants from the equations of Art. 1000, and determined from other remoter data dependent on the fundamental lines in the crystalline form, and the intensity and distribution of the molecular forces within it. Accordingly, Dr. Brewster considers the optic axes as the resultants of others which he terms *polarizing axes*, and from which he conceives to *emanate* polarizing forces producing the phenomena of the rings and of the double refraction and polarization observed. We shall not here stop to examine into the propriety of these terms. The reader who may have doubts on the subject will, in what follows, mentally substitute other and more general phrases in their place expressive of relation and causality, while we proceed to state the assumptions with which he sets out, and the conclusions he very ingeniously deduces from them.

1021. *Postulate 1.* A polarizing axis, when single, has the characters of an axis of no double refraction, and is coincident with the axis of the Huygenian spheroid in such crystals as have but one. A positive axis acts as the axis in quartz, &c. may be supposed to do, and a negative, as that of carbonate of lime, &c.

1022. *Post. 2.* The polarizing force of a single axis in any medium is proportional to, and measured by, the tint developed in the ordinary and extraordinary pencils into which a doubly refracting prism analyzes a polarized ray, which has traversed a given thickness of the medium.

1023. *Corol. 1.* The polarizing force of a single axis in the same medium is as the square of the sine of the angle made by the ray traversing it internally, with the axis.

1024. *Corol. 2.* The same force is also inversely as the thickness necessary to be traversed at a given angle to develop the same or equal tints. This may be regarded as the *intrinsic* polarizing force or *intensity* of the axis.

1025. *Post. 3.* When two axes exist in one medium and operate together, they polarize a tint whose measure (see Art. 906) is the diagonal of a parallelogram whose sides measure, on the same scale, the tints which would be polarized by either, separately, and include between them an angle *double* of the mutual inclination of two planes passing through the ray and either axis respectively.

1026. *Corol. 1.* If t and t' be the numerical measures of the tints polarized by either of two axes separately, T that polarized by their joint action, and C the angle between the planes just described, the tint T will be given by the equation

$$T^2 = t^2 + 2 t t' \cdot \cos 2 C + t'^2.$$

1027. *Corol. 2.* If a and b represent the intensities of the axes, and α and β the angles which the ray makes with each respectively, we have $t = a \cdot \sin \alpha^2$; $t' = b \cdot \sin \beta^2$, and

$$T^2 = (a \cdot \sin \alpha^2)^2 + (b \cdot \sin \beta^2)^2 + 2 a b \cdot \sin \alpha^2 \cdot \sin \beta^2 \cdot (1 - 2 \cdot \sin C^2),$$

$$= \{ a \cdot \sin \alpha^2 + b \cdot \sin \beta^2 \}^2 - 4 a b \cdot \sin \alpha^2 \cdot \sin \beta^2 \cdot \sin C^2,$$

or else

$$T^2 = \{ a \cdot \sin \alpha^2 - b \cdot \sin \beta^2 \}^2 + 4 a b \cdot \sin \alpha^2 \cdot \sin \beta^2 \cdot \cos C^2.$$

1028. If γ be the angle contained between the polarizing axes, since α, β, γ are the sides of a spherical triangle, and C the angle included between the sides a and b , or opposite to γ , we have

$$\cos C = \frac{\cos \alpha \cdot \cos \beta - \cos \gamma}{\sin \alpha \cdot \sin \beta},$$

and if this be written for $\cos C$ in the latter of the expressions above given for T^2 , we find on reduction

$$T^2 = \{ a \cdot \sin \alpha^2 + b \cdot \sin \beta^2 \}^2 - 4 a b \{ 1 - \cos \alpha^2 - \cos \beta^2 - \cos \gamma^2 + 2 \cdot \cos \alpha \cdot \cos \beta \cdot \cos \gamma \}.$$

1029. *Corol.* If the polarizing axes be at right angles to each other, $\gamma = 90^\circ$ and $\cos \gamma = 0$, and the expression for the compound tint becomes

$$T^2 = \{ a \cdot \sin \alpha^2 + b \cdot \sin \beta^2 \}^2 - 4 a b (\sin \alpha^2 - \cos \beta^2).$$

1030. *Proposition.* Two rectangular polarizing axes, either both positive or both negative, being given, two other axes, or fixed lines, may be found, such that calling θ and θ' the angles made with them respectively by a ray traversing a spherical portion of the medium, the tint polarized shall be proportional to $\sin \theta \cdot \sin \theta'$.*

Let $A C$ and $B C$ (fig. 199) be the two polarizing axes including a right angle, of which let $B C$ be the more powerful. Let $O C$ be a ray penetrating the crystal in that direction; and in a plane $P C Q$ perpendicular to $A C B$, draw any two lines $P C, Q C$, making equal angles with $B C$, either of which we will represent by x . Then if a sphere about C as a centre be conceived, it will intersect the planes $A C B, P C Q, O C A, O C B, O C P, O C Q$ in lines of great circles $B A, P B Q, O A, O B, O P, O Q$, and we shall have $P B = Q B = x$, $O A = \alpha$, $O B = \beta$, $O P = \theta$, $O Q = \theta'$; and by Spherical Trigonometry, from the triangle $O B P$, we have

$$\cos O B P \left(= \sin O B A = \sin A O B \cdot \frac{\sin O A}{\sin A B} = \sin \alpha \cdot \sin C, \text{ since } A B = 90^\circ \right)$$

$$= \frac{\cos \beta \cdot \cos x - \cos \theta}{\sin \beta \cdot \sin x},$$

* M. Biot appears to have first noticed the fact announced in this proposition, viz. that Dr. Brewster's hypothesis of polarizing axes leads to a result *mathematically* identical with his own elegant law of the product of the sines. He has, however, suppressed his demonstration. Dr. Brewster's verification of this coincidence of results seems to have been founded on a numerical comparison of Biot's experiments on sulphate of lime with his own theory.

Light, and therefore

$$-\cos \theta = \sin \alpha \cdot \sin \beta \cdot \sin x \cdot \sin C - \cos \beta \cdot \cos x,$$

and similarly from the triangle O B Q, since O B Q = 90° + O B A, we obtain a second relation

$$+\cos \theta' = \sin \alpha \cdot \sin \beta \cdot \sin x \cdot \sin C + \cos \beta \cdot \cos x;$$

and, adding and subtracting, (putting, for brevity's sake, $\cos \theta' = p$, $\cos \theta = q$.)

$$p + q = 2 \cdot \cos \beta \cdot \cos x; \quad p - q = 2 \cdot \sin \alpha \cdot \sin \beta \cdot \sin x \cdot \sin C.$$

These equations express the geometrical relations subsisting between the lines P C, Q C, and the axes A C, B C; and, if combined with the equations of Art. 1028 and 1029, suffice to eliminate α , β , and C, and to express T in terms of x , θ , and θ' alone. To execute this, we have by the equations just demonstrated

$$\left(\frac{p+q}{2 \cdot \cos x}\right)^2 = \cos \beta^2; \quad \left(\frac{p-q}{2 \cdot \sin x}\right)^2 = \sin \alpha^2 \cdot \sin \beta^2 \cdot \sin C^2;$$

and in the latter, putting $1 - \cos C^2$ for $\sin C^2$, and for $\cos C^2$ its value given by Art. 1028, which, since $\gamma = 90^\circ$, becomes simply

$$\sin \alpha^2 \cdot \sin \beta^2 \cdot \cos C^2 = \cos \alpha^2 \cdot \cos \beta^2,$$

we have

$$\begin{aligned} \left(\frac{p-q}{2 \cdot \sin x}\right)^2 &= \sin \alpha^2 \cdot \sin \beta^2 - \cos \alpha^2 \cdot \cos \beta^2, \\ &= \sin \alpha^2 - \cos \beta^2. \end{aligned}$$

Hence we get, for the values of $\sin \alpha^2$ and $\sin \beta^2$,

$$\begin{aligned} \sin \alpha^2 &= \left(\frac{p+q}{2 \cdot \cos x}\right)^2 + \left(\frac{p-q}{2 \cdot \sin x}\right)^2, \\ \sin \beta^2 &= 1 - \left(\frac{p+q}{2 \cdot \cos x}\right)^2; \end{aligned}$$

and, substituting these in the equation of Art. 1029,

$$T^2 = \left\{ b + \frac{a-b}{4 \cdot \cos x^2} (p+q)^2 + \frac{a}{4 \cdot \sin x^2} (p-q)^2 \right\}^2 - \frac{a b}{\sin x^2} (p-q)^2.$$

Such is the general form of the expression for the tint, when referred to arbitrary axes in the manner here supposed, and it is complicated enough; but if we fix the position of the new axes so as to make $\sin x = \frac{a}{b}$

the complication disappears; we have then $\frac{a}{4 \cdot \sin x^2} = \frac{b}{4}$, and $\frac{a-b}{4 \cdot \cos x^2} = \frac{b}{4}$, so that the value of T^2 reduces itself to

$$\begin{aligned} T^2 &= b^2 \left\{ \left(1 - \left(\frac{p+q}{2}\right)^2 + \left(\frac{p-q}{2}\right)^2\right)^2 - (p-q)^2 \right\}, \\ &= b^2 \{ (1-pq)^2 - (p-q)^2 \} = b^2 \{ 1 - p^2 - q^2 + p^2 q^2 \}, \\ &= b^2 (1-p^2) (1-q^2) = b^2 \cdot \sin \theta^2 \cdot \sin \theta'^2, \end{aligned}$$

restoring the values of p and q , or $\cos \theta'$ and $\cos \theta$, consequently

$$T = -b \cdot \sin \theta \cdot \sin \theta'.$$

The negative sign is prefixed for the reason stated further on in Art. 1034.

Thus we see that the combined action of the two axes in the manner here supposed, on Dr. Brewster's principles, will give rise to a series of isochromatic lines arranged in the form of spherio-lemniscates about two poles P, Q, determined by the condition

103.

$$\sin B P = \sin B Q = \sqrt{\frac{\text{intensity of the feebler axis}}{\text{intensity of the stronger}}};$$

and the lines C P, C Q so determined have therefore the character of the optic axes in biaxial crystals, and may be designated with Dr. Brewster by the name of resultant axes. We must be careful, however, not to confound a resultant with a polarizing axis in this theory.

If the polarizing axes be not of the same denomination, as if one be positive and the other negative, the value of $\sin B P$ becomes imaginary, and the tints cannot be so arranged. But if we suppose the new axes to lie in this case in the same plane with the polarizing ones, as in fig. 200, all other things remaining, we have here

1032. Combination of a positive with a negative axis.

$$\cos O B A = + \cos O B Q, \quad \text{and} \quad \cos O B A = - \cos O B P,$$

but

$$\cos O B A = - \frac{\cos \alpha}{\sin \beta}, \quad \text{and} \quad \cos O B Q = \frac{\cos \beta \cdot \cos x - \cos \theta'}{\sin \beta \cdot \sin x},$$

so that we find

$$\cos \theta' = p = \cos \beta \cdot \cos x + \cos \alpha \cdot \sin x;$$

Light. and similarly

$$\cos \theta = q = \cos \beta \cdot \cos x - \cos a \cdot \sin x,$$

whence, by adding and subtracting, we get at once

$$\cos a = \frac{p - q}{2 \cdot \sin x}; \quad \cos \beta = \frac{p + q}{2 \cdot \cos x},$$

which, substituted in the value of T^2 , give

$$T^2 = \left\{ (a + b) - \left(\frac{a}{\sin x^2} + \frac{b}{\cos x^2} \right) \frac{p^2 + q^2}{4} + \left(\frac{a}{\sin x^2} - \frac{b}{\cos x^2} \right) \frac{pq}{2} \right\} \\ - 4ab + \frac{ab}{\sin x^2 \cdot \cos x^2} (p^2 + q^2) + \frac{2ab (\sin x^2 - \cos x^2)}{\sin x^2 \cdot \cos x^2} pq.$$

Now, if in this we suppose $\frac{a}{\sin x^2} + \frac{b}{\cos x^2} = 0$, or $\tan x^2 = -\frac{a}{b}$, it will, on substitution and reduction, take the form

$$T^2 = \frac{(1 - p^2)(1 - q^2)}{\cos x^4} \cdot b^2 = \frac{b^2 \cdot \sin \theta^2 \cdot \sin \theta'^2}{\cos x^4}$$

and

$$T = \frac{-b}{\cos x^2} \cdot \sin \theta \cdot \sin \theta';$$

that is, restoring the value of x , (since $\tan x^2 = -\frac{a}{b}$, and therefore $\cos x^2 = \frac{b}{b - a}$), finally,

$$T = -(b - a) \cdot \sin \theta \cdot \sin \theta'.$$

1033.
Position of
resultant
axes in
this case.

Thus, in this case also, the isochromatic lines are sphero-lemniscates, and the only difference is that their poles lie now in the plane of the polarizing axes, instead of at right angles to it; and that whereas in the former case the semi-angle between them (x) was given by the equation $\sin x = \sqrt{\frac{a}{b}}$, that is, $\cos x =$

$$\sqrt{\frac{b - a}{b}}, \text{ in this it is given by the equation } \cos x = \sqrt{\frac{b}{b - a}}.$$

1034.
Cases of the
resolution
of a single
axis into
two.

Corol. 1. In the case when $a = b$, or when the two polarizing axes are of the same denomination and of equal intensity, we have $\sin x = 1$, or $x = 90^\circ$, so that the angle between the resultant axes being 180° , they form one straight line, the lemniscates become circles, and the single resultant axis has now the characters of a polarizing axis. Hence, *vice versa*, a single polarizing axis, in any direction, may be resolved into two others equal in intensity, at right angles to it and to each other, and of an opposite denomination to the resolved axis. This follows from the negative sign of T , which is prefixed in extracting the square root in Art. 1030 and 1032; because in the case supposed, when the arc AB is 90° the angle C or AOB is necessarily greater than 90° , and $2C$ the angle of the parallelogram of tints $> 180^\circ$; so that the diagonal will be to be measured backwards through the angle, or must be a negative quantity.

1035.
Composition
of three
equal rect-
angular
axes.

Corol. 2. Since a single axis is equivalent to two equally intense axes of an opposite character at right angles to it and to each other, if we superadd to both another equal axis also of the opposite kind, and in the direction of the first, this will destroy the effect of the first, and therefore the combination of three equal and similar axes arising on the other side at right angles to each other, will be equivalent to none at all. Thus, three equal rectangular axes of the same character destroy each other's effects. This is Dr. Brewster's account of the want of polarization and double refraction in crystals whose primitive form is the cube, regular octohedron, &c., and whose secondary forms indicate a perfect symmetry in their molecules with respect to three rectangular axes.

1036.

There is no necessity to pursue further the general subjects of this species of composition of axes and of tints. Indeed, it appears to us that the rule for the parallelogram of tints, as laid down by Dr. Brewster, becomes inapplicable when a third axis is introduced; for this obvious reason, that when we would combine the compound tint arising from two of the axes (A, B) with that arising from the action of the third (C), although the sides of the new parallelogram which must be constructed are given, (*viz.* the compound tint T , and the simple tint t''), yet the wording of the rule leaves us completely at a loss what to consider as its angle, inasmuch as it assigns no single line which can be combined with the axis C in the manner there required, or which *quoad hoc* is to be taken as a resultant of the axes A, B . For further information therefore on this subject we shall content ourselves with referring the reader to his original Paper in the *Transactions of the Royal Society*, 1818.

§ X. Of Circular Polarization.

1037.

The first phenomena referable to the class of facts to whose consideration this section will be devoted, were noticed by M. Arago in his Memoir published among those of the Institute for 1811 on the colours of crystalized plates. He observed that when a polarized ray was made to traverse at right angles a plate of rock crystal

light. (quartz) cut perpendicularly to the axis of double refraction, on analyzing the emergent ray by a doubly refracting prism, the two images had complementary colours, and that these colours changed when the doubly refracting prism was made to revolve; so that in the course of a half revolution, the extraordinary image (for example) which at first was red, became in succession orange, yellow, yellow-green, and violet, after which the same series of tints would of course recur. It is evident that this is just what would take place, supposing the several coloured rays at their emergence from the rock crystal to be polarized in different planes; and to this conclusion M. Arago came in a second Paper, subsequently read to the Institute. The subject was resumed by M. Biot, in a Paper published in the *Mém de l'Inst.*, 1812; and his labours were completed in a second extremely interesting Paper read to that body in September, 1818.

Part IV.
Phœnomena
of circular
polarization.

When a polarized ray is made to traverse the axis of Iceland spar, beril, and other uniaxal crystals, we have seen that it undergoes no change or modification; and that when analyzed at its egress by a doubly refracting prism, having its principal section in the plane of primitive polarization, the ordinary image will contain the whole ray, or the complementary tints will be white and black. Quartz, however, is an exception to this rule. A polarized ray transmitted, however precisely, along its axis, is still coloured and subdivided, and that the more evidently, the *thicker* is the plate. If we place on a proper apparatus, such as that described in Art. 929 and figured in fig. 189, a very thin plate of this body, and turn round the analyzing prism M in its cell, till the extraordinary image is at its minimum of brightness, it will in this position have a sombre violet, or purple tinge, because the yellow or most luminous rays, which are complementary to purple, are now completely extinguished. Let the angle of rotation of the prism in its cell, measured on the divided circle R, and which in this case will be small, be noted; and then let the rock crystal plate be detached, and another cut from the same crystal, but of twice the thickness, be substituted. The tint of the extraordinary image will no longer be violet; but if the prism be made to revolve through an additional equal arc in the same direction, the violet or purple tint will be restored, and the minimum of brightness attained; and, in general, if the thickness of the plate (always supposed cut from the same crystal) be greater or less in any ratio, the angle of rotation through which the prism must be moved in the same direction, to produce a minimum of intensity and a purple tint in the extraordinary image, is increased or diminished in the same ratio. In consequence, if the plate be sufficiently thick, one or more circumferences will be required to be traversed; and as only the excesses over whole circumferences can be read off, this may produce some confusion or doubt, unless we take care to use a succession of thicknesses so gradually increasing as not to allow of a *saltus* of a whole, or a half circumference.

1038.
Rotatory
phenomena
of quartz.

From this experiment we collect, that the plane of polarization of a mean yellow ray which has traversed the axis of a quartz plate, has been turned aside from its original position, through an angle proportional to the thickness of the plate; and, therefore, assumes at its egress a position the same as it would have, had it revolved uniformly in one direction, during every instant of the ray's progress through the plate. The same holds good for all the other homogeneous rays; but to prove it, we must abandon the use of white light, and operate with pure rays of the particular colour we would examine. If we use pure red light, for instance, or defend the eye with a pure red glass, the same will be observed, only that instead of a violet tint and a minimum of light, we shall have a total obliteration of the extraordinary pencil when the prism attains its proper position, thus proving, what in the former mode of observation might have been doubtful, that the polarization of the emergent ray is *complete*.

1039.
Rotation
of the plane
of polarization.

In examining in this way the quantity by which one and the same plate of quartz turns aside the planes of polarization of the different homogeneous rays, M. Biot ascertained that the *more refrangible rays are more energetically acted on than the less*, and have their planes of polarization deviated through a greater arc. According to this eminent philosopher, the constant coefficient, or index, which represents the velocity with which the plane of polarization may be conceived to revolve, is proportional to the square of the length of an undulation of the homogeneous ray under consideration; so that if we call λ the length of an undulation, and t the thickness of the plate, the deviation produced will be equal to $k \cdot \lambda^2 t$, k being a certain constant. The

1040.
Law of ro-
tation of the
plane of
simple
coloured
rays.

value of this constant he assigns at $\frac{18^\circ.414}{(6.18614)^2}$, when t is reckoned in millimètres; and the following is stated by him as the numerical amount of the deviations in degrees (sexagesimal) produced by one millimètre of thickness of rock crystal on the several rays:

Designation of the homogeneous ray.	Arc of rotation corresponding to one millimètre.
Extreme red	17°.4964
Limit of red and orange.....	20°.4798
Limit of orange and yellow ..	22°.3138
Limit of yellow and green....	25°.6752
Limit of green and blue	30°.0460
Limit of blue and indigo	34°.5717
Limit of indigo and violet....	37°.6829
Extreme violet	44°.0827

Light.

1041.
Right and
left handed
quartz.

In the course of these researches M. Biot was led to the very singular discovery of a constant difference subsisting in different specimens of rock crystal, in the *direction* in which this rotation or angular shifting of the plane of polarization of a ray traversing them takes place. In some specimens it is observed to be from right to left, in others from left to right. To conceive this distinction, let the reader take a common cork-screw, and, holding it *with the head towards him*, let him turn it in the usual manner, as if to penetrate a cork. The head will then turn the same way with the plane of polarization of a ray in its progress *from* the spectator through a *right-handed* crystal may be conceived to do. If the thread of the cork-screw were reversed, or what is termed a *left-handed thread*, then the motion of the head as the instrument advanced would represent that of the plane of polarization in a left-handed specimen of rock crystal. It will be observed, that we do not here mean to say that the plane of polarization *does* so revolve in the interior of a crystal, but that the ray at its egress presents the same phenomena as to polarization *as if* it had done so. This is necessary, for we shall see presently that a very different view of the subject may be taken.

1042.
Phenomena
of plagie-
dral crystals

In crystals which present this remarkable difference, when cut and polished, and when the external indications of crystalline form are obliterated, no other difference can be detected. Their hardness, transparency, refractive and double refractive powers are the same; and, with the exception of the direction in which it takes place, their effects in deviating the planes of polarization of the rays which traverse them are alike. Experiments subsequent to M. Biot's researches have, however, established, as a result of extensive induction, a very curious connection between this direction and the crystalline forms affected by individual specimens. In the variety of crystallized quartz, termed by Haüy, *Plagiedral*, there occur faces which (unlike those in all the more common varieties) are unsymmetrical related to the axes and apices of the primitive form, whether regarded as the rhomboid or bipyramidal dodecahedron. Fig. 201 represents such a crystal, in which when the apex A is set upwards, the faces C, C, C, are observed to lean all in one direction, *viz.* to the right, with respect to the axis, as if distorted from a symmetrical position by some cause acting from left to right all round the crystal. When the vertex B is set upwards, the same distortion, and in the same direction, is observed in the plagiedral faces D, D, D, and crystals of quartz are excessively rare, if they exist at all, in which two plagiedral faces leaning opposite ways occur. Now it has been ascertained, that in crystals where one or more of these faces, however minute and even of microscopic dimensions, can be seen, we may thence predict with certainty the direction of rotation in a plate cut from it, which is always that in which the plagiedral face appears to lean with respect to an observer regarding it as the reader does the figure, which represents a *right-handed* crystal. Hence we are entitled to conclude, that whatever be the cause which determines the direction of rotation, the same has acted in determining the direction of the plagiedral faces. Other crystallized minerals, as apatite, &c. also present plagiedral and unsymmetrical faces; but, independent of their extreme rarity, they are not possessed of the property of rotation; so that at present we are unable to say whether this curious law be general, or to conjecture to what principles it will hereafter prove to be referable.

1043.
Superposi-
tion of
plates of
rock
crystal.

When two plates of rock crystal are superposed, if they be both right-handed or both left, their joint rotatory effect will be the sum of their respective ones, *i. e.* each ray's plane of polarization will be shifted through an angle equal to the sum of those through which it would have been shifted by their separate actions. If their characters be opposite, it will be their difference, *i. e.* the *index of rotation* in a right-handed crystal being regarded as positive, it will be negative in a left-handed one.

1044.
Amethyst.

The amethyst (and, possibly, also the agate in some cases) presents the very remarkable and curious phenomenon of these two species of quartz crystallized together in alternate layers of very minute thickness. Accordingly, when a crystal of amethyst is cut at right angles to the axis, and examined by polarized light transmitted exactly along the axis, and analyzed as usual, it offers a striped or fringed appearance, as represented in fig. 202, variegated with different colours, according to the different planes of polarization assumed by the rays emergent at its several points, and presenting, according to the distribution of its elements, the most beautiful combinations and contrasts of coloured fasciæ and spaces. For a particular account of these phenomena, the reader is referred to a Paper by Dr. Brewster, (*Edinburgh Transactions*, vol. xi.) who first observed and publicly described them, though we have reason to believe them to have been known to others by independent observation previous to the publication of his very curious and interesting Memoir. The layers may be distinctly seen *cropping out* to the surface in a fresh fracture of the mineral, and imparting that peculiar undulated fracture which is the chief mineralogical character of this substance by which it is known from ordinary quartz.

1045.
Rotatory
phenomena
in liquids.

But the phenomena of rotation as above described are not confined to quartz. Many liquids, and even vapours exhibit it, a circumstance which would seem very unexpected, when we consider that in liquids and gases the molecules must be supposed unrelated to each other by any crystalline arrangement, and independent of each other; so that to produce any such phenomena, each individual molecule must be conceived as unsymmetrically constituted, *i. e.* as having a right and a left side. M. Biot and Dr. Seebeck appear about the same time to have made this singular and interesting discovery; but the former has analyzed the phenomena with particular care, and it is from his Memoir above cited that we extract the following statements. The liquids in which he observed a right-handed rotatory property, according to our sense of the word above explained, in which the observer is supposed to look in the direction of the ray's motion, are oil of turpentine, oil of laurel, vapour of turpentine oil, and an alcoholic solution of artificial camphor produced by the action of muriatic acid on oil of turpentine. The left-handed rotation was observed by him in oil of lemons, syrup of cane sugar, and alcoholic solution of natural camphor. In all these, the intensity of the action, or the velocity of rotation, was much inferior to quartz. The following are their *indices of rotation*, or the arcs of rotation produced by one millimètre of thickness in the plane of polarization of a certain homogeneous red ray chosen by M. Biot for a standard, as calculated from his data.

Right-handed.	Index of rotation.	Left-handed.	Index of rotation.	Part IV.
Rock crystal	+ 18°.414	Rock crystal.....	- 18°.414	
Oil of turpentine	+ 0°.271	Oil of lemon	- 0°.436	
Ditto, another specimen	+ 0°.251	Concentrated syrup of sugar	- 0°.554	
Ditto, purified by repeated distillations	+ 0°.286			
Oil of laurel.....				
Solution of 1753 parts of artificial } camphor in 17359 of alcohol .. }	+ 0°.018			

It follows further from M. Biot's researches, that when any two or more liquids are mixed together, or combined with plates of rock crystal, the rotation produced by the compound medium will be always the sum of the rotations produced by the several simple ones, in thicknesses equal to their actual thicknesses present in the combination, the thicknesses in mixed liquids being assumed in the ratio of the volumes of each respectively mixed; so that calling T the compound thickness, and R the resulting index of rotation, we shall always have

$$R \cdot T = r \cdot t + r' \cdot t' + r'' \cdot t'' + \&c.$$

where r , r' , &c. are the indices (with their signs) of the elementary ingredients, and t , t' , &c. their thicknesses. Thus, when 66 parts by measure of oil of turpentine, having the index + 0.258 are made to act against 38 of oil of lemon, we have

$$+ 66 \times 0.251 - 38 \times 0.436 = 0.002,$$

so that these thicknesses ought almost exactly to compensate each other; and such was, in fact, the result of M. Biot's experiment, the whole pencil transmitted being found to retain its primitive polarization without the least trace of an extraordinary image. Again, when into two tubes of the same bore, but of very unequal lengths, equal quantities of oil of turpentine were poured, and the rest of their lengths filled with sulphuric ether, which has no rotatory property, or in which $r = 0$, the two compound thicknesses thus differently constituted gave identically the same tints in all positions of the analyzing prism. Thus we see that dilution or mixture which only separate, without decomposing the molecules, do not alter their rotatory power. Nay, even when reduced to vapour, M. Biot found, that oil of turpentine still preserved its property and peculiar character; and, had not the explosion of his apparatus prevented accurate measures, would probably enough have been found to retain the same index of rotation allowing for the change of density. From these circumstances he concludes that the rotatory power is essentially inherent in the molecules of bodies, and carried with them into all their combinations. But this is too rapid a generalization; for neither sugar nor camphor in the solid state possess this property, though examined for it in the same circumstances as quartz is, by transmitting the polarized ray along their optic axes; and, on the other hand, quartz held in solution by potash, or (as Dr. Brewster has found) melted by heat, and thus deprived of its crystalline arrangement, manifests no such property. This obscure part of chemical optics well deserves additional attention.

M. Fresnel's researches have been directed to the rotatory phenomena with the same brilliant success which has distinguished his other inquiries into the nature of light; and he has shown that they may be explained by conceiving the molecules of the ether, which propagate rays along the axis of quartz, or rotatory fluids, instead of vibrating in straight lines, to revolve uniformly in circles, in the manner explained in Art. 627, (where we have shown (Corol.) that such a mode of vibration may subsist, and must arise from the interference of two rectangular vibrations of equal amplitude, but differing in phase by a quarter undulation,) and by admitting that, in virtue of some peculiar mechanism in the molecules of the media in question, such circular vibrations, when performed from right to left, bring into play an elasticity slightly different from that which propagates them forward when performed in the contrary direction. The colours produced by such media he conceives to originate in the interference of two pencils thus *circularly polarized*, and lagging the one behind the other by an interval of retardation proportioned to their difference of velocities.

But to make this last hypothesis admissible, it is incumbent on us to show that the phenomenon which necessarily accompanies a difference of velocities, *viz.* a bifurcation of the pencil in the act of refraction at oblique surfaces, really takes place. This has accordingly been shown by M. Fresnel, by an experiment which, though of great delicacy, is decisive and satisfactory. From a crystal of quartz he procured to be cut a prism having its refracting angle 150° , and its faces equally inclined to the axis; so that a ray traversing it internally parallel to its axis should be incident at equal angles, *viz.* of 75° on either face. As this is too great to allow of the ray's egress, he cemented on the surfaces the two halves of another precisely similar prism cut from another rock crystal of an opposite rotatory character. Thus in fig. 203, ACB is the first prism, and the side CB of the second prism CBE being cemented on to CB, this prism is bisected by the plane BD, and the half of it DBE transferred to the other side, and cemented with its side BC in contact with AC, thus producing the achromatic parallelepiped FABD; so that if a ray be incident on Q in the direction PQ parallel to the base AB, *i. e.* to the axis of the two crystals, it will traverse all three in the direction of the axes of their spheroids of double refraction; and, therefore, so far as the Huygenian law of double refraction is concerned, ought to undergo no division. Now it is evident, that if the ray PQ be at its entry into AFC divided into two circularly polarized in opposite directions, the one (R) moving quicker than the other (L,) then, at quitting the surface AC, a bifurcation must take place, the ray R being *least*, and L *most* refracted. In this state they are incident on the medium ACB, and now the portions R and L, by reason of the opposite nature of the media, exchange velocities; so that R, which at its emergence from the face AC of FAC was *least* refracted upwards, will now

1046.
Law of
rotation in
mixtures.

1047.
Fresnel's
theory of
circular po-
larization.

1048.
Peculiar
double re-
fraction
produced
by circu-
larly polari-
zing media.

Light. be most refracted downwards; and thus the separation of the images will be doubled, and the same will take place at the common face C B. Thus this combination, both from the doubling of the separation, and the greatness of the angles of incidence, is peculiarly well adapted to render sensible any bifurcation, or difference of velocities, however small, which may exist along the axis. Accordingly, with the compound prism, so constructed, a double refraction is produced; and the two rays are really observed to emerge, making a sensible angle with each other.

1049. But it is, moreover, observed, that though thus separated by a real double refraction, the two pencils have not acquired the characters which double refraction usually impresses on the ordinary and extraordinary rays, at their emergence, but very different ones. In common cases of double refraction the two emergent pencils are each wholly polarized in opposite planes, and either of them when examined with a doubly reflecting prism gives two unequal images, one alternately more and less bright than the other, as the prism revolves through successive quadrants. This is not the case with the two pencils produced in the case before us.

Characters of circular polarization.

First, Either of them, when examined with a doubly refracting prism, gives constantly two images of equal intensity, in whatever plane the principal section of the latter be placed. In this respect, then, they present the characters of unpolarized light, and may be regarded as each consisting of two rays polarized at right angles to each other. But

1050. Secondly, They differ from ordinary, or unpolarized light, in a very remarkable property, which was first discovered by Fresnel, and is a chief distinctive character of this kind of polarization. Suppose either of them to be incident at right angles on the surface A B of a parallelepiped of crown glass of the refractive index 1.51, having its angles A B C and A D C each $54\frac{1}{2}^\circ$, it will then be totally reflected at the internal surface B C; and (if the parallelepiped be long enough) again in the same plane at the opposite surface A D, and will emerge at length perpendicularly through the surface B C. But the emergent ray, instead of comporting itself as ordinary light, will now be found to be *completely polarized* in a plane 45° inclined to that in which the reflections were made, whatever may have been the position of that plane. If both the pencils be treated in this manner, it will be found that the one, after its two total reflexions will assume a plane of polarization 45° in azimuth to the right, and the other 45° to the left of the plane of the reflexions.

Other characters of circularly polarized rays.

Thus we see that the effect of double refraction along the axis of quartz is to impress on either of the emergent pencils opposite polarizations, or modifications, of a nature totally distinct from that given to a ray by ordinary reflexion, or by double refraction through Iceland spar, &c.; and, as in the last described experiment, so long as the ray enters perpendicularly into the first surface of the glass parallelepiped, it is indifferent in what plane the two reflexions are operated, and since when presented to a doubly refracting prism in any plane indifferently it always divides itself into two equal pencils, it is evident that the ray thus modified has *no sides*, i. e. no particular relations to certain regions of space; and therefore that the epithet *circular polarization*, apart from all theoretical considerations, may be naturally applied to this peculiar modification. But the characters above described are not the only ones belonging to a ray thus modified, for

Thirdly, Such a ray being transmitted through a thin crystallized lamina, and parallel to its axis, is divided by subsequent double refraction into two rays of complementary colours, thus marking a decided difference between it and a ray of common light; while, on the other hand, these colours are not the same with those which would arise from a ray of light polarized in the usual way and similarly analyzed, but differ from them by an exact quarter of a tint, either in excess or defect, as the case may be.

1051. Fourthly, A ray so modified by this peculiar double refraction, when transmitted again along the axis of rock crystal, or through columns of oil of turpentine, of lemons, &c., and then analyzed by a double refracting prism, gives rise to no phenomena of colour, differing in this from polarized, and agreeing with common light.

1052. Another independent mode of impressing on a ray all this assemblage of characters has been discovered by M. Fresnel. It consists in inverting the process described in Art. 1049. Thus, into the side C D of the glass parallelepiped there mentioned, let a common polarized ray be introduced at a perpendicular incidence, the parallelepiped being so placed that the plane of internal reflexion at the side A D shall be 45° inclined to that of its primitive polarization. Then, after undergoing two total internal reflexions at G and F, it will emerge at E deprived of its characters of ordinary polarization and endowed with those of circular, and being no way distinguishable from one of the pencils produced by double refraction along the axis of rock crystal.

Another mode of producing circular polarization

1053. It remains to show, however, that the characters here described, as impressed on a ray by transmission along the axis of rock crystal, are really those which ought to belong to a ray propagated by circular vibrations. And, first, it follows from Art. 627, that this latter ray is the resultant of two rays polarized at right angles, and differing in their phases by a quarter undulation. It must, therefore, of necessity possess the first character, viz. that of division into two equal pencils by double refraction in any plane, for the same reason that unpolarized light is so divided, the difference of phases having nothing to do with this character.

1054. In the next place, a ray propagated by circular vibrations when incident on rock crystal in the direction of the axis, will (by hypothesis) be propagated along it by that elasticity which is due to the direction of its rotation, the wave then will enter the crystal without further subdivision, and there will be no difference of paths, or interfering rays at its emergence; and, of course, no colours produced on analyzing by double refraction, which is another of the characters in question.

1055. When a ray propagated by circular vibrations is incident on a crystallized lamina it may be regarded as composed of two, one polarized in the plane of the principal section, the other at right angles to it, of equal intensity, and differing in phase by a quarter undulation. Each of these will be transmitted unaltered, and therefore at their emergence and subsequent analysis will comport themselves in respect of their interferences, just as would do the two portions of a ray primitively polarized in azimuth 45° , and divided into two by the

Tints produced by circularly polarized rays.

Light.

double refraction of the lamina, provided that a quarter undulation be added to the phase of one of these latter rays. Now such rays will, as we have shown at length in Art. 969, produce by the interference of their doubly refracted portions, the ordinary and extraordinary tints due to the interval of retardation within the crystallized lamina. Hence, in the present case, the tints produced will be those due to that interval, *plus* or *minus* the quarter of an undulation added to, or subtracted from, the phase of one of the portions; and, consequently, will differ one-fourth of a tint, or order, from that which would arise from the use of a beam of ordinary polarized light incident in azimuth 45° on the lamina.

Part IV.

There remains but one more character of the rays transmitted along the axis of quartz, which we must show to belong to a ray propagated by circular vibrations, *viz.* that described in Art. 1049. But in order to this it will be necessary to state the result of M. Fresnel's researches on the modifications which light undergoes by total reflexion in the interior of transparent bodies.

1056.
Modifica-
tions im-
pressed on
light by
total
reflexion.

When a ray polarized in any azimuth is incident on a reflecting surface which reflects the whole of the incident light, if we decompose it into two, the one having its vibrations performed parallel, and the other perpendicular to the surface, and regard each of these as independent of the other; it is evident that the reflexion of these portions will be performed under very different circumstances, the ethereal molecules having in the former case to glide as it were on the surface, and therefore parallel to the strata in which their density is constant, while in the latter each molecule in the act of vibration will pass into strata of variable density. The reflexions therefore will be performed at different depths in the two cases; and from this cause will arise a difference of route, and a consequent difference of phase in the reflected portions, so that the total reflected ray will no longer be capable of being regarded as one having a single origin, but as two of unequal intensities, oppositely polarized, and differing in phase by a quantity depending on the angle of incidence and the refractive power of the medium. From peculiar considerations, of a delicate nature, and depending on a discussion of the imaginary forms assumed by the general expressions for the intensity of a ray reflected at any angle (Art. 852) when applied to the case of total reflexion, M. Fresnel has been led to the following expression for the difference of phases (δ) of the two portions in question.

$$\cos \delta = \frac{2 \mu^2 \cdot (\sin i)^4 - (\mu^2 + 1) \cdot (\sin i)^2 + 1}{(\mu^2 + 1) (\sin i)^2 - 1},$$

where μ is the index of refraction, and i the angle of internal incidence. This formula, it is to be observed, is given by him, not as strictly demonstrated, but merely as highly probable, as an interpretation of the analytical meaning of the imaginary formula alluded to. The mode of its deduction being, however, independent of experiment, and entirely *a priori*, it is clear that if found verified by careful experiment in circumstances properly varied, it may be received as a physical law, like any other result of the same kind. Now we have already seen, that in the case of crown glass, where $\mu = 1.51$ and $i = 54\frac{1}{2}^\circ$, a polarized ray, having its azimuth 45° , reckoned from the plane of total reflexion, has its polarization destroyed, and becomes resolved into a ray having the other characters of a resultant from two differing 45° in phase, by two total reflexions at this angle, (Art. 1056.) But if in the above formula we make $\mu = 1.51$, and $i = 54^\circ 37'$, we shall find $\delta = 45^\circ$, and $2\delta = 90^\circ$, so that the above equation is verified in this case. M. Fresnel also found that the same effect was produced by three reflexions when the angle of incidence was $69^\circ 12'$, and by four when $74^\circ 42'$, both agreeing with the formula which gives in the former case $\delta = \frac{1}{3} 90^\circ$, and in the latter $\delta = \frac{1}{4} 90^\circ$, for the difference of phase gained or lost by one portion on the other at each reflexion. Similar verifications were obtained by performing two reflexions at the internal surface of glass, and two at the confines of glass and water at angles of $68^\circ 27'$.

It appears, then, that when a ray polarized in azimuth 45° undergoes two total reflexions at the angles, and in the manner described, it becomes circularly polarized; and if *vice versa*, the two elements of a ray so circularly polarized be made to retrace their course, they will reunite into a ray polarized completely in one plane. Thus we see, that all the characters of the rays transmitted along the axis of rock crystal agree with those of a ray so compounded, and possessing *circular polarization*. In order, then, to explain the phenomena presented by a polarized ray when incident on a plate of this substance cut at right angles to its axis, we must first regard the ray as resolved into two others (which we will call A and B) of equal intensity; the one A polarized in a plane 45° inclined to the *right*, the other 45° inclined to the *left* of the vertical, (which, to fix our ideas, we shall take for the plane of primitive polarization.) Now, since by Art. 615 a ray polarized in any plane may be regarded as equivalent to two rays each of half its intensity, differing in their phases by a quarter undulation, let us conceive the ray A as resolved into two, Aa polarized in the plane $+45^\circ$, and having its phase advanced $+\frac{1}{8}$ undulation, and another Ab also polarized at $+45^\circ$, but having its phase retarded, or $-\frac{1}{8}$ undulation, so that Aa and Ab differ $\frac{1}{4}$ undulation in their phases. Similarly, let B be regarded as decomposed into Ba polarized at -45° , and having its phase $+\frac{1}{8}$ undulation, and Bb polarized also at -45° , but having its phase $-\frac{1}{8}$ undulation different from B. Thus will the original ray be resolved into the four Aa, Ab, Ba, Bb. Now, let us combine these two and two in a cross order, then Aa combined with Bb will be equal rays, polarized in opposite planes, and differing $\frac{1}{4}$ undulation in their phases, and will therefore compose one circularly polarized ray, in which the rotation is from *right* to *left*. Similarly, the pair Ab, Ba will compound another equally intense circularly polarized ray having its rotation the contrary way. Now these will (*ex hypothesi*) be transmitted through the quartz with unequal velocities, and thus an interval of retardation will arise, and if the surface of egress or ingress be *oblique* to the axis, a double refraction will take place; and two circularly polarized rays will emerge in different directions, as experiments show they do. If *perpendicular* they will emerge superposed, and will compound one ray. Let us now examine what will be the character and state of polarization of this compound ray. To this end conceive a molecule of ether C to be at once agitated by two circular motions in opposite directions; one in a circle equal and similar to AP in

1057.
Explana-
tion of the
rotatory
phenomena.

Light. the direction A P, the other in a circle equal and similar to B Q, and in the direction B Q, fig. 205. Let A, B be two molecules setting out at once from A, B in these circles with equal velocities, then will the motion of C at any instant be equal to that compounded of the motions of A and B at that instant. When A comes to P let B come to Q, then arc $AP = BQ$, and the motions at P and Q will be each resolved into two, those of which parallel to C D (a perpendicular to P Q) conspire, while those in the directions P D and Q D parallel to P Q oppose, and being equal destroy each other; thus C will move only in virtue of the sum of the two former, and its vibrations will therefore be rectilinear, and in the plane C D perpendicular to P D Q. If the thickness of the plate of quartz were nothing, or such that the interval of retardation were an exact number of undulations, A, B would lie at opposite extremities of a diameter, and C D the new plane of polarization would be perpendicular to A M that diameter, or coincident with the plane of primitive polarization. But if not, the quicker motion will have gained on the other a part of a circumference M B, which is to a whole circumference as the thickness of the plate is to that which would produce a difference of a whole undulation; and at the emergence of the two waves into air, after which they circulate with equal velocity, if we suppose the one molecule to be setting out from A, the other will be setting out, not from M the opposite extremity of the diameter, but from B, and therefore C D the new plane of polarization (which from what has just been shown must always bisect the angle A C B) will no longer be coincident with C N the primitive plane of polarization, at right angles to A M, but will make an angle D C N with it equal to half B C M, and therefore proportional to M B, or to the interval of retardation, *i. e.* to the thickness of the plate. Thus the system of rays emerging from the rock crystal plate will compound one ray polarized in one plane, and in the position the original plane would have had, had it revolved uniformly round the ray as an axis during its passage through the plate. Thus we have a complete and satisfactory explanation of the apparent rotation of the plane of polarization, as observed by Biot in the case of a homogeneous ray.

1058. It is observed, that the spectra formed by the double refraction of rock crystal along its axis are very highly and unequally coloured. The violet rays are most *separated*, and therefore the difference of velocities of the two rotating pencils is much greater for violet than for red rays. Consequently, the apparent velocity of rotation of the plane of polarization will also be greater for the violet rays in the same proportion, and thus arise all the phenomena of coloration observed and described by M. Biot. It is scarcely possible to imagine an analysis of a natural phenomenon more complete, satisfactory, and elegant. With regard to the physical reason of the difference of velocity in the two circular polarized pencils within the quartz, it is true we remain in the dark; but the fact of such difference existing is now shown to be no hypothesis, but a fact demonstrated by their observed difference of refraction, and by the observed characters of the two emergent rays.

§ XI. Of the Absorption of Light by Crystallized Media.

1059. Crystallized media, endowed with the property of double refraction, are found to absorb the differently coloured rays differently, according to their planes of polarization, and the manner in which these planes are presented to the axis of the crystal, and also to exert very different absolute absorbing energies on rays of one colour polarized in different planes. A remarkable instance of this has been already often referred to in the case of the brown tourmaline, a plate of which, cut parallel to the axis, absorbs almost entirely *all* rays polarized in the plane of the principal section, and lets pass only such among oppositely polarized rays as go to constitute a brown colour.

1060. When such a plate, then, is exposed to natural light, since at the entrance of each ray into its substance it is resolved into two, one polarized in the plane of the principal section, and one perpendicular to it, the former is absorbed in its progress by the action of the crystal, while the brown portion of the latter escaping absorption, but retaining at its egress the polarization impressed on it, after traversing the plate, appears with its proper colour, and *wholly polarized in a plane at right angles to the axis*. Thus the curious phenomenon of the polarization of light by transmission through a plate of tourmaline, or other coloured crystal, is explained, or at least resolved into the more general fact of an absorbing energy varying with the internal position of the plane of polarization. The crystal, in virtue of its double refractive property, divides the ray into two, and polarizes them oppositely; and the unequal absorption of these two portions *subsequently* causes the total suppression of one, and the partial of the other of the portions so separated. Thus we see that the polarized beam obtained by transmission through a tourmaline must always be of much less than half the intensity of the incident light.

1061. The destruction of the pencil polarized in the principal section is not, however, sudden; for if the plate of tourmaline be very thin, the emerging pencil will only be partially polarized, indicating the existence in it of rays belonging to the other pencil. This is best shown by cutting a tourmaline into a prism having its refracting edge parallel to the axis, and its angle small, so as to produce a wedge whose thickness increases not too rapidly. If we look through this at a distant candle, we shall see only one image, *viz.* the extraordinary through the back of the wedge, (if thick enough;) but as the eye approaches the edge, the ordinary image appears at first very faint, but increasing in intensity till, at the very edge, it becomes equal to the other. At the same time the colour of the latter, which at first was intense, becomes diluted; and the images approximate not only to equality of light, but to similarity of tint. We see by this, too, that in strictness the ordinary pencil is never *completely* absorbed by any thickness, however great; but as it diminishes in geometrical progression as the thickness increases in arithmetical, the absorption may for all practical purposes be regarded as total at moderate thicknesses.

Light.

The indefatigable scrutiny of Dr. Brewster, to whom we owe nearly all our knowledge on this subject, has shown that the same property is possessed in greater or less perfection by the greater number of coloured doubly refracting media; and the expression of the property may be rendered general by considering all doubly refractive media as possessing two distinct absorbing powers or two separate *scales of absorption* for the two pencils, or (adopting the language of § III. part 2) as having two distinct *types*, or curves expressing the law of absorption throughout the spectrum. If these types be both straight lines parallel to the abscissa, the crystal will be colourless. Such are limpid carbonate of lime, quartz, nitre, &c. If they be similar and equal curves, the medium, although coloured, will present the same colour, and the same intensity of tint, in common as in polarized light. If dissimilar, or if, although similar, their ordinates are in a ratio of inequality, the character, in the former case, and the intensity in the latter, will vary on a variation of the plane of polarization of the incident beam, so that if a plate cut from such a crystal be exposed to a beam of polarized white light, and turned round in its own plane, or otherwise inclined to the beam, its colour will change either in hue or depth or both. Dr. Brewster has remarked such change of colour and the phenomena connected with it in a great variety of crystals both with one and two axes, of which he has given a list in a most interesting Paper on the subject in the *Philosophical Transactions*, 1819, p. 1, which we strongly recommend to the reader's perusal. It may be familiarly seen in a prism of smoked quartz of a pretty deep tinge, which held with its axis in the plane of polarization appears of a purple or amethyst colour, while if held in a direction at right angles to this position, its colour is a yellow brown.

But in order to analyze the phenomena more exactly, we must examine the two pencils separately. To this end Dr. Brewster took a rhomboid of yellow carbonate of lime of sufficient thickness to give two distinct images of a small circular aperture placed close before it, and illuminated with white light, when he observed that the image seen by extraordinary refraction appeared of a deeper colour and less luminous than the other, being an orange yellow, while the ordinary image was a yellowish white. He found, moreover, that the difference of colour was greater as the paths of the refracted rays within the crystal were more inclined to the axis, being 0 when the rays passed along the axis, and a maximum when at right angles to it. If we denote by Y_o and Y_e the ordinates of the curves, expressing the law of absorption as in Art. 490, for the ordinary and extraordinary pencil respectively, these will both therefore decrease as we proceed from the red to the violet end of the spectrum, corresponding to types of the character of that represented in fig. 114; but Y_e being smaller, and decreasing more rapidly than Y_o . Moreover, since $Y_o = Y_e$ in the axis, and since as we recede from the axis Y_o increases (because the colour of the ordinary pencil becomes whiter and more luminous) while Y_e diminishes by the same degrees, (the extraordinary becoming deeper and less bright,) we shall represent both these changes satisfactorily by putting

$$Y_o = Y (1 + k \cdot \sin \theta^2); \quad Y_e = Y (1 - k \cdot \sin \theta^2).$$

These give $Y_o + Y_e = 2Y = \text{constant}$, or independent of θ , which agrees with an observation of Dr. Brewster, that in every situation the combined tints of the two images are exactly the same with the natural colour of the mineral, (which, in this instance, appears to have been alike in all directions.)

In this case, then, the colour of a plate of the crystal of given thickness exposed to natural light will be the same, whether the plate be cut parallel or perpendicular to the axis. But Dr. Brewster has observed, that this is not always the case, but that great differences occasionally exist in this respect. Thus he found, that in some specimens of *sapphire* the colour when viewed *along* the axis was deep blue, and when *across* it yellowish green. In *Idocrase* an orange-yellow tint is seen along the axis, and a yellowish green across it. Specimens of *tourmaline* also are not uncommon in which the tint across the axis is green, while along the axis it is deep red; and, in general, this mineral is always much more opaque in the direction of the axis than in any other; so much so, indeed, that plates of a very moderate thickness cut across the axis are nearly impermeable to light. One of the most remarkable instances of this kind we have met with is a variety of sub-oxysulphate of iron, which crystallizes in regular hexagonal prisms, and which viewed through two opposite sides of the prism is light green, but along the axis, a deep blood red, so intense that a thickness of $\frac{1}{20}$ inch allows scarcely any light to pass. It is obvious, that to such cases the formulæ of the last article do not extend. But a slight modification will enable us to embrace the phenomena in an analytical expression. For if we take

$$y_o = X_o + Y_o \cdot \sin \theta^2; \quad y_e = X_e + Y_e \cdot \sin \theta^2;$$

where X_o , Y_o , &c. as well as y_o , y_e represent functions of λ (the length of an undulation) being the ordinates of so many curves, or types of tints, whose relations are to be determined, we have

$$y_o + y_e = (X_o + X_e) + (Y_o + Y_e) \sin \theta^2.$$

Now this is the tint which a sphere of the medium of a diameter = 1 will exhibit when viewed by natural light along a diameter inclined θ^o to the axis. If we represent by A and B the ordinates of the types of the tints it is observed to exhibit in the directions of the axis, and perpendicular to it, we have, when $\theta = 0$,

$$y_o + y_e = A = X_o + X_e;$$

and when $\theta = 90^o$,

$$y_o + y_e = B = (X_o + X_e) + (Y_o + Y_e),$$

whence we have

$$Y_o + Y_e = B - A;$$

and the tint exhibited by ordinary light at the inclination θ to the axis, will be represented by

$$\begin{aligned} y_o + y_e &= A + (B - A) \cdot \sin \theta^2, \\ &= A \cdot \cos \theta^2 + B \cdot \sin \theta^2. \end{aligned}$$

Part IV.

1062.

Media possess two distinct absorbing powers.

1063.

Absorption of the rays in the two pencils examined in crystals with one axis.

Formulae for the light transmitted in each.

1064.

Cases of two distinct colours.

Investigation of formulae for these cases.

Expression of the colour transmitted in common light.

Light. Thus in the case of our sub-oxysulphate of iron, A is the ordinate of the type of a deep blood-red tint, and B in like manner represents a bright pale green, so that we shall have at any intermediate inclination θ

$$\text{tint} = (\text{deep red}) \times \cos \theta^2 + (\text{light green}) \times \sin \theta^2,$$

which represents faithfully enough the gradual passage of one hue into the other as the inclination changes.

1065. Suppose now the incident beam polarized in any plane, and let the plane in which the ray and the axis of the sphere lie make an angle $= a$ with that plane. Then would $\cos a^2$ and $\sin a^2$ represent the intensities of the ordinary and extraordinary pencils which superposed make up the emergent beam, were the crystal limpid; but in virtue of its absorbent powers, they will be reduced respectively to

$$y_o = \cos a^2 (X_o + Y_o \sin \theta^2), \quad \text{and} \quad y_e = \sin a^2 (X_e + Y_e \sin \theta^2),$$

so that at their emergence they will no longer make up white light, but a variable tint whose type has for its ordinate

$$(X_o \cdot \cos a^2 + X_e \cdot \sin a^2) + (Y_o \cdot \cos a^2 + Y_e \cdot \sin a^2) \cdot \sin \theta^2,$$

in which it will be recollected that $X_o + X_e = A$, and $Y_o + Y_e = B - A$.

To determine the individual values of X_o , &c. however, we must have two more conditions, and these will be found by considering, first, that in the direction of the axis the tint must be independent of a , which gives $X_o \cdot \cos a^2 + X_e \cdot \sin a^2$ independent of a , and therefore $X_o = X_e$, and either of them $= A$. To get another condition, let the tints be noticed which the sphere or crystal exhibits when its axis is perpendicular to the visual ray; and, first, coincident with, next, perpendicular to, the plane of polarization, i. e. when $a = 0$, and $a = 90^\circ$.

These are respectively

$$X_o + Y_o, \quad \text{and} \quad X_e + Y_e;$$

and calling these a and b , we have

$$Y_o = a - X_o = a - A, \quad Y_e = b - A.$$

Hence the final expression for the tint seen in polarized light will be

$$A + \{ (a - A) \cdot \cos a^2 + (b - A) \sin a^2 \} \cdot \sin \theta^2,$$

that is,

$$A \cdot \cos \theta^2 + \{ a \cdot \cos a^2 + b \cdot \sin a^2 \} \cdot \sin \theta^2,$$

in which it will be observed that a and b are complements of each other to the tint B, because

$$a + b = X_o + Y_o + X_e + Y_e = B, \text{ by Art. 1064.}$$

1066. Such is the expression for the apparent hue of crystals with one axis, which exhibit a variable colour in common or polarized light, according to their position with respect to the incident light. The phenomenon in question may be generally termed *dichroism*, though the word has usually been applied only to that particular case where a marked change in the character of the tint takes place, as from red to green, &c.

1067. The dichroism of biaxial crystals differs in many of its phenomena from those having only one optic axis. If we look through a plate, or into a crystal of any biaxial mineral, having the property in question, illuminated by natural light in such a direction that the visual ray within the crystal shall pass along, and in the immediate neighbourhood of, one of the axes, we shall perceive a phenomenon like that represented in fig. 206, consisting of two similar and equal sombre spaces A B one on either side of the pole P, and of the principal section P P', and if we look along the other axis P' a similar pair of spaces will be seen in its neighbourhood. In the mineral called *dichroite* by Haüy, (on account of the striking difference of its colours in different positions,) or *iolite* (from its violet hue) by others,* of which the phenomena have been described by Dr. Brewster in the Paper already cited, these spaces are of a full blue colour, while the intermediate region towards O, along the line O P C, and the space beyond P towards C are yellowish white. In epidote the sombre spaces are *brown*, and the region around O and in the principal section green, of a greater or less degree of dilution. In this latter mineral (at least in some of its more ordinary varieties of crystalline form, viz. in long striated prisms much flattened, and terminated by dihedral summits placed obliquely, so as to truncate two of the angles of the prism) the phenomena are seen without any artificial section, merely by looking in obliquely, across the axis of the prism; and the same is true of many other minerals, as, for instance, the axinite, in which the transition of colour is extremely remarkable and beautiful.

1068. The phenomena of dichroism in biaxial, as well as in uniaxial crystals, are evidently related to the optic axes, and depend on the planes of polarization assumed by the intromitted light, during its transit through the crystal to whose absorptive power it is subjected. Now, if we consider the form and situation of the sombre spaces where the greatest absorptive energy is exerted, we are at once struck by their analogy with those occupied by the more vividly coloured parts of the rays about the axes in the situation of fig. 179. That figure represents (Art. 900) the extraordinary set of rings as seen in a crystal whose principal section is in the plane of primitive polarization. Fig. 207 represents the ordinary or complementary set as seen around either of the axes, the pole P, and the principal section being here occupied with white light, and very bright, in consequence of its containing the whole incident light, while the lateral or coloured portions occupied by the rings are less illuminated, the colours originating in an abstraction of certain rays.

Conceive now a number of such sets of coloured rings not all of exactly the same dimensions, nor having

* Mohs, with his usual contemptuous disregard of, or rather hostility to, all ordinary convenience and received usage, chooses to call this mineral "*prismatic quartz*." Such a nomenclature *must* ere long work out its own destruction, but while it subsists the nuisance is intolerable. We cannot but lament, that such a cause should exist to raise up prejudice against a system in many respects so useful and valuable.

Light.

precisely the same pole, but very nearly so, to be superposed on one another, then would the colours be obliterated and blended into white light by their overlapping, but still the general intensity of the light in the lateral regions would remain much feebler than in the principal section, and the effect would be precisely that of fig. 206, viz. two sombre, cloudy, fan-shaped spaces traversed by a narrow ray of vivid light, opening out from P towards C and O. Such would be the case with a limpid crystal, supposing such a slight degree of confusion of structure as to produce the non-coincidence of the rays from all its molecules. In this case, however, neither of the spaces in question would appear coloured, nor would the phenomena be seen at all without the use of polarized light and its subsequent analysis. But if we conceive the crystal, instead of limpidity, to possess the property of double absorption, the suppressed and transmitted portions will be, not white light, but light of the colour of one or other of the pencils into which it is resolved by double refraction, according to its plane of polarization and the thickness of the medium it has traversed; and the analysis of the emergent ray may be regarded as performed, at least imperfectly by the difference of absorptive powers acting differently on the two pencils. In support of this it may be noticed, that when we examine the system of rings in the usual way, by polarized light, in crystals presenting the above phenomenon, they are usually found to be very irregular, several sets evidently overlapping and interfering with one another, and rendering the non-coincidence of all the axes a matter of ocular demonstration.

In Art. 931 we investigated the law of intensity of the illumination of the polarized rings in different parts of their periphery for uniaxial crystals. As what is there said does not apply, however, to biaxial ones, and as the present subject has led us to the consideration of the more general case, it will not be irrelevant, if we digress at this point, in order to show, what modifications the statement there made must receive to embrace the phenomena of biaxial crystals.

M. Biot has stated the general law of polarization in biaxial crystals, from his elaborate researches on that subject (*Mém. sur les Lois Générales de la Double Réfraction et Polarisation*, &c. *Mém. Acad. Sci.* 1819) to be as follows:

If two planes be drawn through the course of a ray within a crystal and through the two optic axes, and a third plane bisecting the angle included between the two former, this will be the plane of polarization if the ray be an ordinary one—but one perpendicular to it if extraordinary. Thus in fig. 209, CP and CP' being the optic axes, and AC a ray penetrating the crystal, if PA, P'A be joined by arcs of circles on the sphere HOKA having C for its centre, and the angle PAP' be bisected by the arc AN, the plane ACN bisecting the dihedral angle between the planes PCA and P'CA is the plane of ordinary polarization, and a plane perpendicular to it that of extraordinary. This is the law of fixed polarization, and expresses generally the planes of polarization assumed by the two rays at their emergence from doubly refracting crystals. It is a consequence of Fresnel's general theory, (though deducible from it by a train of analytical reasoning far too intricate and refined to allow of its insertion in a treatise like the present,) and, having been experimentally established long before that theory was devised, must be looked on as a strong additional proof of its conformity to nature.

The doctrine of movable polarization, however, which, so far as respects the phenomena of the colours and intensity of the rings, has been shown by M. Biot in the same excellent paper, to represent with fidelity their various affections, whether in uniaxial or biaxial crystals, requires the resulting ray to assume at its emergence a plane of polarization alternately coincident with, and making with the primitive plane of polarization twice the angle which the plane of fixed polarization so determined would make; so that if we draw AM (fig. 208) bisecting the angle PAP', the emergent ray will be affected by subsequent analysis, as if polarized either in the plane of primitive polarization, or making with it an angle equal to twice CMA, and from this it is easy to derive the law of intensity in question, for the ray by which the point A of the rings is formed consists of two portions, of which (A) is affected by subsequent analysis by a prism of Iceland spar, as if it were polarized in a plane making an angle 2CMA = ψ with the plane of primitive polarization, in which we suppose the principal section of the analyzing prism to be placed, and the other, complementary to this (1 - A) retains its primitive polarization. The portion A then will be divided between the ordinary and extraordinary image in the proportion (cos 2ψ)² : (sin 2ψ)², and (considering only the latter,) A being its intensity at its emergence from the crystal, A . (sin 2ψ)² will be its intensity in the extraordinary image, or in the primary set of rings, while the whole of the portion 1 - A will pass into the ordinary or complementary set, as in Art. 932, so that we have only to express this in terms of the azimuth of the crystallized plate itself, and the direction of the ray within the crystal. For this purpose, put α = angle COP = azimuth of the principal section of the plate reckoned from the plane of primitive polarization, θ = AP, θ' = AP', and let us (for simplicity) consider only at present the case when P and P' are near, as in nitre, so that arcs of circles may be regarded as straight lines, and spherical as plane triangles, (see Art. 907.) Now if in fig. 208 we put φ for the angle PNA, or the angle made by the plane of ordinary polarization with the principal section, we shall have ψ = CMA = COP + MNO = COP + PNA

= α + φ. To find φ we have only to consider that sin φ² = $\left(\frac{PA}{PN}\right)^2 \times \sin(PAN = \frac{1}{2} PAP')^2$; but since NA

bisects the angle of the triangle PAP' and cuts the base, PN = PP' × $\frac{PA}{PA + AP'} = \frac{2a\theta}{\theta + \theta'}$, and

$$(\sin \frac{1}{2} PAP')^2 = \frac{1}{2} (1 - \cos PAP') = \frac{4a^2 - (\theta - \theta')^2}{4\theta\theta'}$$

so that

$$\sin \phi^2 = \frac{(\theta + \theta')^2 \{ 4a^2 - (\theta - \theta')^2 \}}{16a^2 \cdot \theta\theta'}$$

Part IV.

Analogy in respect of intensity of illumination

1069.

DIGRESSION Theory of the polarized rings resumed.

1070.

Biot's general law for the planes of polarization in biaxial crystals

1071.

Doctrine of movable polarization applied to biaxial crystals, Fig. 208.

Law of intensity of the rings in different points of their periphery.

Analytically expressed,

Light. A more symmetrical value of ϕ will, however, be had by expressing the value of $\sin 2\phi$, which being equal to $4 \cdot \sin \phi^2 (1 - \sin \phi^2)$ is immediately given by substitution of the foregoing. If we execute the reductions we Part IV.

shall find that, putting S for $\frac{\theta + \theta' + 2a}{2}$ = half the sum of the sides of the triangle PAP'

$$\sin 2\phi = \frac{2(\theta + \theta')(\theta - \theta')}{(2a)^2} \cdot \frac{\sqrt{S(S-\theta)(S-\theta')(S-2a)}}{\theta\theta'}$$

Now $\frac{2\sqrt{S(S-\theta)(S-\theta')(S-2a)}}{\theta\theta'}$ is the well known expression for the sine of the angle PAP' included

between the sides θ, θ' , and therefore calling this angle P , we have

$$\sin 2\phi = \frac{(\theta + \theta')(\theta - \theta')}{(2a)^2} \cdot \sin P.$$

The nature of this expression renders the transition from plane to spherical triangles easy, and we may conclude consequently, that, in crystals where the axes make any angle $2a$, that if we take

$$\sin 2\phi = \frac{\sin(\theta + \theta') \cdot \sin(\theta - \theta')}{(\sin 2a)^2} \cdot \sin P,$$

and $\psi = a + \phi$, we shall still have the intensity of the extraordinary rings represented by $A(\sin 2\psi)^2$, and that of the ordinary by $1 - A + A(\cos 2\psi)^2$, that is, $1 - A(\sin 2\psi)^2$, their sum being, as it ought, unity.

1072.
Form of the
black cross
in biaxial
crystals in-
vestigated

The black cross which divides the system of the primary rings, is too remarkable a feature not to require express notice. Its form, it is evident, must be determined by the condition that the line MA shall be everywhere perpendicular to COD , in which circumstances the locus of A will be a curve marking out its central or blackest portion. The problem then is reduced to a purely geometrical one, Required a curve PA such that a line drawn from A bisecting the angle between lines AP, AP' drawn to two given points P, P' , shall always be perpendicular to a given line COD . To resolve this, retaining the former notation, and putting $OM = x, MA = y, OA = r$, we have

$$\cos AOP = \cos(AOM - a) = \frac{x \cdot \cos a + y \cdot \sin a}{r} = \frac{N}{r},$$

$$\sin AOP = \frac{y \cdot \cos a - x \cdot \sin a}{r} = \frac{M}{r},$$

putting N and M for the respective functions in the numerators of the fractions.

Now since PAM is half the angle PAP' , it is easy to see that we must have $2 \times$ angle $O'AM = PAO - P'AO$.

But,

$$\cos PAO = \frac{\theta^2 + r^2 - a^2}{2r\theta}; \cos P'AO = \frac{\theta'^2 + r^2 - a^2}{2r\theta'};$$

and

$$\sin PAO = \sin AOP \times \frac{PO}{PA} = \frac{aM}{r\theta}; \sin P'AO = \frac{aM}{r\theta'};$$

consequently we have, first,

$$\sin 2OAM, \text{ or } \frac{2xy}{r^2} = \frac{aM}{2r^2} \left\{ \frac{\theta'^2 + r^2 - a^2}{\theta\theta'} - \frac{\theta^2 + r^2 - a^2}{\theta\theta'} \right\} = \frac{aM}{2r^2\theta\theta'} (\theta'^2 - \theta^2), \quad (a)$$

and, secondly,

$$\cos 2OAM = \frac{y^2 - x^2}{r^2} = \frac{1}{r^2} \left\{ \frac{(\theta'^2 + r^2 - a^2)(\theta'^2 + r^2 - a^2)}{4} + a^2 M^2 \right\}.$$

Now we have further,

$$\begin{aligned} \theta^2 = a^2 + r^2 - 2ar \cdot \cos AOP &= a^2 + r^2 - 2aN \}; & \theta'^2 - \theta^2 &= 4aN, \\ \theta'^2 &= a^2 + r^2 + 2aN \} \end{aligned}$$

which substituted in the values of $\sin 2OAM$ and $\cos 2OAM$ above, give the equations

$$\begin{aligned} xy \cdot \theta\theta' &= a^2 \cdot MN, \\ (y^2 - x^2) \cdot \theta\theta' &= r^4 + a^2(M^2 - N^2), \end{aligned}$$

and, eliminating $\theta\theta'$ from these, we obtain

$$a^2(y^2 - x^2) \cdot MN = xy \{ r^4 + a^2(M^2 - N^2) \}.$$

In this it only remains to substitute for M and N their values $y \cdot \cos a - x \cdot \sin a$, and $y \cdot \sin a + x \cdot \cos a$, which done, the whole will be found divisible by r^4 , and will reduce itself to the very simple equation

Light.

$$xy = a^2 \cdot \sin a \cdot \cos a = \frac{a^2}{2} \cdot \sin 2a.$$

Part IV.

When the axes are near it is a hyperbola

The black cross then is an hyperbola, passing through the poles P, P', and having the planes of primitive polarization, and one perpendicular to it (CD and *cd*) for its asymptotes, and which as *a* approaches to 0, or 90°, approaches nearer and nearer to its asymptotes, with which it at last coincides in the limiting case, all which particulars are exactly conformable to fact, and may easily be verified by turning a plate of nitre round between crossed tourmalines. When the inclination of the axes is so considerable, that the rings about both poles cannot be seen at once, there will arise modifications from the substitutions of the sines, &c. of arcs for the arcs themselves, which it is not worth while to enter into.

To return now to the phenomena of dichroism. That portion of the light transmitted by a biaxial coloured medium which has relation to the optic axes, and which forms the sombre brushes of colour (in fig. 206,) and the bright spaces which divide them, have evidently for their analytical expression a function of the form

$$Y \cdot (\cos 2\phi)^2 + B \cdot (\sin 2\phi)^2; \quad (a)$$

where Y and B are functions of λ , and represent the ordinates of the types of two fundamental tints, ϕ representing as before the angle PNA, fig. 208, or the angle made by the plane of ordinary polarization with the principal section. But besides this, the phenomena described by Dr. Brewster, as exhibited by the iolite, require us to admit two other portions, which may be more naturally referred, not to either of the optic axes but to the line CO (fig. 209) bisecting them, and having for its expression a function of the form $a \cdot \cos OA^2 + b \cdot \sin OA^2$. In this mineral, when exposed to common light (or to polarized, provided we place its principal section at right angles to that of polarization,) the lateral brushes A, B, fig. 206, are blue, and the bright rays which divide them, passing through the poles P, P' are white, or yellowish white, and so far the phenomena agree with the expression (a) if we suppose Y to represent a bright yellowish white, and B a blue. But according to that expression alone, the blue spaces should be continued down to the equator C *ab* D, fig. 206, and there ought to be two directions CD and *ab* in which the mineral viewed transversely to the axis of the prism (which is perpendicular to the plane C *ab* D) should appear yellow, and two others, *mn* and *pq*, in which it should transmit a blue colour, while in the direction of the axis O it should appear yellow. Now, on the contrary, the equatorial colour is nearly uniform and pale yellow, while that along the axis O is blue; and in proceeding from the equator toward the axis O of the prism, the yellow diminishes, and the blue gains strength, whether we set out from C and D, or from *a* and *b*, precisely as would be indicated by the other formula

$$y \cdot (\sin OA)^2 + b \cdot (\cos OA)^2,$$

y representing a yellow white and *b* a blue tint. If, therefore, we put $OA = \nu$, the joint expression

$$T = (Y \cdot \cos 2\phi^2 + B \cdot \sin 2\phi^2) + (y \cdot \sin \nu^2 + b \cdot \cos \nu^2); \quad (b)$$

will be found to represent pretty correctly the variations of colour as far as they can be judged of by the eye. Thus, at O where $\nu = 0$, and $\phi = 90^\circ$, we have $T = Y + b$, which may indicate either a yellow, a white, or a blue, according as we suppose Y or *b* to be predominant. The fact being, that the tint at O is blue, we must suppose the latter to express the more decided colour. As we proceed from O along the sections OC, OD, or O*a*, O*b*, in both of which $\sin 2\phi = 0$, we have

$$T = (Y + y \cdot \sin \nu^2) + b \cdot \cos \nu^2 = (Y + b) + (y - b) \cdot \sin \nu^2$$

Now *y* expressing a yellow white and *b* a strong blue, *y* - *b* will express a proportionally vivid yellow, and therefore the blue tint $Y + b$ seen along the axis will be diluted with more and more yellow as we approach the equator; at PP', then, (by a proper assumption of numerical values) it will be rendered nearly neutral, after which the yellow will predominate, and, at the equator, will remain alone sensible, the expression for T then becoming $T = Y + y$, at the points C, *a*, *b*, D. Let us next consider the case when $\cos 2\phi = 0$, or $\phi = 45^\circ$, that is to say, along the axes or most intense lines of the lateral brushes. In this case we have

$$T = B + (y \cdot \sin \nu^2 + b \cdot \cos \nu^2) = (B + b \cdot \cos \nu^2) + y \cdot \sin \nu^2.$$

Now if we suppose B and *b* to represent blue tints, since (in the case of iolite) the angle between the axes or PP' = 62° 50' and OP = 31° 25', we have in the immediate vicinity of the poles, $(\sin \nu)^2 = \frac{1}{4}$ nearly, and $\cos \nu^2 = \frac{3}{4}$, so that in the immediate neighbourhood of P the tint of the most intense part of the brushes will be $B + \frac{3}{4}b + \frac{1}{4}y$, which, on very reasonable suppositions of the numerical values of B, *b* and *y* will denote a full and rich blue. But as we approach the equator at *m*, *n*, *p*, *q*, $\cos \nu^2$ diminishing and $\sin \nu^2$ increasing, the sombre tint B is continually more feebly reinforced by the tint $b \cdot \cos \nu^2$ and more strongly counteracted by $y \cdot \sin \nu^2$, till at length it will be overpowered, and the colour in these points, as in C, *a*, D, *b*, will be yellow only somewhat less decided than in the latter, its tint being represented by $T = y + B$ instead of $y + D$.

In general, if we put A for the tint transmitted along the axis O of the prism, P for that seen along the poles, L for that of the lateral branches at their origin close to the poles, and E for the mean equatorial tint, we shall have for determining Y, *y*, B, *b*, the equations

$$A = Y + b, \quad 2E = 2y + B + Y,$$

$$P = Y + y \cdot \sin a^2 + b \cdot \cos a^2; \quad L = B + y \cdot \sin a^2 + b \cdot \cos a^2,$$

on elimination from these, it will appear that there is an equation of condition to be satisfied, viz.

$$2(A - P) = (2A - 2E - P + L) \cdot \sin a^2; \quad (c)$$

1073.

Empirical formula expressing the phenomena of dichroism

1074.

Determination of the coefficients from the observed tints.

Light. and that supposing it satisfied, one of the tints, as y , will (so far as these conditions are concerned) remain arbitrary, and the others will be given by the equation Part IV.

$$\left. \begin{aligned} 2Y &= 2E + P - L - 2y \\ 2B &= 2E - P + L - 2y \\ 2b &= 2A - 2E - P + L + 2y \end{aligned} \right\}; (d)$$

in which y must, however, be such as to render Y, B, b real tints, *i. e.* expressed by positive numbers.

1075. Application to the iolite. To apply this, for example's sake, to the case of the iolite, let us regard every white ray as consisting of two complementary rays of bright yellow and bright blue of equal efficacy; and suppose that by observation we have ascertained its equatorial tint E to be a very pale but strongly luminous yellow white, consisting of 110 such yellow rays, and 99 such blue ones, producing a joint intensity = 209. Moreover, let the tint seen along the axis of the prism (A) be a blue, of a good colour, but considerably less intensity, represented by 10 such yellow + 20 such blue rays = 30. That seen along the optic axes (P) to be a white represented by 36 yellow + 36 blue = 72, and that of the most intensely coloured portions of the lateral brushes = L to be a stronger blue than that seen in the axis of the prism, such as may be represented by 28 yellow + 66 blue = 94. These numbers are chosen so as to satisfy the equation of condition, taking $a = 30^\circ$, and if we substitute them we shall find

$$y + y = 114 \text{ yellow} + 84 \text{ blue}; \quad B + y = 106 \text{ yellow} + 114 \text{ blue}; \quad y - b = 104 \text{ yellow} + 64 \text{ blue},$$

y remaining indeterminate; if we suppose its composition to be m yellow + n blue, we may determine m and n by the two conditions that b shall (as we have before supposed) represent, a pure blue without any mixture of yellow, and Y a very pale yellow, such as would result from a mixture of yellow and blue in the ratio of 10 to 9. These conditions are satisfied by taking $m = 104$ and $n = 75$; so that we have, finally,

$$\begin{aligned} Y &= 10 \text{ yellow} + 9 \text{ blue}; & B &= 2 \text{ yellow} + 39 \text{ blue}; \\ y &= 104 \text{ yellow} + 75 \text{ blue}; & b &= 0 \text{ yellow} + 11 \text{ blue}; \end{aligned}$$

and these being taken for the values of the coefficients in the expression (b) Art. 1073, it will be found on trial to reproduce the tints actually observed. In fact, the extreme equatorial tints being $y + Y$ and $y + B$, will be respectively represented by 114 yellow + 84 blue, and 106 yellow + 114 blue; the former is a very pale yellow, but highly luminous, being equivalent to 30 rays of yellow diluted with 168 of white; while the latter is a blue so pale as to be undistinguishable from white, and also highly luminous, being equivalent to 8 rays of blue diluted with 212 of white.

1076. Phenomena exhibited by various crystals. The reader will perceive that the formula in question is merely empirical, and that more numerous experiments than we possess will be required to establish or disprove it. It is unfortunately, however, difficult to meet with biaxial crystals sufficiently dichromatic for the purposes of decisive experiment, and at the same time large and transparent enough to admit of being cut into the forms and examined in the directions required, through a thickness sufficient for a full development of their colours. Such are indeed hardly less rare than the most precious gems; and this circumstance is a great obstacle to the advancement of our knowledge in one of the most interesting branches of optical inquiry, which that of dichroism certainly deserves to be considered. Among artificial crystals, however, there is room to suppose that subjects fit for such experiments may be met with. One remarkable instance of dichroism among these has been mentioned in the sub-oxy-sulphate of iron. To this we may add the potash-muriate of palladium, which exhibits along the axis of the four-sided prism in which it crystallizes a deep red, and in a transverse direction a vivid green. (Wollaston, *Phil. Trans.* 1804. On a new metal in Crude Platina.) The curious property of the purpurates of ammonia, potash, &c. described by Dr. Prout, (*Phil. Trans.* 1808,) which by transmitted light exhibit an intense red, and by reflected, on one surface, a dull reddish brown, and on another a splendid green, appears referable, not so much to the principles of dichroism properly so called, as to some peculiar conformation of the green surfaces, producing what may be best termed a *superficial colour*, or one analogous to the colour of thin plates, and striated or dotted surfaces. A remarkable example of such superficial colour, differing from the transmitted tints, is met with in the green fluor of Alston-moor, which on its surfaces, whether natural or artificial, exhibits, in certain lights, a deep blue tint, not to be removed by any polishing.

1077. Unequal effects of heat on the colours of the two pencils. Dr. Brewster has shown that the action of heat often modifies in a very remarkable manner the colour of doubly refracting crystals, producing a permanent change in the scale of absorption of the crystals as affecting one of the pencils and not the other. Thus, having selected several crystals of Brazilian topaz which displayed no change of colour by exposure to polarized light, (and in which, of course, the types of both absorptions must have been alike,) and bringing them to a red heat, or even boiling them in olive oil, or mercury, they experienced a permanent change, and had acquired the property of absorbing polarized light unequally. He then took a topaz in which one of the pencils was yellow and the other pink; and by exposing it to a red heat, he found the extraordinary pencils more powerfully acted on than the ordinary, the yellow colour being discharged entirely from the one, while only a slight change was produced in the pink tint of the other. This change of colour in the topaz by heat (though not its intimate nature) is well known to jewellers, who are in the habit of thus developing in this gem a colour more highly prized. It is remarkable, that while hot the topaz is perfectly colourless, and acquires the pink colour gradually in cooling. By the repeated action of very intense heat Dr. Brewster was never able to modify or remove this permanent pink tint. How far violent compression, slow application, and abstraction of the heat, or other modifying circumstances, might prevent its development, it

Light. would be interesting to examine; since we cannot help being otherwise struck by the force of the argument Part IV.
geologists may draw, from the existence in rocks of a mineral which mere elevation of temperature unaccompanied
with change of composition, thus irrevocably alters.

One general character of all dichroite bodies is, that when natural light is transmitted through a plate of 1078.
sufficient thickness, in any direction not coincident with one of the optic axes, the emergent beam is wholly or General
partially polarized by reason of the unequal action of the medium on the two pencils, and the consequent sup- character of
pression of one of them. And, in general, whatever cause tends to interfere unequally with their free trans- dichroite
mission through a medium, will produce a similar effect. Thus, for example, if the continuity of a doubly crystals.
refracting medium be interrupted by a film of any uncrystallized substance, since the two pencils by reason of Effects of
their angular separation are incident on this film at different angles; and since, moreover, their relative refractive uncrystal-
indices, with respect to the medium composing the film, differ, they will undergo partial reflexion at the film in lized inter-
different proportions, and thus an inequality will arise in the parts transmitted. If the refractive index of the rupting
film be precisely equal to the ordinary refractive index of the crystal (supposed, for simplicity, to be uniaxial) films.
the ordinary ray, it is evident, will undergo no disturbance or diminution, while the extraordinary ray will be
changed in direction and diminished in intensity by partial reflexion at its ingress and egress, at every such film
which may exist in the medium. If the films be extremely numerous, and if, moreover, they be not disposed
in planes, but in undulatory or irregular surfaces through the medium, this will make no difference, so far as the
ordinary ray is concerned, which will still pass undisturbed through the system, (except so far as any opacity in
the matter of the films may extinguish a portion of it;) but the extraordinary ray will be rendered confused,
and dispersed, its egress from the films not being performed (by reason of their curvature) at the same angles
as its ingress, and that irregularly, according to their varying inclination. Hence will arise a phenomenon pre- Phenomena
cisely such as is presented by the agate, and other irregularly laminated bodies, through plates of which, if a of agate.
luminary be viewed, it is seen distinctly, but as if projected on a curtain of nebulous light; and if ex-
amined with a tourmaline, or doubly refracting prism, the distinct image, and the nebulous light, are found to
be oppositely polarized. If we examine a piece of agate with a magnifier, the laminated structure and unequal
refraction of the laminae are very apparent; it appears wholly composed of a set of exceedingly close layers,
not arranged in planes, but in undulating or crinkled lines like a number of figures of 333333 placed close
together. The planes of polarization of the nebulous and distinct image are parallel and perpendicular to
the general direction of the layers, which through any very small portion of the substance is generally pretty
uniform.

But the film interposed may, itself, be crystallized, and inserted between adjacent portions of a regular crystal, 1079.
according to the crystallographic laws which regulate the juxtaposition of the molecules at the common surfaces Action of a
of macle or hemitrope crystals. Let A D E F (fig. 210) be such a plate interrupted by a crystallized lamina crystallized
B C E F, bounded by parallel planes, and let us consider what will happen to a ray S a incident at a. It is interrupting
evident, that were the crystallized lamina away, or were its molecules homologously situated with those of the film.
portions on either side of it; in the latter case, we should have an uninterrupted crystal; in the former, two Fig. 210.
prisms disposed with their principal sections parallel, and acting in opposition to each other; in either case, the
emergent ordinary and extraordinary pencils separated by double refraction at the first surface will emerge
parallel to the incident ray, and therefore to each other. But the principal section of the crystallized film being
non-coincident with those of the two prisms A B E, C F G, it will alter the polarization of the portions a b, a c;
and in place of their being, as in the former case, each refracted singly by the second prism C F G, they will now
each be refracted doubly, so that in place of two emergent rays there will now be four. The subdivision of the
rays within the interposed lamina may evidently be disregarded, for they will be refracted in passing from the
film into the second prism in the same direction, where contiguous, as they would were an infinitely thin plate of
air interposed. Now, in that case, they would emerge from the film in pairs respectively parallel to the incident
rays a b, a c, and therefore to each other. Hence the refraction at the second prism will be precisely the
same as if the lamina were suppressed, and in its place the rays a b, a c had received at a the polarizations
they acquire by its action. Now, these being in opposite planes, it is evident that each of the rays a b, a c
would undergo both an ordinary and an extraordinary refraction. Let us denote these four emergent pencils
so arising by O O, O E, E O, E E, and suppose a b to be the direction taken by the ordinary refracted portion
of S a, and a c that of the extraordinary. Then, since O O has been refracted ordinarily by the prism C F G,
and was incident on it in the direction of the ordinary ray a b, its direction on emerging will be parallel to S a.
Similarly, E E is refracted extraordinarily, and being incident in the direction b c of the extraordinary portion
of S a, it also will emerge parallel to S a, and thus the two rays O O, E E will emerge parallel, and their
systems of waves will be superposed. But the portions O E and E O, the one being incident in the ordinary
direction, but refracted extraordinarily, the other incident in the extraordinary direction and refracted ordinarily,
will neither emerge parallel to the original ray S a, nor to each other; and this will give rise to two lateral
images, one on each side of the central or direct image, which will have, moreover, an intensity equal (except
in extreme cases) to the sum of those of the lateral images.

If the film E B C F be very thin, or if either of its optic axes be nearly coincident with the direction in which 1080.
the light traverses it, the difference of paths and velocities within it will give rise to an interference of the pairs Phenomena
of rays going to form either pencil emergent from the film, and thus will arise the colours of the rings in each of inter-
image. Those on either side the central one will be consequently tinged with the respective colours of the rupted
primary and complementary set of rings; while the central image, being formed by the precise superposition of Iceland
two similar complementary pencils will appear white. spar.

All these phenomena actually occur, and have been described by Dr. Brewster, and explained by him on the
principles here laid down, in certain not uncommon specimens of Iceland spar, which are interrupted by such

Light. hemitrope films, passing through the longer diagonals of opposite faces of the primitive rhomb. If we look at a candle through such an interrupted rhomb, it will be seen accompanied by a pair of lateral images such as here described, and exhibiting frequently the complementary tints with great splendour.

1081. If the luminary from which the ray $S\alpha$ issues be small, the lateral images will be separated by a dark interval from each other and from the central one, but if large they will overlap. If infinite (as where the uniform light of the sky is viewed) all the images will be superposed. But the field of view will not necessarily be uniform and white. The central image will form an intense white screen, or ground, on which will be projected the lateral ones. Now, if the film be so constituted as to have within the visible field of view of one only of the lateral images the pole of one of its sets of rings, (which will be the case whenever one of its optic axes is not very remote from perpendicularity to the surface of the plate AD , so as to admit of one of the rays OE or EO traversing the film in the direction of its axis,) that set of rings will not be seen projected centrally on the corresponding set complementary to it of the other lateral image, by reason of the angular separation of these two images. Of course its colours will not be neutralized, and it will be visible *per se*, though very faint, being diluted by the whole white light of the central image (OO , EE) and by the whole visible and nearly uniform portion of the other lateral one (OE .)

1082. This is not the only way in which a crystal perfectly colourless may exhibit its sets of rings by exposure to common daylight without previous polarization, or without subsequent analysis of the transmitted pencil. The general mass of the crystallized plate may have one of its optic axes in the direction of the visual ray, as in fig. 211, and the portion of it $CDdc$ included between two films $BCcb$ and DdE will then form precisely such a combination as that above described, and will exhibit a set of rings feeble in proportion to the rarity and minuteness of the films, and the consequently small area of their outcropping surfaces BC , DE . These are not hypothetical cases. Dr. Brewster states himself to have met with specimens of nitre exhibiting their rings *per se*. Such are rare. But in the bicarbonate of potash it is an accident of continual occurrence; and, indeed, almost universal. The films in both cases are easily recognised, and their position and that of the system of rings seen leave no doubt of the correctness of the explanation here given. Such crystals, of which more will no doubt be hereafter recognised, may be termed *idiocyclophanous* till a better term can be thought of.

§ XII. On the effects of Heat and Mechanical Violence in modifying the action of Media on Light, and on the application of the Undulatory Theory to their explanation.

1083. It was ascertained independently, and about the same time by Dr. Seebeck and Dr. Brewster, that when glass, which in its ordinary state offers none of the phenomena of double refracting media, is heated or cooled unequally, it loses this character of indifference, and presents phenomena of coloration, &c. analogous, in many respects, to those exhibited by doubly refracting crystals. If the heat communicated be below the temperature at which glass softens, the effect is transient, and vanishes when the glass attains a uniform temperature throughout its substance, whether by the equable distribution of the caloric throughout its mass, or by its abstraction in cooling. But if the temperature communicated be so high as to allow the molecules of the glass to yield to the mechanical forces of dilatation and contraction produced in the act of cooling and take a new arrangement, the effect is permanent, and glass plates so prepared have many points of resemblance with crystallized bodies. Dr. Brewster afterwards ascertained, that mechanical compression or dilatation applied to glass, jellies, gums, and singly refractive crystals (such as fluor spar, &c.) is capable of imparting to them the same characters. If the medium to which the pressure is applied be perfectly elastic, like glass, the effect, like that of heat, is transient. But if during the continuance of the compression or dilatation, the particles of the medium are allowed to take their own arrangement and state of equilibrium, then when the external force is withdrawn a permanent polarizing character will be found to exist.

1084. As periodical colours are not produced in phenomena of this class without a resolution of the incident light into two pencils moving with different velocities, and as a difference of velocities is invariably accompanied with a difference of refraction at inclined surfaces, it might be expected that media thus under the influence of heat or pressure should become doubly refractive. This has been verified by direct experiment by M. Fresnel, who has shown that a peculiar species of double refraction is thus produced.

1085. As the unusual heating or cooling of glass and other substances, is well known to produce in the parts heated or cooled a corresponding inequality of bulk, and thus to bring the parts adjacent into a state of strain in all respects analogous to that arising from mechanical violence, and as, in fact, the effects of heat in communicating double refraction to glass, whether transient or permanent, are all, as we shall see, (with one very obscure and doubtful exception) commensurate with the amount of the strain thus transiently or permanently induced, we have little hesitation in regarding the inequality of temperature as merely the remote, and the mechanical tension or condensation of the medium as the proximate cause of the phenomena in question, and are very little disposed to call in the agency of a peculiar crystallizing fluid, endowed with properties analogous to those of magnetism, electricity, &c., to account for the phenomena, still less to regard media under the influence of heat or pressure as in any way thereby rendered more crystalline than in their natural state of equilibrium.

1086. In gasiform, or fluid media, no such phenomena are observed to be developed by either heat or pressure; the reason is obvious, the pressure is equally distributed in all directions, and the elasticity of the ether (on the undulatory hypothesis) preserves its uniformity.

But in solids the case is different. The molecules cannot shift their places one among the other, and the

Light.

effect of a compression in any direction is, *first*, to urge contiguous particles nearer together in that direction, and thereby to call into action their repulsive forces, more than in the natural state, to maintain the equilibrium; secondly, but much more slightly to urge contiguous particles in a direction perpendicular to that of the pressure laterally *asunder*, by reason of the increase of the oblique repulsive force developed by the approach of the molecules in the line of pressure to those which lie obliquely to that line. But this action, which in fluids would cause a motion of the lateral particles out of the way, in solids is ultimately equilibrated by an increase of the attractive forces of the adjacent molecules in a line perpendicular to the line of pressure; and thus we see that every external force applied to a solid is accompanied with a condensation of its particles in the direction of the force and a dilatation in a perpendicular direction. It is probable, however, that this latter is extremely minute, on account of the rapid diminution of the molecular forces by increase of distance, rendering the diagonal action insensible. But the former may easily be conceived to produce in the ether, in virtue of its connection (whatever it be) with the molecules of refracting media, a difference of elasticity in the two directions in question, accompanied with all the necessary concomitants of interfering pencils, periodical colours, and double refraction. The effect of dilatation will be the converse of that of compression, the direction of maximum elasticity in the one case being that of minimum in the other.

Part IV.

Mode of action of pressure on the molecules of solids.

These views are in perfect accordance with the experiments described by Brewster and Fresnel on compressed and dilated glass. According to the former (*Phil. Trans.* 1816. vol. 106) the effect of pressure on the opposite edges of a parallelepiped of glass is to develop in it "neutral" and "depolarizing axes," the former parallel and perpendicular to the direction of the pressure, the latter 45° inclined to them; in other words, a parallelepiped of glass so compressed, will when exposed to a ray polarized in the plane parallel or perpendicular to the sides to which the pressure is applied, produce no change in its polarization and develop no periodical colours, while if polarized in 45° of azimuth with respect to those sides, it will develop a tint, descending in the scale of the coloured rings as the pressure increases.

1087

Effects of compression described.

In this case, if the pressure be uniformly applied over the whole length of each opposite side, the elasticity of the ether in every point of the plate will be uniform in either direction at every point of the plate, being a maximum in one, and a minimum in that at right angles to it. The incident light therefore if polarized in azimuth a° will resolve itself into two pencils of unequal intensity (*viz.* $\cos a^\circ$ and $\sin a^\circ$) polarized in these two planes, and differing at their egress by an interval of retardation proportional to $t \times (v' - v)$, where t is the thickness traversed, and $v' - v$ the difference of velocities of the pencils, which when received on a double refracting prism will (as in the case of a crystallized plate (Art. 969) give rise to complementary periodical tints in the two images, the extraordinary image vanishing when $a = 0$, or 90, and the contrast being a maximum at 45°. It is, of course, extremely difficult to give such a perfect equality of pressure, so that we must not be surprised if a perfect uniformity of tint over the whole surface of the glass should not take place. In the experiment, however, described by Dr. Brewster (Prop. I. of the Memoir cited) this seems to have been the case.

1088.

Explanation on the undulatory doctrine.

If we suppose the elasticity of the ether in compressed glass *less* in the direction of the force applied (and where consequently the medium is densest, according to the general law) than in the perpendicular, the contrary will be the case in dilated. Hence, supposing the forces equal, in two similar plates, the extraordinary waves, or those whose vibrations are performed in the direction of the pressure, and which are therefore polarized at right angles to that direction, will advance most rapidly in the former case, the ordinary in the latter. Consequently, if we regard the interval of retardation or the tint, $t(v' - v)$ as negative in the former case, it will be positive in the latter; and the tints in the two cases will present the opposite characters of those exhibited by doubly refracting crystals of the two classes described in Art. 940, *et seq.* see also Art. 803, as negative and positive, or repulsive and attractive. Two such plates, therefore, placed homologously, or with the directions of the forces coincident, ought to neutralize each other, and if crossed at right angles should reinforce each other; and in general, if t be the thickness and f the compressing force applied to any plate (supposing the difference of velocities to be proportional to the force, and regarding dilating forces as negative) we shall have for homologously situated plates

1089.

Opposite effects of compression and dilatation.

$T = \text{tint polarized by any number of plates}$

$$= (f \cdot t + f' \cdot t' + f'' \cdot t'' + \&c.)$$

Law of superposition

In the case of crossed plates the thicknesses of those placed transversely are to be regarded as negative, just as in the case of the superposition of crystallized plates. All these results are conformable to the experiments of Dr. Brewster.

The phenomena of contracted and dilated glass may most easily and conveniently be produced by bending a long parallel plate of glass having its longer edges polished, and passing the light through them across its breadth. In this case, as in all cases of flexure, the convex surface is in a state of dilatation, and the concave of compression, while there exists a certain intermediate line or boundary between these oppositely affected regions in which the substance is in its natural state of equilibrium, and on both sides of which neutral line the degree of strain increases as we *recede* from it towards either surface. Fig. 212 is a section of such a bent plate, much exaggerated, through which light, polarized in a plane 45° inclined to its length, has been passed and analyzed as usual. The neutral line is marked by a divided black stripe, and the tints on either side of it descend in Newton's scale, being arranged in stripes disposed according to the lines 11, 22, 33, 44, &c. The tints, however, on opposite sides of the neutral line have opposite colours, being positive on the side of the dilatation, or towards the convexity, and negative on the compressed or concave side. In a plate of glass 1.5 inch broad, 0.28 thick and six inches long, Dr. Brewster developed seven orders of colours before the glass broke with the bending force applied. This experiment affords an exceedingly beautiful illustration of the action of compressing and bending forces on solids, and furnishes ocular evidence of the state of strain into which their several parts

1090.

Tints produced by bending a glass plate

Fig. 212.

State of strain ascertained by the tints

Light. are brought by external violence. The ingenuity of Dr. Brewster has not overlooked its application to the useful and important object of ascertaining the state of strain and pressure on the different parts of architectural structures, as stone bridges, timber framings, &c., by the use of glass models actually put together as the buildings themselves. We must recollect always, however, that the information thus afforded will only be distinct when the load intended to be sustained is many times the weight of the materials.

1091. If a plate of glass be subjected to several distinct compressions and dilatations in different directions, Dr. Brewster finds, that its action will be the same as the combined action of several plates each subjected to one of the forces employed. Thus a square of glass compressed equally on all its four edges exerts no polarizing action.

1092. If a pressure be applied at a single point of a mass of glass, or rather at two opposite points, it will diverge from these points in all directions into the mass, and the lines of equal pressure, which are in fact the isochromatic lines, must have their form determined in some measure by the figure of the compressing screw or tool at its point of contact with the glass, for this figure regulates the form and curvature of the indentation immediately under it. Dr. Brewster has figured several of the curves produced by the application of such pressure to different parts of the same parallelepiped of glass, for which the reader is referred to his Paper, as well as for a variety of beautiful figures produced by crossing plates differently strained.

1093. M. Biot has observed, that in some instances glass maintained in a state of vibration by the action of a bow or otherwise, depolarizes light, *i. e.* restores the vanished pencil. This is a necessary consequence of the alternate compressions and dilatations which follow each other in rapid succession in all the vibrating molecules. Nodal lines (see *ACOUSTICS*) being exempt from such variations of density ought to be marked by black bands, and may thus, perhaps, be rendered evident to the eye.

1094. When masses of jelly (especially of isinglass) are pressed between plates they acquire a polarizing action. If dilated by proper management, and in that state allowed to dry and harden, the character so impressed, according to Dr. Brewster, is permanent when the dilating force is removed; to explain which, we must consider that the exterior coats indurate more rapidly than the interior, and when they have acquired the consistency of a solid, they will be capable of resisting the subsequent contraction of the interior portions and keeping them in a dilated state, even when the original dilating force is removed. That force only served to determine the figure and dimensions of the exterior crust, and when once that crust is fully formed and indurated, it becomes capable of maintaining them without the further aid of the cause which gave them rise. The polarizing power of isinglass thus developed is very great, and even exceeds that of some doubly refractive crystals, such as beryl; a plate of isinglass whose thickness is 624 polarizing the tint which would be reflected by a plate of air whose thickness is unity, while a plate of beryl parallel to the axis, to polarize the same tint, will require a thickness = 720. Glass compressed, or dilated, by an equal force, would require a thickness (according to Dr. Brewster) = 12580 to produce the same tint.

1095. We come now to consider the transient effects of unequal temperature below the softening point of glass. The immediate effect of an increase or diminution of temperature in one point of a piece of glass, is to produce a mechanical strain on all the surrounding part, which if the difference of temperature is considerable, is of the utmost violence, and capable of breaking asunder the thickest pieces of glass; an effect with which every one is familiar. Now, as we know that strain alone develops a polarizing action, the rule of philosophy, "*non plures causas admitti debere*," &c. which forbids the admission of a second cause when one adequate to the effect is known to be in action, will hardly justify us in attributing a peculiar action to the caloric, independent of its power of altering the dimensions of matter.

1096. When a heated iron bar is applied along the edge of a parallelepiped of glass held in a polarized beam, analyzed as usual, the vanished image is restored in various degrees of intensity in different parts of the glass. The *neutral axes* are parallel and perpendicular to the heated edge, and the axes in whose azimuth the tint polarized is the strongest, at 45° of inclination. If held in that azimuth, the first effect of the heat is to produce a line, or, as it were, a wave of white light at the heated edge, which advances gradually upon the glass, driving before it a dark and undefined wave. Nearly at the same instant, and *long before the slightest increase of temperature can have reached the further extremity of the glass plate*, a similar but fainter white wave advances from the edge opposite to the heated one, driving before it a similar undefined dark wave; and at no perceptible interval of time another white fringe appears in a very diluted state about the centre of the plate, advancing equally towards the heated edge on one side and that most remote on the other, and thus condensing the two undefined dark waves into two black fringes. The white tints are succeeded by tints of a lower order in the scale of colour, yellow, red, purple, blue, &c., till at length the whole scale of the colours of thin plates is seen arranged in four sets of fringes parallel to the heated edge, and having for their origins the black fringes above mentioned. At the same time, other lateral fringes are produced along the edge perpendicular to the heated one. Thus in all six sets are seen; two *exterior*, *viz.* those parallel to the heated edge, and *outside* of the black fringes; two *interior*, in the same direction, but between the black fringes; and two *terminal*, along the lateral edges. The whole phenomena is as represented in fig. 213. The fringes along the heated edge A B are most distinct and numerous, those along the opposite, C D, less so, and the interior and terminal fringes least of all.

1097. As glass is an extremely bad conductor of heat, and as culinary heat is propagated through glass *entirely* by conduction, it follows, that the sudden application of an elevated temperature to the edge A B must produce a dilatation in it, not participated in by the rest of the glass. If, therefore, the stratum of molecules A B were detached from the rest of the glass, it would elongate itself so as to project at its two ends beyond the edges A C, D B. When the heat of this stratum communicated itself to the next, that also would elongate itself, but in a less degree; and thus after a very long time, during which the heat had penetrated to the further extremity of the glass, its outline would assume the form *a C D b*, the lines *a C*, *b D* being certain curves depending on the law of propagation and the time elapsed. This would be the state of things were the glass plate composed

Fig. 213.

Light. of discrete strata, each of which could dilate independently of all the rest. And since in each of these (regarded as infinitely thin) the temperature and strain would be uniform, there would arise no polarizing action. But, in reality, the case is quite different; every stratum is indissolubly connected along its whole extent with the strata adjacent, and can neither expand nor contract without forcing them to participate in its change of dimension. In so far, then, as two adjacent strata participate in the change of temperature they expand together; but when one is hotter than the other, the former is found to expand *less*, and the other *more* than if they were independent. Now the strain thus induced on any stratum is not, like the caloric which causes it, confined by the conducting power of the medium, but propagates itself *instantly* (with diminished energy) to the strata beyond, by reason of the mutual action of the molecules.

Part IV.

The general problem, then, to investigate the actual state of strain of any molecule at any moment is one of some complexity, inasmuch as it depends at once on the laws of the slow propagation of heat, and the instantaneous but variable participation of change of figure necessary to establish among the particles a momentary equilibrium under the circumstances of temperature at the time; but, without attempting minutely to analyze the effects, if we content ourselves with acquiring a general idea how they arise, we shall find little difficulty. For in fig. 214, if we conceive the stratum $ABba$ adjacent to the border AB to be dilated by the heat, the rest of the glass retaining its original temperature; if this stratum could expand separately, its edges Aa , Bb would project out beyond the general edges Ca , $D\beta$; and if we regard two terminal strata $CAEG$, $DBFH$, as detached from the interior portion $CD\beta a$, and free to move by the force applied at their extremities A , B , they would be raised by the dilatation of the portion $ABba$ into the situation represented in the figure, turning round C , D as fulcrums, and leaving triangular intervals $Ca a$, $D b \beta$ vacant, and in these circumstances there would be no strain on any part of the system. But the cohesion of the glass prevents the formation of these vacancies, and the bars or levers $CAEG$, $DBFH$ cannot move into this situation without dragging with them, and therefore distending the strata of $CD\beta a$. Let PQ be any such stratum, and let it be distended to $p q$. Then by its elasticity it will tend to draw the bars $CAEG$ and $BDHF$ together; and its action will therefore tend, first, to produce a pressure on the fulcrums C , D , urging the points CD together, and therefore bringing the stratum CD into a state of compression. Secondly, to produce also a pressure on Aa , Bb , or a resistance to the dilatation of $ABba$, which its increased temperature would naturally produce. It will therefore tend to compress back the strata of $ABba$ into a smaller length than what would be natural to them in their heated state, *i. e.* to bring them also into a relatively compressed state. Thirdly, the tension of $p q$ being sustained at C , D and A , B , will tend to bend inwards the levers $ACGE$, $BDFH$, rendering them concave at the edges GE , HF , and convex at CA , DB , and thus distending the lines CA , DB , and compressing the strata adjacent to EG , HF .

1098.
State of the various regions of the plate as to strain determined. Fig. 214.

From this reasoning it is clear, that the glass, in consequence of these various strains, will assume a figure concave on all its edges, but chiefly so at the lateral ones AC , DB , as in fig 215; and that the state of strain of its various parts will be as there expressed, all the edges being compressed, but principally AB and CD , and the interior distended. The limit between the distended and compressed portions parallel to AB must necessarily be marked by neutral lines ab , cd on either side of which the strain will increase, being a maximum in the middle and on or near the edges. Consequently, it ought to polarize four sets of fringes, having ab , cd for their origins, and of which the two external (or those between these lines to the edge) ought to have a character opposite to those of the internal, the portion of the intromitted pencil polarized parallel to AB being propagated faster than that parallel to AC in the one case, and slower in the other. This opposition of characters is conformable to Dr. Brewster's observations, who states (*Phil. Trans.* 1816) that the parts of the glass which exhibit the two exterior sets of fringes (adjacent to the edges AB , CD) have "the structure of" attractive crystals, while the parts which exhibit the interior and terminal sets have that of repulsive ones; meaning, of course, in the language of the undulatory doctrine, that the order of velocities of the doubly refracted pencils is reversed in passing from one region of the glass to the other, for of its actual structure we can know nothing. That the terminal fringes ought (as observed) to have the same character as the interior is a necessary consequence of the above reasoning, for the terminal regions DB , AC are compressed in directions parallel to their edges, and therefore perpendicular to the direction in which the central portion is distended; and we have already seen that compression in one direction is equivalent (so far as the character of the tints produced is concerned) to distension in that perpendicular to it.

1099.
Production of fringes of opposite characters. Fig. 215.

The terminal fringes.

Lastly, the black lines separating the terminal fringes from the interior ones, arise from the combined action of the tension of the interior region parallel to AB (fig. 214) exerting itself on any point as q on the inner border of the terminal portion $DBFH$, (which we have regarded as an elastic bar, or lever,) and the distension of the line DB also exerting itself at q , and arising from the convexity given to this line. In virtue of these two forces, every point q in a certain line at a proper distance from the extreme edge HF , will be equally distended in opposite directions, and will therefore be in a neutral state, as to polarization, and, of course, appear black. The terminal fringes are less developed than the rest, because they arise simply from the flexure of the edges HF , GE , which is an indirect effect of the principal force, and is very small, (owing to the small dilatability of glass by heat, and consequent minuteness of the versed sine of the curve into which they are distorted,) a line of indifference separating them from the others lies near the edges; for the same reason, the tension of the convex line DB being small, and therefore putting itself in equilibrium with that of the distended column $p q$ at a point q near its extremity, where it is evident that the strain parallel to $p q$ must be much diminished; the greater portion of the whole tension of $p q$ being resisted by the spring of laminae situated still further from the edge than DB .

1100.
Neutral lines separating adjacent sets of fringes.

If a lamina of glass, uniformly heated, be suddenly cooled at one of its edges, the reverse of all these effects will arise; the outer column $ABab$ (fig. 214) will suddenly contract and compress violently the columns

1101

Light. beyond $\alpha\beta$, from which no heat has yet been abstracted, and drag inwards the ends of the terminal levers EAGC, BFHD, which will thus be violently pressed on the parts βQ and αP as fulcræ; and their action being thus transmitted to the opposite edge CD will tend to lengthen it, and thus bring it, as well as the edge AB, into a distended state. The terminal edges will also be sprung outwards. The strain on every point will be exactly the reverse of what is expressed in fig. 215, and a corresponding inversion of the characters of the tints will take place; all which is agreeable to Dr. Brewster's observation, (Prop. 14 of the *Memoir* cited.)

1102. When a crack takes place in a piece of unequally heated glass, the directions and intensities of the straining forces in every part, which depend wholly on the cohesion of its molecules, and the continuity of the levers, springs, &c. into which it may be mentally conceived to be divided, is suddenly altered; and the fringes are accordingly observed to take instantly a new arrangement, and assume forms related to the figure of that part of the glass which preserves its continuity. To analyze the modifications arising from variations of external figure and different applications of the heat, would be to involve ourselves unnecessarily in a wilderness of complexity. One simple case may, however, be noticed, in which the centre of a circular piece of glass is heated. Each exterior annulus of this will be placed in a state of distension parallel to its circumference, and will compress all within it by a force parallel to the radius. The central point will be neutral, being equally confined in all directions, and the annuli adjacent to the centre will in like manner be compressed both radially and circumferentially. The radial strain continues as we recede from the centre, but the circumferential diminishes, and at length, as already said, changes to a state of distension, and of course passes through a neutral state, thus giving rise to a black circle and concentric fringes of opposite characters, the whole of which will be intersected by the arms of a black cross parallel and perpendicular to the plane of primitive polarization, and which of course remains fixed while the plate is turned round in its own plane.

1103. There is only one experiment of Dr. Brewster which seems hostile to the theory here stated. He made a partial crack with a red-hot iron in a very thick piece of glass, and allowed it to close by long standing, which it did, so as to disappear entirely. In this state, the glass, when unequally heated, exhibited the same fringes, as if no crack had existed; but the moment the crack was opened by a slight heat applied near it, they suddenly changed their figure, and assumed that due to the portion having the crack for a part of its outline. It seems, however, that a very great adhesive force takes place between the surfaces of glass when thus in optical contact; and to those who are aware how the free expansion and contraction of dissimilar metallic bars may be commanded, and the bars in consequence made to ply on change of temperature by mere forcible juxtaposition, without soldering, till the difference of expansion has reached a certain point, when they give way with a snap and regain their state of equilibrium, the anomaly will not appear in the light of a radical objection. (*We think it not improbable that the musical sounds said to issue at sunrise from certain statues, may originate in some pyrometrical action of the kind here alluded to. We have often been amused by a similar effect produced in the bars of the grate of a fire place.*)

1104. Such are, in general, the transient effects of a heat below the softening point of glass, unequally distributed through its substance. But if a mass of glass be heated up to, or beyond that point, so as to allow its molecules to glide with more or less freedom on one another, and adapt themselves to any form impressed on the mass, and then suddenly cooled, either by plunging into water, or by exposure to cold air, the heat is abstracted from its external strata with so much greater rapidity than it can be supplied by conduction from within, that they become rigid, while the inner portions are still soft and yielding. At this instant, there is therefore no strain in any part; but, the abstraction of the heat still going on, the internal parts at length become solid, and tend, of course, to contract in their dimensions. In this, however, they are prevented by the external crust already formed, which acts as an arch or vault, and keeps them distended, at the same time that these latter portions themselves are to a certain extent forced to obey the inward tension, and are strained inwards from their figure of equilibrium. Glass in this state is said to be unannealed. If the cooling has been sudden, and the mass considerable, it either splits in the act of cooling, or flies to pieces, when cold, spontaneously, or on the slightest scratch which destroys the continuity of its surface; and the pieces when put together again (which, however, is seldom practicable, as it usually flies into innumerable fragments, or even to powder, as is familiarly shown in the glass tears called Rupert's drops, which exhibit a very high polarizing energy from their intense strain, and which burst with a violence amounting to explosion, on the rupture of their long slender tails) are found not to fit, but to leave a slight vacancy; thus satisfactorily proving the state of unnatural and violent distension in which its interval parts have been held. The case is precisely analogous to that of a gelatinous substance allowed to indurate under the influence of dilating forces. (See Art. 1094.)

1105. If the cooling be less sudden, and carefully managed, the glass, though much more brittle than ordinary annealed glass, is yet susceptible (with great caution) of being cut and polished; and in this state, if polarized light be passed through it, it exhibits coloured phenomena of astonishing variety and splendour, forming fringes, irises, and patterns of exquisite regularity and richness, according to the form and size of the mass, and the degree of strain to which it is subjected. In all these cases if the external form be varied, the pattern varies correspondingly, as it is easy to perceive it ought; for if any part of the exterior crust be removed, that part of the strain which it sustained will fall on the remainder, and on the new surface produced. Figures 216, 217, and 218, represent the patterns exhibited by a circular, a square, and a rectangular plate of about $\frac{1}{2}$ inch thick, the two latter being placed so as to have one side parallel to the plane of primitive polarization. Figure 219 and 220 represent the patterns shown by the two latter in azimuth 45° , and fig. 221 that arising from the crossing of two plates equal and similar to fig. 220, each being in azimuth 45° . In all these cases the laws of superposition of Art. 1089 are observed, when similar points of similar plates are laid together. If symmetrically, the tints polarized is the same as would be polarized by one plate whose thickness is their sum; if crosswise, their difference.

Case of a circular plate heated in the centre.

Singular effect of a crack allowed to close.

Phenomena of unannealed glass

Rupert's drops.

Patterns exhibited by circular, square, and rectangular unannealed plates.
Fig 216—221.

Light.

If a square or rectangular plate be turned about in its own plane, from azimuth 0° , the arms of the black cross dividing it into four quarters become curved, as in fig. 222, and pass in succession over every part of the disc; thus showing that the positions of the axes of elasticity of the molecules vary for every different point of the plate, and in different parts of it have every possible situation. We shall not here attempt to analyze the mechanical state of the molecules in any case, as it would lead us too far; but merely mention an experiment of Dr. Brewster, which is sufficient to show the conformity of our theory of these figures with fact. According to this excellent observer, the fringes parallel to the edge AB of the rectangle (fig. 220) are similar in their character to those produced by setting the corresponding edge of a similar plate of annealed glass on a hot iron. Now, in the latter case, the exterior fringes adjacent to AB, CD arise from a compressed state of the columns parallel to AB; and the interior, from a distended. And, in the unannealed plate the distribution of the forces is almost exactly similar to that described in Art. 1098 and 1099. In fact, such a plate may be likened, in some respects, to a frame of wood over which an elastic surface is stretched like a drum. The four sides will all be curved inwards by its tension, and they will all be compressed in the direction of their length by the direct tension, independent of the secondary effect produced by their curvature. The terminal fringes in the articles referred to arise solely from the secondary forces thus developed; but the analogy between the cases would be complete, if, instead of supposing the annealed plate heated at one edge only, the heat were applied at all the four simultaneously, by surrounding it with a frame of hot iron. For a farther account of the beautiful and interesting phenomena produced by unannealed glass, we must refer the reader to Dr. Brewster's curious Paper already cited.

M. Fresnel has succeeded in rendering sensible the bifurcation of the pencils produced by glass subjected to pressure, by an ingenious combination of prisms having their refracting angles turned opposite ways, and of which the alternate ones are compressed in planes at right angles to each other, thus (as in the case of the double refraction along the axis of quartz) doubling the effect produced.

The effects produced by unequal heat and pressure on crystallized bodies, in altering their relations to light transmitted through them, are less sensibly marked than in uncrystallized, being masked by the more powerful effects produced by the usual doubly refractive powers. In crystals, however, where these powers are feeble, or in which they do not exist in any sensible degree, (as in fluor spar, muriate of soda, and other crystals which belong to the tessular system, Dr. Brewster has shown that a polarizing and doubly-refractive action is developed by these causes just as in uncrystallized ones; and M. Biot, by applying violent pressure to crystallized substances while viewing through them their systems of rings in the immediate vicinity of their axes where the polarizing action is very weak, has succeeded in producing an evident distortion of the rings from the regularity of their form, thus rendering it manifest, that it is only the extreme feebleness of the polarizing action so induced in comparison with the ordinary action of the crystal, which prevents its becoming sensible in all directions.

In applying what is here said to heat, however, we consider only its indirect action, or that arising from its unequal distribution, inducing a strain, and thus resolving itself into pressure, as above shown. But Professor Mitscherlich in a most interesting series of researches (which we hope, ere long to see embodied in a regular form, but of which at present only the most meagre and imperfect details have reached us) has shown that the action of heat on crystallized bodies, even when uniformly distributed, so that the whole mass shall be at one and the same temperature, is totally different from what obtains in uncrystallized. In the latter (as well as in crystals of the tessular system) an elevation of temperature, common to the whole mass, produces an equal dilatation in all directions, the mass merely increases in dimensions, without change of figure. In crystals, however, not belonging to the tessular system, *i. e.* whose forms are not symmetrical relative to three rectangular axes, the dilatation caused by increase of temperature is so far from being the same in all directions, that in some cases a dilatation in one direction is accompanied with an actual contraction in another.

Of this important fact, (the most important, doubtless, that has yet appeared in pyrometry.) M. Mitscherlich has adduced a remarkable and striking instance in the ordinary Iceland spar, (carbonate of lime.) This substance when heated, *dilates* in the direction of the axis of the obtuse rhomboid which is the primitive form of its crystals, and *contracts* in every direction at right angles to that axis, so that there must exist an intermediate direction, in which this substance is neither lengthened nor contracted by change of temperature. A necessary consequence of such inequality of pyrometric action is, that the angles of the primitive form will undergo a variation, the rhomboid becoming less obtuse as the temperature increases, and this has been ascertained to be the case by direct measurement; M. Mitscherlich having found, that an elevation of temperature from the freezing to the boiling point of water produced a diminution of $8' 30''$ in the dihedral angle at the extremities of the axis of the rhomboid, (*Bulletin des Sciences publié par la Société Philomatique de Paris*, 1824, p. 40.)

M. Mitscherlich assured himself of the fact in question by direct measurement of a plate of Iceland spar parallel to the axis, at different temperatures, by the aid of the "Spherometer," a delicate species of *calibre* contrived by M. Biot for measuring the thickness of any laminar solid by the revolution of a screw whose point is just brought into light contact with the surface, and by which the 10,000th of an inch is readily appreciated and measured. The experiment is necessarily one of great delicacy, but our readers may assure themselves at least of the *general fact* of unequal change of dimension by change of temperature, by a very simple experiment requiring almost no apparatus. Let a small quantity of the sulphate of potash and copper, (an anhydrous salt easily formed by crystallizing together the sulphates of potash and of copper,) be melted in a spoon over a spirit lamp. The fusion takes place at a heat just below redness, and produces a liquid of a dark green colour. The heat being withdrawn, it fixes into a solid of a brilliant emerald green colour, and remains solid and coherent till the temperature sinks nearly to that of boiling water, when all at once its cohesion is destroyed; a commotion takes place throughout the whole mass, beginning from the surface, each molecule, as if animated

Part IV

1106.

Effect of turning round an unannealed plate in its own plane. Fig. 222.

Relation of these phenomena to those of transiently heated annealed plates.

1107.

1108.

Effects of unequal heat and pressure on crystallized bodies.

1109.

Mitscherlich's researches on the dilatation of crystals by heat.

1110.

Pyrometrical properties of Iceland spar.

1111.

Mode of ascertaining them.

Pyrometrical property of sulphate of potash and copper.

Light.

starting up and separating itself from the rest, till, in a few moments, the whole is resolved into a heap of incoherent powder, a result which could evidently not take place, had all the minute and interlaced crystals of which the congealed salt consisted contracted equally in all directions by the cooling process, as in that case their juxtaposition would not be disturbed. Phenomena somewhat similar, and referable to the same principles, have (if we remember right) been encountered by M. Achard in the fusion of various frits for glasses, &c.

1112.

Double refraction of crystals variable with temperature. Remarkable property of sulphate of lime.

The relation of the optical and crystallographical characters of bodies is so intimate, that no change can be supposed to take place in the latter without a corresponding alteration in the former. As the rhomboid of Iceland spar becomes less obtuse by heat, and therefore approximates nearer to the cube, in which the double refraction is nothing, it might be expected that the power of double refraction should diminish, and this result has been verified by M. Mitscherlich by direct measurement. More recently, the same distinguished chemist and philosopher has ascertained the still more remarkable and striking fact, that the ordinary sulphate of lime or gypsum, which, at common temperatures, has two optic axes in the plane of its laminae, inclined at 60° to each other, undergoes a much greater change by elevation of temperature; the axes gradually approaching each other, collapsing into one, and (when yet further heated) actually opening out again in a plane at right angles to the laminae, thus affording a beautiful exemplification of Fresnel's theory of the optic axes as above explained.

1113

This singular result we cite from memory, having in vain searched for the original source of our information; but it might have been expected, from the low temperature at which the chemical constitution of this crystal is subverted, by the disengagement of its water, that the changes in its optical relations by heat would be much more striking than in more indestructible bodies. We have not, at this moment, an opportunity of fully verifying the fact; but we observe, that the tints developed by a plate of sulphate of lime now before us, exposed as usual to polarized light, rise rapidly in the scale when the plate is moderately warmed by the heat of a candle held at some distance below it, and sink again when the heat is withdrawn, which, so far as it goes, is in conformity with the result above stated. Mica, on the contrary, similarly treated, undergoes no apparent change in the position of its axes or the size of its rings, though heated nearly to ignition. The subject is in the highest degree interesting and important, and lays open a new and most extensive field for optical investigation. It is in excellent hands, and we doubt not will, ere long, form a conspicuous feature in the splendid series of crystallographical discovery which has already so preeminently distinguished its author.

§ XIII. *Of the Use of Properties of Light in affording Characters for determining and identifying Chemical and Mineral Species, and for investigating the intimate Constitution and Structure of Natural Bodies.*

1114.

Relation between the refractive powers and chemical composition of bodies.

Newton, who "looked all nature through," was the first to observe a connection between the refractive powers of transparent media and their chemical properties. His well known conjecture of the inflammable nature of the diamond, from its high refractive power, so remarkably verified by the subsequent discovery of its one and only chemical constituent, (*carbon*), was, perhaps, less remarkable for its boldness, at a period when Chemistry consisted in a mere jargon, in which *salt*, *sulphur*, *earth*, *oil*, and *mercury* might be almost indifferently substituted for one another, than it would have been fifty years later. His divination of the inflammable nature of one of the constituents of water is at least equally striking as an instance of sagacity, and even more remarkable, for the important influence which its verification has exercised over the whole science of Chemistry. These instances suffice to show the value of the refractive index, either taken in conjunction with the specific gravity of a medium, or separately as a physical character. The refractive indices of a vast variety of bodies have been ascertained by the labours of Newton and later experimenters, among whom Dr. Brewster and Dr. Wollaston have been the largest contributors to our knowledge. They may be grouped together in a general way, in order of magnitude, as follows:

1115.

Classification of bodies according to their refractive densities.

Class 1. Gases and vapours. Refractive index from 1.000 to 1.002, under ordinary circumstances of pressure and temperature.

Class 2. $\mu = 1.05 \dots \mu = 1.45$. Comprising the condensed gases; ethereal, spirituous, and aqueous liquids; acid, alkaline, and saline solutions, (not metallic.)

Class 3. Comprising, first, almost all unctuous, fatty, waxy, gummy, and resinous bodies; camphors, balsams, vegetable and animal inflammables, and all the varieties of hydro-carbon. Secondly, stones and vitreous compounds, in which the alkalis and lighter alkaline earths in combination with silica, alumina, &c. are the predominant ingredients. Thirdly, saline bodies not having the heavy metals, or the metallic acids predominant ingredients. $\mu = 1.40 \dots 1.60$.

Class 4. Pastes, (glasses with much lead,) and, in general, compounds in which lead, silver, mercury, and the heavy metals, or their oxides abound. Precious stones, simple combustibles in the solid state, including the metals themselves.

$\mu = 1.60$ and upwards.

These classes, however, admit of so many exceptions and anomalies, and are themselves so vague and indefinite, that we shall not attempt to distribute the observed indices under any of them, but rather prefer, for convenience of reference, presenting the whole list in the form of a Table, arranged in order of magnitude, in which all these classes are mingled indiscriminately—a form, in some measure, consecrated by usage.

Light.

Part IV.

Table of Refractive Indices, or Values of μ for Rays of Mean Refrangibility, (unless expressed to the contrary.)
Dr. Wollaston's results, however, are all (according to Dr. Young, Philosophical Transactions, vol. xcii. p. 370.) to be regarded as belonging to the Extreme Red Rays.

1116.

N. B. In this Table the authorities are referred to as follows :

Br. Brewster, *Encyclop. Ed.*, and *Treatise on New Philosophical Instruments.* Bos. Bosovich.
 B. Y. Dr. Young's Calculations of Dr. Brewster's Unreduced Observations. *Quarterly Journal*, vol. xxii.
 Bi. Biot. F. Faraday. Du. Dulong. M. Malus. N. Newton. Fr. Fraunhofer.
 W. Wollaston, *Phil. Trans.* He. From our own observation. Eul. Euler the younger.
 C. and H., authorities cited by Dr. Young in his *Lectures*.

Vacuum 1.000000

GASES,

at the freezing temperature and pressure = $29^{\text{in}}.922 = 0^{\text{m}}.76$.

Hydrogen	1.000138	Du.
Oxygen	1.000272	Du.
Atmospheric air	1.000294	Bi.
Azote	1.009300	Du.
Nitrous gas	1.000303	Du.
Carbonic oxide	1.000340	Du.
Ammonia	1.000385	Du.
Carburetted hydrogen	1.000443	Du.
Carbonic acid	1.000449	Du.
Muriatic acid	1.000449	Du.
Hydrocyanic acid	1.000451	Du.
Nitrous oxide	1.000503	Du.
Sulphuretted hydrogen	1.000644	Du.
Sulphurous acid	1.000665	Du.
Olefiant gas	1.000678	Du.
Chlorine	1.000772	Du.
Protophosphuretted hydrogen	1.000789	Du.
Cyanogen	1.000834	Du.
Muriatic ether	1.001095	Du.
Phosgene	1.001159	Du.
Vapour of sulphuret of carbon	1.001500	Du.
Vapour of sulphuric ether (boiling point at 35° centig.)	1.001530	Du.

LIQUIDS AND SOLIDS.

Ether expanded by heat to three times its volume	$\mu = 1.0570$	Br.
Tabasheer from Vellore, a yellowish transparent variety	1.1111	Br.
First new fluid discovered by Dr. Brewster in cavities in topaz	1.1311	Br.
Tabasheer, transparent, from Nagpore	1.1454	Br.
Ditto ditto ditto another specimen	1.1503	Br.
Ditto, whitest variety, from Nagpore	1.1825	Br.
New fluid discovered by Dr. Brewster in amethyst, at 83 $^{\circ}$ Fahr.	1.2106	Br.
Second new fluid discovered by Dr. Brewster in topaz, at 83° Fahr.	1.2946	Br.
Nitrous oxide liquefied by pressure	{ much less than water }	F.
Muriatic acid gas ditto ditto	{ both much less than water }	F.
Carbonic acid gas ditto ditto	{ 1.307 Br. 1.3085 Br. 1.3100 W. }	
Ice		
Chlorine liquefied by pressure	{ rather less than water }	F.
Cyanogen liquefied by pressure	{ perhaps less than water }	F.
Ditto ditto	1.316	Br.
Sulphurous acid liquefied by pressure	{ equal to water }	F.
Water	1.336	{ N. W. Br. }
Sulphuretted hydrogen liquefied by pressure ..	{ rather greater than water }	F.
Ammonia liquefied by pressure	{ greater than water, and greater than all the other liquefied gases }	F.
Aqueous humour of the eye	1.3366	Br.
Ditto of the haddock	1.341	B. Y.
Vitreous ditto	1.336	W.

Vitreous humour of the haddock	1.3394	Br.
Ditto	1.340	B. Y.
Ditto of the lamb	1.345	B. Y.
Ditto of the pigeon	1.353	B. Y.
Saliva	1.339	B. Y.
Expectorated mucus	1.339	B. Y.
Salt water (1 sea water)	1.343	Br.
Cryolite	{ 1.344 } Br.	
Vinegar (distilled)	1.349	
Ditto	1.344	Eul.
Vinegar	1.372	H.
Acetic acid (? strength)	1.347	B. Y.
Jelly fish (Medusa <i>Æquora</i>)	1.396	Br.
White of egg	1.345	Br.
Port wine	1.351	Eul.
Human blood	1.351	B. Y.
Saturated aqueous solution of alum	1.354	He.
Oil of box-wood	1.356	B. Y.
Ether	{ 1.358 } W.	
Albumen	1.374	B. Y.
White of a hen's egg	1.360	W.
Brandy	{ 1.361 } Br.	
Rum	1.359	B. Y.
Oil of ambergris	1.360	B. Y.
Alcohol	{ 1.368 } Br.	
Ditto (S. G. 0.866)	1.379	B. Y.
Ditto	1.37	W.
Ditto (rectified spirits)	1.370	N.
Ditto	1.371	C.
Ditto	1.372	He.
Ditto	1.374	Br.
Ditto	1.377	B. Y.
Saturated solution of salt	1.377	B. Y.
Muriatic acid (? S. G.)	1.375	C.
Ditto (S. G. 1.134)	1.376	Br.
Ditto	1.392	He.
Ditto	1.395	B. Y.
Ditto (strong)	1.401	Br.
Ditto (highly concentrated)	1.4098	Bi.
Oil of wine	1.379	B. Y.
Sweet spirit of nitre	1.384	He.
Cornea of a lamb	1.386	B. Y.
Malic acid	1.395	Br.
Pus	1.395	B. Y.
Nitrous (?) acid (? strength)	{ 1.396 } Br.	
Nitric acid (? strength)	1.404	B. Y.
Crystalline lens of the eye (human?) outer coat	1.406	Br.
Ditto ditto middle coat	1.3767	Br.
Ditto ditto centre	1.3786	Br.
Ditto of the lamb's eye, outer coat	1.3990	Br.
Ditto ditto middle coat	1.386	B. Y.
Ditto ditto centre	1.428	B. Y.
Ditto of the haddock's eye, outer coat	1.436	B. Y.
Ditto ditto middle coat	1.410	B. Y.
Ditto ditto centre	1.439	B. Y.
Ditto of the ox	{ 1.380 } W.	
Ditto ditto	1.447	
Ditto of the pigeon	1.463	Eul.
Juice of orange peel	1.406	B. Y.
Solution of potash, S. G. 1.416, (ray E)	1.403	B. Y.
Nitric acid (S. G. 1.48)	1.40563	Fr.
Nitric acid	1.410	B. Y.
Hydrate of soda melted by heat	1.410	W.
Hydrophosphoric acid ditto	1.412	C.
Phosphoric acid (fluid)	1.411	B. Y.
	1.423	B. Y.
	1.426	Br.

Light.	Gluten of wheat, dried	1.426	B. Y.	Oil of spearmint	1.481	Br.	Part IV.
	Fresh yolk of an egg	1.428	B. Y.		1.496	B. Y.	
	Sulphuric acid (S. G. 1.7)	1.429	N.	Oil of lemon	1.481	Br.	
	Ditto ditto (? S. G.)	1.430	He.		1.489	B. Y.	
	Ditto	1.435	W.	Carbonate of potash (?)	1.482	Br.	
	Ditto	1.440	Br.	Oil of pennyroyal	1.482	Br.	
	Fluor spar	1.433	W.	Ditto	1.485	B. Y.	
		1.436	Br.	Linseed oil (S. G. 0.932)	1.482	N.	
		1.433	Br.	Linseed oil	1.485	W.	
	Oil of rhue	1.449	B. Y.	Ditto	1.487	B. Y.	
	Phosphorous acid	1.437	B. Y.	Oil of savine	1.482	Br.	
		1.441	B. Y.		1.482		
	Hydrophosphoric acid, cold	1.442	Br.	Oil of juniper	1.491	B. Y.	
		1.446	W.		1.483	Br.	
	Spermaceti (melted)	1.454	B. Y.	Sulphate of ammonia and magnesia	1.483	B. Y.	
		1.452	C.	Train oil	1.485		
	Oil of wax	1.453	Br.	Oil of wormwood	1.489	B. Y.	
	Oil of wormwood	1.453	Br.	Ditto	1.489		
	Bees wax, melted	1.453	B. Y.	Castor oil	1.485	B. Y.	
	Oil of chamomile	1.457	Br.		1.490	Br.	
	Ditto	1.476	B. Y.	Florence oil	1.485	B. Y.	
	Oil of lavender	1.457	Br.	Oil of thyme	1.486	B. Y.	
	Ditto	1.467	W.	Oil of dill seed	1.487	B. Y.	
	Ditto	1.475	B. Y.	Oil of feugreek (? fenugreek)	1.487	Br.	
	Alum.	1.457	W.	Ditto	1.488	B. Y.	
	Ditto (S. G. 1.714)	1.458	N.		1.487	W.	
	Ditto	1.488	B. Y.	Camphor	1.496	B. Y.	
	Tallow (melted)	1.460	W.		1.500	C.	
	White wax (melted)	1.462	B. Y.		(S. G. = 0.996)	1.500	N.
		1.467	Br.	Oil of hyssop	1.487	Br.	
	Oil of poppy	1.483	B. Y.		1.495	B. Y.	
	Sulphate of magnesia (double? least refraction)	1.465	Br.	Windsor soap	1.487	B. Y.	
		1.467	N.	Obsidian	1.488	Br.	
	Borax, (S. G. = 1.714)	1.467	C.	Iceland spar...weakest refraction	1.488	W.	
		1.475	Br.		1.519	B.	
	Oil of peppermint	1.468	W.	Ditto	1.657	W.	
		1.473	B. Y.	Ditto	1.665	Br.	
		1.469	Br.	Ditto	1.6543	M.	
	Oil of rosemary	1.472	B. Y.	Ditto	1.4833	M.	
		1.470	Br.	Ditto	(S. G. = 2.72)	1.667	N.
	Oil of spermaceti	1.473	B. Y.	Sulphate of magnesia (? greatest refraction)	1.488	Br.	
		1.469		Nut oil (perhaps impure)	1.490	He.	
	Oil of almonds	1.470	W.	Ditto	1.507	Br.	
		1.481	B. Y.	Oil of castor	1.490	Br.	
	Ditto	1.483	Br.		1.49	W.	
	Oil of turpentine, rectified	1.470	W.	Tallow (cold)	1.492	B. Y.	
	Spirit of turpentine, (S. G. 0.874)	1.471	N.		1.483	B. Y.	
	Oil of turpentine	1.475	Br.	Oil of caraway seed (carui seminis)	1.491	Br.	
	Ditto	1.476	B. Y.		1.490	B. Y.	
	Ditto (common)	1.476	W.	Oil of marjoram	1.491	Br.	
	Ditto	1.482	C.		1.491	B. Y.	
	Ditto	1.485	B. Y.	Oil of nutmeg	1.497	W.	
	Ditto (common)	1.486	He.		1.491	B. Y.	
	Ditto, S. G. = 0.885, (ray E)	1.47835	Fr.	Nut oil	1.507	Br.	
		1.467*	N.		1.491	B. Y.	
		1.469	W.	Oil of Angelica	1.493	Br.	
	Oil of olives	1.470	Br.		1.492	B. Y.	
		1.4705	He.	Bees wax, cold	1.492	B. Y.	
		1.476	B. Y.	Bees wax	1.507	B. Y.	
		1.471	Br.	Ditto 14° Reaum.	1.5123	M.	
	Oil of bergamot	1.473	B. Y.	Ditto, melting	1.4503	M.	
		1.471	Br.	Ditto, boiling	1.4416	M.	
	Oil of beech, misprinted? oil of brick	1.471	Br.	Ditto	1.542	W.	
		1.470	B. Y.	Ditto (white wax, cold)	1.535	W.	
	Oil of brick, distilled from spermaceti	1.471	Br.	Sulphate of iron, greatest refraction	1.494	Br.	
		1.473	Br.		1.494	B. Y.	
	Oil of juniper	1.473	Br.	Balsam of sulphur	1.497	Br.	
	Butter, cold	1.474	B. Y.		1.475	W.	
		1.480	W.	Sulphate of potash	1.509	Br.	
	Palm oil	1.475	B. Y.		1.495	B. Y.	
		1.475	{ B. Y.	Honey	1.495	B. Y.	
	Oil of rape seed	1.475	Br.	Rochelle salt (mean green rays)	1.4985	He.	
		1.475	B. Y.	Ditto (mean red)	1.4929	He.	
	Naphtha	1.476	W.	Ditto (tartrate of potash and soda)	1.515	Br.	
	Essence of lemon	1.476	N.	Treacle	1.500	W.	
	Gum arabic (S. G. = 1.375)	1.476	N.	Yolk of an egg (dry)	1.500	B. Y.	
	Oil of dill seed	1.477	Br.	Oil of beech nut	1.500	Br.	
	Oil of thyme	1.477	Br.		1.500	Br.	
		1.478	B. Y.	Oil of rhodium	1.503	B. Y.	
	Oil of cajeput	1.483	Br.		1.505	Br.	
		1.479	Br.	Glass, plate and crown, various specimens:			
	Opal (partly hydrophanous)	1.479	Br.	Ditto English plate	1.500	W.	
	Naples soap	1.479	B. Y.	Ditto French plate	1.504	W.	
	Oil of mace, melted	1.481	B. Y.	Ditto English plate (extreme red)	1.5133	He.	
				Ditto plate	1.514	Bos.	

* The S. G. of Newton's specimen was 0.913.

Light.

Part IV.

Glass, Dutch plate	1.517	W.
Ditto crown, common	1.525	W.
Ditto plate	1.526	Bos.
Ditto crown, a prism by Dollond, (extra red)	1.526	He.
Ditto plate	1.527	Br.
Ditto ditto	1.529	Bos.
Ditto crown, another prism by Dollond, (extra red)	1.5301	He.
Ditto Fraunhofer's crown, No. 13, (ray E.)		
S. G. 2.535	1.5314	Fr.
Ditto yellow plate, S. G. 2.52	1.532	C.
Ditto Fraunhofer's crown, No. 9, (ray E.)		
S. G. 2.535	1.5330	Fr.
Ditto Radcliffe crown	{ 1.533 } { 1.536 }	W.
Ditto crown	1.534	Br.
Ditto plate	1.538	Bos.
Ditto ditto	1.542	Bos.
Ditto St. Gobin	1.543	W.
Ditto crown	1.544	Br.
Ditto old plate	1.545	W.
Ditto vulgar	1.550	N.
Ditto Fraunhofer's crown, M, S. G. 2.756, (ray E)	1.5631	Fr.
Ditto plate, (S. G. 2.76)	1.573	C.
Ditto bottle	1.582	Br.
N. B. It is probable that the more refractive specimens of this list are low flint glasses, containing lead.		
Spermaceti, cold	{ 1.503 } { 1.535 }	B. Y.
Oil of pimento	{ 1.503 } { 1.507 } { 1.510 }	B. Y. Br. B. Y.
Starch (dry)	1.504	B. Y.
Muriate of antimony, dry	1.504	B. Y.
Ditto, two days exposed	1.613	B. Y.
Oil of amber	{ 1.505 } { 1.507 }	W. B. Y.
Birdlime	1.506	B. Y.
Oil of sweet fennel seed	{ 1.506 } { 1.507 } { 1.507 }	Br. B. Y. W.
Balsam of Copaiba	{ 1.514 } { 1.516 }	B. Y.
Stilbite	1.528	Br.
Oil of cummin	1.508	Br.
Scammony	{ 1.508 } { 1.578 }	Br. B. Y.
Oil of mace	1.510	B. Y.
Gum Arabic	{ 1.512 } { 1.526 }	B. Y.
Ditto (not quite dry)	1.512	Br.
Ditto	1.513	B. Y.
Ditto	1.514	W.
Human cuticle	{ 1.514 } { 1.517 }	W.
Nitre, greatest index	1.514	Br.
Ditto, least	1.335	Br.
Ditto	1.524	C.
"Niter" (?) S. G. 1.9	1.524	N.
Dantzic vitriol (sulphate of iron)	1.515	N.
Nadelstein from Faroe	1.5153	Br.
Mesotype, least index	1.516	Br.
Ditto, greatest	1.522	Br.
Sulphate of zinc, ordinary refraction	1.517	Br.
Myrrh	{ 1.517 } { 1.524 }	B. Y. Br.
Tartaric acid, least refraction	1.518	Br.
Ditto, greatest	{ 1.529 } { 1.575 }	Br. Br.
Gum dragon (Tragacanth)	1.520	Br.
Glass of borax 1, sillex 2	{ 1.66 } { 1.522 }	W. Br.
Gum lac, or Shell lac	{ 1.52+ } { 1.525 }	W. Br.
Oil of sassafras	1.528	B. Y.
	{ 1.522 } { 1.532 }	B. Y. Br.
	1.536	W.
	1.544	Eul.
	1.524	W.
Caoutchouc	{ 1.534 } { 1.557 }	Br. B. Y.

Selenite	1.525	W.
Ditto, greatest refraction	1.536	Br.
"A selenites, S. G. 2.252"	1.488	N.
Citric acid	1.527	Br.
Leucite	1.527	Br.
	1.528	W.
Canada balsam	{ 1.532 } { 1.549 }	B. Y. Br.
Balsam of Gilead	1.529	B. Y.
Crystalline of ox (dried) and of a fish	1.530	W.
	1.531	W.
Pitch	{ 1.581 } { 1.586 } { 1.588 }	B. Y. Br. B. Y.
Sulphate of copper, least refraction	1.531	Br.
Ditto ditto, greatest	1.552	Br.
Olibanum	{ 1.532 } { 1.544 }	B. Y. Br.
Brazil pebble, (S. G. 2.62)	1.532	C.
Glass of phosphorus (fused phosphoric acid)	1.532	Br.
Solid phosphoric acid	1.544	Br.
Glass of borax (fused borax)	1.532	Br.
Manna	1.533	B. Y.
Ditto, burned	{ 1.547 } { 1.565 }	B. Y.
Arragonite, extraordinary index	1.5348	M.
Ditto, ordinary	1.6931	M.
Elemi	{ 1.535 } { 1.547 }	W. Br.
	1.550	B. Y.
Mastic	{ 1.535 } { 1.539 }	W. B. Y.
	1.560	Br.
Arseniate of potash	1.535	W.
Gum anime	{ 1.535 } { 1.546 }	W. B. Y.
	1.535	W.
Copal	{ 1.549 } { 1.553 }	Br. B. Y.
Oil of cloves	1.535	W.
White sugar	1.539	B. Y.
Ditto	1.535	W.
Ditto (melted)	1.541	B. Y.
Ditto (after melting)	1.545	B. Y.
Felspar	1.555	Br.
Oil of cashew nut	1.536	Br.
Oil of aniseed	1.536	B. Y.
Mellite	{ 1.601 } { 1.538 }	Br. Br.
Ditto	1.556	Br.
Gum juniper	{ 1.538 } { 1.541 }	Br. B. Y.
Carbonate of barytes, least refraction	1.540	Br.
Box-wood	1.542	W.
Colophony	1.543	W.
Apophyllite, the variety which exhibits white and black rings (<i>Leucocyclite</i>)	1.5431	He.
Carbonate of strontia, least refraction	1.543	Br.
Ditto ditto greatest	1.700	Br.
Dichroite (iolite)	1.544	Br.
Petroleum	1.544	B. Y.
Sal gemmæ, S. G. 2.143 (rock salt)	1.545	N.
Ditto (rock salt)	1.557	Br.
Chio turpentine	{ 1.545 } { 1.557 }	B. Y. Br.
Gum sagapenum	1.545	B. Y.
Turpentine	1.545	B. Y.
Burgundy pitch	{ 1.546 } { 1.560 }	B. Y. Br.
Gum thus	{ 1.546 } { 1.554 }	B. Y. Br.
Oil of tobacco	1.547	Br.
Rock crystal	1.547	W.
Quartz, ordinary refractive index	1.5484	M.
Ditto, extraordinary	1.5582	M.
Amethyst	1.562	W.
Rock crystal (double)	1.562	Br.
Crystal of the rock (S. G. 2.65)	1.563	N.
Rock crystal	{ 1.568 } { 1.575 }	C.

Light.				Part IV.	
Amber	1.547	W.	Sulphate of barytes	1.6468	M.
Ditto, (S. G. 1.04)	1.556	N.	Ditto ditto ordinary refraction (along the		
	1.548	B. Y.	axis) for yellow green rays	1.6460	He.
Resin	1.552	B. Y.	Ditto, another specimen, ditto, red rays	1.6459	He.
	1.559	Br.	Ditto ditto for yellow green	1.6491	He.
Guaiacum	1.550	B. Y.	A "pseudo topazius" (S. G. 4.27) sulphate of		
Glue, nearly hard	1.553	B. Y.	baryta	1.643	N.
Chalcedony	1.553	Br.	Sulphate of barytes	1.646	W.
Comptoonite	1.553	Br.	Ditto ditto double, greater refraction	1.664	Br.
	1.559	B. Y.		1.624	B. Y.
Opium	1.57—	W.	Oil of cassia	1.631	B. Y.
Hyposulphate of lime (mean red)	1.5611	He.		1.641	Br.
Ditto, mean yellow green	1.566	He.	Muriate of ammonia	1.625	Br.
Dragon's blood	1.562	B. Y.	Aloes	1.634	B. Y.
	1.565	Br.	Opal coloured glass	1.635	Br.
Horn	1.58—	W.	Euclase, ordinary index	1.6429	Bi.
Pink, coloured glass	1.570	Br.	Ditto, extraordinary	1.6630	Bi.
Assafoetida	1.575	B. Y.	Sulphate of strontia	1.644	Br.
	1.576	Br.	Hyacinth red glass	1.647	Br.
Flint glass (various specimens)	1.578	He.	Mother of pearl	1.653	Br.
	1.583	W.	Spargelstein	1.657	Br.
Ditto, a prism by Dollond (extreme red)	1.584	He.	Epidote, least refraction	1.661	Br.
Ditto, (extreme red)	1.585	He.	Ditto, greatest	1.703	Br.
Ditto, another specimen	1.586	W.	Tourmaline	1.668	Br.
Ditto	1.590	Bos.	Cryolite, least refraction	1.668	Br.
Ditto	1.593	Bos.	Ditto, greatest	1.685	Br.
Ditto	1.594	Bos.	Chloruret of sulphur	1.67—	He.
Ditto	1.596	Br.	Nitrate of bismuth, least refraction, about	1.67—	He.
Ditto, a prism by Dollond (extreme red)	1.601	He.	Ditto, greatest, about	1.89—	He.
Ditto ditto, marked "heavy," (extreme red)	1.602	He.	Sulphuret of carbon	1.678	Br.
	1.604	Br.	Orange coloured glass	1.695	Br.
Ditto, another specimen	1.604	Bus.	Boracite	1.701	Br.
Ditto	1.605	M.	Glass tinged red with gold	1.715	Br.
Ditto, Fraunhofer's No. 3 (ray E)	1.6145	Fr.	Glass, lead 1, flint 2	1.724	Br.
Ditto, another variety	1.616	Br.	Deep red glass	1.729	
Ditto ditto	1.625	Bos.	Nitrate of silver, least refraction	1.729	Br.
Ditto, Fraunhofer's No. 30 (ray E)	1.6374	Fr.	Ditto, greatest	1.788	Br.
Ditto ditto No. 23 (ray E)	1.6405	Fr.	Glass, lead 3, flint 4	1.732	Br.
Ditto ditto No. 13 (ray E)	1.6420	Fr.	Hypophosphite of soda and silver, least refraction	1.735	He.
Anhydrite, ordinary index	1.5772	Bi.	Ditto, greatest	1.785	He.
Ditto, extraordinary	1.6219	Bi.	Axinite	1.735	Br.
	1.578	B. Y.	Nitrate of lead	1.758	Br.
Gum ammoniac	1.592	Br.	Cinnamon stone	1.759	Br.
Hyposulphite of lime, least refraction	1.583	He.	Chrysoberyl	1.760	Br.
Ditto, greatest	1.628	He.		1.756	He.
Balsam of styrax	1.584	Br.	Spinelle	1.761	Br.
Emerald	1.585	Br.		1.812	W.
	1.586	W.	Felspar	1.764	Br.
Benzoin	to 1.596		Sapphire, (white)	1.768	W.
	1.589	B. Y.	Ditto, (blue)	1.794	Br.
Oil of cinnamon	1.604	B. Y.	Rubellite	1.768	He.
	to 1.632			1.779	Br.
Tortoise shell	1.591	Br.	Ruby	1.779	Br.
	1.593	B. Y.	Jargon (orange coloured)	1.782	Br.
Balsam of Peru	1.597	Br.	Glass, lead 1, flint 1 (Zeiber)	1.787	Ze.
	1.605	B. Y.	Pyrope	1.792	Br.
	1.596	W.	Labrador hornblende	1.80±	He.
Guaiacum	1.600	B. Y.	Muriate of antimony (variable) about	1.8	W.
	1.619	Br.	Arsenic	1.811	W.
Beryl	1.598	Br.	Carbonate of lead, least refraction	1.813	Br.
	1.60—	W.	Ditto ditto greatest	2.084	Br.
Balsam of Tolu	1.610	B. Y.	Borate of lead, fused and cooled (extreme red)	1.866	He.
	1.627	B. Y.	Sulphate of lead	1.925	Br.
	1.628	Br.	Glass, lead 2, sand 1	1.987	W.
Ruby red glass	1.601	Br.	Zircon	1.95	W.
Essential oil of bitter almonds	1.603	Br.	Ditto, least refraction	1.961	Br.
Meionite	1.606	Br.	Ditto, greatest	2.015	Br.
Purple coloured glass	1.608	Br.	Sulphur (Haüy)	1.958	Ha.
Resin of jalap	1.608	B. Y.	Ditto	2.008	B. Y.
Hyposulphite of strontia, least refraction	1.608	He.	Ditto	2.04	W.
Ditto ditto greatest	1.651	He.	Ditto, native	2.115	Br.
Colourless topaz	1.6102	Bi.	Ditto, melted	2.148	Br.
Bluish topaz (clairgorm)	1.624	Br.	Calomel	1.970	Br.
Brazilian topaz, ordinary index	1.6325	Bi.	Tungstate of lime, least refraction	1.970	Br.
Ditto ditto, extraordinary	1.6401	Bi.	Ditto, greatest	2.129	Br.
Blue topaz, Aberdeen	1.636	Br.		1.980	W.
Yellow topaz	1.638	Br.	Glass of antimony	2.216	Br.
Red topaz	1.652	Br.	Glass, lead 3, flint 1 (by Zeiber)	2.028	Z.
Green coloured glass	1.615	Br.	Scaly oxide of iron	2.1—	V.
	1.620	B. Y.	Silicate of lead, atom to atom, extreme red	2.123	He.
Castor	1.626	Br.		2.125	B. Y.
Sulphate of barytes, ordinary	1.6201	Bi.	Phosphorus	2.224	Br.
Ditto, extraordinary	1.6352	M.		2.260	Br.

Light.

Nitrite of lead (biaxial, ?quadro-nitrite) in six-sided prisms, ordinary refraction	2.322	He.
Diamond (S. G. = 3.4)	2.439	N.
Ditto	2.470	Br.
Ditto (brown coloured)	2.487	Br.
Ditto (examined by Rochon)	2.755	Ro.
Plumbago	{ from 2.04	W.
	{ to... 2.44	W.

Chromate of lead	least refraction	{ 2.479	Br.
	greatest	{ 2.500	Br.
Octohedrite		2.503	Br.
Realgar, artificial		2.926	Br.
Red silver ore		2.974	Br.
		2.500	Br.
		2.549	Br.
		2.564	Br.
Mercury (prchable, see Art. 594)		5.829	

Part IV.

In casting our eyes down the foregoing Table, we cannot but be struck with the looseness and vagueness of those results which refer to bodies whose chemical nature is in any respect indeterminate. The refractive indices assigned to the different oils, acids, &c. though no doubt accurately determined for the particular specimens under examination, are yet, as scientific data, deprived of most of their interest from the impossibility of stating precisely what *was* the substance examined. Most of the fixed oils are probably (as appears from the researches of Chevreul) compounds, in very variable proportions of two distinct substances, a solid, concrete matter, (stearine,) and a liquid, (elaine,) and it is presumeable, that no two specimens of the same oil agree in the proportions. This is, probably, peculiarly the case with the oil of anise seed, which congeals almost entirely with a very moderate degree of cold. An accurate reexamination of the refractive and dispersive powers of natural bodies of strictly determinate chemical composition, and identifiable nature, though doubtless a task of great labour and extent, would be a most valuable present to optical science. Fraunhofer's researches have shown to what a degree of refinement the subject may be carried, as well as the important practical uses to which it may be applied. The high refractive power of oil of cassia, accompanied by a corresponding dispersion, has led Dr. Brewster to conceive the existence in it of some peculiar chemical element not yet cognisable by analysis. The low refractions of the oils of box-wood and ambergris are not less remarkable. It is among the artificial salts, however, that the widest field is open for the application of precise research, and one in which a rich harvest of important results would, in all probability, amply repay the trouble of the investigation, whether considered in an optical, a chemical, or a crystallographical point of view.

1117.
Remarks on
the Table
of Refractive
Indices.

The fraction $P = \frac{\mu^2 - 1}{s}$, where μ is the refractive index, and s the specific gravity of the medium, expresses (in the doctrine of emission) the intrinsic refractive energy of its molecules, *supposing the ultimate atoms of all bodies equally heavy*. The following results have been stated by various authors, as its values for bodies most widely differing in their chemical and mechanical relations.

1118.
Table of
intrinsic
Refractive
Powers.

I. Gases, taking the value of P for atmospheric air as unity. (From Biot's *Précis Élémentaire*, ii. 224.)

Oxygen	0.86161	Azote	1.03408	Carburetted hydrogen..	2.09270
Air	1.00000	Muriatic gas	1.19625	Ammonia.....	2.16251
Carbonic acid	1.00476	Supercarburetted hydrogen	1.81860	Hydrogen	6.61436

II. Direct values of P given by the formula.

Those marked Dulong are computed from the refractive indices of Dulong in the last table.

Tabasheer	0.0976	Brewster.	Muriatic acid glass	0.5514	Dulong.	Vapour of sulph. ether..	0.9138	Dulong.
Cryolite	0.2742	Brewster.	Sulphuric acid	0.6124	Newton.	Protophosphuretted hydr.	0.9680	Dulong.
Flour spar	0.3426	Brewster.		0.6424	Malus.	Ammonia	1.0032	Dulong.
Oxygen	0.3799	Dulong.	Calcareous spar	{ 0.6536	Newton.	Rectified spirits of wine	1.0121	Dulong.
Sulphate of barytes ..	{ 0.3829	Dulong.	Sal gem	0.6477	Newton.	Carbonate of potash ..	1.0227	Brewster.
	{ 0.3979	Newton.	*Muriate of soda.....	1.2086	Brewster.	Chromate of lead	1.0436	Brewster.
Sulphurous acid gas ..	0.44548	Dulong.	Alum	0.6570	Newton.	Olefant gas.....	1.0654	Dulong.
Nitrous gas	0.44911	Dulong.	Nitric acid	0.6676	Brewster.	*Muriate of ammonia..	1.1290	Brewster.
	{ 0.4528	Dulong.	Borax	0.6716	Newton.	Carburetted hydrogen	1.2204	Dulong.
Air	{ 0.4530	Biot.	Niter	0.7079	Newton.	Camphor	1.2551	Newton.
	{ 0.5208	Newton.	*Nitre	1.1962	Brewster.	Olive oil	1.2607	Newton.
Carbonic acid	0.45372	Dulong.	Hydrocyanic acid	0.7366	Dulong.	Linseed oil	1.2819	Newton.
Azote	0.4734	Dulong.	Ruby	0.7389	Brewster.	Bees wax	1.3308	Malus.
Chlorine	0.48133	Dulong.	Dantzic vitriol, (sul. iron)	0.7551	Newton.	Spirit of turpentine ..	1.3222	Newton.
Glass of antimony	0.4864	Newton.	Muriatic ether (vapour)	0.7552	Dulong.	Amber	1.3654	Newton.
Nitrous oxide	0.5078	Dulong.	Brazilian topaz	0.7586	Brewster.	Octohedrite.....	1.3816	Brewster.
Phosgen	0.5188	Dulong.	Rain water	0.7845	Newton.	Diamond	1.4566	Newton.
Selenite	0.5386	Newton.	Flint glass (mean) ..	0.7986	Brewster.	Realgar	1.6666	Brewster.
Carbonic oxide	0.5387	Dulong.	Cyanogen	0.8021	Dulong.	Amhergris	1.7000	Brewster.
Quartz	0.5415	Malus.	Sulphuretted hydrogen .	0.8419	Dulong.	Mercury (probable)....	2.4247	
Rock crystal	{ 0.5450	Newton.	Gum Arabic	0.8574	Newton.	Sulphur	2.2000	Brewster.
	{ 0.6536	Brewster.	Vapour of sulphuret of			Phosphorus.....	2.8857	Brewster.
Vulgar glass	0.5436	Newton.	carbon	0.8743	Dulong.	Hydrogen	3.0953	Dulong.

The results marked with an asterisk in this table have probably originated in some miscalculation. As hydrogen stands highest in this scale, so it is probable that fluorine, should we ever obtain it in an insulated state, would prove the lowest. The optical properties of tabasheer, in all points of view, are strange anomalies.

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Remarks on
this Table.

It will be observed, that the function $\frac{\mu^2 - 1}{s}$ only expresses the intrinsic refractive power on the hypothesis of the infinite divisibility of matter, and the equal gravitating power of every infinitesimal molecule. But if, as modern Chemistry indicates, material bodies consist of a finite number of atoms, differing in their actual weight for every differently compounded substance, the intrinsic refractive energy of the atoms of any given medium will be the product of the above function by the atomic weight. This will alter totally the order of media from what obtains in the foregoing table. Thus, the weight of the atom of hydrogen being the least, and that of mercury one among the

Light.

greatest in the chemical scale, such multiplication will depress the rank of the former, and exalt that of the latter, so as to separate them entirely from the proximity they now hold. A distinction, too, will require to be regarded between compound and simple atoms. But as these considerations are peculiar to the system of emission, we shall not prosecute them farther in detail.

The dispersive powers of bodies afford another very interesting and distinctive character. Of these, Dr. Brewster, in his *Treatise on New Philosophical Instruments*, has given the following extensive table, almost entirely from his own observation.

TABLE OF DISPERSIVE POWERS.

Column 1 contains the name of the medium; column 2 the value of the function $\frac{\delta \mu}{\mu - 1}$; column 3, that of $\mu \delta$ simply, $\delta \mu$ being the difference of refractive indices of extreme red and violet rays.

Dispersive Powers.	$\frac{\delta \mu}{\mu - 1}$	$\delta \mu$.	Au- thor.	Dispersive Powers.	$\frac{\delta \mu}{\mu - 1}$	$\delta \mu$.	Au- thor.
Chrom. lead, greatest estimated	0.400	0.770	B.	Oil brick	0.046	0.021	B.
Ditto greatest exceeds	0.296	0.570	B.	Flint glass, (Boscov. lowest)	0.0457		B.
Realgar, melted, different kind	0.267	0.394	B.	Nitric acid	0.045	0.019	B.
Chrom. lead, least refraction .	0.262	0.388	B.	Oil lavender	0.045	0.021	B.
Realgar melted	0.255	0.374	B.	Balsam of sulphur	0.045	0.023	B.
Oil cassia	0.139	0.089	B.	Tortoise shell	0.045	0.027	B.
Sulphur after fusion	0.130	0.149	B.	Horn	0.045	0.025	B.
Phosphorus	0.128	0.156	B.	Canada balsam	0.045	0.024	B.
Balsam Tolu	0.103	0.065	B.	Oil marjorum	0.045	0.022	B.
Balsam Peru	0.093	0.058	B.	Gum olibanum	0.045	0.024	B.
Carb. lead, greatest	+0.091	+0.091	B.	Nitrous acid (?)	0.044	0.018	B.
Barbadoes aloes	0.085	0.058	B.	Cajeput oil	0.044	0.021	B.
Oil aniseed	0.074	0.044	B.	Oil hyssop	0.044	0.022	B.
Balsam styrax	0.069	0.039	B.	Oil rhodium	0.044	0.022	B.
Guaiacum	0.066	0.041	B.	Pink coloured glass	0.044	0.025	B.
Carb. lead, least refraction . .	0.066	0.056	B.	Oil savine	0.044	0.021	B.
Oil cummin	0.065	0.033	B.	Oil poppy	0.044	0.020	B.
Gum ammoniac	0.063	0.037	B.	Jargon, greatest refraction . .	0.044	0.045	B.
Oil Barbadoes tar	0.062	0.032	B.	Mur. acid	0.043	0.016	B.
Oil cloves	0.062	0.033	B.	Copal	0.043	0.024	B.
Green glass	0.061	0.037	B.	Nut oil	0.043	0.022	B.
Sulphate lead	0.060	0.056	B.	Burgundy pitch	0.043	0.024	B.
Deep red glass	0.060	0.044	B.	Oil turpentine	0.042	0.020	B.
Oil sassafras	0.060	0.032	B.	Oil rosemary	0.042	0.020	B.
Opal coloured glass	0.060	0.038	B.	Felspar	0.042	0.022	B.
Resin	0.057	0.032	B.	Glue	0.041	0.022	B.
Oil sweet fennel seed	0.055	0.028	B.	Balsam capivi	0.041	0.021	B.
Oil spearmint	0.054	0.026	B.	Oil nutmeg	0.041	0.021	B.
Orange glass	0.053	0.042	B.	Stilbite	0.041	0.021	B.
Rock salt	0.053	0.029	B.	Amber	0.041	0.023	B.
Flint glass, (Boscov. greatest) .	0.0527		Bos.	Oil peppermint	0.040	0.019	B.
Caoutchouc	0.052	0.028	B.	Spinelle	0.040	0.031	B.
Oil pimento	0.052	0.026	B.	Carb. lime, greatest refraction	0.040	0.027	B.
Flint glass	0.052	0.032	B.	Oil rape seed	0.040	0.019	B.
Deep purple glass	0.051	0.031	B.	Bottle glass	0.040	0.023	B.
Oil Angelica	0.051	0.025	B.	Gum elemi	0.039	0.021	B.
Oil thyme	0.050	0.024	B.	Sul. iron	0.039	0.019	B.
Oil fen (? fenu) greek	0.050	0.024	B.	Diamond	0.038	0.056	B.
Oil wormwood	0.049	0.022	B.	Oil olives	0.038	0.018	B.
Oil pennyroyal	0.049	0.024	B.	Gum mastic	0.038	0.022	B.
Oil caraway	0.049	0.024	B.	White of egg	0.037	0.013	B.
Oil dill seed	0.049	0.023	B.	Oil rhue	0.037	0.016	B.
Oil bergamot	0.049	0.023	B.	Gum myrrh	0.037	0.020	B.
Flint glass	0.048	0.029	B.	Beryl	0.037	0.022	B.
Chio turpentine	0.048	0.028	B.	Obsidian	0.037	0.018	B.
Gum thus	0.048	0.028	B.	Ether	0.037	0.012	B.
Flint glass	0.048	0.028	B.	Selenite	0.037	0.020	B.
Oil lemon	0.048	0.023	B.	Alum	0.036	0.017	B.
Oil juniper	0.047	0.022	B.	Oil castor	0.036	0.018	B.
Oil chamomile	0.046	0.021	B.	Sulphur copper	0.036	0.019	B.
Gum juniper	0.046	0.025	B.	Crown glass, very green	0.036	0.020	B.
Carb. strontia, greatest refrac.	0.046	0.032	B.	Gum Arabic	0.036	0.018	B.

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Table of
Dispersive
Powers.

Part IV.

Light.

Part IV.

Dispersive Powers.	$\frac{\delta \mu}{\mu - 1}$	$\delta \mu$.	Au- thor.	Dispersive Powers.	$\frac{\delta \mu}{\mu - 1}$	$\delta \mu$.	Au- thor.
Sugar, cooled from fusion ..	0.036	0.020	B.	Sulphuric acid	0.031	0.014	B.
Jelly fish (<i>medusa æquora</i>) body	0.035	0.013	B.	Apophyllite (leucocyclite)....	0.031	0.017	He.
Water	0.035	0.012	B.	Tartaric acid	0.030	0.016	B.
Aqueous humour, haddock eye	0.035	0.012	B.	Borax	0.030	0.014	B.
Vitreous humour, haddock eye	0.035	0.012	B.	Axinite	0.030	0.022	B.
Citric acid	0.035	0.019	B.	Alcohol	0.029	0.011	B.
Rubellite ..	0.035	0.027	B.	Sulph. barytes	0.029	0.019	B.
Leucite ..	0.035	0.018	B.	Tourmaline.	0.028	0.019	B.
Epidote	0.035	0.024	B.	Crown glass, Leith, (Robi- son,) cited by Brewster....	0.027		Rob.
Common glass, Boscovich's highest, (Brewster)	0.0346		Bos.	Carb. strontia, least refraction	0.027	0.015	B.
Garnet	0.033	0.027	B.	Rock crystal	0.026	0.014	B.
Common glass, Boscovich's lowest, cited by Brewster ..	0.033		Bos.	Emerald	0.026	0.015	B.
Pyrope	0.033	0.026	B.	Carb. lime, least refraction ..	0.026	0.016	B.
Chrysolite	0.033	0.022	B.	Blue sapphire	0.026	0.021	B.
Crown glass	0.033	0.018	B.	Bluish topaz, cairngorm	0.025	0.016	B.
Oil ambergris	0.032	0.012	B.	Chrysoberyl	0.025	0.019	B.
Oil of wine	0.032	0.012	B.	Blue topaz, Aberdeenshire ..	0.024	0.025	B.
Phosphoric acid, solid prism .	0.032	0.017	B.	Sulph. strontia	0.024	0.015	B.
Plate glass	0.032	0.017	B.	Fluor spar	0.022	0.010	B.
				Cryolite	0.022	0.007	B.

Respecting the results in this table, the remark applied to that of refractive indices may be yet more strongly urged. The whole stands in need of a radical reinvestigation. Those only, however, who have had some experience of the difficulties in the way of a strict scientific examination of dispersive powers, can appreciate either the labour of such a task, or the merit of Dr. Brewster in his researches, which we must not be understood as in the slightest degree depreciating by this remark. But the refinements of modern science are every day carrying us beyond all that could be contemplated in its earlier stages, and it is matter of congratulation, rather than disappointment, to every true philosopher, to see his methods replaced by others more powerful, and his results rendered obsolete by the more exact conclusions of his successors. What is now chiefly wanted is a knowledge of the whole series of refractive indices for the several definite rays throughout the spectrum, under uniform circumstances, and for all media whose chemical and other characters are sufficiently definite and constant to enable us to identify and reproduce them in the same state, at all times. The researches of Fraunhofer and Arago have shown that accuracy in the determination of refractive indices sufficient for the purpose, may be attained, and we trust, therefore, that this great desideratum will not long remain unsupplied.

To the substances in the table many important remarks apply. *In general*, high refractive is accompanied by high dispersive power; but exceptions are endless, especially among the precious stones, of which diamond affords a striking instance. Particular bodies seem to carry their dispersive as well as their refractive powers with them into their compounds, and that more *evidently*, because by the peculiar mode in which the dispersion is represented, the state of condensation is eliminated. Thus, fluorine, and even oxygen, appear to exercise a very lowering influence on the dispersive powers of their compounds, while hydrogen, sulphur, and especially lead, act with great energy in the opposite sense. The contrast between the oils of ambergris and cassia, is at least as remarkable in point of dispersive as of refractive power. The following experiment would seem to point out the hydrogen of the latter oil, as the principle to which its extraordinary dispersion is due, and is otherwise instructive, as exemplifying strongly the independence of the two powers *inter se*. A stream of chlorine was passed through oil of cassia till it refused to act any farther. The oil was at first 'greatly deepened in colour, but as the action proceeded, it changed to a much lighter ruddy yellow, which it retained till the action was complete, (and which in a few days changed to a fine rose red.) Copious fumes of muriatic acid gas were given off during the whole process, indicating the abstraction of abundance of hydrogen, and at length the oil was converted into a viscous mass, drawing out into long threads, having entirely lost its peculiar perfume, and acquired a pungent, penetrating scent, and an acrid, astringent taste, totally unlike its former aromatic flavour. It was inflammable, though less than before, burning with a flame green at the edges, indicating the presence of chlorine. Its refractive power was very little diminished. A drop being placed in the angle of two glass plates, and close to it a drop of unaltered oil of cassia, the spectrum of a line of light was viewed at once with the same eye through both the media. They still formed a continuous line, the spectrum of the unaltered oil being more refracted by only about one-fourth the breadth of that of the *altered* specimen. But the dispersive power of the latter was most remarkably diminished, being brought below not only that of the unaltered oil, but even below that of flint glass. When the dispersion of the unaltered oil was corrected by flint glass, that of the altered was found to be much more than corrected; and when the angle of the glass plates was such that the dispersion of the latter was just corrected by a prism of Dollond's "heavy" flint, whose refracting angle = about 25°, the uncorrected spectrum of the former was about equal to that of the flint prism. The dispersion, then, had been diminished to half its former amount, while the refraction had suffered hardly any appreciable change. (October 7, 1825.)

The angle of complete polarization of a ray reflected at the surface of a medium, affords a most valuable character in mineralogy, as it gives at once an approximation to the refractive index, sufficient in a great variety

1121.
Remarks on
the Table of
Dispersive
Powers.

1122.

Experiment
on oil of
cassia.

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Light.

Use of the polarizing angle as a physical character.

Action of crystallized surfaces on reflected light.

1124.

Table of angles between the optic axes of crystals.

of cases to decide between two substances, which might be otherwise confounded together, and inasmuch as it can be measured on any single surface sufficiently polished to give a regular reflexion, thus enabling us to apply this character to minute fragments, or to specimens set as jewels, or otherwise too precious to be sacrificed; to opaque bodies, and to a variety of other cases where a direct measure of the refraction would be impracticable. It has not escaped the acute and careful observation of Dr. Brewster, that the polarizing angle on the surfaces of crystallized media is not absolutely the same in all planes of incidence; and the deviation, though excessively small when the natural reflexion is used, becomes very sensible, and even enormous, when the reflexion is weakened by covering the surface with a cement of a refraction approaching that of the medium, so as to allow only those rays to reach the eye which have penetrated, as it were, to some minute depth, and undergone some part of the action of the crystal *as such*. The point is among the most curious and interesting in the doctrine of reflexion, and we regret that our limits, as well as the obscurity still hanging over it, and which it will require much elaborate research to dissipate, prevent our devoting a section to it, but we must be content to refer the reader to an excellent paper on the subject by that Philosopher, *Philosophical Transactions*, 1819.

The angles included between the optic axes of biaxial crystals is a physical character of the first rank, both on account of its distinctness, its extent of range, (indifferently over the whole quadrant,) and its immediate and intimate connection with the state in which the molecules of the crystals subsist, and what may, loosely speaking, be termed their structure. It is, however, a character by no means easily determined: both axes rarely lying within one field of view, capable of being examined through natural surfaces, and requiring, in almost all cases, the production of artificial sections; at least, this is the only safe way for observations of the tints, for the angles at which, in a thin parallel plate, the several successive orders of colours are produced in situations remote from the axes, are for the most part far too vague to lead to any accurate conclusion as to the position of these lines within the plate, not to speak of the sources of fallacy highly coloured, or dichroite, crystals obviously present. With these considerations before us, we cannot but be struck with surprise and admiration at the unwearied assiduity, which could produce, almost unassisted, a table of results so extensive and so valuable as the following.

Table of the Inclinations of the Optic Axes in various Crystals.

I. UNIAXAL CRYSTALS. Inclination = 0.

Negative Class.

Carbonate of lime, (Iceland spar.)	Corundum.	Idocrase, (Vesuvian.)	Arseniate of potash.
Carbonate of lime and magnesia, (bitter spar.)	Sapphire.	Wernerite.	Muriate of lime.
Carbonate of lime and iron, (brown spar.)	Ruby.	Mica from Kariat.	Muriate of strontia.
Tourmaline.	Emerald.	Phosphate of lead.	Subphosphate of potash.
Rubellite.	Beryl.	Phosphato-arseniate of lead.	Sulphate of nickel and copper.
	Apatite.	Hydrate of strontia.	

Positive Class.

Zircon.	Tungstate of lime.	Apophyllite.	Hydrate of magnesia.
Quartz.	Titanite.	Sulphate of potash and iron.	Ice.
Oxide of iron.	Boracite.	Supracetate of copper and lime.	

Unclassed.

Hyposulphate of lime.

Oxysulphate of iron.

II. BIAxIAL CRYSTALS.

Names of crystals.	Character of the principal axis according to Dr. Brewster's system.	Inclination of optic axes.	Names of crystals.	Character of the principal axis according to Dr. Brewster's system.	Inclination of optic axes.
Sulphate of nickel, certain specimens ..	+	3° 0'	Mica	—	45° 0'
Carbonate of lead	—	5 15	Lepidolite	—	45 0
Sulphato-carbonate of lead	—	—	Benzoate of ammonia	+	45 8
Carbonate of strontia	—	6 56	Sulphate of magnesia and soda	46 49
Carbonate of baryta	—	—	Sulphate of ammonia	+	49 42
Nitrate of potash	—	5 20	Brazilian topaz (Brewster and Biot)	49 r 50°
Mica, certain specimens	—	6 0	Sugar	50° 0'
Talc	—	7 24	Sulphate of strontia	+	50 0
Mother of pearl	—	11 28	Muriosulphate of magnesia and iron ..	—	51 16
Hydrate of baryta	—	13 18	Sulphate of ammonia and magnesia..	+	51 22
Mica, certain specimens	—	14 0	Phosphate of soda	—	55 20
Arragonite	—	18 18	Comptonite... ..	+	56 6
Prussiate of potash (? Ferrocyanate) ..	+	19 24	Sulphate of lime	+	60 0
Mica, certain specimens	25 0	Oxynitrate of silver	+	62 16
Cymophane	+	27 51	Iolite	—	62 50
Anhydrite	+	28 7	Felspar	—	63 0
Borax	+	28 42	Topaz (Aberdeenshire)	+	65 0
	..	30 0	Sulphate of potash	+	67 0
	..	31 0	Carbonate of soda	—	70 1
Mica, various specimens examined by	..	32 0	Acetate of lead	—	70 25
M. Biot	34 0	Citric acid	+	70 29
	..	37 0	Tartrate of potash	—	71 20
Biaxial apophyllite	—	35 8	Tartaric acid	—	79 0
Sulphate of magnesia	—	37 24	Tartrate of potash and soda	+	80 0
Sulphate of barytes	+	37 42	Carbonate of potash	80 30
Spermaceti (about)	+	37 40	Cyanite	+	81 48
Tineal (native borax)	—	38 48	Chlorate of potash	82 0
Nitrate of zinc (estimated)	40 0	Epidote, about	84 19
Stilbite	+	41 42	Muriate of copper	84 30
Sulphate of nickel	+	42 4	Peridot	87 56
Carbonate of ammonia	—	43 24	Crystallized Cheltenham salts	88 14
Sulphate of zinc	—	44 28	Succinic acid, estimated at about	90 0

Light.

Among crystals with one axis, Dr. Brewster has enumerated the Idocrase, or Vesuvian, and correctly. Had he noticed, however, in the specimens examined by him the very striking inversion of the tints of Newton's scale exhibited in the rings of that now before us, he would doubtless have made mention of it. We insert here the scale of colours exhibited by a plate cut from the specimen in question, (a fine large crystal,) as affording another remarkable case in addition to that of the hyposulphate of lime, and the several varieties of uniaxial apophyllite already mentioned, of such inversion.

Part IV.
1125.
Remarks.
Inverted
tints of
Vesuvian.

Table of the tints exhibited by a plate of Vesuvian, thickness = 0.11035 inch, cut a little obliquely to a perpendicular to the axis.

Angle of Incidence.	Ordinary Image.	Extraordinary Image.	$n =$	Angle of Refraction ϵ .
+ 66° +'	No light passed	No light passed.		
+ 66 0	Brick red	Dull pale green.		
+ 64 0	Orange red	Fine blue green.		
+ 60 0	Tolerable orange pink	Fine bluish green.		
+ 52 0	Pale yellow pink	Pale yellowish green.		
+ 47 0	Pink, with a dash of purple. .	Pretty bright yellow.		
+ 42 0	Pale neutral purple	Good yellow	$\frac{1}{2}$	- 25° 56'
+ 37 0	Bluish white	Yellow less bright.		
+ 30 0	Very pale yellowish white. . .	Sombre brownish yellow.		
+ 15 0	Yellowish white.	Very sombre yellow brown . .	} 0	- 6 31 ±
+ 10 0	Yellowish white.	Almost totally extinct		
+ 3 0	Yellowish white.	Very sombre purplish brown		
± 0 0	Yellowish white.	Dusky brownish yellow.		
- 9 0	Bluish white	Rather dull yellow.		
- 12 0	Dull purplish blue	Bright yellow	$\frac{1}{2}$	+ 7 48
- 16 0	Ruddy purple	Pale yellow.		
- 19 0	Pink, verging to brick red . .	Imperfect green.		
- 22 0	Yellowish red	Tolerable bluish green.		
- 26 0	Yellow, inclining to orange . .	Rich greenish blue.		
- 28 0	Bright yellow	Blue purple.		
- 28 30	Bright yellow	Neutral purple	1	+ 18 10
- 29 0	Bright yellow	Ruddy purple.		
- 30 0	Yellow green.	Crimson.		
- 32 0	Good green	Good pink.		
- 35 0	Greenish blue	Orange pink.		
- 37 30	Blue purple	Pale yellow.		
- 38 30	Neutral purple	Pale yellow	$\frac{1}{2}$	+ 24 0
- 39 15	Ruddy purple	Greenish yellow.		
- 41 30	Good pink.	Good green.		
- 45 0	Pink yellow	Fine greenish blue.		
- 47 20	Yellowish white	Blue purple.		
- 47 30	Yellowish white	Neutral purple	2	+ 28 48
- 48 0	Very pale green.	Ruddy purple.		
- 49 30	Fine green.	Good pink.		
- 53 0	Fine blue green.	Orange pink.		
- 54 0	Greenish blue	Yellow.		
- 54 -	No light passed.	No light passed.		

The first ring, it will be observed, in calculating from this table, is contracted beyond what is due to the law of the sines, probably from the section examined not passing precisely over their common centre, and gives a polarizing power greater than that deduced from the angles corresponding to $n = 1$, $n = \frac{3}{2}$, $n = 2$, all which agree in assigning 41.35 nearly as the measure of the power in question. See Art. 1126.

It follows from this series, that of the two images formed by double refraction in Vesuvian, and other similar crystals, the *most refracted* should be the *least dispersed*, a peculiarity we have not yet had an opportunity of verifying by direct observation. It follows, however, immediately from the theory of the rings above delivered, since the smaller the diameters of the rings for any coloured ray, the greater the separation of its pencils by double refraction. Hence, in the present case, the red rays will be separated by a greater interval than the violet in the two spectra; and, consequently, the least refracted spectrum will be the longest. In the variety of apophyllite exhibiting white and black rings, (*leucocyclite*) the two dispersions should be almost exactly equal, and the only difference between the two spectra ought to consist in a slight variation in the proportional breadths of the several coloured spaces in them.

Another very important optical character is the *intensity* of the polarizing, or doubly refractive energy. This may be concluded by measuring the actual angular separation of the images; but this is usually too small to

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Light.
Polarizing
powers con-
sidered as a
physical
character
of media.

admit of being determined with sufficient precision, in such very imperfect specimens as are usually subjected to examination for the purpose of identification, and a much better course is to make the tint developed at a perpendicular incidence, by a plate of given thickness in a direction at right angles to both the optic axes, the object of determination. This tint (which we shall term the equatorial tint) may be derived immediately from observations of tints at any angle, by the formula

$$N = \frac{n}{t} \cdot \frac{\cos \rho}{\sin \theta \cdot \sin \theta'},$$

where N is the tint in question, numerically expressed as usual, and where n is the tint, (also similarly expressed) developed at an angle of incidence whose corresponding angle of refraction is ρ , on a plate whose thickness is t , (expressed in English inches and decimals) and where θ, θ' are the angles made by the ray in traversing the plate with the two axes. This value of N is the same with $\frac{1}{h}$ in the equation of Art. 907. The following list of a very few substances will suffice to show the great range the value of N admits, and its consequent utility as a physical character, considerations which we hope will induce observers to extend the list itself, as well as to give it all possible exactness.

UNIAXIAL CRYSTALS.

	For mean yellow rays.	
	$N =$	$h = \frac{1}{N}$
Iceland spar	35801	0.000028
Hydrate of strontia (assuming $\mu = 1.5$)	1246	0.000802
Tourmaline	851	0.001175
Hyposulphate of lime	470	0.002129
Quartz	312	0.003024
Leucocyclite (uniaxial apophyllite, 1st variety)	109	0.009150
Camphor	101	0.009856
Vesuvian	41	0.024170
Uniaxial apophyllite, 2d variety	33	0.030374
Ditto, 3d variety	3	0.366620

BIAXIAL CRYSTALS.

	For mean yellow rays.	
	$N =$	$h = \frac{1}{N}$
Nitre	7400	0.000135
Anhydrite (angle between axes $= 43^\circ 48'$)	1900	0.000526
Mica (angle between axes $= 45^\circ$)	1307	0.000765
Sulphate of baryta	521	0.001920
Heulandite (white;—angle between axes $= 54^\circ 17'$)	249	0.004021

1127. But the phenomena of refraction, reflexion, and polarization, may not only be applied by the aid of these and similar tables of registered results, to the examination and identification of substances in the gross, they are also of use in detecting peculiarities of structure in individual specimens, or in certain species which would otherwise escape observation. The singular structure of amethyst has been already explained, and a variety of cases of hemitropism might be noticed, in which the juxtaposition of the parts is rendered evident by the test of polarized light. Of these, however, by far the most curious and interesting are those in which the juxtaposed parts combine to form a regular whole, and to produce a species of pseudo-crystal, built up as it were of several individuals, arranged with a regard to symmetry, and forming a structure of more or less complication. Such instances have been more particularly noticed in nitre, arragonite, topaz, apophyllite, sulphate of potash, analcime, harmotome, &c.

1128. The usual form of the crystals of nitre, when large and well developed, is the regular hexagonal prism; but a section of this, cut at right angles to the axis, is very commonly found to consist of two or more portions, in which the optic meridians are 60° inclined to each other; but the plane of division often intersects one of the lateral faces of the prism, without any visible external mark of a breach of continuity, so that but for the test of polarized light, the macled structure would never be discerned. The phenomena of arragonite, in this respect, are very similar to those of nitre.

1129. If a plate of Brazilian topaz, cut at right angles to the axis of the rhombic prism in which it crystallizes, be examined by polarized light, it will occasionally be found to consist of a central rhomb, surrounded by a border in which the optic meridians of the alternate sides are inclined at $\frac{1}{4}$ of a right angle to that of the central compartment, and $\frac{1}{2}$ a right angle to each other. In consequence, when such a rhombic plate is held with its long diagonal in the plane of primitive polarization, two opposite sides of the border appear bright, the other two black, and the central compartment of intermediate brightness. Such specimens often exhibit the phenomena of dichroism in the central compartment, while the border is colourless in all positions.

1130. But it is in the apophyllite of the variety named by Dr. Brewster, Tesselite, that this enclosure of one crystal in a case as it were of another, is exhibited in the most regular and extraordinary manner. In one of the varieties of this singular body, whose form is the right rectangular prism with flat summits, slices taken off from either summit were found by him to be of uniform structure; but, when these were detached, every subsequent slice was

Use of polarized light in detecting complex structures.

Compound crystals of nitre.

Arragonite.

Topaz.

Tesselite.
Fig. 223.

Light.

found to consist of a rectangular border enclosing no less than nine several compartments, arranged as in fig. 223, and separated from each other, and from the border, by delicate lines or films as there marked. Each of these compartments possesses its own peculiar crystallographic structure, and polarizes its peculiar tints, the law of symmetry being observed. In some specimens the triangular spaces *p q r s* were wanting, while in others they seem to have consisted of two portions, separated by an imaginary prolongation of the line joining their obtuse angles with the central lozenge.

The terminal plates, the central lozenge, and the minute stripes dividing the compartments from each other (which are sections of laminae or films parallel to the axis of the crystal, and running its whole length) consist of that uniaxal variety, in speaking of which we have used the term leucocyclite, from the whiteness of its rings. The rectangles R V, S T, (with the exception of the portions occupied by the lozenge and partitions) consist of a *biaxal* medium, having its axes 34° inclined to each other, and its optic meridian parallel to the axis of the prism, and passing through the diagonals R V, S T of these rectangles. The other rectangles are composed of a similar medium, but with its optic meridian at right angles to the former, or passing through the diagonals R T, S V.

A still more remarkable and artificial structure has been observed by Dr. Brewster, in a variety of the Faroe apophyllites of a greenish white hue. When a complete prism of this variety is exposed to polarized light, with its axis in 45° of azimuth, the light being transmitted perpendicularly through two opposite sides, the pattern represented in fig. 224 is seen, in which the central curvilinear area is red, and its complements to the surrounding rectangle green. The squares immediately adjacent on either side in the direction of the axis are also vivid red in their centres, fading into white, while the rest of the pattern consists in a most brilliant succession of red, green, and yellow, bands, for a coloured figure of which we must refer the reader to the original most curious and interesting memoir, (*Edinburgh Transactions*, vol. ix. part ii.) where, as also in the *Edinburgh Philosophical Journal*, vol. i. he will find the phenomena described in full detail.

The sulphate of potash offers another very remarkable example of compound structure. This salt occurs in hexagonal prisms, and occasionally in bipyramidal dodecahedrons. But besides these forms it also occurs in rhombic prisms of 114° and 66° . These Dr. Brewster found to have two axes, while the hexagonal prisms have but one; thus affording another instance of dimorphism in addition to those of arragonite, sulphur, &c. On examining the dodecahedrons, however, he found them to consist of six equilateral triangular prisms, of the biaxal variety, grouped together, and having their optic meridians all converging to the common axis; the molecules being so disposed in each opposite pair of individuals as to make the angle between the opposite faces of either pyramid (114°) equal to the obtuse angle of the rhomboid.

The structure and mode of action of the analcime, described by Dr. Brewster in vol. x. of the *Edinburgh Transactions*, part i. p. 187, are so extremely singular, that it is difficult to say whether it should be regarded as a grouped crystal, consisting of independent portions adhering together, or as a mass the distribution of the ether in whose parts is governed by a general and uniform law; the latter, however, is probably the truth. The form of this crystal is the icositetrahedron, contained by twenty-four similar and equal trapezia, and may be regarded as derived from the cube by the truncation of each of its angles by three planes symmetrically related to the edges including it. If we conceive from the centre of this cube, (in its natural situation with respect to the derived figure) planes to pass through each of the edges, and through each of the diagonals of the six faces, they will divide the cube into twenty-four irregular tetrahedra; and of these, all the faces which pass through edges of the cube will also pass through edges of the derived figure, while those which pass through diagonals of faces of the cube will also pass through diagonals of the faces of its derivative, bisecting their obtuse angles. Now it appears from Dr. Brewster's observations, that all the molecules situated in any part of any one of these planes are devoid of the power of double refraction and polarization; and that in proportion as a molecule is distant from all such planes, its polarizing power is greater. In this respect it differs entirely from all crystals hitherto examined, every particle of which, wherever situated, so long as they belong to one and the same crystalline system, being equally endued with the polarizing virtue. Nor is there a closer analogy between the mode of action in question, and that of unannealed glass and similar bodies; for in these a change of external form is always accompanied with a change of the polarizing powers, while in the analcime each particular portion, whether detached from the mass, or in its natural connection with the adjacent molecules, possess the very same optical properties. The action too of the portions which possess a polarizing power is not related to axes given only in direction, and passing through every molecule, but to planes given both in direction and in place within the mass, (the planes above mentioned;) the tint developed at any point of a plate being as the square of the distance from the nearest of such planes, and the isochromatic lines being, in consequence, straight fringes of colour arranged parallel to the dark bands marked out by the intersection of such planes with the plate examined. The phenomena described are accompanied with a sensible double refraction. The reader is referred to the memoir already cited (which is one of the most interesting to which we can direct his attention) for further details: and to a work understood to be forthcoming from the pen of the eminent author here and so often before cited, on optical mineralogy, for what we are sure will prove a treasure of valuable information on every point connected with this important application of optical science.

§ XIV. On the Colours of Natural Bodies.

It was our intention to have devoted a considerable share of these pages to the explanation of such natural phenomena as depend on optical principles, but the great length to which this essay has already extended, renders it necessary to confine what we have to say on such subjects within very narrow limits, and to points of promi-

Part IV

1131.

Another variety.

Fig. 224.

1132.

Sulphate of potash.

1133.

Analcime.

1134.

- Light.** **Newton's theory of the colours of natural bodies.** **1135.** **Postulates.** **1136.** **1137.** **1138.** **Cause of opacity.** **1139.** **Origin of natural colours.** **1140.** **Objections.** **1141.** **Apparent exceptions considered.** **1142.** **Case of transparent coloured media.**
- ment importance. Among these there is certainly none more entitled to consideration than the phenomena of colour, as exhibited by natural objects, which strike us wherever we turn our eyes, and it is impossible to pass in total silence the theory devised by Newton to account for them; a theory of extraordinary boldness and subtilty, in which great difficulties are eluded by elegant refinements, and the appeal to our ignorance on some points is so dexterously backed by the weight of our knowledge on others, as to silence, if not refute, objections which at first sight appear conclusive against it. The postulates on which this theory rests are essentially as follows:
1. All bodies are porous; the pores or intervals vacant of ponderable matter, occupying a very much larger portion of the whole space filled by the body, than the solid particles of which it essentially consists.
 2. These solid particles have a certain size (and perhaps figure) essential to them as particles of that particular medium, and which cannot be changed by any mechanical action, or by any means not involving a change in the chemical nature or condition of the medium. They are, in short, the ultimate *atoms*; to break which, is to destroy their essence, and resolve them into other forms of matter, having other properties.
 3. These atoms are perfectly transparent, and equally permeable to light of all refrangibilities, which, having once passed their surfaces, is in the act of pursuing its course through their substances. Newton, indeed, makes his atoms only "in some measure transparent." But he never refers to this limitation, and his theory depends essentially on their *perfect* transparency, as is indeed obvious from his account of opacity, which is contained in the next postulate.
 4. *Opacity in natural bodies arises from the multitude of reflexions caused in their internal parts.* It is obvious, therefore, that unless we admit a cause of opacity in *atoms* different from that which, on this hypothesis, causes it in their aggregates constituting natural bodies, the former cannot be otherwise than absolutely pellucid, since no reflexions can take place where there are no intervals, and no change of medium. Of the sufficiency of this cause, either in natural bodies or atoms, however, we confess there does appear to us some room for doubt, as it seems difficult so to conceive these internal reflexions, that the rays subjected to them shall be *all* and *for ever* retained, entangled as it were, and running their rounds from atom to atom, without a possibility of reaching the surface and escaping; which, were they to do, it is evident that every body so constituted, receiving a beam of light, would in fact only disperse it in all directions in the manner of a self luminous one.
 5. *The colours of natural bodies are the colours of thin plates, produced by the same cause which produces them in thin laminae of air, glass, &c.* viz. the interval between the anterior and posterior surfaces of the atoms, which, when an odd multiple of half the length of a fit of easy reflexion and transmission for any coloured ray moving within the medium, obstructs its penetration of the second surface, and when an even, ensures it, (see Art. 655.) The *thickness*, therefore, of the atoms of a medium, and of the interstices between them, determines the colour they shall reflect and transmit at a perpendicular incidence. Thus, if the molecules and interstices be less in size than the interval at which total transmission takes place, or less than that which corresponds to the edge of the central black spot in the reflected rings, a medium made up of such atoms and interstices will be perfectly transparent. If greater, it will reflect the colour corresponding to its thickness.
- It may be objected to this, that all natural colours do not of necessity find a place in the scale of tints of thin plates, even those of bodies whose chemical composition is uniform; but to this we may answer, that the colours reflected from the first layer only of molecules next the surface ought to be *pure* tints, those from lower layers having to make their way to the eye through the upper strata, and thus undergoing other analyses, by transmissions and reflexions among the incumbent atoms. Besides which, whatever *shape* we attribute to the atoms, it is impossible that all rays shall penetrate them so as to traverse the same thickness of them, unless we regard them as mere *laminae* without angles or edges, and of enormous refractive power.* The same answer must be made to the objection, equally obvious, that the transmitted tint ought to be in all cases complementary to the reflected one, and that therefore cases like that of leaf gold, opalescent glass, and infusion of *lignum nephriticum*, all which reflect one tint and transmit another, but in all which this condition is violated, form exceptions to the theory. But, in reality, the transmitted rays have traversed the whole thickness of the medium, and have therefore undergone, many more times, the action of its atoms, than those reflected, especially those near the first surface, to which the brighter part of the reflected colour is due.
- The infusion of *lignum nephriticum* is a very singular case, and its peculiar properties have been explained by Dr. Young, on the supposition of minute particles of definite magnitude suspended in it. Though very transparent, it yet *reflects* a bluish green colour, while the light transmitted is yellow or wine-coloured, in this respect offering almost the exact converse of leaf gold. It is, however, no doubt a case of opalescence, and is exactly imitated by certain yellow glasses, in which a very visible thin film of opalescent matter near the surface reflects to the eye a bluish green tint, while yet the colour transmitted has the yellow tint belonging to the glass. The reflexion proceeds from particles which have nothing to do with the transmitted light.
- But, in fact, the objection (as appears to us) is not yet fully answered. Transparent coloured media (*clear* liquids in which no floating particles exist,) have *no* reflected colour. When examined by pouring them into an opaque vessel, blackened internally and filled to the brim, and when the colourless reflexion from their upper surface is destroyed by reflexion in an opposite plane at the polarizing angle, it is seen at once that no light is reflected from within the medium, either near the surface, or at greater depths; and if this mode of examination be regarded as objectionable, as perhaps destroying the internal as well as external reflexion, it is equally satisfactory to observe, that the image of a white object reflected from the surface of a fluid in a black opaque vessel is always purely white, whatever be the colour of the reflecting fluid. We are not aware that the objection so put has been sufficiently considered, or even propounded. To us its weight appears considerable,

* Newton appears to have been fully aware of the necessity of taking this into consideration. Prop. vii. book ii. *Opt. versus finem.*

Light.

and we cannot but believe that some other cause besides mere internal reflexions must interfere to prevent the complementary colour from reaching the eye; and that absorption, with its kindred phenomenon, or rather its extreme case, opacity, is not satisfactorily accounted for in this theory, but must rather be admitted as (at present,) an ultimate fact, of which the cause is yet to seek.

If this be granted, the colours of all bodies may be distinguished into *true*, viz., those which arise from rays which have actually entered their substance and undergone their absorptive action, (as the colours of powders of transparent coloured media, cinnabar, red lead, Prussian blue, those of flowers, &c.) and *false*, or *superficial*, or those which originate obviously in the law of interference; thus, the variable colours of feathers, insects' wings, striated surfaces, oxidated steel, and a variety of cases to which the Newtonian doctrine strictly applies, for there is no denying that cases of colour, not *merely* superficial, do occur, in which the Newtonian doctrine, to say the least, is highly probable. To instance one or two only. If a few drops of an extremely weak solution of nitrate of silver be added to a very dilute solution of hyposulphite of lime, a precipitate is formed of an opalescent whiteness and extreme tenuity. If more of the nitrate be added, the precipitate increases in weight and aggregation, and at the same time changes its colour, becoming first yellow, then yellow brown, then a rich orange brown, then a purplish brown, and, finally, a deep brown black. The precipitate, meanwhile, continually acquires density, and, finally, sinks rapidly to the bottom. It is impossible, in this series, not to trace the tints of the first order of reflected rings, produced by the thickening of the minute particles in the act of aggregation, but equally impossible not to recognise the agency of a cause totally different, acting to increase the opacity of the compound by an absorptive action far superior to, and independent of, the action of the particles as thin plates. The phenomena of *Hematine*, described by Chevreul and cited by Dr. Brewster, (*Encyc. Edin. Optics*, p. 623; see also Biot, *Traité de Phys.* tom iv. p. 134, there referred to,) afford too close an approximation to this case also. The diffused light and blue colour of the clear sky, affords another very satisfactory instance. This blue is, no doubt, a blue of the first order, reflected from minute aqueous particles in the air. The proof is, that at 74° distance from the sun, it is *completely polarized* in a plane passing through the sun's centre.

Cases in which Newton's theory applies.

Another objection, no less obvious, to the Newtonian doctrine, has been successfully answered by Newton himself. A change of obliquity of incidence, it may be urged, should cause a change of colour, as a plate of given thickness reflects a different tint at oblique and perpendicular incidences. But this variation is less, the greater the refractive power of the medium; and as the refractive power increases with the density, that of the dense ultimate atoms of bodies must be exceeding great, so that the tint reflected from them will vary little with a change of incidence, (art. 669.) The colours of oxidated steel afford an excellent case in point. The refractive power of this oxide, though great, (2.1), is, doubtless, not to be compared with that of the ultimate atoms of bodies, yet the tints on the surface of blued steel vary but little with a change of obliquity. We may add, too, that the colour exhibited by any body of sensible magnitude, is in reality an average of the colours reflected from all its molecules at all possible incidences, so that no change of incidence ought to be expected to affect it.

1144. Another objection. Answered.

Of the extreme tenuity of the ultimate molecules of bodies, Newton seems to have had but an inadequate idea, as he supposed that they might be seen through microscopes magnifying three or four thousand times.* We have viewed an object *without utter indistinctness*, through a microscope by Amici, magnifying upwards of three thousand times in linear measure, and had no suspicion that the object seen was even approaching to resolution into its primitive molecules. But it should rather seem that Newton regarded his colorific molecules as divisible *groupes* of atoms of a yet more delicate kind, and yet more dense, and these again as still further resolvable till the last stage of indivisibility be reached. M. Biot has given a striking, and, we may almost term it, picturesque account of this doctrine, in his *Traité de Physique*.

1145. Newton's ideas of the size of the particles of bodies.

§ XV. Of the Calorific and Chemical Rays of the Solar Spectrum.

It has long been a matter of everyday observation, that solar light exercises a peculiar influence in altering the colours of bodies exposed to it, either by deepening or discharging them, even when totally secluded from air, and that various metallic salts and oxides, especially those of silver, are speedily blackened and reduced when freely exposed to direct sunshine, or even to the ordinary light of a bright day. Whether these effects were owing to the heat of the rays, or to some other cause, remained long uninquied. The first step was

1146.

* The passage, however, is in the highest tone of a refined philosophy, and, independent of its theoretic bearings, we extract it, as indicating a scrutinizing spirit of observation far beyond the age he lived in.

"In these descriptions I have been the more particular, because it is not impossible but that microscopes may at length be improved to the discovery of the particles of bodies on which their colours depend, if they are not already in some measure arrived to that degree of perfection. For if those instruments are or can be so far improved as with sufficient distinctness to represent objects five or six hundred times bigger than at a foot distance they appear to our naked eyes, I should hope that we might be able to discover some of the greatest of those corpuscles. And by one that would magnify three or four thousand times perhaps they might all be discovered, but those which produce blackness. In the mean while I see nothing material in this discourse that may rationally be doubted of, excepting this position: That transparent corpuscles of the same thickness and density with a plate, do exhibit the same colour. And this I would have understood not without some latitude, as well because those corpuscles may be of irregular figures, and many rays must be obliquely incident on them, and so have a shorter way through them than the length of their diameters, as because the straitness of the medium put in on all sides within such corpuscles may a little alter its motions or other qualities on which the reflection depends. But yet I cannot much suspect the last, because I have observed of some small plates of Muscovy glass which were of an even thickness, that through a microscope they have appeared of the same colour at their edges and corners where the included medium was terminated, which they appeared of in other places. However, it will add much to our satisfaction, if those corpuscles can be discovered with microscopes; which if we shall at length attain to, I fear it will be the utmost improvement of this sense. For it seems impossible to see the more secret and noble works of nature within the corpuscles, by reason of their transparency."

Light. made by Scheele, who ascertained that muriate of silver is much more powerfully blackened in the violet rays than in any other part of the spectrum. (*Traité de l'Air et du Feu*, § 66.) The experiments of Sir W. Herschel, on the heating power of the several prismatic rays, on the other hand; which appeared in 1800, showed satisfactorily that the more refrangible rays possess very little heating power, the calorific effect being at its maximum for the extreme red rays, and even extending considerably beyond the limits of the spectrum in that direction. This remarkable discovery, which established the independence of the heating and illuminating effects of the solar rays, led Professor Ritter, of Jena, in 1801, to examine whether a similar extension beyond the limits of the visible spectrum might not also have place in the chemical or deoxidating rays, and on exposing muriate of silver in various points within and without the spectrum, he found the maximum of effect to lie beyond the visible violet rays, the action being less in the violet itself, still less in the blue, and diminishing with great rapidity as he proceeded towards the less refrangible end. Dr. Wollaston independently arrived at the same conclusion.

1147. The solar rays, then, possess at least three distinct powers: those of heating, illuminating, and effecting chemical combinations or decompositions, and these powers are distributed among the differently refrangible rays, in such a manner as to show their complete independence on each other. Later experiments have gone a certain way to add another power to the list—that of exciting magnetism. Without calling in question the accuracy of the observations which are directed to establish this point, we may be permitted to hope that further researches will, ere long, explain the causes of failure in those numerous cases where such effects have not been produced.

1148. The calorific rays appear, from experiments of Berard, to obey the laws of polarization and double refraction, like those of light. Those of interference could not be made without excessive difficulty. In the case of the chemical rays, the same difficulty is not experienced; and Dr. Young, and after him, by more delicate means, M. Arago, have satisfactorily demonstrated that these conform to the same laws of interference, whether polarized or otherwise, that are obeyed by the luminous rays similarly circumstanced. Thus, a set of fringes formed by the interference of two solar pencils with a common origin, being kept very steadily projected for a long time on one and the same part of a sheet of paper rubbed with muriate of silver, a series of black lines became traced on it, the intervals of which were smaller than those of the dark and luminous fringes formed by homogeneous violet light.

1149. Dr. Wollaston having observed that gum guaiacum is turned green by exposure to solar light in contact with air, took two specimens of paper coloured with a yellow solution of this gum in alcohol, and exposed one of them to air and sunshine, the other to air in the dark. The former was turned perceptibly green in five minutes, and the change was complete in a few hours, while the latter was no way discoloured after many months. He then concentrated the violet rays on paper so coloured, by a lens, and the change was speedily performed, while in the most luminous there was no change of colour, and, in the red rays, the green colour was not only produced, but when induced by exposure to the violet, was again destroyed, and the original yellow colour restored. This seems, however, to have been merely an effect of the heat, as the warmth from the back of a heated silver spoon discharged the green colour just as effectually.

1150. Mr. Faraday has observed that glass tinged purple with manganese, has its hue much deepened by the passage of solar light through it, and that two portions of the same plate, one preserved in the dark, the other exposed freely, after some time differ materially in intensity of colour.

The direct action of solar light, or, possibly, of its heat also, produces other chemical effects, such as the immediate combination of the elements of phosgen, the explosion of an atomic mixture of chlorine and hydrogen, and other phenomena, all indicative of powers resident in this wonderful agent, of which we have but a very imperfect notion at present. The green colour of plants, and the brilliant hues of flowers, depend entirely on it. Tansies which had grown in a coal pit, were found totally destitute either of colour or of their peculiar and powerful flavour, and the bleaching and sweetening of celery by the exclusion of light, is another familiar instance of the same cause. How far the differently coloured rays are concerned in these effects, has never yet been *accurately* investigated, though attempts have been made; but we hope, from the distinguished ability of an eminent individual who has recently taken up this most interesting inquiry, that our stock of knowledge will soon receive material accessions.

1152. We cannot close this Essay without an expression of regret, that the Memoir of Professor Airey, on the Spherical Aberration of the Eyepieces of Telescopes, just on the point of publication in the *Transactions of the Cambridge Philosophical Society*, reached us too late to allow of our attempting to condense its valuable contents, and we can only recommend it to the notice of our readers in lieu of, and in preference to, anything we could ourselves say on that subject. A similar expression of regret applies to the interesting "Theory of Systems of Rays," by Professor Hamilton of Dublin, a powerful and elegant piece of analysis, communicated to the Royal Irish Academy in 1824, and only now in the course of impression, but of which enough has reached us, by the kindness of its Author, to make us fully sensible of the benefit we might have derived from its perusal at an earlier period of our undertaking.

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ERRATA ET CORRIGENDA.

N B. The reader is requested to correct in advance the following Errata, and to strike out the passages here referred to

Page.	Line.	Error.	Correction.
311,	39,	existences,	existence.
do.	31,	more,	most.
317,	2,	line,	sine.
319,	26,	as the sun's surface,	at the sun's surface
369,	5 from bott.	$-\frac{y^2}{2\mu}$ {	$-\frac{y^2}{2\mu I_1}$ {
399,	47,	axis,	axes.
400,	15,	act,	art.
401,	15,	P E,	P C.
402,	30,	R, R O p.	P, P O p.
do.	35,	p,	Q.
410,	22 from bott.	dele " see <i>Micrometer</i> , in a subsequent part of this Article."	
414,	44,	dele " by the writer of these pages."	
415,	3 from bott.	by water,	into water.
420,	30,	spectra of distortion,	subordinate spectra.
428,	17,	secondary,	second.
431,	33,	R r, V v,	R R', V V'.
do.	36,	R N V,	R' N V'.
431,	27,	from experiments,	from other experiments.
454,	28,	P B, P Q, or A B,	P B - P Q, or A B.
461,	11 from bott.	two vibrations,	two rectangular vibrations.
476,	14 from bott.	dele all that relates to the fringes on the wings of the <i>Papilio Idas</i> , being founded on a mistake.	
480,	32,	limits,	limit.
509,	38,	falls,	falls.
521,	1,	produce,	to produce.
524,	margin, Art. 925.	polarization,	depolarization.
528,	22,	positive class,	attractive class.
531,	16, add as follows:—	With respect to this third law, however, it must be confessed that it appears to require a stricter examination, as, if admitted in its full extent, it seems to controvert the fundamental principles of the doctrine of interference	
564,	19,	dele what is said about the nodal lines.	
566,	32 from bott.	after disprove it, insert as follows: Instead of the expression (b.) Art. 1073, we might otherwise assume	

$$T = (Y \cdot \cos 2\phi^2 + B \cdot \sin 2\phi') \cdot (y \cdot \cos v^2 + b \cdot \sin v^2),$$

and determining the coefficients accordingly, obtain another expression for the tint.

Plate I

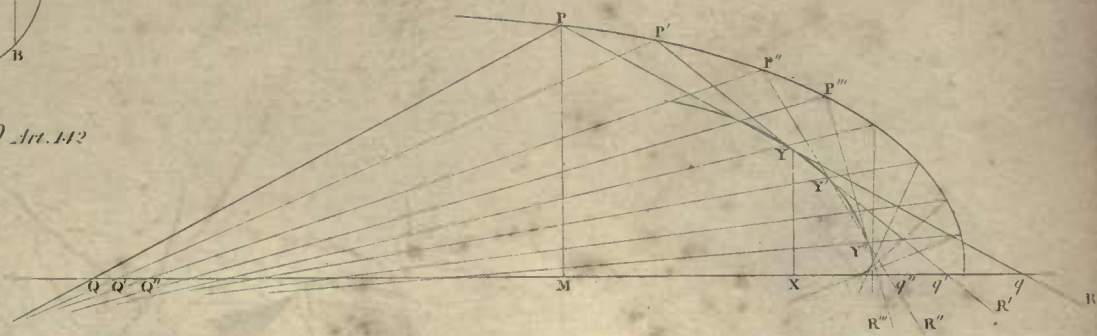
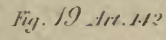
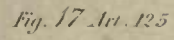
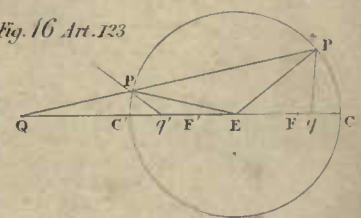
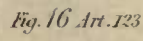
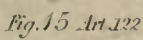
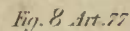
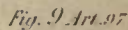
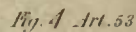


Fig. 20
Art. 145

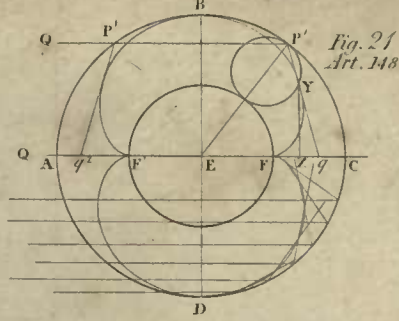
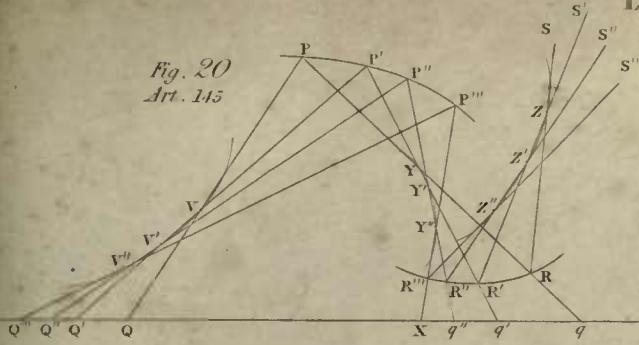


Fig. 21
Art. 148

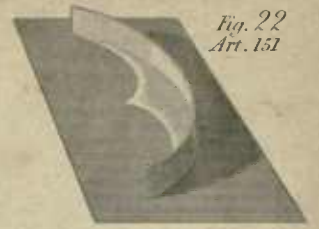


Fig. 22
Art. 151

Fig. 23
Art. 154

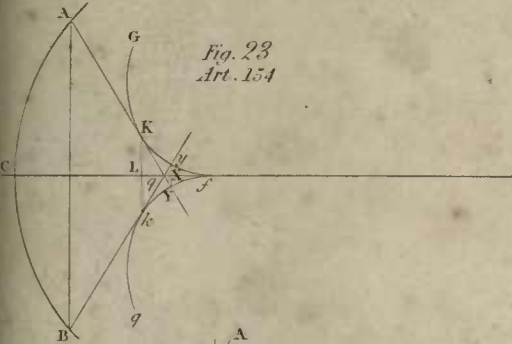


Fig. 24
Art. 160

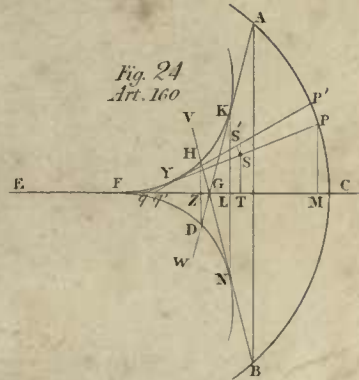


Fig. 25
Art. 163

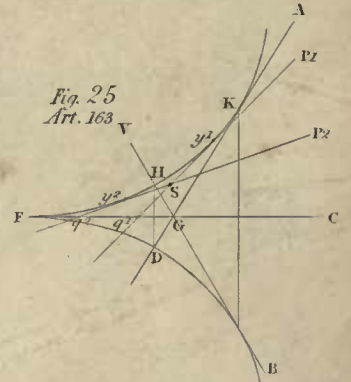


Fig. 26
Art. 164

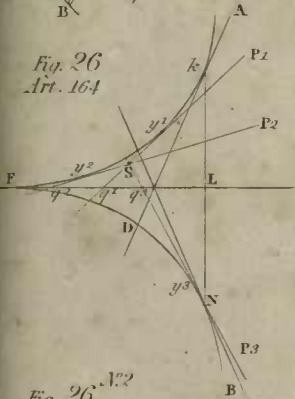


Fig. 26
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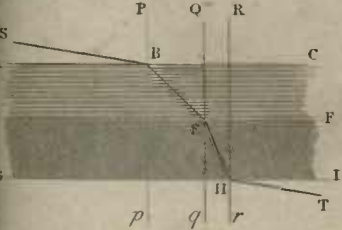


Fig. 23
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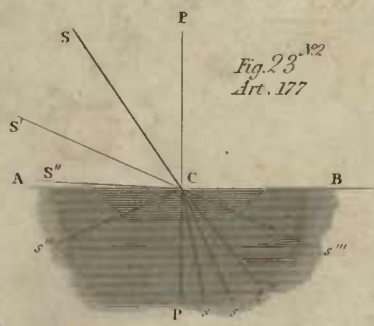


Fig. 27
Art. 198

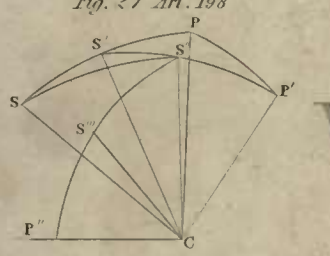


Fig. 24
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Fig. 25
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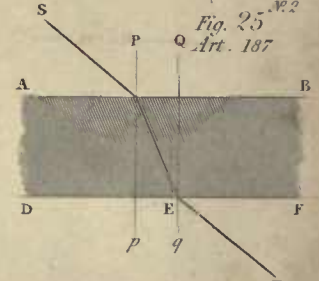


Fig. 29
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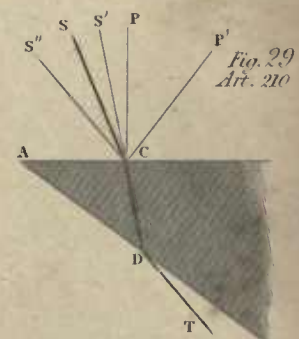


Fig. 30
Art. 210

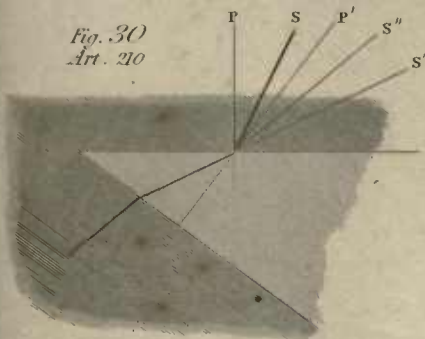


Fig. 31
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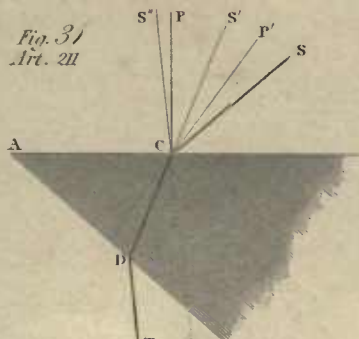


Fig. 32
Art. 211

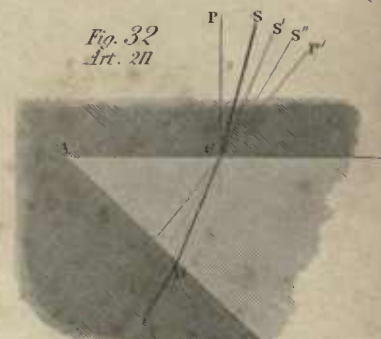


Fig. 33. Art. 221.

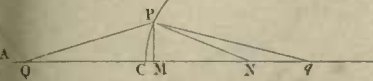


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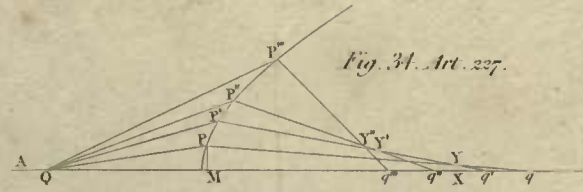


Fig. 36. Art. 233 & 303.

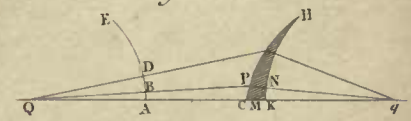


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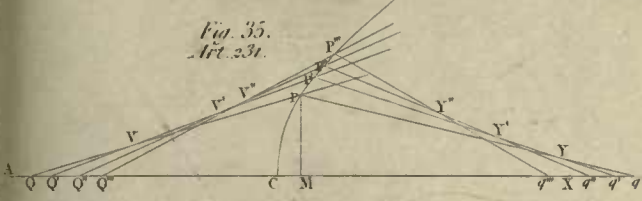


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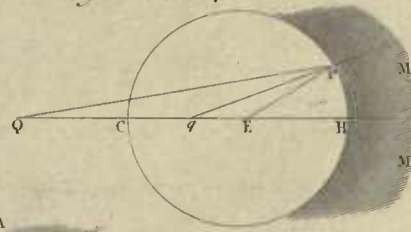


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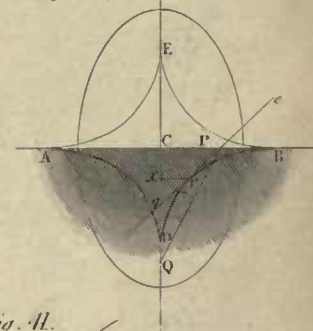


Fig. 38. Art. 236 & 303.



Fig. 40. Art. 238.

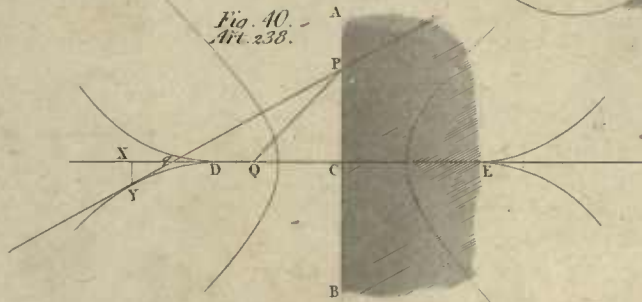


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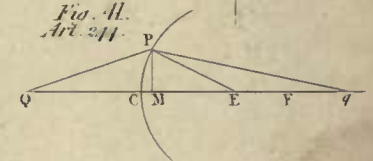


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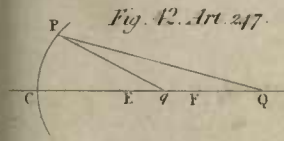
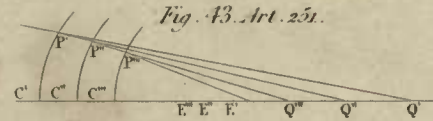


Fig. 43. Art. 251.



Article 259.

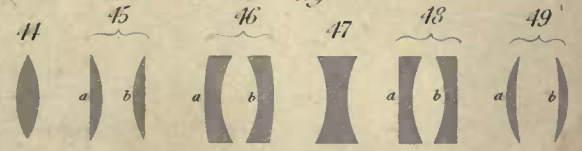


Fig. 50. Art. 269.



Fig. 51. Art. 271.

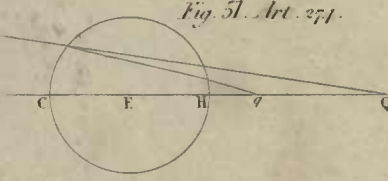


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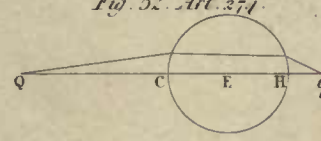


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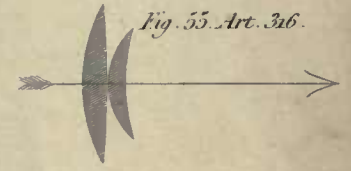


Fig. 53. Art. 287 & 301.

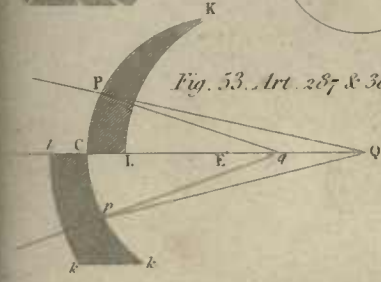


Fig. 54. Art. 289.

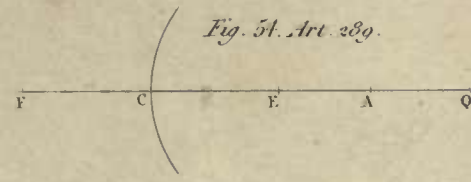


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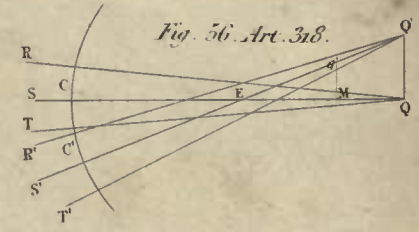


Fig. 57. Art. 321.

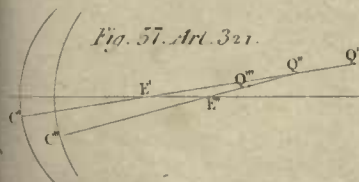


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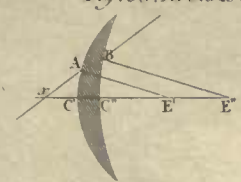


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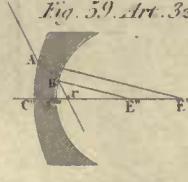


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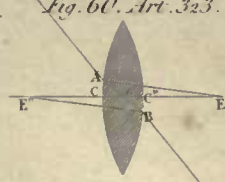


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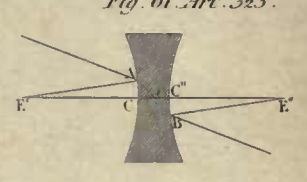


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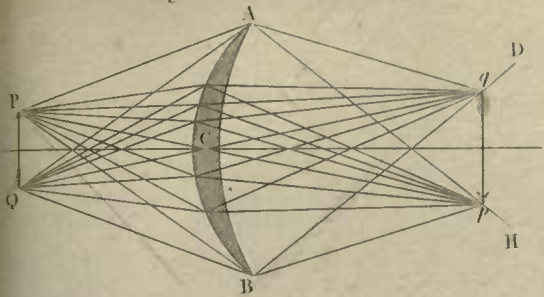


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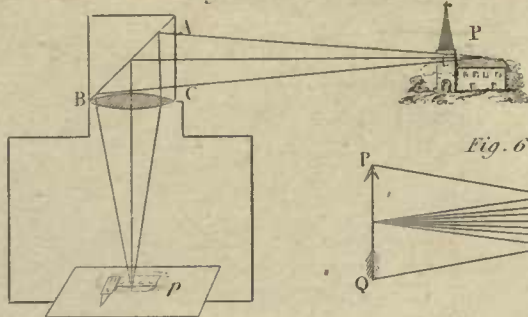


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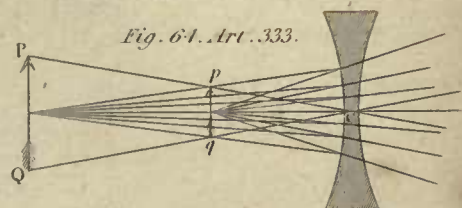


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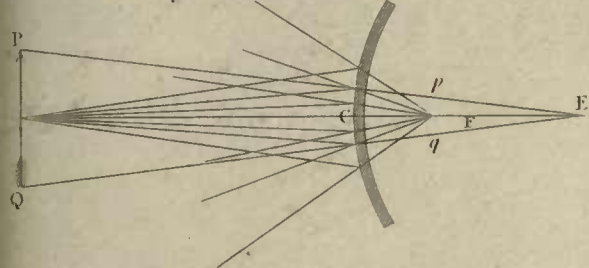


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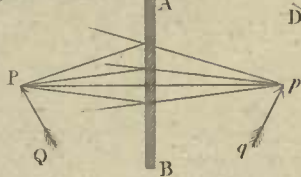


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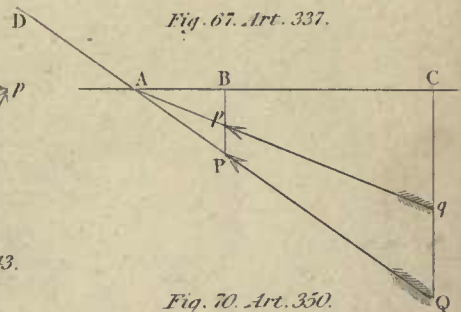


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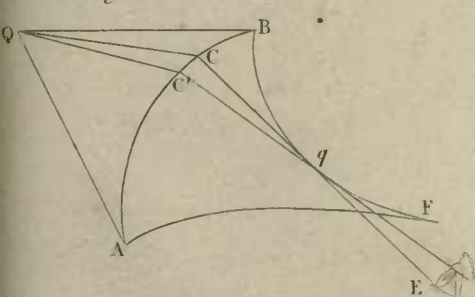


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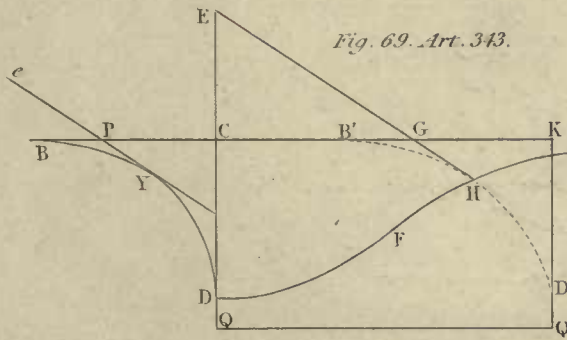


Fig. 70. Art. 350.

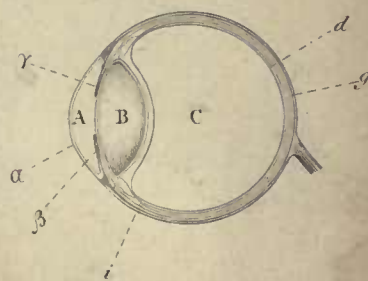


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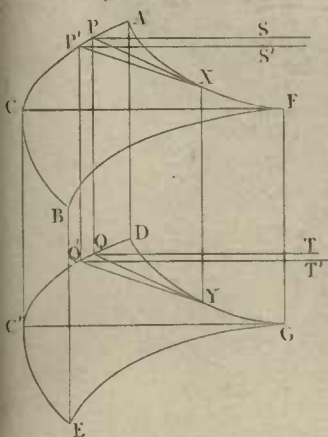


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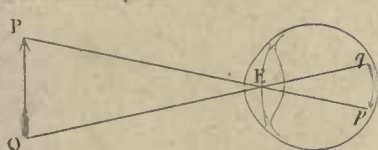


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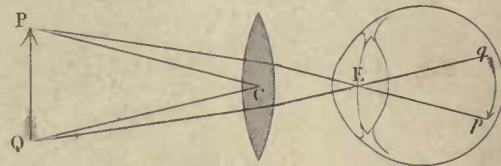


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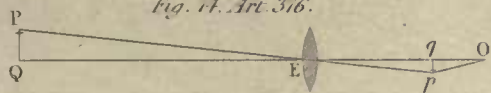


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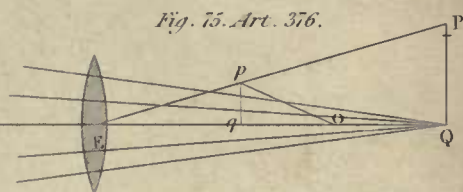


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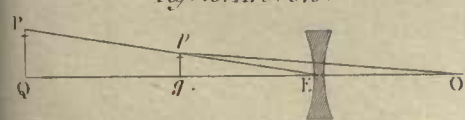


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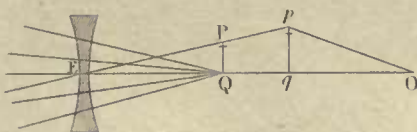


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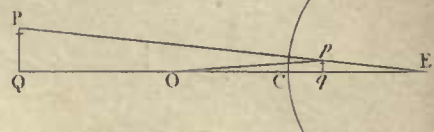


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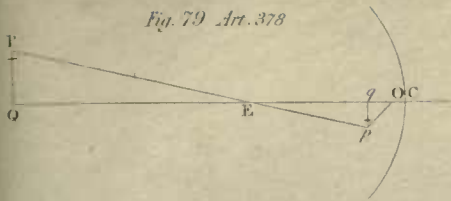


Fig. 80 Art. 381

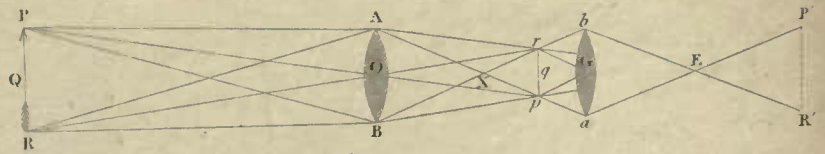


Fig. 81 Art. 387

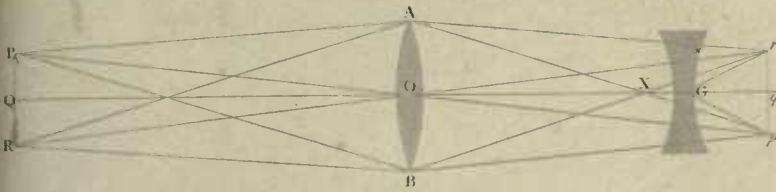


Fig. 82 Art. 389

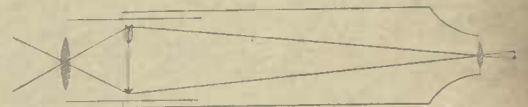


Fig. 83 Art. 391

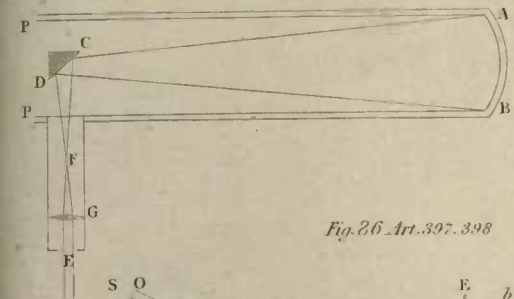


Fig. 84 Art. 392

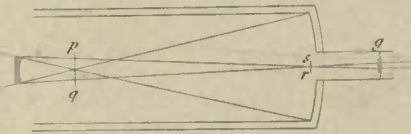


Fig. 85 Art. 395

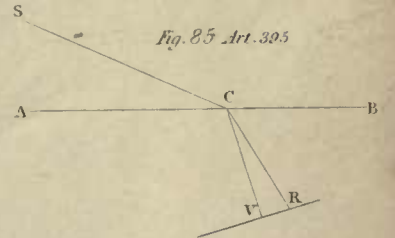


Fig. 86 Art. 397, 398

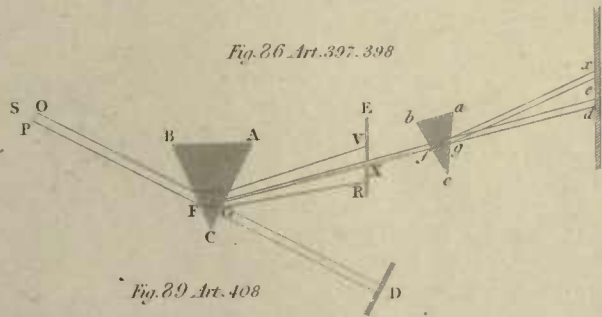


Fig. 87 Art. 406



Fig. 88 Art. 406

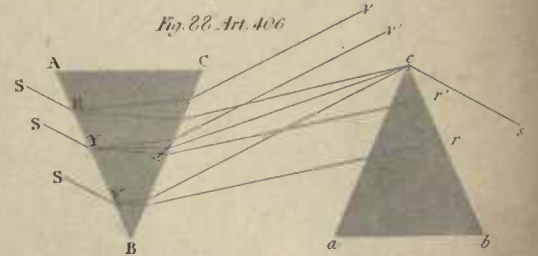


Fig. 89 Art. 408

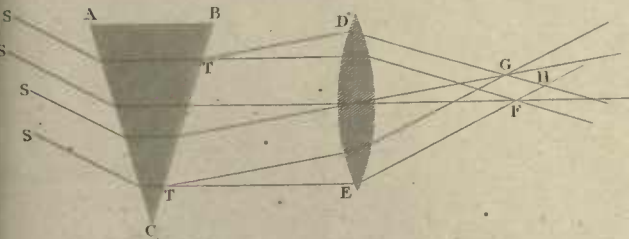


Fig. 89 N° 2 Art. 412

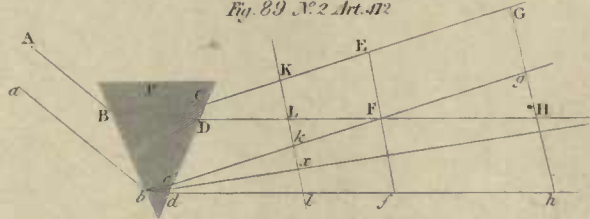


Fig. 92 Art. 415

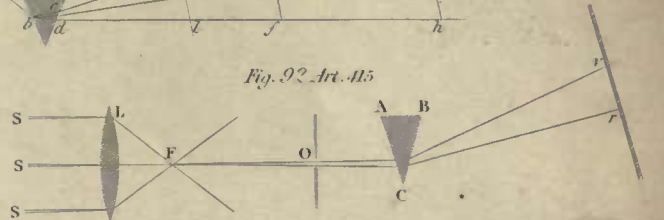


Fig. 90 Art. 413

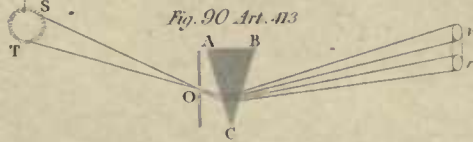


Fig. 93 Art. 416

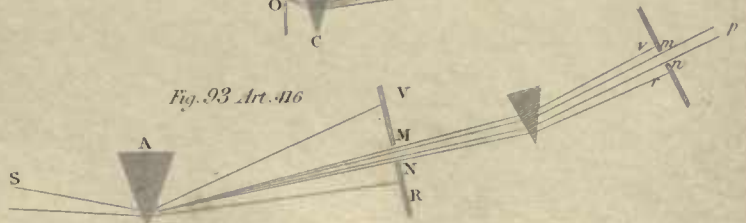
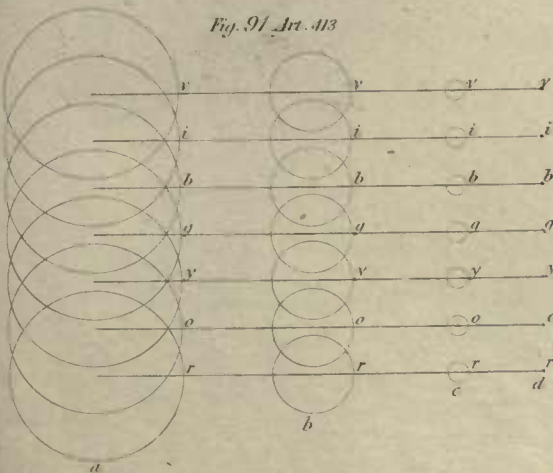


Fig. 91 Art. 413



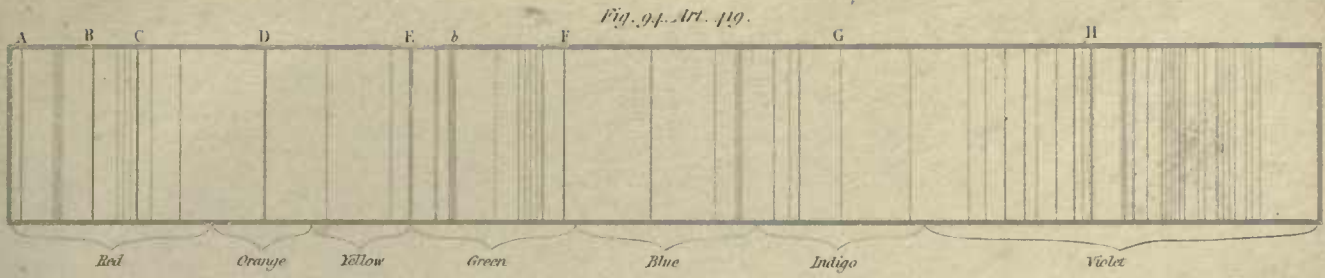


Fig. 95. Art. 422.

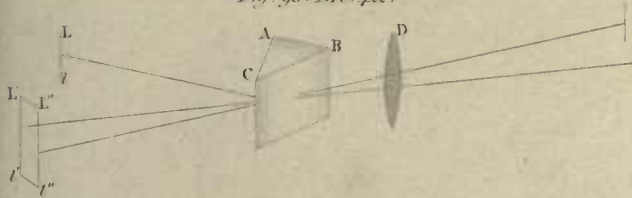


Fig. 96. Art. 423.

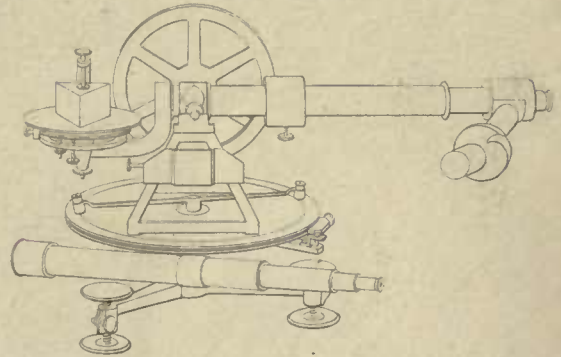


Fig. 97. Art. 426.

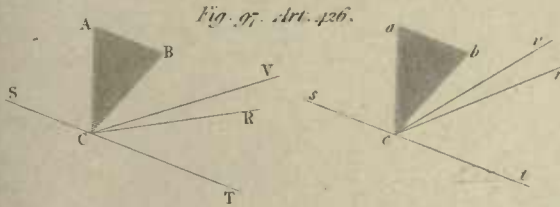


Fig. 98. Art. 426.



Fig. 99. Art. 429.

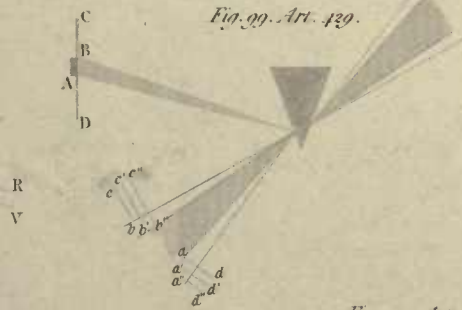


Fig. 100. Art. 431.

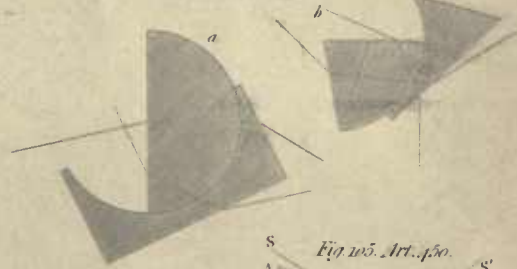


Fig. 101. Art. 432.

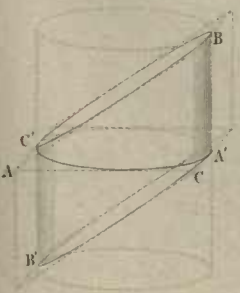


Fig. 103. Art. 434.

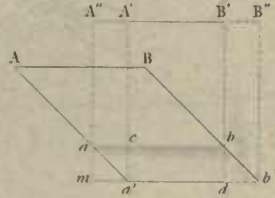


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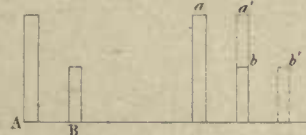


Fig. 106. Art. 451.

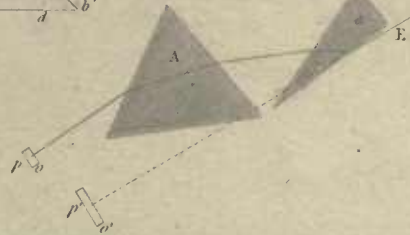


Fig. 108. Art. 470, 472.



Fig. 111. Art. 476.



Fig. 107. Art. 457.

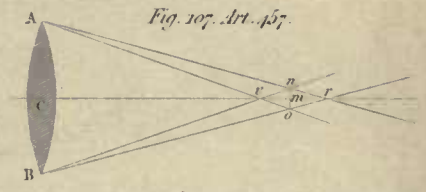


Fig. 109. Art. 475.



Fig. 110. Art. 475.



Fig. 112. Art. 490.

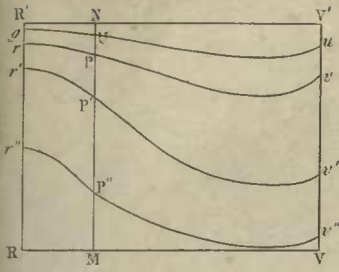


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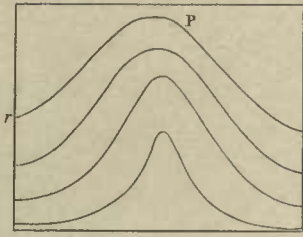


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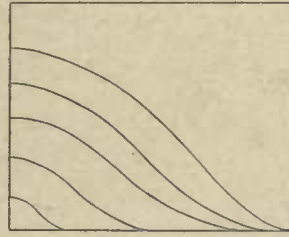


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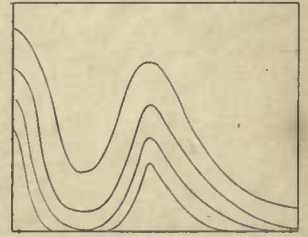


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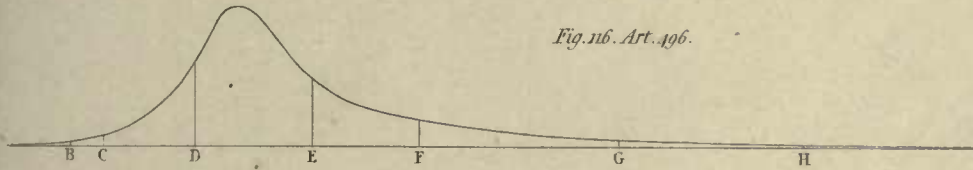


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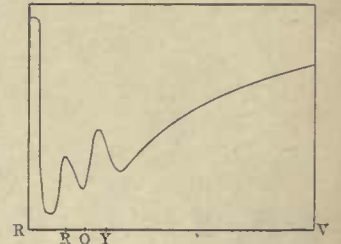


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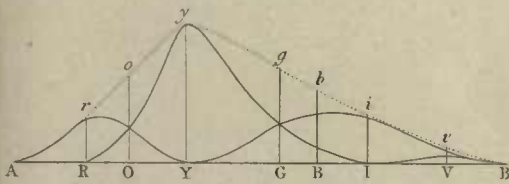


Fig. 119. Art. 541.

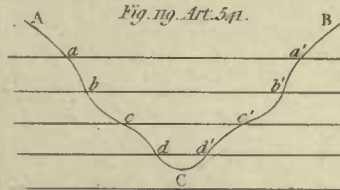


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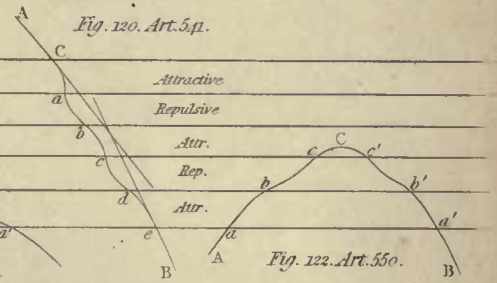


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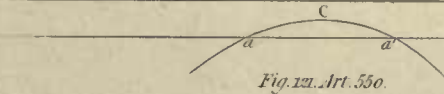


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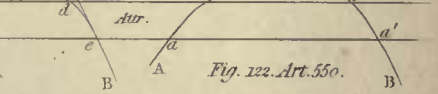


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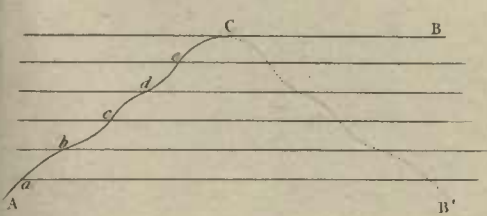


Fig. 124. Art. 554.

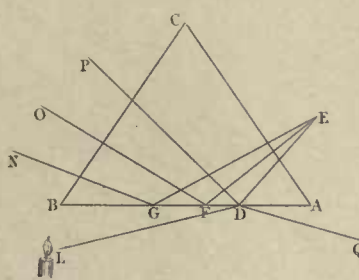


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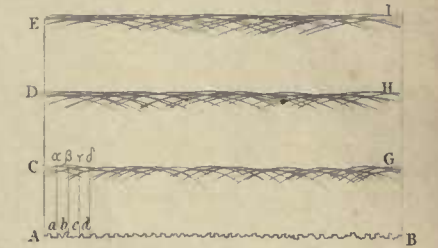


Fig. 126. Art. 579.

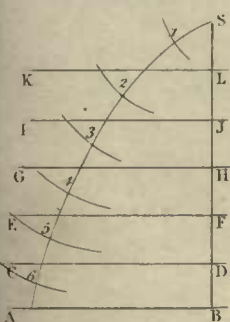


Fig. 127. Art. 584.

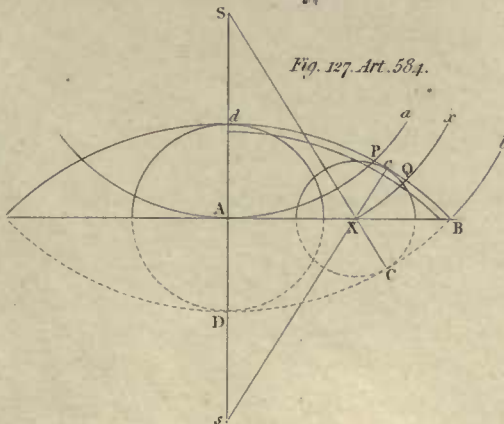


Fig. 128. Art. 585.

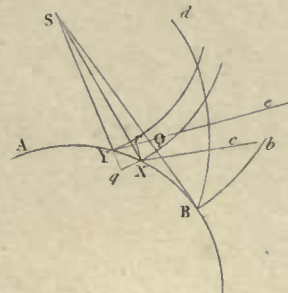
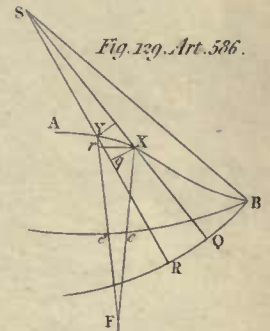


Fig. 129. Art. 586.



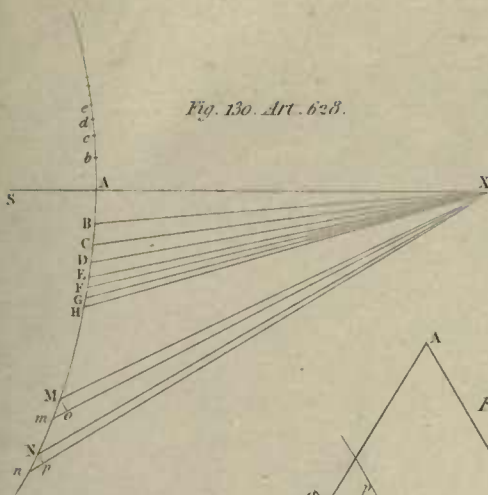


Fig. 130. Art. 628.

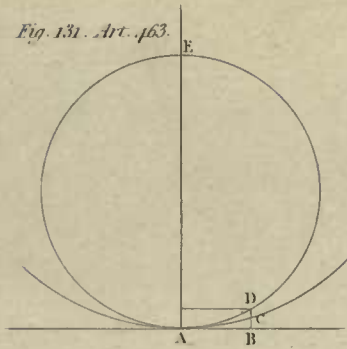


Fig. 131. Art. 463.

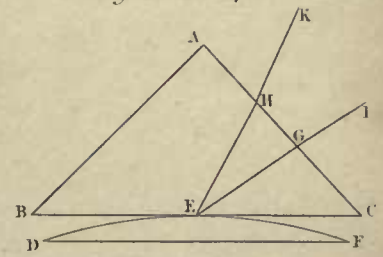


Fig. 132. Art. 640.

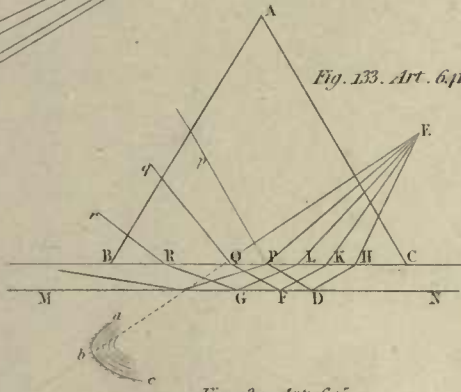


Fig. 133. Art. 641.

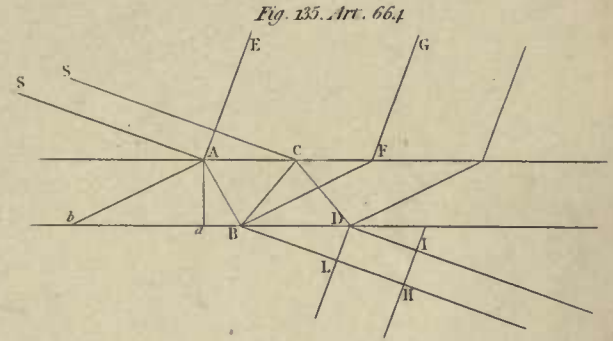


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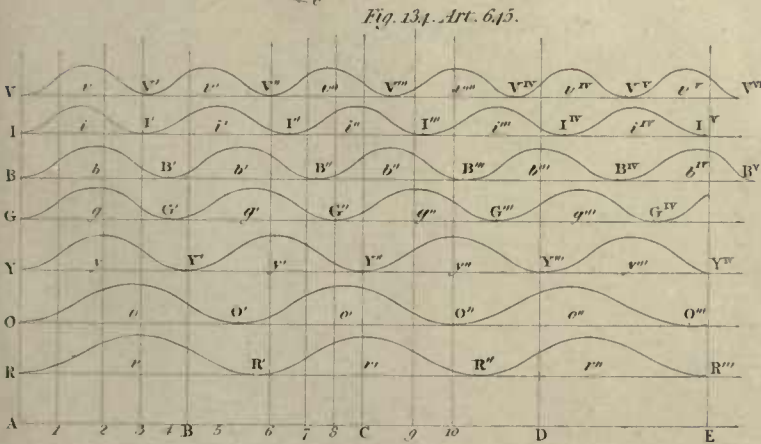


Fig. 134. Art. 645.

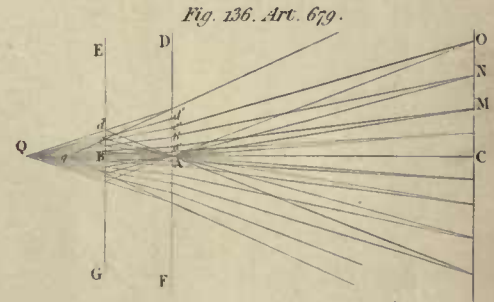


Fig. 136. Art. 679.

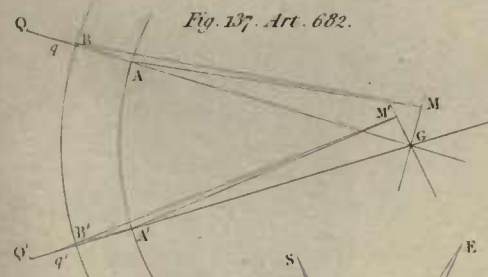


Fig. 137. Art. 682.

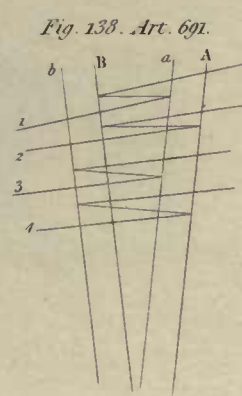


Fig. 138. Art. 691.

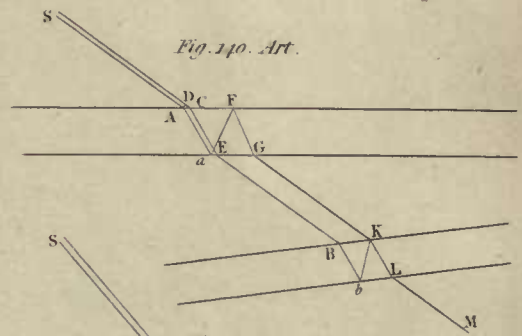


Fig. 140. Art.

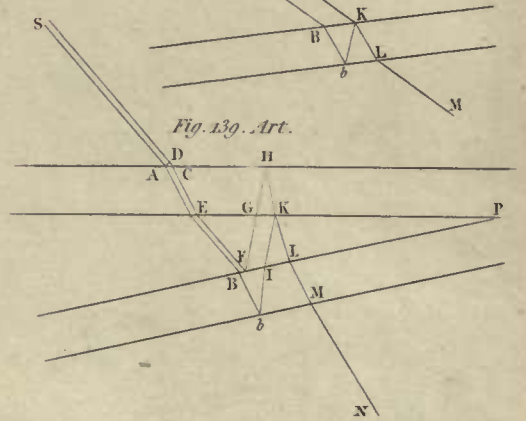


Fig. 139. Art.

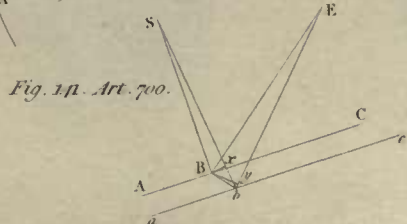


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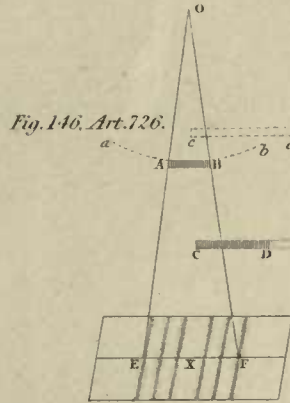
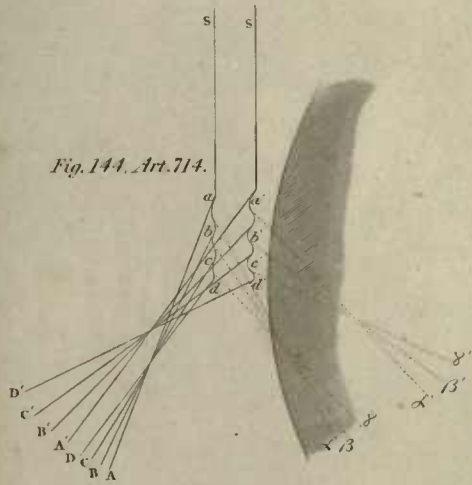
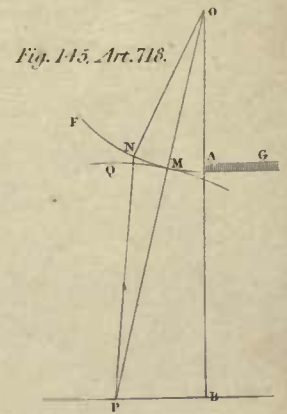
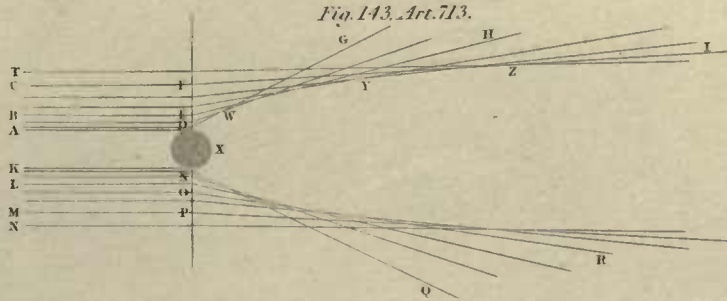
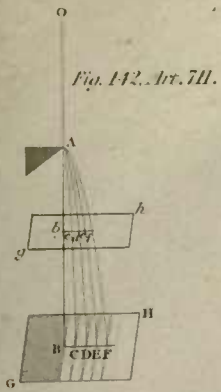


Fig. 147, Art. 735.

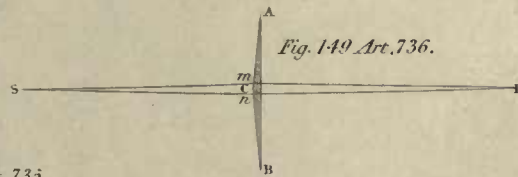
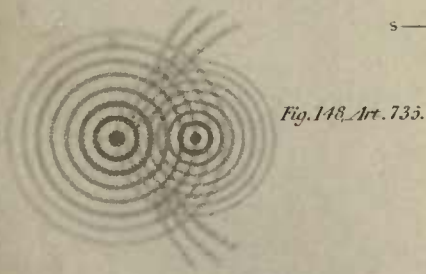


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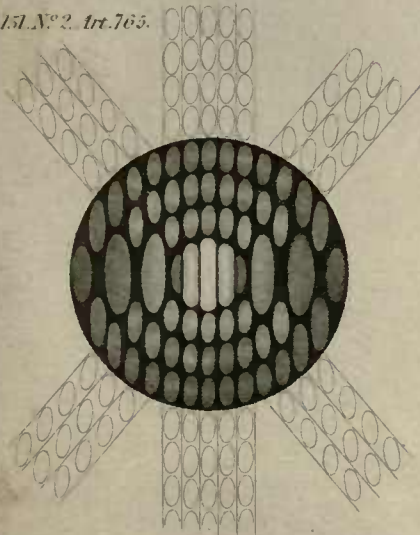


Fig. 153, Art. 770.



Fig. 154, Art. 770.



Fig. 155, Art. 770.



Fig. 156, Art. 770.

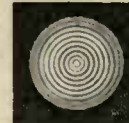


Fig. 157, Art. 770.



Fig. 159. Art. 771.



Fig. 158. Art. 771.

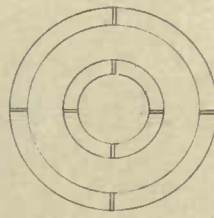


Fig. 151. N^o 3. Art. 742.

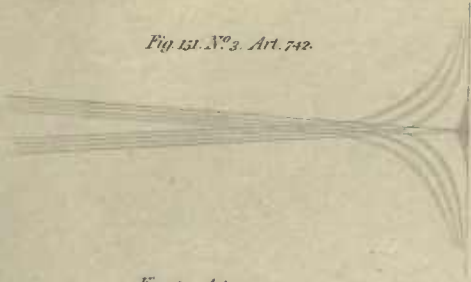


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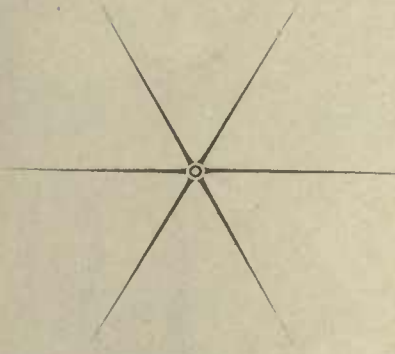


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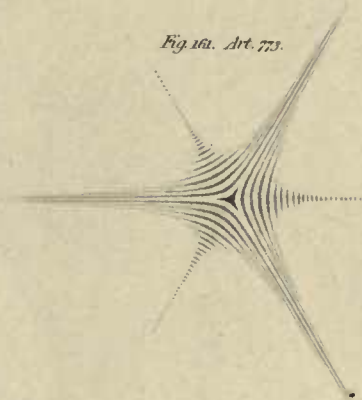


Fig. 162. Art. 775.



Fig. 163. Art. 775.

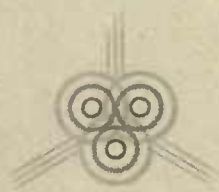


Fig. 164. Art. 775.



Fig. 166. Art. 777.

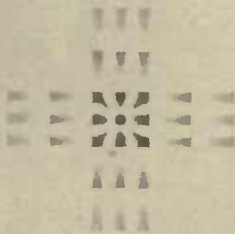


Fig. 167. Art. 778.

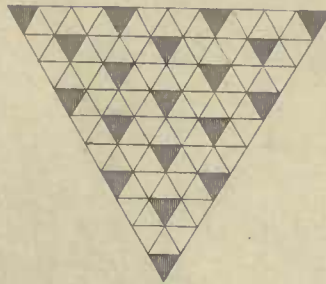


Fig. 165. Art. 776.

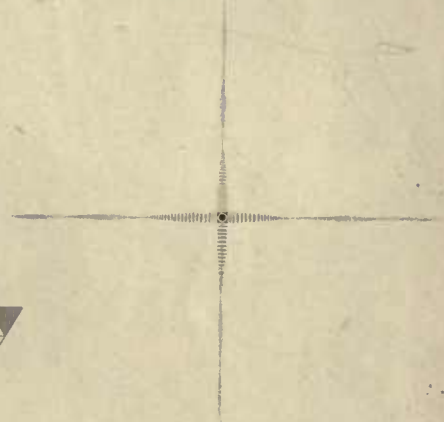


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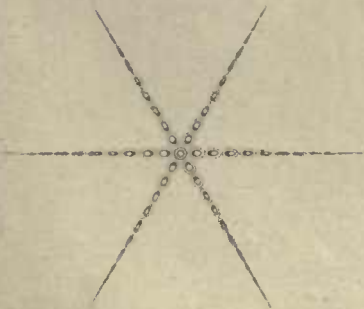


Fig. 170. Art. 836.

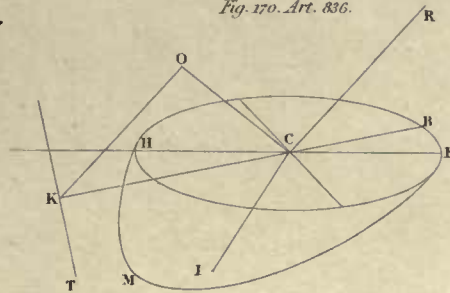


Fig. 169. Art. 790.

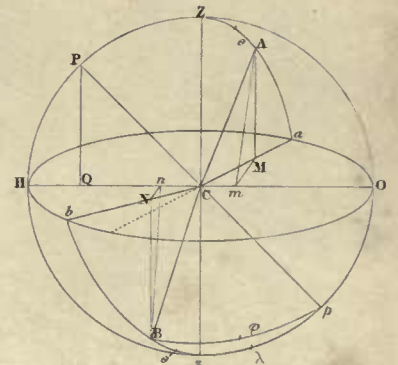


Fig. 171. Art. 825.

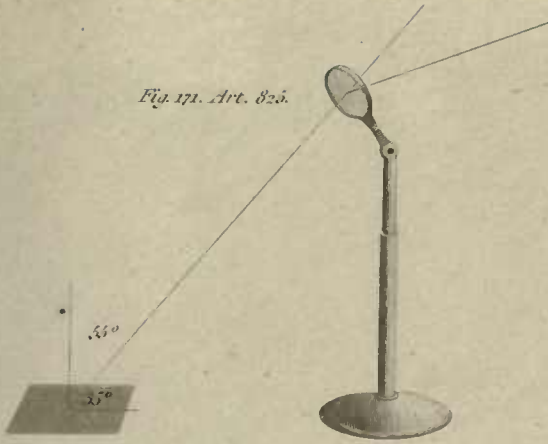


Fig. 172. Art. 826.

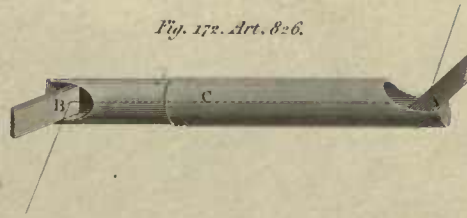


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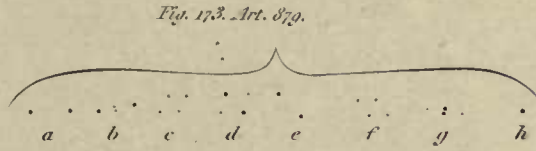


Fig. 174. Art. 881.

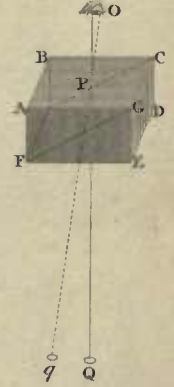


Fig. 175. Art. 882.

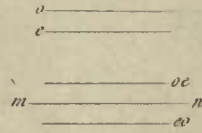
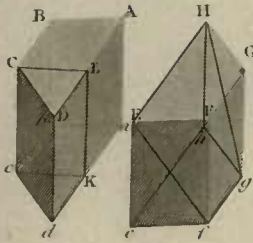


Fig. 176. Art. 892.



Fig. 177. Art. 893.

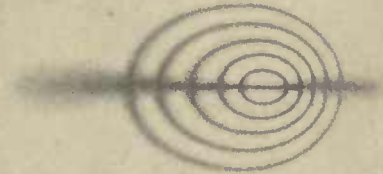


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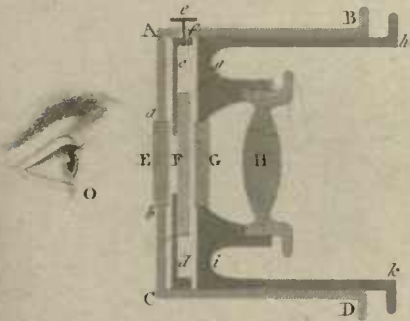


Fig. 179. Art. 900.



Fig. 180. Art. 901.



Fig. 181. Art. 900.



Fig. 182. Art. 900.



Fig. 183. Art. 902.

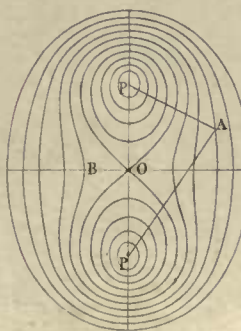


Fig. 184. Art. 910.



Fig. 185. Art. 901.

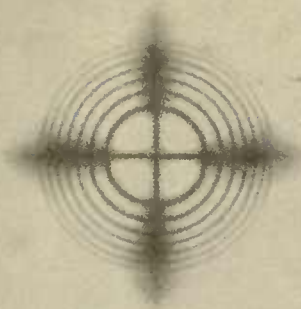


Fig. 186. Art. 926.



Fig. 191. Art. 935.



Fig. 186. Art. 922.



Fig. 187. Art. 922.

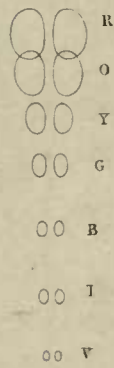


Fig. 189. Art. 929.

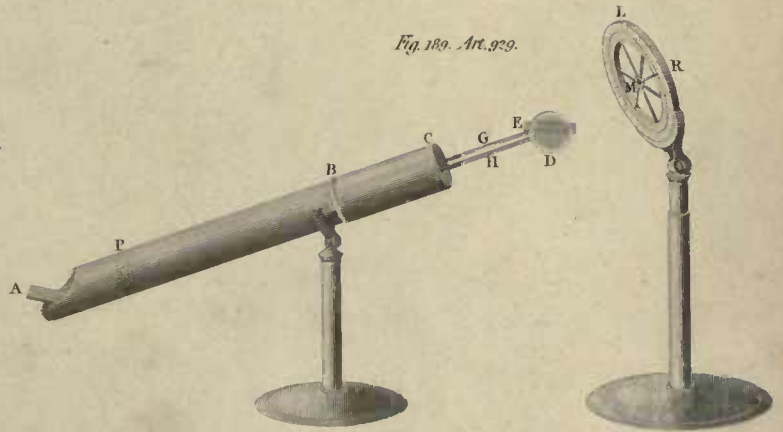


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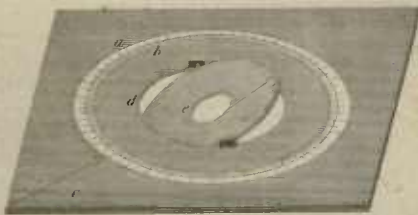


Fig. 192. Art. 939.



Fig. 193. Art. 939.



Fig. 194. Art. 969.

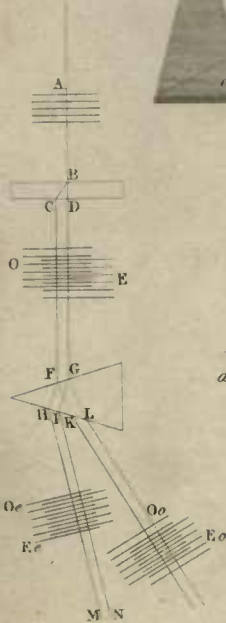


Fig. 195. a.



Fig. 196. Art. 986.

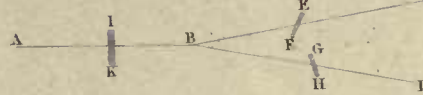


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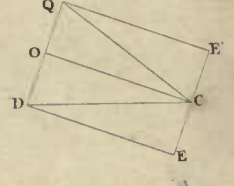


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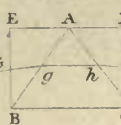


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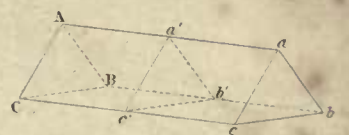


Fig. 199 Art. 1030

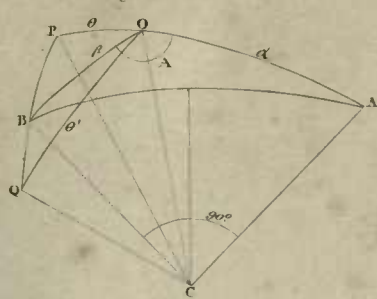


Fig. 200 Art. 1032

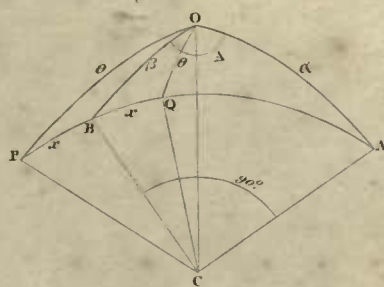


Fig. 202 Art. 10.42

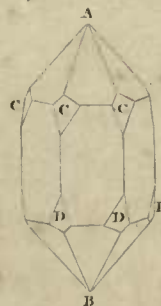


Fig. 202 Art. 1041

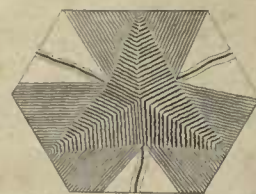


Fig. 203 Art. 10, 18

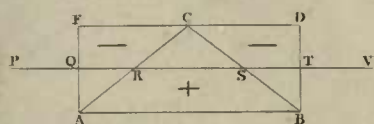


Fig. 204 Art 1050

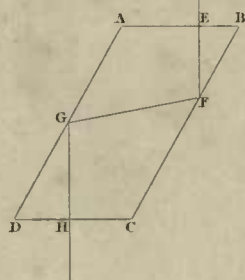


Fig. 205 Art. 1057

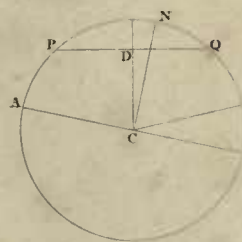


Fig. 206 Art 1067.



Fig. 207 Att 1068



Aug. 208 Art. 1071

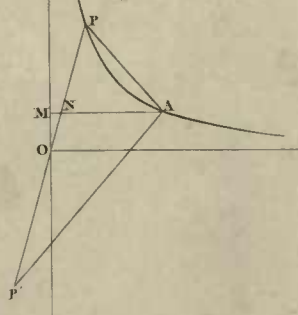


Fig. 209 Art 2170

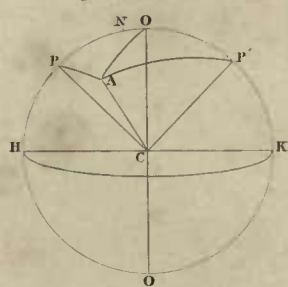


Fig. 210 Ar. 1081 & 1079.

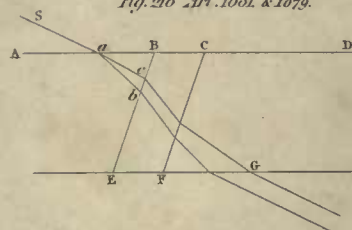


Fig. 2n Art 1082

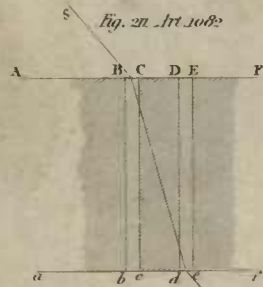


Fig. 222 Art. 1090



Fig. 23 Art. 1096

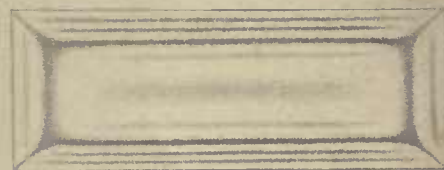


Fig. 26 Art. 1105

Fig. 215 Art. 1099

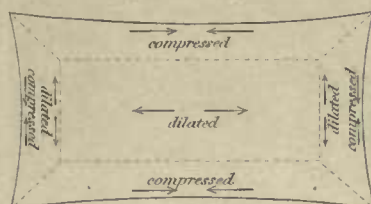
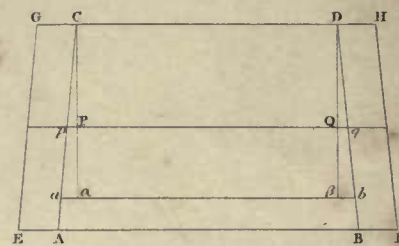


Fig. 24 Art. 1098



LIGHT.

Fig. 217 Art. 1105

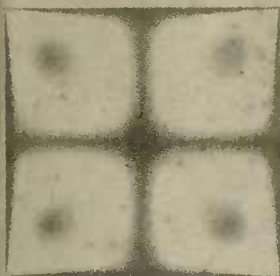


Fig. 219 Art. 1105

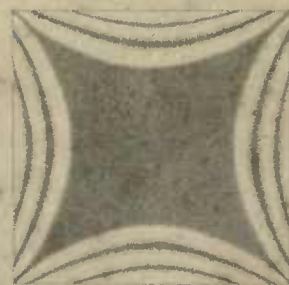


Fig. 221 Art. 1105

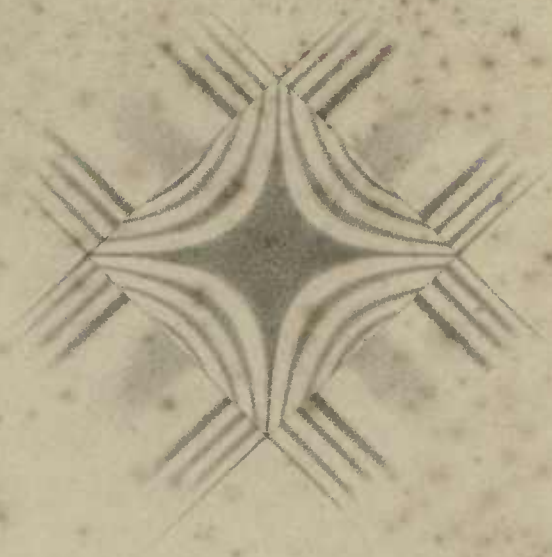


Fig. 218 Art. 1105

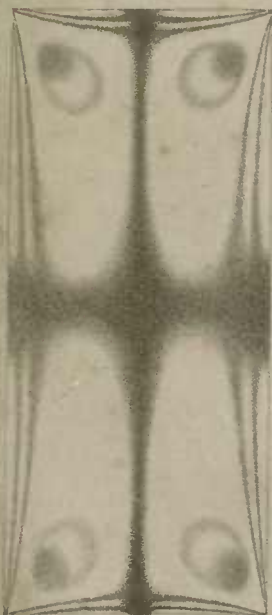


Fig. 220 Art. 1105



Fig. 222 Art. 1106

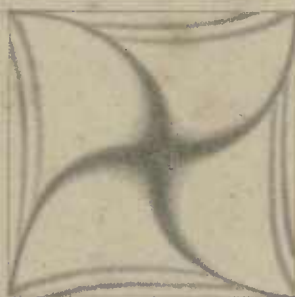


Fig. 223

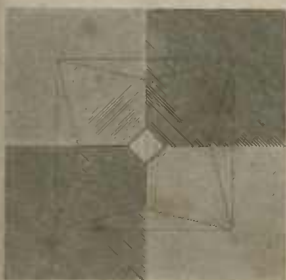
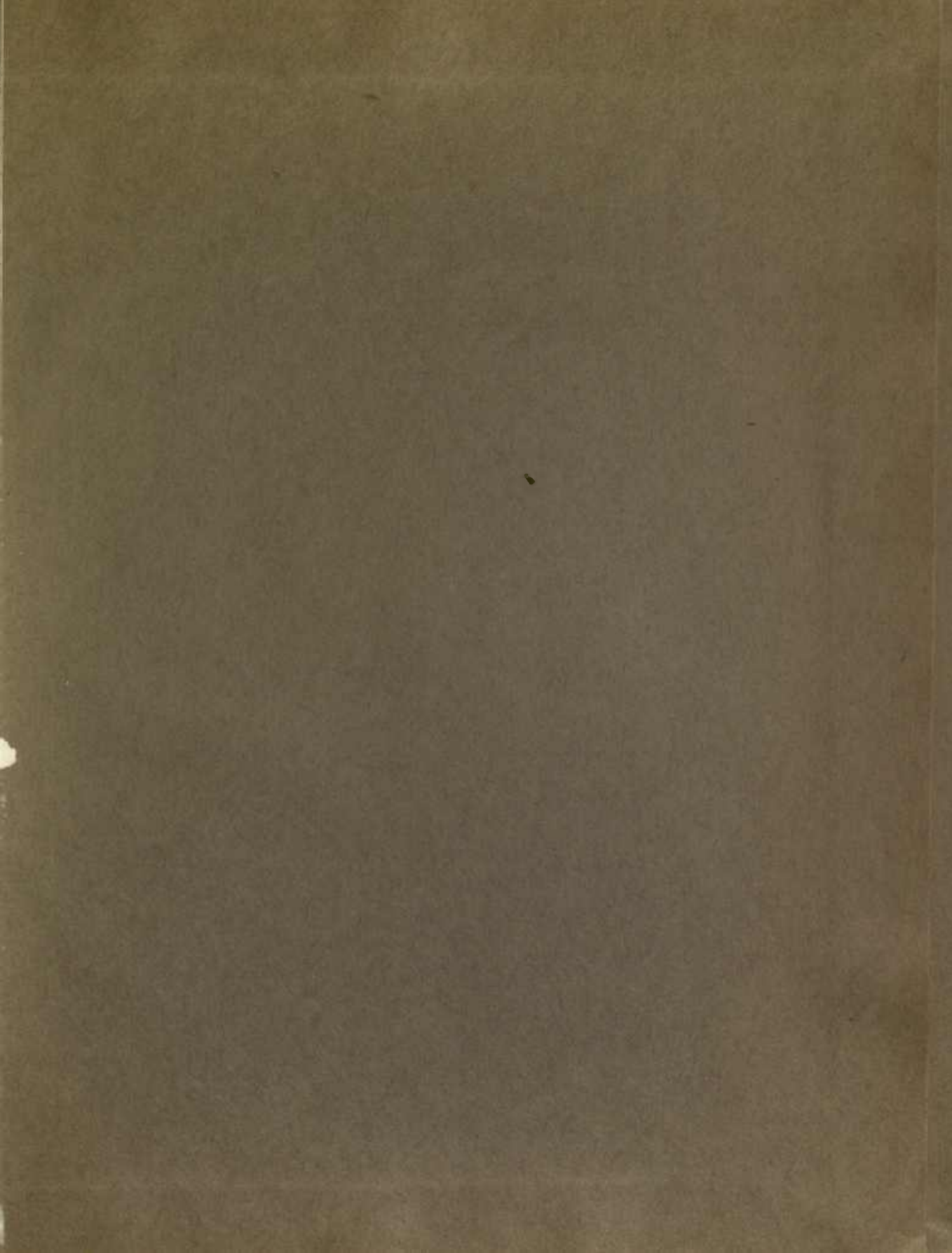


Fig. 224





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